Possible K-physics tests of LHCb anomalies

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Main line of argument based on Glashow, DG, Lane, PRL 2015



LHCb has performed several new measurements of $b \rightarrow s$ modes. Agreement with the SM is less than perfect. **Motivation:** LHCb's $b \rightarrow s$ data

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VS.

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Note

- muons are among the most reliable objects within LHCb
- the electron channel would be an obvious culprit (brems + low stats).
 But there is no disagreement

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0 + 0 + 6

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(within large errors)

There seems to be BSM LFNU

and the effect is in $\mu\mu$, not ee

Motivation 2

4

Actually, after some effective-theory insights, two further pieces of info support the above picture

 P'_{5} deficit in angular $B \rightarrow K^{*} \mu \mu$ data







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Basic observation:

• Without further assumptions, LFNU at a non-SM level implies LFV at a non-SM level.

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In fact:

Consider a new, LFNU interaction above the EWSB scale, e.g. with •

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• In what basis are quarks and leptons in the above interaction?

Generically, it's not the mass eigenbasis. (This basis doesn't yet even exist. We are above the EWSB scale.)

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Rotating q and l to the mass eigenbasis generates LFV interactions.



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Let's now turn to Q1: Can we (easily) make sense of data **1** to **5** ?

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Consider the following Hamiltonian

$$H_{\rm SM+NP}(\bar{b} \rightarrow \bar{s} \mu \mu) = -\frac{4 G_F}{\sqrt{2}} V_{tb}^* V_{ts} \frac{\alpha_{\rm em}}{4\pi} \left[\bar{b}_L \gamma^\lambda s_L \cdot \left(C_9^{(\mu)} \bar{\mu} \gamma_\lambda \mu + C_{10}^{(\mu)} \bar{\mu} \gamma_\lambda \gamma_5 \mu \right) \right]$$

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purely vector

purely axial

Let's now turn to Q1: Can we (easily) make sense of data **0** to **5** ? It is highly non-trivial that a simple consistent BSM picture exists to describe the above data 0 to 9 Consider the following Hamiltonian purely vector purely axial $H_{\rm SM+NP}(\bar{b} \rightarrow \bar{s} \mu \mu) = -\frac{4G_F}{\sqrt{2}} V_{tb}^* V_{ts} \frac{\alpha_{\rm em}}{4\pi} \left[\bar{b}_L \gamma^\lambda s_L \cdot \left(C_9^{(\mu)} \bar{\mu} \gamma_\lambda \mu \right) + C_{10}^{(\mu)} \bar{\mu} \gamma_\lambda \gamma_5 \mu \right]$ lepton current Note: $C_9^{\text{SM}}(m_b) \approx +4.2$ $C_{10}^{\text{SM}}(m_b) \approx -4.4$ $C_9^{\text{SM}}(m_b) \approx -C_{10}^{\text{SM}}(m_b)$ (i.e. in the SM also the lepton current has nearly V – A structure [Bobeth, Misiak, Urban, 99] [Khodjamirian et al., 10]















• Recalling our full Hamiltonian

$$H_{\rm SM+NP}(\bar{b} \rightarrow \bar{s}\mu\mu) = -\frac{4G_F}{\sqrt{2}} V_{tb}^* V_{ts} \frac{\alpha_{\rm em}}{4\pi} \left[\bar{b}_L \gamma^\lambda s_L \cdot \left(C_9^{(\mu)} \bar{\mu} \gamma_\lambda \mu + C_{10}^{(\mu)} \bar{\mu} \gamma_\lambda \gamma_5 \mu \right) \right]$$



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the shift to the C_{9} Wilson coeff. in the $\mu\mu$ -channel becomes

$$k_{\rm SM} C_9^{(\mu)} = k_{\rm SM} C_{9,\rm SM} + \frac{G}{2} (U_L^d)_{33}^* (U_L^d)_{32} |(U_L^t)_{32}|^2$$

Explaining b
$$\rightarrow$$
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The NP contribution has opposite sign than the SM one if

$$G\left(U_{L}^{d}\right)_{32} < 0$$

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• On the other hand, in the ee-channel

$$k_{\rm SM} C_9^{(e)} = k_{\rm SM} C_{9,\rm SM} + \frac{G}{2} (U_L^d)_{33}^* (U_L^d)_{32} |(U_L^t)_{31}|^2$$

Explaining b
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 s data
• Recalling our full Hamiltonian

$$H_{SM=NP}(\bar{b} \rightarrow \bar{s} \mu \mu) = \left[-\frac{4 G_F}{\sqrt{2}} V_{1b}^* V_B \frac{\alpha_{mn}}{4\pi} \right] [\bar{b}_L \gamma^\lambda s_L \cdot (C_9^{(\mu)} \bar{\mu} \gamma_\lambda \mu + C_{10}^{(\mu)} \bar{\mu} \gamma_\lambda \gamma_5 \mu)]$$
the shift to the C_9 Wilson coeff. in the $\mu\mu$ -channel becomes

$$k_{SM} C_9^{(\mu)} = \left[k_{SM} C_{9,SM} + \left[\frac{G}{2} (U_L^d)_{33}^* (U_L^d)_{32} | (U_L^f)_{32} |^2 \right]$$

$$= \beta_{SM} + \beta_{NP}$$
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the shift to the C_9 Wilson coeff. in the $\mu\mu$ -channel becomes

$$k_{SM} C_9^{(u)} = \left[k_{SM} C_{9,SM} + \frac{G}{2} (U_{L33}^{(u)} (U_{L32}^{(t)}) | U_{L32}^{(t)} |^2 \right]$$

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The NP contribution has opposite sign than the SM one if

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• On the other hand, in the ee-channel

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Explaining
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 data
• The above shifts to the $C_{u,vo}$ Wilson coeffs. imply

$$R_{\kappa} \approx \frac{|C_{9}^{(u)}|^{2} + |C_{10}^{(u)}|^{2}}{|C_{9}^{(e)}|^{2} + |C_{10}^{(e)}|^{2}} = \frac{2 \cdot (\beta_{SM} + \beta_{NP})^{2}}{2 \cdot \beta_{SM}^{2}}$$







D. Guadagnoli, B- and K-physics LFV



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EV model signatures

$$\mathbb{I} \quad \frac{BR(B^* \to K^* \mu e)}{BR(B^* \to K^* \mu \mu)} = \frac{\beta_{NP}^2}{(\beta_{SM} + \beta_{NP})^2} \cdot \frac{|(U_L^t)_{31}|^2}{|(U_L^t)_{32}|^2} \cdot 2$$







Analogous considerations hold for purely leptonic LFV decays

$$\boxed{ \frac{BR(B_s \rightarrow \mu e)}{BR(B_s \rightarrow \mu \mu)} = \frac{\beta_{\rm NP}^2}{(\beta_{\rm SM} + \beta_{\rm NP})^2} \cdot \frac{|(U_L^{\ell})_{31}|^2}{|(U_L^{\ell})_{32}|^2}}$$

 $\blacksquare Again, B_s \rightarrow \mu \tau \text{ would be even more promising}$















LFV model signatures

 $\mathbf{\nabla}$

An interesting signature outside B physics would be $K \rightarrow (\pi) \ell \ell'$

Note that, while at LHCb lots of K mesons are produced, they decay too late for the detector size (except the K_s)

The "K-physics analogue" of R_κ:



is long-distance dominated [see D'Ambrosio et al., 1998] hence potentially less promising

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See however Crivellin et al., 1601.00970 for a recent reappraisal





The interaction advocated in GGL

$$H_{\rm NP} = G \, \bar{b}'_L \gamma^{\lambda} b'_L \, \bar{\tau}'_L \gamma_{\lambda} \tau'_L$$

can also manifest itself in $K \to (\pi) \ell \ell'$, for example





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- $K_L^0 \rightarrow e^{\pm} \mu^{\mp}$ $K^+ \rightarrow \pi^+ e^{\pm} \mu^{\mp}$

Exp limits •

$$BR(K_{L}^{0} \rightarrow e^{\pm}\mu^{\mp}) < 4.7 \times 10^{-12}$$

$$BNL \ E871 \ Collab., \ PRL \ 1998$$

$$BR(K^{+} \rightarrow \pi^{+}\mu^{+}e^{-}) < 1.3 \times 10^{-11}$$

$$BNL \ E865 \ Collab., \ PRD \ 2005$$

$$BNL \ E865 \ Collab., \ PRL \ 2000$$

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$$\beta^{(K)} = \frac{G(U_L^d)_{32}^* (U_L^d)_{31} (U_L^t)_{31}^* (U_L^t)_{32}}{\frac{4G_F}{\sqrt{2}} V_{us}^*}$$





$$\beta^{(K)} = \frac{G(U_L^d)_{32}^* (U_L^d)_{31} (U_L^\ell)_{31}^* (U_L^\ell)_{32}}{\frac{4G_F}{\sqrt{2}} V_{us}^*}$$

$$|\beta^{(K)}|^2 = 2.15 \times 10^{-14}$$

(within model A of DG, Lane, PLB 2015)

I obtain

$$\frac{\Gamma(K_L^0 \rightarrow e^{\pm} \mu^{\mp})}{\Gamma(K^+ \rightarrow \mu^+ \nu_{\mu})} = \left|\beta^{(K)}\right|^2$$

$$\frac{\Gamma(K^{+} \rightarrow \pi^{+}\mu^{\pm}e^{\mp})}{\Gamma(K^{+} \rightarrow \pi^{0}\mu^{+}\nu_{\mu})} = 4 |\beta^{(K)}|^{2}$$



$$3^{(K)} = \frac{G(U_L^d)_{32}^* (U_L^d)_{31} (U_L^t)_{31}^* (U_L^t)_{32}}{\frac{4G_F}{\sqrt{2}} V_{us}^*}$$

$$\left|\beta^{(K)}\right|^{2} = 2.15 \times 10^{-14}$$
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I obtain

$$\frac{\Gamma(K_L^0 \rightarrow e^{\pm} \mu^{\mp})}{\Gamma(K^+ \rightarrow \mu^+ \nu_{\mu})} = |\beta^{(K)}|^2$$

$$BR(K_L^0 \rightarrow e^{\pm} \mu^{\mp}) \approx 6 \times 10^{-14}$$
with
$$BR(K^+ \rightarrow \mu^+ \nu_{\mu}) \approx 64\%$$

$$\Gamma(K^+)/\Gamma(K_L^0) \approx 4.2$$

$$\frac{\Gamma(K^{+} \rightarrow \pi^{+}\mu^{\pm}e^{\mp})}{\Gamma(K^{+} \rightarrow \pi^{0}\mu^{+}\nu_{\mu})} = 4 \left|\beta^{(K)}\right|^{2}$$



$$\beta^{(K)} = \frac{G(U_L^d)_{32}^* (U_L^d)_{31} (U_L^t)_{31}^* (U_L^t)_{32}}{\frac{4G_F}{\sqrt{2}} V_{us}^*}$$

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I obtain

$$\frac{\Gamma(K_{L}^{0} \rightarrow e^{\pm}\mu^{\mp})}{\Gamma(K^{+} \rightarrow \mu^{+}\nu_{\mu})} = |\beta^{(K)}|^{2} \qquad \qquad \frac{\Gamma(K^{+} \rightarrow \pi^{+}\mu^{\pm}e^{\mp})}{\Gamma(K^{+} \rightarrow \pi^{0}\mu^{+}\nu_{\mu})} = 4 |\beta^{(K)}|^{2}$$

$$\mathbb{R}(K_{L}^{0} \rightarrow e^{\pm}\mu^{\mp}) \approx 6 \times 10^{-14} \qquad \qquad \mathbb{R}(K^{+} \rightarrow \pi^{+}\mu^{\pm}e^{\mp}) \approx 3 \times 10^{-15}$$
with
$$\mathbb{R}(K^{+} \rightarrow \mu^{+}\nu_{\mu}) \approx 64\%$$

$$\Gamma(K^{+})/\Gamma(K_{L}^{0}) \approx 4.2$$
with
$$\mathbb{R}(K^{+} \rightarrow \pi^{0}\mu^{+}\nu_{\mu}) \approx 3\%$$



Frequently made objection: what about the SM? It has LFNU, but no LFV

Take the SM with zero v masses.

 Charged-lepton Yukawa couplings are LFNU, but they are diagonal in the mass eigenbasis (hence no LFV)

Or more generally, take the SM plus a minimal mechanism for v masses.

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• Physical LFV will appear in W couplings, but it's suppressed by powers of $(m_y / m_w)^2$

Bottom line: in the SM+v there is LFNU, but LFV is nowhere to be seen (in decays)

 But nobody ordered that the reason (=tiny m_v) behind the above conclusion be at work also beyond the SM

So, BSM LFNU \implies BSM LFV (i.e. not suppressed by m_{i})






LFV predictions in one of the two scenarios of [DG, Lane]

	$B^{\scriptscriptstyle +} o K^{\scriptscriptstyle +} \ \mu^{\scriptscriptstyle \pm} \ au^{\scriptscriptstyle \mp}$	$B^+ \rightarrow K^+ e^{\pm} \tau^{\mp}$	$B^{\scriptscriptstyle +} ightarrow K^{\scriptscriptstyle +} e^{\pm} \mu^{\mp}$
	$1.14 imes10^{-8}$	$3.84 imes10^{ ext{10}}$	$0.52 imes10^{-9}$
Exp:	$< 4.8 imes 10^{-5}$	$< 3.0 imes 10^{-5}$	$< 9.1 imes 10^{-8}$

	$B_{s} \rightarrow \mu^{\pm} \tau^{\mp}$	$B_{\rm s} \to e^{\pm} \tau^{\mp}$	$B_{\rm s} \rightarrow e^{\pm} \mu^{\mp}$
	$1.37 imes10^{-8}$	$4.57 imes10^{-10}$	$1.73 imes10^{-12}$
Exp:	—	—	$< 1.1 imes 10^{-8}$

All predictions are phase-space corrected.