The dilepton results at beam energy scan from STAR Electron-positron tomography of hot, dense Medium

> Lijuan Ruan (Brookhaven National Laboratory)



- Electron-positron tomography
- Chiral symmetry
- Our experimental approach and results
- Summary

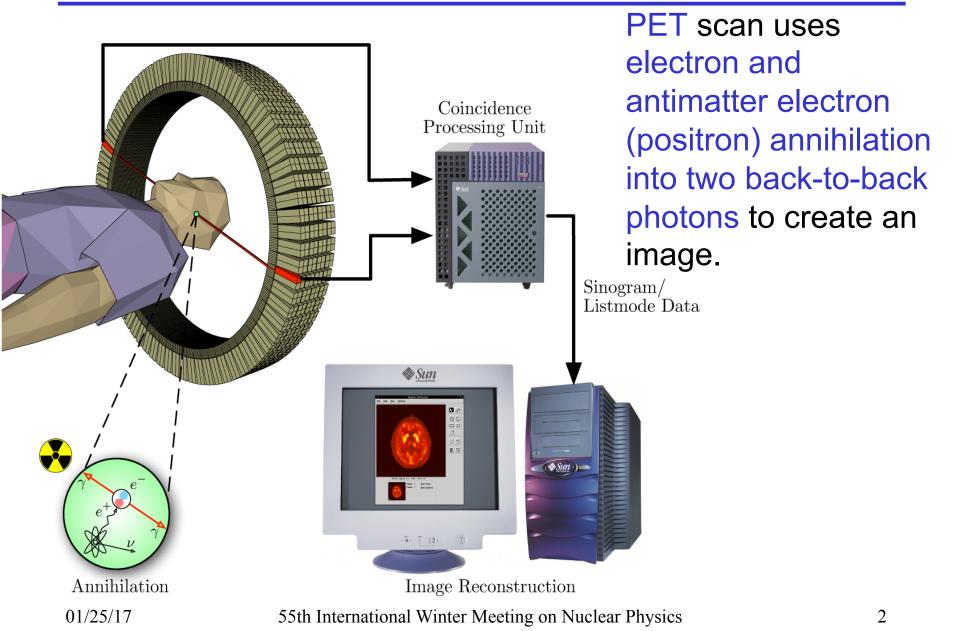


a passion for discovery



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Traditional Positron-emission Tomography (PET)

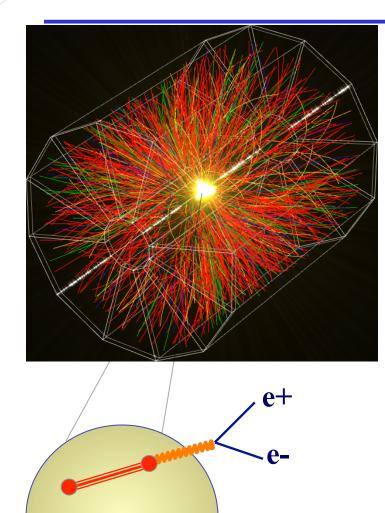




Rapp

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EM probe: electron-positron tomography



- In our method, we detect electron and positron pairs from quark-antiquark annihilation.
- Electron-positron pairs are penetrating probes and can provide information deep into the system and early time.
- Using electron-positron tomography, we would like to study the symmetry of the Quark-Gluon Plasma.

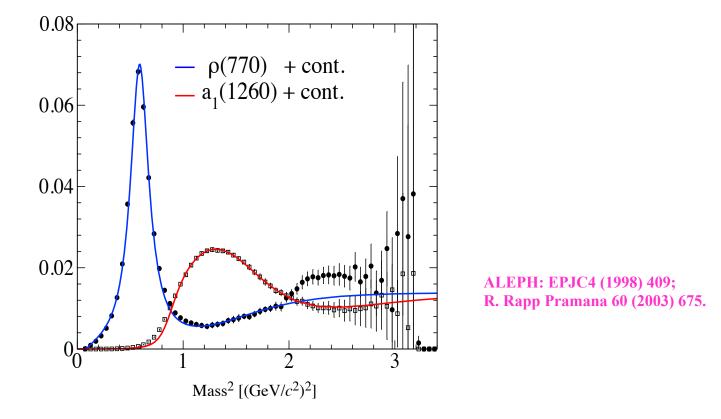


Microscopic picture:

- quark condensate: left-handed quark and righthanded antiquark attract each other through the exchange of gluons. Generate 99% of visible mass in the universe.
- electron condensate: electrons attract each other through the vibration of the crystal at low temperature. Generate superconductivity in the metal.



ρ and a1 resonance (spectrum function) in vacuum

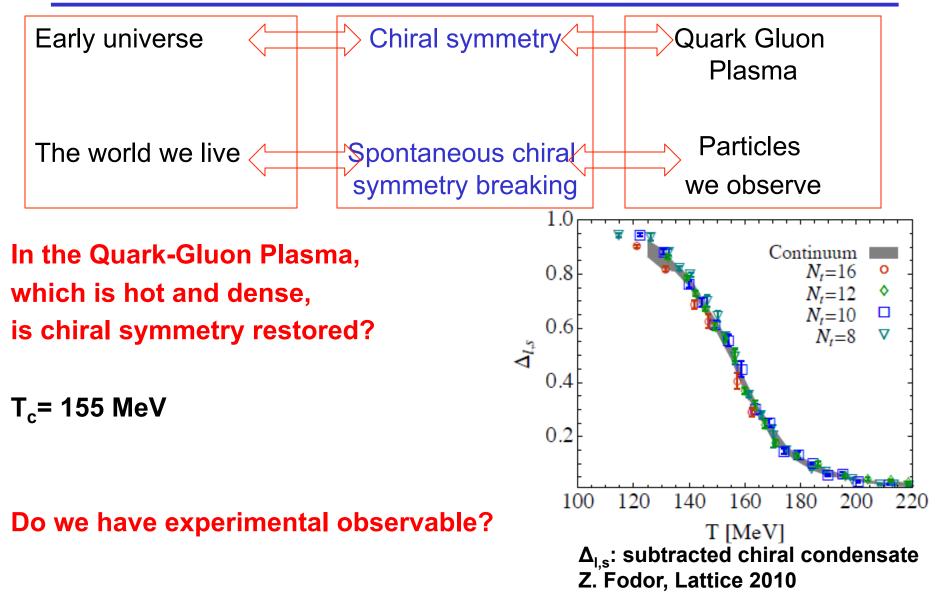


Spontaneous chiral symmetry breaking: mass distributions are different

Chiral symmetry restoration: mass difference disappears

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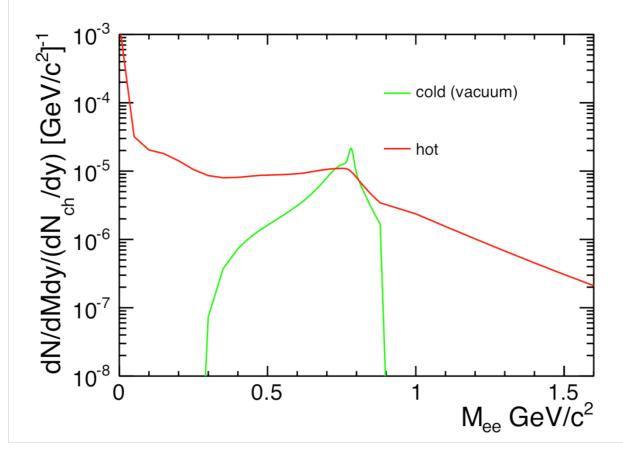
Is chiral symmetry restored in Quark-Gluon Plasma?



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The p resonance mass spectrum function



Observable for chiral symmetry restoration:

a modified (broadened) p spectral function

Model: Rapp & Wambach, priv. communication Adv. Nucl.Phys. 25, 1 (2000); Phys. Rept. 363, 85 (2002)

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BROOKHAVEN RHIC @ Brookhaven National Laboratory





A heavy-ion collision event



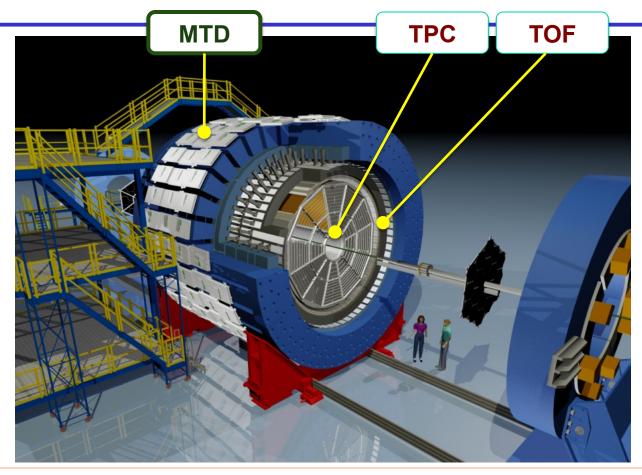




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The STAR Detector

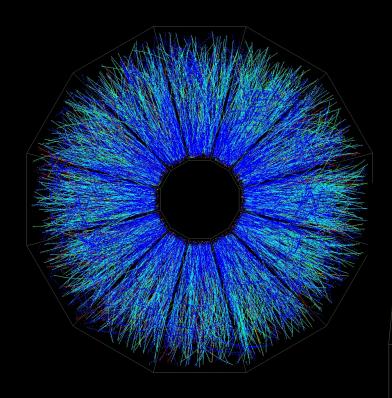


<u>Solenoidal Tracker at RHIC (1200 tons)</u> Time Projection Chamber

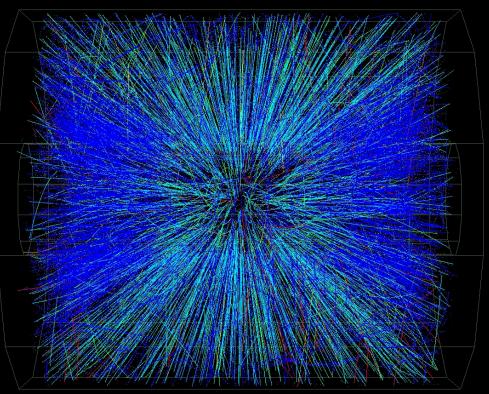
- 1. Second largest device of its kind ever built
- 2. Measure ionization energy loss (dE/dx) and momentum

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¹⁹⁷Au + ¹⁹⁷Au Collisions at RHIC



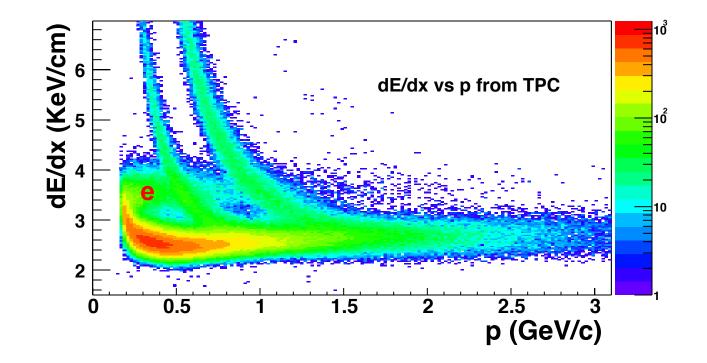
Central Event







Particle identification

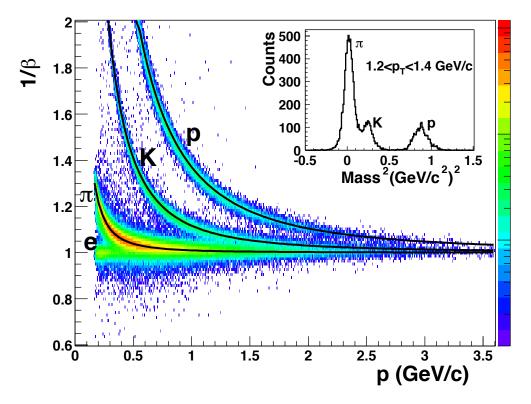


Electrons are highly contaminated by other particles.

Need new experimental tool to clearly identify electrons!

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BROOKHAVEN Particle identification from Time of Flight Detector



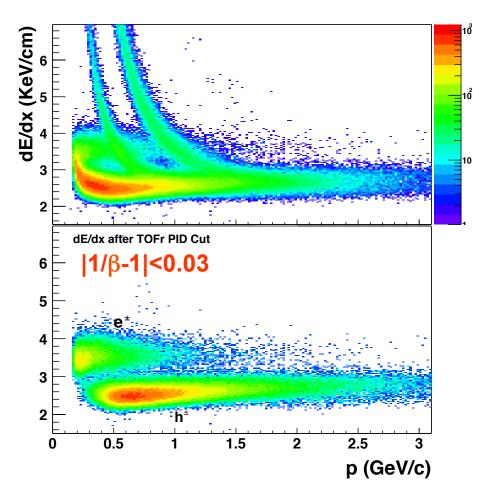
STAR Collaboration, PLB616(2005)8

Hadron identification: proton up to 3 GeV/c, kaon and pion up to 1.6 GeV/c

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Electron identification

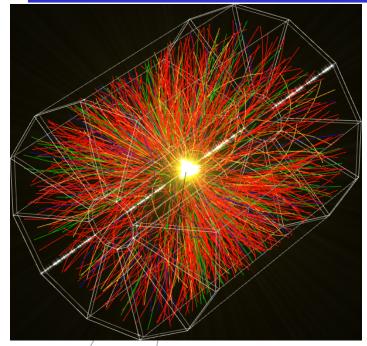


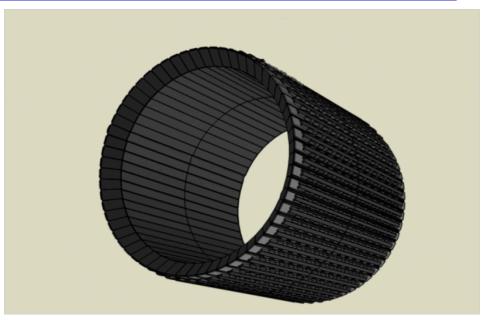
Combining information from the TPC and TOF, we obtain clean electron samples at $p_T < 3$ GeV/c.

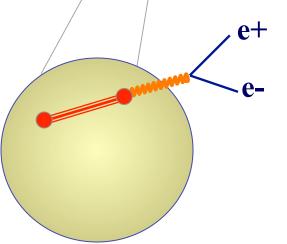
STAR Collaboration, PRL94(2005)062301



The electron-positron tomography tools





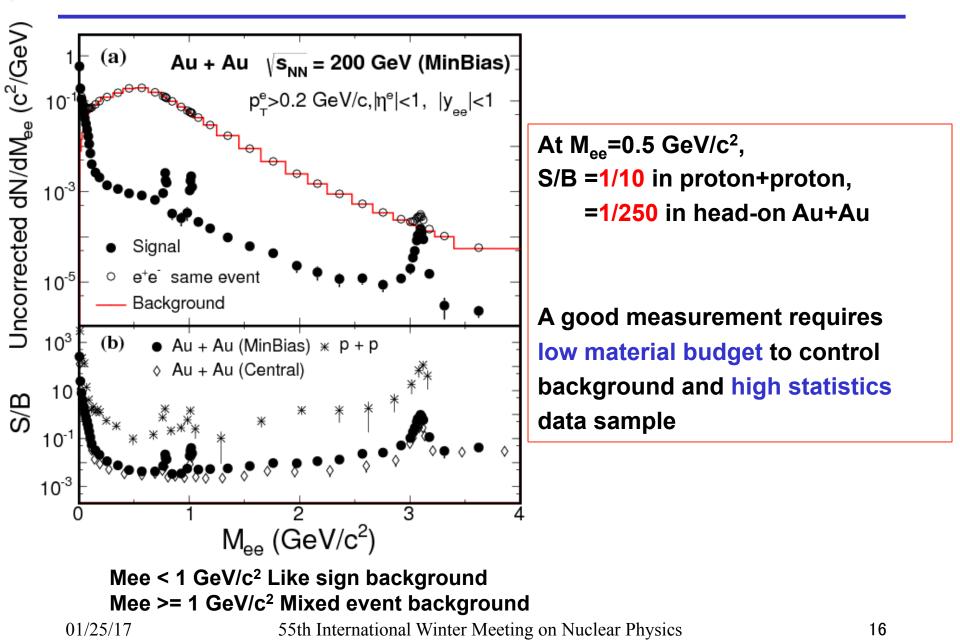


The Time of Flight Detector completes the experimental tool for electron-positron tomography: clean electron identification and large acceptance.

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Electron-positron invariant mass distribution



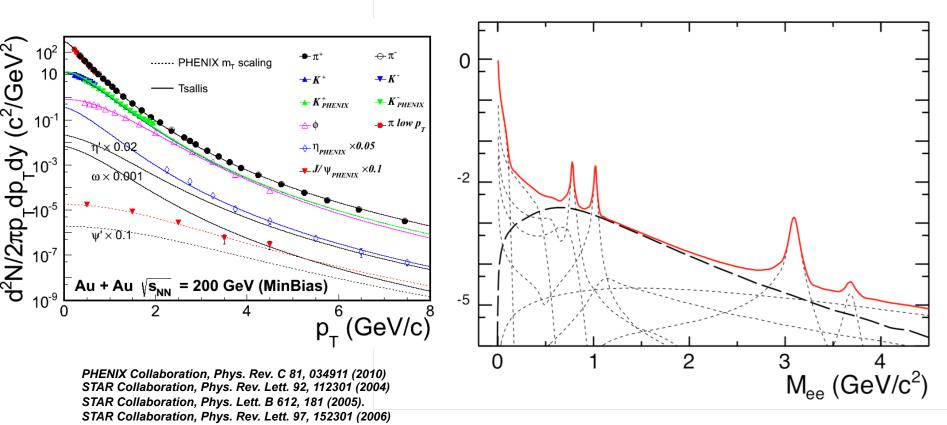


Electron-positron signal: e+e- pairs from light flavor meson and heavy flavor decays (charmonia and open charm correlation): Pseudoscalar meson Dalitz decay: π^0 , η , $\eta' \rightarrow \gamma e^+e^-$ Vector meson decays: ρ^0 , ω , $\phi \rightarrow e^+e^-$, $\omega \rightarrow \pi^0 e^+e^-$, $\phi \rightarrow \eta e^+e^-$ Heavy flavor decays: $J/\psi \rightarrow e^+e^-$, $ccbar \rightarrow e^+e^- X$, bbbar $\rightarrow e^+e^- X$ Drell-Yan contribution

In Au+Au collisions, we search for QGP thermal radiation at 1.1<M_{ee}<3.0 GeV/c² (intermediate mass range) Vector meson in-medium modifications at M_{ee}<1.1 GeV/c² (low mass range)



Electron-positron emission mass spectrum



Z. Tang et al. Phys. Rev. C 79, 051901 (2009)

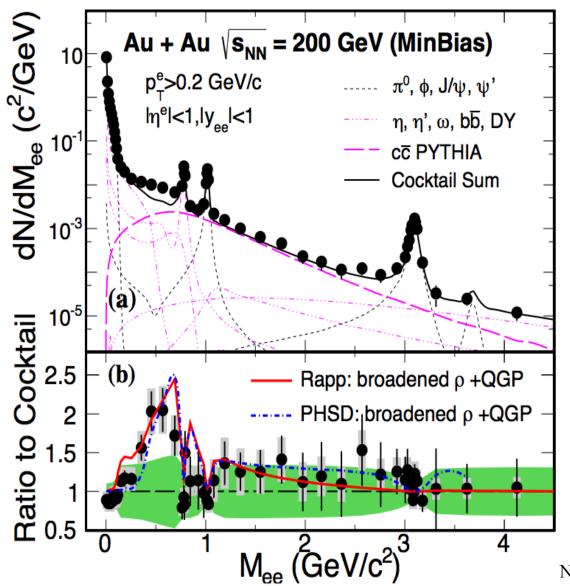
Electron-positron mass spectrum from known hadronic sources without hot, dense medium contribution.

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Electron positron emission mass spectrum in 200 GeV Au+Au

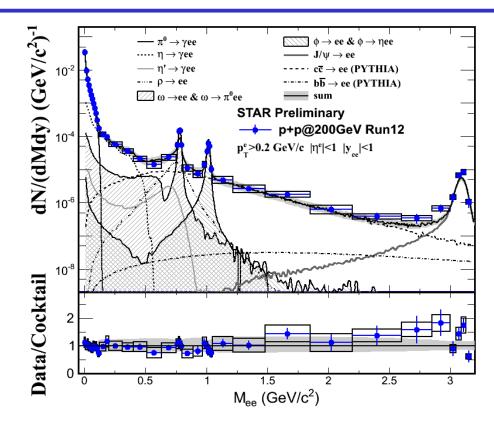
Phys. Rev. Lett. 113 (2014) 22301



Significant excess is observed for $0.3 < M_{ee} < 0.8 \text{ GeV/c}^2$, representing the hot, dense medium contribution.



Electron positron emission mass spectrum in 200 GeV proton+proton collisions



STAR: QM2014

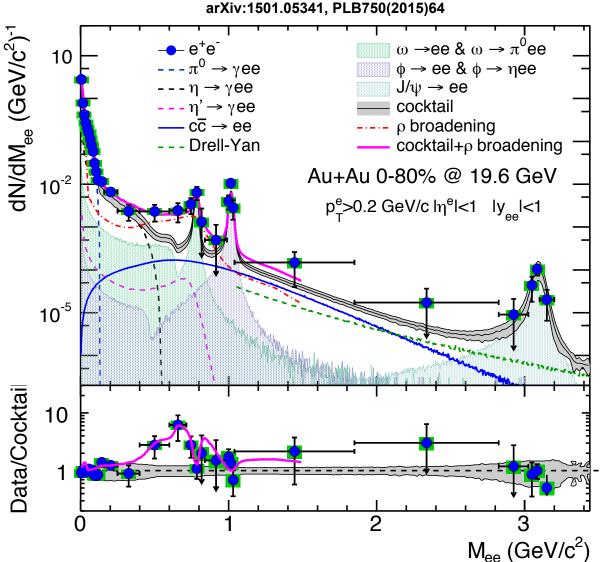
The cocktail simulation with expected hadronic contributions, is consistent with data in proton+proton collisions.

No hot, dense medium, no excess!

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Electron positron emission mass spectrum in 19.6 GeV Au+Au

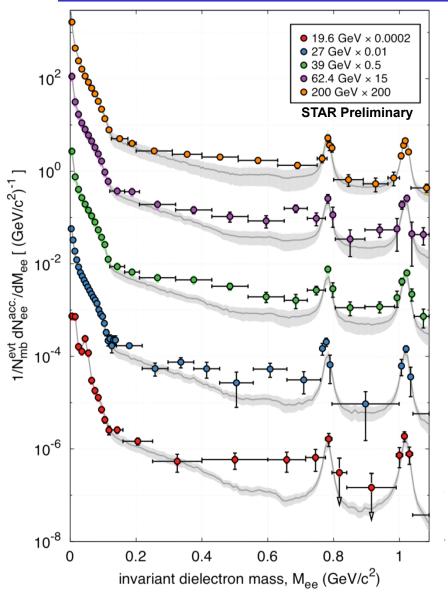


Significant excess is observed in $0.3 < M_{ee} < 0.8 \text{ GeV/c}^2$, representing the hot, dense medium contribution.

55th International Winter Meeting on Nuclear Physics

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Electron-positron emission at lower energies

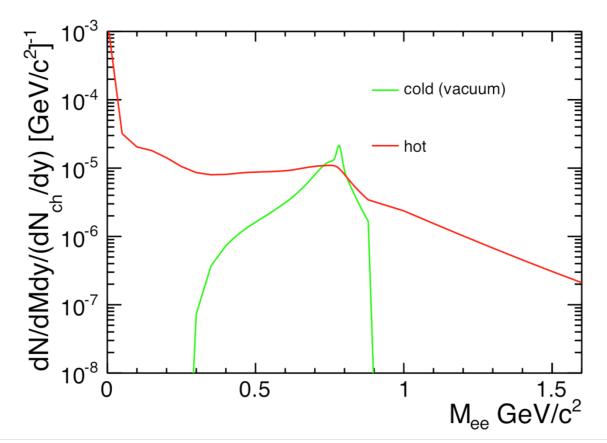


Low-mass excess is observed for 19.6, 27, 39, 62.4, and 200 GeV Au+Au collisions!

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The mass distribution from hot, dense medium in 200 GeV Au+Au



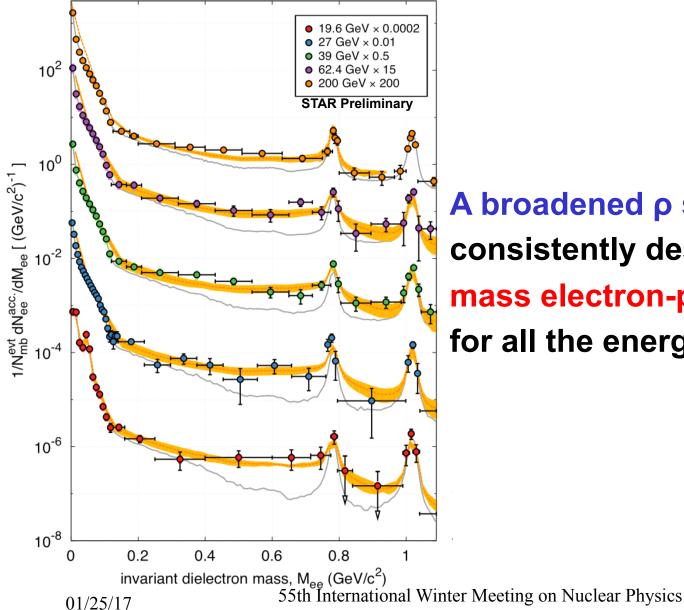
Red: a broadened p spectrum function Green: vacuum-like spectrum function

Model: Rapp & Wambach, priv. communication Adv. Nucl.Phys. 25, 1 (2000); Phys. Rept. 363, 85 (2002)

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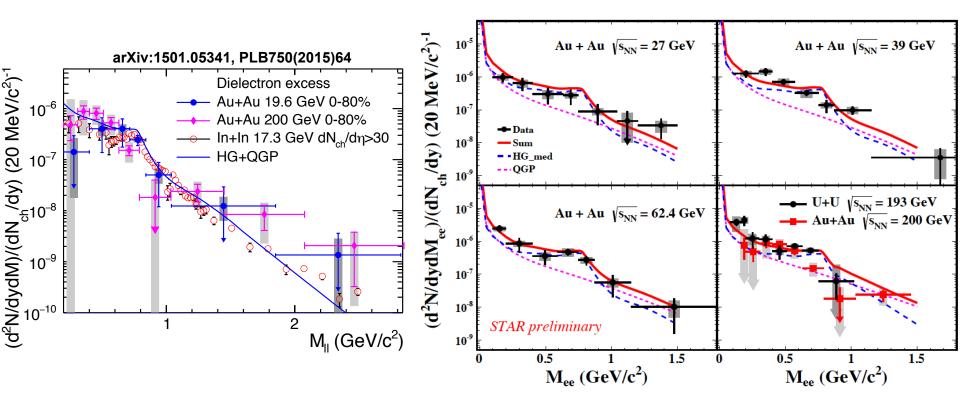


Electron-positron emission at lower energies



A broadened ρ spectrum function consistently describes the low mass electron-positron excess for all the energies 19.6-200 GeV.

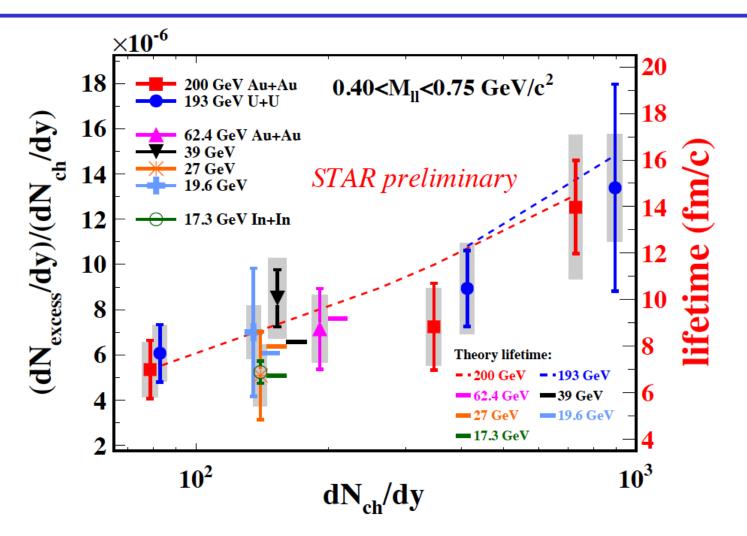
BROOKHAVEN The p resonance spectrum function: broadened



A broadened ρ spectrum function consistently describes the low mass electron-positron excess for all the energies 19.6-200 GeV.

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THOOKHEVEN The low mass measurements: lifetime indicator

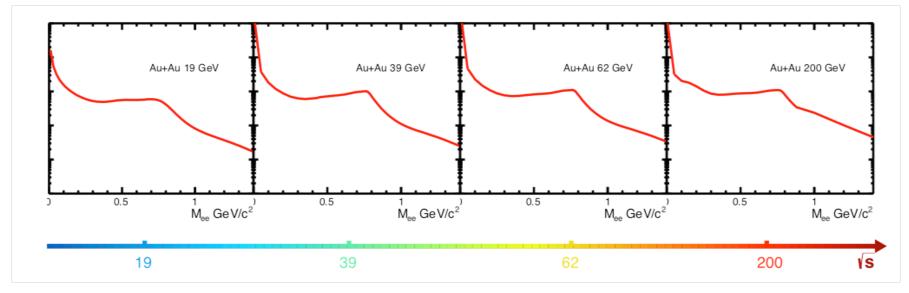


Low-mass electron-positron production, normalized by dN_{ch}/dy , is proportional to the life time of the medium from 17.3 to 200 GeV.

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The contribution from hot, dense medium



The electron-positron spectrum from hot, dense medium is consistent with a broadened ρ resonance in medium.

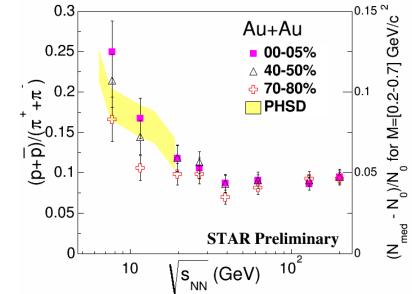
The production yield normalized by dN_{ch}/dy is proportional to lifetime of the medium from 17.3 to 200 GeV. Why?



The contribution from hot, dense medium from 17.3 to 200 GeV

Low-mass electron-positron emission depends on T, total baryon

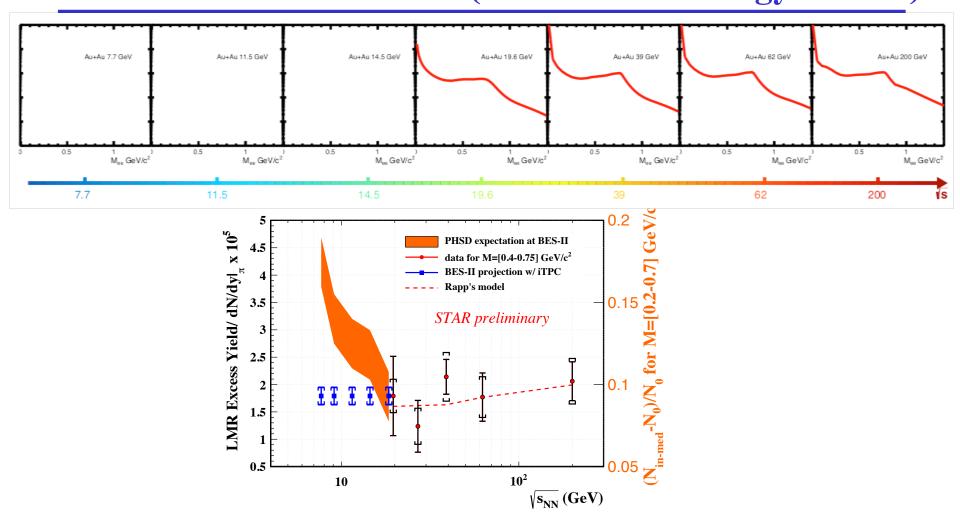
- density, and lifetime
- Coupling to the baryons plays an essential role to the modification of ρ spectral function in the hot, dense medium.



Normalized low-mass electron-positron production, is proportional to the life time of the medium from 17.3 to 200 GeV, given that the total baryon density is nearly a constant and that the emission rate is dominant in the Tc region.

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Probe total baryon density effect 7.7 GeV to 19.6 GeV (RHIC beam energy scan II)



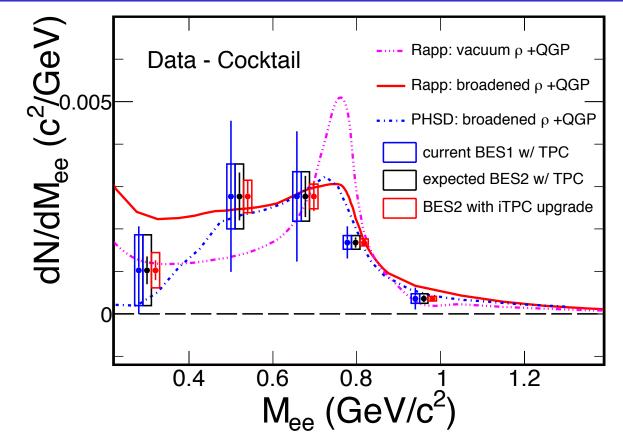
Broader and more electron-positron excess down to 7.7 GeV collision energy? Beam Energy Scan II provides a unique opportunity to quantify the total baryon density effect on the ρ broadening!

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ATIONAL LABORATORY

THOMAL LABORATORY Distinguish the mechanisms of rho broadening

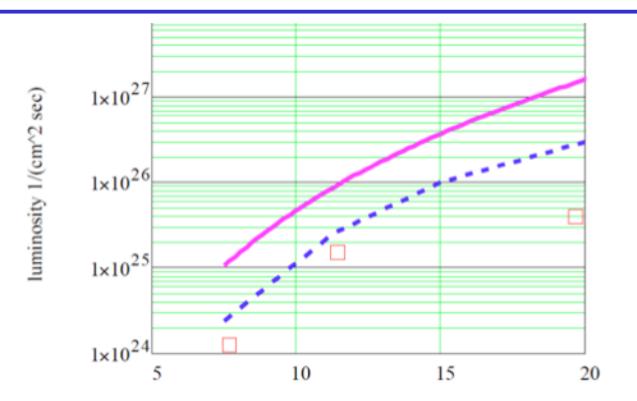


Knowing the mechanism that causes in-medium rho broadening and its temperature and baryon-density dependence is fundamental to our understanding and assessment of chiral symmetry restoration in hot QCD matter !

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Beam Energy Scan II in 2019-2020



RHIC is unique to study chiral symmetry restoration:

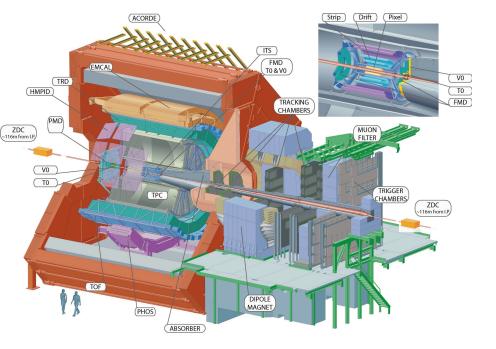
Beam energy scan II: collision energies 7.7, 9.1, 11.5, 14.5, 19.6 GeV.

Electron cooling from CAD will increase collision rate from 3-10.

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World-wide interest





• World interest: SPS, PHENIX, LHC, HADES, FAIR, NICA, KEK

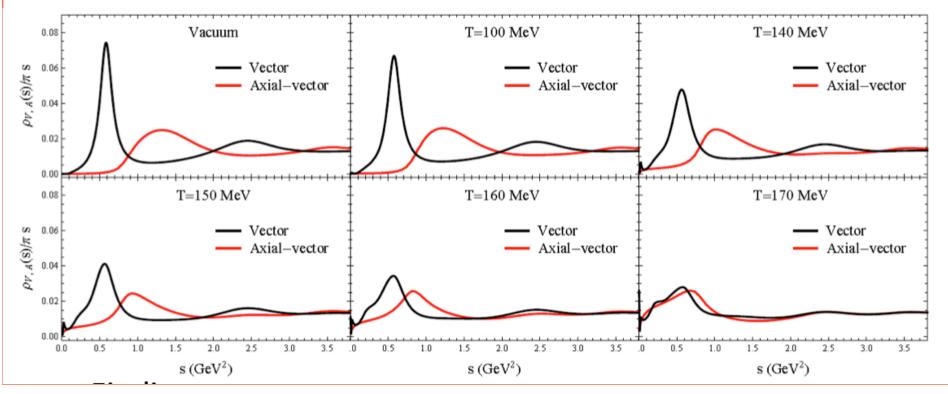
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The future electron-positron program

To link electron-positron measurements to chiral symmetry restoration need more precise measurement at $\mu_B = 0$:

- Lattice QCD calculation is reliable at $\mu_B = 0$.
- Theoretical approach: derive the a1(1260) spectral function by using the broadened rho spectral function, QCD and Weinberg sum rules, and inputs from Lattice QCD; to see the degeneracy of the rho and a1 spectral functions (Hohler and Rapp 2014).

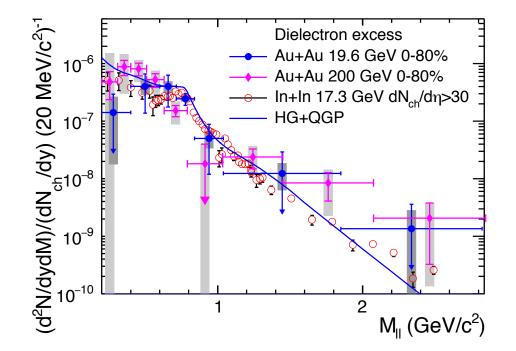


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The future electron-positron program

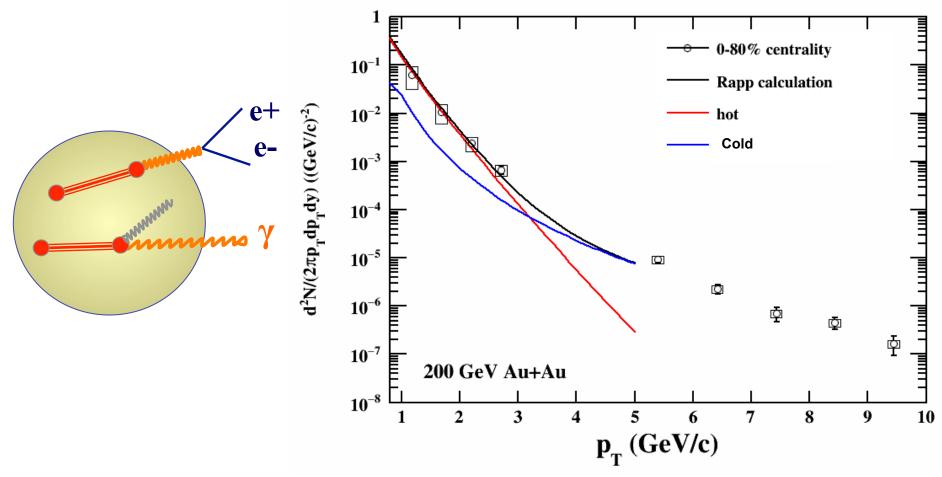
 The slope in the intermediate mass region represents the true average temperature T of the medium.



 Low-mass electron-positron emission depends on T, total baryon density, and life time, and enables systematic life-time measurements.



Photon emission

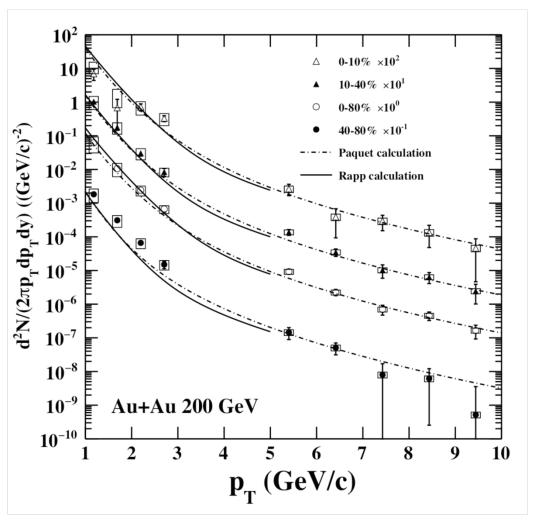


Hot contribution observed in the photon energy spectrum!

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Photon emission



Quark-Gluon Plasma emission spectrum: photon energy a few 10⁹ electron volts

Sun emission spectrum: Photon energy a few electron volts.

Hottest matter in the universe: a few trillion degree Celsius!

STAR Collaboration, arXiv: 1607.01447, submitted to PLB.



Electron-positron tomography is enabled by the Time of Flight Detector at STAR.

A broadened ρ spectrum function consistently describes the low mass electron-positron excess in Au+Au collisions for all the energies 19.6, 27, 39, 62.4 and 200 GeV.

Beam Energy Scan II (7.7-19.6 GeV) will provide a unique opportunity to quantify the effect of Chiral Symmetry Restoration via total baryon density effect on the ρ broadening.

Enable unique measurements of the temperature and lifetime of hot, dense medium

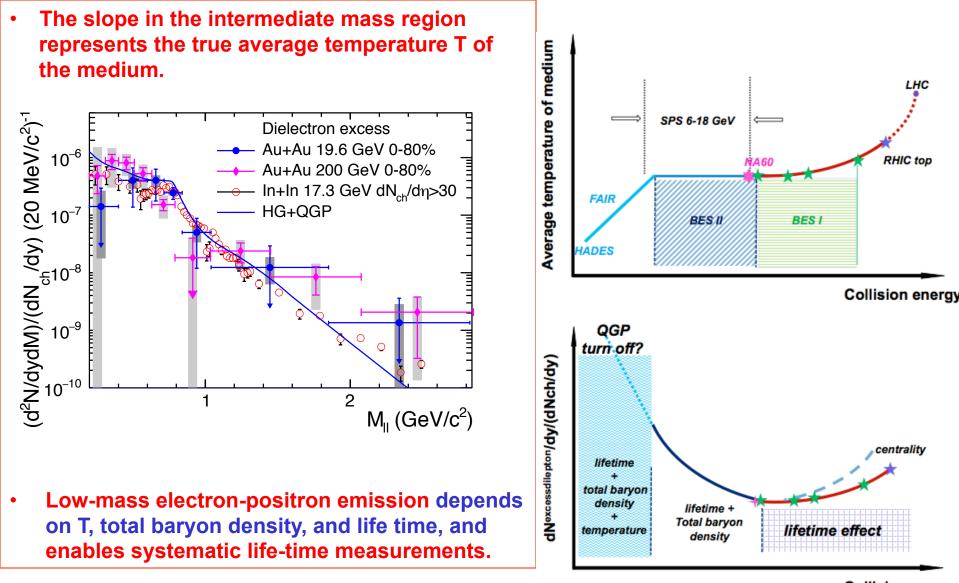
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The future electron-positron program



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