The End of the Beginning Jorge Piekarewicz - Florida State University



beginning of the end. But it is, perhaps, the end of the beginning.

Neutron Skins of Nuclei: From Laboratory to Stars May 4-7, 2015





The Beginning of the End May 17-27, 2016

Thank you all for Coming!

ster inform

la Repubblica ocietà

Check in Check out

Xavier Roca Maza. Il liceo scientifico a Barcellona, poi un lavoro da elettricista, infine l'università e il trasloco a Milano

SEGUE DALLA PRIMA DI MILANO

ER FORTUNA Xavier non solo conosce anche la lingua di noi comuni esseri mortali, ma è pure dotato di una buona dose di santa pazienza. Come potrebbe altri-menti studiare per anni e anni sempre e solo la struttura del nucleo? Non annoia? «Assolutamente no! La materia è talmente complessa che n c'è la possibilità di stancar

c'e la ponde sorridendo — E l ______conferma il fatto che è dal pr mo decennio del XX secole da quando cioè fu scoperta l struttura interna all'atom che siamo in molti a studia lo, senza annoiarci». Tocca crederci. Anche per

ché lui Fisica se l'è proprio an data a cercare. Terminato il li ceo scientifico, infatti, aveva iniziato a lavorare. «Era inizia to come lavoretto estivo - ri corda — Un'azienda di mate-

riali elettrici, niente di che. Se non che ci avevo preso velocemente gusto ad avere la mia indipendenza economica. Dopo quattro anni — però — l'idea che la mia vita fosse tutta li mi ha spinto a rispolverare il mio antico interesse per il funzionamento delle cose e mi-sono iscritto a Fisica». Pura passione perché - come dice sorridendo — «fisica si studia solo per de vozione. Non si sceglie certo per i sol-

Eadesso, eccolo qui. Dal 2010 presso il dipartimento di Fisica dell'università Statale di Milano. «Prima con una borsa di studio dell'Istituto Nazio nale di fisica nucleare, da due anni ricercatore a tempo determinato spiega — Mi occupo di fisica teorica delle interazioni fondamentali, conduco le esercitazioni di elettromagnetismo». Direttamente da Barcellona. Cervellino in entrata. Perché? Dubito che in Italia i ricercatori abbiano vita più facile che in Spagna. Infatti. La precarietà come destino. Ma - siccomeanche i fisici hanno un cuore - lui qui ci è arrivato per amore che, nel suo caso, prende il nome di Benedetta, di Roma ma conosciuta a Barcellona mentre lui sta facendo il dottorato e lei il suo post dottorato. Fisica, pure

IL CUORE In Spagna ho conosciuto Benedetta, anche lei fisica, el'ho seguita a Torino e poi a Roma

ILLAVORO

Alla base c'è una

materia si studia solo

pura passione

perché questa

per devozione,

nonperisoldi



Va'dove tiporta ilprotone

per le due ruote se solo ci fossero più

piste ciclabili e meno rotaie del

tram», fa lunghe passeggiate a parco

In laboratorio serve il cervello mail fisico ha anche un cuore

lei. Questione di elettroni, immagino. tuto Italiano di Tecnologia e da un pa-«Quando Benedetta ha trovato lavoio di anni ci possiamo godere la città». ro in Italia ho deciso di seguirla». Pec-Qui si muove rigorosamente in bici («Milano sarebbe una città perfetta cato che lei lo trovi prima a Torino e poi a Roma, mentre lui finisce in via Celoria. «Per due anni ho fatto il pendolare Torino-Milano, per altri due ho diviso le mie settimane tra qui e Ro-Lambro e se proprio deve usare l'auma. Poi, finalmente, anche Benedetto, utilizza «l'ottimo servizio di carta ha trovato lavoro a Milano, presso sharing». «Qui sto bene: questa è una il Center for Genomic Science dell'Isti- città che funziona, è vivace, giovane e

c'è la possibilità di trovare lavoro di ottimo livello — racconta — Peccato il cielo sia spesso così grigio».

Nessuna nostalgia di Barcellona? «Ovviamente mi mancano familiari e amici, la luce e un clima che non conosce veramente l'inverno. Ma qui ho trovato un gruppo con il quale lavoro molto bene». Possiamo considerarti dei nostri? «Chi può dirlo? - risponde con lo sguardo di chi ha quotidianamente a che fare con relatività spazio temporale - Per me, ma forse non so lo, non è il luogo il criterio prioritario della scelta: conta l'amore e il lavoro». Insomma, va' dove ti porta il cuore (e il protone).

NSKINS-2015

Just before NSKINS-2016

Why is this Meeting Different from any other Meeting?

Neutron Skins of Nuclei

Dear Colleague:

We are planning a two-week program at Mainz during the May 17-27, 2016 period that will gather the stakeholders interested in the determination of the neutron skins of nuclei (e.g., Ca and Pb) and their impact on the density dependence of the symmetry energy. As you may have heard, a successful one-week workshop aimed at defining the goals and expectations for the upcoming two-week program took place in Mainz in May, 2015.

Although various meetings of this sort have already been organized, it is the primary goal of the workshop to establish quantitatively the strengths and limitations of the various experimental techniques. This requires a detailed analysis of the systematic errors. Moreover, given that in most instances theory must be used to connect the measured experimental observable to the neutron skin, it also essential to quantify the statistical and systematic errors associated with the given theory. Finally, we aim to design a suite of experiments and the precision required to determine neutron skins in a manner that will provide meaningful constraints on the density dependence of the symmetry energy — and ultimately on astrophysical observables. Thus, our interest for you to present your results and actively shape the success of the program.

Given that the Mainz Institute for Theoretical Physics (MITP) limits the number of participants to the program, we urge you to apply as soon as possible by visiting the program webpage at: <u>https://indico.mitp.uni-mainz.de/conferenceDisplay.py?confld=47</u>. There, you will be able to find material of relevance to the program — which will be constantly updated — including some of the slides from the May, 2015 meeting.

Kindest Regards

Chuck Horowitz, Jorge Piekarewicz, Concettina Sfienti, and Mark Vanderhaeghen

From Measurable Observables to the Neutron Skin

What is actually measured?

Cross section, asymmetry, spin observables, ...

- How is the measured observable connected to the neutron skin?
- What are the assumptions implicit in making this connection? Impulse approximation, off-shell ambiguities, distortion effects, ...
- How sensitive is the extraction of the neutron radius/skin to these assumptions?
- Quantitative assessment of both statistical and systematic errors

All observables are equal, but some observables are more equal than others ... Pedigree!





Theory Informing Experiment

Quantitative assessment of both statistical and systematic errors; theory must provide error bars!

Uncertainty quantification and covariance analysis (theoretical errors & correlations)

- Precision required in the determination of the neutron radius/skin?
 - As precisely as "humanly possible" fundamental nuclear structure property
 - To strongly impact Astrophysics?
 - What astrophysical observables to benchmark?
- Is there a need for a systematic study over "many" nuclei? PREX, CREX, SREX, ZREX, ...

Is there a need for more than one Q-square point? Radius and diffuseness ... the whole form factor?





Uncertainty Quantification



Need both; Systematic and Statistical Theoretical Errors

Nuclear Structure, Error Bars, and Correlations: Statistical

Reinhard-Nazarewicz, PRC81 (2010) 051303(R) Fattoyev-Piekarewicz, PRC86 (2012) 015802; PRC84 (2011) 064302

Empirical constants determined from optimization of a quality measure

$$\chi^{2}(\mathbf{p}) = \sum_{n=1}^{N} \left(\frac{\mathcal{O}_{n}^{(\text{th})}(\mathbf{p}) - \mathcal{O}_{n}^{(\text{exp})}}{\Delta \mathcal{O}_{n}} \right)^{2} = \chi^{2}(\mathbf{p}_{0}) + \mathbf{x}^{T} \hat{\Sigma}^{-1} \mathbf{x} + \dots$$

- Predictions accompanied by meaningful theoretical errors
- Ovariance analysis least biased approach to uncover correlations

 $\operatorname{Cov}(A,B) = \partial_i A \hat{\Sigma}_{ij} \partial_j B;$

Cov(A, B) $\operatorname{Corr}(A, B) = -$ Var(A)Var(B)



The Enormous Reach of the Neutron Skin: Covariance Analysis

- Neutron skin as proxy for neutron-star radii ... and more!
- Calibration of nuclear functional from optimization of a quality measure
- New era: predictability typical uncertainty quantification demanded
- Neutron skin strongly correlated to a myriad of neutron star properties: Radii, Enhanced Cooling, Moment of Inertia, ...





A Path Forward in Uncertainty Quantification



Castello di Trento ("Trint"), watercolour, 19,8 x 27,7, painted by A. Dürer on his way back from Venice (1495)

Information and Statistics in Nuclear Experiment and Theory ISNET-3 Trento, November 16-20, 2015

Main Topics

Estimation of statistical uncertainties of calculated quantities, assessment of systematic errors, validation and verification of extrapolations, information content of observables with respect to current theoretical models, statistical tools of nuclear theory and planning of future experiments, Bayesian methods and computational techniques, novel methods of optimization

Key Speakers

Anatoli Afanasjev (Mississippi State University, USA), Enrique Ruiz Arriola (University of Granada, Spain), Julia Bliss (Technical University of Darmstadt, Germany), Rick Casten (Yale University, USA), Gianluca Colo (University of Milan and INFN, Italy), Andreas Ekstrom (University of Tennessee, USA), Christian Forssen (Chalmers University of Technology, Sweden), Dick Furnstahl (Ohio State University, USA), Krzysztof Granada, Spain), Julia Bliss (Technical University of Jyväskylä, Finland), Dave Ireland (University of Glasgow, UK), Yannen Jaganathen (Michigan State University, USA), Markus Kortelainen (University of Jyväskylä, Finland), Dave Ireland (University of Glasgow, UK), Yannen Jaganathen (Michigan State University, USA), Markus Kortelainen (University of Jyväskylä and Helsinki Institute of Physics, Finland), Amy Lovell (Michigan State University, USA), Rodrigo Navarro-Perez (Lawrence Livermore National Laboratory, USA), Witold Nazarewicz (Michigan State University, USA), Nils Paar (University of Basel, Switzerland), Alessandro Pastore (University of York, UK), Jorge Piekarewicz (Florida State University, USA), Soctt Pratt (Michigan State University, USA), Nils Paar (University of Milan and INFN, Italy), Jan Ryckebusch (Ghent University of Frangen, Germany), David Richards (Jefferson Laboratory, USA), Xavier Roca-Maza (University of Milan and INFN, Italy), Jan Ryckebusch (Ghent University of Notre Dame, USA), Bartomic Jzate (Iniversity of Nuclear Physics PAN - Krakow), Sarah Wesolowski (Ohio State University, USA), Stefan Wild (Argonne National Laboratory, USA)

Organizers

David Ireland (University of Glasgow) Witold Nazarewicz (FRIB/NSCL - Michigan State University) Bartlomiej Szpak (Insitute of Nuclear Physics PAN - Krakow)

Director of the ECT*: Professor Wolfram Weise (ECT*)

The ECT* is sponsored by the "Fondazione Bruno Kessler" in collaboration with the "Assessorato alla Cultura" (Provincia Autonoma di Trento), funding agencies of EU Member and Associated States and has the support of the Department of Physics of the University of Trento.

For local organization please contact: Gianmaria Ziglio - ECT* Secretariat - Villa Tambosi - Strada delle Tabarelle 286 - 38123 Villazzano (Trento) - Italy Tel.:(+39-0461) 314721 Fax:(+39-0461) 314750, E-mail: ect@ectstar.eu or visit http://www.ectstar.eu INT Program INT-16-2a Bayesian Methods in Nuclear Physics June 13 - July 8, 2016



Electroweak Measurements

A huge, predicted atomic parity violation has now been observed in ytterbium, further aiding tabletop experimental searches for physics beyond the standard model that complement ongoing efforts at high-energy colliders.



PV elastic e-scattering

Atomic PV measured in long-chains (e.g., Yb) to eliminate uncertainties in atomic theory

Enormously Clean ... Extraordinarily Expensive!

- Weak FF determined in a model independent way (exactly as the Charge FF)
- Very strongly coupled to nSkin …
- "Mild" model dependence going from FF to nSkin (Theory Homework)
- Measuring the FF at two (or more) points (Experiment/Theory Homework)
- Measuring the FF in several nuclei? (Experiment/Theory Homework)

Understanding dispersive corrections (Theory Homework)

From Dark Matter to Neutron Stars

- Coherent elastic ν -Nucleus scattering has never been observed!
- Predicted shortly after the discovery of weak neutral currents
- Enormously challenging; must detect exceedingly slow recoils
- CEvNS (pronounced "7s") are backgrounds for DM searches
- CEvNS is coherent ("large") as it scales ~N²
- "Piggybacking" on the enormous progress in dark-matter searches





Coherent Elastic ν -Nucleus Scattering at the Spallation Neutron Source (ORNL) may become possible in the "not-so-distant" future



Electric Dipole Polarizability





Enormous progress in sight ...

- High quality data on a variety of nuclei at RCNP & GSI, such as Pb, Sn-isotopes, Ni, Ca... (Experiment Homework)
 - K,J, L, ... are not experimental observables! Extract K by reproducing data on GMR Extract L by reproducing data on GDR (Theory Homework)

One single compelling theoretical picture!

 Impedance matching between theory and experiment; e.g., quasi-D contribution (Experiment/Theory Homework)

Measure the full dynamic response to learn about FF (Experiment Homework)

 $S(q,\omega)$

Coherent Elastic Pion and Proton Scattering

PREX-II and CREX as Anchors for FRIB Physics

"One of the main science drivers of FRIB is the study of nuclei with neutron skins 3-4 times thicker than is currently possible. FRIB will provide rare isotopes to explore the properties of halos and skins. JLab uses parity violation to measure the neutron radius of stable lead and calcium nuclei. Studies of neutron skins at JLab and FRIB will help pin down the behavior of nuclear matter at densities below twice typical nuclear density" 2013 Subcommittee Report to NSAC

The Traditional Approach: Proton-Nucleus Scattering

- FRIB will scatter protons from radioactive nuclei in inverse kinematics
- Large and uncontrolled uncertainties in the reaction mechanism
- Enormous ambiguities yield an energy dependent neutron skin
- FRIB must use PREX-II and CREX as calibrating anchors!





Uncontrolled Uncertainties

Impulse Approximation
 Off-shell Ambiguities
 Distortion Effects

Potentially very useful observables Difficult Experiment/Theory Homework. *Very large skins of unstable neutron-rich nuclei a critical new direction!*

 $\mathcal{M}_{PWIA} \sim A_1(s,t)\rho_T(Q)/Q \approx A_1(s,t)\rho_V(Q)$

What else ?

Extra Credit Homework and Planning

What worked in 2015; what did not?
Deliverables for May 2016?
Breadth vs Depth for May 2016?
Goals for two weeks in May 2016?

	RAPID COMMUNICAT
PHYSICAL REVIEW C 92, 031305(R) (2015)	
Dipole polarizability of ¹²⁰ Sn and nuclear energy density functionals	
T. Hashimoto, ^{1,*} A. M. Krumbholz, ² PG. Reinhard, ³ A. Tamii, ¹ P. von Neumann-Cosel, ^{2,†} T. Adachi, ⁴ N. Aoi, ¹ C. A. Bertulani, ⁵ H. Fujita, ¹ Y. Fujita, ¹ E. Ganioğlu, ⁶ K. Hatanaka, ¹ E. Ideguchi, ¹ C. Iwamoto, ¹ T. Kawabata, ⁷ N. T. Khai, ⁸ A. Krugmann, ² D. Martin, ² H. Matsubara, ⁹ K. Miki, ¹ R. Neveling, ¹⁰ H. Okamura, ¹ H. J. Ong, ¹ I. Poltoratska, ² V. Yu. Ponomarev, ² A. Richter, ² H. Sakaguchi, ¹ Y. Shimbara, ¹¹ Y. Shimizu, ¹² J. Simonis, ² F. D. Smit, ¹⁰ G. Süsoy, ⁶ T. Suzuki, ¹ J. H. Thies, ¹³ M. Yosoi, ¹ and J. Zenihiro ¹²	

 PHYSICAL REVIEW C 92, 064304 (2015)

 Image: Second Structure
 Image: Second Structure

 Neutron skin thickness from the measured electric dipole polarizability in ⁶⁸Ni, ¹²⁰Sn, and ²⁰⁸Pb

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So Far!

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Neutron and weak-charge distributions of the ⁴⁸Ca nucleus

G. Hagen^{1,2*}, A. Ekström^{1,2}, C. Forssén^{1,2,3}, G. R. Jansen^{1,2}, W. Nazarewicz^{1,4,5}, T. Papenbrock^{1,2}, K. A. Wendt^{1,2}, S. Bacca^{6,7}, N. Barnea⁸, B. Carlsson³, C. Drischler^{9,10}, K. Hebeler^{9,10}, M. Hjorth-Jensen^{4,11}, M. Miorelli^{6,12}, G. Orlandini^{13,14}, A. Schwenk^{9,10} and J. Simonis^{9,10}

The power of two: Assessing the impact of a second measurement of the weak-charge form factor of ²⁰⁸Pb

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The power of two: Assessing the impact of a second measurement of the weak-charge form factor of ²⁰⁸Pb

One Point: PREX/PREX-II/SPREX





Two Points: Hyper-PREX



$R_{ch} = 5.501 2(13) \, \text{fm}$



Charge FF as a proxy for weak FF

- The End of the Beginning: MITP, May 4-7, 2015
 The Beginning of the End: MITP, May 17-27, 2016
 The End: Sometime around the beginning of 2017 Culminating with a Topical Review in JPG
- What is actually measured? Cross section, asymmetry, spin observables, ...
- How is the measured observable connected to the neutron skin?
- What are the assumptions implicit in making this connection? Impulse approximation, off-shell ambiguities, distortion effects
- How sensitive is the extraction of the neutron radius/skin to these assumptions?
- Quantitative assessment of both statistical and systematic errors





Let's untangle the knots!







Our gratitude to MITP and all of you for coming; let's have a fun and productive meeting!