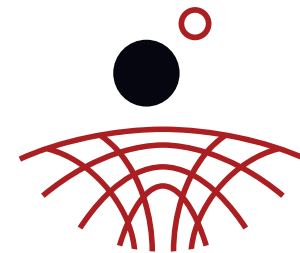
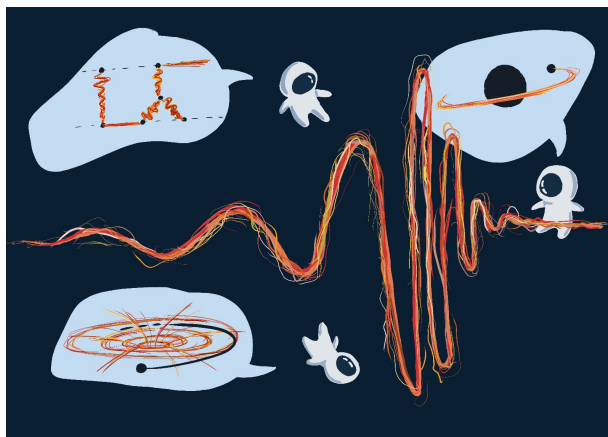


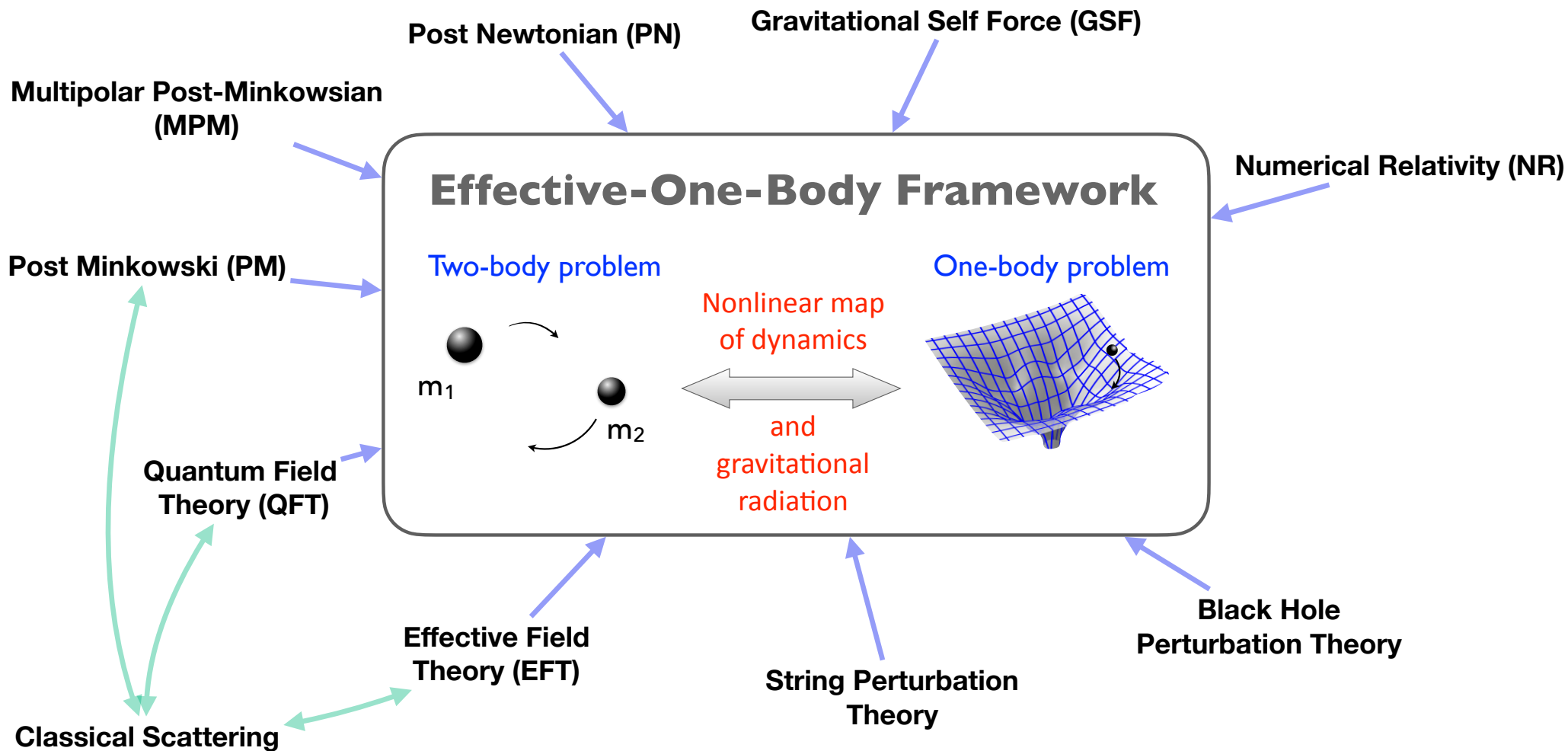
Efficient Eccentric EOB Dynamics via Near-Identity Averaging Transformations

Philip Lynch

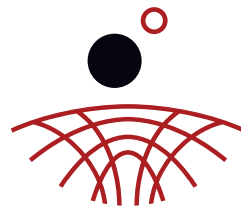
Contributors: Alessandra Buonanno, Maarten van de Meent,
Aldo Gamboa



**Astrophysical and
Cosmological Relativity**
Max Planck Institute
for Gravitational Physics



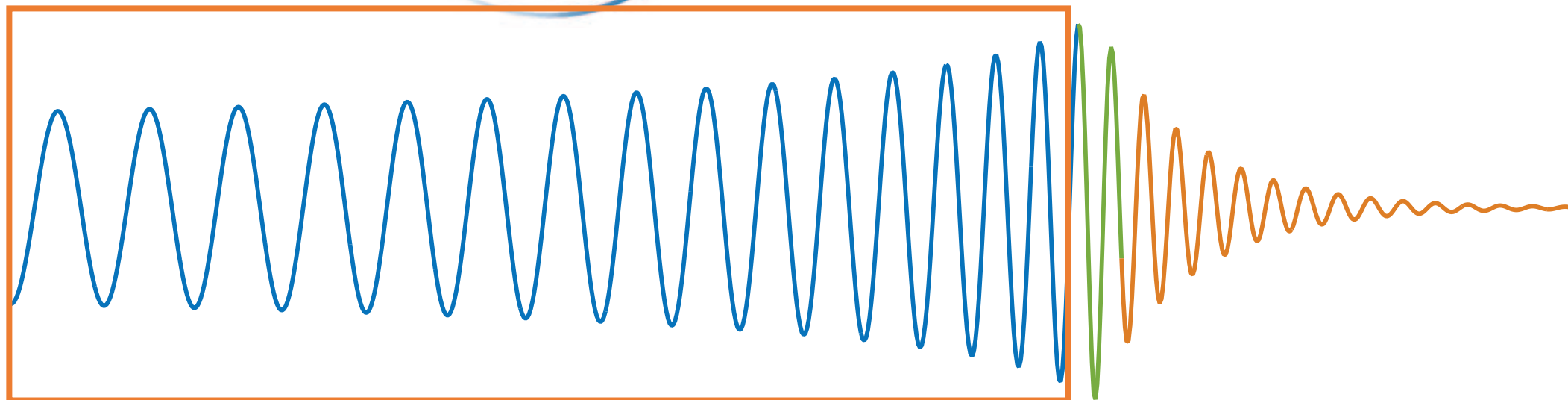
[LISA WF Whitepaper: [arXiv: 2311.01300](https://arxiv.org/abs/2311.01300)]



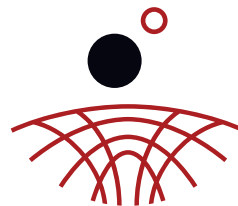
Inspiral

Merger

Ringdown

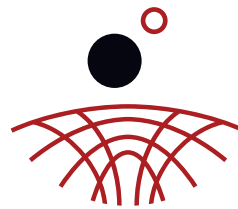


[Antelis & Moreno 1610.03567]



Motivation

- Need to model long eccentric **MBHB** signals for **LISA**
- **Quasicircular** dynamics are fast w/ "**Post-Adiabatic expansion**"
 - [Nagar & Rettengo 18],[Rettengo +19] [Reimenschneider+21], [Mihaylov+21]
- **Eccentric dynamics** are slow
- Borrow techniques from **EMRI** modelling:
 - Speed up the **inspiral** section
 - **Osculating Orbital Elements** (OOEs)
 - **Near Identity Averaging Transformations** (NITs)
- Proof of concept:
 - h_{22} for *non-spinning eccentric* binaries



EOB Inspiral Dynamics

$$H_{\text{EOB}} = M \sqrt{1 + 2\nu(\hat{H}_{\text{eff}} - 1)}$$

$$\hat{H}_{\text{eff}} = \sqrt{p_{r_*}^2 + A(\nu, r) \left(1 + \frac{p_\phi^2}{r^2} + Q(\nu, r, p_{r_*}) \right)}$$

$$p_{r_*} = A(\nu, r) \sqrt{\bar{D}(\nu, r)} p_r$$

$$\frac{dr}{dt} = \frac{\partial p_{r_*}}{\partial p_r} \frac{\partial H_{\text{EOB}}}{\partial p_{r_*}} \quad \frac{d\phi}{dt} = \frac{\partial H_{\text{EOB}}}{\partial p_\phi}$$

$$\frac{dp_{r_*}}{dt} = -\frac{\partial p_{r_*}}{\partial p_r} \frac{\partial H_{\text{EOB}}}{\partial r} + \mathcal{F}_r(\nu, r, p_r, p_\phi)$$

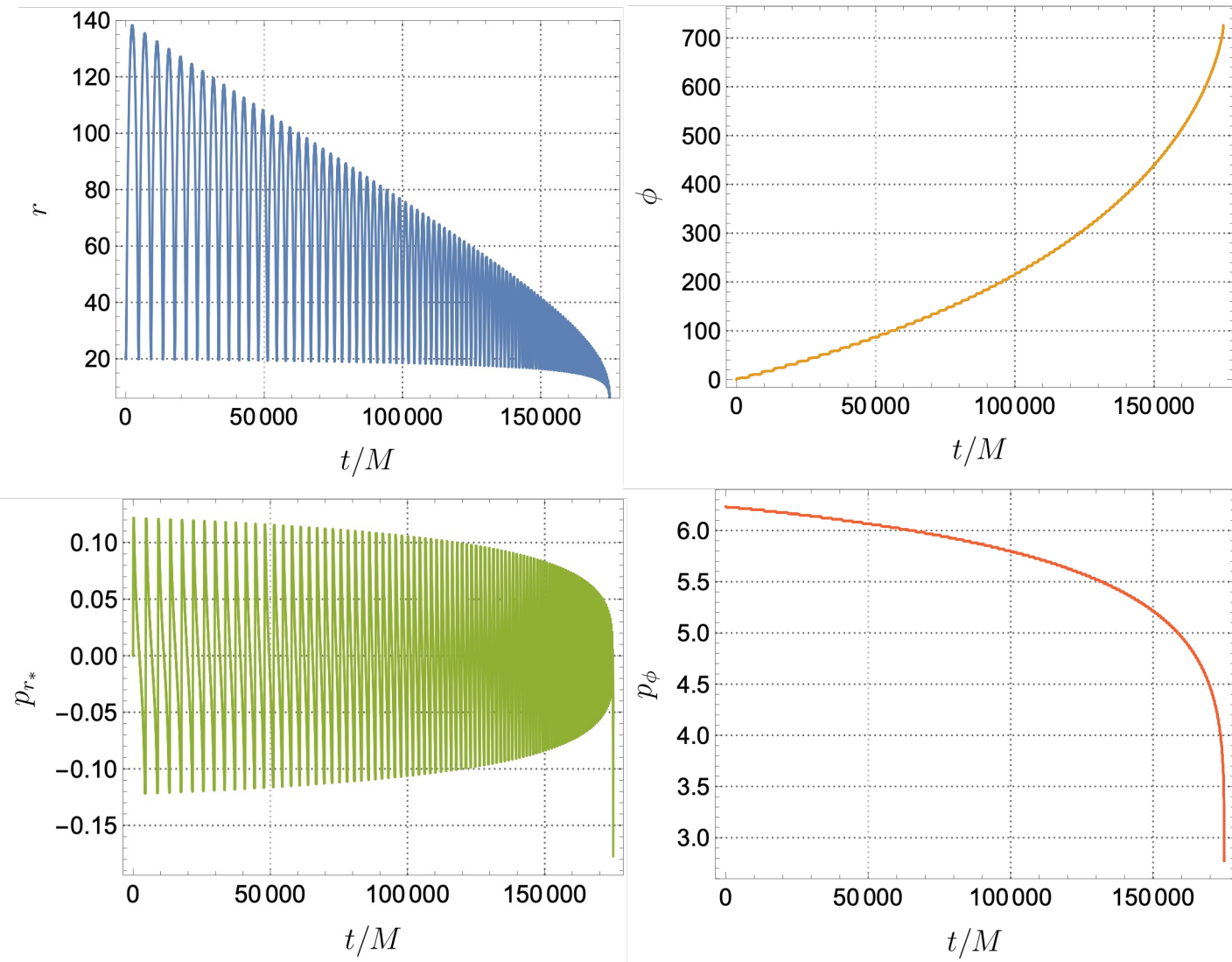
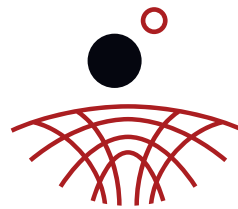
$$\frac{dp_\phi}{dt} = \mathcal{F}_\phi(\nu, r, p_r, p_\phi)$$

$$h(t) = \frac{1}{d_L} \sum_{\ell=2}^{\ell_{\text{max}}} \sum_{m=-\ell}^{\ell} h_{\ell m}(\nu, r, \phi, p_r, p_\phi) {}_{-2}Y_{\ell m}(\Theta, \Phi)$$

- Potentials: SEOBNRv5

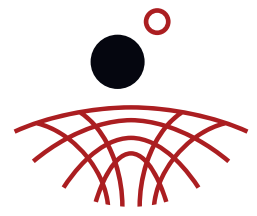
- RR. Forces: 3PN Eccentric

- WF Modes: 3PN Eccentric



- $m_1 = m_2 = 30 M_\odot @ 10\text{Hz}$
- $(v, p_0, e_0 \xi_0) = (0.25, 34.94, 0.75, 0)$
- Using RK4 adaptive time stepper
- Steps taken: 86,567
- Dynamics time: 3.34s

$(r(t), \phi(t), p_r(t), p_\phi(t))$



$h_{22}(t)$

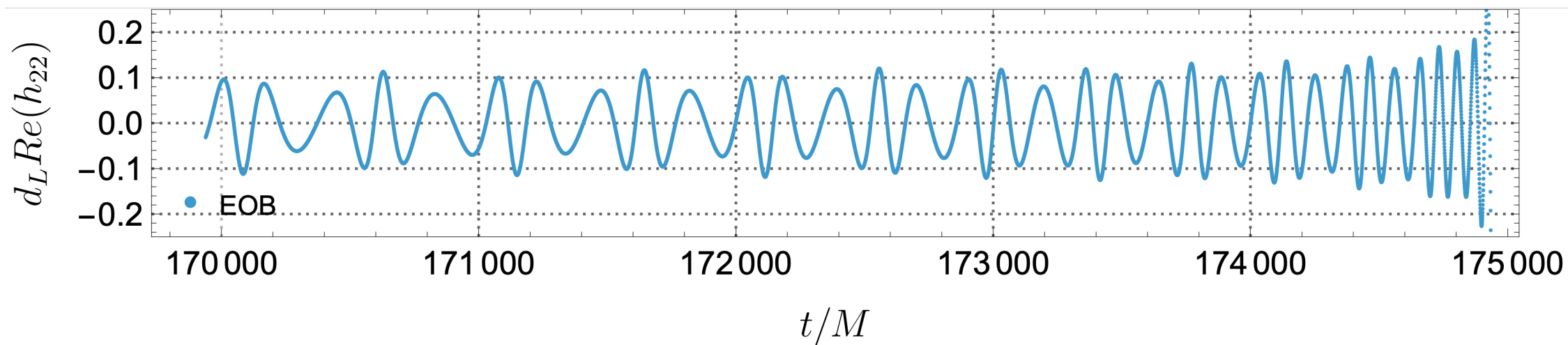
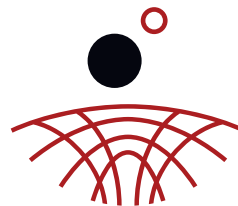


$h_{22}(N \Delta t)$

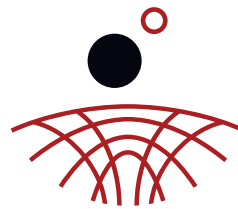


$$\Delta t \leq \frac{1}{2 f_{max}}$$

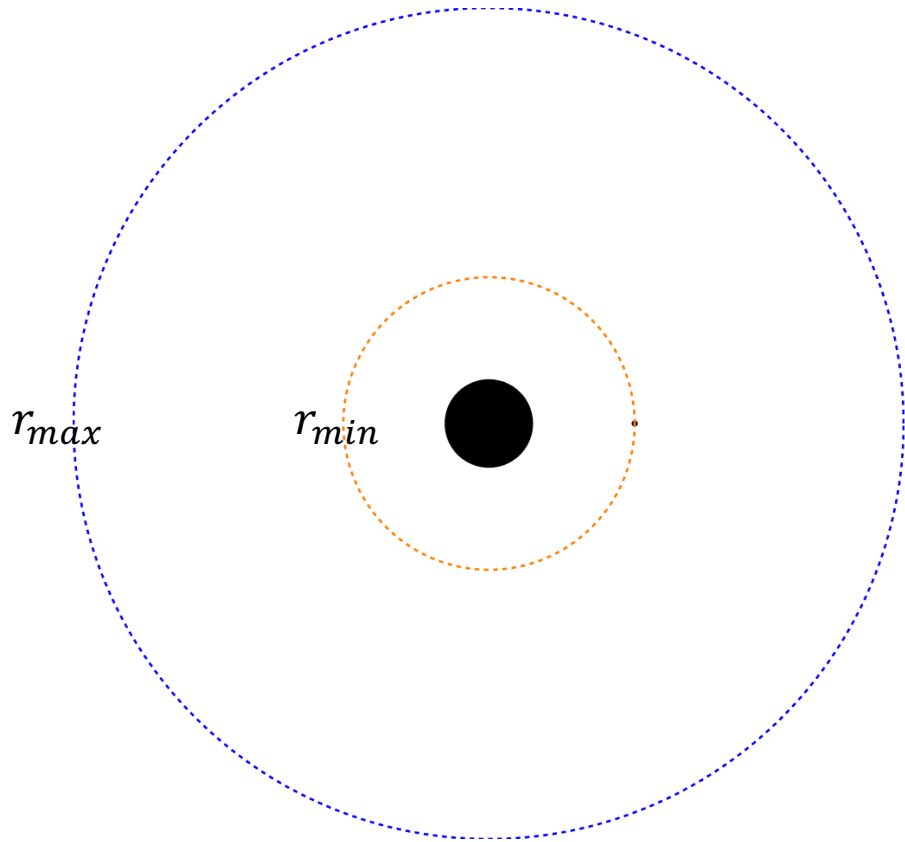
$\hat{h}_{22}(f)$



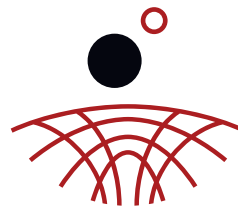
- WF sampled $\Delta t = 0.5 M$
- WF Generation time: $\sim 2.68s$
- Time to beat: $6.02s$



Quasi-Keplerian Parametrization



- Semilatus Rectum: $p = \frac{2r_{max}r_{min}}{r_{max}+r_{min}}$
- Eccentricity: $e = \frac{r_{max}-r_{min}}{r_{max}+r_{min}}$
- Relativistic anomaly: $\xi = \chi - \chi_0$
- $r = \frac{p}{1+e \cos \xi}$
- $H_{eff}(p, e), p_\phi(p, e), p_r(p, e, \xi)$
 - [Hinderer & Babak 17]



Osculating Orbital elements

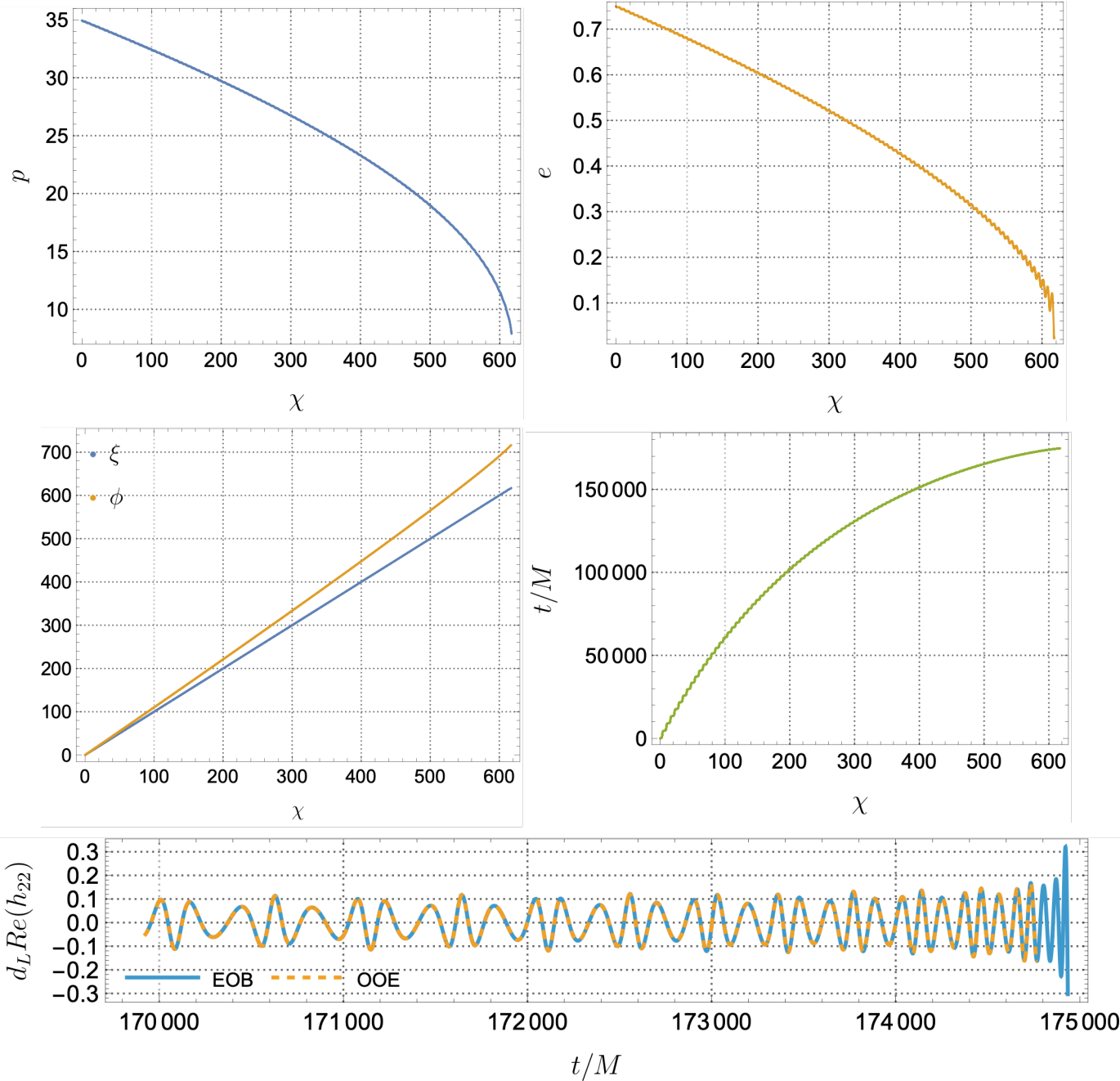
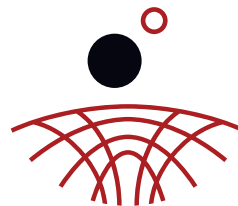
- Constants of motion p & e initial phase χ_0
- Orbital Elements: $I^A = \{p, e, \chi_0\} \rightarrow I(t)^A$
- Osculating Conditions:

1. $x^\alpha(t) = x_{cons}^\alpha(I^A(t), t)$

2. $u^\alpha(t) = u_{cons}^\alpha(I^A(t), t)$

$$\begin{aligned} \frac{dp}{d\chi} &= \epsilon F_p & \frac{de}{d\chi} &= \epsilon F_e \\ \frac{d\xi}{d\chi} &= 1 + \epsilon f_\xi^{(1)} \\ \frac{d\phi}{d\chi} &= f_\phi & \frac{dt}{d\chi} &= f_t \end{aligned}$$

- Bookkeeping parameter: ϵ
- Functions are **interpolants** of (ν, p, e)



- $(\nu, p_0, e_0, \xi_0) = (0.25, 34.94, 0.75, 0)$

- Dynamics time: 218.9s

- Steps taken: 3,202

- WF Generation time: ~ 1.13 s

- Mismatch: 8.01×10^{-6}



Near-Identity (Averaging) Transformations

Transform to new variables in series exp. in ϵ

$$\tilde{p} = p + \epsilon Y_p^{(1)} + \epsilon^2 Y_p^{(2)} + \mathcal{O}(\epsilon^3)$$

$$\tilde{e} = e + \epsilon Y_e^{(1)} + \epsilon^2 Y_e^{(2)} + \mathcal{O}(\epsilon^3)$$

$$\tilde{\xi} = \xi + \epsilon X_\xi^{(1)} + \mathcal{O}(\epsilon^2)$$

$$\tilde{\phi} = \phi + Z_\phi^{(0)} + \epsilon Z_\phi^{(1)} + \mathcal{O}(\epsilon^2)$$

$$\tilde{t} = t + Z_t^{(0)} + \epsilon Z_t^{(1)} + \mathcal{O}(\epsilon^2)$$

To obtain equations of motion independent of ξ

$$\frac{d\tilde{p}}{d\chi} = \epsilon \tilde{F}_p^{(1)} + \epsilon^2 \tilde{F}_p^{(2)} + \epsilon^3 \tilde{F}_p^{(3)} + \mathcal{O}(\epsilon^4)$$

OPA

$$\frac{d\tilde{e}}{d\chi} = \epsilon \tilde{F}_e^{(1)} + \epsilon^2 \tilde{F}_e^{(2)} + \epsilon^3 \tilde{F}_e^{(3)} + \mathcal{O}(\epsilon^4)$$

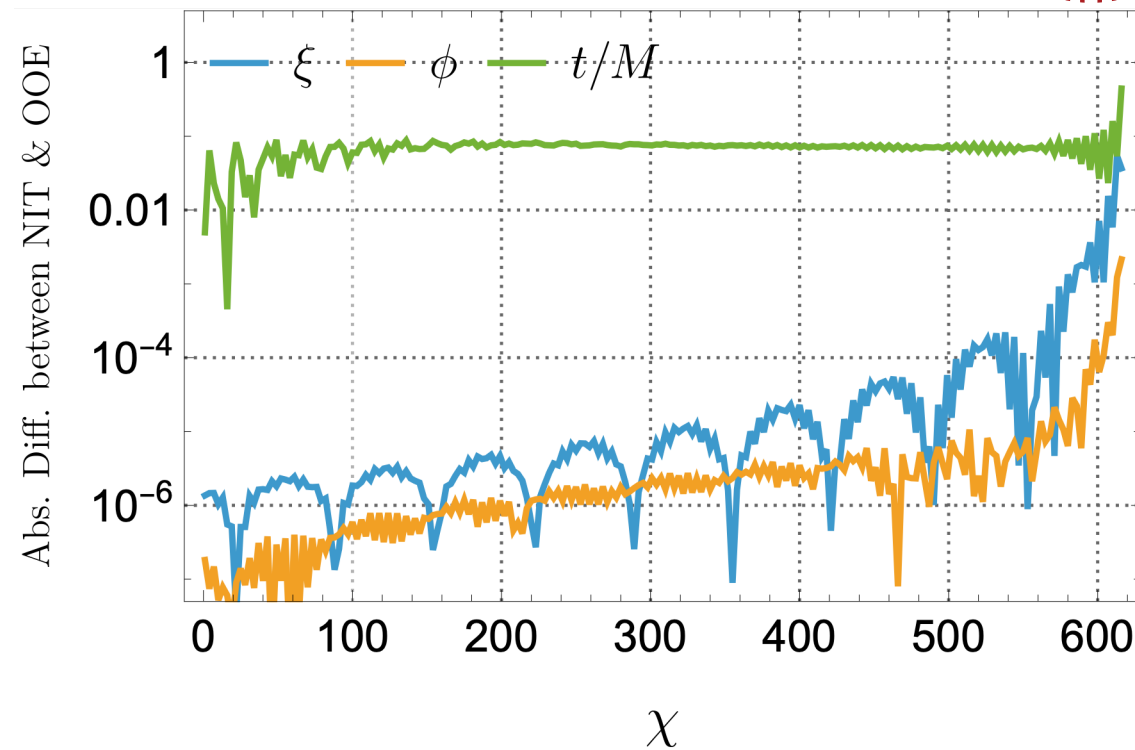
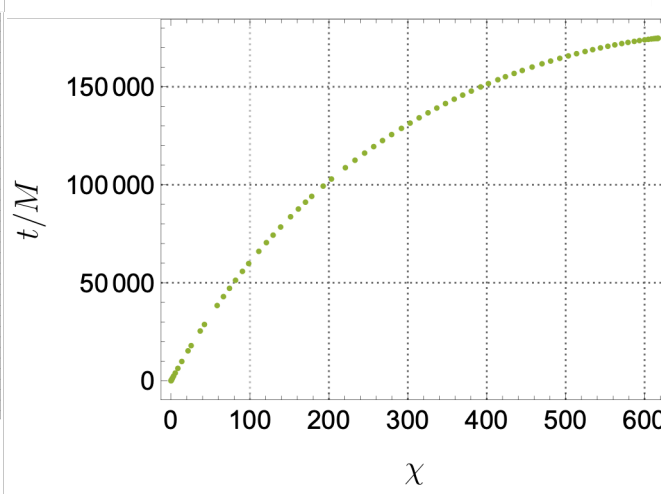
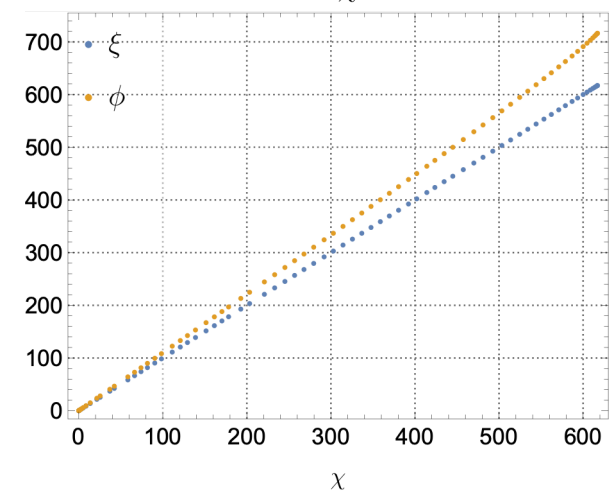
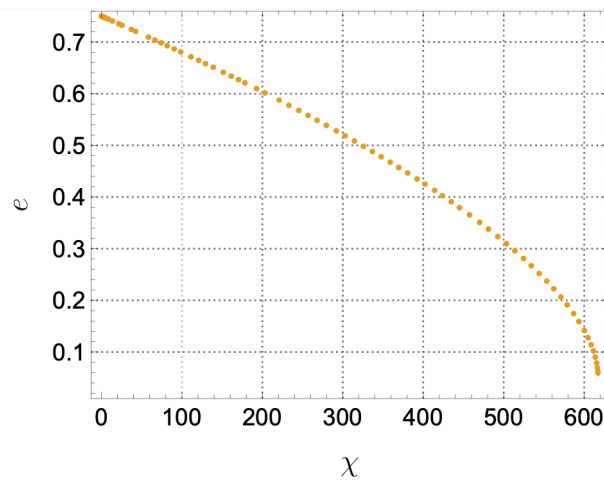
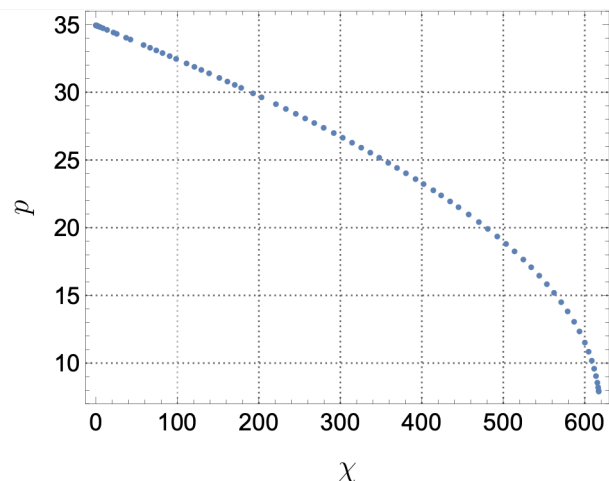
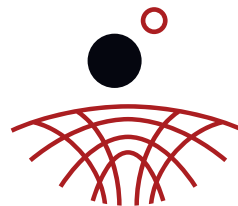
$$\frac{d\tilde{\xi}}{d\chi} = 1 + \epsilon f_\xi^{(1)} + \epsilon^2 f_\xi^{(2)} + \mathcal{O}(\epsilon^3)$$

1PA

$$\frac{d\tilde{\phi}}{d\chi} = \tilde{f}_\phi^{(0)} + \epsilon \tilde{f}_\phi^{(1)} + \epsilon^2 \tilde{f}_\phi^{(2)} + \mathcal{O}(\epsilon^3)$$

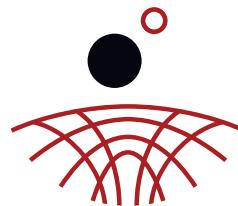
$$\frac{d\tilde{t}}{d\chi} = \tilde{f}_t^{(0)} + \epsilon \tilde{f}_t^{(1)} + \epsilon^2 \tilde{f}_t^{(2)} + \mathcal{O}(\epsilon^3)$$

2PA

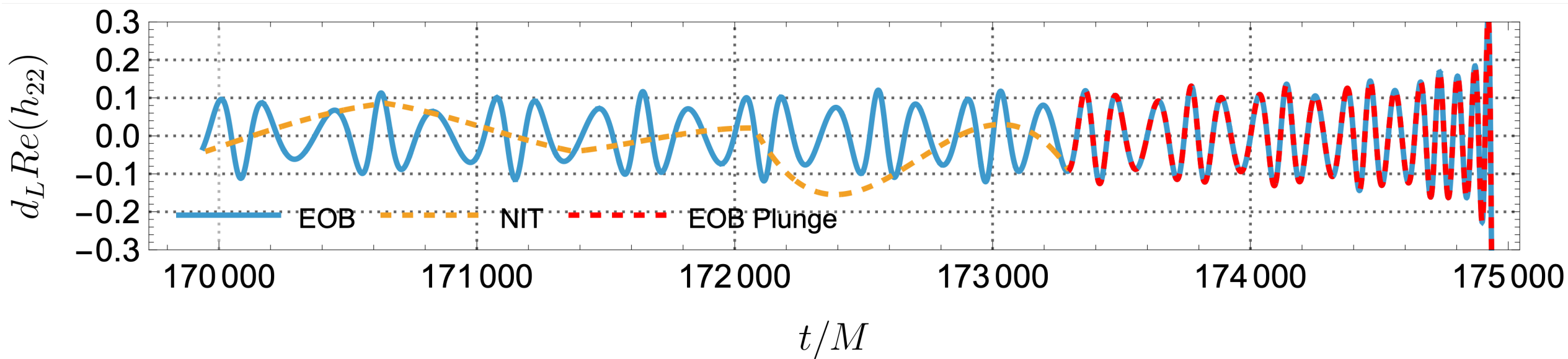


- Steps taken: 75
- Dynamics time: $\sim 0.058s$
- 2 orders of magnitude of speed up

- NIT breaks down close to LSO
- Stop NIT when $Q = \frac{\Omega^2}{\dot{\Omega}} < 100$
- Use EOB for Transition to Plunge



Generating Waveforms with NIT dynamics



- Steps: 66

$$(\tilde{p}(\chi), \tilde{e}(\chi), \tilde{\chi}(\chi), \tilde{\phi}(\chi))$$



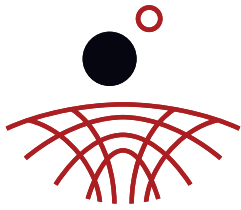
$$h_{22}(\tilde{t}(\chi) - Z_t^{(0)} - \epsilon Z_t^{(1)})$$

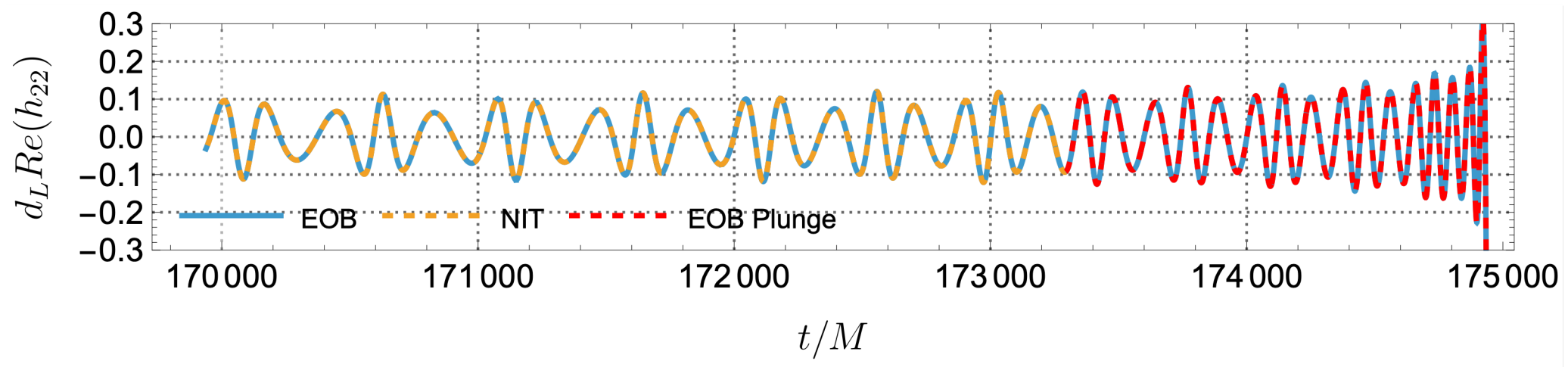
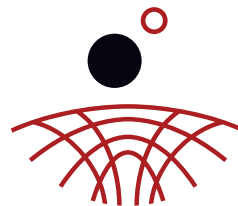


$$h_{22}(N \Delta t)$$



$$\hat{h}_{22}(f)$$





WF Evaluations

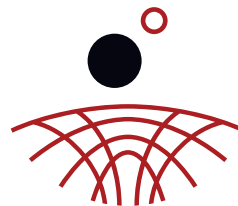
- NIT: 6,406
- EOB: 86,567

Total time (Dynamics + WF)

- NIT 2.2s
- EOB: 6.02s
- Speed-up: $\sim \times 2.74$

Mismatch

- 2.48×10^{-6}



What order PA Order do we Need?

Transform to new variables in series exp. in ϵ

$$\tilde{p} = p + \epsilon Y_p^{(1)} + \epsilon^2 Y_p^{(2)} + \mathcal{O}(\epsilon^3)$$

$$\tilde{e} = e + \epsilon Y_e^{(1)} + \epsilon^2 Y_e^{(2)} + \mathcal{O}(\epsilon^3)$$

$$\tilde{\xi} = \xi + \epsilon X_\xi^{(1)} + \mathcal{O}(\epsilon^2)$$

$$\tilde{\phi} = \phi + Z_\phi^{(0)} + \epsilon Z_\phi^{(1)} + \mathcal{O}(\epsilon^2)$$

$$\tilde{t} = t + Z_t^{(0)} + \epsilon Z_t^{(1)} + \mathcal{O}(\epsilon^2)$$

To obtain equations of motion independent of ξ

$$\frac{d\tilde{p}}{d\chi} = \epsilon \tilde{F}_p^{(1)} + \epsilon^2 \tilde{F}_p^{(2)} + \epsilon^3 \tilde{F}_p^{(3)} + \mathcal{O}(\epsilon^4)$$

$$\frac{d\tilde{e}}{d\chi} = \epsilon \tilde{F}_e^{(1)} + \epsilon^2 \tilde{F}_e^{(2)} + \epsilon^3 \tilde{F}_e^{(3)} + \mathcal{O}(\epsilon^4)$$

$$\frac{d\tilde{\xi}}{d\chi} = 1 + \epsilon f_\xi^{(1)} + \epsilon^2 f_\xi^{(2)} + \mathcal{O}(\epsilon^3)$$

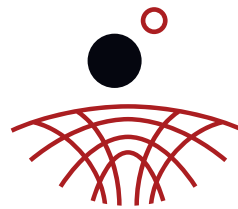
$$\frac{d\tilde{\phi}}{d\chi} = \tilde{f}_\phi^{(0)} + \epsilon \tilde{f}_\phi^{(1)} + \epsilon^2 \tilde{f}_\phi^{(2)} + \mathcal{O}(\epsilon^3)$$

$$\frac{d\tilde{t}}{d\chi} = \tilde{f}_t^{(0)} + \epsilon \tilde{f}_t^{(1)} + \epsilon^2 \tilde{f}_t^{(2)} + \mathcal{O}(\epsilon^3)$$

OPA

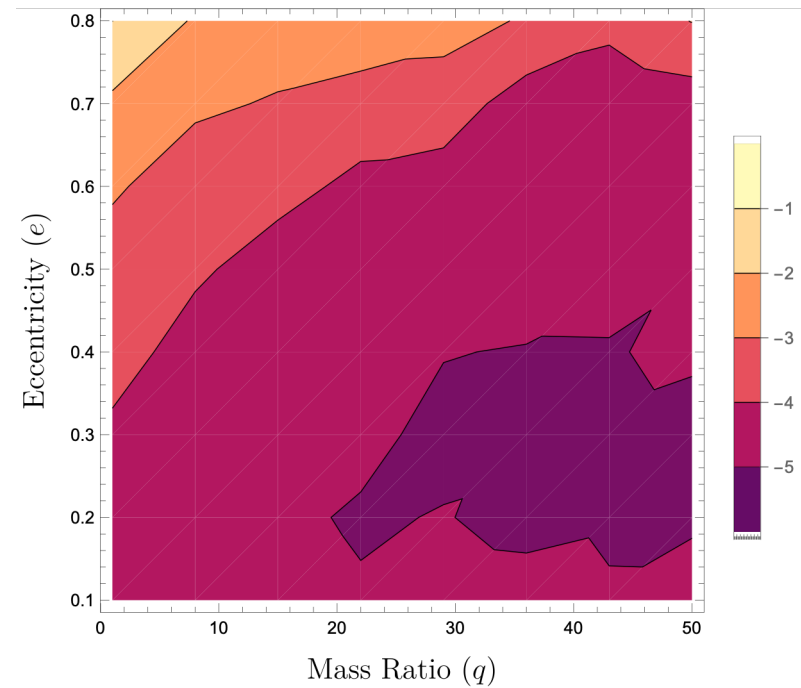
1PA

2PA

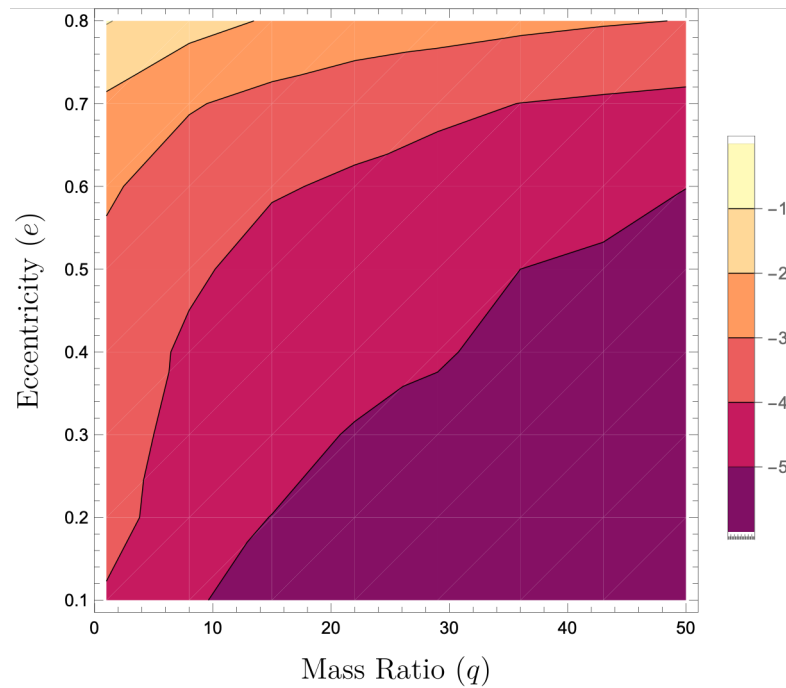


What order PA Order do we Need?

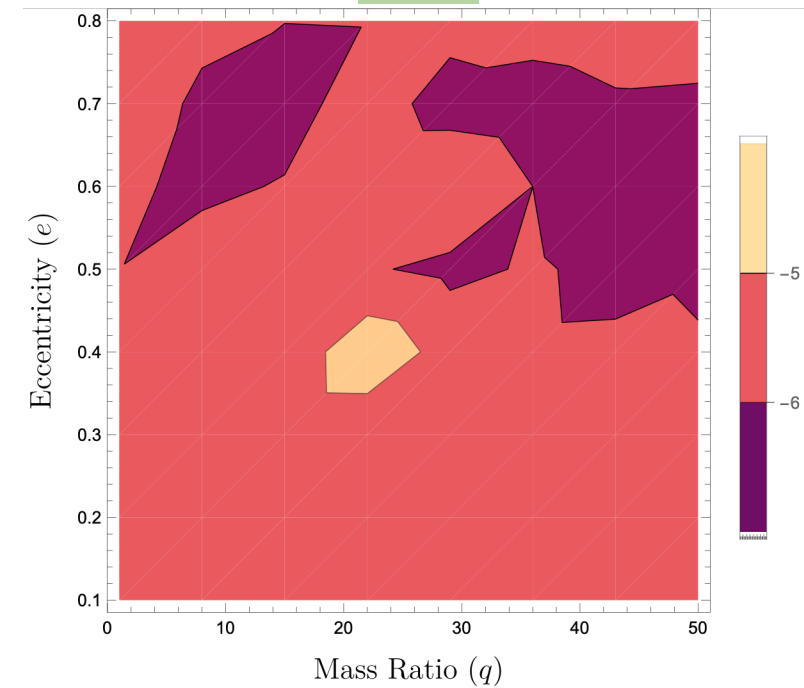
OPA



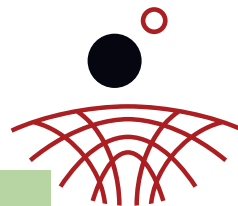
1PA



2PA

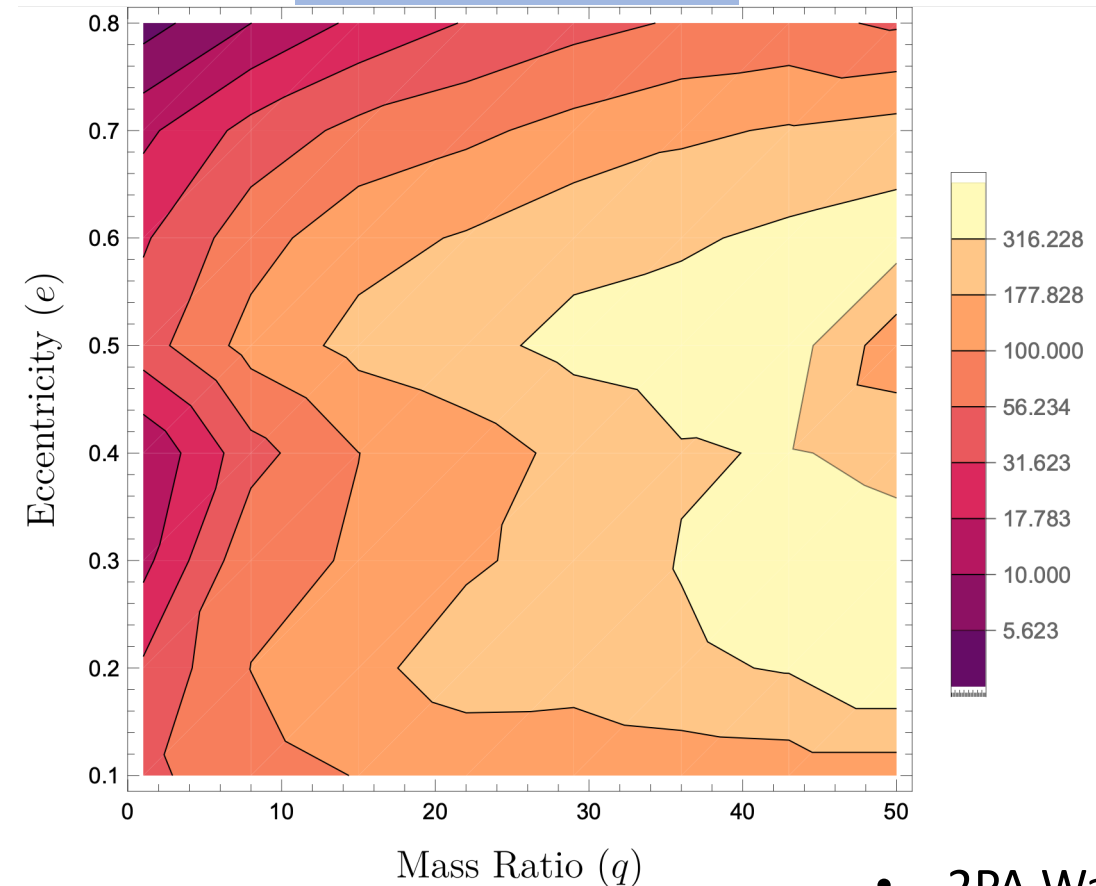


$M = 60 @ 10 \text{ Hz}$

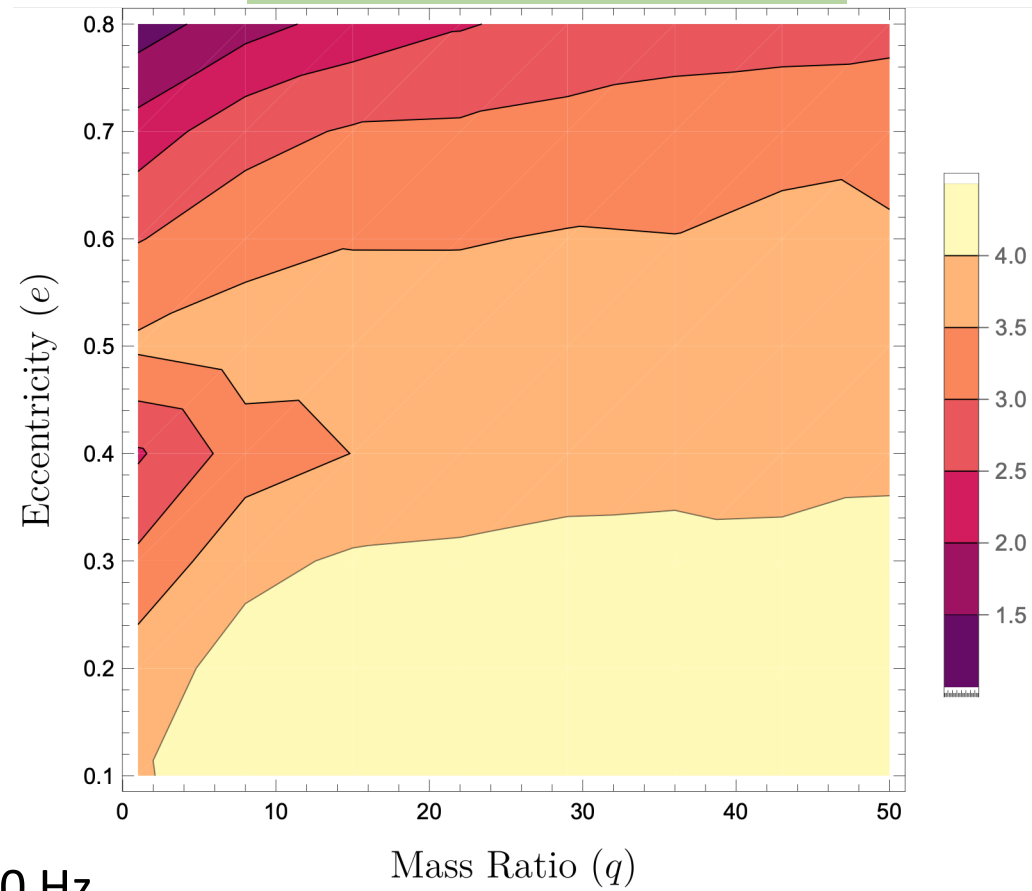


Speedup?

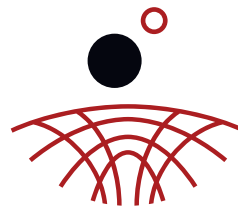
Dynamics



Dynamics + Waveform Generation



- 2PA Waveforms 60M @ 10 Hz
- Dynamics: 4- 540 times speedup
- Walltime: 1.2 – 5 times speedup



Future Work

Near term

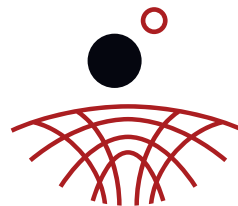
- Add **higher modes**
- **Small e** needs different orbital elements
- More rigorous testing

Medium-term

- Adding **aligned spins**
- Need **fast multidimensional interpolation** method
- Copy tech. from **FastEMRIWaveforms** Package

Long-term

- Combine NIT with ν expansion
- **Hybrid EOB-GSF 1PA** model for EMRIs
- **Calibrate EOB fluxes** to ecc. 0PA fluxes



Summary

- Using **OOEs and NITs** speeds up the inspiral **dynamics** by **a factor of q**
- WF generation sped up by $\sim 1.2 - 5$
- 0PA and 1PA give not accurate enough, but 2PA gives $\mathcal{M} \sim 10^{-5}$
- Opens up new **synergies** between **EOB and EMRI WF** modelling