Intro to collider physics

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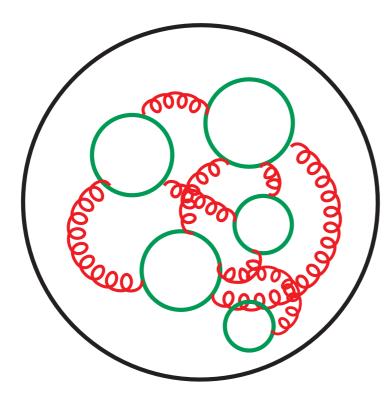
Before we start

- This is a huge subject.
 - Focus more on intuitive understanding, generic feature, less on specifics.
 - Only a (small) subset.
- Focus on methodology, rather than specific models.

Hopefully, this serves as the starting point of your further study.

Many good references, such as Tao Han, TASI lecture, hep-ph/0508097



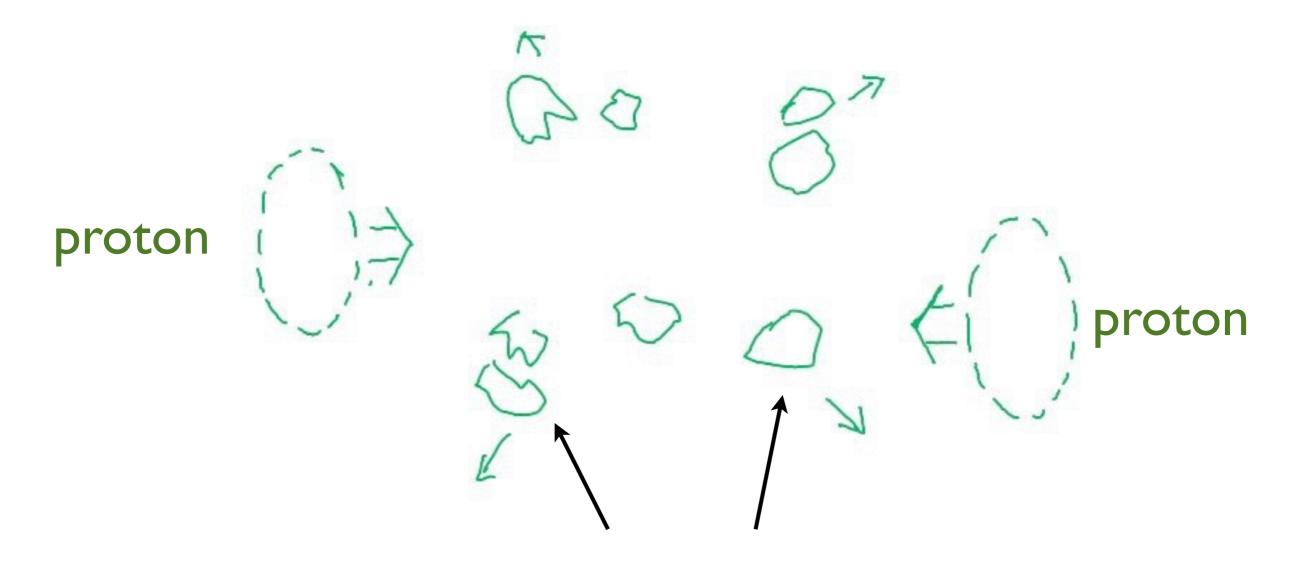




gluon valence: u, d "sea": qbar, s sbar, c, cbar, b, bbar

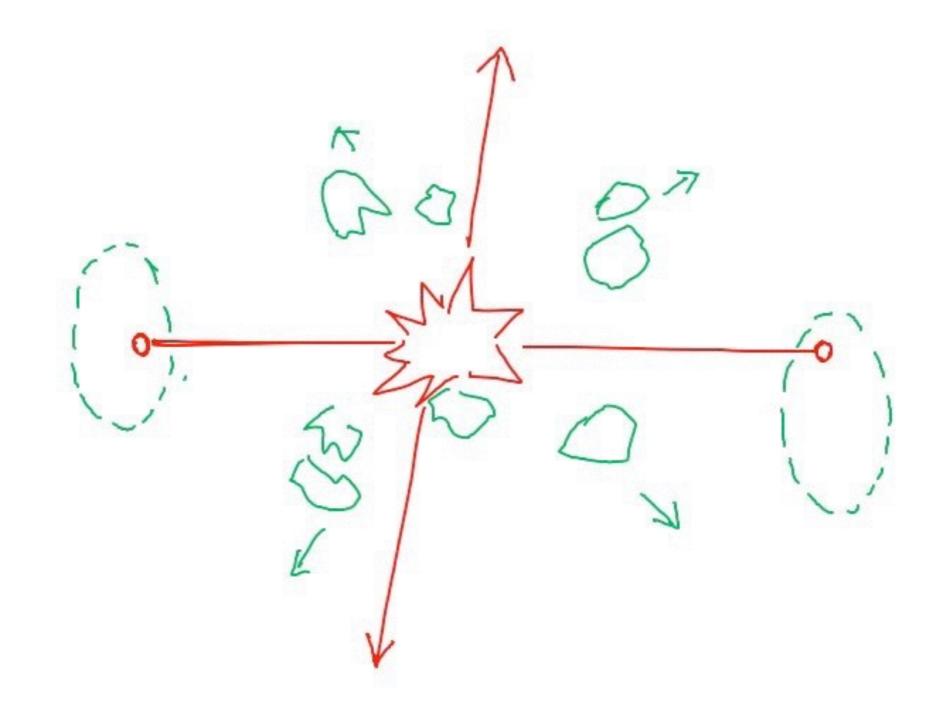
binding energy ~ GeV

Most of the time

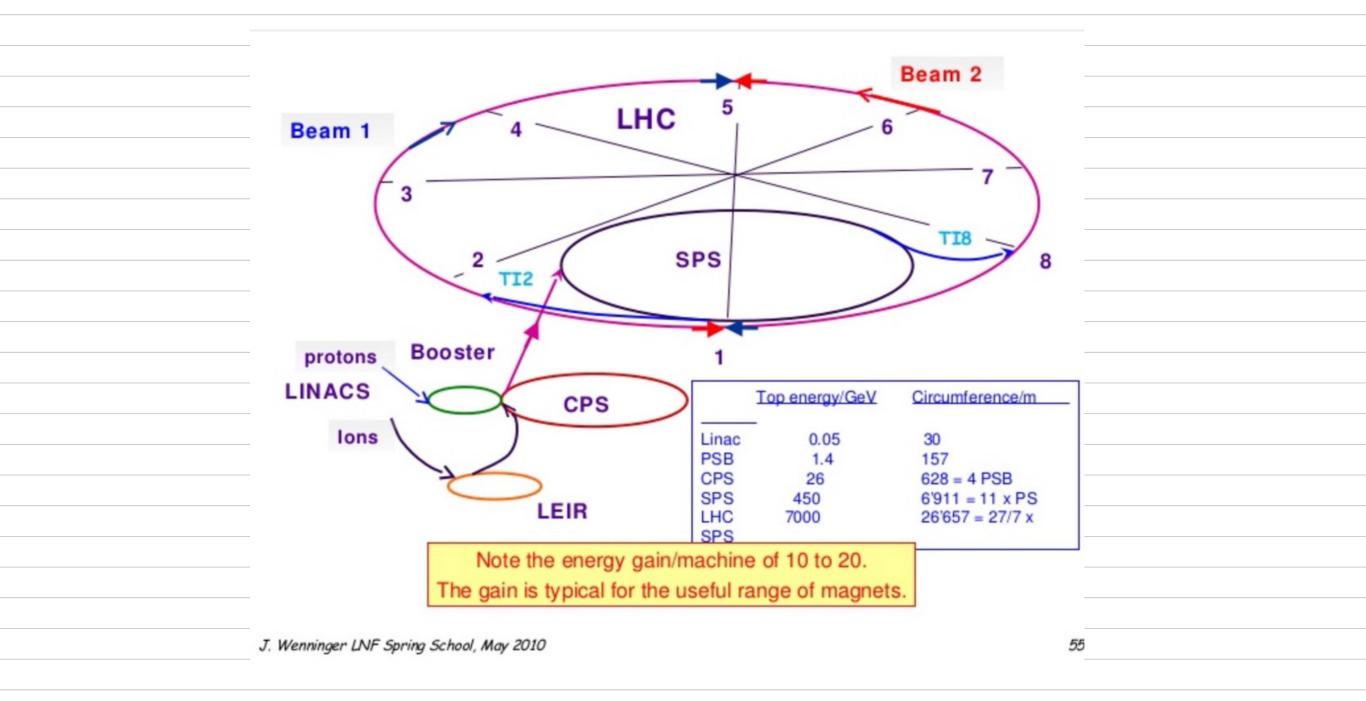


low energy fragments: $\rm E \sim GeV$

High energy collision rare

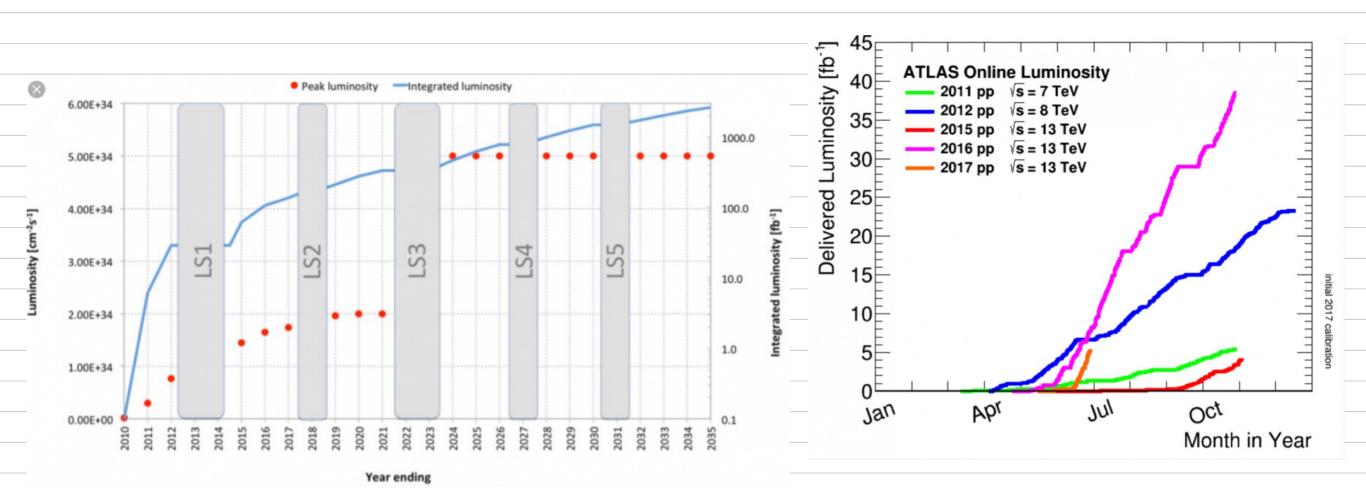


The Large Hadron Collider (LHC)



Luminosity $L \times K N^2 f/a$ f: revolution freq f~11.25 KHZ N: # of protons in a bunch. N~10" K: # of bunches K=2808 a: bean size Larea) A ~ TT (16 µm)² T. cross section. # events •

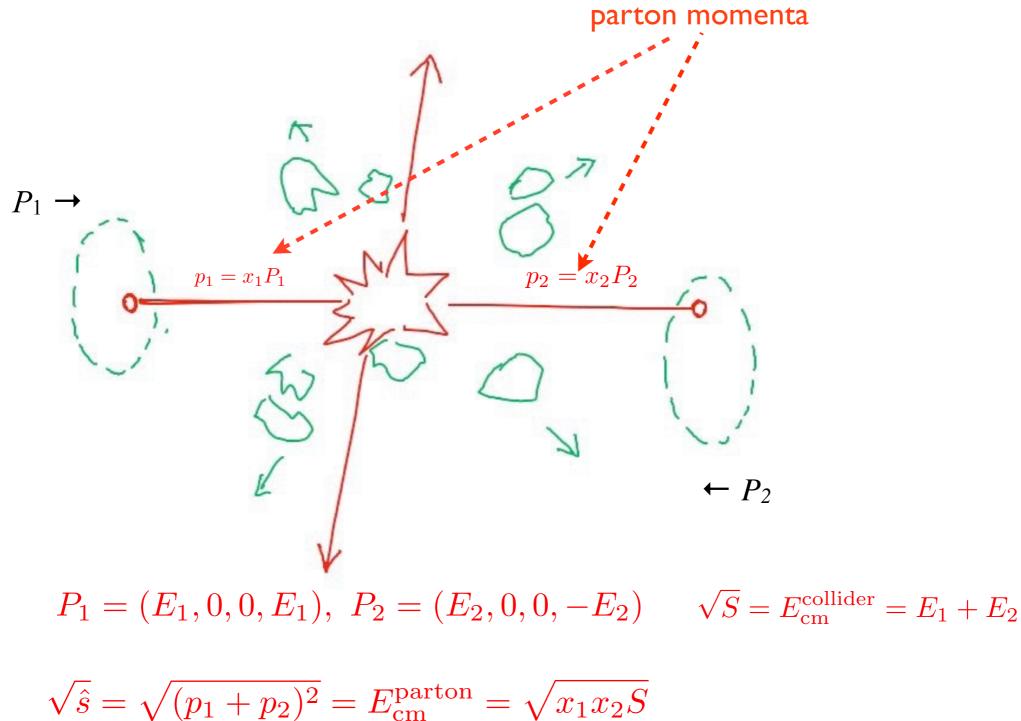
LHC Luminosity



$$|mb = 10^{-27} cm^2 = 2.56 (GeV)^{-2}$$

 $|0^{34} cm^{-2} s^{-1} \sim 100 Fb^{-1} / Yr$

Kinematics



Rapidity

Define rapidity

$$y = \frac{1}{2} \ln \frac{E + p_z}{E - p_z}$$
$$p^{\mu} = (E_T \cosh y, p_T \sin \phi, p_T \cos \phi, E_T \sinh y), \quad E_T = \sqrt{p_T^2 + m^2}$$

Under boost along z-direction

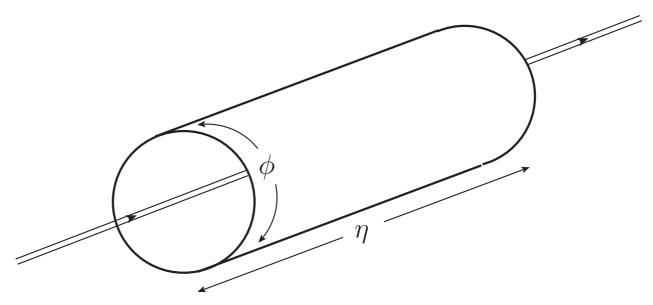
$$y' = \frac{1}{2} \ln \frac{E' + p'_z}{E' - p'_z} = \frac{1}{2} \ln \frac{(1 - \beta_0)(E + p_z)}{(1 + \beta_0)(E - p_z)} = y - y_0$$
$$\to \frac{d}{dy} = \frac{d}{dy'}$$

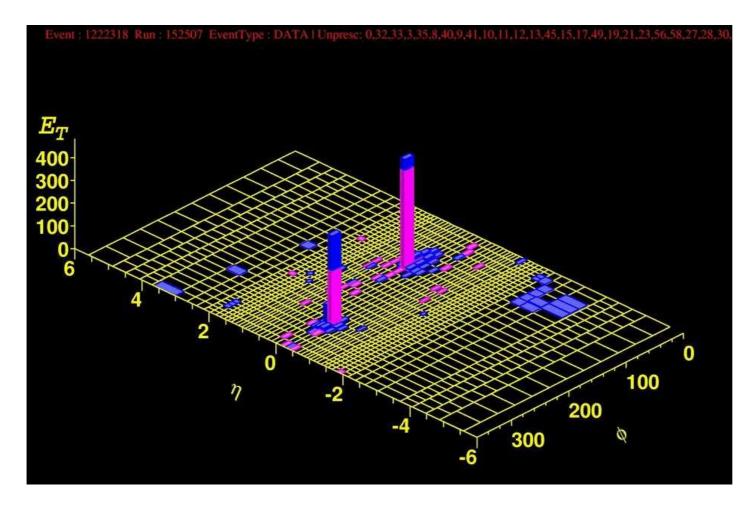
In the massless limit : pseudo-rapidity

$$y \to \frac{1}{2} \ln \frac{1 + \cos \theta}{1 - \cos \theta} = \ln \cot \frac{\theta}{2} \equiv \eta$$

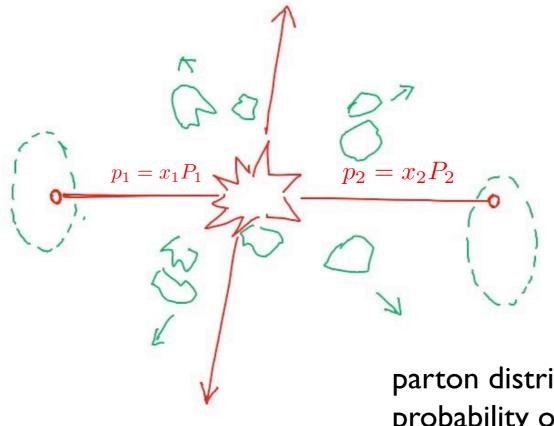
Coordinate System

$$\eta = -\ln\left[\cot\left(\frac{\theta}{2}\right)\right]$$





Parton Distribution Function (PDF)

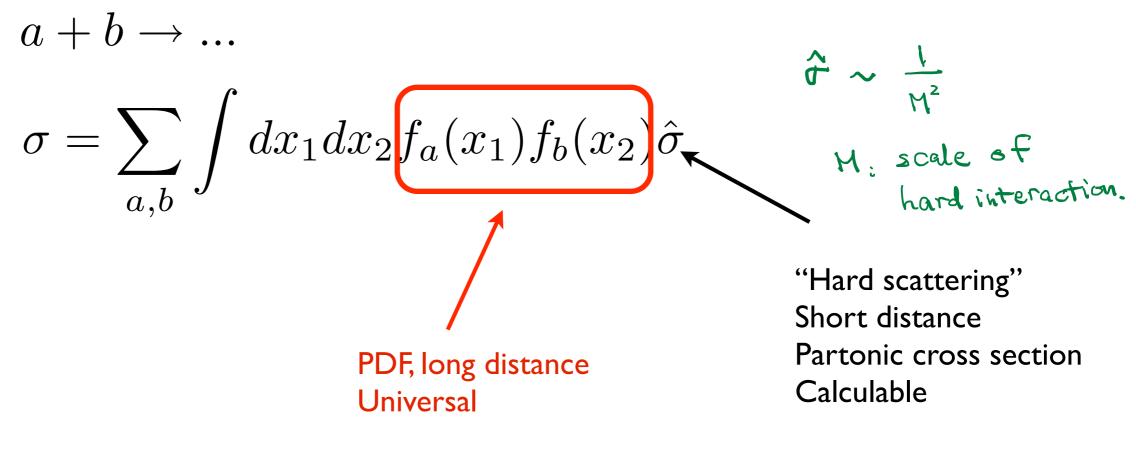


Partons can be gluon, or different flavors of quarks, labelled by a, b...

parton distribution function $f_a(x)$: probability of finding parton a with momentum fraction x

- $f_a(x)$ can not be computed.
- However, we can measure them using certain processes.
- They are universal! Can be used everywhere!

Prediction for hadron collisions



Factorization!

Intuitively, make sense:

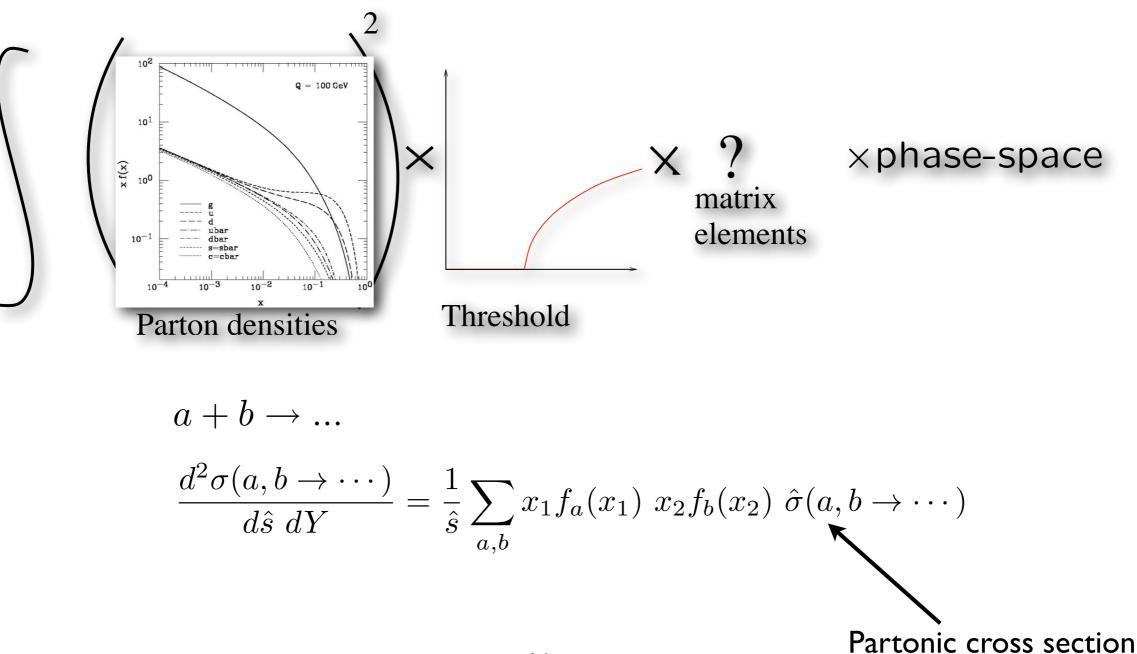
short distance physics should not "know" about long distance physics.

In practice, very difficult to prove.

However, it is used anyway (otherwise we cannot calculate anything). And, it works very well.

Production.

- Schematics of production at hadron colliders.
 - Dominated by parton densities and thresholds (mass and cut).



A useful representation

$$P_1 = (E, 0, 0, E), P_2 = (E, 0, 0, -E)$$
 $p_1 = x_1 P_1, p_2 = x_2 P_2$

Define Parton center of mass rapidity: $Y e^Y = \sqrt{\frac{x_1}{x_2}}$

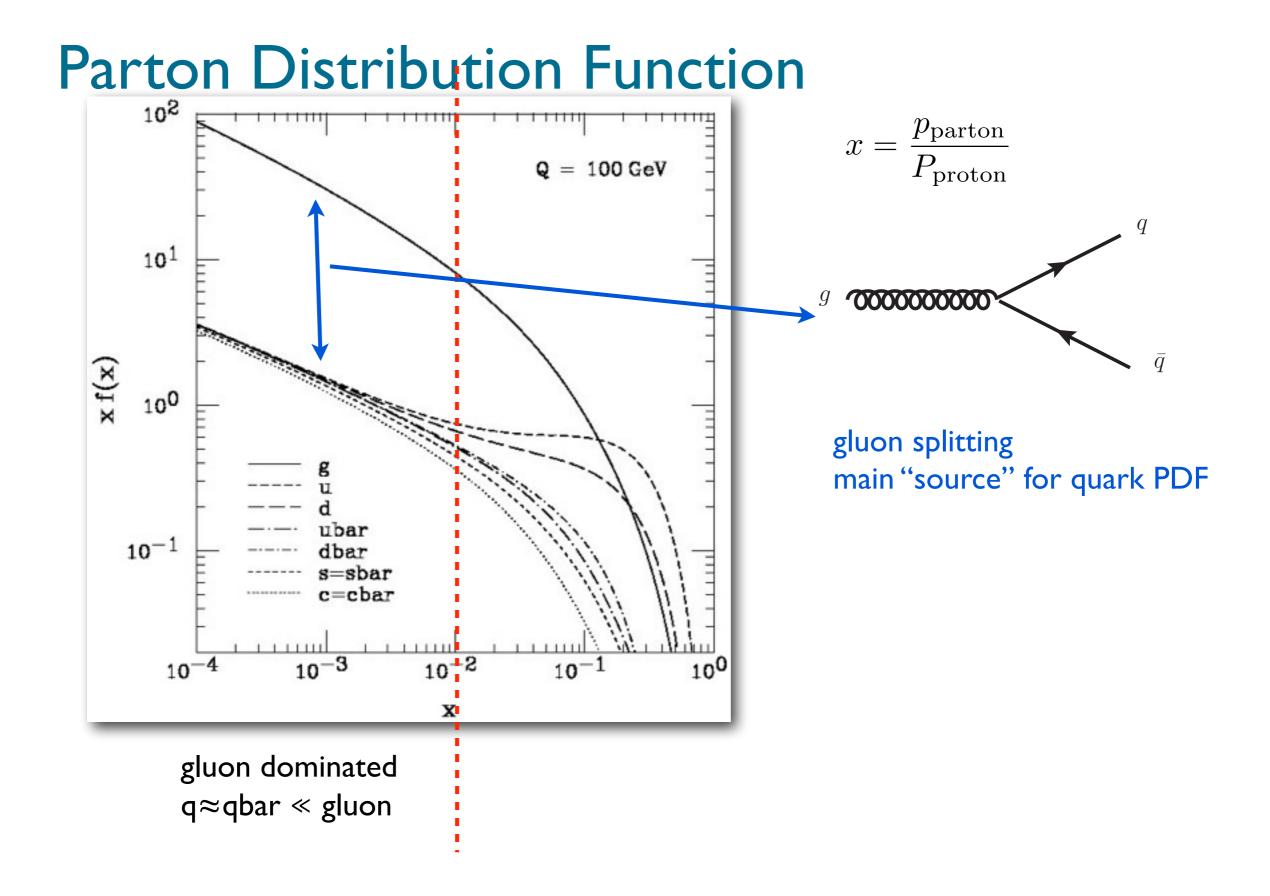
We can verify
$$\cosh Y = \frac{(x_1 + x_2)E}{\sqrt{\hat{s}}} \implies \text{boost of parton c.o.m frame}$$

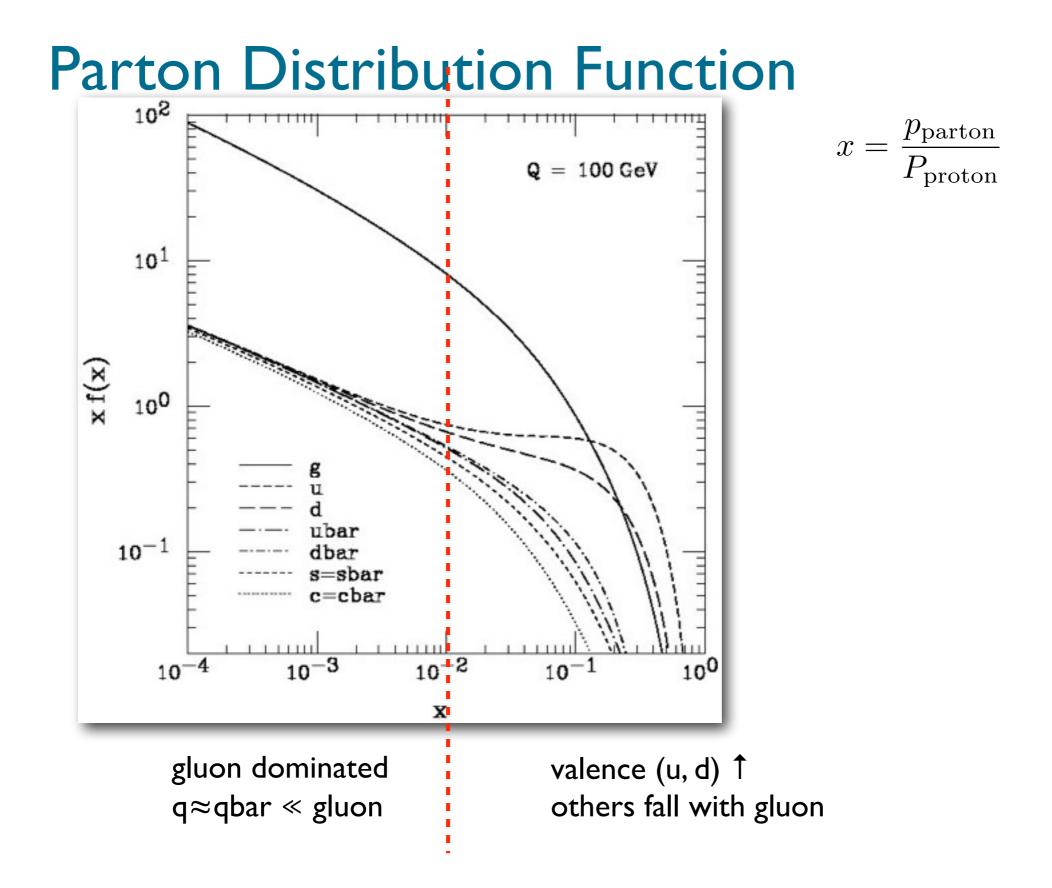
Starting with
$$\frac{d^2\sigma(a,b\to\cdots)}{dx_1dx_2} = \sum_{a,b} f_a(x_1)f_b(x_2)\hat{\sigma}(a,b\to\cdots)$$

Using Jacobian:
$$\frac{\partial |\hat{s}, Y|}{\partial |x_1, x_2|} = \frac{\hat{s}}{x_1 x_2}$$

We obtain:

$$\frac{d^2\sigma(a,b\to\cdots)}{d\hat{s}\ dY} = \frac{1}{\hat{s}}\sum_{a,b} x_1 f_a(x_1) x_2 f_b(x_2) \ \hat{\sigma}(a,b\to\cdots)$$





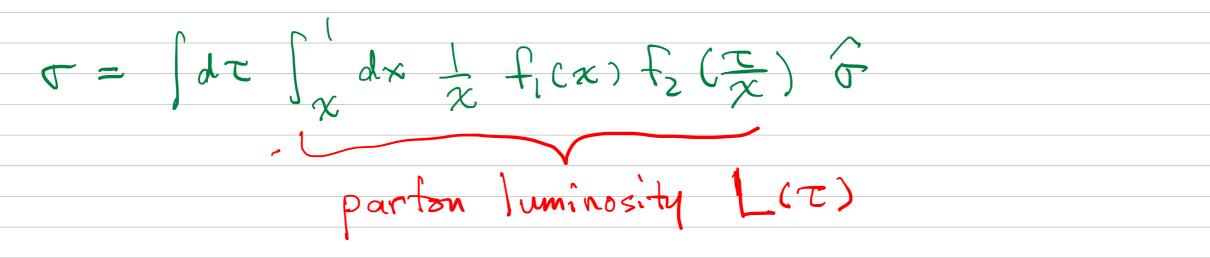
Parton luminosity

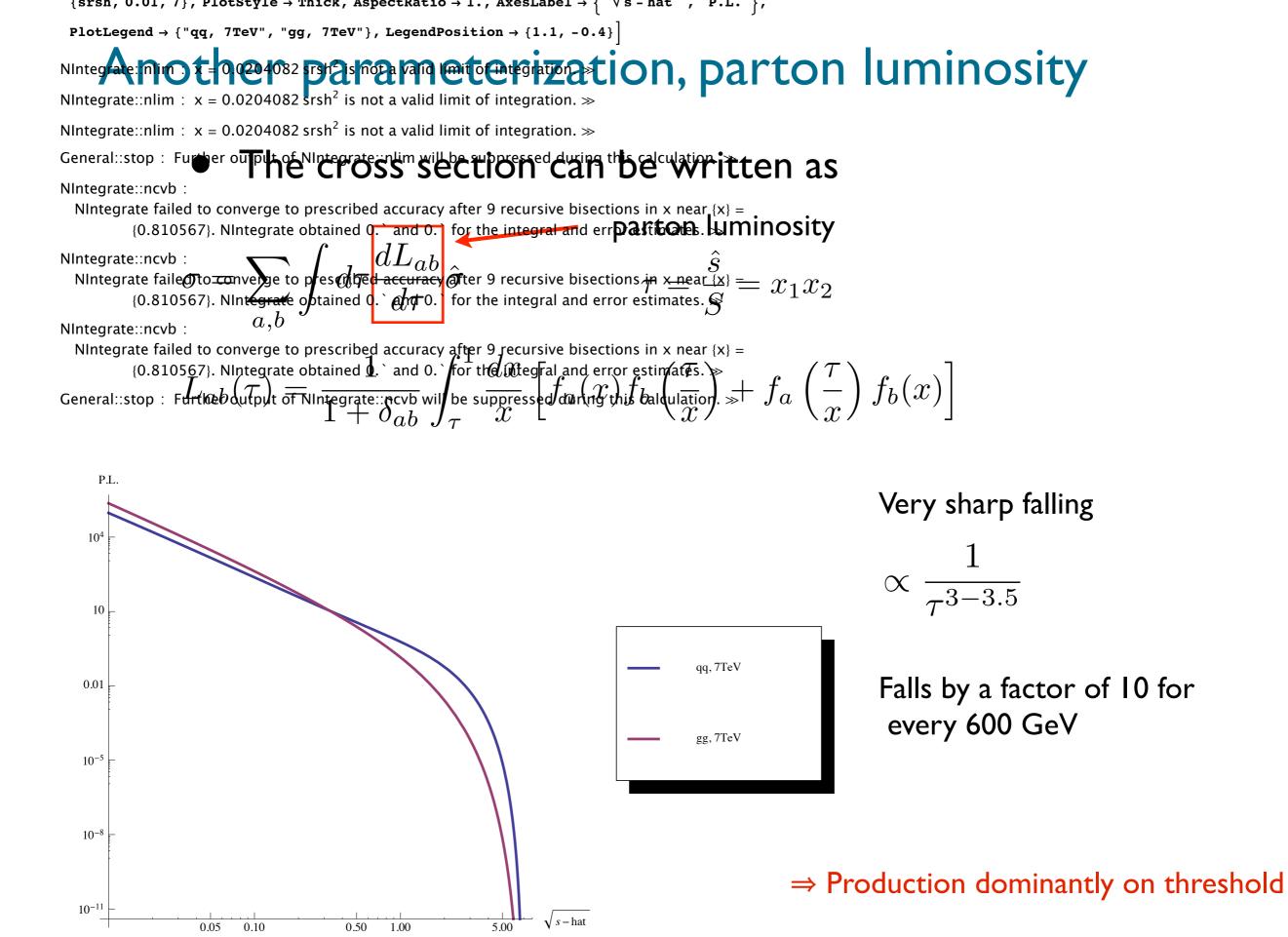
 $= \int dx_1 \int dx_2 f(x_1) f_2(x_2) f$

define $\tau = \chi_1 \chi_2 \qquad \chi = \chi_1$ 5 = 7

Jacobian for variable change

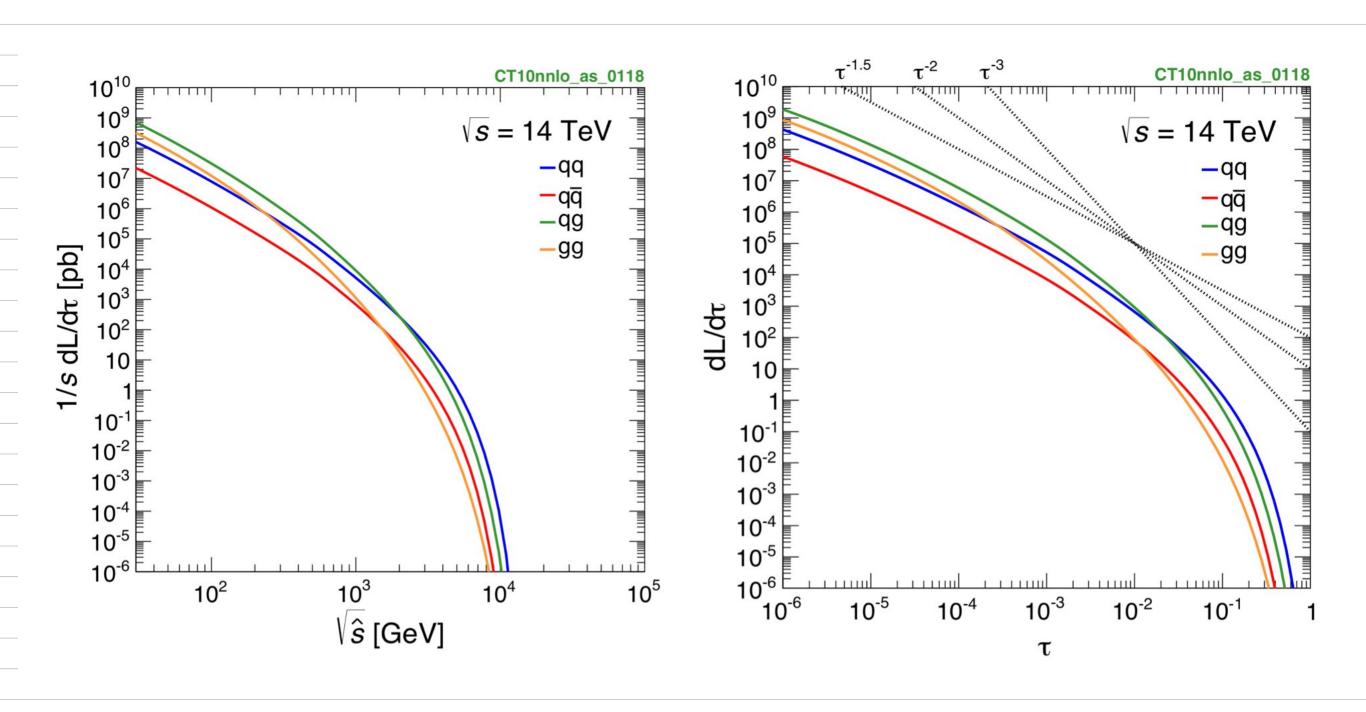


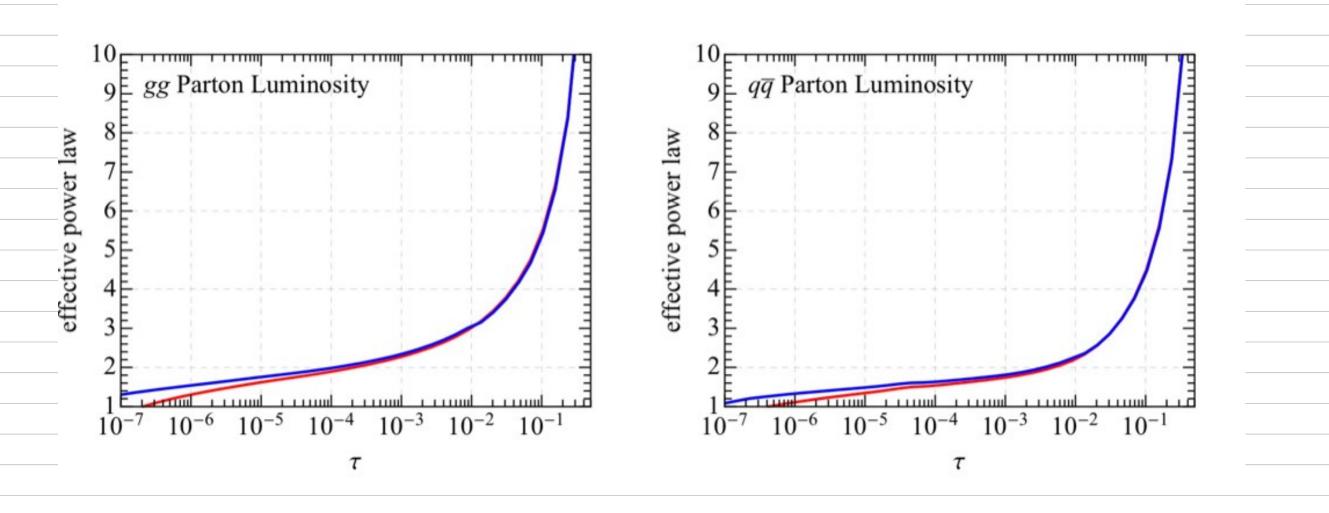




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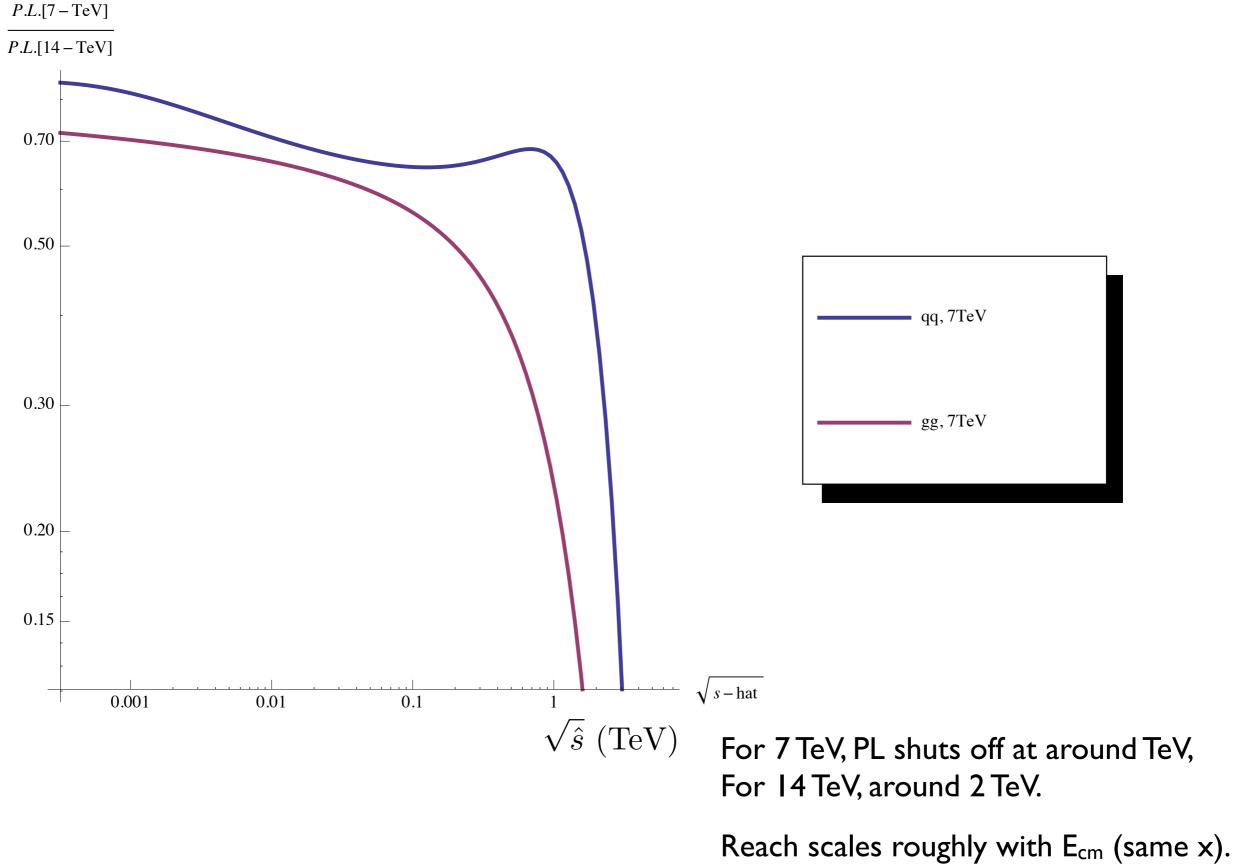
TeV, log scale





 $\texttt{PlotLegend} \rightarrow \{\texttt{"qq, 7TeV", "gg, 7TeV"}\}, \texttt{LegendPosition} \rightarrow \{\texttt{1.1, -0.4}\}, \texttt{Joined} \rightarrow \texttt{True}$





Rough estimates of discovery reach

$$\sigma \sim L_p \cdot \hat{\sigma} \sim \frac{1}{\tau^a} \hat{\sigma}$$

 L_p : parton luminosity, $\hat{\sigma}$: parton cross section

Production of new physics particle of mass M

Fast falling parton luminosity \Rightarrow

dominant contribution from parton cross section near threshold

$$\hat{s} \sim M^2 \rightarrow \tau \sim \frac{M^2}{S}$$
$$\hat{\sigma} \sim \frac{1}{M^2}$$

Number of new physics particle produced:

 $N = \sigma \cdot \mathcal{L}$ \mathcal{L} : luminosity

Discovery reach

 $E_2 > E_1$

Reach for new physics at these 2 colliders Collider I: M_1 . Collider 2: M_2 .

Assume the reach is obtained from the same number of signal events

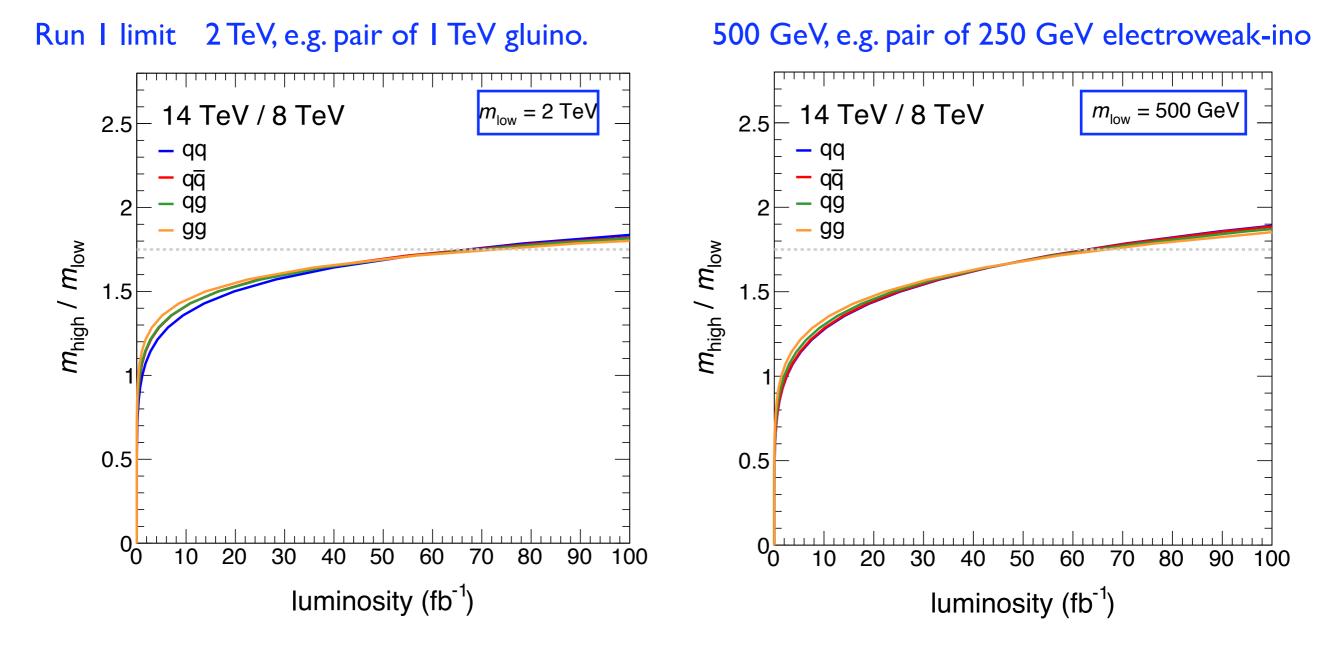
$$\frac{1}{\tau_1^a} \frac{1}{M_1^2} \mathcal{L}_1 = \frac{1}{\tau_2^a} \frac{1}{M_2^2} \mathcal{L}_2 \qquad \text{used} \quad \hat{\sigma} \sim \frac{1}{M^2}$$

We have

$$\frac{M_2}{M_1} = \left(\frac{S_2}{S_1}\right)^{1/2} \left(\frac{S_1}{S_2}\frac{\mathcal{L}_2}{\mathcal{L}_1}\right)^{\frac{1}{2a+2}} \qquad \text{used} \quad \hat{s} \sim M^2 \to \tau \sim \frac{M^2}{S}$$



As data accumulates



Rapid gain initial 10s fb⁻¹, slow improvements afterwards. Reaching the "slow" phase after Moriond 2017

Phase space

• General phase space factor:

$$d\Pi_n = \Pi_f \left(\int \frac{d^3 p_f}{(2\pi)^3} \frac{1}{2E_f} \right) (2\pi)^4 \delta^{(4)} (p_a + p_b - \sum p_f)$$

• One additional final state particle

~ an additional factor of
$$\frac{1}{16\pi^2}$$

• For example

2-body

$$X = A_{\mu} + b_{\mu}.$$

$$d TT_{2} = (2\pi)^{4} \delta^{(4)} (X_{\mu} - a_{\mu} - b_{\mu}) \frac{d^{2}a^{1}}{(2\pi)^{5} 2a_{0}} \frac{d^{2}b^{1}}{(2\pi)^{5} 2b_{0}}$$

$$= \frac{1}{8\pi} \lambda^{V_{2}} (1, \frac{a^{2}}{X^{2}}, \frac{b^{2}}{X^{2}}) \frac{d\Omega}{4\pi} = \frac{1}{4\pi} \frac{|\vec{P}a|}{X} \frac{d\Omega}{4\pi}$$

$$\lambda(x, y, z) = \chi^{2} + \eta^{2} + z^{2} - 2xy - 2yz - 2xz$$

$$Assume : a^{2} = b^{2} = m^{2} \quad \text{for simplicity}$$

$$\lambda^{V_{2}} = (1 - \frac{4m^{2}}{X^{2}})^{V_{2}}$$
Near Hureshold
$$\chi^{2} \doteq (2m + 8)^{2}$$

$$\lambda^{V_{2}} \approx (\frac{\delta}{2m})^{1/2} + \cdots$$

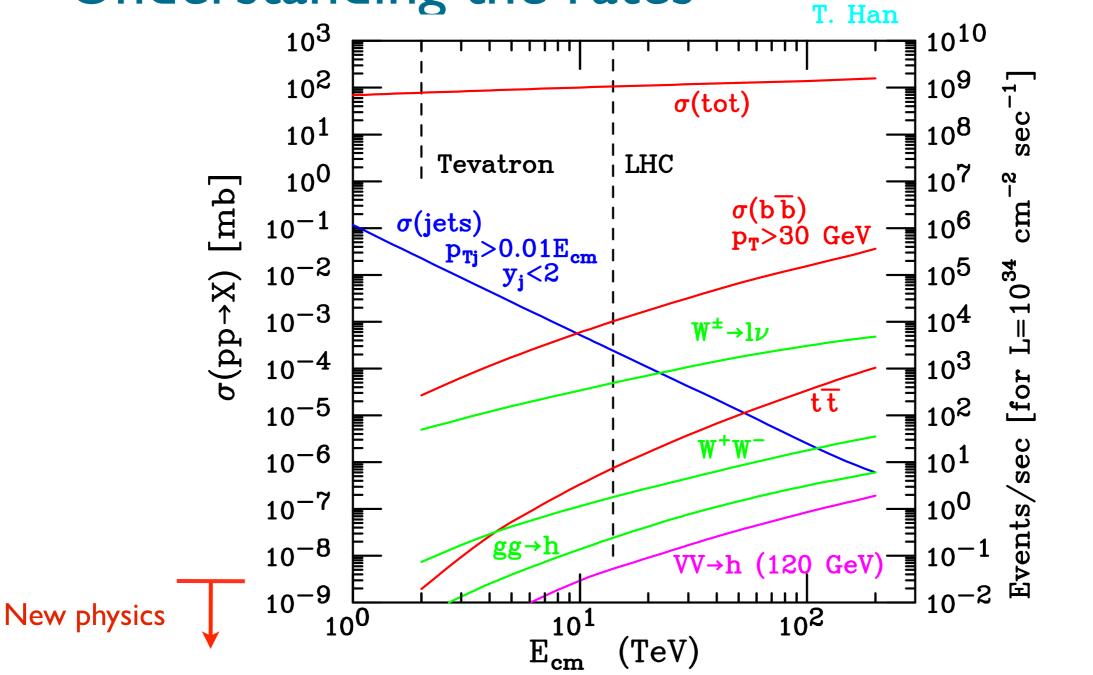
$$dT_{2} \asymp \delta^{1/2}$$

3 body. $y_n = P_{i,n} + P_{2,n} + P_{3,n}$ $dT_3 = (2\pi)^4 s^{(4)} (y - P_i - P_2 - P_3) \frac{3}{11} \frac{dP_i}{dP_i}$ $i = i (2\pi)^3 2P_{i0}$ Decompose $Y = X + P_3$ $X = P_1 + P_2$ $dT_{3} = \int_{2\pi} dT_{2} (Y \rightarrow X P_{3}) dT_{2} (X \rightarrow P_{1}, P_{2}) dX^{2}$ $m_{1} + m_{2} \leq \sqrt{X^{2}} \leq \sqrt{Y^{2}} - m_{3}$ Way above threshold, energy is the only dim-ful quantity dTI3 ~ The E2 dTI2 suppressed w.r.t. 2-body Near twreshold. Y2~ (3m+8)² $d\pi_2(\gamma \Rightarrow \chi P_3) \sim d\pi_2(\chi \Rightarrow P_1, P_2) \sim \delta^{1/2}$ $dX^2 \sim mS$ dTT3 x 82 open slower than 2-body

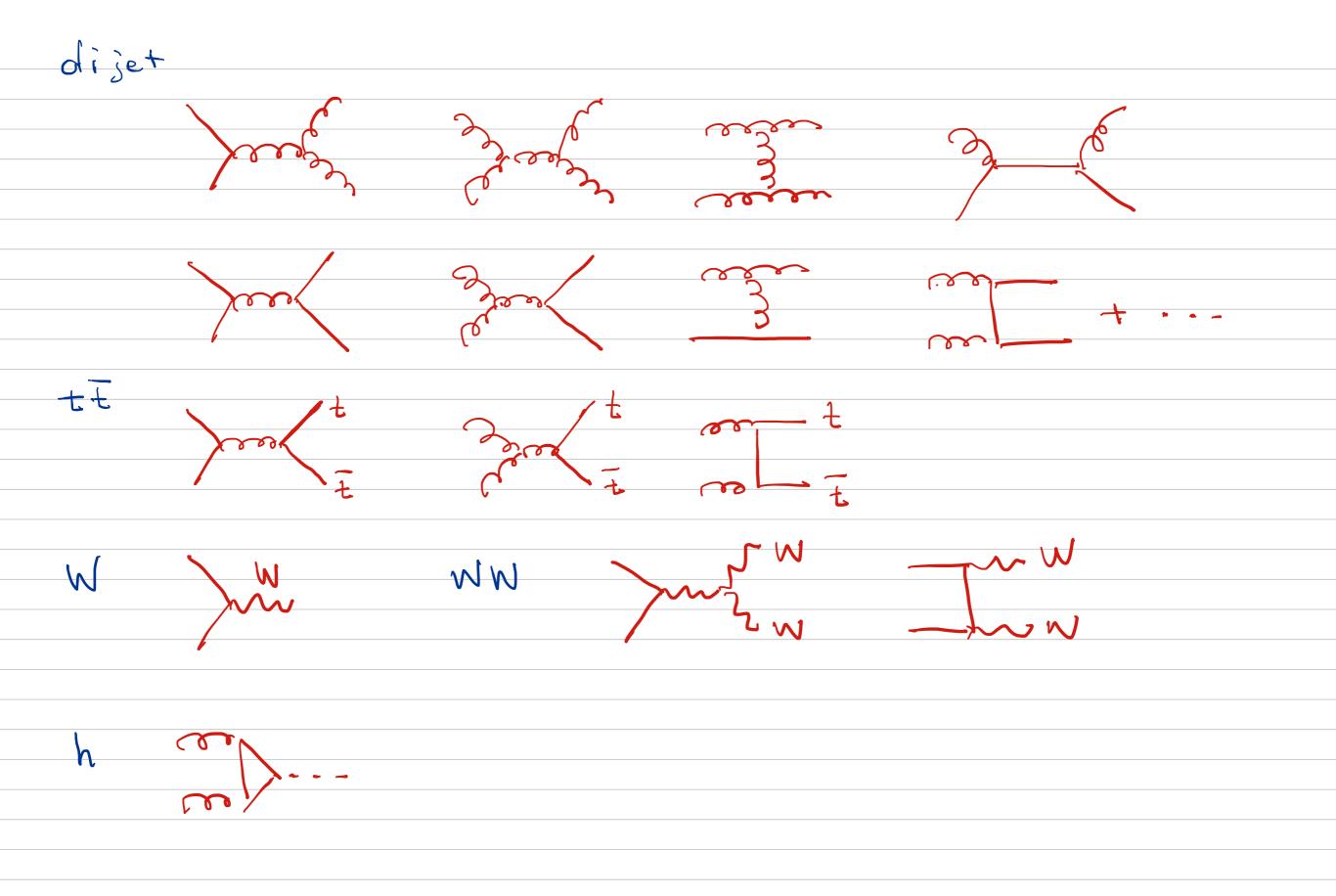
Rate also depends on

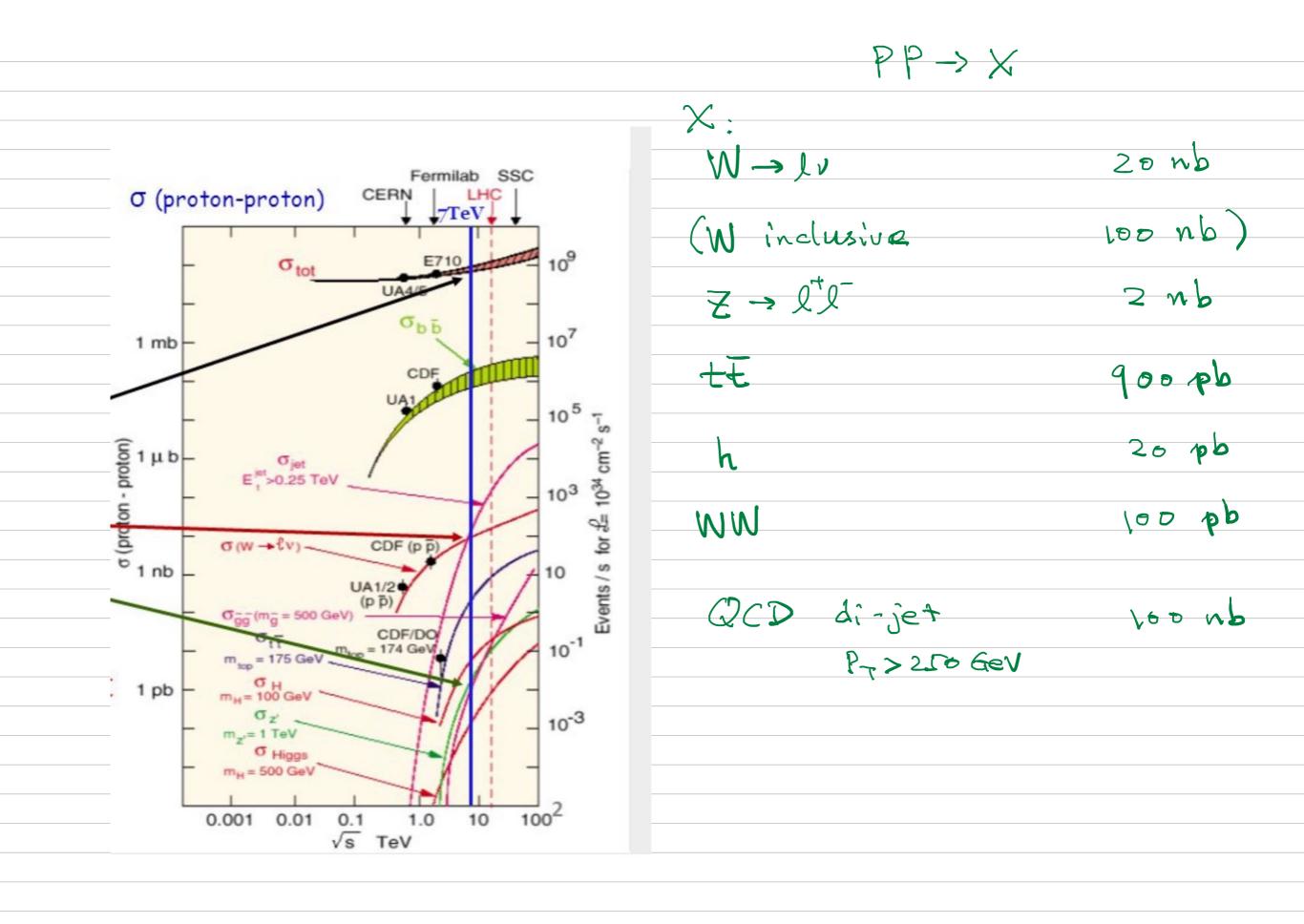
- Coupling constants
 - More final state particles, higher power of coupling constants.
 - QCD process dominates over weak processes.
- Singularities (enhancements) of matrix elements
 - Resonances.
 - Collinear and soft regime...

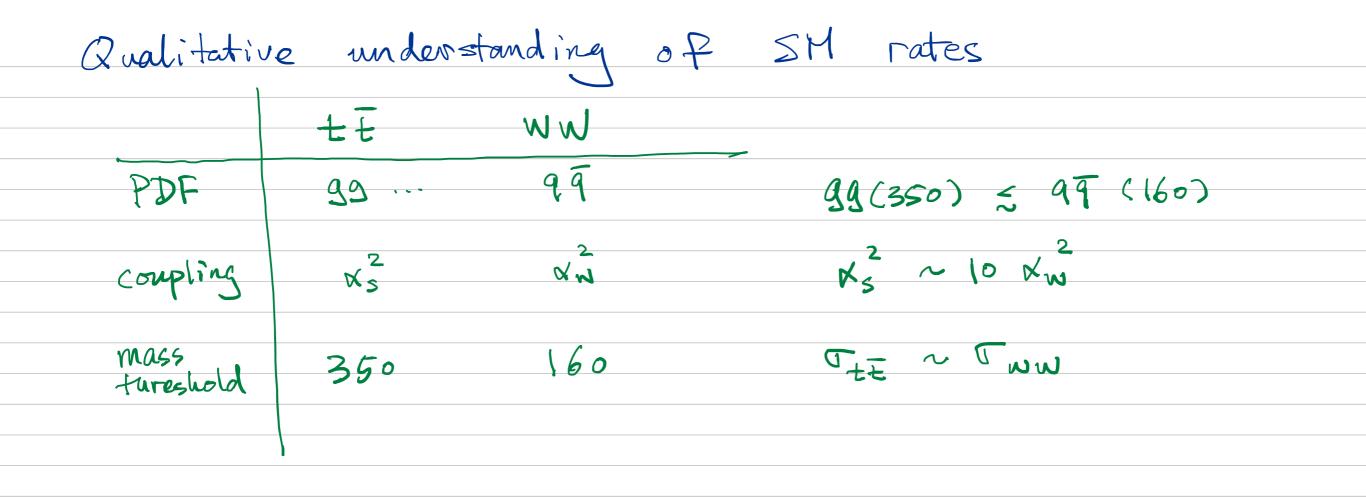
Understanding the rates

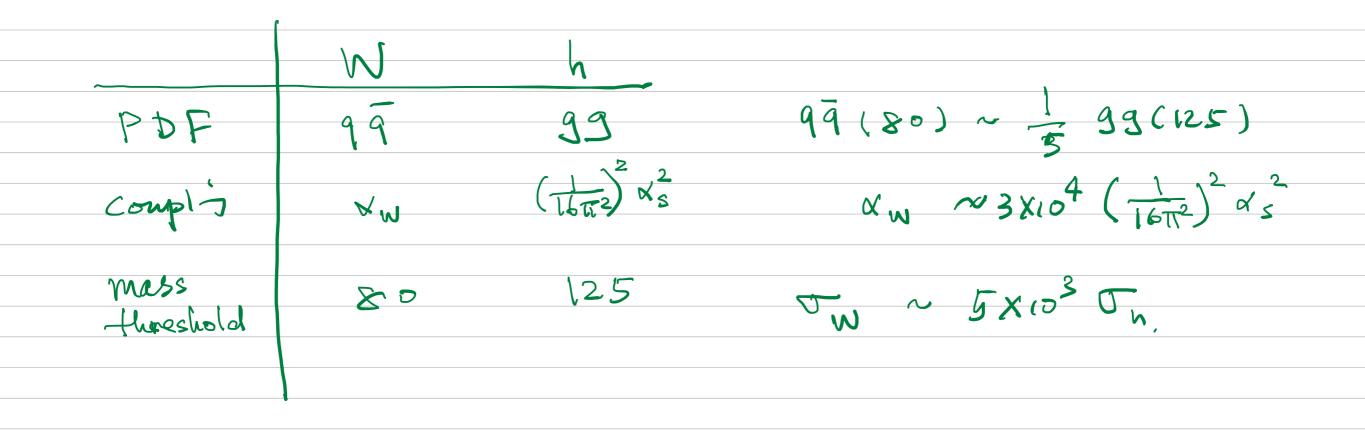


Example: considering ttbar vs W⁺W⁻, The relevant factors are: top is twice as heavy as W (2 times higher threshold) α_s^2 vs α_w^2 ttbar is gg dominated, WW is qqbar.









dijet vs
$$\pm \overline{E}$$

 $\sigma(dijet, p_{\tau}^{2} > 250) \sim 100 \ \overline{\sigma_{et}}$
 $\cdot Many more diagrams for di-jet. $\mathcal{O}(10)$ enhancement
 $\cdot \text{ Forward pingubarity in di-jet}$
 $\overline{I} \qquad \overline{\sigma_{etc}} \qquad \text{etc.}$$

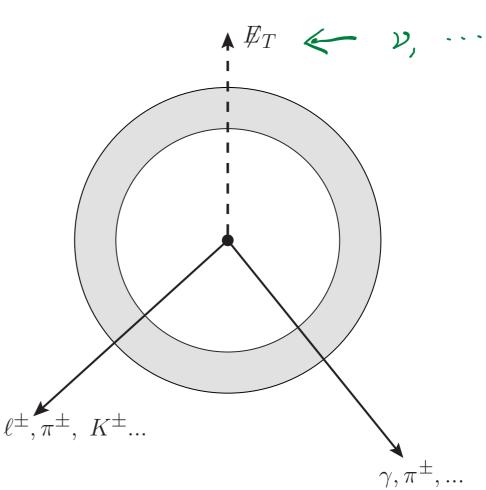
Why is it hard to discover TeV-scale new physics at the LHC

- p p collider, "prefers" to produce lighter states.
- Production rates scale roughly as $\sigma_{pp\to M} \sim \frac{1}{M^6}$
- TeV new physics $M_{\rm NP} \sim 5 10 \times M_{\rm SM(W,Z,t,...)}$
 - $\sigma_{\rm SM} \ge 10^6 \times \sigma_{\rm NP}$
- Dominated by QCD: A messy environment.
- Need:
 - Precise knowledge of the SM processes.
 - Anticipation of potential new physics states and their properties.

Being produced does not mean we can see them!

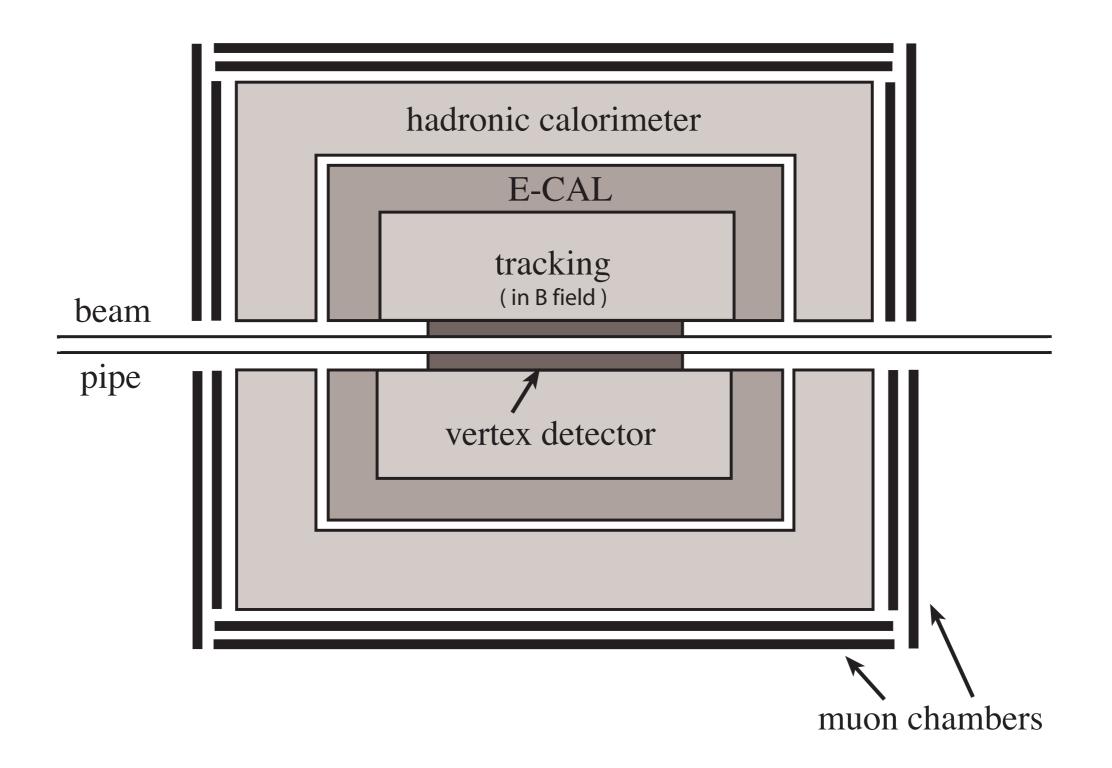
Final state Objects

- Colored particles: cluster of hardonic energy, jet
- Leptons: electron, muon
- Photon
- Heavy flavor: bottom (charm) $\tau_{b} \sim 1^{0^{-12}} \leq \tau_{c} \sim 1^{0^{-13}} \leq 1^{0^{-13}}$
 - Missing energy (MET)

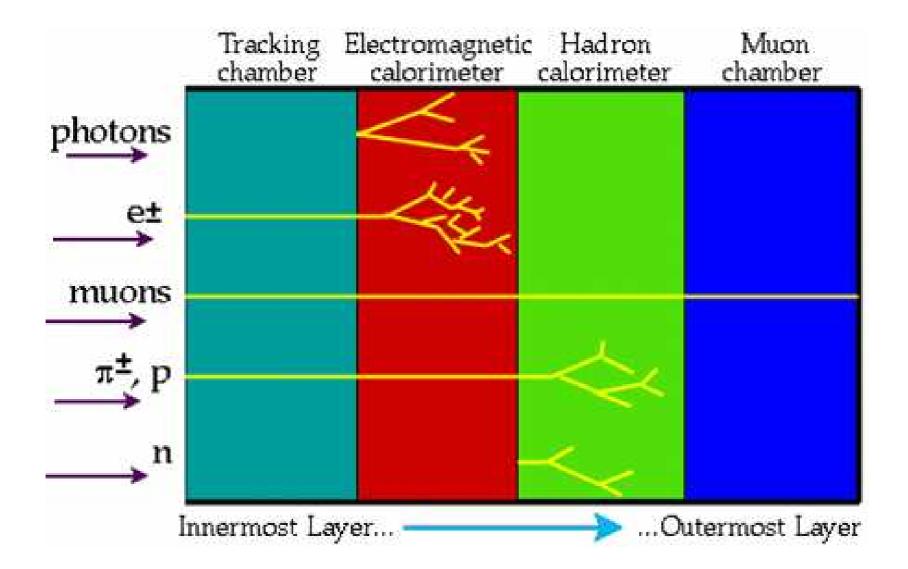


~ 9,9 K -

Modern detector (cartoon)



Identifying particles



From SM processes

- QCD: quark, gluon→ jets
- QCD heavy flavor: b, c.
- Z: $Z \to (q\bar{q}, \ell^+\ell^-, \nu\bar{\nu}) \to \text{jets}$, lepton pair, $\not\!\!\!E_T$
- $W: W^{\pm} \to (q\bar{q'}, \ell^{\pm}\nu) \to \text{jets}, \text{lepton} + \not\!\!\!E_T$
- Top: $t \to b + (W \to q\bar{q}' \text{ or } \bar{\ell}\nu)$
- Tau lepton: narrow jet(s), lepton.

SM Rates at 7 TeV:

- **QCD di-jet:** $p_T^j > 100 \text{ GeV}, 300 \text{ nb}$
- Heavy flavor: $b\overline{b}, p_T^b > 100 \text{ GeV}, 1 \text{ nb}$
- W+...: $W^{\pm} \to \ell \nu$, 14 nb $W^{\pm}(\to \ell \nu) + 1 \text{ jet}, \ p_T^j > 100 \text{ GeV}, \ 70 \text{ pb}$

one lepton + jets + MET

 $W^{\pm}(\rightarrow \ell \nu) + 2 \text{ jet}, \ p_T^j > 100 \text{ GeV}, \ 2 \text{ pb}$ $W^{\pm}(\rightarrow \ell \nu) + 1$ jet, $p_T^j > 200$ GeV, 5 pb • **Z** + ...: $Z(\to \ell^+ \ell^-)$, 1.4 nb

di-lepton + jets

 $Z(\to \ell^+ \ell^-) + 1 \text{ jet}, \ p_T^j > 100 \text{ GeV}, 10 \text{ pb}$

New Physics: ~ pb

SM rates at 7 TeV

• di-boson: $W^+W^-: 30 \text{ pb}$ di-lepton + MET, ~ 1.2 pb

 $W^+W^- + 1$ jet, $p_T^j > 100$ GeV, 2 pb di-lepton+jet+MET ~ 0.1 pb $W^+Z: 7$ pb, $W^-Z: 3.7$ pb

tri-lepton + MET ~ 0.1 pb

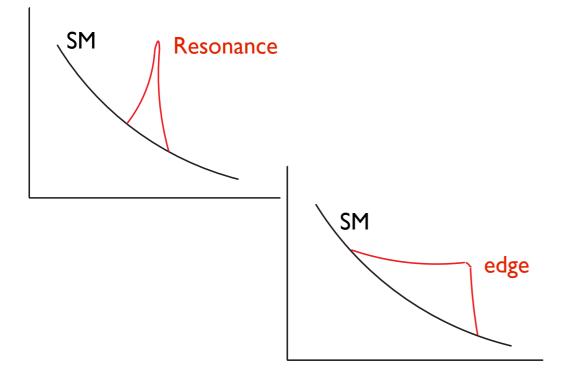
• top pair: 160 pb! Always has 6 objects.

 $t\bar{t} \rightarrow bbW^+W^- \rightarrow bbjj\ell\nu, bb\ell\nu\ell\nu, bbjjjj$

- (MET+lepton+Jet 40%, Heavy flavor...)
- Looks like new physics, pair production of a massive particle followed by a decay cascade.

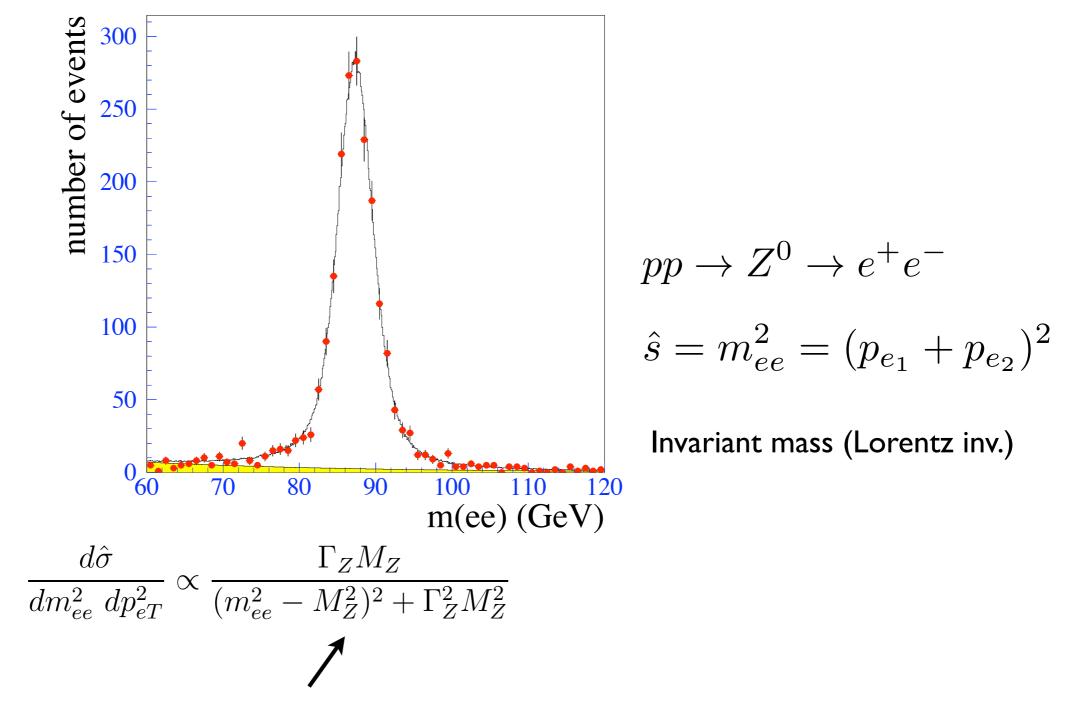
Two possible ways of discovery:

final state	rate estimate		
begin with \geq 2 hard jets	10 ⁵ Hz		
in addition			
hard jet	10 ² Hz		
or $ ot\!$	$\sim 10^2$ Hz		
or 1 lepton	10 ² Hz		
or 2 lepton	1 Hz		
or $2\ell = e^{\pm} + \mu^{\pm}$	10 ⁻⁴ Hz		



 Special kinematical features, resonances, edges, ...

Resonance



From matrix element: Breit-Wigner

Narrow width approximation

$$\frac{1}{X} = \frac{1}{1} \frac{1}{(S_{x} - m_{X}^{2})^{2} + \Gamma_{x}^{2} m_{X}^{2}} \approx \frac{\pi}{m_{x}\Gamma_{X}} Scs - m_{X}^{2}) \text{ if } F_{X} em_{X}}$$

$$\frac{1}{(S_{x} - m_{X}^{2})^{2} + \Gamma_{X}^{2} m_{X}^{2}} \approx \frac{\pi}{m_{x}\Gamma_{X}} Scs - m_{X}^{2}) \text{ if } F_{X} em_{X}}$$

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$$\frac{1}{(S_{x} - m_{X})^{2} + \Gamma_{X}^{2} m_{X}^{2}} = \pi Gcx)$$

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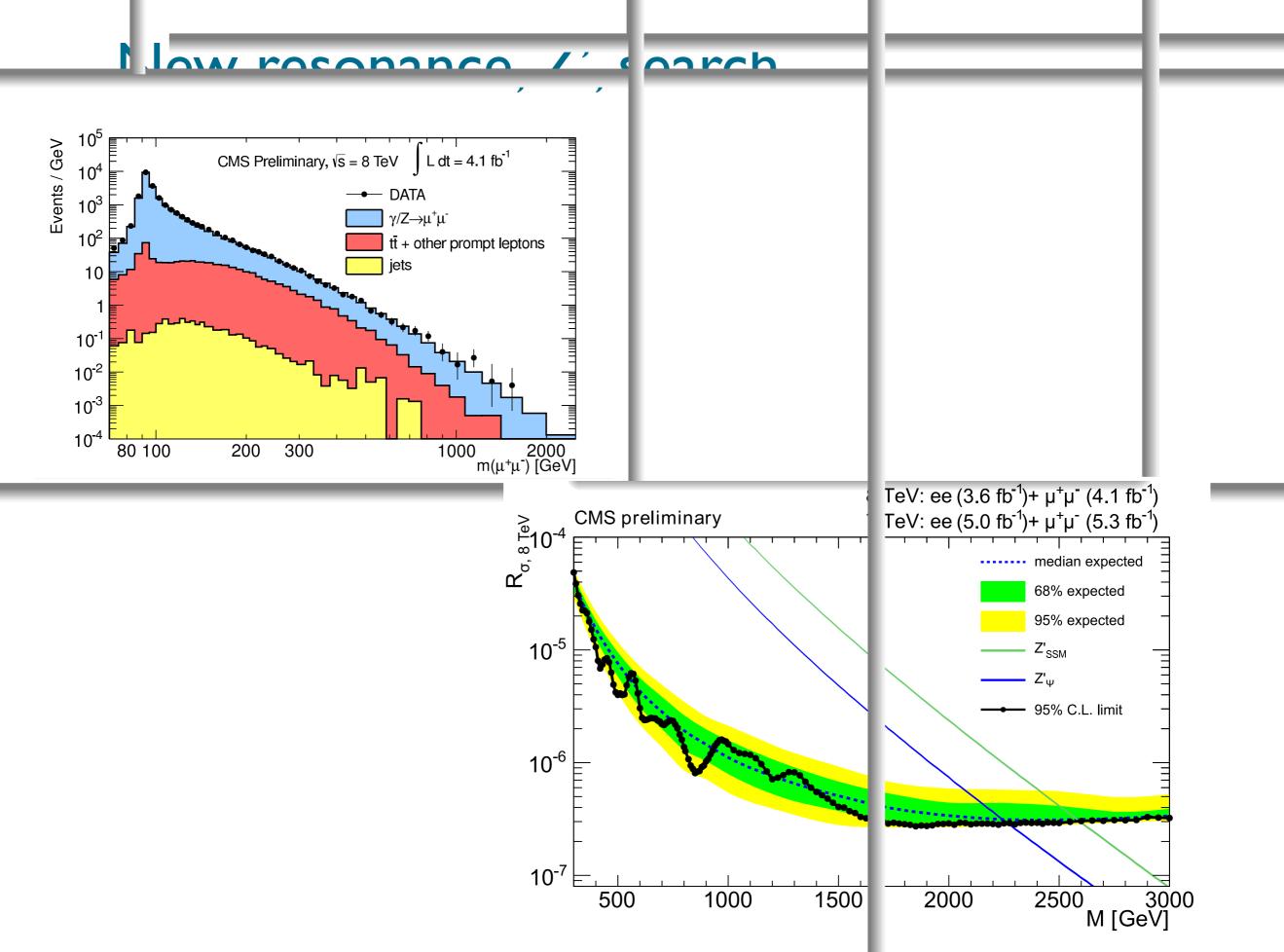
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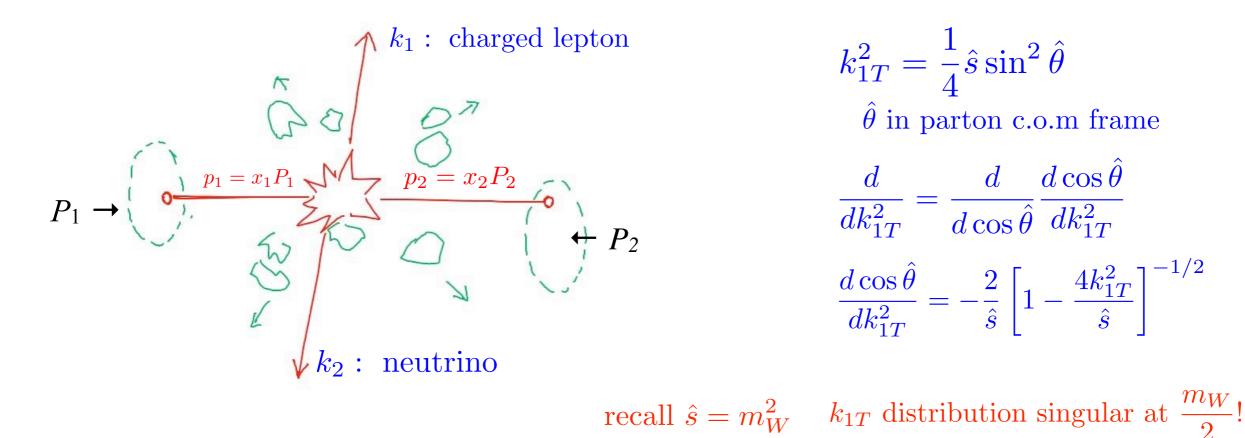
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Almost a resonance:

- What if we don't observe all the final state particles. For example, conspider $W \to \ell \nu$
- Cannot form an interesting Lorentz invariant variable.
 - At least can look for something invariant under boost along z-direction, e.g., transverse component



Jacobian peak

Transverse Mass

 $M_{12}^{2} = (E_{e} + E_{v})^{2} - (k_{1T} + k_{2T})^{2} - (k_{1Z} + k_{2Z})^{2}$

Define

Define transverse mass

 $M_{T} = (E_{eT} + E_{vT})^{2} - (k_{1T} + k_{2T})^{2}$

End boint

 $m_{12}^2 > m_T^2$ end point at $M_T = M_{12} = M_N$

Proof for the end point.

$$M_{12}^{2} = (E_{e} + E_{v})^{2} - (\overline{k_{1T}} + \overline{k_{2T}}) - (k_{12} + k_{22})^{2} \qquad E_{e} = \int \overline{E_{eT}}^{2} + \overline{k_{12}}^{2}$$

$$M_{T}^{2} = (E_{eT} + E_{vT})^{2} - (\overline{k_{1T}} + \overline{k_{2T}})^{2}$$

$$(E_{e} + E_{v})^{2} = E_{eT}^{2} + \overline{E_{vT}}^{2} + k_{12}^{2} + k_{22}^{2} + 2\int \overline{E_{vT}}^{2} + k_{12}^{2}$$

$$m_{T}^{2} - m_{T}^{2}$$

$$= 2\sqrt{E_{vT}}^{2} + \overline{k_{12}}^{2} \int \overline{E_{vT}}^{2} + \overline{k_{23}}^{2} - 2E_{eT}E_{vT} \geq 0$$

$$M$$

$$= (\int \int \int \frac{2}{\sqrt{2}} = E_{eT}^{2} + \overline{E_{vT}}^{2} + k_{12}^{2} + E_{eT}^{2} + \frac{2}{k_{22}} + 2 \int \overline{E_{vT}}^{2} + \overline{k_{12}}^{2}$$

$$(E_{e} + E_{v})^{2} = E_{eT}^{2} + E_{vT}^{2} + k_{12}^{2} + E_{vT}^{2} + E_{vT}^{2} + E_{vT}^{2}$$

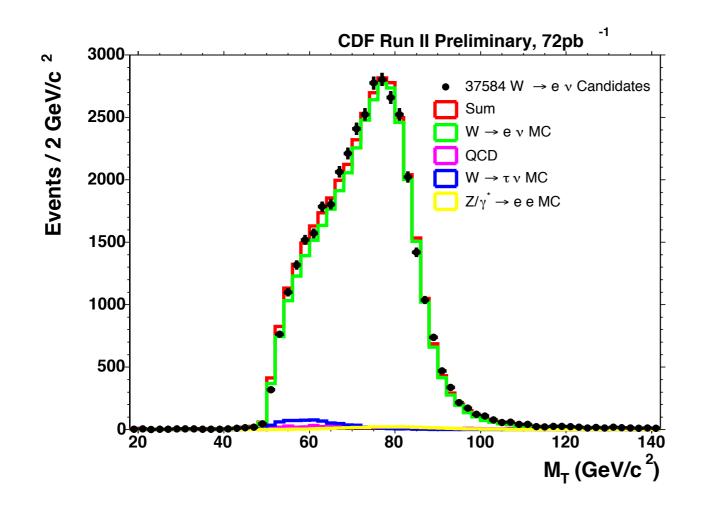
$$M_{T}^{2} - M_{T}^{2}$$

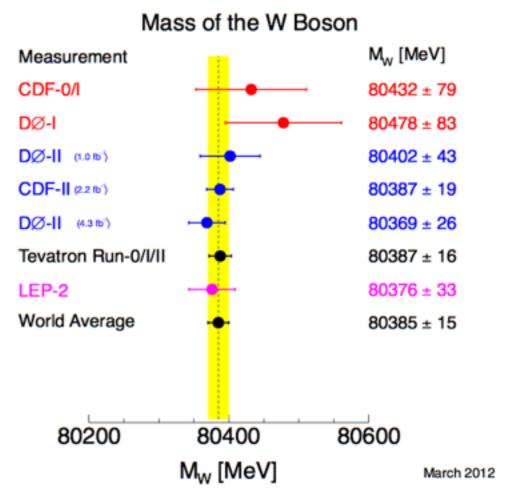
$$= 2\sqrt{E_{eT}^{2} + k_{12}^{2}} \int \overline{E_{vT}}^{2} + E_{vT}^{2} + k_{12}^{2} + E_{vT}^{2} + E$$

Jacobian peak in
$$M_T$$

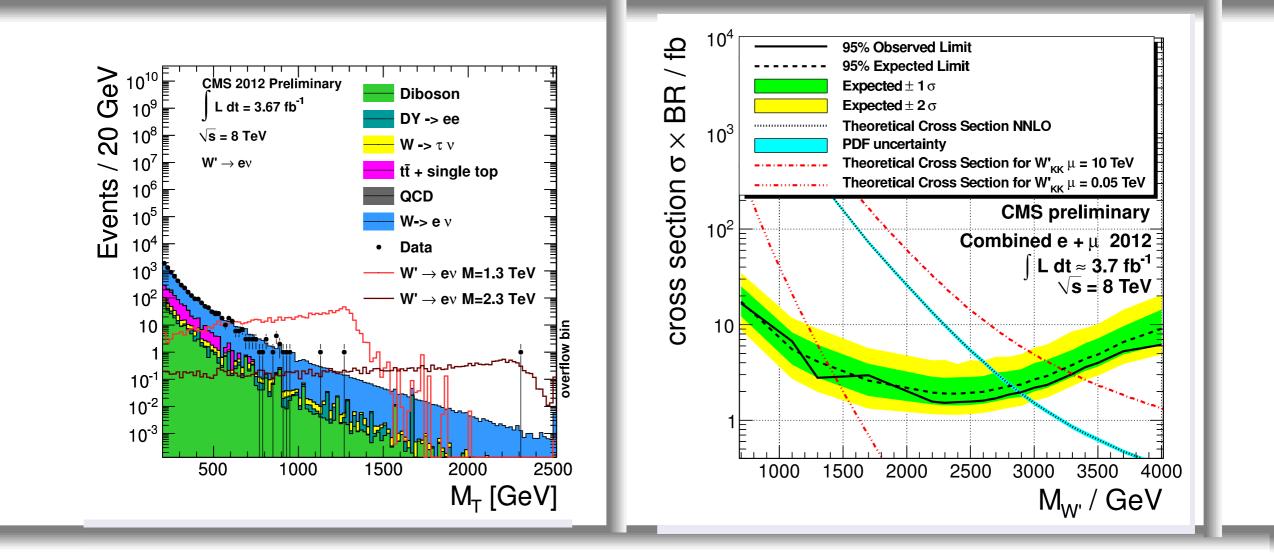
If N produced without transverse boost.
 $M_T = 2|k_{1T}| = 2|k_{2T}|$
 $\frac{d}{dm_T^2} = \frac{d}{dk_{1T}} \rightarrow Jacobian peak at $M_T^2 = M_W^2$
 $\frac{d\hat{\sigma}}{dm_T^2} \sim \frac{1}{4\pi} \frac{(G_F W_W)^2}{(S-m_W^2)^2 + (F_W W_W)^2} \frac{2-m_T^2/s}{(1-m_T^2/s)^{1/2}}$
Position of Jacobian peak smeared by width, resolution
Shape of Jacobian peak shanged by transverse boost.$

Measuring the W mass

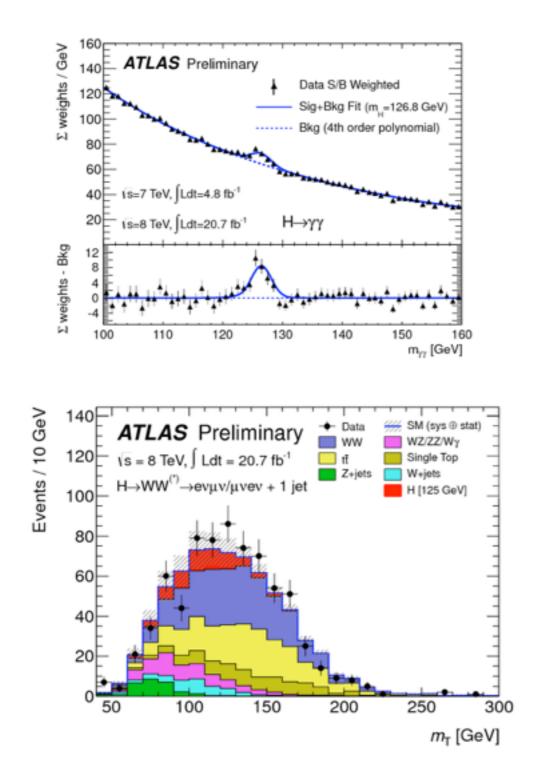


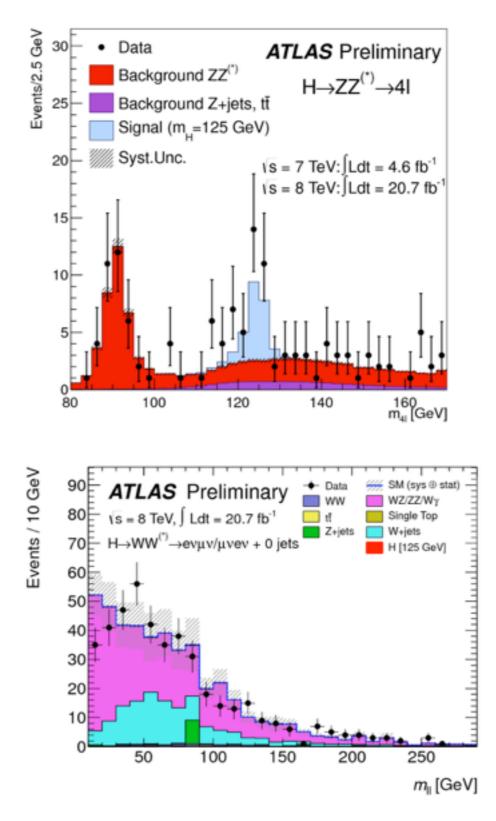


M/ coarch



Seeing Higgs





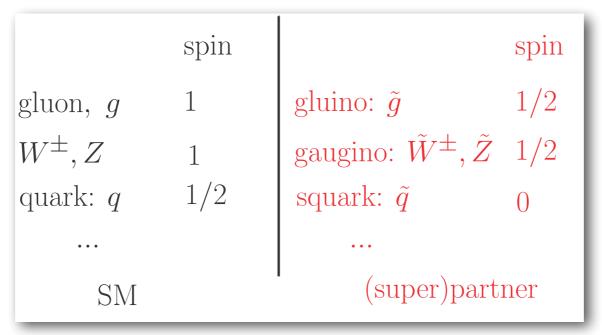
Complicated New physics signals

Partners:

New physics states with similar interactions to those of the Standard Model particles, such as the superpartners in Supersymmetry.

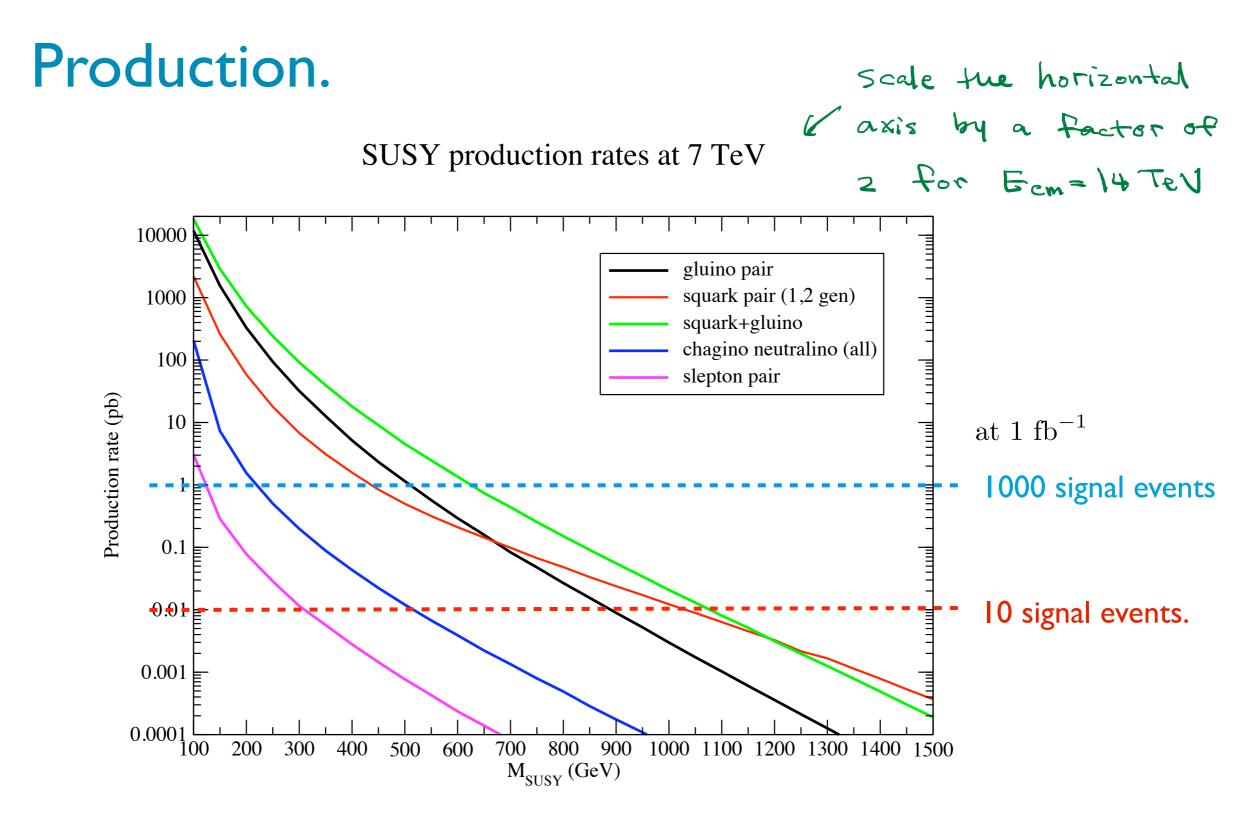
TeV Supersymmetry (SUSY)

- Supersymmetry. $|boson\rangle \Leftrightarrow |fermion\rangle$
- An extension of spacetime symmetry.
- New states: "Partners"



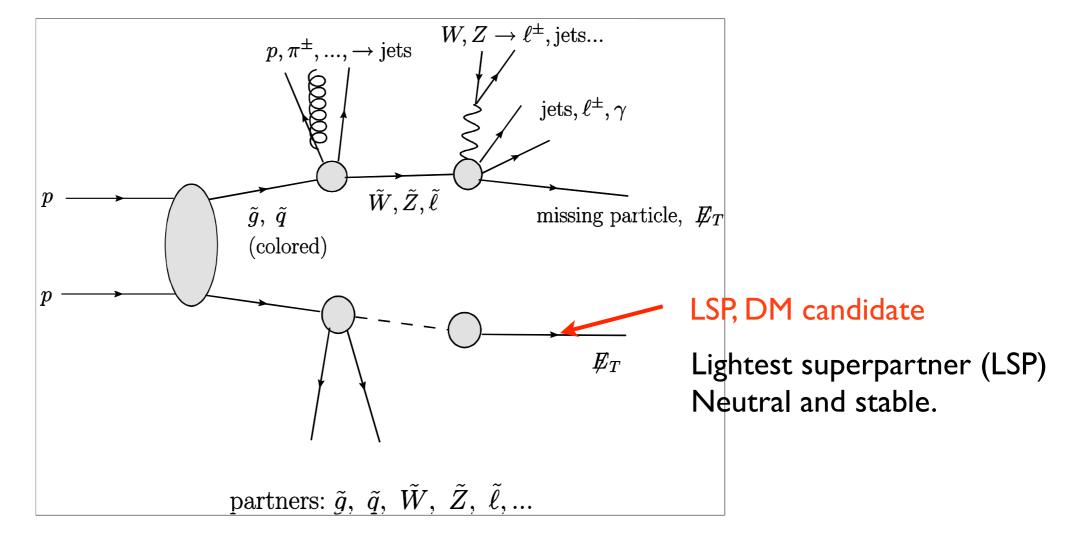
- Couplings relate to SM interactions via supersymmetry.
 - ~ same strength.

Review: S. Martin "A Supersemmtrage Primer", hep-ph/9709356



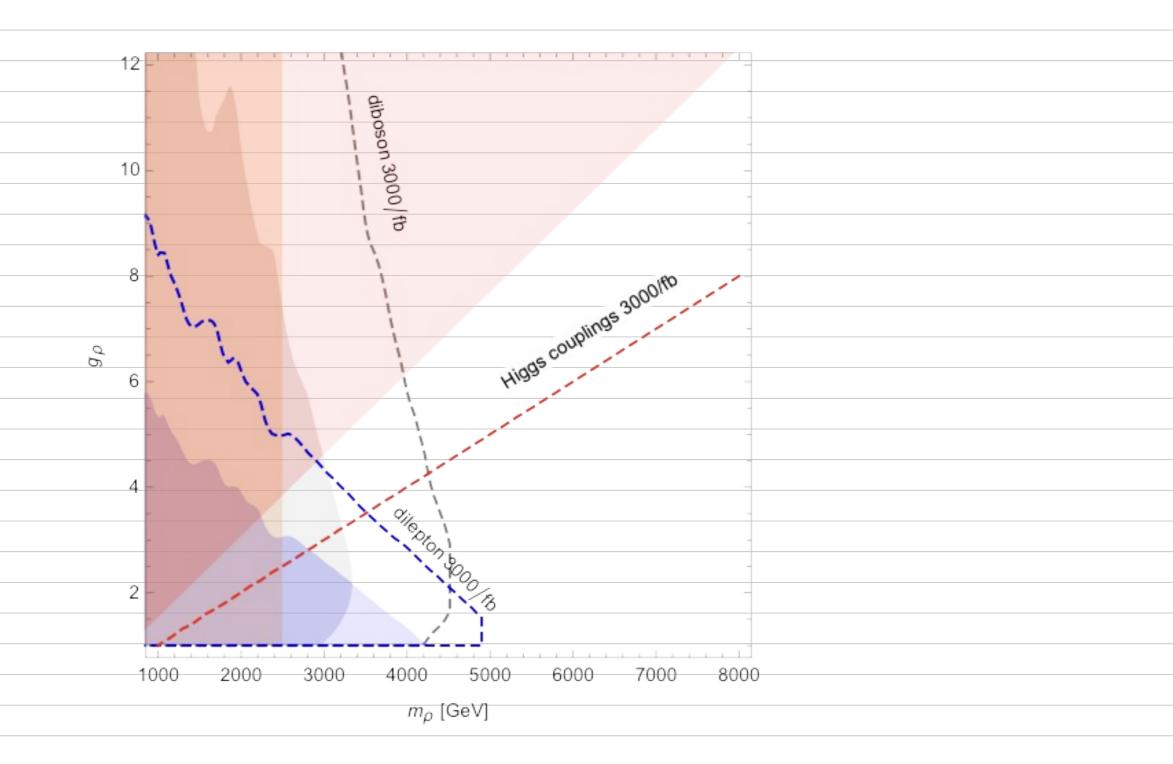
Dominated by the production of colored states. Similar pattern for other scenarios. Overall rates scaled by spin factors.

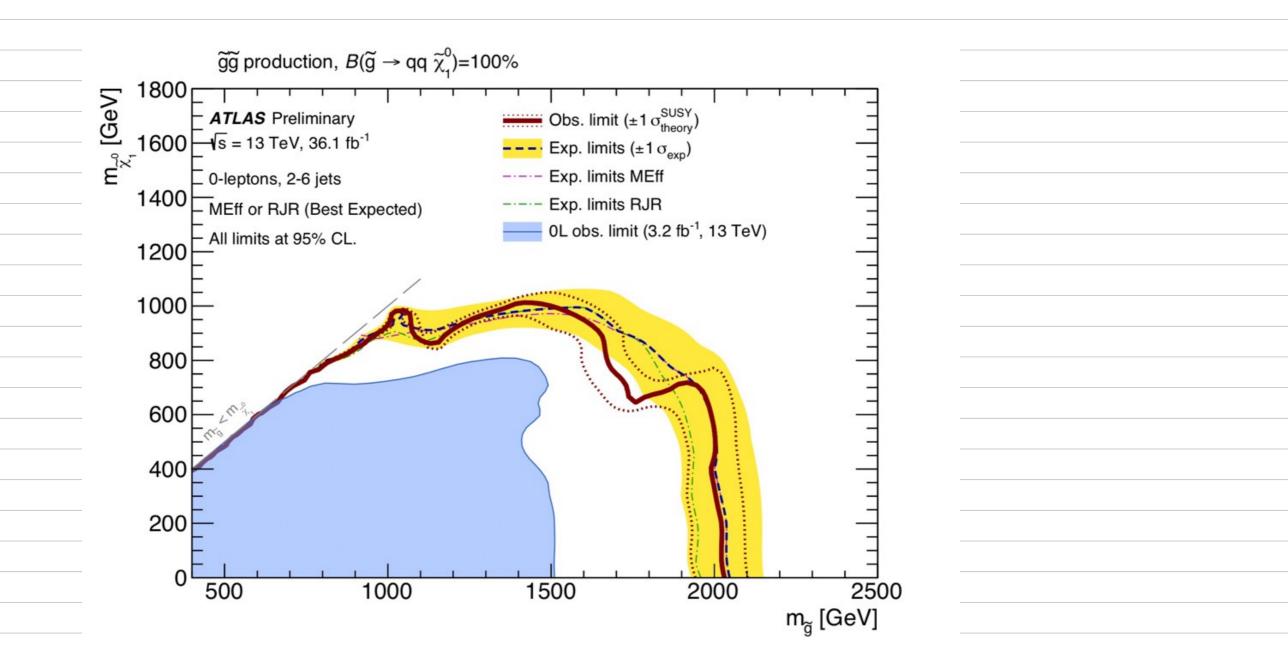
SUSY at colliders



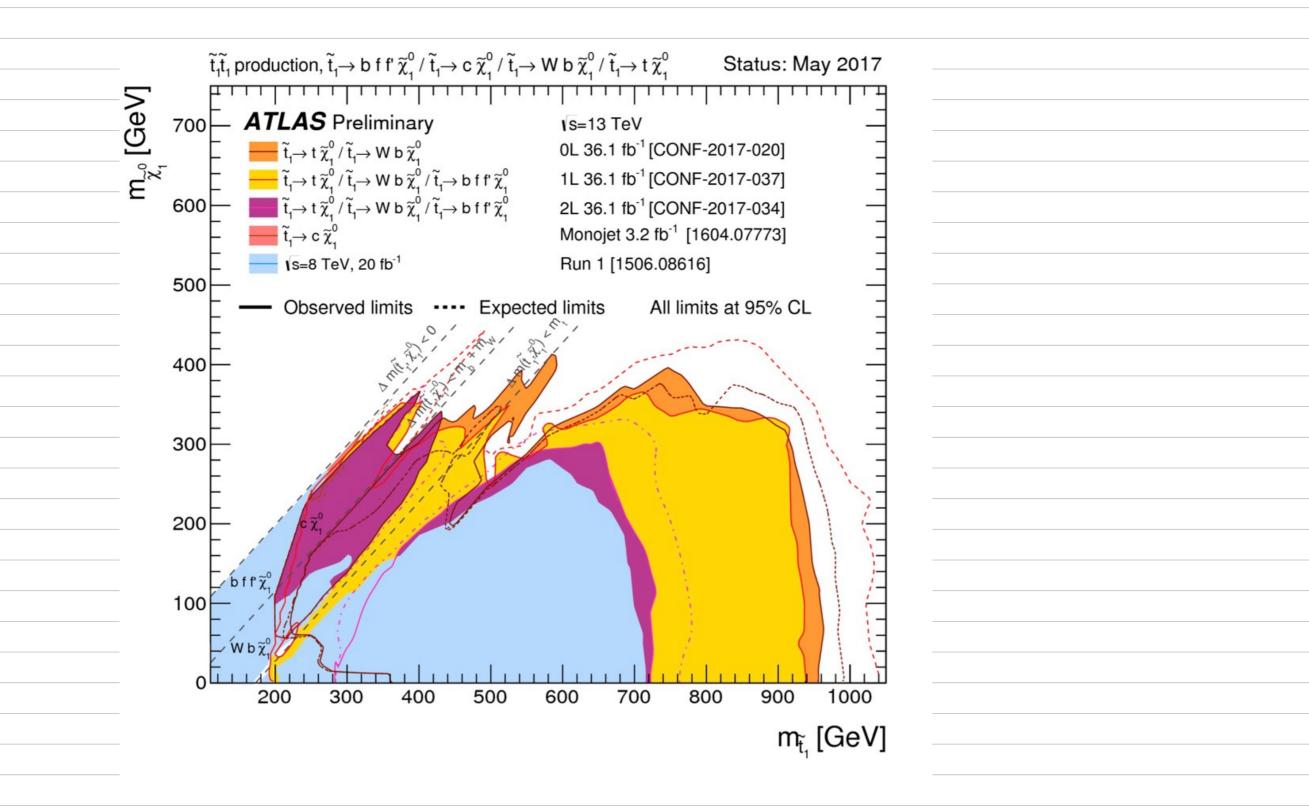
- long decay chain.
- jets, leptons, missing E_T
- Nice signal, good discovery potential.

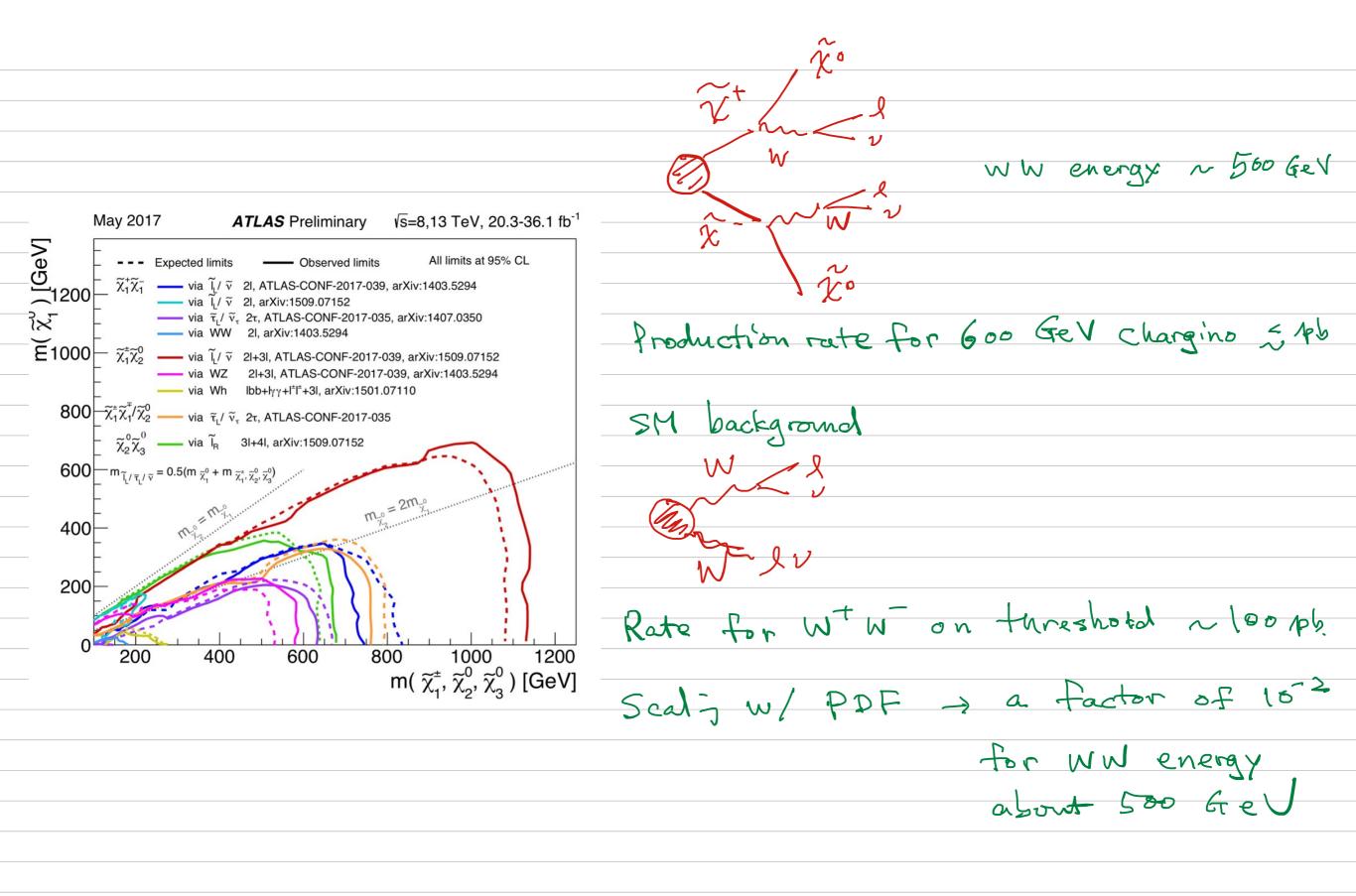
May 2017 Model	e, μ, τ, γ	Jets	$E_{\rm T}^{\rm miss}$	∫£ dı[ħ	Mass limit	$\sqrt{s} = 7, 8$	TeV $\sqrt{s} = 13$ TeV	$\sqrt{s} = 7, 8, 13 \text{ TeV}$ Reference	
$\begin{array}{c} MSUGRA(CMSSM\\ \tilde{q}\tilde{q}, \tilde{q} \rightarrow q \tilde{x}_{1}^{0} \\ \tilde{q}\tilde{q}, \tilde{q} \rightarrow q \tilde{x}_{1}^{0} \\ \tilde{q}\tilde{q}, \tilde{q} \rightarrow q \tilde{x}_{1}^{0} \\ \tilde{z}\tilde{s}, \tilde{s} \rightarrow q \tilde{q} \tilde{x}_{1}^{0} \\ \tilde{s}\tilde{s}, \tilde{s} \rightarrow q q \tilde{x}_{1}^{0} \\ \tilde{s}\tilde{s}, \tilde{s} \rightarrow q q \tilde{x}_{1}^{0} \\ \tilde{s}\tilde{s}, \tilde{s} \rightarrow q q \tilde{x}_{1}^{0} \\ \tilde{g}\tilde{s}, \tilde{s} \rightarrow q q \mathcal{W} \tilde{x}_{1}^{0} \\ \tilde{g}\tilde{s}, \tilde{s} \rightarrow q \tilde{s} \mathcal{H} \tilde{s} \tilde{s} \\ \tilde{g}\tilde{s}, \tilde{s} \rightarrow q \tilde{s} \mathcal{H} \tilde{s} \tilde{s} \\ \tilde{g}\tilde{s}, \tilde{s} \rightarrow q \tilde{s} \mathcal{H} \tilde{s} \tilde{s} \\ \tilde{g}\tilde{s}, \tilde{s} \rightarrow q \tilde{s} \mathcal{H} \tilde{s} \tilde{s} \\ \tilde{g}\tilde{s}, \tilde{s} \rightarrow q \tilde{s} \tilde{s} \tilde{s} \\ \tilde{g}\tilde{s}, \tilde{s} \rightarrow q \tilde{s} \tilde{s} \tilde{s} \tilde{s} \tilde{s} \tilde{s} \tilde{s} \tilde{s}$	$\begin{array}{c} 0.3 \ e, \mu/1-2 \ \tau \\ 0 \\ mono-jet \\ 0 \\ 3 \ e, \mu \\ 0 \\ 1-2 \ \tau + 0 - 1 \ e \\ 2 \ \gamma \\ \gamma \\ 2 \ e, \mu \left(Z \right) \\ 0 \end{array}$	2-6 jets 1-3 jets 2-6 jets 2-6 jets 4 jets 7-11 jets	Yes Yes Yes Yes Yes Yes Yes Yes Yes	20.3 36.1 36.1 36.1 36.1 36.1 3.2 3.2 20.3 13.3 20.3 20.3	4.2 3 3 4 5 5 5 5 5 5 5 5 5 5 5 5 5	1.57 TeV m 2.02 TeV 2.01 TeV 1.825 TeV 1.8 TeV 2.0 TeV 1.65 TeV 1.85 TeV 1.8 TeV	$\begin{split} &m(\hat{q}) \!\!=\!\!m(\hat{g}) \\ &t(\hat{k}_1^0) \!\!<\!\! 200 \mathrm{GeV}, m(1^w \mathrm{gen.} \hat{q}) \!\!=\!\!m(2^{nd} \mathrm{gen.} \hat{q}) \\ &m(\hat{q}_1^0) \!\!<\!\! 200 \mathrm{GeV}, \\ &m(\hat{k}_1^0) \!\!<\!\! 200 \mathrm{GeV}, \\ &m(\hat{k}_1^0) \!\!<\!\! 200 \mathrm{GeV}, \\ &m(\hat{k}_1^0) \!\!<\!\! 400 \mathrm{GeV} \\ &m(\hat{k}_1^0) \!\!<\!\! 400 \mathrm{GeV} \\ &cr(NLSP) \!\!<\!\! 0.1 mm \\ &m(\hat{k}_1^0) \!\!<\!\! 560 \mathrm{GeV}, cr(NLSP) \!\!<\!\! 0.1 mm, \mu \!\!<\!\! 0 \\ &m(\hat{k}_1^0) \!\!>\!\! 430 \mathrm{GeV} \\ &m(\hat{k}_2^0) \!\!>\!\! 430 \mathrm{GeV} \\ &m(\hat{k}_2^0) \!\!>\!\! 580 \mathrm{GeV}, cr(NLSP) \!\!<\!\! 0.1 mm, \mu \!\!>\!\! 0 \\ &m(NLSP) \!\!+\!\! 3430 \mathrm{GeV} \\ &m(\hat{d}) \!\!>\!\! 1.8 \times 10^{-4} \mathrm{eV}, m(\hat{q}) \!\!=\!\! 1.5 \mathrm{TeV} \end{split}$	1507.05525 ATLAS-CONF-2017-022 1804.07773 ATLAS-CONF-2017-022 ATLAS-CONF-2017-022 ATLAS-CONF-2017-020 ATLAS-CONF-2017-030 ATLAS-CONF-2017-033 1807.05979 1606.09150 1507.05493 ATLAS-CONF-2016-065 1503.03290 1502.01518	
$\begin{array}{c} & \tilde{g}\tilde{g}, \tilde{g} \rightarrow b \tilde{b} \tilde{\chi}_{1}^{0} \\ & \tilde{g}\tilde{g}, \tilde{g} \rightarrow t \tilde{t} \tilde{\chi}_{1}^{0} \\ & \tilde{g}\tilde{g}, \tilde{g} \rightarrow b \tilde{t} \tilde{\chi}_{1}^{1} \end{array}$	0 0-1 e,μ 0-1 e,μ	3 b 3 b 3 b	Yes Yes Yes	36.1 36.1 20.1	2 2 2	1.97 TeV	m(k ⁰ ₁)<500 GeV m(k ⁰ ₁)<200 GeV m(k ⁰ ₁)<300 GeV	ATLAS-CONF-2017-021 ATLAS-CONF-2017-021 1407.0600	
$ \begin{array}{c} \tilde{b}_1 \tilde{b}_1, \tilde{b}_1 \rightarrow b \tilde{x}_1^0 \\ \tilde{b}_1 \tilde{b}_1, \tilde{b}_1 \rightarrow t \tilde{x}_1^1 \\ \tilde{b}_1 \tilde{b}_1, \tilde{b}_1 \rightarrow t \tilde{x}_1^1 \\ \tilde{r}_1 \tilde{r}_1, \tilde{r}_1 \rightarrow b \tilde{x}_1^\pm \\ \tilde{r}_1 \tilde{r}_1, \tilde{r}_1 \rightarrow \tilde{r}_1 \tilde{x}_1^0 \\ \tilde{r}_1 \tilde{r}_1 (natural GMSB) \\ \tilde{r}_1 \tilde{r}_2 \tilde{r}_2, \tilde{r}_2 \rightarrow \tilde{r}_1 + Z \end{array} $	0 2 e, µ (SS) 0-2 e, µ 0-2 e, µ 0 2 e, µ (Z) 3 e, µ (Z)	2 b 1 b 1-2 b 0-2 jets/1-2 mono-jet 1 b 1 b	2 b Yes 2	36.1 36.1 4.7/13.3 20.3/36.1 3.2 20.3 36.1	ã₁ 950 GeV ẵ₁ 275-700 GeV ĩ₁ 117-170 GeV 200-720 GeV 205-950 GeV ĩ₁ 90-323 GeV ĩ₁ 90-323 GeV ĩ₁ 90-323 GeV ĩ₂ 290-790 GeV		$\begin{array}{l} m(\tilde{\xi}_1^0)\!\!<\!\!420GeV \\ m(\tilde{\xi}_1^0)\!\!<\!\!420GeV, m(\tilde{\xi}_1^0)\!\!+\!m(\tilde{\xi}_1^0)\!\!+\!100GeV \\ m(\tilde{\xi}_1^0)\!\!+\!\!2m(\tilde{\xi}_1^0)\!\!+\!\!55GeV \\ m(\tilde{\xi}_1^0)\!\!+\!\!1GeV \\ m(\tilde{\xi}_1^0)\!\!+\!\!1GeV \\ m(\tilde{\xi}_1^0)\!\!+\!\!150GeV \\ m(\tilde{\xi}_1^0)\!\!+\!\!150GeV \\ m(\tilde{\xi}_1^0)\!\!+\!\!160GeV \\ m(\tilde{\xi}_1^0)\!\!+\!160GeV \\ m(\tilde{\xi}_1^0$	ATLAS-CONF-2017-038 ATLAS-CONF-2017-030 1209.2102, ATLAS-CONF-2016-077 1506.08616, ATLAS-CONF-2017-020 1604.07773 1403.5222 ATLAS-CONF-2017-019	
$\begin{array}{c} \tilde{i}_{2}\tilde{i}_{2},\tilde{i}_{2}\rightarrow\tilde{i}_{1}+\hbar \\ \\ \tilde{i}_{L,R}\tilde{i}_{L,R},\tilde{\ell}\rightarrow\ell\tilde{x}_{1}^{0} \\ \tilde{x}_{1}^{+}\tilde{x}_{1}^{-},\tilde{x}_{1}^{+}\rightarrow\tilde{\ell}\nu(\ell\tilde{v}) \\ \tilde{x}_{1}^{+}\tilde{x}_{1}^{-},\tilde{x}_{2}^{+}\rightarrow\tilde{v}(\tau\tilde{v}),\tilde{x}_{2}^{0}\rightarrow\tilde{\tau}\tau(\tau) \\ \tilde{x}_{1}^{+}\tilde{x}_{2}^{-}\rightarrow\tilde{w}_{1}^{-}v\tilde{\ell}_{1}(\ell\tilde{v}),\ell\tilde{v}\tilde{\ell}_{L}(\ell\tilde{v}) \\ \tilde{x}_{1}^{+}\tilde{x}_{2}^{0}\rightarrow\tilde{w}_{1}^{-}v\tilde{\ell}_{1}(\tilde{v}),\tilde{v}\tilde{v}\tilde{\ell}_{L}\ell(\tilde{v}) \\ \tilde{x}_{1}^{+}\tilde{x}_{2}^{0}\rightarrow\tilde{w}_{1}^{-}v\tilde{t}_{1}\tilde{x}_{1}^{0},\hbar\rightarrow\tilde{b}\tilde{b}/WW/\tau \\ \tilde{x}_{2}^{+}\tilde{x}_{2}^{0}\rightarrow\tilde{w}_{1}^{-}\tilde{u}h\tilde{x}_{1}^{0},\hbar\rightarrow\tilde{b}\tilde{b}/WW/\tau \\ \tilde{x}_{2}^{+}\tilde{x}_{2}^{-}\tilde{x}_{2}^{-}\tilde{\tau}\ell\ell \\ GGM (wino NLSP) weak prod. \\ GGM (bino NLSP) weak prod. \end{array}$	$3 e, \mu$ $2 \cdot 3 e, \mu$ $\tau/\gamma\gamma \qquad e, \mu, \gamma$ $4 e, \mu$ $, \tilde{\chi}_{1}^{0} \rightarrow \gamma \tilde{G} \qquad 1 e, \mu + \gamma$	4 b 0 0-2 jets 0-2 b 0 -	Yes Yes Yes Yes Yes Yes Yes Yes Yes	36.1 36.1 36.1 36.1 36.1 20.3 20.3 20.3 20.3	\vec{r}_1 320-880 GeV \vec{x} 90-440 GeV \hat{x}_1^{\pm} 710 GeV \hat{x}_1^{\pm} 760 GeV \hat{x}_1^{\pm} 760 GeV \hat{x}_1^{\pm} 760 GeV \hat{x}_1^{\pm} 580 GeV \hat{x}_{1x}^{\pm} 580 GeV \hat{x}_{1x}^{\pm} 635 GeV \hat{w} 115-370 GeV \hat{w} 590 GeV	m(č ²)=m($\begin{split} m(\tilde{\epsilon}_{1}^{n}) &= 0 \text{ GeV } \\ m(\tilde{\epsilon}_{1}^{n}) &= 0, m(\tilde{\epsilon}, \tilde{\nu}) &= 0.5(m(\tilde{\epsilon}_{1}^{n}) + m(\tilde{\epsilon}_{1}^{n})), \\ m(\tilde{\epsilon}_{1}^{n}) &= 0, m(\tilde{\epsilon}, \tilde{\nu}) = 0.5(m(\tilde{\epsilon}_{1}^{n}) + m(\tilde{\epsilon}_{1}^{n})), \\ \tilde{\epsilon}_{2}^{n}), m(\tilde{\epsilon}_{1}^{n}) &= 0, m(\tilde{\epsilon}, \tilde{\nu}) = 0.5(m(\tilde{\epsilon}_{1}^{n}) + m(\tilde{\epsilon}_{1}^{n})), \\ m(\tilde{\epsilon}_{1}^{n}) &= m(\tilde{\epsilon}_{2}^{n}), m(\tilde{\epsilon}_{2}^{n}) = 0, \tilde{\epsilon} \text{ decoupled } \\ m(\tilde{\epsilon}_{1}^{n}) &= m(\tilde{\epsilon}_{2}^{n}), m(\tilde{\epsilon}_{2}^{n}) = 0, \tilde{\epsilon} \text{ decoupled } \\ \tilde{\epsilon}_{2}^{n}), m(\tilde{\epsilon}_{1}^{n}) &= 0.5(m(\tilde{\epsilon}_{2}^{n}) + m(\tilde{\epsilon}_{1}^{n})), \\ c_{1} < 1 \text{ mm } \end{split}$	ATLAS-CONF-2017-019 ATLAS-CONF-2017-039 ATLAS-CONF-2017-039 ATLAS-CONF-2017-035 ATLAS-CONF-2017-039 ATLAS-CONF-2017-039 1501.07110 1405.5086 1507.05493 1507.05493	
Direct $\tilde{k}_{1}^{+}\tilde{k}_{1}^{-}$ prod., long-lived \tilde{k} Direct $\tilde{k}_{1}^{+}\tilde{k}_{1}^{-}$ prod., long-lived \tilde{k} Stable, stopped \tilde{g} R-hadron Metastable \tilde{g} R-hadron Metastable \tilde{g} R-hadron GMSB, stable $\tilde{\tau}, \tilde{k}_{1}^{0} \rightarrow \tau(\tilde{\epsilon}, \tilde{\mu}) + \tau$ GMSB, $\tilde{k}_{1}^{0} \rightarrow \tau G$, long-lived \tilde{k}_{1}^{0} $\tilde{g}\tilde{g}, \tilde{k}_{1}^{0} \rightarrow eev/euv/\mu\mu\nu$ GGM $\tilde{g}\tilde{g}, \tilde{k}_{1}^{0} \rightarrow Z\tilde{G}$	* dE/dx trk 0 trk dE/dx trk	1-5 jets - - -	Yes Yes - - Yes -	36.1 18.4 27.9 3.2 19.1 20.3 20.3 20.3	\$\bar{x}_1^a\$ 430 GeV \$\bar{x}_1^a\$ 495 GeV \$\bar{x}\$ 850 GeV \$\bar{x}\$ 850 GeV \$\bar{x}\$ 537 GeV \$\bar{x}\$ 537 GeV \$\bar{x}\$ 440 GeV \$\bar{x}\$ 1.0 TeV \$\bar{x}\$ 1.0 TeV	1.58 TeV 1.57 TeV	$\begin{split} & m(\tilde{k}_1^0) - 160 \ \text{MeV}, \tau(\tilde{k}_1^0) = 0.2 \ \text{ns} \\ & m(\tilde{k}_1^0) - 180 \ \text{MeV}, \tau(\tilde{k}_1^0) < 151 \ \text{ns} \\ & m(\tilde{k}_1^0) = 100 \ \text{GeV}, \ 10 \ \mu \text{s} < \tau(\tilde{k}) < 1600 \ \text{s} \\ & m(\tilde{k}_1^0) = 100 \ \text{GeV}, \ \tau > 10 \ \text{ns} \\ & 10 < \tan \theta < 50 \\ & 1 < \tau(\tilde{k}_1^0) < 30 \ \text{GeV}, \ \tau > 10 \ \text{ns} \\ & 10 < \tan \theta < 50 \\ & 1 < \tau(\tilde{k}_1^0) < 31 \ \text{ns}, \ \text{SPS8 model} \\ & 7 < \tau(\tilde{k}_1^0) < 740 \ \text{nm}, \ m(\tilde{k}) = 1.3 \ \text{TeV} \\ & 6 < c\tau(\tilde{k}_1^0) < 480 \ \text{nm}, \ m(\tilde{k}) = 1.1 \ \text{TeV} \end{split}$	ATLAS-CONF-2017-017 1506.05332 1310.6584 1606.05129 1604.04520 1411.6795 1409.5542 1504.05162 1504.05162	
$ \begin{array}{l} LFV pp \rightarrow \bar{v}_r + X, \bar{v}_r \rightarrow e \mu / e \tau / \mu \tau \\ Bilinear \ RPV \ CMSSM \\ \bar{x}_1^+ \bar{x}_1^-, \bar{x}_1^+ \rightarrow W \bar{x}_1^0, \bar{x}_1^0 \rightarrow e e \tau, e \mu \nu, \\ \bar{x}_1^+ \bar{x}_1^-, \bar{x}_1^+ \rightarrow W \bar{x}_1^0, \bar{x}_1^0 \rightarrow e \tau \nu, e \tau \nu \\ \bar{x}_{\bar{n}}, \bar{x}_{\bar{n}} \rightarrow q q q \\ \bar{x}_{\bar{n}}, \bar{x}_{\bar{n}} \rightarrow q q \bar{x}_1^0, \bar{x}_1^0 \rightarrow q q q \\ \bar{x}_{\bar{n}}, \bar{x}_{\bar{n}} \rightarrow t \bar{x}_1^0, \bar{x}_1^0 \rightarrow q q q \\ \bar{x}_{\bar{n}}, \bar{x}_{\bar{n}} \rightarrow t \bar{x}_1^0, \bar{x}_1^0 \rightarrow q q q \\ \bar{x}_{\bar{n}}, \bar{x}_{\bar{n}} \rightarrow t \bar{x}_1^0, \bar{x}_1 \rightarrow b s \\ \bar{i}_1 \bar{i}_1, \bar{i}_1 \rightarrow b s \\ \bar{i}_1 \bar{i}_1, \bar{i}_1 \rightarrow b \ell \end{array} $	$2e, \mu$ (SS) $\mu\mu\nu$ $4e, \mu$ 7 $3e, \mu + \tau$ 0 40 $41e, \mu 8$	0-3 b 4-5 large-R jd 4-5 large-R jd 8-10 jets/0-4 8-10 jets/0-4 2 jets + 2 b 2 b	jots - 4 b - 4 b -	3.2 20.3 13.3 20.3 14.8 14.8 36.1 36.1 15.4 36.1	\$\vec{r}\$, 4.8 \$\vec{x}_1^*\$ \$\vec{x}_1^*\$	eV 1.55 TeV 2.1 TeV 1.65 TeV	$\begin{split} &\mathcal{X}_{311}^{*}=0.11, \mathcal{X}_{132/133/233}=0.07 \\ &m(\tilde{q})=m(\tilde{g}), c\tau_{12,p}<1 \ mm \\ &m(\tilde{t}_{1}^{*})>400 GeV, \mathcal{X}_{12k}\neq0 \ (k=1,2) \\ &m(\tilde{t}_{1}^{*})>0.2 svm(\tilde{t}_{1}^{*}), \mathcal{X}_{133}\neq0 \\ &BR(t)=BR(t)=BR(t)=OF(t) \\ &m(\tilde{t}_{1}^{*})=BOO \ GeV \\ &m(\tilde{t}_{1}^{*})=BOO \ GeV \\ &m(\tilde{t}_{1}^{*})=I \ TeV, \mathcal{X}_{122}\neqO \\ &m(\tilde{t}_{1})=I \ TeV, \mathcal{X}_{323}\neqO \\ &BR(\tilde{t}_{1}\rightarrowbc/\mu)>20\% \end{split}$	1607.08079 1404.2500 ATLAS-CONF-2016-075 1405.5086 ATLAS-CONF-2016-057 ATLAS-CONF-2016-057 ATLAS-CONF-2017-013 ATLAS-CONF-2017-013 ATLAS-CONF-2017-013 ATLAS-CONF-2016-022, ATLAS-CONF-2016-084 ATLAS-CONF-2017-036	
er Scalar charm, $\tilde{c} \rightarrow c \tilde{\chi}_1^0$	0	2 c	Yes	20.3	2 510 GeV		m(k ⁰)<200 GeV	1501.01325	





multi-jet channel need very good modeling of background





Interactions.

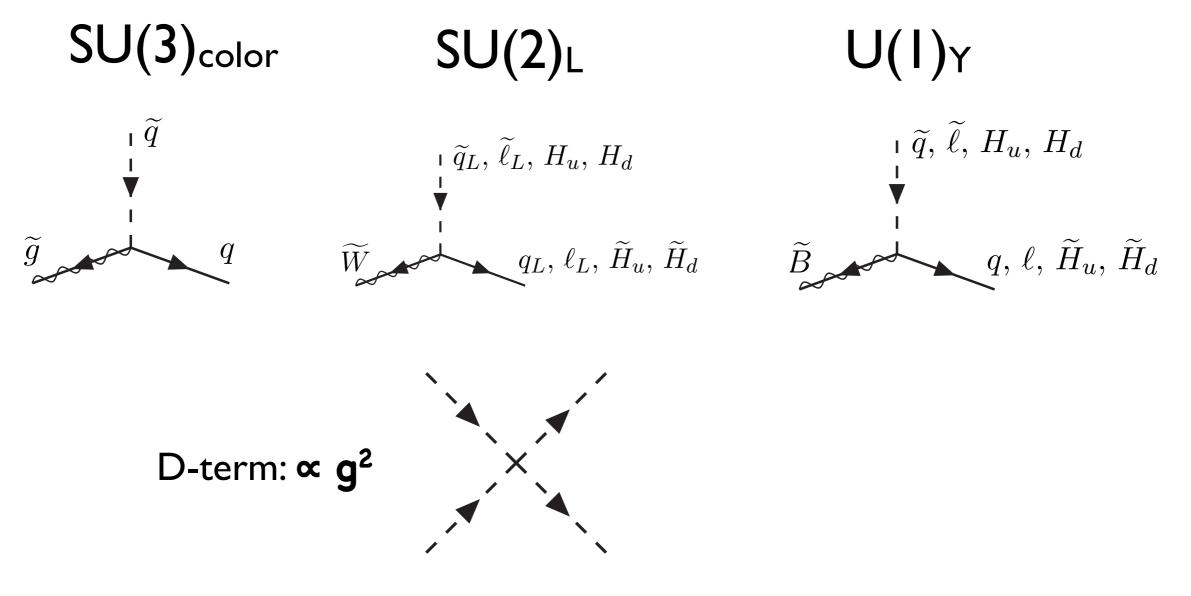
More details: for example, S. Martin "Supersymmetry Primer"

- Superpartners have the same gauge quantum numbers as their SM counter parts.
- Similar gauge interactions. \triangleright G_{μ}, W, Z, γ \overline{q} G_{μ} non-Abelian

Interactions.

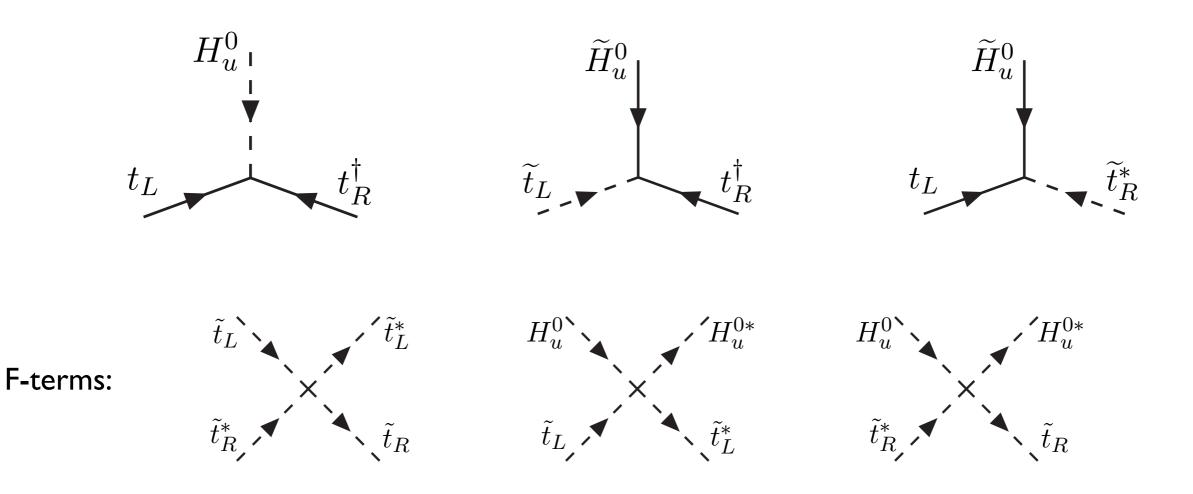
- SUSY \Rightarrow additional couplings

strength fixed by corresponding gauge



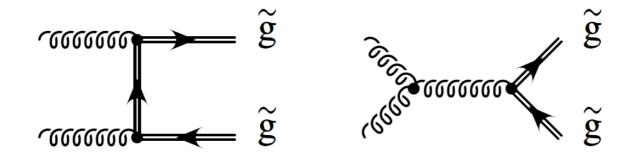
Interactions.

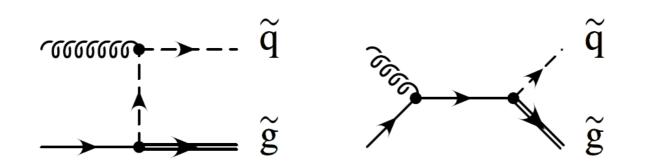
 SM fermions (such as the top quark) receive masses by coupling to the Higgs boson.



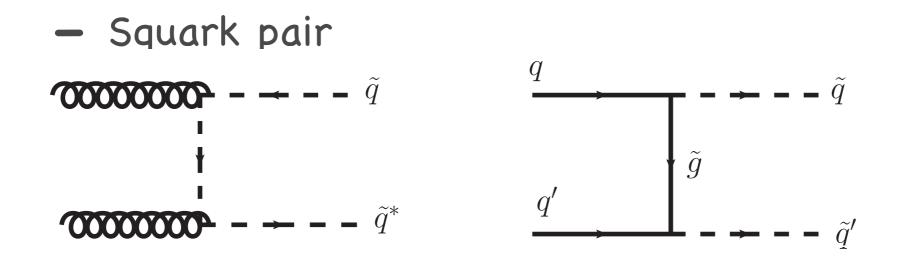
Examples of production: colored

• Squark and gluino production.



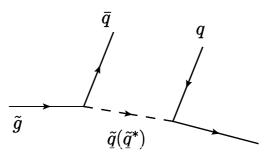


Examples of production

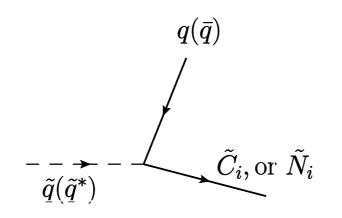


Decay of squark and gluino

- Gluino always decays into squark (on or off-shell).
 - Glunino -> squark + Jets

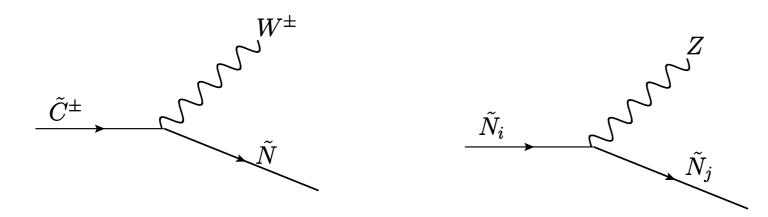


- Squark decay.
 - Jet +
 - To gluino, then go through off-shell squark.
 - To chargino or neutralino.



Next steps

• To W or Z (maybe Higgs.)



- Lepton (suppressed by W/Z-> lepton BR.)
 1 or 2 leptons.
- Jets (softer, constrained by W and Z mass).

Simple rules.

- Typically, there are many channels through which a superpartner can decay.
- 2 body mode (almost) always dominate over 3-body mode.

 \blacktriangleright A factor 1/100 suppression from phase space.

- Charge channel often bigger than the neutral channels.
- Higgsino prefers 3rd generation.
- Wino prefers left-handed. •
- Typically, only one or two modes dominates.
 - Signature easier to understand.

Exercise:

Choose a SUSY spectrum, such as one of the so called SNOWMASS Points and Slopes (SPS) benchmarks, <u>http://arxiv.org/abs/hep-ph/0202233</u>

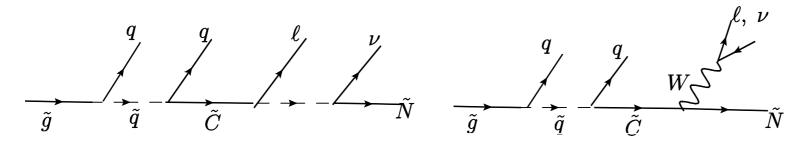
Use a spectrum and coupling calculator such as SUSPECT, SoftSUSY, or just PYTHIA... Understand the output.

Long decay chains

- Putting the pieces together.
- Many channels, many final states.

$$\underbrace{\tilde{g}}_{\tilde{q}} \underbrace{\tilde{q}}_{\tilde{N}} \underbrace{\tilde{q}}_{\tilde{N}} \underbrace{\tilde{q}}_{\tilde{N}} \underbrace{\tilde{g}}_{\tilde{q}} \underbrace{\tilde{q}}_{\tilde{N}} \underbrace{\tilde$$

2-lepton chain



1-lepton chain

$$\begin{split} \tilde{g} &\to q_1[\tilde{q}] \to q_1 q_2 \tilde{N}_0 \\ \tilde{g} &\to q_1[\tilde{q}] \to q_1 q_2 [\tilde{N}_i] \to q_1 q_2 [Z] \tilde{N}_0 \to q_1 q_2 q_3 q_4 \tilde{N}_0 \\ \tilde{g} &\to q_1[\tilde{q}] \to q_1 q_2 [\tilde{C}_i] \to q_1 q_2 [W] \tilde{N}_0 \to q_1 q_2 q_3 q_4 \tilde{N}_0 \\ \tilde{g} &\to q_1[\tilde{q}] \to q_1 q_2 [\tilde{N}_i] \to q_1 q_2 [Z] \tilde{N}_0 \to q_1 q_2 \ell^+ \ell^- \tilde{N}_0 \\ \tilde{g} &\to q_1[\tilde{q}] \to q_1 q_2 [\tilde{N}_i] \to q_1 q_2 q_3 q_4 (\ell^+ \ell^-) \tilde{N}_0 \end{split}$$

Exercise: draw diagrams for tri-lepton, same sign di-lepton

Typical variables I: counts.

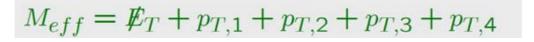
IVY A

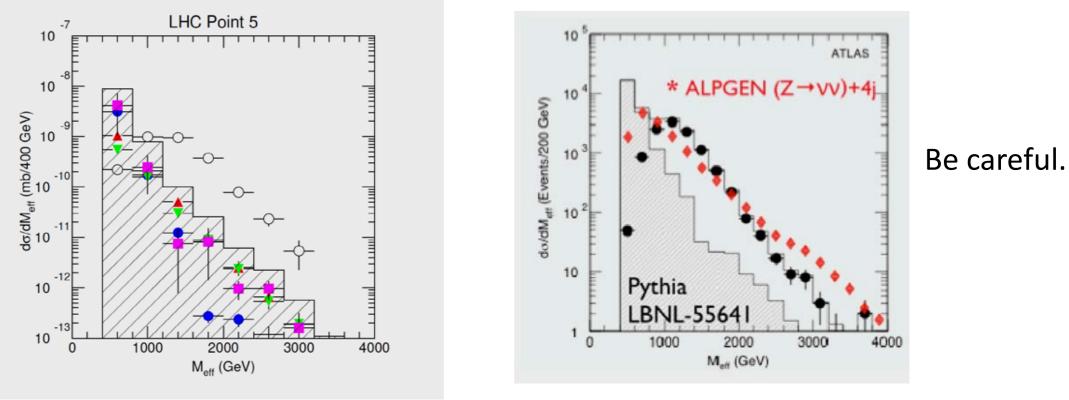
Inclusive counts. Useful for signal >> backrgound.

$n_j imes$ jet +	b-jet non-b-jet
$n_\ell imes$ lepton +	ℓ all flavor and charge combo: e.g. $2\ell \rightarrow 21$ comb.
$n_{\gamma} \times \gamma$	

Kinematical features: transverse variables.

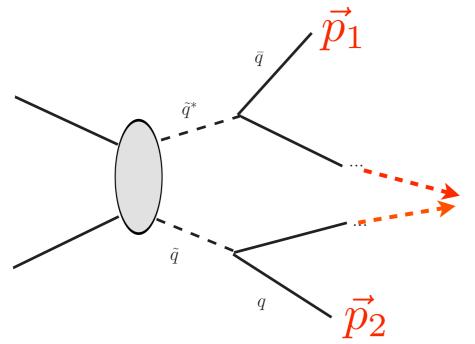
- Multiple hard objects.
- No resonance.
- Transverse variables made of several energetic objects. $M_{\rm eff}~H_{\rm T}$





Gianotti and Mangano, 2005

Another example: α_T



momenta labelled so that $p_{1T} \ge p_{2T}$

missing particles, total momentum $ec{p_3}$

$$\vec{p}_{1T} + \vec{p}_{2T} + \vec{p}_{3T} = 0$$

Define:
$$\alpha_T = \frac{p_{2T}}{m_T}$$
 $m_T = \sqrt{(p_{1T} + p_{2T})^2 - (\vec{p}_{1T} + \vec{p}_{2T})^2}$

Define p_T fractions
$$x_i = \frac{p_{iT}}{\sum_{i=1,3} p_{iT}}, x_i \le 1 \text{ and } \sum_{i=1,3} x_i = 2$$

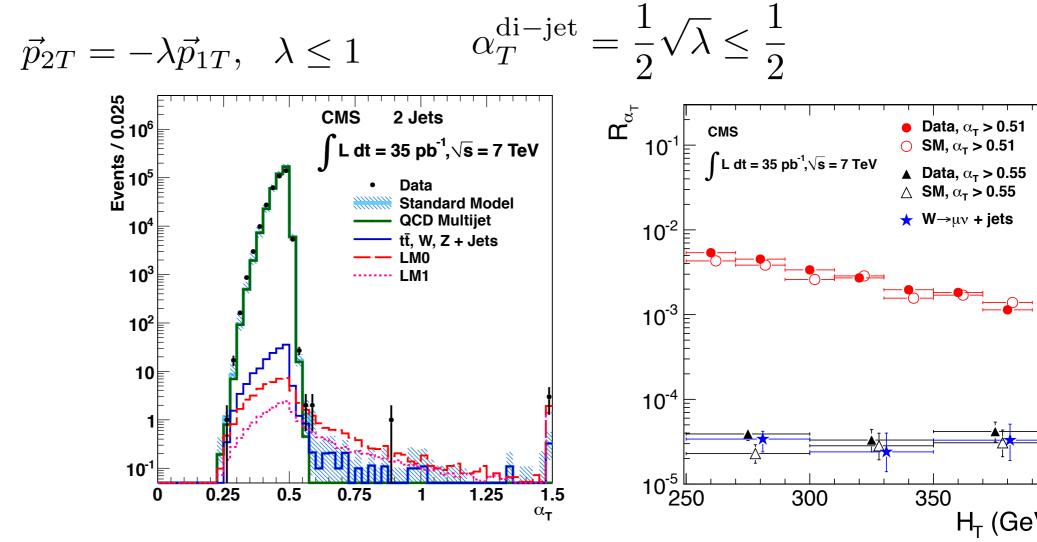
We obtain
$$\alpha_T = \frac{1}{2} \frac{x_2}{\sqrt{1 - x_3}}$$

 α_T can be either <1/2 (more often), or > 1/2

For a nice review, see Michael Peskin, "Razor and Scissors"

Another example: α_T

 In comparison, consider QCD di-jet, with one of the jet (say p_{2T}) energy miss measured.



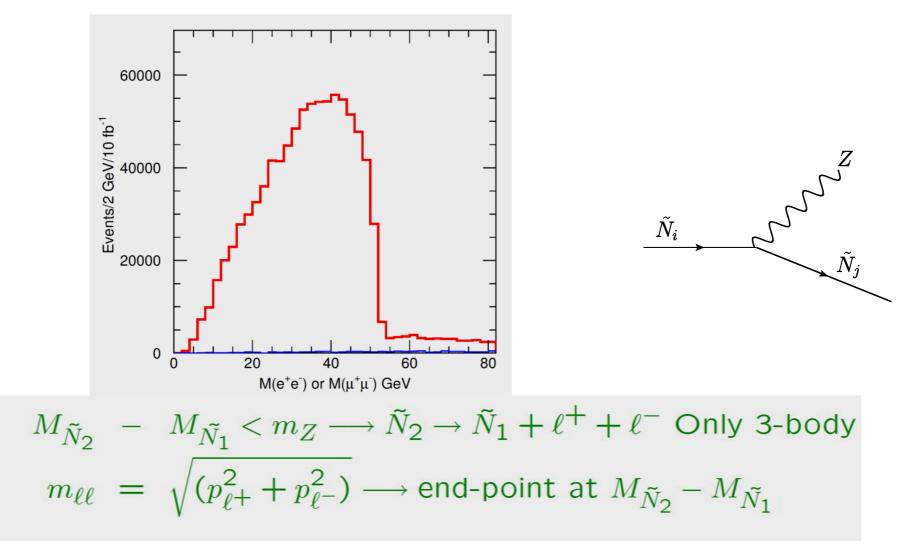
Many additional transverse variables: M_{T2} , Razor,

Kinematical variables: invariant masses

- Most useful: di-lepton edges and endpoints. (Mentioned earlier in neutralino decay).
 - Clean.
- Invariant mass distribution also carry spin information. Probably needs high statistics.
 For a review: See LW and J. Yavin, 2008
- More complicated invariant masses in longer decay chains possibly useful, but feature is less sharp. May need high statistics as well.

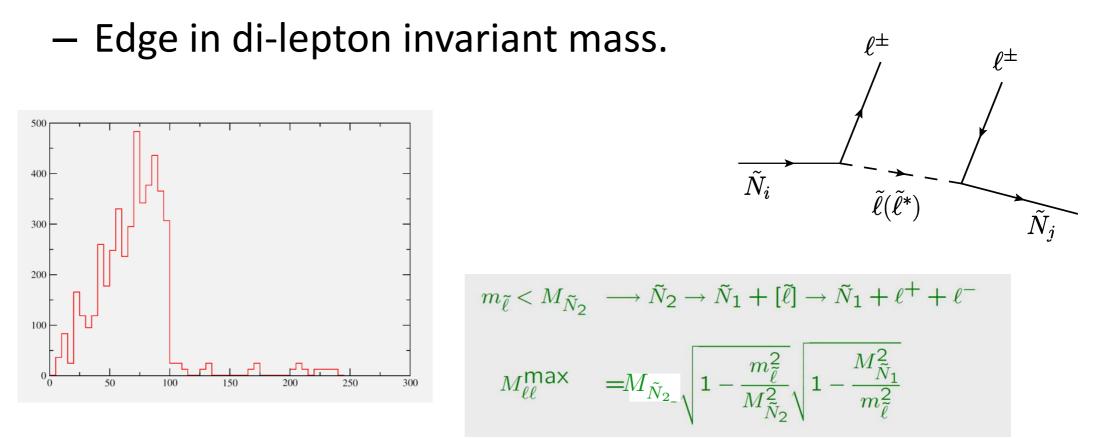
For example, see Miller and Osland. A set of papers.

- 3-body. End-point in di-lepton invariant mass.
 - Same flavor di-lepton.
 - Combinatorials can be suppressed with flavor subtraction.



More leptons if we are lucky

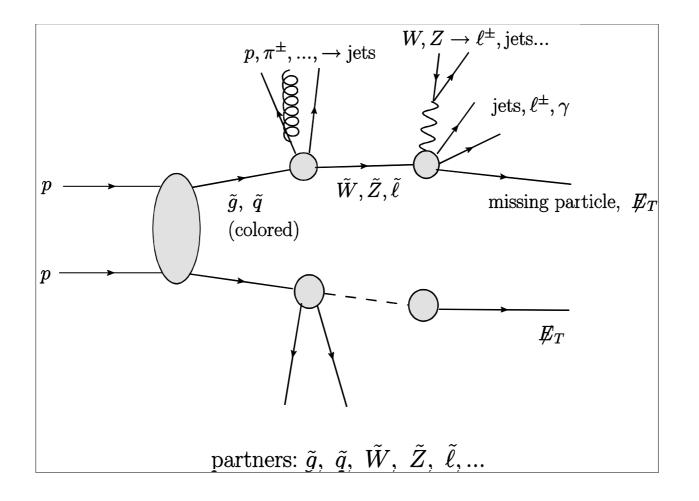
- A lot of leptons. No branching ratio suppression.
- On shell slepton, very distinctive feature.



More complicated edges useful, but need high statistics. See several papers by: Miller, Osland.



Topology: model independent approach



partners:

Same gauge interactions as the $ilde{g}, \ ilde{q}, \ ilde{W}, ilde{Z}, \ ilde{\ell}...$ SM particles Similar signatures.

 $q^{\text{KK}}, q^{\text{KK}}, W^{\text{KK}}, Z^{\text{KK}}, \ell^{\text{KK}}...$

http://indico.cern.ch/conferenceOtherViews.py?view=standard&confId=94910 http://www.lhcnewphysics.org/web/Overview.html

A promising, and complicated, scenario.

$$> \text{TeV} \qquad \underbrace{ \begin{array}{c} & \widetilde{u}, \ \widetilde{d}, \ \ldots \\ & \widetilde{t}, \ \widetilde{b} \end{array} }_{\tilde{t}}$$

$$p \ p \to \tilde{g}\tilde{g} \to t\bar{t}t\bar{t}(\text{or }t\bar{t}b\bar{b}, t\bar{t}t\bar{b} \dots)$$

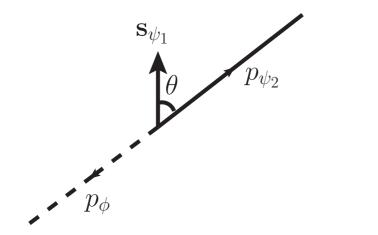
 $\tilde{g} \to t\bar{t}(b\bar{b}) + \tilde{N}, \text{ or }t\bar{b} + \tilde{C}^- \ t \to b\ell^+\nu$

The Dominant channel

- Multiple b, multiple lepton final state.
 - Good early discovery potential.
 - Challenging to interpret: top reconstruction
 A new method of fitting branching ratio to various final states
 Acharya, Grajek, Kane, Kuflik, Suruliz, Wang, arXiv:0901.3367

An example of a challenging measurement: spin or distinguishing SUSY with others.

Spin of new resonances

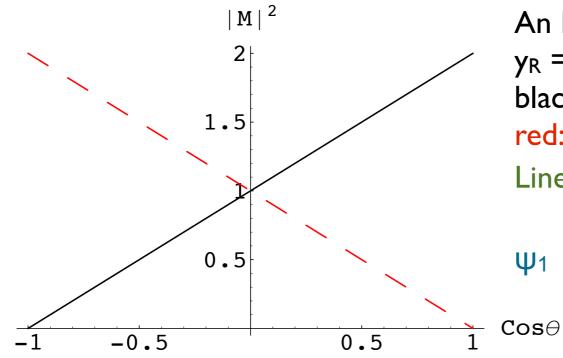


 $\psi_1 \rightarrow \psi_2 + \phi$

 $y_L\phi\bar{\psi}_2P_L\psi_1 + y_R\phi\bar{\psi}_2P_R\psi_1$

- Eample spin of fermion.
 - In the rest frame of the fermion.
 - Define angle θ of the decay product w.r.t. the polarization axis of ψ_1 .
 - Coupling could be chiral if $y_L \neq y_R$

Fermion spin

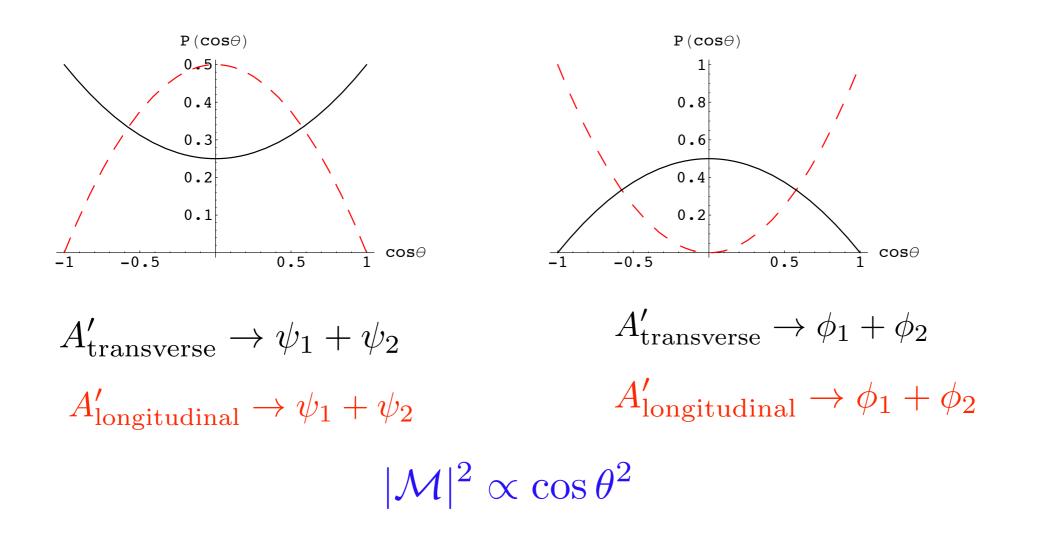


An Example $y_R = 0$ black: ψ_1 right-handed, red: ψ_1 left-handed Linear in $\cos\theta$

 ψ_1 not polized, no correlation, no spin information

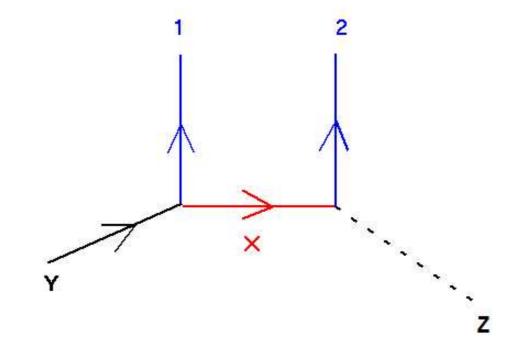
- Go to the rest frame.
- Coupling chiral.
- Ψ_1 polarized.

Spin-1



In general: $|\mathcal{M}|^2 \propto \cdots + \cos \theta^{2J_{\text{mother}}}$

Example of spin measurement



1 and 2 are observable particles, q, ℓ , W^{\pm}

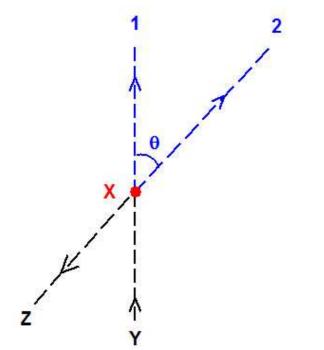
We are interested in the spin of X (on-shell).

We choose to use

$$t_{12} = (p_1 + p_2)^2.$$

In general, can not reconstruct the rest frame of X

Consider the rest frame of X



 $t_{12} \propto (1 - \cos \theta)^2$

Direction of $\,Y$ and 1 can be chosen to define the polarization of X For X with spin J_X

$$\frac{d\Gamma}{dt_{12}} = a \ t_{12}^{2J_X} + b \ t_{12}^{2J_X-1} + \cdots$$

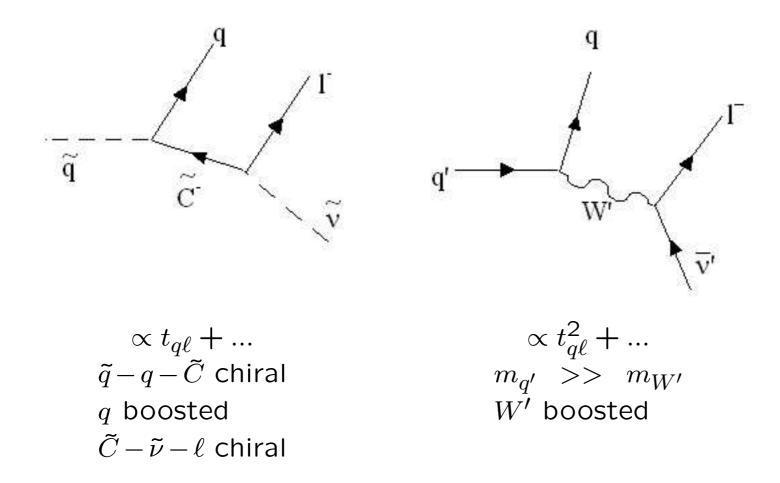
In principle, fitting the degree of this polynomial tells the the spin of X.

In practice, whether the coefficient a, b, ... are non-zero depends on the chirality of the coupling between X and I, 2, Z, Y, and the mass differences between them.

Interpreting the results correctly depending on our understanding the spectrum and couplings.

Example: SUSY vs spin-1 partner

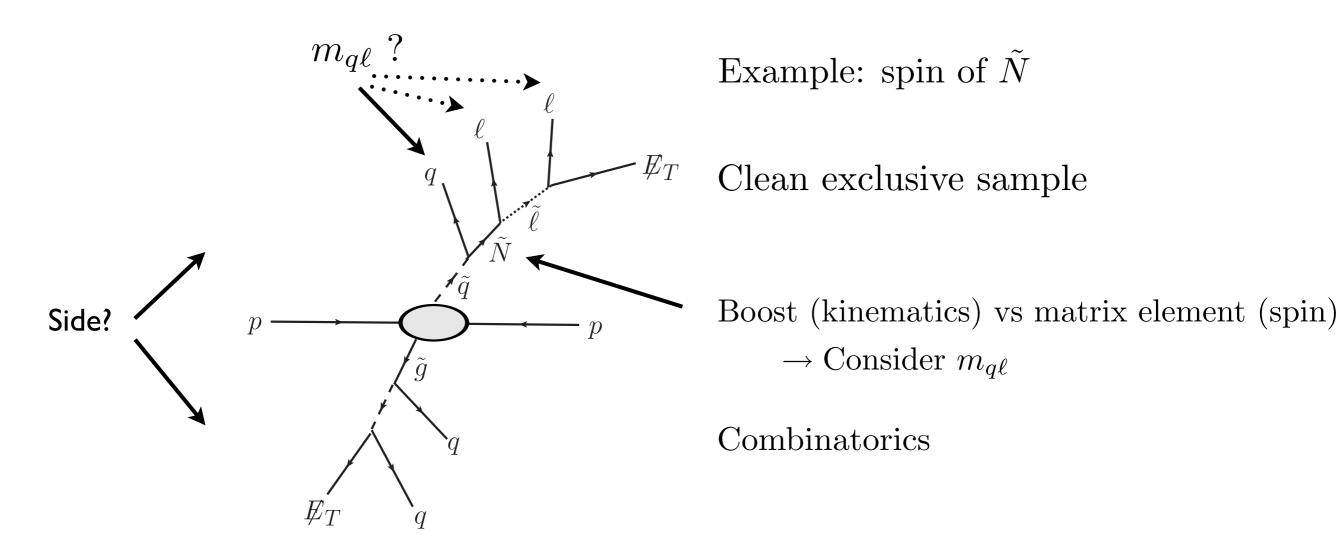
Decay through charged partners $\tilde{\chi}^{\pm}$, $W^{\prime\pm}$...



Usually there are more leptons in the decay chain.

Near/far lepton has to be separated.

Spin measurements. Supersymmetry?



- No universally applicable method. Different strategies will be used in different scenarios.
 A review: LTW and Yavin, arXiv:0802.2726
- More information of the signal, masses and underlying processes, is crucial.

Lepton colliders

- Fixed c.o.m.
- Much cleaner environment.
- Energy not as high.