

Neutrino quantum kinetics: Theory and (some) results

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IFIC, Valencia*

**Open Quantum Systems
Workshop**

29/04/2026



Classical transport

Distribution function

$$f(\mathbf{x}, \mathbf{p}, t)$$

Boltzmann equation

$$\frac{\partial f}{\partial t} + \mathbf{v} \cdot \frac{\partial f}{\partial \mathbf{x}} + \dot{\mathbf{p}} \cdot \frac{\partial f}{\partial \mathbf{p}} = \mathcal{C}[f]$$

Collision term:

$$\mathcal{C}[f_1] = \frac{1}{2p_1} \int \frac{d^3 \mathbf{p}_2}{(2\pi)^3 2E_2} \frac{d^3 \mathbf{p}_3}{(2\pi)^3 2E_3} \frac{d^3 \mathbf{p}_4}{(2\pi)^3 2E_4} (2\pi)^4 \delta^{(4)}(p_1 + p_2 - p_3 - p_4) \langle |\mathcal{M}|^2 \rangle F(1, 2, 3, 4)$$

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$$f(\mathbf{x}, \mathbf{p}, t)$$

Boltzmann equation

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$$F(1, 2, 3, 4) = (1 - f_1)(1 - f_2)f_3f_4 - (1 - f_3)(1 - f_4)f_1f_2$$

$$f_i = f_{\phi_i}(\mathbf{x}, \mathbf{p}_i, t)$$

Electrons, nucleons, neutrinos...

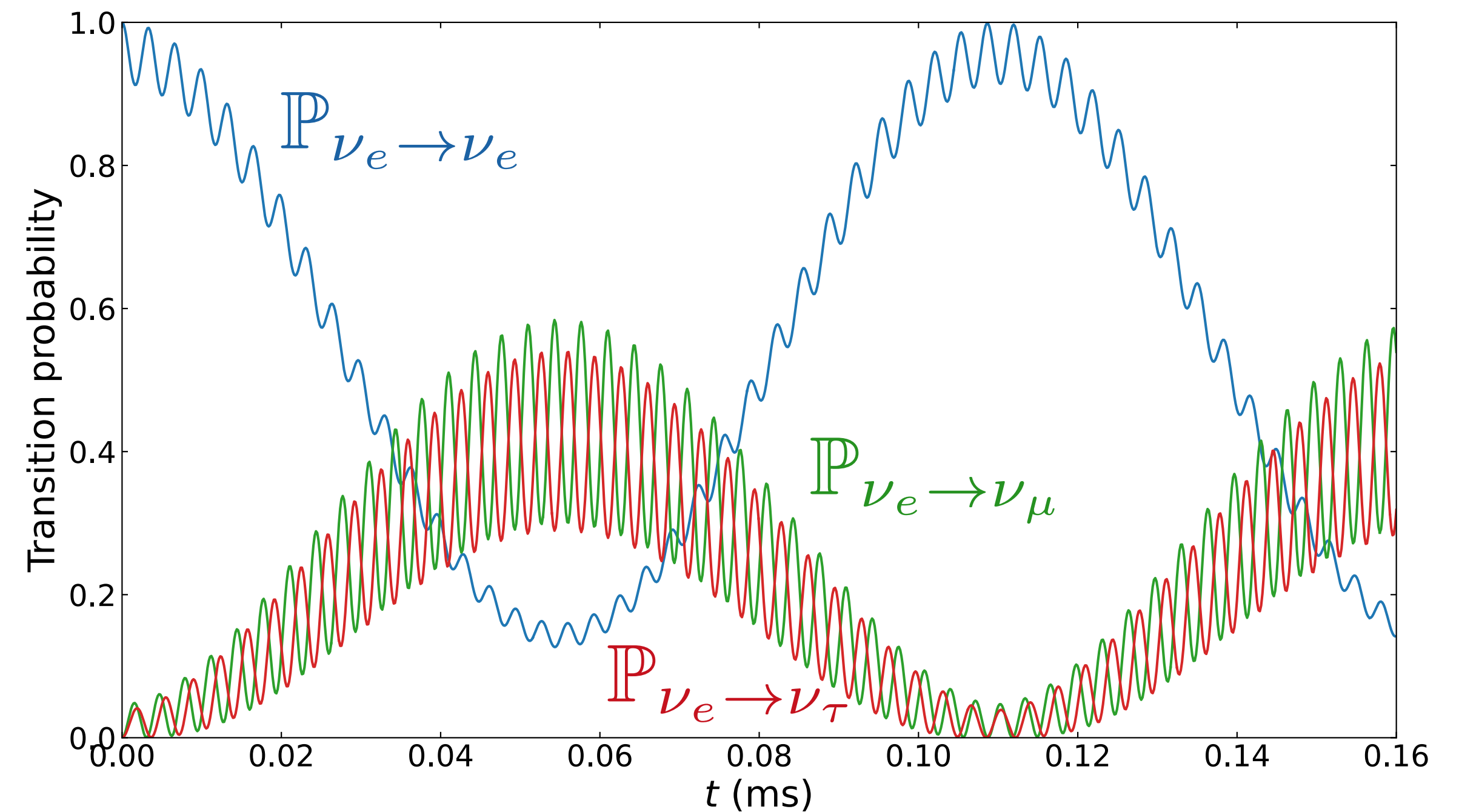
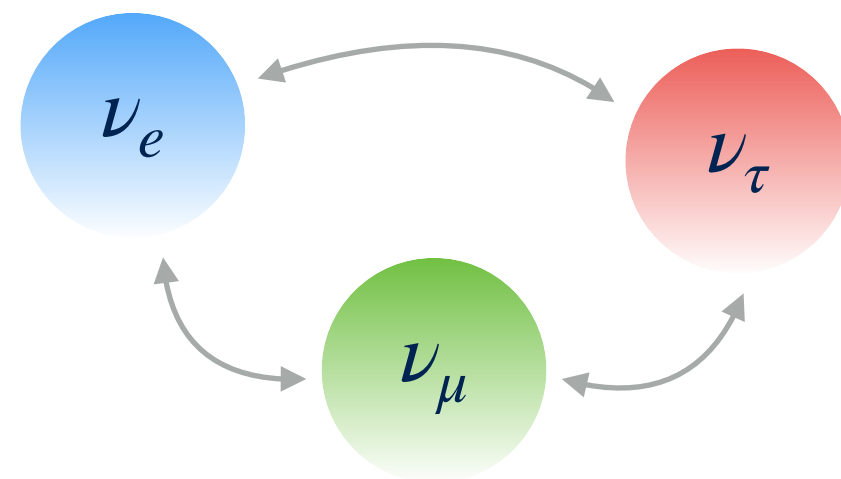
Neutrino flavor oscillations

- Standard Model: 3 species of massless neutrinos ν_L
- Experimental evidence in the second half of the 20th century
→ massive neutrinos

Flavor states $\begin{pmatrix} \nu_e \\ \nu_\mu \\ \nu_\tau \end{pmatrix} = U \begin{pmatrix} \nu_1 \\ \nu_2 \\ \nu_3 \end{pmatrix} = \begin{pmatrix} U_{e1} & U_{e2} & U_{e3} \\ U_{\mu1} & U_{\mu2} & U_{\mu3} \\ U_{\tau1} & U_{\tau2} & U_{\tau3} \end{pmatrix} \begin{pmatrix} \nu_1 \\ \nu_2 \\ \nu_3 \end{pmatrix}$ Mass states

PMNS mixing matrix

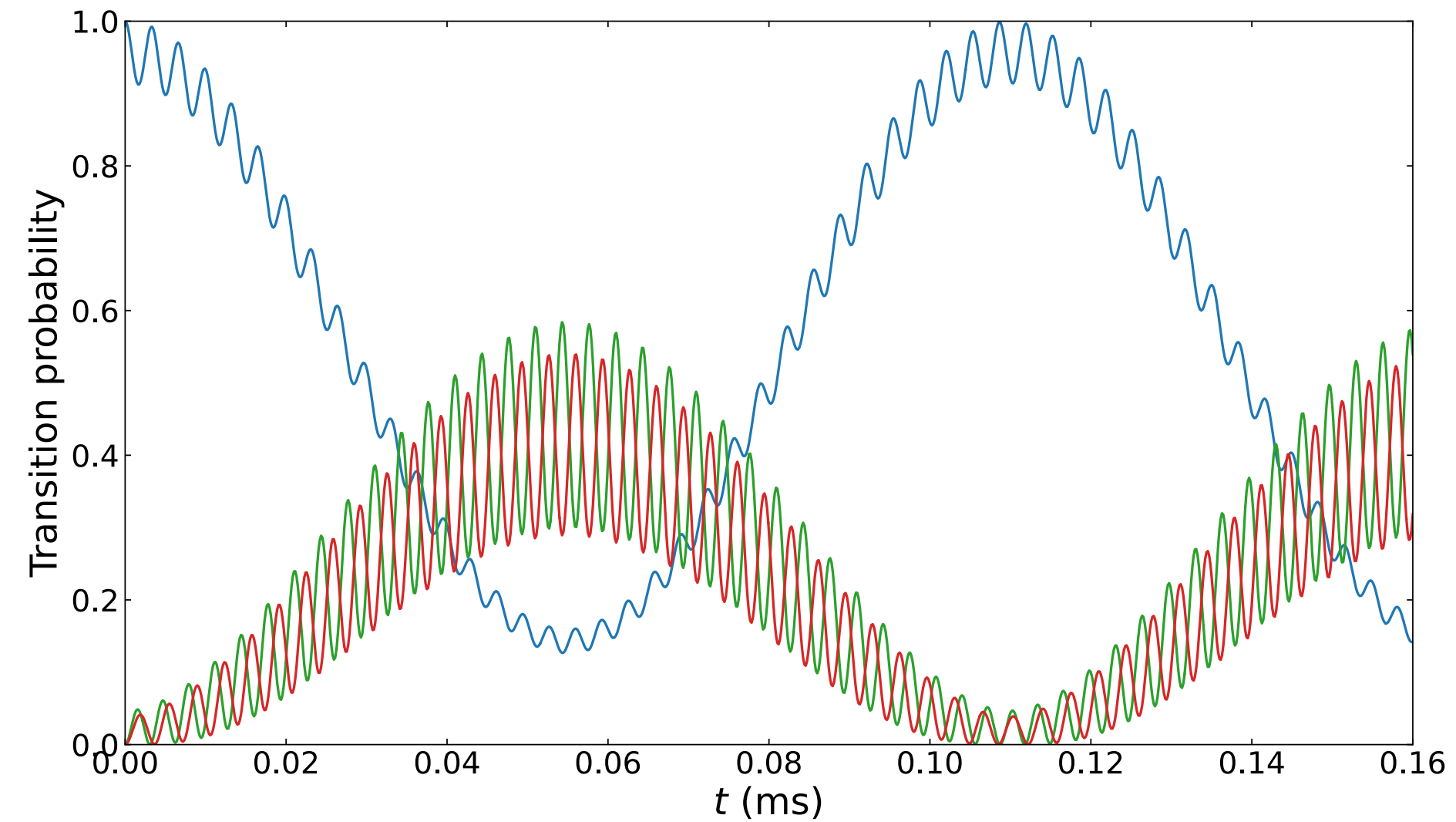
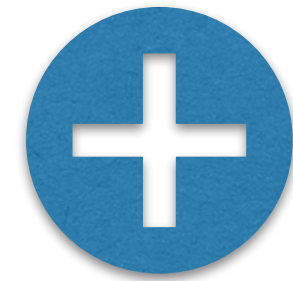
⇒ neutrino oscillations



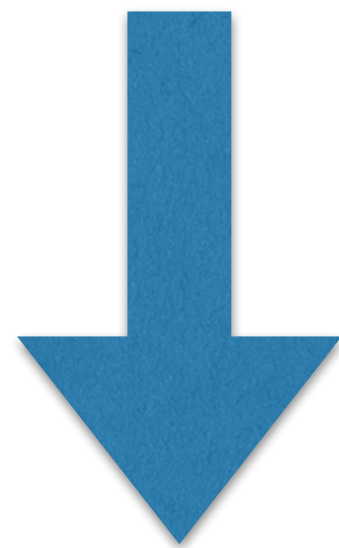
Neutrino quantum kinetic theory?

Boltzmann equation

$$\begin{aligned}\frac{\partial f_{\nu_e}}{\partial t} + \mathbf{v} \cdot \frac{\partial f_{\nu_e}}{\partial \mathbf{x}} + \dot{\mathbf{p}} \cdot \frac{\partial f_{\nu_e}}{\partial \mathbf{p}} &= \mathcal{C}_{\nu_e} \\ \frac{\partial f_{\nu_\mu}}{\partial t} + \mathbf{v} \cdot \frac{\partial f_{\nu_\mu}}{\partial \mathbf{x}} + \dot{\mathbf{p}} \cdot \frac{\partial f_{\nu_\mu}}{\partial \mathbf{p}} &= \mathcal{C}_{\nu_\mu} \\ \frac{\partial f_{\nu_\tau}}{\partial t} + \mathbf{v} \cdot \frac{\partial f_{\nu_\tau}}{\partial \mathbf{x}} + \dot{\mathbf{p}} \cdot \frac{\partial f_{\nu_\tau}}{\partial \mathbf{p}} &= \mathcal{C}_{\nu_\tau}\end{aligned}$$



**Superposition
of flavor
states**



**Quantum Kinetic
Equation**

G. Sigl and G. Raffelt, *Nucl. Phys. B* **406** (1993)
T. Stirner, G. Sigl and G. Raffelt, *JCAP* **05** (2018)
D. Fiorillo, G. Raffelt and G. Sigl, *PRL* **133** (2024)

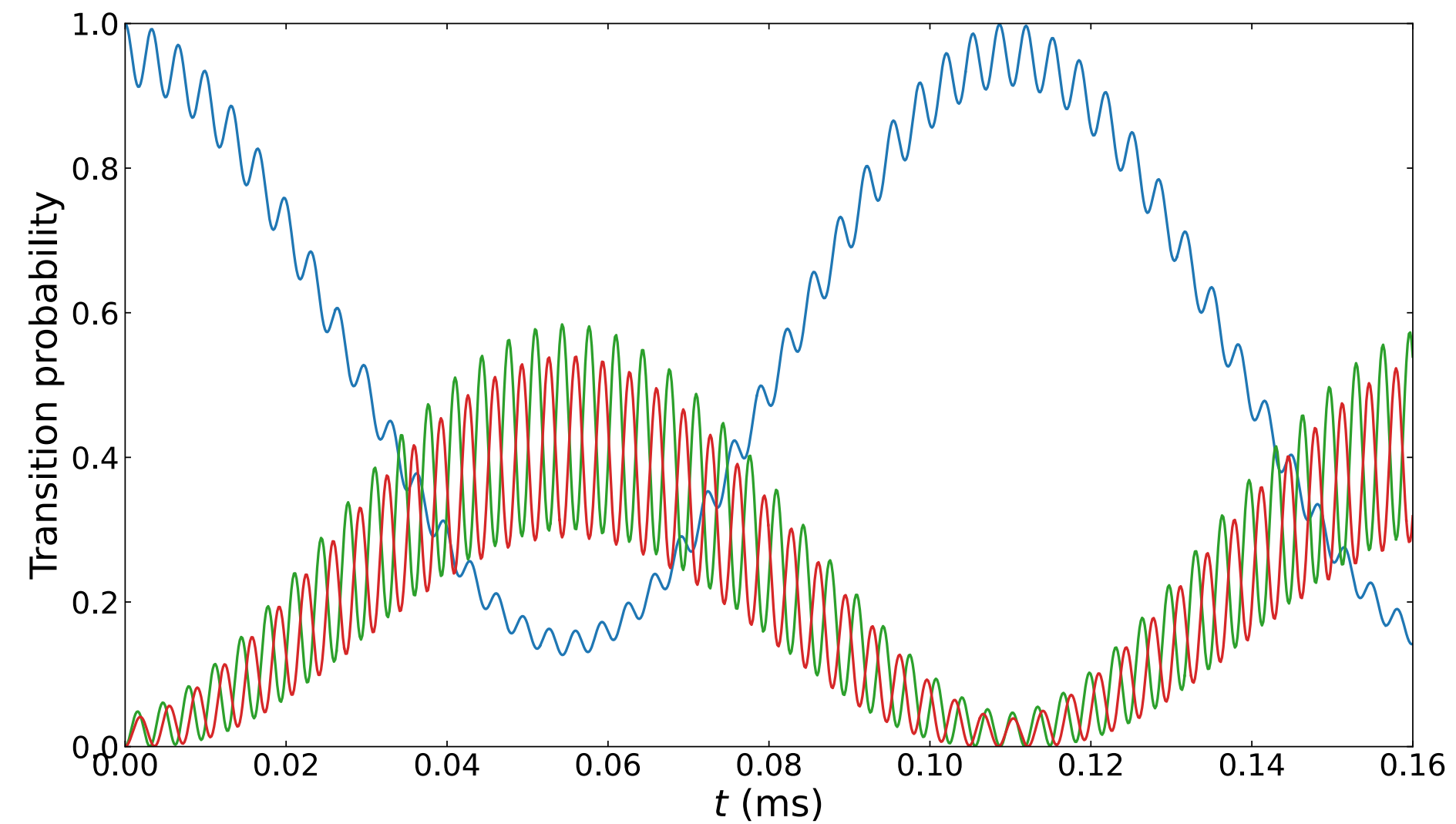
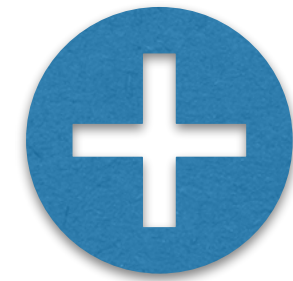
A. Vlasenko, G. Fuller and V. Cirigliano, *PRD* **89** (2014)
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BBGKY approach

**Quantum Kinetic
Equation**

Blackboard:

- * Hamiltonian (second quantization)
- * One-body, two-body... density matrices
- * Ehrenfest equation and BBGKY hierarchy
- * Mean-field approximation and beyond: separation of scales and collision term

Example: $\nu - \nu$ interactions

Neutrino-neutrino interaction Hamiltonian

$$H_{\nu\nu} = \frac{G_F}{4\sqrt{2}} \sum_{\alpha,\beta} \int d^3\mathbf{x} [\bar{\psi}_{\nu_\alpha} \gamma_\mu (1 - \gamma_5) \psi_{\nu_\alpha}] [\bar{\psi}_{\nu_\beta} \gamma^\mu (1 - \gamma_5) \psi_{\nu_\beta}]$$

$$\psi_{\nu_\alpha}(\mathbf{x}) = \sum_h \int \frac{d^3\mathbf{p}}{(2\pi)^3 2E_p} \left[a_{\nu_\alpha}(\mathbf{p}, h) u_{\mathbf{p}}^{(h)} e^{i\mathbf{p}\cdot\mathbf{x}} + b_{\nu_\alpha}^\dagger(\mathbf{p}, h) v_{\mathbf{p}}^{(h)} e^{-i\mathbf{p}\cdot\mathbf{x}} \right]$$

Example: $\nu - \nu$ interactions

Neutrino-neutrino interaction Hamiltonian

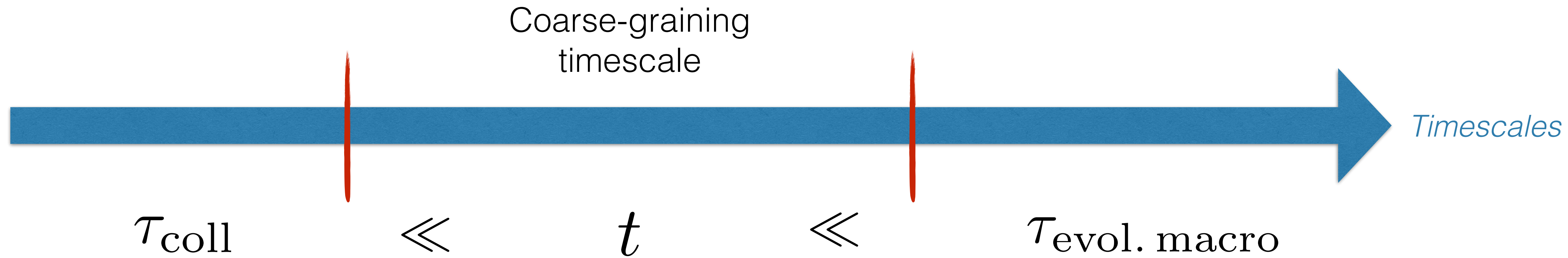
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One obtains the interaction coefficient:

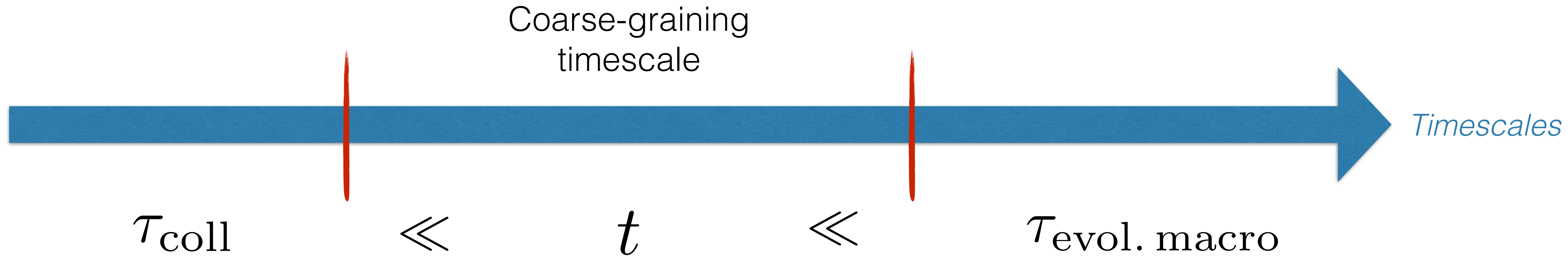
$$v_{\nu_\alpha(3)\nu_\beta(4)}^{\nu_\alpha(1)\nu_\beta(2)} = \frac{G_F}{2\sqrt{2}} (2\pi)^3 \delta^{(3)}(\mathbf{p}_1 + \mathbf{p}_2 - \mathbf{p}_3 - \mathbf{p}_4) \left[\bar{u}_{\mathbf{p}_1}^{(h_1)} \gamma_\mu (1 - \gamma_5) u_{\mathbf{p}_3}^{(h_3)} \right] \left[\bar{u}_{\mathbf{p}_2}^{(h_2)} \gamma^\mu (1 - \gamma_5) u_{\mathbf{p}_4}^{(h_4)} \right]$$

A problem of many scales



A problem of many scales

Typical energy ~ temperature ~ 10 MeV

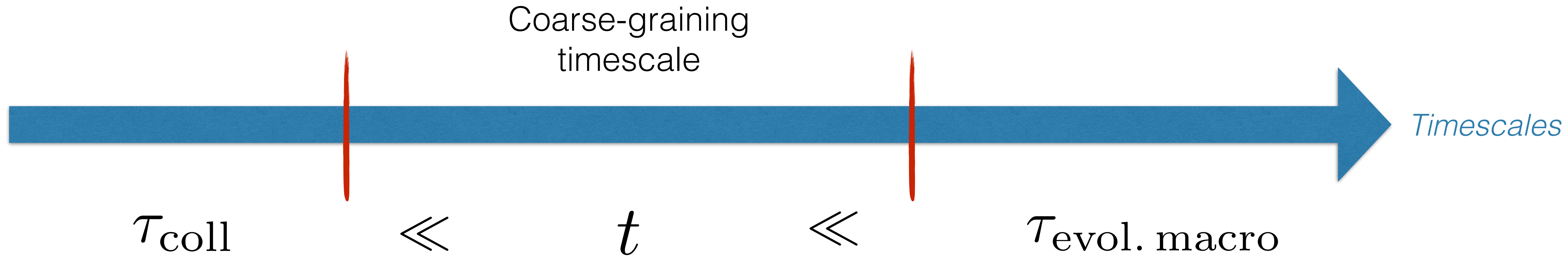


Interaction length ~ de Broglie wavelength

$$\lambda_{\text{dB}} \sim \frac{1}{10 \text{ MeV}} \sim 10^{-13} \text{ m} \longleftrightarrow \tau_{\text{coll}} \sim 10^{-21} \text{ s}$$

A problem of many scales

Typical energy \sim temperature ~ 10 MeV



Interaction length \sim de Broglie wavelength

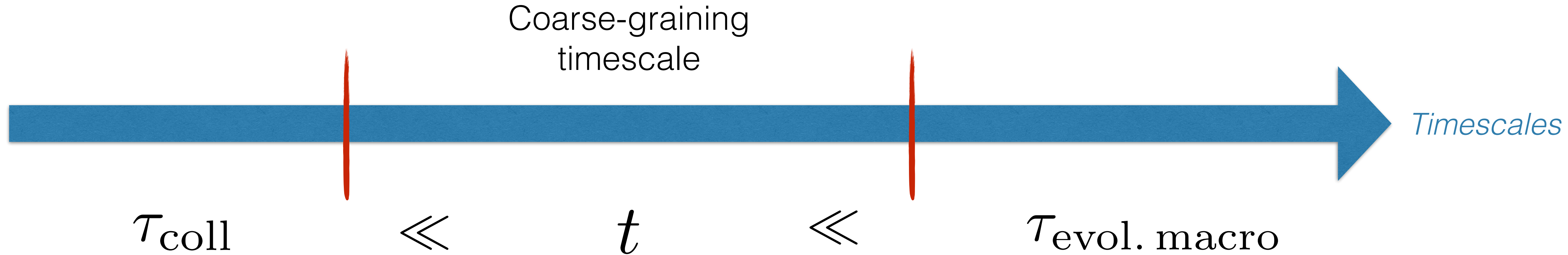
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Duration between collisions = Mean-free path

$$\tau_{\text{mfp}} \sim (G_F^2 T^5)^{-1} \sim 10^{-4} \text{ s}$$

A problem of many scales

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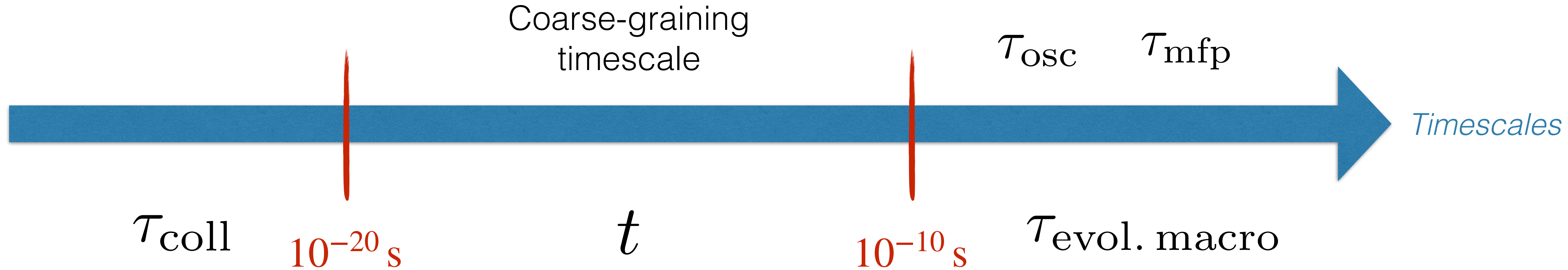
Oscillation timescales

$$\tau_{\text{osc, vac}} \sim \left(\frac{\Delta m^2}{2E} \right)^{-1} \sim \left(\frac{2 \times 10^{-3} \text{ eV}^2}{2 \times 10 \text{ MeV}} \right)^{-1} \sim 10^{-5} \text{ s}$$

$$\tau_{\text{osc, self}} \sim \left(\sqrt{2} G_F n_\nu \right)^{-1} \sim 10^{-10} \text{ s}$$

A problem of many scales

Typical energy ~ temperature ~ 10 MeV



Interaction length ~ de Broglie wavelength

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Homogeneous Quantum Kinetic Equation

$$i \frac{d\rho_j^i}{dt} = [t + \Gamma, \rho]_j^i + i C_j^i$$

$$C_{i'_1}^{i_1} \propto \frac{1}{4} \left(\tilde{v}_{i_3 i_4}^{i_1 i_2} \rho_{j_3}^{i_3} \rho_{j_4}^{i_4} \tilde{v}_{j_1 j_2}^{j_3 j_4} (\hat{1} - \rho)_{i'_1}^{j_1} (\hat{1} - \rho)_{i_2}^{j_2} - \tilde{v}_{i_3 i_4}^{i_1 i_2} (\hat{1} - \rho)_{j_3}^{i_3} (\hat{1} - \rho)_{j_4}^{i_4} \tilde{v}_{j_1 j_2}^{j_3 j_4} \rho_{i'_1}^{j_1} \rho_{i_2}^{j_2} \right. \\ \left. + (\hat{1} - \rho)_{j_1}^{i_1} (\hat{1} - \rho)_{j_2}^{i_2} \tilde{v}_{j_3 j_4}^{j_1 j_2} \rho_{i_3}^{j_3} \rho_{i_4}^{j_4} \tilde{v}_{i'_1 i_2}^{i_3 i_4} - \rho_{j_1}^{i_1} \rho_{j_2}^{i_2} \tilde{v}_{j_3 j_4}^{j_1 j_2} (\hat{1} - \rho)_{i_3}^{j_3} (\hat{1} - \rho)_{i_4}^{j_4} \tilde{v}_{i'_1 i_2}^{i_3 i_4} \right)$$

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Gain

Loss

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$\langle |\mathcal{M}|^2 \rangle$

Homogeneous ensemble of ultrarelativistic neutrinos:

$$\rho_{\nu\beta}^{\nu\alpha}(\mathbf{p}, h) = (2\pi)^3 2E_{\mathbf{p}} \delta^{(3)}(\mathbf{p} - \mathbf{p}') \delta_{h-h'} \rho_{\alpha\beta}(\mathbf{p})$$

Homogeneous Quantum Kinetic Equation

$$i \frac{d\rho}{dt} = [\mathcal{H}_{\text{vac}} + \mathcal{H}_{\text{mat}} + \mathcal{H}_{\text{self}}, \rho] + i\mathcal{C}$$

$$\mathcal{H}_{\text{vac}} = U \frac{M^2}{2|\mathbf{p}|} U^\dagger$$

Vacuum term

$$\mathcal{H}_{\text{mat}} = \sqrt{2}G_F (n_{e^-} - n_{e^+}) \begin{pmatrix} 1 & 0 & 0 \\ 0 & 0 & 0 \\ 0 & 0 & 0 \end{pmatrix}$$

Matter mean-field potential

$$\mathcal{H}_{\text{self}} = \sqrt{2}G_F \int \frac{d^3\mathbf{q}}{(2\pi)^3} (1 - \hat{p} \cdot \hat{q}) [\rho(\mathbf{q}) - \bar{\rho}(\mathbf{q})]$$

Self-interaction mean-field potential

Homogeneous Quantum Kinetic Equation

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Self-interaction mean-field potential

Rich phenomenology of flavor conversion mechanisms, including:

Vacuum oscillations

MSW resonance

Matter-neutrino resonance

Slow instabilities

Fast instabilities

Weak inhomogeneities and advection term



Homogeneous case $\rho_{ij}(\mathbf{p}) = \langle a_j^\dagger(\mathbf{p}) a_i(\mathbf{p}) \rangle$

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Wigner transform



$$\rho_{ij}(\mathbf{x}, \mathbf{p}) = \langle \mathcal{D}_{ij}(\mathbf{x}, \mathbf{p}) \rangle = \int \frac{d^3 \Delta}{(2\pi)^3} e^{i\Delta \cdot \mathbf{x}} \left\langle a_j^\dagger \left(\mathbf{p} - \frac{\Delta}{2} \right) a_i \left(\mathbf{p} + \frac{\Delta}{2} \right) \right\rangle$$

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Weak inhomogeneities: $\ell \gg \lambda_{\text{dB}} \sim |\mathbf{p}|^{-1} \sim 10^{-13} \text{ m}$

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lengthscale of variation of ρ

$$|\Delta| \sim \ell^{-1} \ll |\mathbf{p}|$$

Weak inhomogeneities and advection term



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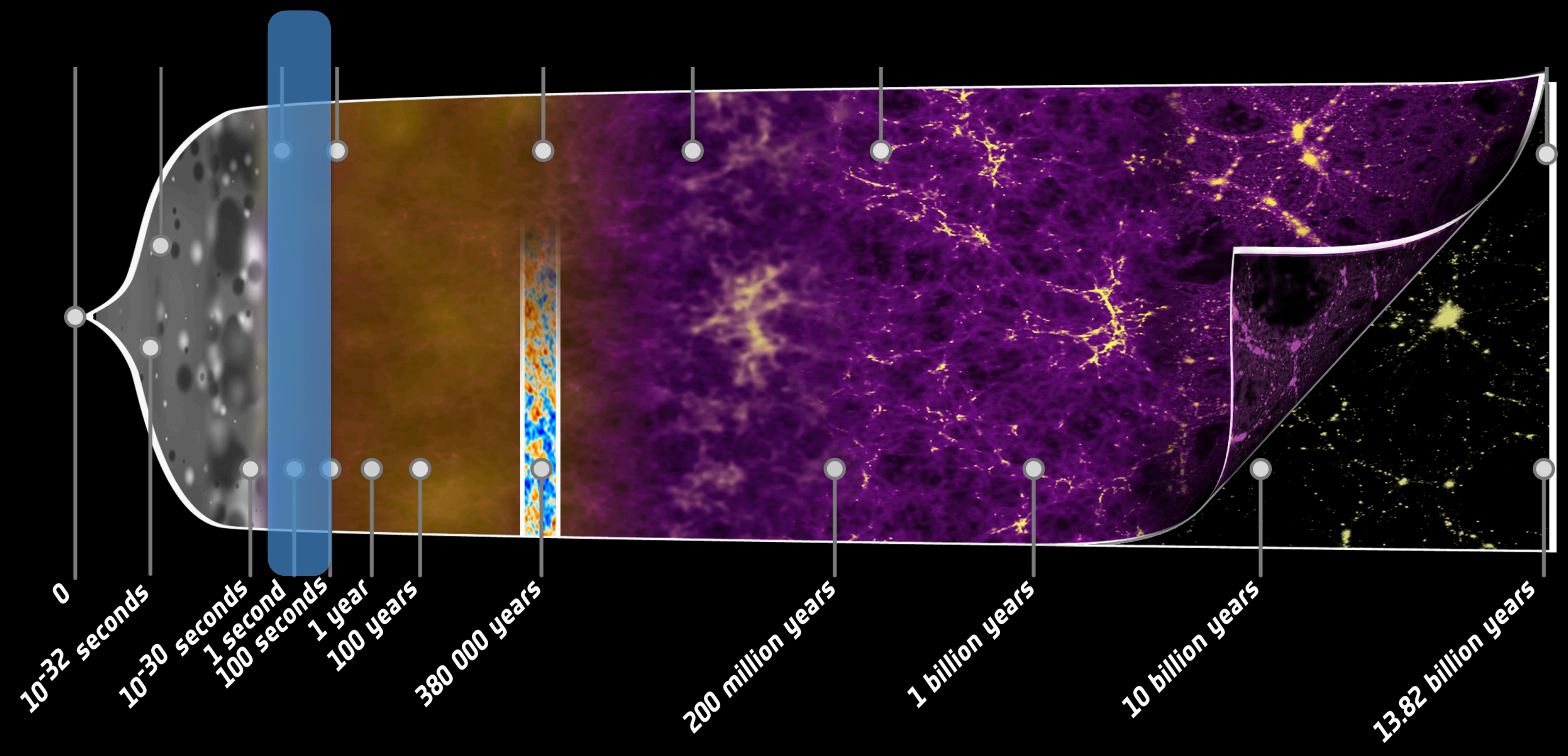
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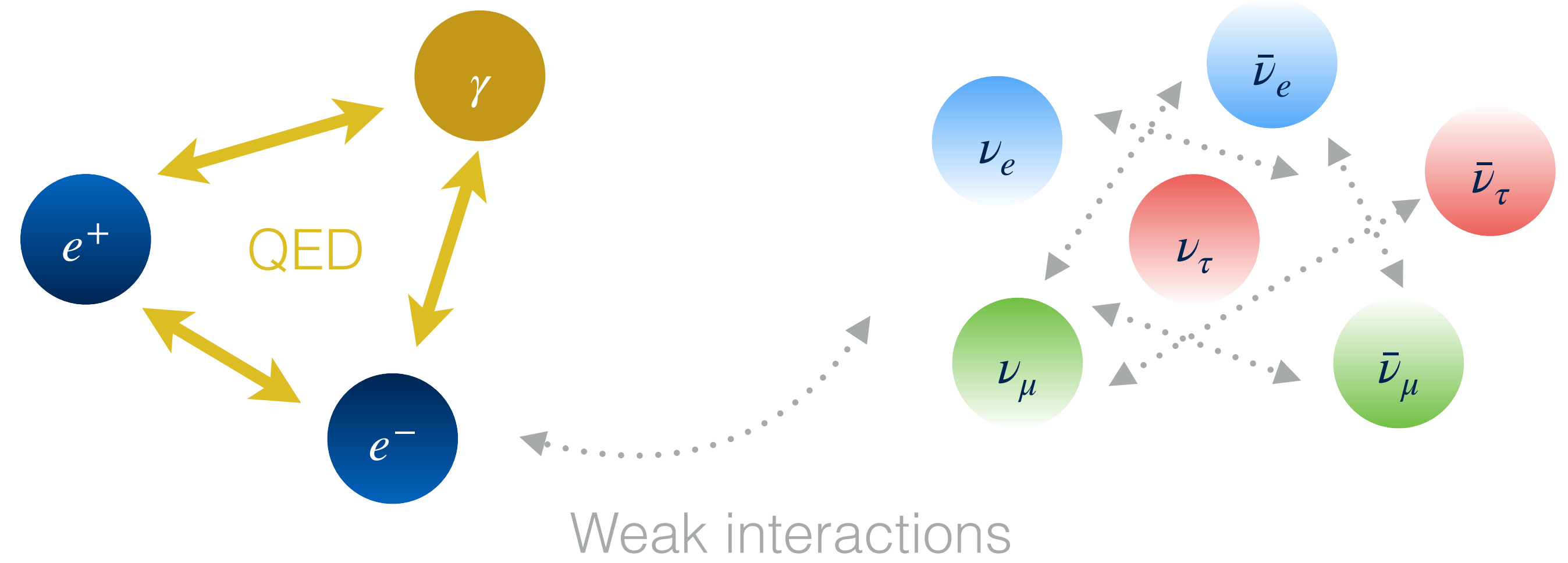
$$\longrightarrow \langle [\mathcal{D}_{ij}(\mathbf{x}, \mathbf{p}), H_0] \rangle \simeq -i \mathbf{v} \cdot \partial_{\mathbf{x}} \rho_{ij}(\mathbf{x}, \mathbf{p})$$

$$k_B T \sim 1 \text{ MeV} \sim 10^{10} \text{ K}$$

'MeV age'



Neutrinos in the MeV age

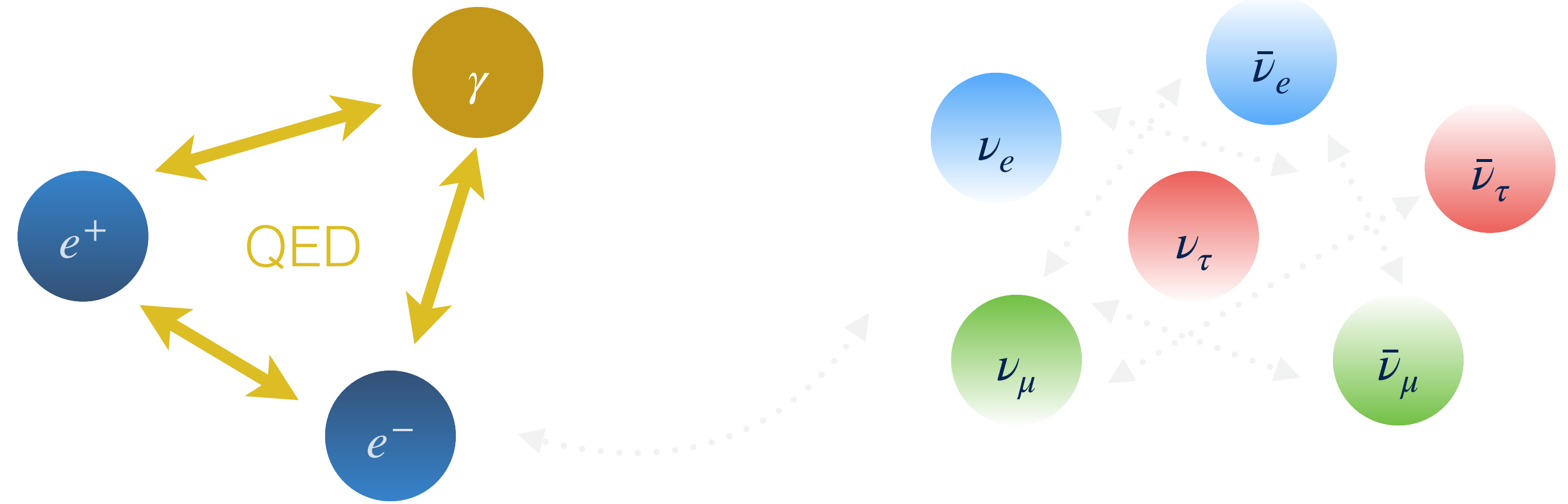


Neutrinos in the MeV age

NEUTRINO DECOUPLING

$$\frac{\text{collision rate}}{\text{expansion rate}} \quad \frac{\Gamma}{H} \simeq \frac{G_F^2 T^5}{\sqrt{g_*} T^2 / M_{\text{Pl}}}$$

$$\Gamma/H = 1 \iff T_{\text{dec}} \simeq 1 \text{ MeV}$$

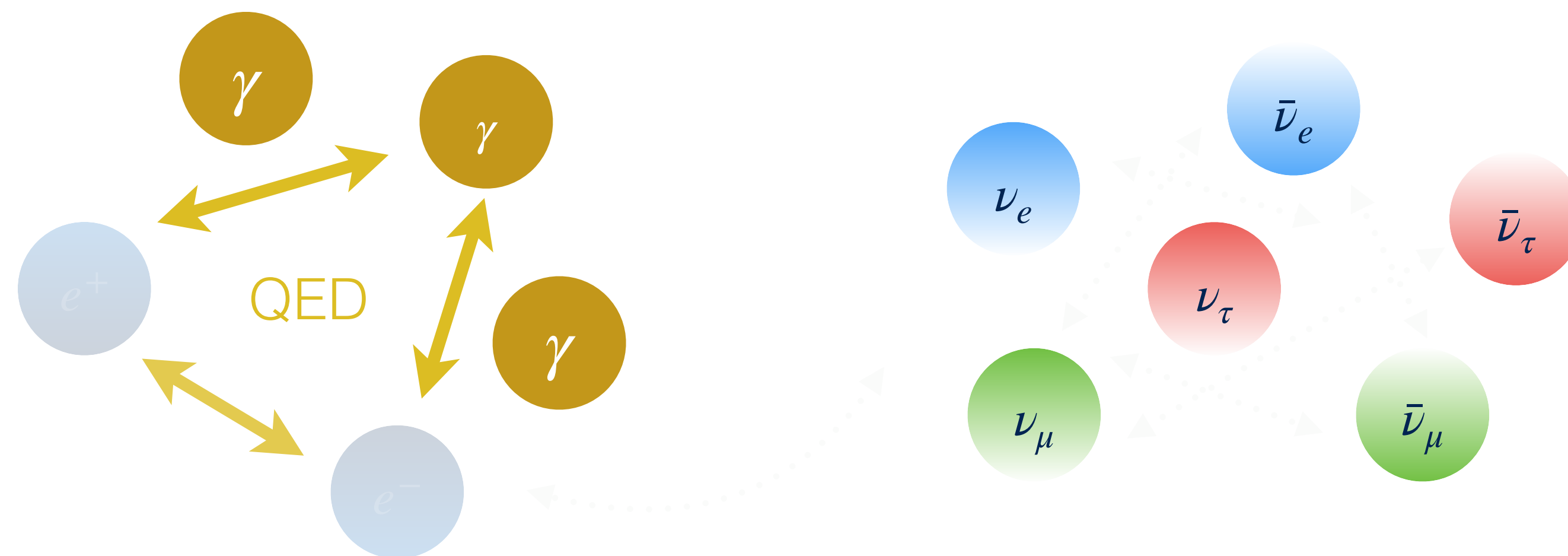


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ELECTRON-POSITRON ANNIHILATIONS

→ entropy release

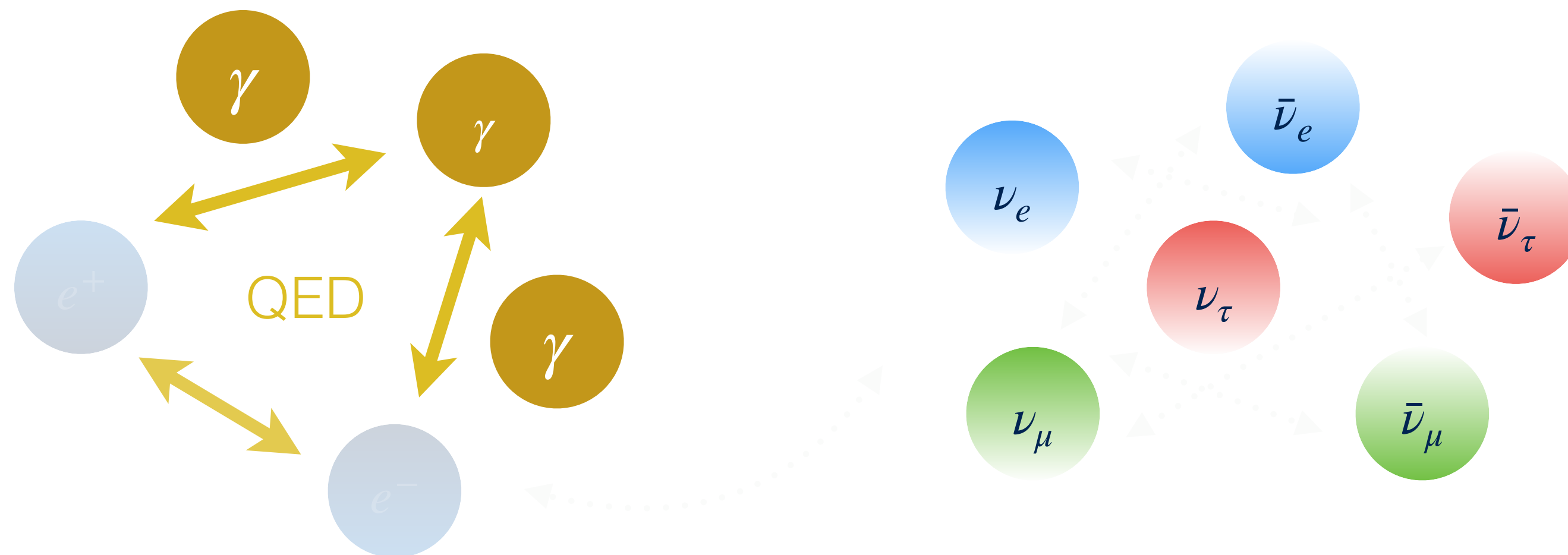
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NEUTRINO OSCILLATIONS

$$\Omega_{\text{vac}} = \frac{\Delta m^2}{2E}$$

$$\Omega_{\text{matt}} = \frac{\sqrt{2}G_F}{m_W^2} E [\rho_{e^\pm} + P_{e^\pm}] = \frac{7\pi^2}{45} \frac{\sqrt{2}G_F}{m_W^2} E T^4$$

MSW transition for $\Omega_{\text{vac}} = \Omega_{\text{matt}}$

$$T_{\text{MSW}} \in [12 \text{ MeV}, 2.8 \text{ MeV}]$$

ELECTRON-POSITRON ANNIHILATIONS

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Neutrino transport in the early Universe

$$\frac{\partial \varrho}{\partial t} - Hp \frac{\partial \varrho}{\partial p} = -i [\mathcal{H}_{\text{vac}} + \mathcal{H}_{\text{mat}} + \mathcal{H}_{\nu\nu, \varrho}] + \mathcal{C}$$

- Calculation of **standard neutrino decoupling**: $N_{\text{eff}} = 3.044$

$$\rho_{\text{tot}} = \rho_{\gamma} + \rho_{\nu, \bar{\nu}} = \left[1 + N_{\text{eff}} \times \frac{7}{8} \left(\frac{4}{11} \right)^{4/3} \right] \rho_{\gamma}$$



JF, C. Pitrou, M.C. Volpe [[2008.01074](#)]

J. Bennett *et al.* [[2012.02726](#)]

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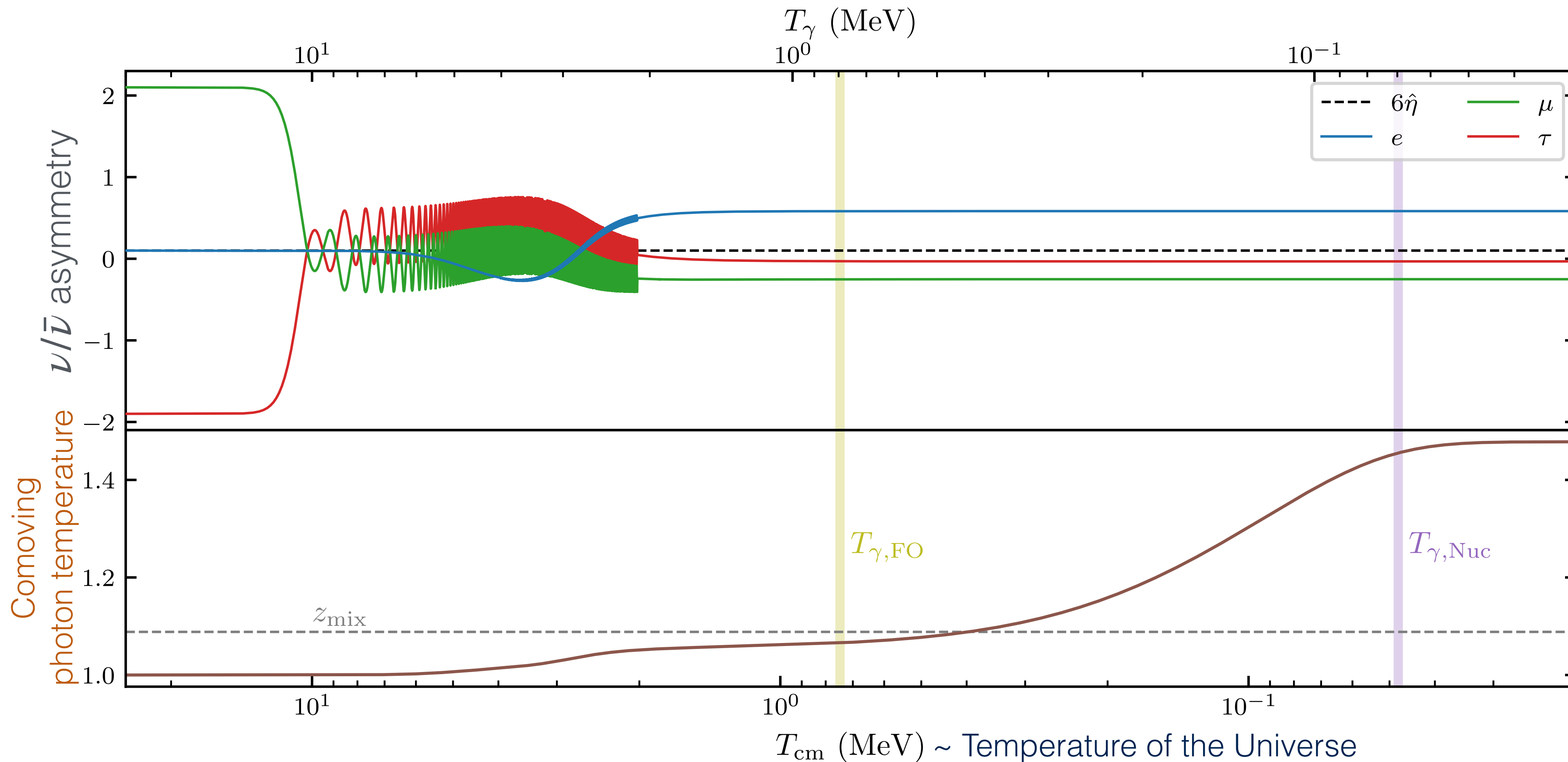


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$$\eta_{\alpha} \equiv \frac{n_{\nu_{\alpha}} - n_{\bar{\nu}_{\alpha}}}{T^3}$$

Initial values

$$\eta_{\text{av}} = \eta_e = 0.1$$

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JF, C. Pitrou [[2110.11889](#)]

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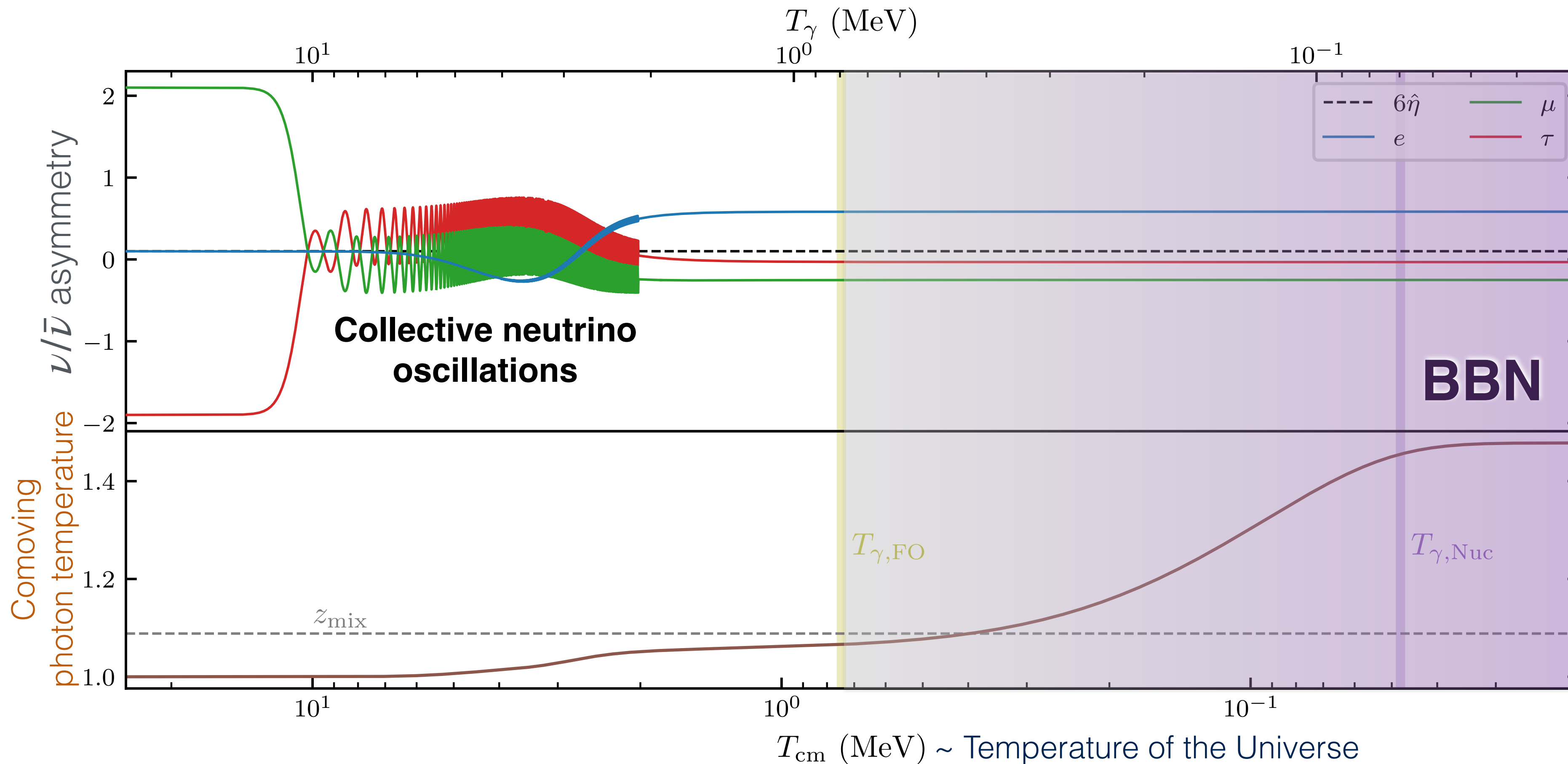


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Summary and prospects

- The **quantum kinetic approach** consists in describing the evolution **at the one-body level**.

Formally similar to OQS equations with system = 'one neutrino' and environment = 'the rest'

- It relies on a **separation of scales**:

- quasi-instantaneous scatterings compared to a 'long' macroscopic evolution $\tau_{\text{coll}} \ll \tau_{\text{macro evol.}}$
- Weak inhomogeneities, on large scales compared to the de Broglie wavelength $|\mathbf{p}|^{-1} \ll \ell$

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- Ongoing debate: **entanglement is neglected** in the QKE approach. *Is that always justified?*