### Relativistic hydrodynamics as an asymptotic series

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103.3452, 1302.0697, 1409.5087,
1503.07514, 1603.05344, 1609.04803
& 1610.02023 (partial review / viewpoint)

# Hydrodynamization

M. P. Heller, R. A. Janik and P. Witaszczyk, Phys. Rev. Lett. 108, 201602 (2012), **1103.3452** 

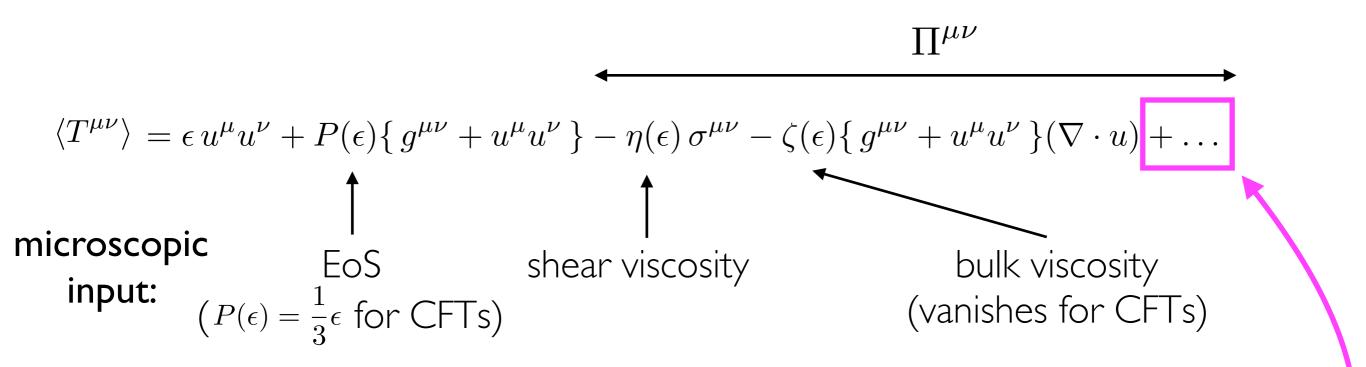
M. P. Heller, A. Kurkela and M. Spalinski 1609.04803

### Relativistic hydrodynamics

hydrodynamics is

an EFT of the slow (?) evolution of conserved currents in collective media close to equilibrium (?)

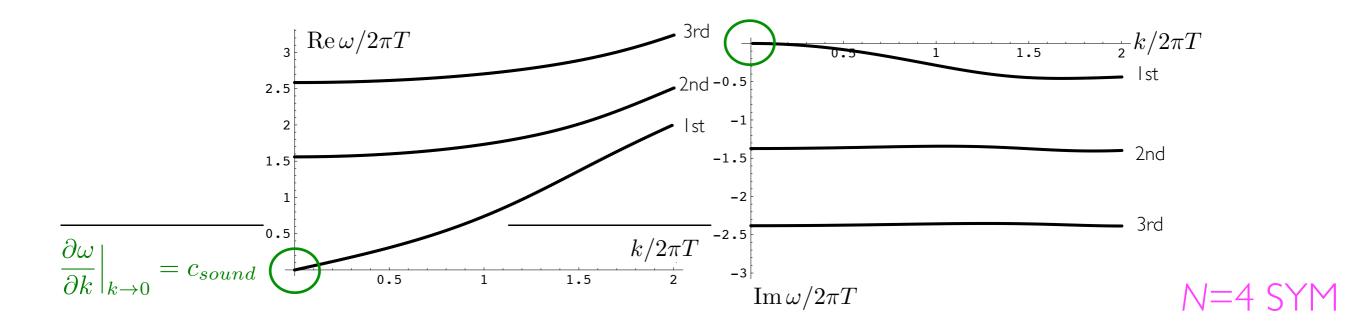
**DOFs**: always local energy density  $\epsilon$  and local flow velocity  $u^{\mu}$   $(u_{\nu}u^{\nu} = -1)$ **EOMs**: conservation eqns  $\nabla_{\mu} \langle T^{\mu\nu} \rangle = 0$  for  $\langle T^{\mu\nu} \rangle \underline{\text{expanded in gradients}}$ 



This talk: behaviour of the gradient expansion at large orders in the number of  $abla ^{\prime}$ 

### Excitations in strongly-coupled plasmas

see, e.g. Kovtun & Starinets [hep-th/0506184]



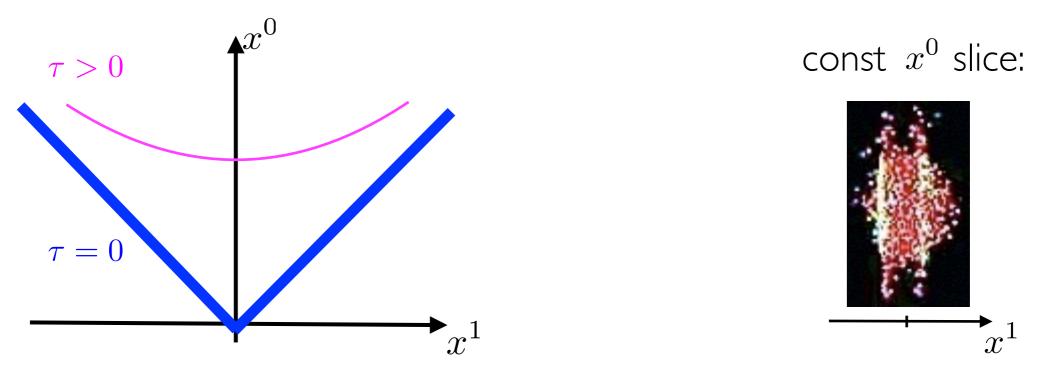
 $\omega(k) \rightarrow 0$  as  $k \rightarrow 0$ : slowly dissipating modes (hydrodynamic sound waves)

all the rest: far from equilibrium (QNM) modes damped over  $t_{therm} = O(1)/T$ 

Linear response theory:

$$\delta \langle T^{\mu\nu} \rangle = \int d^3k \sum_{modes} e^{-i\,\omega_{mode}(k)\,t + i\,\vec{k}\cdot\vec{x}} \left[ \left( \operatorname{Res}_{\omega=\omega_{mode}(k)} G_R(\omega,\vec{k}) \right) \cdot \delta g(\omega_{mode}(k),\vec{k}) \right] \,^{\mu\nu}$$

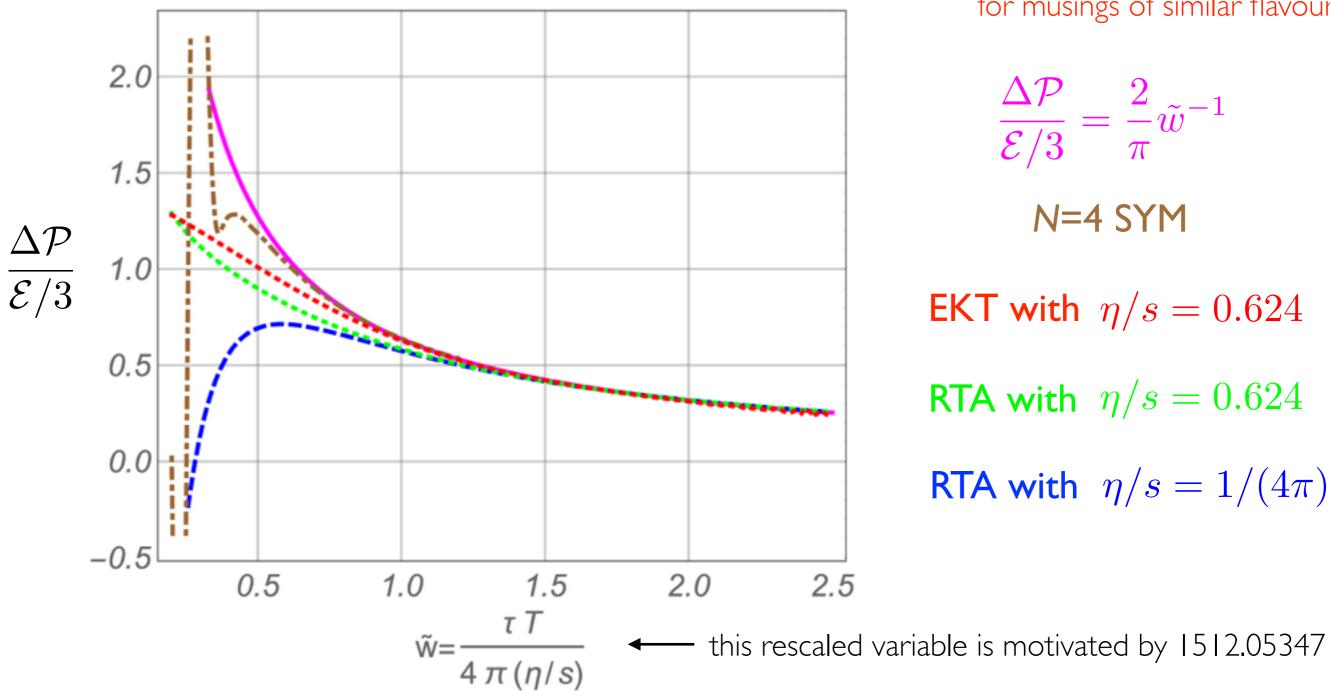
## Boost-invariant flow [Bjorken 1982]



Boost-invariance: in  $(\tau \equiv \sqrt{x_0^2 - x_1^2}, y \equiv \operatorname{arctanh} \frac{x_1}{x_0}, x_2, x_3)$  coords no y-dep

In a CFT: 
$$\langle T^{\mu}_{\nu} \rangle = \text{diag} \left\{ -\mathcal{E}(\tau), -\mathcal{E} - \tau \dot{\mathcal{E}}, \, \mathcal{E} + \frac{1}{2}\tau \dot{\mathcal{E}}, \, \mathcal{E} + \frac{1}{2}\tau \dot{\mathcal{E}} \right\}$$
  
and via scale-invariance  $\frac{\Delta \mathcal{P}}{\mathcal{E}/3}$  is a function of  $w \equiv \tau T$   
Gradient expansion: series in  $\frac{1}{w}$ .

### Hydrodynamization (across conformal theories) 1103.3452 & 1609.04803



see Andrei Starinets' talk for musings of similar flavour

$$\frac{\Delta \mathcal{P}}{\mathcal{E}/3} = \frac{2}{\pi} \tilde{w}^{-1}$$

$$N=4 \text{ SYM}$$

EKT with  $\eta/s = 0.624$ 

RTA with  $\eta/s = 0.624$ 

RTA with  $\eta/s = 1/(4\pi)$ 

Viscous hydrodynamics works despite huge anisotropy captured by  $-\eta\,\sigma^{\mu
u}$ 

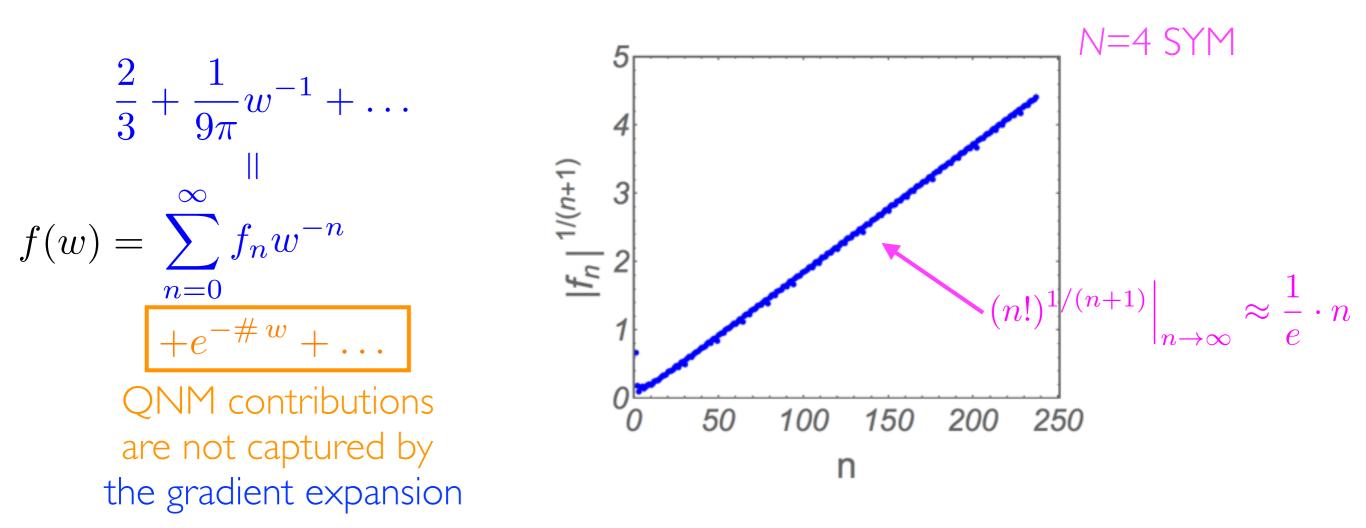
see Paul Romatschke's talk for possible implication for HIC pheno

# Why can hydrodynamization occur?

M. P. Heller, R. A. Janik and P. Witaszczyk, Phys. Rev. Lett. 110, 211602 (2013), **1302.0697** 

### Hydrodynamic gradient expansion is divergent

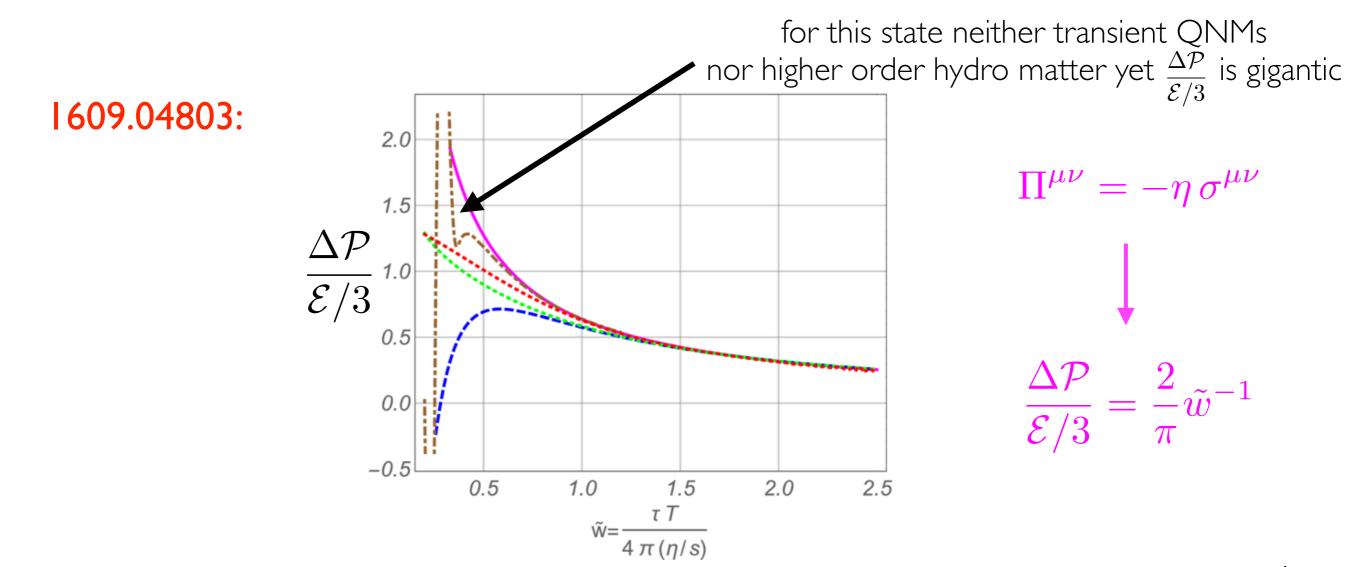
In I302.0697 we computed  $f(w) \equiv \frac{2}{3} + \frac{1}{6} \frac{\Delta \mathcal{P}}{\mathcal{E}}$  up to  $O(w^{-240})$ :



The gradient expansion cannot converge as I 302.0697 explicitly demonstrated

### Outlook: why can hydrodynamization occur?

Divergent series are better than convergent series since their applicability is not limited by the series itself but whether one can neglect effects not captured by it



This does not mean that second (or third) order hydro is not important  $\longrightarrow$  other flows? Note that for CFTs the bulk viscosity term can also be very large at hydro threshold Attems et al. 1604.06439 Hydrodynamization with anomalous transport?

# What is the hydro gradient expansion?

M. P. Heller, R. A. Janik and P. Witaszczyk, Phys. Rev. Lett. 110, 211602 (2013), **1302.0697** 

M. P. Heller, M. Spaliński, Phys. Rev. Lett. 115, 072501 (2015), **1503.07514** 

> A. Buchel, M. P. Heller and J. Noronha PRD RC in press, **1603.05344**

> > M. P. Heller, 1610.02023

### Borel transform, gradient expansion and QNMs

Analytic continuation of  $f_B(\xi) \approx \sum_{n=0}^{240} \frac{1}{n!} f_n \xi^n$  revealed the following singularities: 302.0697 40 20 m(ξ) N=4 SYM: -40-20 -10 0 10 20  $\operatorname{Re}(\xi)$ Branch cuts start at  $\frac{3}{2}i\hat{\omega}_{QNM_1} = \frac{3}{2}iT \times \omega_{QNM_1}(k=0)$ . Inverse trafo ambiguous:  $\left(\int_{\mathcal{C}_i} d\xi - \int_{\mathcal{C}_i} d\xi\right) \left[w \, e^{-\xi \, w} f_B(\xi)\right] \sim w^{\alpha_{QNM_1}} e^{-\frac{3}{2}i \,\hat{\omega}_{QNM_1} \, w} \times (\dots) + \dots$  $e^{-i\hat{\omega}_{QNM_1}Tt} \longrightarrow e^{-i\hat{\omega}_{QNM_1}\int T\,d\tau}$ 

## Lesson from MIS 1503.07514

$$(\tau_{\Pi} u^{\alpha} \nabla_{\alpha} + 1) \Pi_{\mu\nu} = -\eta \sigma_{\mu\nu} + \dots$$

$$\int C_{\eta} = \frac{\eta}{s} \text{ and } C_{\tau_{\Pi}} \equiv \hat{\omega}_{MIS} = \tau_{\Pi} T$$

$$f' = -\frac{1}{C_{\tau_{\Pi}}} + \frac{2}{3C_{\tau_{\Pi}}f} + \frac{16}{3w} - \frac{16}{9wf} + \frac{4C_{\eta}}{9C_{\tau_{\Pi}}wf} - \frac{4f}{w} + \dots$$

$$\int \text{gradient expansion and I transient QNM:}$$

$$f = \sum_{n=0}^{\infty} \frac{f_n}{w^n} + \delta f + \mathcal{O}(\delta f^2)$$

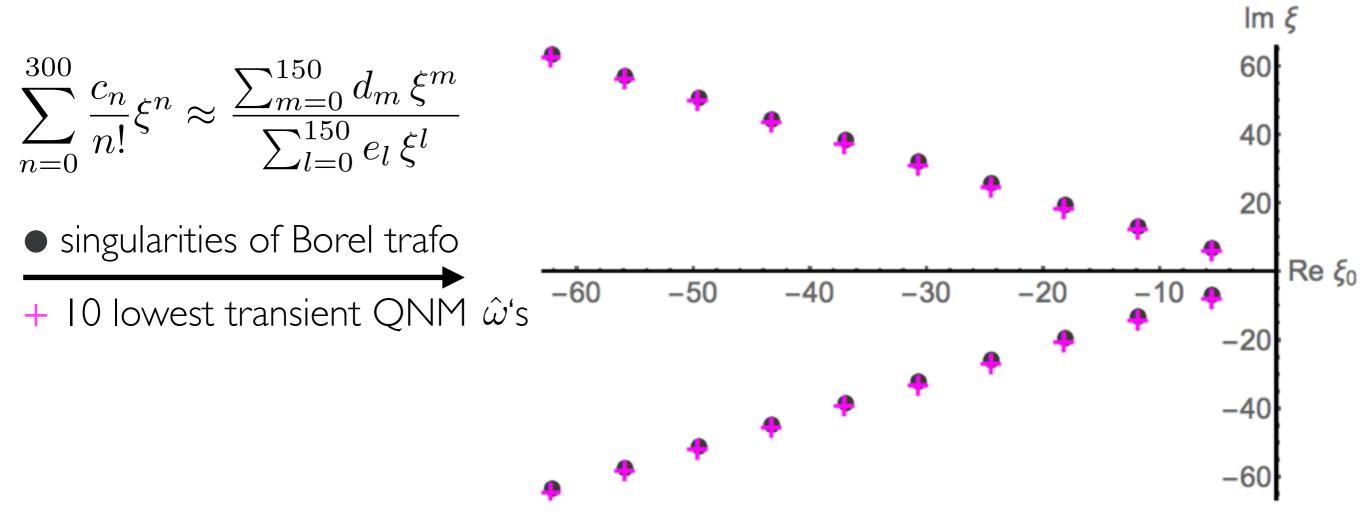
$$\exp(-\frac{3}{2C_{\tau_{\Pi}}}w) \times \dots$$

$$\int \text{transseries:}$$

$$f = \sum_{m=0}^{\infty} (c_{amb} + r)^m \left\{ w^{\frac{C_{\eta}}{C\tau_{\Pi}}} \exp\left(-\frac{3}{2C_{\tau_{\Pi}}}w\right) \right\}^m \sum_{n=0}^{\infty} a_{m,n} w^{-n}$$
resurgence

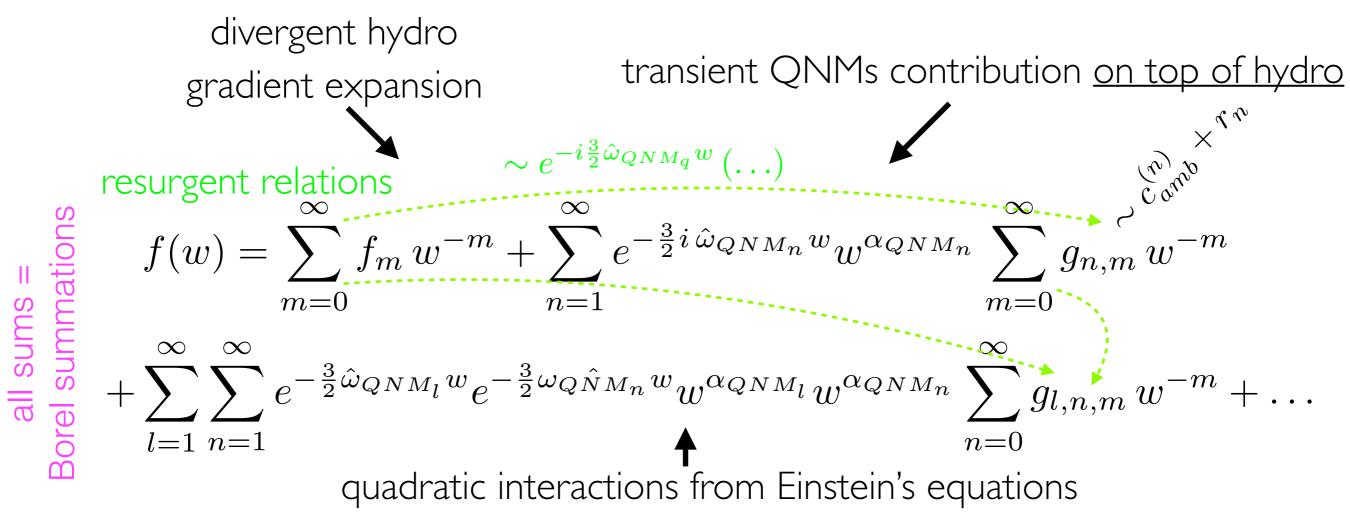
# Lesson from cosmology 1603.05344

$$\frac{d \operatorname{Entropy}}{dt} = V \times \left(\sum_{n=0}^{\infty} c_n \xi^n\right)^2 + \dots \text{ with } \xi = \frac{H}{T} \text{ for a h} \mathcal{F} \mathsf{f} \text{ in } -dt^2 + e^{2Ht} d\vec{x}^2$$



Hydrodynamic gradient expansion knowns about all transient QNMs

## Transseries for N=4 SYM<sub>1610.02023</sub>



	non-equilibrium physics	<b>QM</b> with $V = -\frac{1}{2}x^{2}(1 - \sqrt{g}x)^{2}$
Appealing analogy:	gradient expansion in $\frac{1}{w}$	perturbative series in $g$
	transient QNMs $e^{-i\frac{3}{2}\hat{\omega}_{QNM_n}w}(\ldots)$	instanton $e^{-1/(3g)}(\ldots)$

Open problem: covariantize to  $\Pi^{\mu\nu} = \eta \, \sigma^{\mu\nu} + \ldots + e^{i \, \hat{\omega}_{QNM_1} \int dx^{\alpha} T \, u_{\alpha}} \times (\ldots) + \ldots$ see also Mukund Rangamani's talk 10/13 10/13

# "Weak" coupling

M. P. Heller, A. Kurkela and M. Spalinski 1609.04803

see also Gabriel Denicol's talk

### **RTA** kinetic theory

Natural language to talk about weakly coupled media is the Boltzmann equation:

$$p^{\mu}\partial_{\mu}g(x,p) = C[g(x,p)]$$
 with  $\langle T^{\mu\nu}\rangle(x) = \int_{\text{momenta}} g(x,p) p^{\mu}p^{\nu}$ 

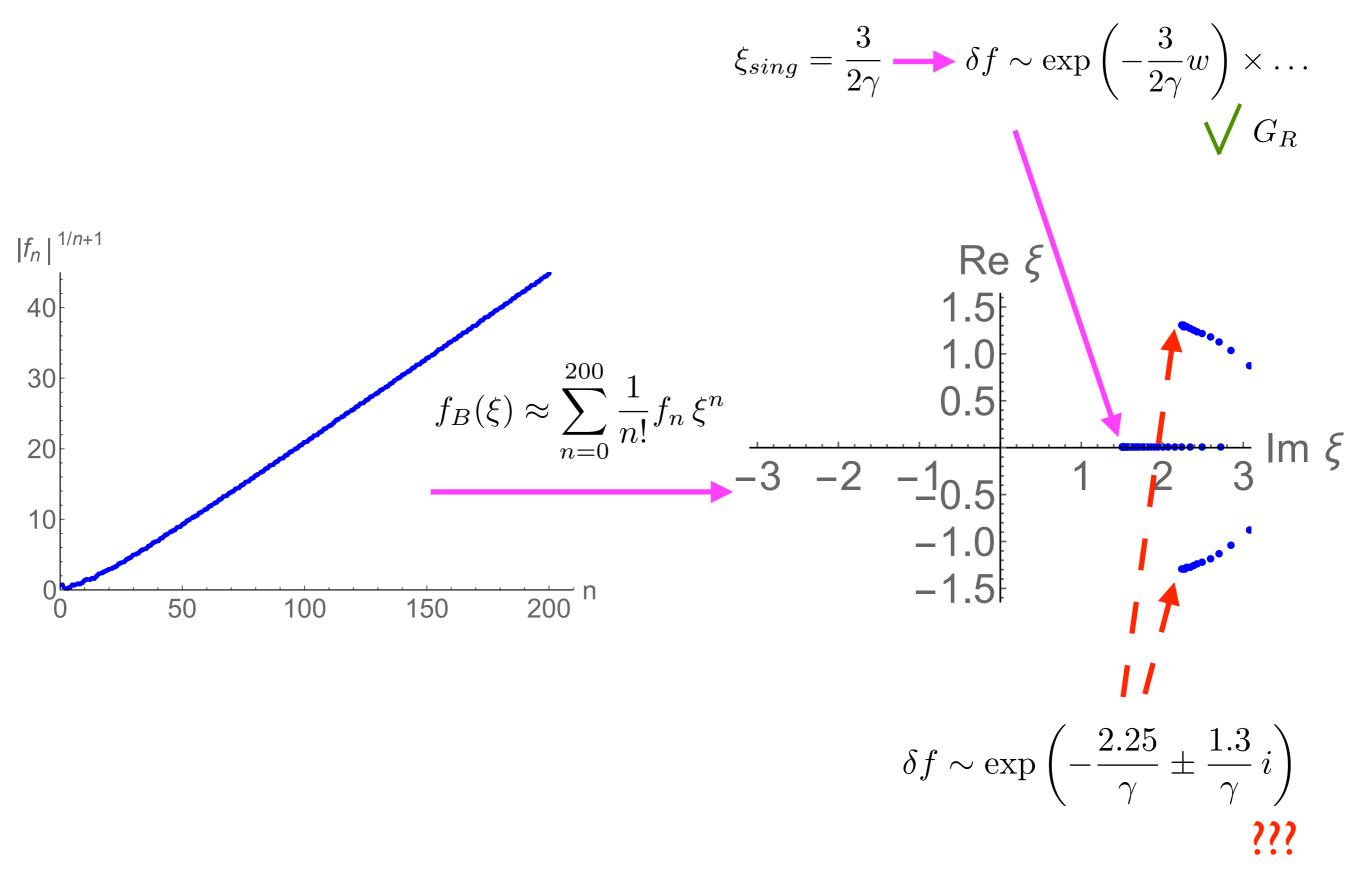
LO C[g(x, p)] for gauge theories is complicated. We will use instead

$$C[g(x,p)] = -\frac{p^{\mu}u_{\mu}}{\hat{\mathcal{T}}} \Big\{ g(x,p) - g_0(x,p) \Big\} \text{ with } g_0(x,p) = e^{\frac{u_{\mu}p^{\mu}}{T}}$$

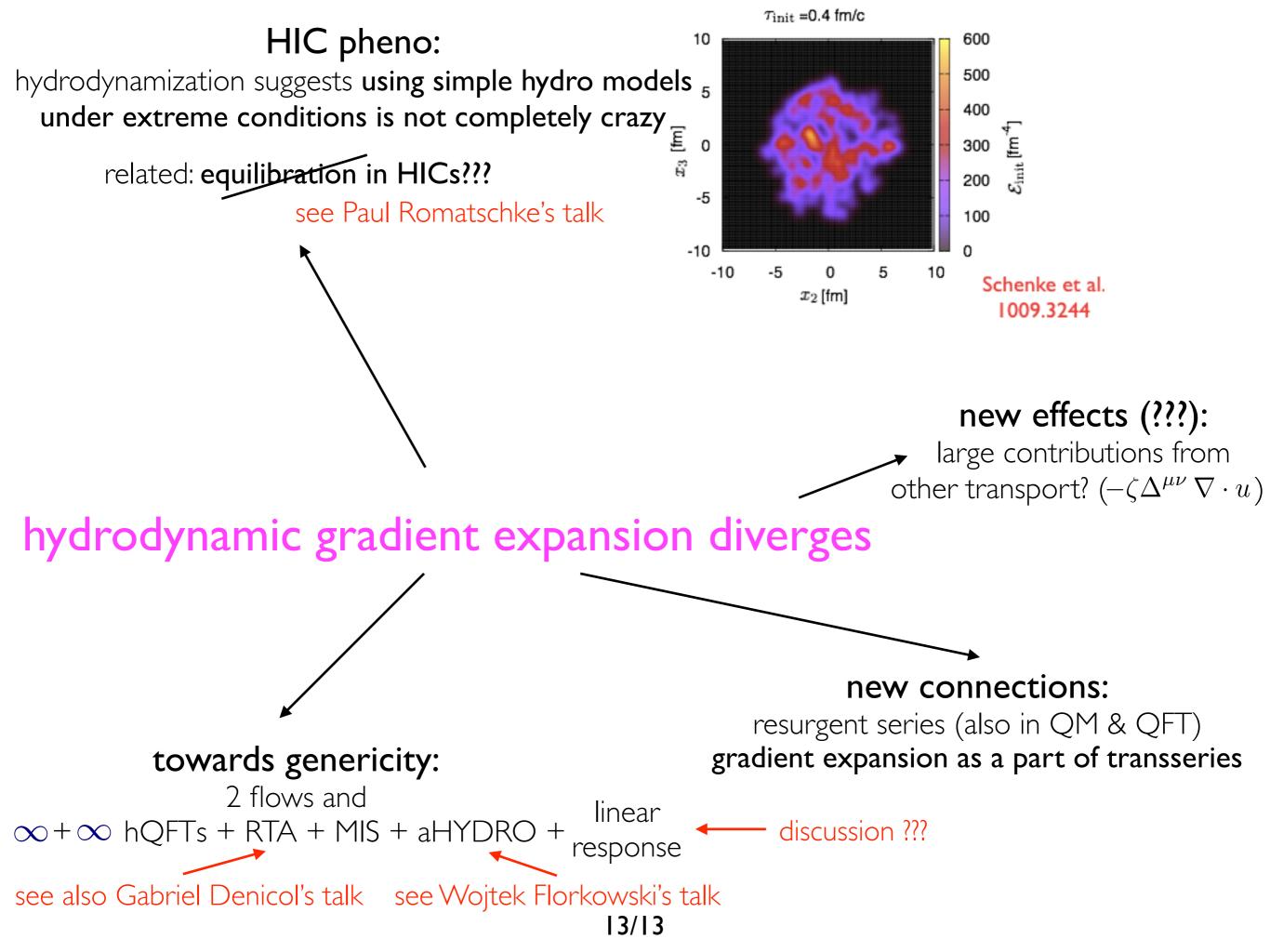
This equation is, typically, highly nonlinear due to  $\langle T^{\mu\nu} \rangle u_{\nu} = -\mathcal{E}(T) u^{\mu\nu}$ 

CFTs: 
$$p^{\mu}p_{\mu} = 0$$
 and  $\hat{\mathcal{T}} = \frac{\gamma}{T}$ . Note that  $\gamma$  can be scaled-away (we set it to I).

QNM in kinetic theory



## Executive summary



Support

### Evolution equations for relativistic viscous fluids

$$\nabla_{\mu} \{ \epsilon u^{\mu} u^{\nu} + P(\epsilon) \{ g^{\mu\nu} + u^{\mu} u^{\nu} \} - \eta(\epsilon) \sigma^{\mu\nu} \} = 0 \text{ is acausal.}$$

### Remedy: make $\Pi^{\mu\nu} = \langle T^{\mu\nu} \rangle - (\epsilon u^{\mu}u^{\nu} + P(\epsilon)\{g^{\mu\nu} + u^{\mu}u^{\nu}\})$ a new DOF, e.g. $(\tau_{\Pi}\mathcal{D} + 1) \Pi^{\mu\nu} = -\eta\sigma^{\mu\nu}$

Small perturbations obey Maxwell-Cattaneo equation

$$\partial_t^2 \delta u_z - \frac{\eta}{s} \frac{1}{\tau_{\Pi} T} \partial_x^2 \delta u_z + \frac{1}{\tau_{\Pi}} \partial_t \delta u_z = 0$$

Take it seriously:

$$\label{eq:star} \begin{split} \omega &= -i \frac{\eta}{sT} k^2 + \dots \quad \text{and} \quad \omega = -i \frac{1}{\tau_{\Pi}} + i \frac{\eta}{sT} k^2 + \dots \\ \text{hydrodynamics} \quad \text{purely imaginary "quasinormal mode"} \end{split}$$

Generalization that adds  $\operatorname{Re}(\omega_{QNM})$ :  $\left((\frac{1}{T}\mathcal{D})^2 + 2\omega_I \frac{1}{T}\mathcal{D} + |\omega|^2\right) \Pi^{\mu\nu} = -\eta |\omega|^2 \sigma^{\mu\nu}$ 

1409.5087