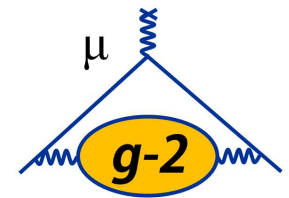


$(g-2)_\mu$ overview



Thomas Teubner



- Introduction
- QED and weak contributions
- a_μ^{had} : HVP (and HLbL) status, KNT17 HVP update
- BSM?
- Outlook

Introduction & motivation:

SM `too' successful, but incomplete:

- ν masses (small) and mixing point towards some high-scale (GUT) physics
- Need to explain dark matter & dark energy
- Not enough CP violation in the SM for matter-antimatter asymmetry
- And: $a_{\mu}^{\text{EXP}} - a_{\mu}^{\text{SM}}$ at $\sim 3-4 \sigma$ plus other deviations e.g. in the flavour sector

Is there a common New Physics (NP) explanation for all these puzzles?

- Uncoloured leptons are particularly clean probes to establish and constrain/distinguish NP, complementary to high energy searches at the LHC
- No direct signals for NP from LHC so far:
 - some models like CMSSM are in trouble already when trying to accommodate LHC exclusion limits and to solve muon $g-2$
 - is there any **TeV scale** NP out there? Or unexpected new **low scale** physics?

The key may be provided by **low energy observables**

including precision QED, EDMs and LFV → see talks by R. Szafron and A. de Gouvea

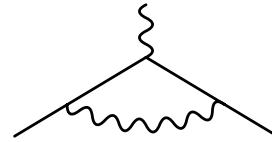
Introduction

- Dirac equation (1928): g is 2 for fundamental fermions $\vec{\mu} = g \frac{Qe}{2m} \vec{s}$
- 1947: small deviations from predictions in hydrogen and deuterium hyperfine structure; Kusch & Foley propose explanation with $g_s = 2.00229 \pm 0.00008$

- 1948: Schwinger calculates the famous radiative correction:

that $g = 2(1+a)$, with

$$a = (g-2)/2 = \alpha/(2\pi) = 0.001161$$



“ If you can’t join ‘em, beat ‘em “

This explained the discrepancy and was a crucial step in the development of perturbative QFT and QED

- The anomaly a (Anomalous Magnetic Moment) is from the Pauli term:

$$\delta \mathcal{L}_{\text{eff}}^{\text{AMM}} = -\frac{Qe}{4m} a \bar{\psi}(x) \sigma^{\mu\nu} \psi(x) F_{\mu\nu}(x)$$

This is a dimension 5 operator, non-renormalisable and hence not part of the fundamental (QED) Lagrangian. But it occurs through radiative corrections and is calculable in perturbation theory.

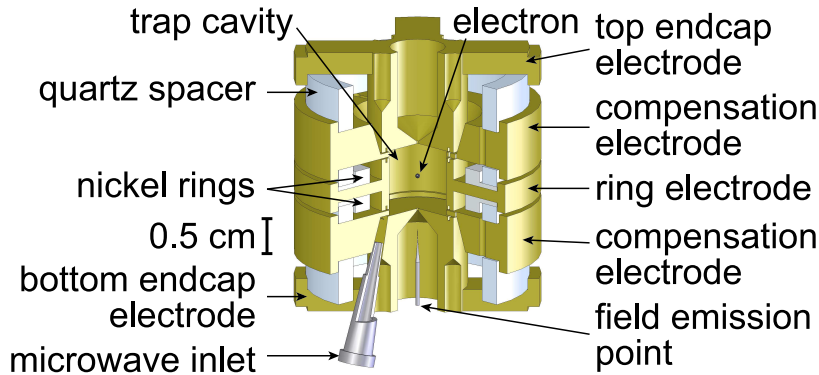
Magnetic Moments: a_e vs. a_μ

$$a_e = 1\,159\,652\,180.73 (0.28) \cdot 10^{-12} \quad [0.24\text{ppb}]$$

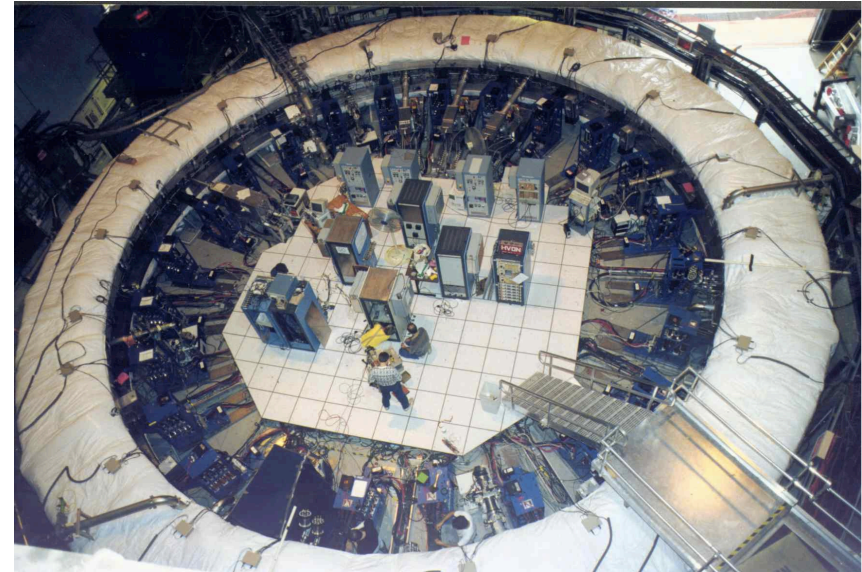
Hanneke, Fogwell, Gabrielse, PRL 100(2008)120801

$$a_\mu = 116\,592\,089(63) \cdot 10^{-11} \quad [0.54\text{ppm}]$$

Bennet et al., PRD 73(2006)072003



one electron quantum cyclotron



- a_e^{EXP} more than 2000 times more precise than a_μ^{EXP} , but for e^- loop contributions come from very small photon virtualities, whereas muon `tests' higher scales

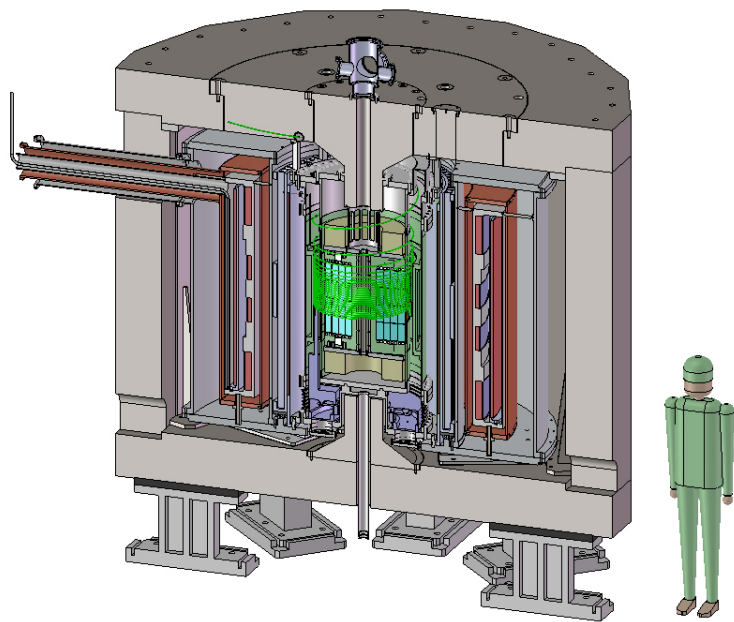
- dimensional analysis: **sensitivity to NP** (at high scale Λ_{NP}): $a_\ell^{\text{NP}} \sim C m_\ell^2 / \Lambda_{\text{NP}}^2$

→ μ wins by $m_\mu^2 / m_e^2 \sim 43000$ for NP, but a_e provides best determination of α

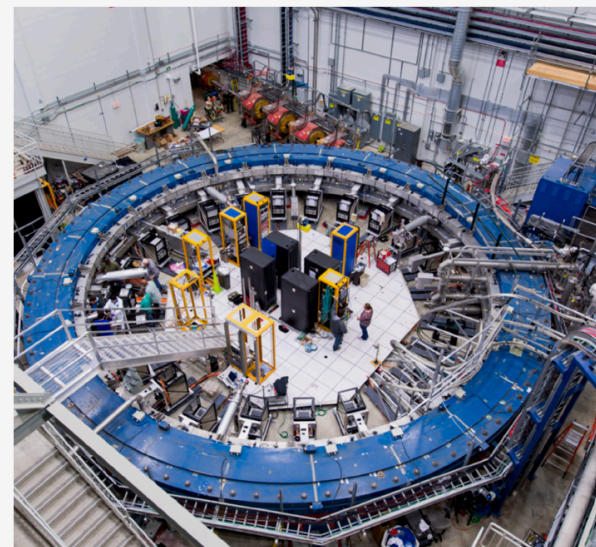
a_μ : back to the future

- CERN started it nearly 40 years ago
- Brookhaven delivered 0.5ppm precision
- E989 at FNAL and J-PARC's $g-2$ /EDM experiments are happening and should give us certainty

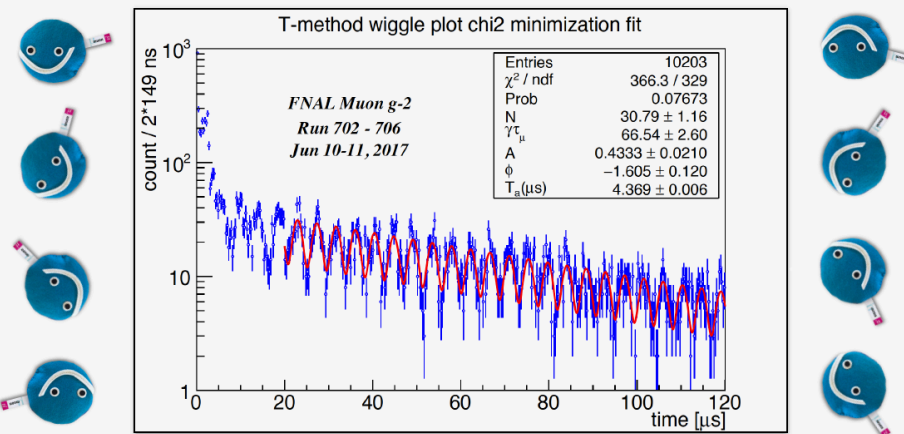
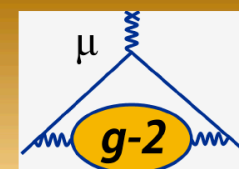
→ talks by Lee Roberts and Tsutomu Mibe



JUNE
12
2017



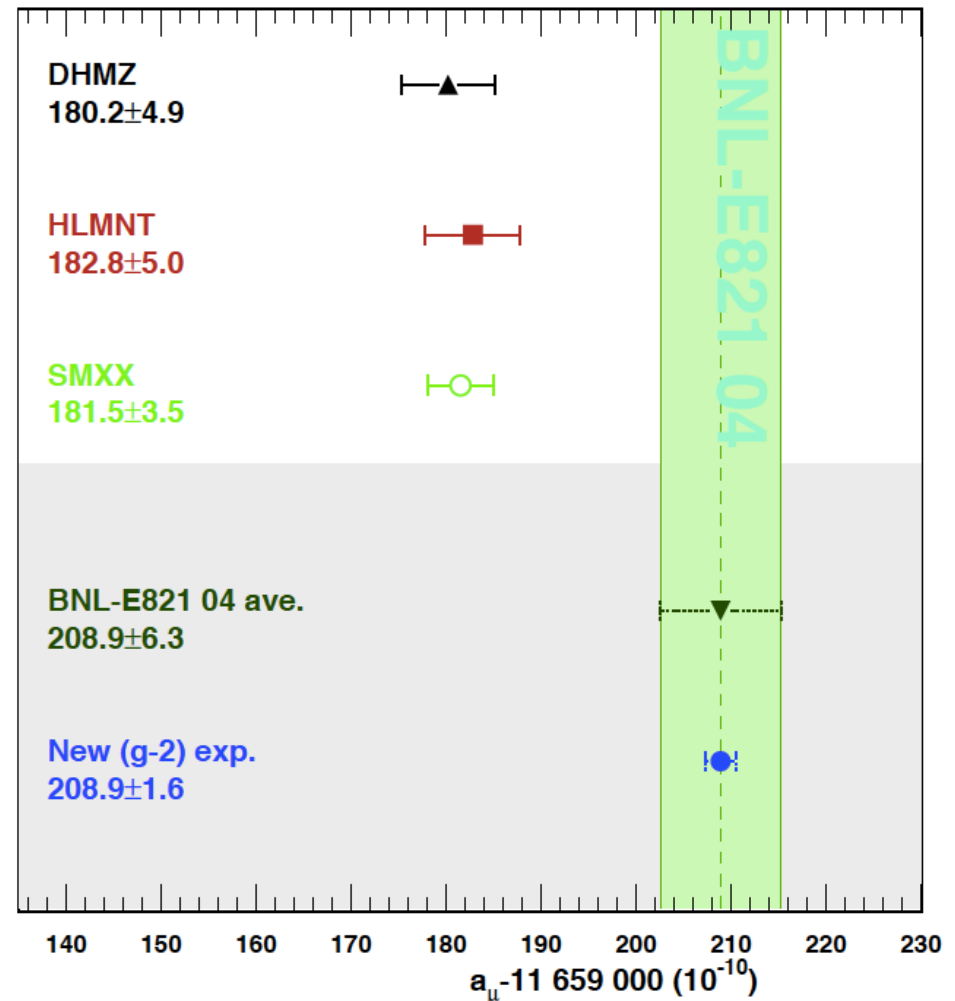
**MUON G-2
WIGGLE PARTY**
5 pm @ Fermilab Users' Center



a_μ : Status and future projection → charge for SM TH

$$a_\mu = a_\mu^{\text{QED}} + a_\mu^{\text{EW}} + a_\mu^{\text{hadronic}} + a_\mu^{\text{NP?}}$$

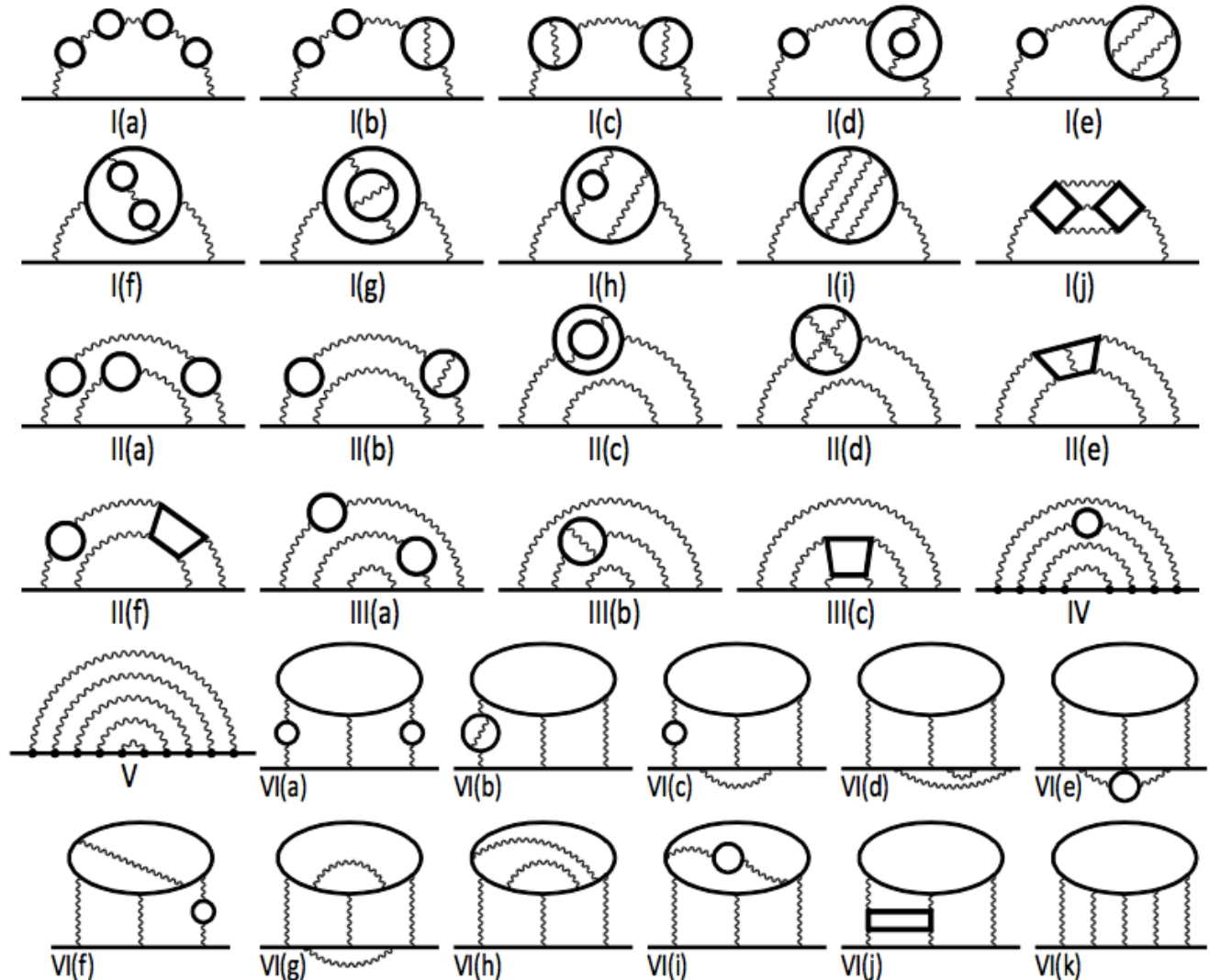
- if mean values stay and with **no** a_μ^{SM} improvement:
5 σ discrepancy
- if also EXP+TH can improve a_μ^{SM}
'as expected' (consolidation of L-by-L on level of Glasgow consensus, about factor 2 for HVP): NP at 7-8 σ
- or, if mean values get closer, very strong exclusion limits on many NP models (extra dims, new dark sector, xxxSSSM)...



T. Aoyama, M. Hayakawa,
T. Kinoshita, M. Nio (PRLs, 2012)

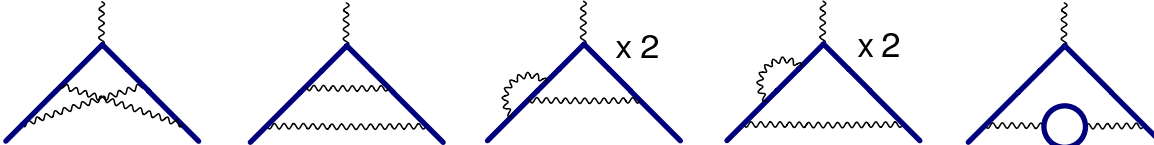
A triumph for perturbative QFT and computing!

10th
12672
diagrams



- code-generating code, including renormalisation
- multi-dim. numerical integrations

- **Schwinger 1948:** 1-loop $a = (g-2)/2 = \alpha/(2\pi) = 116\,140\,970 \times 10^{-11}$

- 2-loop graphs:
 

- 72 3-loop and 891 4-loop diagrams ...

- **Kinoshita et al. 2012:** 5-loop completed numerically (12672 diagrams):

$$a_{\mu}^{\text{QED}} = 116\,584\,718.951 (0.009) (0.019) (0.007) (0.077) \times 10^{-11}$$

errors from: lepton masses, 4-loop, 5-loop, α from ^{87}Rb

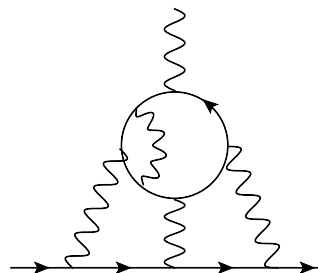
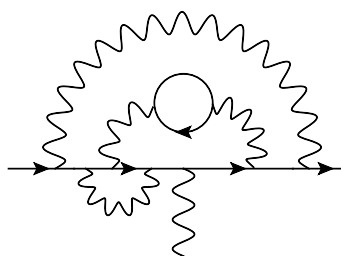
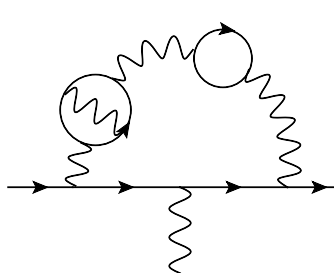
- QED extremely accurate, and the series is stable: $a_{\mu}^{\text{QED}} = C_{\mu}^{2n} \sum_n \left(\frac{\alpha}{\pi}\right)^n$

$$C_{\mu}^{2,4,6,8,10} = 0.5, 0.765857425(17), 24.05050996(32), 130.8796(63), 753.29(1.04)$$

- Could a_{μ}^{QED} still be wrong?

Some classes of graphs known analytically ([Laporta](#); [Aguilar, Greynat, deRafael](#)),

- ... but 4-loop and 5-loop rely heavily on numerical integrations
- Recently several independent checks of 4-loop and 5-loop diagrams:
Baikov, Maier, Marquard [NPB 877 (2013) 647], Kurz, Liu, Marquard, Smirnov AV+VA, Steinhauser [NPB 879 (2014) 1, PRD 92 (2015) 073019, 93 (2016) 053017]:
- all 4-loop graphs with internal lepton loops now calculated independently, e.g.

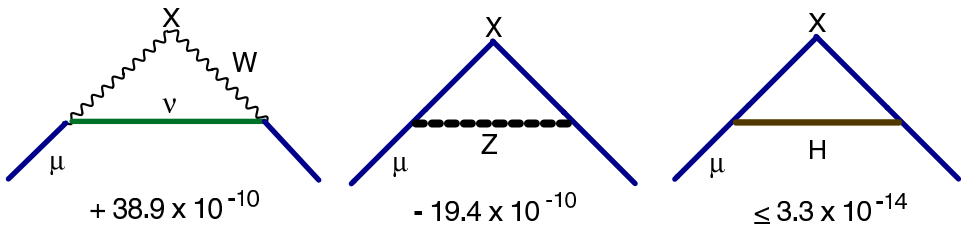


(from Steinhauser et al., PRD 93 (2016) 053017)

→ talk by Matthias Steinhauser

- 4-loop universal (massless) term calculated semi-analytically to 1100 digits (!) by Laporta, arXiv:1704.06996, also new numerical results by Volkov, 1705.05800
- all agree with Kinoshita et al.'s results, so **QED is on safe ground** ✓

- Electro-Weak 1-loop diagrams:



$$a_\mu^{\text{EW}(1)} = 195 \times 10^{-11}$$

- known to 2-loop (1650 diagrams, the first EW 2-loop calculation):
Czarnecki, Krause, Marciano, Vainshtein; Knecht, Peris, Perrottet, de Rafael

- agreement, a_μ^{EW} relatively small, 2-loop relevant: $a_\mu^{\text{EW}(1+2 \text{ loop})} = (154 \pm 2) \times 10^{-11}$

- Higgs mass now known, update by Gnendiger, Stoeckinger, S-Kim,

PRD 88 (2013) 053005

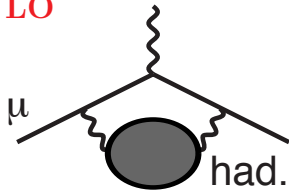
$$a_\mu^{\text{EW}(1+2 \text{ loop})} = (153.6 \pm 1.0) \times 10^{-11} \quad \checkmark$$

compared with $a_\mu^{\text{QED}} = 116\,584\,718.951(80) \times 10^{-11}$

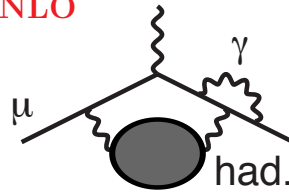
- Hadronic: **non-perturbative**, the limiting factor of the SM prediction **X** \Rightarrow **✓**

$$a_\mu^{\text{had}} = a_\mu^{\text{had,VP LO}} + a_\mu^{\text{had,VP NLO}} + a_\mu^{\text{had,Light-by-Light}}$$

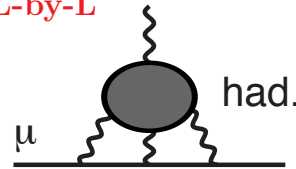
LO



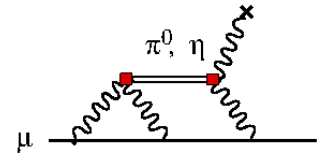
NLO



L-by-L



e.g.



- L-by-L**: - so far use of **model calculations** (+ form-factor data and pQCD constraints), \rightarrow overview talk by A. Nyffeler
 - also good news from **lattice QCD**, and \rightarrow talks by H. Wittig and C. Lehner
 - new **dispersive** approaches \rightarrow talks by B. Kubis, V. Pauk and G. Colangelo

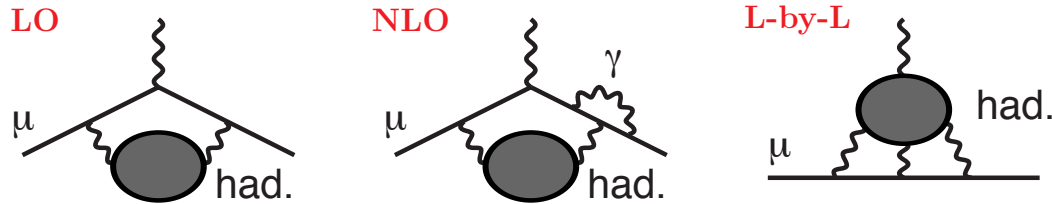
- Below I will use the '**updated Glasgow consensus**':
(original by Prades+deRafael+Vainshtein)

$$a_\mu^{\text{had,L-by-L}} = (98 \pm 26) \times 10^{-11}$$

- so far no indication for a big surprise
- expect that L-by-L prediction can be improved further
- with new results & progress, tell politicians/sceptics: L-by-L can be predicted!**

$a_\mu^{\text{had, VP}}$: Hadronic Vacuum Polarisation

$$a_\mu^{\text{had}} = a_\mu^{\text{had, VP LO}} + a_\mu^{\text{had, VP NLO}} + a_\mu^{\text{had, Light-by-Light}}$$

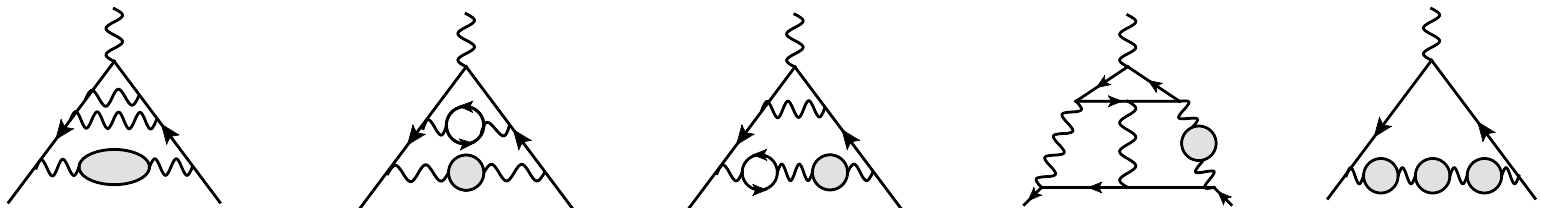


HVP: - most precise prediction by using e^+e^- hadronic cross section (+ tau) data and well known dispersion integrals → for space-like HVP see talk by L. Trentadue

- done at LO and NLO (see graphs) → see HVP talks by Z. Zhang and F. Jegerlehner

- and recently at NNLO [Steinhauser et al., PLB 734 (2014) 144, also F. Jegerlehner]

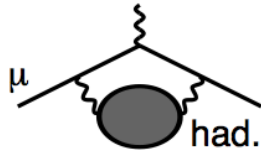
$a_\mu^{\text{HVP, NNLO}} = + 1.24 \times 10^{-10}$ not so small, from e.g.:



- Alternative: lattice QCD, but need QED and iso-spin breaking corrections
Lots of activity by several groups, errors coming down, QCD+QED started

Hadronic Vacuum Polarisation, essentials:

Use of data compilation for HVP:



pQCD not useful. Use the **dispersion relation** and the **optical theorem**.

$$\text{had.} = \int \frac{ds}{\pi(s-q^2)} \text{Im had.}$$

$$2 \text{Im had.} = \sum_{\text{had.}} \int d\Phi \left| \text{had.} \right|^2$$

$$a_{\mu}^{\text{had,LO}} = \frac{m_{\mu}^2}{12\pi^3} \int_{s_{\text{th}}}^{\infty} ds \frac{1}{s} \hat{K}(s) \sigma_{\text{had}}(s)$$

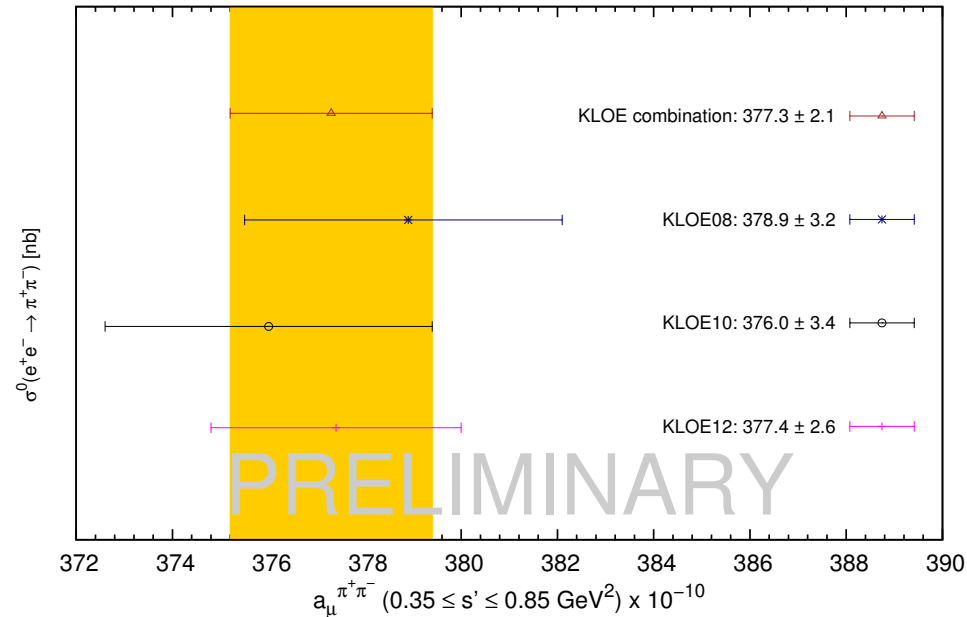
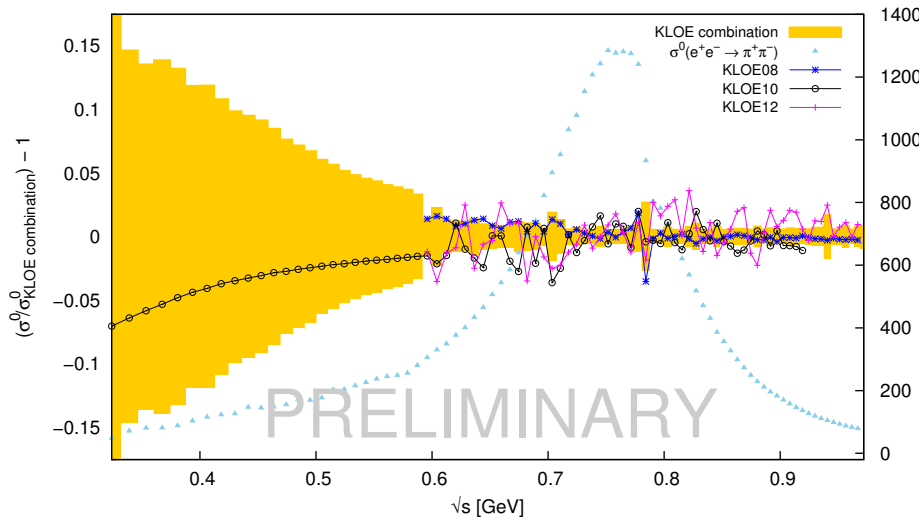
- Weight function $\hat{K}(s)/s = \mathcal{O}(1)/s$
 \Rightarrow **Lower** energies **more important**
 $\Rightarrow \pi^+\pi^-$ channel: 73% of total $a_{\mu}^{\text{had,LO}}$

How to get the most precise σ_{had}^0 ? **e^+e^- data**:

- Low energies: **sum ~ 30 exclusive channels**, $2\pi, 3\pi, 4\pi, 5\pi, 6\pi, KK, KK\pi, KK\pi\pi, \eta\pi, \dots$, use iso-spin relations for missing channels
- Above ~ 1.8 GeV: can start to use **pQCD** (away from flavour thresholds), supplemented by narrow resonances ($J/\psi, \Upsilon$)
- Challenge of **data combination (locally in \sqrt{s})**: many experiments, different energy bins, stat+sys errors from different sources, **correlations**; must avoid **inconsistencies/bias**
- traditional '**direct scan**' (tunable e^+e^- beams) vs. '**Radiative Return**' [+ τ spectral functions]
- σ_{had}^0 means 'bare' σ , but WITH FSR: **RadCorrs** [HLMNT '11: $\delta a_{\mu}^{\text{had, RadCor VP+FSR}} = 2 \times 10^{-10}$!]

HVP: KLOE 2 π combination [preliminary]

⇒ Combination of KLOE08, KLOE10 and KLOE12 gives 85 distinct bins between $0.1 \leq s \leq 0.95 \text{ GeV}^2$



→ Covariance matrix now correctly constructed

⇒ a **positive semi-definite matrix**

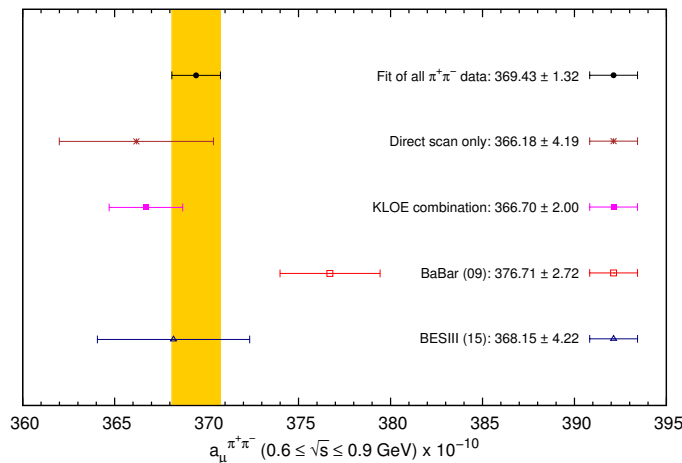
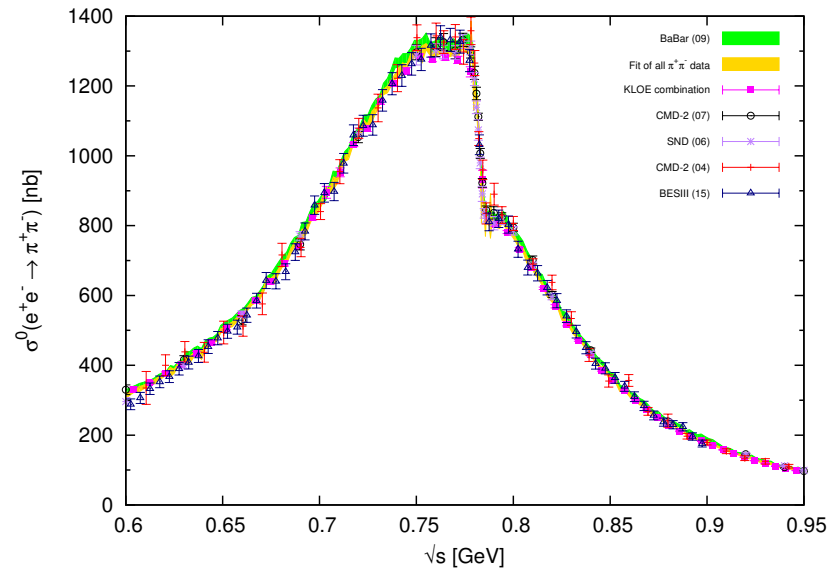
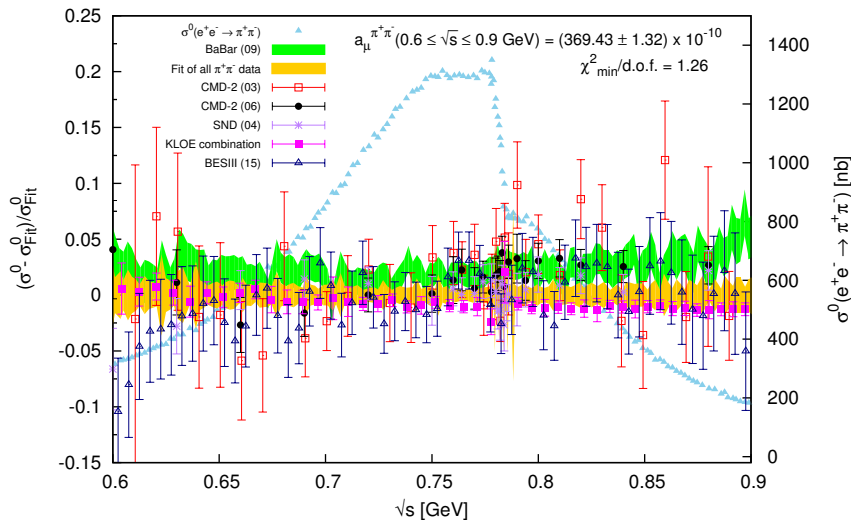
→ **Non-trivial influence of correlated uncertainties** on resulting mean value

$$a_{\mu}^{\pi^{+}\pi^{-}}(0.1 \leq s' \leq 0.95 \text{ GeV}^2) = (489.9 \pm 2.0_{\text{stat}} \pm 4.3_{\text{sys}}) \times 10^{-10}$$

HVP: complete 2π combination by Keshavarzi+Nomura+T

⇒ Large improvement for 2π estimate

→ BESIII [Phys.Lett. B753 (2016) 629-638] and KLOE combination provide **downward influence** to mean value



⇒ **Correlated & experimentally corrected** $\sigma_{\pi\pi}^0(\gamma)$ data now entirely **dominant**

$a_\mu^{\pi^+\pi^-}$ ($0.305 \leq \sqrt{s} \leq 2.00$ GeV):

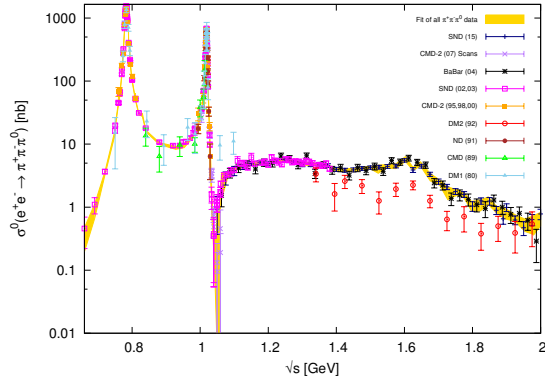
HLMNT11: 505.77 ± 3.09

KNT17: 502.85 ± 1.93 (!!)

(no radiative correction uncertainties)

HVP: other notable exclusive channels

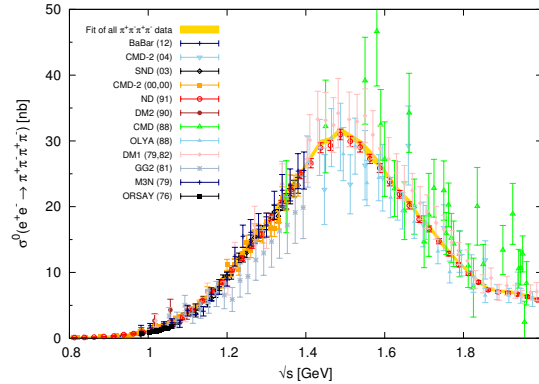
$$\pi^+ \pi^- \pi^0$$



HLMNT11: 47.51 ± 0.99

KNT17: 47.68 ± 0.70

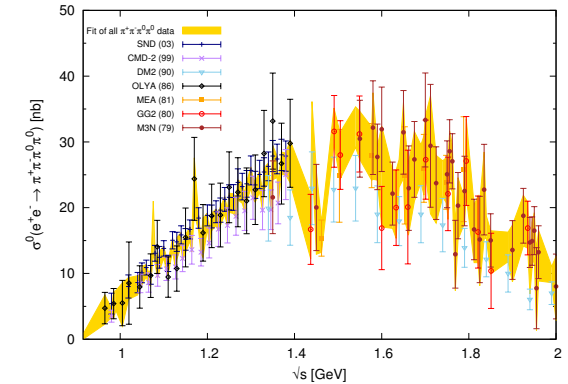
$$\pi^+ \pi^- \pi^+ \pi^-$$



HLMNT11: 14.65 ± 0.47

KNT17: 15.18 ± 0.14

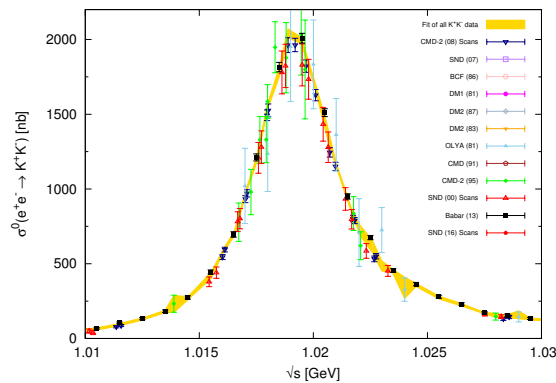
$$\pi^+ \pi^- \pi^0 \pi^0$$



HLMNT11: 20.37 ± 1.26

KNT17: 20.07 ± 1.19

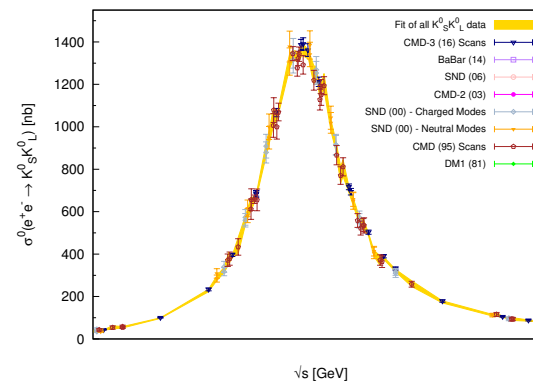
$$K^+ K^-$$



HLMNT11: 22.15 ± 0.46

KNT17: 22.76 ± 0.22

$$K_S^0 K_L^0$$

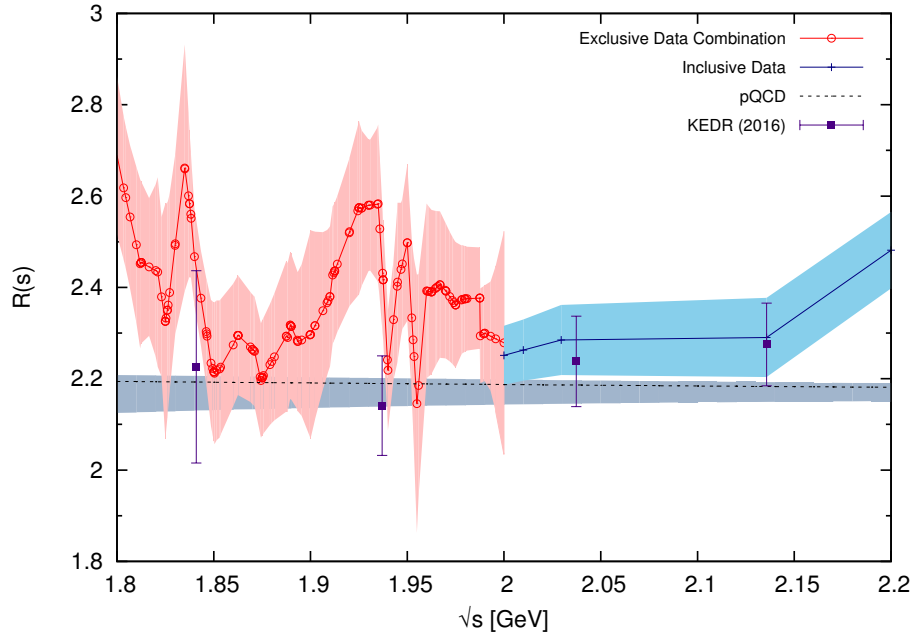


HLMNT11: 13.33 ± 0.16

KNT17: 13.09 ± 0.12

HVP: region of inclusive data/pQCD

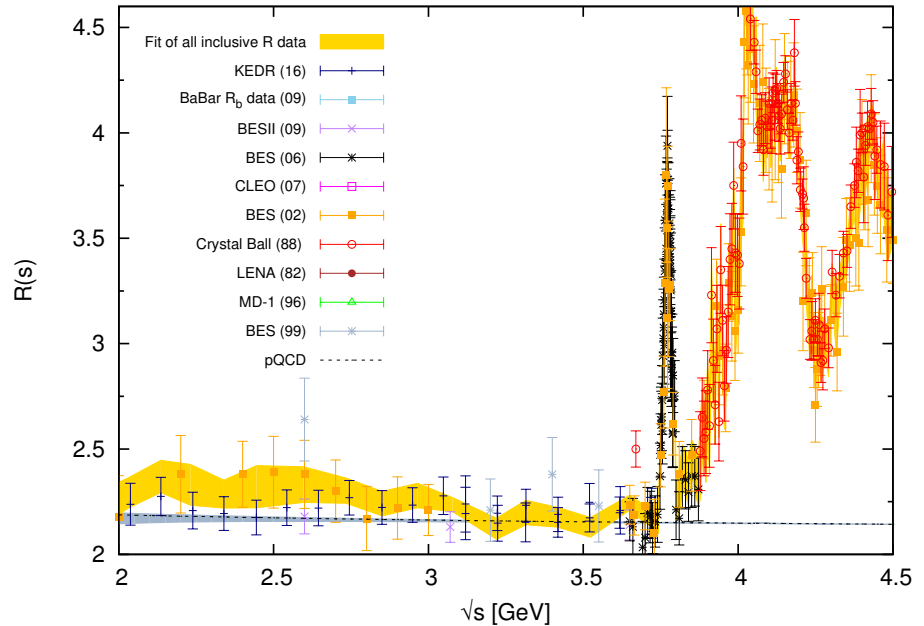
⇒ **New KEDR inclusive R data** ranging $1.84 \leq \sqrt{s} \leq 3.05$ GeV [Phys.Lett. B770 (2017) 174-181] and $3.12 \leq \sqrt{s} \leq 3.72$ GeV [Phys.Lett. B753 (2016) 533-541]



$a_{\mu}^{\text{had, LOVP}} (1.84 \leq \sqrt{s} \leq 2.00 \text{ GeV}):$

pQCD : 6.42 ± 0.03

Data : 6.88 ± 0.25



$a_{\mu}^{\text{had, LOVP}} (2.60 \leq \sqrt{s} \leq 3.73 \text{ GeV}):$

pQCD (inflated errors) : 10.82 ± 0.38

Data : 11.20 ± 0.14

⇒ **Choose to adopt entirely data driven estimate from threshold to 11.2 GeV**

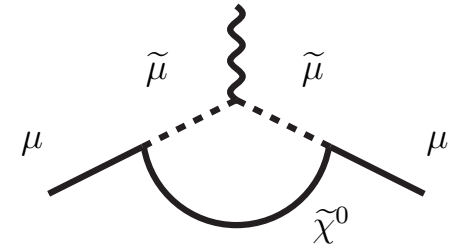
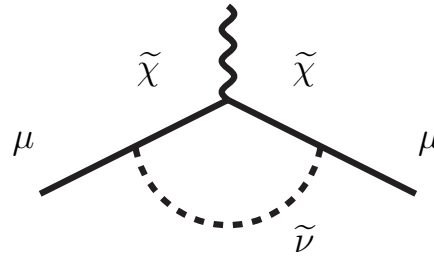
a_μ^{SM} : update HLMNT11 \rightarrow KNT17 presented @ TGM2

	<u>2011</u>	\rightarrow	<u>2017</u>	*to be discussed
QED	11658471.81 (0.02)	\rightarrow	11658471.90 (0.01)	[Phys. Rev. Lett. 109 (2012) 111808]
EW	15.40 (0.20)	\rightarrow	15.36 (0.10)	[Phys. Rev. D 88 (2013) 053005]
LO HLbL	10.50 (2.60)	\rightarrow	9.80 (2.60)	[EPJ Web Conf. 118 (2016) 01016]*
NLO HLbL			0.30 (0.20)	[Phys. Lett. B 735 (2014) 90]*
<hr/>				
	<u>HLMNT11</u>		<u>KNT17</u>	
LO HVP	694.91 (4.27)	\rightarrow	692.23 (2.54)	this work*
NLO HVP	-9.84 (0.07)	\rightarrow	-9.83 (0.04)	this work*
<hr/>				
NNLO HVP			1.24 (0.01)	[Phys. Lett. B 734 (2014) 144] *
<hr/>				
Theory total	11659182.80 (4.94)	\rightarrow	11659181.00 (3.62)	this work
Experiment			11659209.10 (6.33)	world avg
Exp - Theory	26.1 (8.0)	\rightarrow	28.1 (7.3)	this work
<hr/>				
Δa_μ	3.3 σ	\rightarrow	3.9 σ	this work

a_μ : New Physics?

- Many BSM studies use $g-2$ as constraint or even motivation
- SUSY could easily explain $g-2$

- Main 1-loop contributions:



- Simplest case:

$$a_\mu^{\text{SUSY}} \simeq \text{sgn}(\mu) 130 \times 10^{-11} \tan \beta \left(\frac{100 \text{ GeV}}{\Lambda_{\text{SUSY}}} \right)^2$$

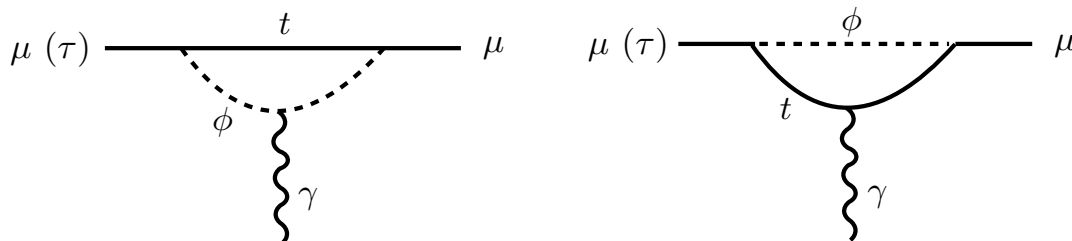
- Needs $\mu > 0$, 'light' SUSY-scale Λ and/or large $\tan \beta$ to explain 281×10^{-11}
- This is already excluded by LHC searches in the simplest SUSY scenarios (like CMSSM); causes large χ^2 in simultaneous SUSY-fits with LHC data and $g-2$
- However:
 - * SUSY does not have to be minimal (w.r.t. Higgs),
 - * could have large mass splittings (with lighter sleptons),
 - * be hadrophobic/leptophilic,
 - * or not be there at all, but don't write it off yet...

New Physics? just a few of many recent studies

- Don't have to have full MSSM (like coded in GM2Calc [by Athron, ..., Stockinger et al., EPJC 76 (2016) 62], which includes all latest two-loop contributions), and
 - **extended Higgs sector** could do, see, e.g. Stockinger et al., JHEP 1701 (2017) 007, 'The muon magnetic moment in the 2HDM: complete two-loop result'
- lesson: 2-loop contributions can be highly relevant in both cases; one-loop analyses can be misleading

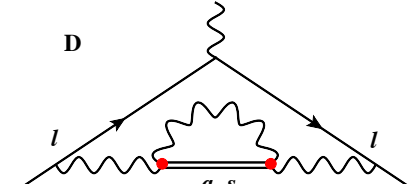
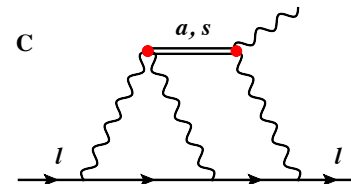
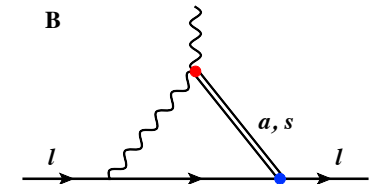
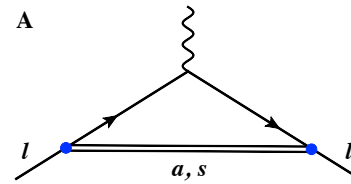
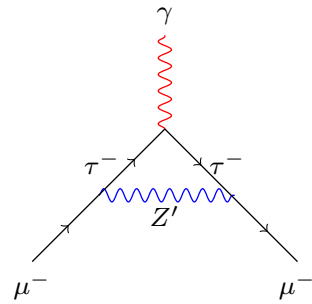
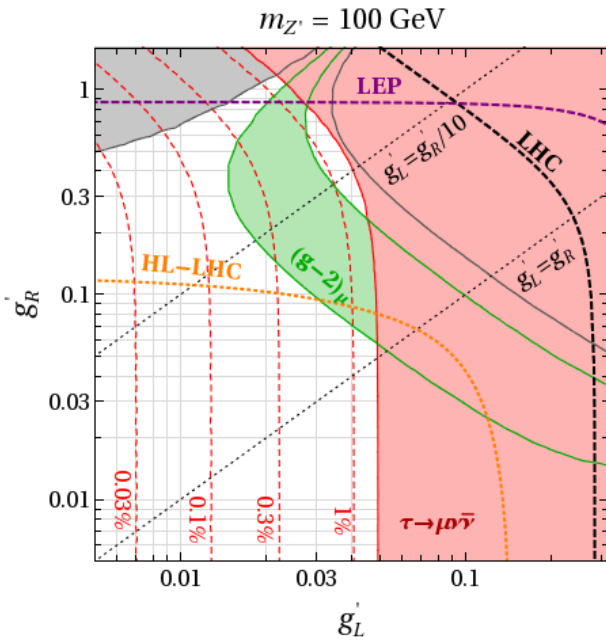
- **1 TeV Leptoquark** Bauer + Neubert, PRL 116 (2016) 141802

one new scalar could explain several anomalies seen by BaBar, Belle and LHC in the flavour sector (e.g. **violation of lepton universality** in $B \rightarrow K\ell\ell$, enhanced $B \rightarrow D\tau\nu$) and solve $g-2$, while satisfying all bounds from LEP and LHC



New Physics? just a few of many recent examples

- **light Z'** can evade many searches involving electrons by non-standard couplings preferring heavy leptons (but see BaBar's direct search limits in a wide mass range, PRD 94 (2016) 011102), or invoke flavour off-diagonal Z' to evade constraints [Altmannshofer et al., PLB 762 (2016) 389]



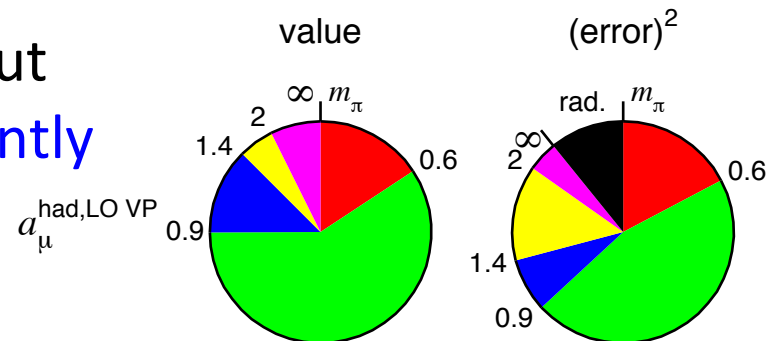
- **axion-like particle (ALP)**, contributing like π^0 in HLbL [Marciano et al., PRD 94 (2016) 115033]
- **'dark photon'** - like fifth force particle [Feng et al., PRL 117 (2016) 071803]

→ see talks by M. Pospelov, A. Filippi, I. Jaegle

Conclusions/Outlook:

- All sectors of the **Standard Model prediction** of $g-2$ have been **scrutinised** a lot in recent years

- The basic picture has not changed, but **recent data**, many from ISR, **significantly improve the prediction for a_μ^{HVP}**



- **Discrepancy $\sim 3-4\sigma$ is consolidated**
- With **further hadronic data** in the pipeline, also on FFs for HLbL, and **efforts from lattice**, the goal of **squeezing Δa_μ^{SM} is in reach**
- Many approaches to explain the discrepancy with NP, linking $g-2$ with other precision observables, the flavour sector, dark matter and direct searches, but so far NP only (con)strained...