

Solving Einstein equation using recursions

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@Loop-the-Loop

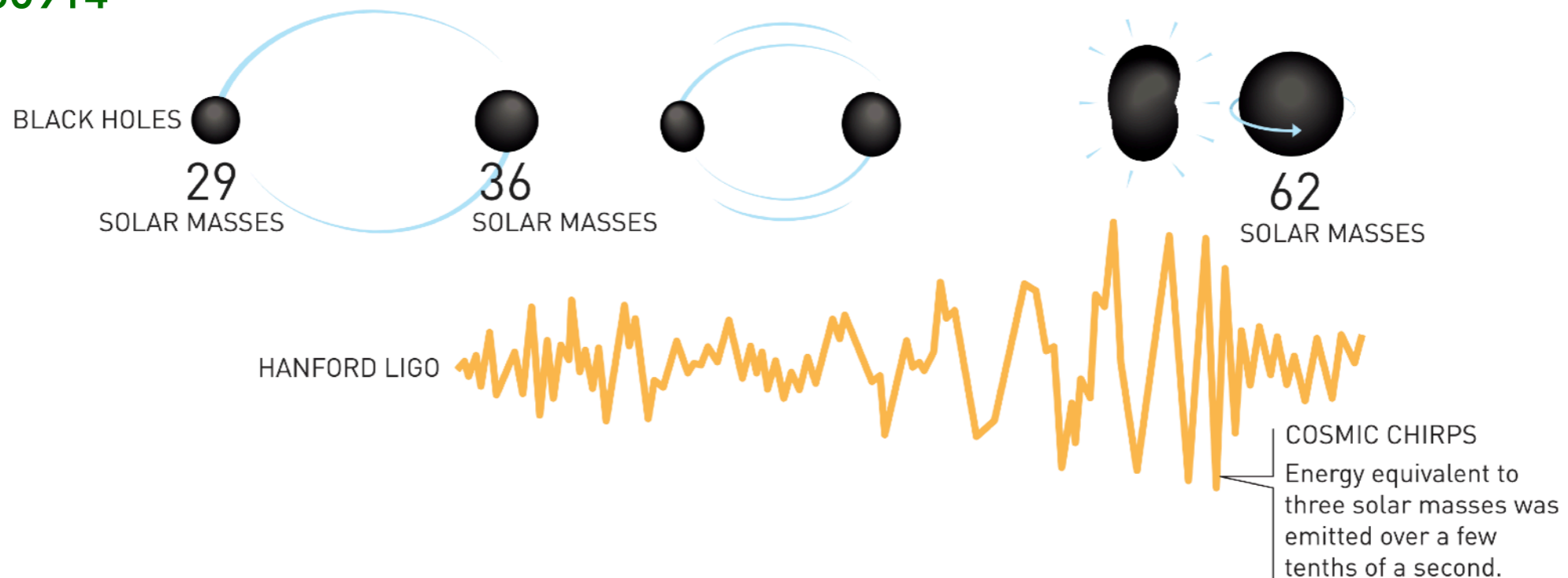
Based on

KL, Damgaard 2403.13216

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Gravitational wave from Binary BH mergers

GW150914



- **Gravitational wave: new window to probe our Universe**
- How do we describe this system? \implies **Solve Einstein equation** (perturbatively)
- *Are the theoretical tools we have powerful enough to solve this problem?*

Toy model — Schwarzschild BH solution

Solving perturbative Einstein Equation

1. **Solve Einstein Equation directly** (old and brute force approach)

- ▶ Green function method
- ▶ Perturbative GR is notorious for its complexity
- ▶ Leading order correction is the practical limit [Florides, Synge 61] [Westpfahl, 85]

2. **Scattering amplitude Approach** (since 2018)

- ▶ Modern techniques in **QFT/Quantum Gravity**

Generalized unitarity [Bern, Dixon, Dunbar, Kosower, hep-ph/9403226](#)

On-shell recursion [Britto, Cachazo, Feng, Witten, hep-th/0501052](#)

Color-kinematics duality and double copy [Bern, Carrasco, Johansson, 0805.3993, 1004.0476](#)

- ▶ Issues — convergence of the series, loop integrals, etc

3. **Go back to the Einstein equation again** (armed with new techniques)

[Damgaard, KL `24]

In this talk

- returning to the solving Einstein equation explicitly
- **Two main ideas**
 - **good variable** — By doubling the fields, the perturbative Einstein equation is drastically simplified. We can hide the ugly infinite expansion.
 - **off-shell recursion** — A new methodology for solving perturbative Einstein Equation Remarkably, all the “higher-loop integrals” are represented by iterations of one-loop **bubble integrals**.
- For the Schwarzschild BH case, we derived **all-order results** — first derivation!
 - **Efficiency** — fixed number of terms, recursions and simple loop integrals...
 - **Universality** — binary black holes & rotating black holes, branes etc
- Recently, the similar results are derived from the amplitude point of view

[Mougiakakos, Vanhove `24]

**Perturbative GR
and
doubling prescription**

Tensor density representation

➤ **Two sources** of the infinite expansion: g^{-1} and $\sqrt{-g}$

➤ **Field redefinition - tensor density** [Landau & Lifshitz book]:

$$\sigma^{\mu\nu} = \sqrt{-g} g^{\mu\nu}, \quad \sigma_{\mu\nu} = \frac{1}{\sqrt{-g}} g_{\mu\nu},$$

➤ Why? There is no $\sqrt{-\sigma}$. The number of σ^{-1} is always greater than σ due to derivatives.

➤ **EH action** (up to total derivative) in terms of the **tensor density**

$$S_{\text{EH}} = \int d^D x \left[\frac{1}{4} \sigma^{\mu\nu} \partial_\mu \sigma^{\rho\sigma} \partial_\nu \sigma_{\rho\sigma} - \frac{1}{2} \sigma^{\mu\nu} \partial_\mu \sigma^{\rho\sigma} \partial_\rho \sigma_{\nu\sigma} + (D-2) \sigma^{\mu\nu} \partial_\mu \hat{d} \partial_\nu \hat{d} \right], \quad \partial_\mu \hat{d} = -\frac{1}{4} \sigma^{\rho\sigma} \partial_\mu \sigma_{\rho\sigma}$$

➤ Substitute the metric perturbation [Cheung, Remmen 18], [Deser, 70], [Capper, Leibbrandt, Medrano, 73]

$$\sigma^{\mu\nu} = \eta^{\mu\nu} - \kappa h^{\mu\nu}, \quad \sigma_{\mu\nu} = \eta_{\mu\nu} + \sum_{n=1}^{\infty} \kappa^n (h^n)_{\mu\nu}.$$

➤ Provides **the simplest form** of the perturbative GR [Cho, Kim, Lee, 23]

- general n-th order terms of the EH action and Einstein eq.
- Three minimal building blocks

Doubling prescription

➤ **Idea:** do not substitute metric perturbations from the beginning!

➤ Let us treat the **metric** (σ) and the **inverse metric** (σ^{-1}) on **equal footing**

[Gomez, Lipinski Jusinskas, Lopez-Arcos, Quintero Velez '22]

➤ **Remove metric** (σ) and **introduce an auxiliary field** $\tilde{\sigma}$. **on-shell value of** $\tilde{\sigma}_{\mu\nu} = \sigma_{\mu\nu}$

➤ Impose a **constraint**:

$$\tilde{\sigma}_{\mu\nu}\sigma^{\nu\rho} = \delta_{\mu}^{\rho}$$

➤ perturbative expansions:

$$\sigma^{\mu\nu} = \eta^{\mu\nu} - h^{\mu\nu}, \quad \tilde{\sigma}_{\mu\nu} = \eta_{\mu\nu} + \tilde{h}_{\mu\nu}.$$

➤ Then \tilde{h} satisfies the constraint $\implies \tilde{h}^{\mu\nu} = h^{\mu\nu} + \tilde{h}^{\mu}_{\rho} h^{\rho\nu}$,

$$\tilde{h}_{(n)}^{\mu\nu} = h_{(n)}^{\mu\nu} + \sum_{m=1}^{n-1} \tilde{h}_{(n-m)}^{\mu\rho} h_{(m)}^{\rho\nu}$$

Source of the Schwarzschild BH

➤ Consider pure gravity with a matter

$$S = \int d^4x \left[\frac{1}{2\kappa^2} \sqrt{-g} R + \frac{1}{2} j_{\mu\nu}(x) g^{\mu\nu}(x) \right]$$

where $j_{\mu\nu}(x)$ is an external source (density) without metric dependence,

➤ Relation to the energy-momentum tensor $T_{\mu\nu}$

$$\sqrt{-g} T_{\mu\nu} = j_{\mu\nu}$$

➤ Schwarzschild BH is not a vacuum solution — point mass source

➤ **Energy-momentum tensor** for a point mass traveling on a worldline $x^\mu(\tau)$

$$T^{\mu\nu}(y^\sigma) = 8\pi GM \int \left[\frac{\delta^{(4)}(y^\sigma - x^\sigma(\tau))}{\sqrt{-g}} \right] \frac{dx^\mu}{d\tau} \frac{dx^\nu}{d\tau} d\tau$$

➤ Source of Schwarzschild BH — a **static point mass** placed at the origin, $\mathbf{x} = 0$

$$j_{\mu\nu}(x) = 8\pi GM v_\mu v_\nu \delta^3(\mathbf{x}), \quad v^\mu = \frac{dx^\mu}{d\tau} = (-1, 0, 0, 0).$$

Field equation

➤ Einstein tensor (density)

$$\begin{aligned} \mathcal{G}^{\mu\nu} = & \frac{1}{2} \sigma^{\rho\sigma} \left[\partial_\rho \partial_\sigma \sigma^{\mu\nu} + \partial_\rho \sigma^{\kappa\mu} \partial_\sigma \tilde{\sigma}_{\kappa\lambda} \sigma^{\nu\lambda} \right] - \sigma^{\rho(\mu} \left[\partial_\rho \partial_\sigma \sigma^{\nu)\sigma} + \partial_\rho \sigma^{|\kappa\lambda} \partial_\kappa \tilde{\sigma}_{\lambda\sigma} \sigma^{\sigma|\nu)} \right] \\ & + \sigma^{\mu\kappa} \sigma^{\nu\lambda} \left[\frac{1}{4} \partial_\kappa \sigma^{\rho\sigma} \partial_\lambda \tilde{\sigma}_{\rho\sigma} + (D-2) \partial_\kappa \hat{d} \partial_\lambda \hat{d} \right] \\ & + \frac{1}{2} \left[\partial_\rho \sigma^{\rho\sigma} \partial_\sigma \sigma^{\mu\nu} - \partial_\sigma \sigma^{\rho\mu} \partial_\rho \sigma^{\sigma\nu} \right] + \sigma^{\mu\nu} \left[\partial_\kappa \left(\sigma^{\kappa\lambda} \partial_\lambda \hat{d} \right) \right], \end{aligned}$$

➤ Einstein equation

$$\sqrt{-\sigma} \mathcal{G}^{\mu\nu} = j^{\mu\nu}$$

$$\mathcal{G}^{\mu\nu} = \sum_{n=1}^{\infty} G^n \mathcal{G}_{(n)}^{\mu\nu}, \quad j^{\mu\nu}(x) = G j_{(1)}^{\mu\nu}(x)$$

➤ $j^{\mu\nu}$ contributes to the G^1 -order only

$$\mathcal{G}_{(1)}^{\mu\nu} = -\frac{1}{2} \square h_{(1)}^{\mu\nu} = j^{\mu\nu}$$

$$\mathcal{G}_{(n)}^{\mu\nu} = 0, \quad n > 1$$

Harmonic vs de Donder gauge

- One of the most straightforward gauge choices is the **harmonic** or **de Donder gauge**

$$g^{\mu\nu}\Gamma_{\mu\nu}^{\rho} = 0 \quad \text{or} \quad \partial_{\mu}h^{\mu\nu} - \frac{1}{2}\eta^{\mu\nu}\partial_{\mu}h^{\rho}_{\rho} = 0 \quad \text{for } g_{\mu\nu} = \eta_{\mu\nu} + h_{\mu\nu}$$

harmonic gauge **de Donder gauge**

- **Linearized harmonic gauge = de Donder gauge**, but not in higher orders

- However, in the tensor density perturbations, these are equivalent

$$g^{\mu\nu}\Gamma_{\mu\nu}^{\rho} = \partial_{\mu}(\sqrt{-g}g^{\mu\rho}) = \partial_{\mu}\sigma^{\mu\rho} = \partial_{\mu}h^{\mu\rho} = 0$$

- In our perturbation convention,

harmonic gauge = de Donder gauge

- If we obtained a solution using amplitude, in what coordinates do we get the result?

\iff **gauge choice**

- **However, it is not obvious in actual computation...**

Schwarzschild metric in harmonic coordinates

➤ The usual form of the Schwarzschild metric

$$ds^2 = - \left(1 - \frac{2GM}{r} \right) dt^2 + \left(1 - \frac{2GM}{r} \right)^{-1} dr^2 + r^2 d\Omega^2$$

➤ In the **harmonic coordinates**, the metric

$$ds^2 = - \frac{r - GM}{r + GM} dt^2 + \frac{r + GM}{r - GM} dr^2 + (r + GM)^2 d\Omega, \quad \text{obtained by } r \rightarrow r + GM$$

➤ The **tensor density** $\sigma^{\mu\nu}$ for this metric ($\sigma^{\mu\nu} = \sqrt{-g} g^{\mu\nu}$),

$$\sigma^{\mu\nu} \partial_\mu \partial_\nu = - \frac{(r + GM)^3}{r^2 (r - GM)} \partial_t^2 + \left(\delta^{ij} - \frac{G^2 M^2 x^i x^j}{r^4} \right) \partial_i \partial_j.$$

➤ The corresponding metric perturbations $h^{\mu\nu}$

$$h^{00} = -1 + \frac{(r + GM)^3}{r^2 (r - GM)} = \frac{4GM}{r} + \frac{7G^2 M^2}{r^2} + \frac{8G^3 M^3}{r^3} + \frac{8G^4 M^4}{r^4} + \dots,$$
$$h^{ij} = \frac{G^2 M^2 x^i x^j}{r^4}.$$

Coefficients of h^{00} is fixed by "8" while h^{ij} truncates at the second order

Remained ambiguity [Fromholz, Poisson, Will '13 The Schwarzschild metric: It's the coordinates, stupid!]

- Even after the harmonic gauge, the form of the metric is not fixed yet
- Most general solution in harmonic coordinate — a new parameter C

$$\sigma^{00} = -1 - \frac{4M}{r} - \frac{7M^2}{r^2} - \frac{8M^3}{r^3} - \frac{8M^4 - 2CM/3}{r^4} + \mathcal{O}(r^{-5})$$

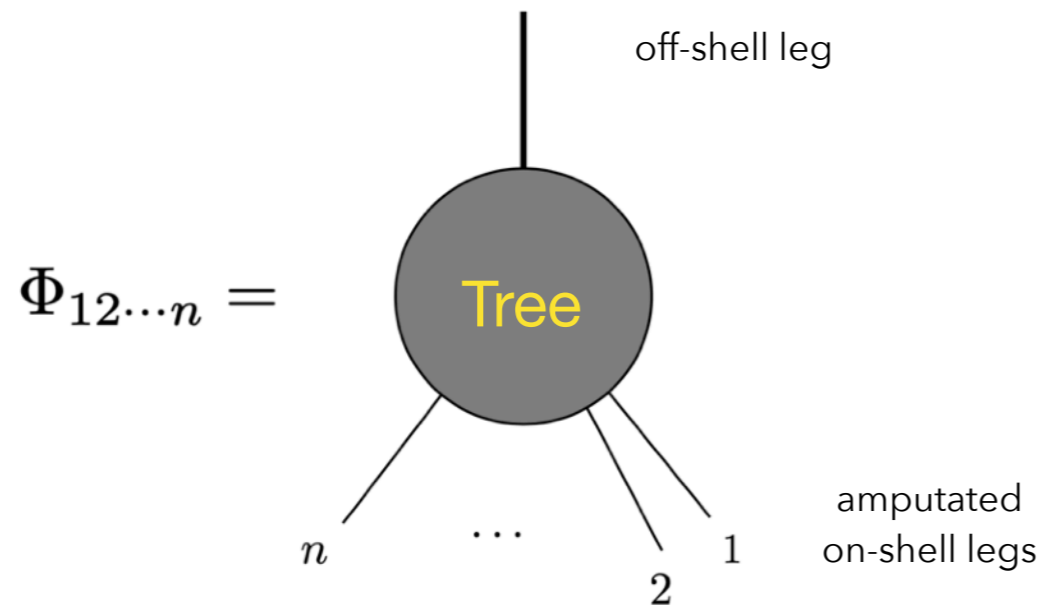
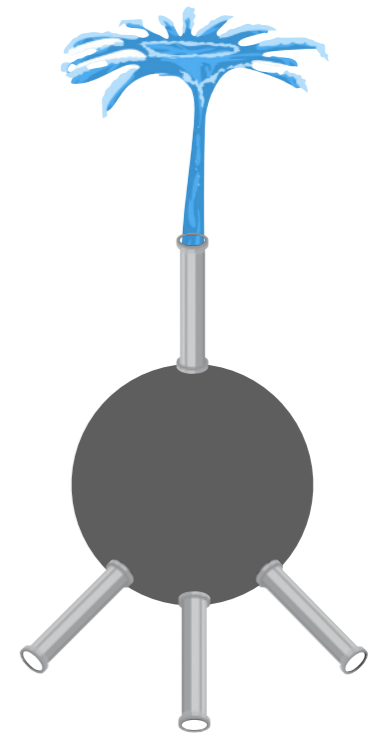
$$\sigma^{ij} = \left(1 - \frac{C}{3r^3} - \frac{2CM^2}{5r^5} + \mathcal{O}(r^{-6}) \right) \delta^{ij} + \left(-\frac{G^2 M^2}{r^2} + \frac{C}{r^3} + \frac{2G^2 M^2 C}{3r^3} + \mathcal{O}(r^{-6}) \right) \frac{x^i x^j}{r^2}$$

- Solving the de Donder gauge, $\partial_\mu h^{\mu\nu} = 0$, admits an **integration constant C**
- If we turn off C , the solution returns to the previous metric expansion.
- The existence of the parameter has recently been observed in the differential equation.
- *How can we interpret this ambiguity in **the field theory context**?*

Recursion Relation for perturbative GR solutions

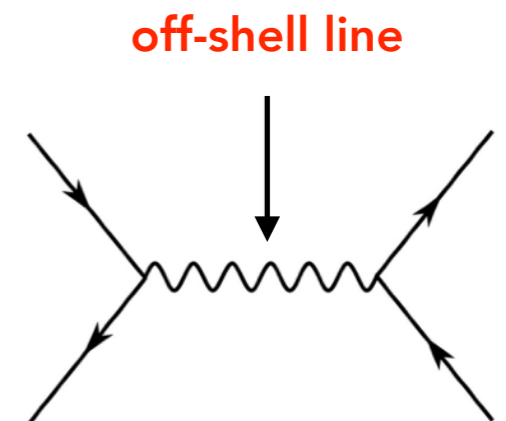
Off-shell Currents

- **Off-Shell recursions:** recursions for **off-shell currents** [Berends, Giele '87] for gluon amplitude at **tree-level**
- **Rank- n Off-shell currents:** sum of all $(n + 1)$ -point Feynman diagrams
- Diagrammatic representation



The off-shell line satisfy the conservation law $\partial_\mu J^\mu_{12\dots n} = 0$ without EoM — **Ward identity**

- Off-shell lines can be **glued** in a specific way (interaction vertices)
Intermediate states are off-shell

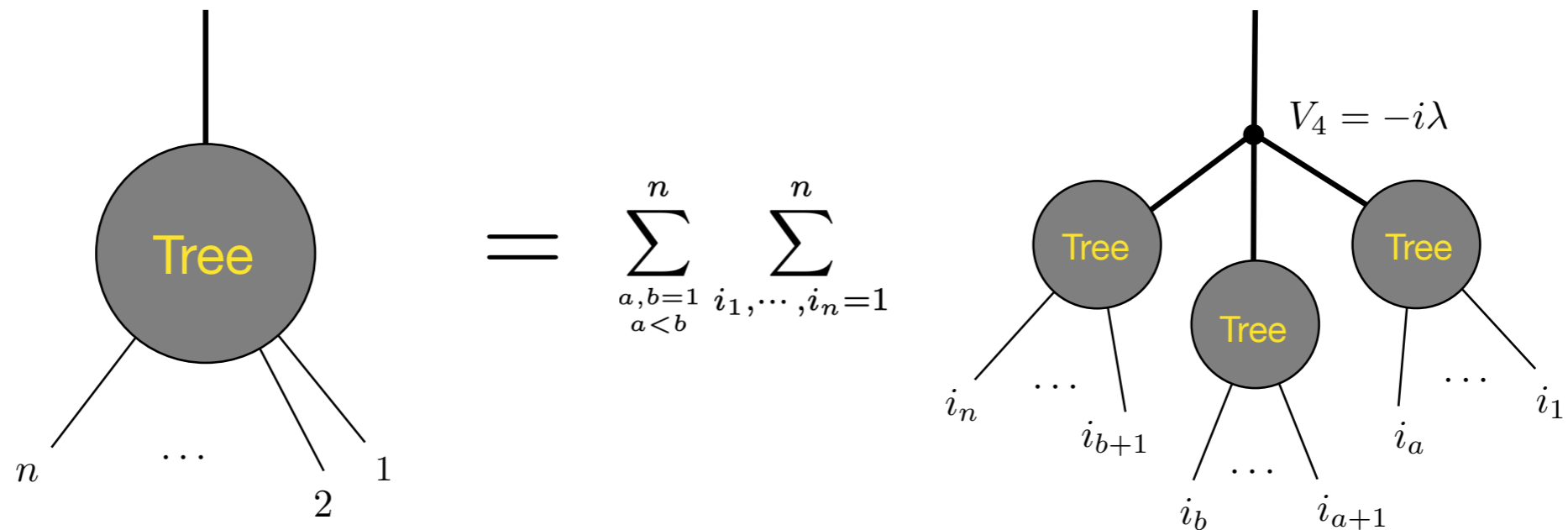


Off-shell Recursion [Berends, Giele '87]

➤ **Recursions: hidden self-similarity** — finite number of interaction vertices (**patterns**)

➤ Identifying the **Hierarchy** for off-shell currents: **# of on-shell legs**

➤ ϕ^4 theory:



➤ **Efficiency:**

- Do not treat individual diagrams
- **Recycling** calculations - never repeat the same calculations!

➤ **Gravity** — infinite number of vertices (No patterns)



Perturbative expansion [Rosly, Selivanov '96,'97], [KL '22]

- **Modern derivation:** substituting the **perturbative expansion** into the classical **EoM**
 \implies connects **solutions of EoM** and **tree-level amplitudes**
- **The classical field** in the quantum effective action formalism — 1-point function in the presence of the source $j^{\mu\nu}$

$$h^{\mu\nu}(x) = \sum_{n=1}^{\infty} \frac{1}{n!} \int_{y_1, y_2, \dots, y_n} \langle 0 | T[h_x^{\mu\nu} h_{y_1}^{\kappa_1 \lambda_1} \dots h_{y_n}^{\kappa_n \lambda_n}] | 0 \rangle_c \frac{i j_{y_1}^{\kappa_1 \lambda_1}}{\hbar} \dots \frac{i j_{y_n}^{\kappa_n \lambda_n}}{\hbar}.$$

- The field corresponds to a different physical quantity depending on the sources:

- **Inverse propagator:** $j_x^{\mu\nu} = \sum_{i=1}^N \int_{y_i} \mathbf{K}_{xy_i}^{\mu\nu, \rho\sigma} e^{-ik_i \cdot y_i} \implies$ scattering amplitude.
- **Plane-wave:** $j_x^{\mu\nu} = \sum_{i=1}^N \int_{y_i} e^{-ik_i \cdot y_i} \implies$ Correlation function.
- **Point-mass source:** $j_x^{\mu\nu} = M v^\mu v^\nu \int_{\mathcal{L}} e^{-i\ell \cdot x} \implies$ **solutions of EoM.** $v^\mu = \frac{dx^\mu}{d\tau} = (-1, 0, 0, 0).$

Perturbative expansion for classical solutions

➤ Substituting the external sources:
$$h^{\mu\nu}(x) = \sum_{n=1}^{\infty} \frac{1}{n!} \int_{\ell_1, \ell_2, \dots, \ell_n} J^{\mu\nu}_{\ell_1 \ell_2 \dots \ell_n} e^{-i\ell_{12\dots n} \cdot x},$$

➤ It is convenient to shift the loop momenta, $\ell_1 \rightarrow -\ell_{12\dots n}$

$$h^{\mu\nu}(x) = \sum_{n=1}^{\infty} \int_{\ell_1} e^{i\ell_1 \cdot x} \int_{\ell_2, \dots, \ell_n} \frac{1}{(n-1)!} J^{\mu\nu}_{-\ell_{12\dots n} \ell_2 \dots \ell_n} = \sum_{n=1}^{\infty} \int_{\ell_1} e^{i\ell_1 \cdot x} J^{\mu\nu}_{(n)|\ell_1},$$

$J^{\mu\nu}_{(n)|\ell_1} = \int_{\ell_2, \dots, \ell_n} \frac{1}{(n-1)!} J^{\mu\nu}_{-\ell_{12\dots n} \ell_2 \dots \ell_n}$

➤ Compare with the **amplitude perturbation** — **A continuous limit**
finite # of particles cannot generate the classical solutions

$$h^{\mu\nu} = \sum_{\mathcal{P}} J^{\mu\nu}_{\mathcal{P}} e^{-ik_{\mathcal{P}} \cdot x}$$

➤ We call the number of the loop momenta of an off-shell current as **rank**.

Here the rank is equivalent to the powers of coupling G

$$h^{\mu\nu} = \sum_{n=0}^{\infty} G^n h^{\mu\nu}_{(n)} \quad \text{and} \quad h^{\mu\nu}_{(n)} = \int_{\ell} J^{\mu\nu}_{(n)|\ell} e^{i\ell \cdot x}$$

Structure of loop integrals

➤ Substituting the perturbative expansion into the EoMs

$$h_{(n)}^{\mu\nu} = \int_{\ell} e^{i\ell \cdot x} J_{(n)|\ell}^{\mu\nu} \quad \text{and} \quad \tilde{h}_{(n)}^{\mu\nu} = \int_{\ell} e^{i\ell \cdot x} \tilde{J}_{(n)|\ell}^{\mu\nu}$$

➤ **Perturbative Einstein eq**

$$h_1(x)h_2(x)\cdots h_n(x) \implies \int_{\ell_1} e^{i\ell_{12\dots n} \cdot x} \int_{\ell_2, \ell_3, \dots, \ell_n} J_{1|\ell_1} J_{2|\ell_2} \cdots J_{n|\ell_n} = \int_{\ell_1} e^{-i\ell_1 \cdot x} \int_{\ell_2, \ell_3, \dots, \ell_n} J_{1|-\ell_{12\dots n}} J_{2|\ell_2} \cdots J_{n|\ell_n}$$

$$\int_{\ell_3, \dots, \ell_n} \left(\int_{\ell_2} J_{1|-\ell_{12\dots n}} J_{2|\ell_2} \right) J_{3|\ell_3} \cdots J_{n|\ell_n}$$

One-loop bubble integral

$$J'_{1|\ell_1 \ell_3 \dots \ell_n}$$

$$\int_{\ell_4, \dots, \ell_n} \left(\int_{\ell_3} J'_{1|-\ell_{13\dots n}} J_{3|\ell_3} \right) J_{4|\ell_4} \cdots J_{n|\ell_n}$$

➤ **Fourier integrals** \iff **loop integrals**: number of loops = number of fields - 1

➤ **Integral Factorization** — iterative structure of loop integrals.

➤ This implies that **only bubble integrals are required**

Deriving and Solving the recursions

Recursions and currents at rank 1

➤ Rank-1 EoM — **Poisson equation**

$$\Delta h_{(1)}^{\mu\nu} = -2j^{\mu\nu} = -2Mv^\mu v^\nu \int_{\mathbf{k}} e^{i\mathbf{k}\cdot\mathbf{x}}$$

➤ Substituting the perturbative expansion $h_{(1)}^{\mu\nu} = \int_{\boldsymbol{\ell}} J_{\boldsymbol{\ell}}^{\mu\nu} e^{-i\boldsymbol{\ell}\cdot\mathbf{x}}$, we obtain the initial condition of the off-shell recursion relation

$$J_{(1)|\boldsymbol{\ell}}^{\mu\nu} = \frac{2\kappa^2 M}{|\boldsymbol{\ell}|^2} v^\mu v^\nu = \frac{16\pi GM}{|\boldsymbol{\ell}|^2} v^\mu v^\nu,$$

Or equivalently

$$J_{(1)|\boldsymbol{\ell}}^{00} = \frac{16\pi GM}{|\boldsymbol{\ell}|^2}, \quad J_{(1)|\boldsymbol{\ell}}^{0i} = 0, \quad J_{(1)|\boldsymbol{\ell}}^{ij} = 0.$$

➤ Since we are assuming an asymptotically flat metric, J^{ij} cannot be a plane wave.

➤ After Fourier transformation, we have the Newton potential — consistent with the metric expansion

$$h_{(1)}^{00} = \frac{4GM}{r} \quad h_{(1)}^{0i} = 0 \quad h_{(1)}^{ij} = 0$$

Recursions and currents at rank 2

➤ The corresponding recursion is

$$J_{(2)|-\ell_1}^{00} = \frac{\kappa}{|\ell_1|^2} \int_{\ell_2} \left[\frac{5}{4} |\ell_2|^2 - \frac{7}{8} \ell_{12} \cdot \ell_2 \right] J_{(1)|-\ell_{12}}^{00} J_{(1)|\ell_2}^{00},$$

$$J_{(2)|-\ell_1}^{ij} = \frac{\kappa}{|\ell_1|^2} \int_{\ell_2} \left[\frac{\ell_{12}^{(i} \ell_2^{j)}}{4} - \frac{\delta^{ij} \ell_{12} \cdot \ell_2}{8} \right] J_{(1)|-\ell_{12}}^{00} J_{(1)|\ell_2}^{00}.$$

➤ 1-loop **bubble integrals**

$$J_{(2)|-\ell_1}^{00} = \frac{(16\pi GM)^2}{|\ell_1|^2} \int_{\ell_2} \frac{1}{|\ell_2|^2 |\ell_{12}|^2} \left[\frac{5}{4} |\ell_2|^2 - \frac{7}{8} \ell_{12} \cdot \ell_2 \right] = \frac{14\pi^2 G^2 M^2}{|\ell_1|}.$$

$$J_{(2)|-\ell_1}^{ij} = \frac{(16\pi GM)^2}{8|\ell_1|^2} \int_{\ell_2} \left[\frac{2\ell_1^{(i} \ell_2^{j)} + 2\ell_2^i \ell_2^j - \delta^{ij} \ell_1^k \ell_2^k}{|\ell_2|^2 |\ell_{12}|^2} + \frac{\delta^{ij}}{2} \frac{1}{|\ell_{12}|^2} \right] = \pi^2 G^2 M^2 \left[-\frac{\ell_1^i \ell_1^j}{|\ell_1|^3} + \frac{\delta^{ij}}{|\ell_1|} \right]$$

➤ The Fourier transformation gives the correct perturbed metric

Recursions and currents at rank 3

➤ Rank-3 recursion

$$|\ell_1|^2 J_{(3)|-\ell_1}^{00} = - (GM)^3 \left[\ell_1^i X_{(3)|-\ell_1}^i + \ell_1^i \int_{\ell_2} \ell_{12}^j J_{(1)|-\ell_{12}}^{00} J_{(2)|\ell_2}^{ij} \right].$$

$$|\ell_1|^2 J_{(3)|-\ell_1}^{ij} = \int_{\ell_2} \left[8d_{(2)|-\ell_{12}}^{(i} d_{(1)|\ell_2}^{j)} - 2h_{(2)}^{ij} \ell_2^k d_{(1)|\ell_2}^k \right] + 2\delta^{ij} \ell_1^k d_{(3)}^k + \frac{1}{2} W_{(3)}^{ij}$$

$$X_{(n)|-\ell_1}^i = \int_{\ell_2} \ell_2^i \sum_{m=1}^{n-1} \tilde{J}_{(n-m)|-\ell_{12}}^{00} J_{(m)|\ell_2}^{00}, \quad Y_{(n)|-\ell_1}^i = \int_{\ell_2} \ell_2^i \tilde{J}_{(n-2)|-\ell_{12}}^{kl} J_{(2)|\ell_2}^{kl},$$

➤ Again, we need only 1-loop bubble integrals.

➤ In dimensional regularization

Scaleless integral
vanishes in dim. Reg.

$$J_{(3)|-\ell_1}^{00} = \frac{(GM)^3}{|\ell_1|^{d-3}} 2^{d+1} \pi^{\frac{d-1}{2}} \Gamma\left(\frac{d-3}{2}\right), \quad J_{(3)|\ell}^{ij} = - \int_{\ell_2} \frac{16\pi^3 \delta^{ij}}{3 |\ell_1|^2 |\ell_2|} = 0.$$

➤ The only place where divergences arise!

➤ **other regularization scheme** — it does not vanish, and the solution should be modified!

➤ This explains the ambiguity, C factor

All-order Currents

➤ From $n \geq 5$ cases, the forms of the EoM/Recursion are fixed.

➤ In the harmonic gauge, the Landau-Lifshitz variables are extremely simple

$$h^{00} = -1 + \frac{(r + GM)^3}{r^2(r - GM)} = \frac{4GM}{r} + \frac{7G^2M^2}{r^2} + \frac{8G^3M^3}{r^3} + \frac{8G^4M^4}{r^4} + \dots,$$

$$h^{ij} = \frac{G^2M^2x^ix^j}{r^4}.$$

➤ One can read off the currents arbitrary order in G from the Fourier transformation

$$J_{(1)|\ell}^{00} = \frac{4(GM)2^{D-1}\pi^{\frac{D}{2}}\Gamma\left[\frac{D-1}{2}\right]}{\Gamma\left[\frac{1}{2}\right]} \frac{1}{|\ell|^{D-1}},$$

$$J_{(2)|\ell}^{00} = 7(GM)^2 2^{D-2}\pi^{\frac{D}{2}}\Gamma\left[\frac{D-2}{2}\right] \frac{1}{|\ell|^{D-2}},$$

$$J_{(n)|\ell}^{00} = \frac{8(GM)^n \pi^{\frac{D}{2}} \Gamma\left[\frac{D-n}{2}\right]}{2^{n-D} \Gamma\left[\frac{n}{2}\right]} \frac{1}{|\ell|^{D-n}}, \quad \text{for } n \geq 3$$

$$J_{(2)|\ell}^{ij} = (GM)^2 \pi^{\frac{D}{2}} 2^{D-3} \left[-\frac{2\Gamma\left[\frac{D}{2}\right]\ell^i\ell^j}{|\ell|^D} + \frac{\Gamma\left[\frac{D-2}{2}\right]\delta^{ij}}{|\ell|^{D-2}} \right].$$

➤ One can show the followings by using the **induction**

Arbitrary rank $n \geq 5$ — J^{00}

➤ We can show that the off-shell currents at an arbitrary order n by **induction**.

➤ The corresponding recursion: $J_{(n)|\ell}^{00} = \mathcal{E}_{(n)|\ell}^{[1]} - \mathcal{E}_{(n)|\ell}^{[2]}$,

even

$$\mathcal{E}_{(2n)|-\ell_1}^{[1]} = (GM)^{2n} \frac{\ell_1^i}{|\ell_1|^2} \left(-X_{(2n)|-\ell_1}^i + Y_{(2n)|-\ell_1}^i \right),$$

$$\mathcal{E}_{(2n)|-\ell_1}^{[2]} = (GM)^{2n} \frac{\ell_1^i}{|\ell_1|^2} \int_{\ell_2} \left(-X_{(2n-2)|-\ell_{12}}^j + Y_{(2n-2)|-\ell_{12}}^j - \ell_{12}^j J_{(2n-2)|-\ell_{12}}^{00} \right) J_{(2)|\ell_2}^{ij}.$$

odd

$$\mathcal{E}_{(2n+1)|-\ell_1}^{[1]} = - (GM)^{2n+1} \frac{\ell_1^i}{|\ell_1|^2} X_{(2n+1)|-\ell_1}^i,$$

$$\mathcal{E}_{(2n+1)|-\ell_1}^{[2]} = (GM)^{2n+1} \frac{\ell_1^i}{|\ell_1|^2} \int_{\ell_2} \left(-X_{(2n-1)|-\ell_{12}}^j - \ell_{12}^j J_{(2n-1)|-\ell_{12}}^{00} \right) J_{(2)|\ell_2}^{ij}.$$

➤ Performing the bubble integrals, we have

$$J_{(2n)}^{00} = \frac{8(GM)^{2n} \pi^{\frac{D}{2}} 2^{D-2n} \Gamma[\frac{D}{2} - n]}{\Gamma[n]} \frac{1}{|\ell|^{D-2n}},$$

$$J_{(2n+1)}^{00} = \frac{8(GM)^{2n} \pi^{\frac{D}{2}} 2^{D-2n-1} \Gamma[\frac{D-2n-1}{2}]}{\Gamma[n + \frac{1}{2}]} \frac{1}{|\ell|^{D-2n-1}},$$

Arbitrary rank $n \geq 5$ — J^{ij}

➤ The EoM for the spatial components

$$\Delta h_{(n)}^{ij} = \sum_{m=1}^{n-1} 4d_{(n-m)}^i d_{(m)}^j + 2\sigma^{ij} \partial_k d^k + Z_{(n)k}^{k(ij)} - 2Z_{(n)k}^{(i|k|j)} + \frac{1}{2} Z_{(n)k}^{(i|k|j)} + \frac{1}{2} W_{(n)}^{(ij)} - \left(Z_{(n-2)kl}^{k(i} - 2Z_{(n-2)kl}^{(i|k|} + \frac{1}{2} Z_{(n-2)lk}^{(i|k|} + \frac{1}{2} W_{(n-2)}^{(i|l|} \right) h_{(2)}^{j)l}$$

➤ Divide the EoM into **3 sectors**: d-sector, W-sector and Z-sector

➤ Interestingly, these three sectors vanish individually (**induction**).

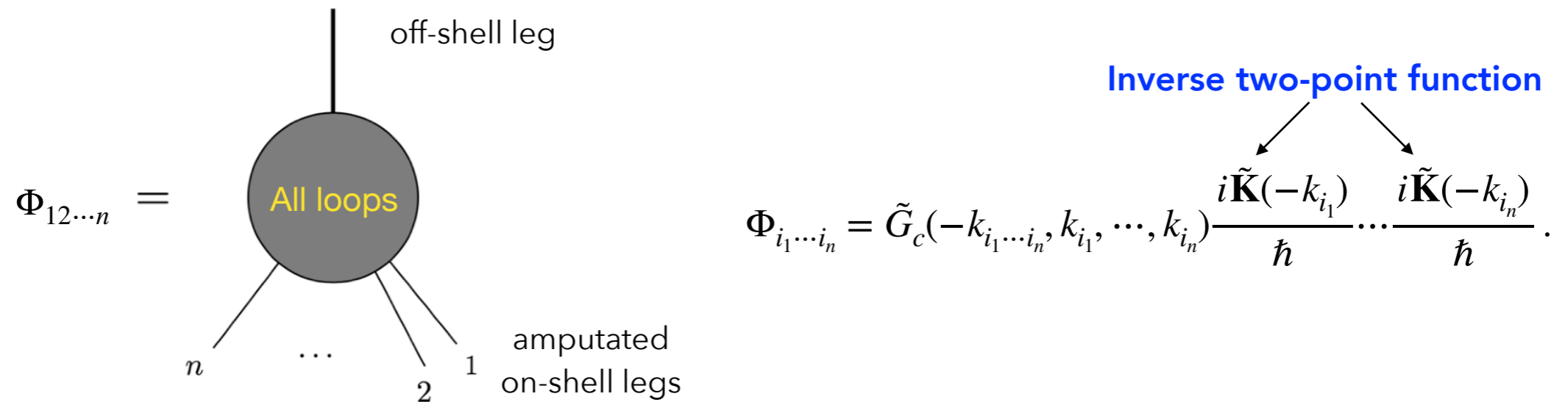
➤ This implies $J_{(n)|\mathcal{E}}^{ij} = 0$, as we expected

➤ This shows the all-order perturbative expansion satisfies the Einstein equation.

Quantum Generalization

Quantum Perturbative Method [KL '22]

➤ **Quantum off-shell currents:** sum of all $(n + 1)$ -point all-loop Feynman diagrams



➤ **Fields** in quantum effective action formalism — **1pt function** with the external source $j_x \equiv j(x)$

$$\varphi_x = \frac{\delta W[j]}{\delta j_x} = \sum_{n=1}^{\infty} \frac{i^n}{\hbar^n n!} \int_{y_1, y_2, \dots, y_n} G_c(x, y_1, y_2, \dots, y_n) j_{y_1} j_{y_2} \dots j_{y_n}$$

➤ Choice of the external source for **amplitude** $j_x = \sum_{i=1}^N \int_{y_i} K_{xy_i} e^{-ik_i \cdot x} = \sum_{i=1}^N \tilde{K}(-k_i) e^{-ik_i \cdot x}$

➤ **Quantum perturbative expansion:** $\varphi_x = \sum_{\mathcal{P}} \Phi_{\mathcal{P}} e^{-ik_{\mathcal{P}} \cdot x}$

Substituting into the "quantum" EoM?

Dyson-Schwinger equation

➤ Quantum analogous of classical EoM: **Dyson-Schwinger equation**

➤ **Quantization** \iff **deformation** of a field to an **operator**

$$\phi_x \mapsto \hat{\phi}_x = \phi_x + \frac{\hbar}{i} \frac{\delta}{\delta j_x}$$

➤ **DS equation** for phi-4 theory:

$$\int_y K(x, y) \phi_y + \frac{\lambda}{3!} \phi_x^3 = j_x - \frac{\lambda \hbar}{2 i} \phi_x \frac{\delta \phi_x}{\delta j_x} + \hbar^2 \frac{\lambda}{3!} \frac{\delta^2 \phi_x}{\delta j_x \delta j_x}.$$

➤ **Strategy:** Treat the functional derivatives ϕ_x as **new independent field variables**

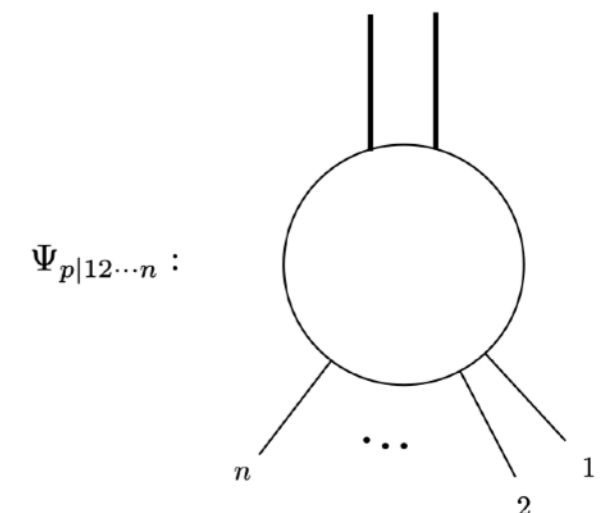
➤ **Descendant fields:** higher point functions with external sources, multiple off-shell legs

$$\text{1st : } \psi_{xy} = \frac{\delta \phi_x}{\delta j_x} = \frac{\delta^2 W[j]}{\delta j_x \delta j_y}, \quad \text{2nd : } \psi'_{xyz} = \frac{\delta^2 \phi_x}{\delta j_x \delta j_x} = \frac{\delta^3 W[j]}{\delta j_x \delta j_y \delta j_z}, \quad \dots,$$

➤ **Derive the perturbative expansion for the descendant fields**

$$\psi_{x,y} = \int_p \Psi_{p|\emptyset} e^{ip \cdot (x-y)} + \sum_{\mathcal{P}} \int_p \Psi_{p|\mathcal{P}} e^{ip \cdot (x-y)} e^{-ik_{\mathcal{P}} \cdot x},$$

$$\psi'_{x,y,z} = \sum_{\mathcal{P}} \int_{p,q} \Psi'_{p,q|\mathcal{P}} e^{ip \cdot (x-y) + iq \cdot (x-z)} e^{-ik_{\mathcal{P}} \cdot x},$$



Descendant equations [KL '22]

➤ Derive the **descendant equations**: acting $\frac{\delta}{\delta j_x}$ on the DS eq.

$$\begin{aligned} \psi_{x,z} = & D_{xz} - \frac{\lambda}{2} \int_y D_{xy} \phi_y^2 \psi_{y,z} + i\hbar \frac{\lambda}{2} \int_y D_{xy} (\phi_y \psi'_{y,y,z} + \psi_{y,z} \psi_{y,y}) \\ & + \hbar^2 \frac{\lambda}{3!} \int_y D_{xy} \psi''_{y,y,y,z}, \end{aligned}$$

$$\begin{aligned} \psi'_{x,z,w} = & -\frac{\lambda}{2} \int_y D_{xy} (2\phi_y \psi_{y,w} \psi_{y,z} + \phi_y^2 \psi'_{y,z,w}) \\ & + i\hbar \frac{\lambda}{2} \int_y D_{xy} (\psi_{y,w} \psi'_{y,y,z} + \phi_y \psi''_{y,y,z,w} + \psi'_{y,z,w} \psi_{y,y} + \psi_{y,z} \psi'_{y,y,w}) \\ & + \hbar^2 \frac{\lambda}{3!} \int_y D_{xy} \psi'''_{y,y,y,z,w}. \end{aligned}$$

➤ However, new descendant fields arise ψ'' and ψ'''

➤ How to truncate them?

\hbar expansion and recursions

➤ Up to now, all the equations are exact

➤ \hbar expansion

$$\varphi_x = \sum_{n=0}^{\infty} \left(\frac{\hbar}{i}\right)^n \varphi_x^{(n)}, \quad \psi_{x,y} = \sum_{n=0}^{\infty} \left(\frac{\hbar}{i}\right)^n \psi_{x,y}^{(n)}, \quad \psi'_{x,y,z} = \sum_{n=0}^{\infty} \left(\frac{\hbar}{i}\right)^n \psi'_{x,y,z}{}^{(n)}$$

➤ We can truncate the new descendant fields **because these are from higher \hbar -order terms**

➤ **1-loop DS equations and tree-level descendant equation**

$$\phi_x^{(1)} = \int_y D_{xy} \left[j_y^{(1)} - \frac{\lambda}{2} \left(\left(\phi_y^{(0)} \right)^2 \phi_y^{(1)} + \phi_y^{(0)} \psi_{y,y}^{(0)} \right) \right]$$

$$\psi_{x,z}^{(0)} = D_{xz} - \frac{\lambda}{2} \int_y D_{xy} \left(\phi_y^{(0)} \right)^2 \psi_{y,z}^{(0)}$$

➤ Substitute the perturbiner expansion into the DS equation

$$\Phi_{\mathcal{P}}^{(1)} = -\frac{\lambda}{2} \frac{1}{(k_{\mathcal{P}})^2 + m^2} \left(\sum_{\mathcal{P}=\mathcal{Q}\cup\mathcal{R}\cup\mathcal{S}} \Phi_{\mathcal{Q}}^{(0)} \Phi_{\mathcal{R}}^{(0)} \Phi_{\mathcal{S}}^{(1)} + \sum_{\mathcal{P}=\mathcal{Q}\cup\mathcal{R}} \int_p \Phi_{\mathcal{Q}}^{(0)} \Psi_{p|\mathcal{R}}^{(0)} \right)$$

$$\Psi_{p|\mathcal{P}}^{(0)} = -\frac{\lambda}{2} \sum_{\mathcal{P}=\mathcal{Q}\cup\mathcal{R}\cup\mathcal{S}} \frac{1}{(p - k_{\mathcal{P}})^2 + m^2} \Phi_{\mathcal{Q}}^{(0)} \Phi_{\mathcal{R}}^{(0)} \Psi_{p|\mathcal{S}}^{(0)}$$

reduction of higher loop integrals

➤ 2-loop recursion

$$\Phi_{\mathcal{P}}^{(2)} = -\frac{\lambda}{2} \frac{1}{k_{\mathcal{P}}^2 + m^2} \left(\sum_{\mathcal{P}=\mathcal{Q}\cup\mathcal{R}\cup\mathcal{S}} \left(\Phi_{\mathcal{Q}}^{(0)} \Phi_{\mathcal{R}}^{(0)} \Phi_{\mathcal{S}}^{(2)} + 2\Phi_{\mathcal{Q}}^{(0)} \Phi_{\mathcal{R}}^{(1)} \Phi_{\mathcal{S}}^{(1)} \right) + \sum_{\mathcal{P}=\mathcal{Q}\cup\mathcal{R}} \int_p \left(\Phi_{\mathcal{Q}}^{(0)} \Psi_{p|\mathcal{R}}^{(1)} + \Phi_{\mathcal{Q}}^{(1)} \Psi_{p|\mathcal{R}}^{(0)} \right) + \frac{1}{3} \int_{p,q} \Psi_{p,q|\mathcal{P}}^{(0)} \right),$$

$$\Psi_{p|\mathcal{P}}^{(1)} = -\frac{\lambda}{2} \frac{1}{(p - k_{\mathcal{P}})^2 + m^2} \sum_{\mathcal{P}=\mathcal{Q}\cup\mathcal{R}\cup\mathcal{S}} \left(2\Phi_{\mathcal{Q}}^{(0)} \Phi_{\mathcal{R}}^{(1)} \Psi_{p|\mathcal{S}}^{(0)} + \Phi_{\mathcal{Q}}^{(0)} \Phi_{\mathcal{R}}^{(0)} \Psi_{p|\mathcal{S}}^{(1)} \right) - \frac{\lambda}{2} \frac{1}{(p - k_{\mathcal{P}})^2 + m^2} \sum_{\mathcal{P}=\mathcal{Q}\cup\mathcal{R}} \int_q \left(\Phi_{\mathcal{Q}}^{(0)} \Psi_{p,q|\mathcal{R}}^{(0)} + \Psi_{p|\mathcal{Q}}^{(0)} \Psi_{q|\mathcal{R}}^{(0)} \right),$$

$$\Psi_{p,q|\mathcal{P}}^{(0)} = -\frac{\lambda}{2} \frac{1}{(p + q - k_{\mathcal{P}})^2 + m^2} \sum_{\mathcal{P}=\mathcal{Q}\cup\mathcal{R}\cup\mathcal{S}} \left(2\Phi_{\mathcal{Q}}^{(0)} \Psi_{p|\mathcal{R}}^{(0)} \Psi_{q|\mathcal{S}}^{(0)} + \Phi_{\mathcal{Q}}^{(0)} \Phi_{\mathcal{R}}^{(0)} \Psi_{p,q|\mathcal{S}}^{(0)} \right).$$

➤ General structure

- Theories with n -point vertex — $(n - 2)$ loop reducible
- pure YM/pure GR — 3pt vertices using the first-order formalism (1-loop reducible)

Steps of deriving the recursions

1. Write down the EoM
2. Constructing the **Dyson-Schwinger equation** from the EoM by the deformation
3. Substituting the **perturbative expansion**
4. \hbar -expansion and truncate the higher \hbar order terms
5. Deriving the off-shell recursion relation
6. Solve them!

Applied to

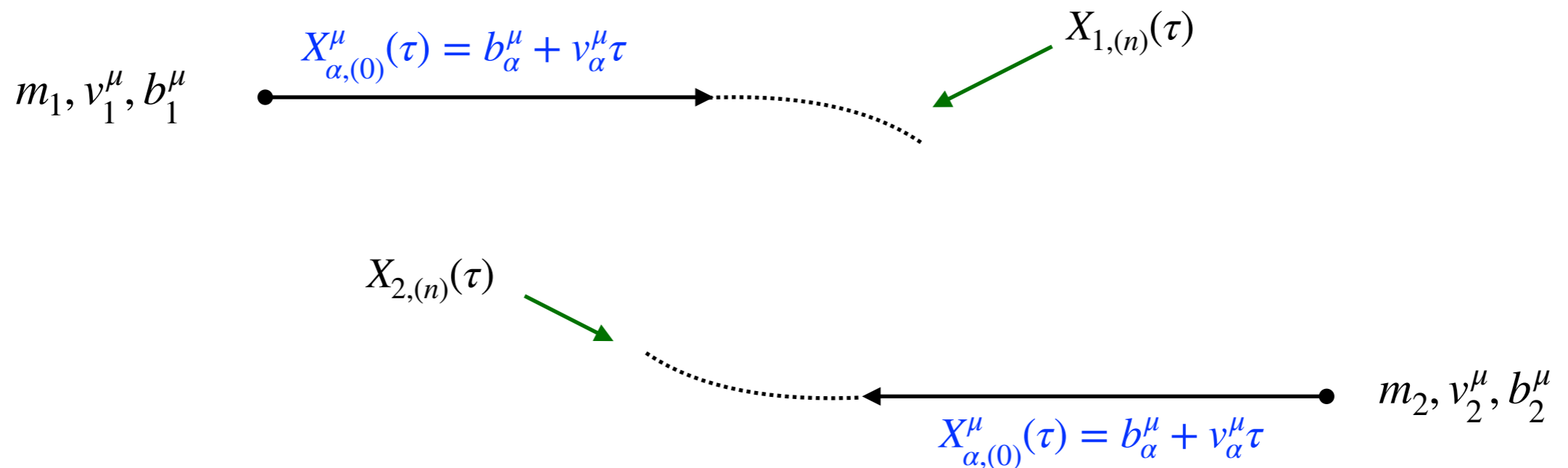
- phi-4 theory: 2-loop
- Pure Yang-Mills theory: 1-loop 4pt
- Einstein-scalar theory (for binary BH system): 2-loop 4pt

Generalization to Binary black holes

Kinematics

- In the inspiral phase, we may treat the spinless BHs as **point particles**
- Considering two body motions (massive point particles)

$$X_{\alpha}^{\mu}(\tau) = X_{\alpha,(0)}^{\mu}(\tau) + \sum_{n=1}^{\infty} G^n X_{\alpha,(n)}^{\mu}(\tau) \quad \alpha = 1,2$$



- We will consider the conservative potential — leading order

- **Goal:** Compute the **momentum kick** order by order in G , $\Delta P_{1,2}^\mu = \int_{-\infty}^{\infty} d\tau \dot{X}_{1,2}^\mu(\tau)$

Action/EoM for two point masses

➤ Change the notation:

$$\mathfrak{g}^{\mu\nu} := \sqrt{-g} g^{\mu\nu}, \quad \mathfrak{g}_{\mu\nu} := \frac{1}{\sqrt{-g}} g_{\mu\nu}$$

➤ Action:

$$S[\mathfrak{g}, j] = S_{\text{EH}}[\mathfrak{g}] + \frac{1}{16\pi G} \int d^4x j^{\mu\nu}(x) \frac{1}{\sqrt{-\mathfrak{g}}} \mathfrak{g}_{\mu\nu}(x)$$

The external source/Energy momentum tensor

$$j^{\mu\nu}(x) = 8\pi G \sum_{\alpha=1}^2 m_{\alpha} \int d\tau \frac{dX_{\alpha}^{\mu}(\tau)}{d\tau} \frac{dX_{\alpha}^{\nu}(\tau)}{d\tau} \delta^4(x^{\mu} - X_{\alpha}^{\mu}(\tau))$$

➤ Einstein equation:

$$\delta_{\mathfrak{g}} S = \frac{1}{16\pi G} \int d^Dx \delta \mathfrak{g}_{\mu\nu} \left[-\mathcal{G}^{\mu\nu} + \frac{1}{\sqrt{-\mathfrak{g}}} \left(j^{\mu\nu} - \frac{1}{2} \mathfrak{g}^{\mu\nu} j^{\rho\sigma} \mathfrak{g}_{\rho\sigma} \right) \right]$$

➤ Geodesic equation:

$$\frac{d}{d\tau} \left[\frac{1}{\sqrt{-\mathfrak{g}}} \tilde{\mathfrak{g}}_{\rho\sigma} \dot{X}_{\alpha}^{\sigma} \right] = \frac{1}{2} \partial_{\rho} \left[\frac{1}{\sqrt{-\mathfrak{g}}} \tilde{\mathfrak{g}}_{\mu\nu} \right]_{x \rightarrow X_{\alpha}} \dot{X}_{\alpha}^{\mu} \dot{X}_{\alpha}^{\nu}$$

Perturbiner expansion

- The n -th order metric fluctuations $h_{(n)}^{\mu\nu}$ is a function of coordinates x^μ , as well as the impact parameter b_α **implicitly**

$$h_{(n)}^{\mu\nu}(x) = h_{(n)}^{\mu\nu}(x; b_1, b_2)$$

- **Perturbiner expansion** ($x_\alpha^\mu = x^\mu - b_\alpha^\mu$, $\alpha = 1, 2$, ℓ_α are the Fourier dual of x_α)

$$n = 1 : h_{(1)}^{\mu\nu}(x_\alpha) = \int_{\ell_1} e^{il_1 \cdot x_1} J_{(1)|[\ell_1, 0]}^{\mu\nu} + \int_{\ell_2} e^{il_2 \cdot x_2} J_{(1)|[0, \ell_2]}^{\mu\nu}$$

$$n > 1 : h_{(n)}^{\mu\nu}(x_\alpha) = \int_{\ell_1} e^{il_1 \cdot x_1} J_{(n)|[\ell_1, 0]}^{\mu\nu} + \int_{\ell_2} e^{il_2 \cdot x_2} J_{(n)|[0, \ell_2]}^{\mu\nu} + \int_{\ell_1, \ell_2} e^{il_1 \cdot x_1 + il_2 \cdot x_2} J_{(n)|[\ell_1, \ell_2]}^{\mu\nu}.$$

$$n = 1 : X_{\alpha, (1)}^\rho(\tau) = \int_{\ell_1} X_{\alpha, (1)|[\ell_1, 0]}^\mu e^{-il_1 \cdot X_{\alpha, 1, (0)}} + \int_{\ell_2} X_{\alpha, (1)|[0, \ell_2]}^\mu e^{-il_2 \cdot X_{\alpha, 2, (0)}},$$

$$n > 1 : X_{\alpha, (n)}^\rho(\tau) = \int_{\ell_1} X_{\alpha, (n)|[\ell_1, 0]}^\mu e^{-il_1 \cdot X_{\alpha, 1, (0)}} + \int_{\ell_2} X_{\alpha, (n)|[0, \ell_2]}^\mu e^{-il_2 \cdot X_{\alpha, 2, (0)}} \\ + \int_{\ell_1, \ell_2} X_{\alpha, (n)|[\ell_1, \ell_2]}^\mu e^{-il_1 \cdot X_{\alpha, 1, (0)} - il_2 \cdot X_{\alpha, 2, (0)}},$$

$$X_{\alpha, \beta, (0)}^\mu = b_\alpha^\mu - b_\beta^\mu + v_\alpha^\mu \tau.$$

Summary and future directions

- ▶ **Established a new computational framework for perturbative GR**
 - ▶ defined a “good” variable — tensor density & doubled metric
 - ▶ derived a recursion relation in a remarkably simple form — no infinite expansion
 - ▶ Showed the integral factorization occurs — only bubble integrals arise
 - ▶ Derived Schwarzschild BH solution all order in Newton constant
- ▶ **Applications**
 - ▶ Extension to binary black holes — two moving point masses
 - ▶ Kerr BH — massive higher spin or worldline SUSY
 - ▶ Finding interesting unknown solutions — physically intuitive setup.
 - ▶ Computing scattering amplitudes for QCD/SM

Thank you!