# Radiative capture and photodisintegration reactions for the synthesis of the *p* nuclei

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#### Outline

#### Introduction

- Nucleosynthesis of heavy elements and the *p* nuclei
- the γ-process reaction network

#### **Experimental measurements of cross-sections**

- photo-induced reaction measurements
- charged-particle induced reaction studies

#### **Experimental results**

- testing the E1-strength in  ${}^{90}$ Zr via  ${}^{89}$ Y(p, $\gamma$ )
- total and partial cross sections for <sup>92</sup>Mo(p,γ)

#### The synthesis of the p nuclei

#### *p* nuclei

- 30-35 neutron deficient isotopes
- cannot be produced by neutroncapture reactions
- relatively low isotopic abundances in comparison to *s*- and *r*-isotopes
- originally thought to be produced via proton-capture
- temperatures would lead to immediate photodisintegration



M. Arnould et al., Phys. Rep. 450 (2007) 97



T. Rauscher et al., Rep. Prog. Phys. 76 (2013) 066201

#### The synthesis of the *p* nuclei



### $\gamma$ process reaction-network

- huge photodisintegration reaction-network
- at temperatures between 1.5 GK and 3 GK in ccSN or type Ia SN
- starting from stable seed nuclei formed in the *s* or *r*-process
- $\gamma$ -process path proceeds first via ( $\gamma$ ,n) reactions
- branching for A < 130 mainly via ( $\gamma$ ,p)
- above A > 130 ( $\gamma$ , $\alpha$ ) get more important

#### The synthesis of the p nuclei

#### reaction-network calculations



- γ-process network calculations cannot reproduce solar system abundance
- other contributions from rp-,  $\alpha$  or other processes?
- problems with photoinduced reaction cross sections?

S.E. Woosley and W. M. Howard, ApJSS 36 (1978) 285

#### Photodisintegration



- measuring cross sections via direct detection of ejectiles or via photoactivation
- using either monochromatic γ-ray beams or Bremsstrahlung

#### **Ground-state contributions**

- measured cross sections cannot directly used for astrophysics
- for γ-induced reaction the ground-state contribution is almost zero
- larger contribution from excited states in the stellar plasma (T<sub>9</sub> > 1.5)
- reaction rates are obtained from the inverse reactions via reciprocity theorem



T. Rauscher, ApJSS **201** (2012) 26

#### **Experimental measurements of cross sections**

#### **Statistical model**

- cross sections in the Gamow window are small ( < µb)</li>
- most of the reactions are not accessible in the laboratory
- reaction rates are calculated mostly in the scope of the statistical model
- cross-section measurements to improve nuclear physics inputparameters:
  - $\gamma$ -strength function (also via ( $\gamma$ , $\gamma'$ ))
  - particle + nucleus optical model potentials



$$\sigma_{jk}^{\mu}(E) = \pi \lambda_j^2 \frac{1}{(2J_I^{\mu} + 1)(2J_j + 1)} \sum_{j^{\mu}} (2J + 1) \frac{T_j^{\mu}(J^{\pi})T_k(J^{\pi})}{T_{tot}(J^{\pi})}$$

#### **Charged-particle induced reaction cross sections**

## **Activation technique**

- widely used technique for measureing cross-sections
- temporal and spatial separation of irradiation and spectroscopy
- no access to reactions involving stable reaction products
- feasible half-lives neccessary



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## $4\pi$ summing crystal method

- complete deexcitation is summed up in one peak
- access to stable reaction products
- need for very different Q-values for competing reactions
- no access to partial cross-sections





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## **In-beam method with HPGe detectors**

- de-excitation of the entry state
  - determination of partial cross sections
  - very sensitive on the γ-ray strength function



- de-excitation of the entry state
  - determination of partial cross sections
  - very sensitive on the γ-ray strength function
- transitions to the ground state
  - determination of the total cross section



 10 MV FN-Tandem ion accelerator

#### HORUS γ-ray spectrometer

- 14 HPGe detectors
  - High resolution
    ≈ 2 keV @ 1332 keV
  - High total efficiency
    ≈ 2% @ 1332 keV
- 5 different angles with respect to beam axis
  - determination of angular distributions
- BGO shields





L. Netterdon et al., NIM A 754 (2014) 94-100

#### **Target chamber**

- cooling trap
- tantalum coating
- independent current readouts
- δ-electron suppression
- built-in detector for Rutherford Backscattering Spectrometry (RBS)





L. Netterdon et al., NIMA 754 (2014) 94-100



- reaction in a region which is normally underproduced in reaction network calculations
- total cross section was measured twice before
- γ-ray strength function in <sup>90</sup>Zr was measured before
- natural yttrium target (583µg/cm<sup>2</sup>)
- beam currents between 1nA and 60nA
- five different proton energies between 3.65 MeV and 4.70 MeV, i.e. γ-ray energies between 7.71 MeV and 12.98 MeV (Q-Value: 8353.4 keV)







L. Netterdon et al., PLB 744 (2015) 358

<sup>89</sup>Y(p,γ) partial cross sections: huge deviations



• using  $\gamma$ -strength function from ( $\gamma$ , $\gamma$ ') measurement:



R. Schwengner *et al.*, PRC **78** (2008) 064314

adjusting γ-strength to measured data



#### Impact on <sup>90</sup>Zr(γ,p)<sup>89</sup>Y

- 15 % larger reaction rate than
  based on E1-strength from (γ,γ')
- twice as large as BRUSLIB reaction rate (QRPA strength)
- three times smaller than
  NONSMOKER rates (Lorentziantype E1-strength)



R. Schwengner *et al.*, PRC **78** (2008) 064314L. Netterdon *et al.*, PLB **744** (2015) 358P. Scholz, AG Zilges, University of CologneRadiative capture and photodisintegration reactions



<sup>86</sup>Sr

- <sup>92</sup>Mo is the most abundant *p* nuclei and its origin is highly debated
- total cross-section were measured before with different techniques at energies below 3.5 MeV
- isotopically enriched <sup>92</sup>Mo target (94 %)
- beam currents between 50 nA and 350 nA
- proton energies between 3.7 MeV and 5.4 MeV
  - extending measurement towards higher energies
  - sensitive to higher energy γ-ray transitions

#### In-beam measurement of the <sup>92</sup>Mo(p,γ)

#### Problem: metastable state @ 391 keV

- Significant half-life
- Electron capture branching to <sup>93</sup>Mo

#### Solution:

- Determine σ<sub>gs</sub>
- Determine  $\sigma_{\rm m}$



#### **Total cross sections**

- Previously measured cross sections show fluctuating behavior
- TALYS calculations
  - Unsatisfactory reproduction with default settings
  - modified γ-strengths
    - M1 shell model
    - Gogny-HFB + QRPA
    - Skyrme-HFB + QRPA



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#### **Partial cross sections**



J. Mayer et al., PRC, accepted

#### **Summary**

- the origin of the *p* nuclei is still unclear due to astrophysical and nuclear physics uncertainties
- direct measurements of (γ,x) reaction rates can in most cases not directly applied to reaction network calculations
- reaction rates are usually calculated within the statistical model and via reciprocity theorem from the inverse reactions
- charged-particle induced reaction studies can be used for the improvement of models for statistical properties of nuclei
  - γ-ray strength functions
  - > particle+nucleus optical model potentials





#### **Uncertainties for γ process nucleosynthesis**

#### **Astrophysical uncertainties**

- seed abundances from *s* or *r*-process or chemical evolution
- temperature and density profiles
- contribution from, for instance, α-process in neutrino-driven wind scenarios or *rp* process in Type Ia X-ray bursts?





S.E. Woosley and W. M. Howard, ApJSS **36** (1978) 285 T. Rauscher *et al.*, Rep. Prog. Phys. **76** (2013) 066201

#### **Uncertainties for γ process nucleosynthesis**

#### **Nuclear physics uncertainties**

- sensitivity study of the γ process in ccSN
- all *p* isotope abundances are sensitive to (γ,n) reaction rates
- only the lighter *p* nuclei are sensitive to (γ,p)
- (γ,a) especially important for the production factors of the heavier *p* isotopes



Fig. 10.—Ratio of *p*-abundances calculated with modified rates and the currently accepted HF rates for all (*a*) *n*-induced, (*b*) *p*-induced, and (*c*)  $\alpha$ -induced reactions and their inverse processes. Squares and crosses denote results obtained with rates 3 times smaller and larger, respectively.

W. Rapp et al., ApJ 653 (2006) 474

#### **Experimental measurements of cross sections**

#### **Monochromatic** γ-ray beams

Example: HIγS @ Duke, TERAS @ Tsukuba

- using Laser Compton-Backscattering to produce (quasi-) monochromatic γ-rays
- measuring (γ,n) cross sections point-by-point until the neutron treshold



T. Kondo et al., Phys. Rev. C 86 (2012) 014316

#### **Experimental measurements of cross sections**

#### Bremstrahlung

- simulating a planck distribution with many Bremsstrahlung spectra using different endpoint energies
- direct measurement of (γ,n) reaction rate at a specific temperature



P. Mohr et al., PLB 488 (2000) 127

## 4*π*-summing crystal



SuN – NaI  $4\pi$  summing detector [NSCL]



FIG. 1. (Color online) Experimental spectra from the SuN detector for measurements at  $E_{\alpha} = 7.7$  MeV. The spectra correspond to the <sup>58</sup>Ni target (solid black), thick tantalum backing (dotted blue), and normalized room background (dotted-dashed red). The inset shows a zoom around the sum-peak region of the <sup>58</sup>Ni( $\alpha, \gamma$ )<sup>62</sup>Zn reaction.

S. Quinn et al., Phys. Rev. C 89 (2014) 054611

### In-beam method with HPGe detectors



Nuclear astrophysics @ HORUS



L. Netterdon et al., NIM A 754 (2014) 94

#### **Target chamber**

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L. Netterdon et al., NIMA 754 (2014) 94-100

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