

# Timing properties of T2K ND280Upgrade detector

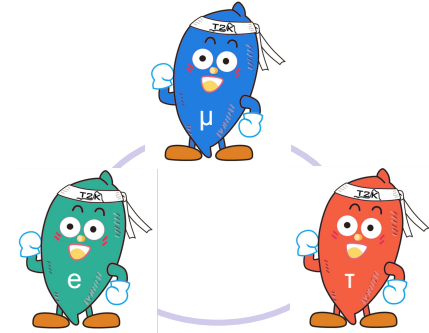
Gioele Reina (he/him)

MPA retreat 2024

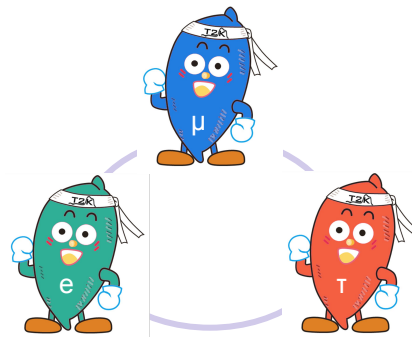


- ❑ **Neutrino oscillations**
- ❑ **T2K experiment and its near detector ND280**
- ❑ **Neutron reconstruction in ND280**
- ❑ **Timing characterization of the detector**

Neutrino are produced in a specific **flavour eigenstate** ( $\nu_e, \nu_\mu, \nu_\tau$ ) but they travel through the **mass eigenstate** ( $\nu_1, \nu_2, \nu_3$ ) and so can be detected in a different flavour



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$$\begin{pmatrix} \nu_e \\ \nu_\mu \\ \nu_\tau \end{pmatrix} = \begin{pmatrix} 1 & 0 & 0 \\ 0 & c_{23} & s_{23} \\ 0 & -s_{23} & c_{23} \end{pmatrix} \begin{pmatrix} c_{13} & 0 & s_{13}e^{-i\delta_{CP}} \\ 0 & 1 & 0 \\ -s_{13}e^{i\delta_{CP}} & 0 & c_{13} \end{pmatrix} \begin{pmatrix} c_{12} & s_{12} & 0 \\ -s_{12} & c_{12} & 0 \\ 0 & 0 & 1 \end{pmatrix} \begin{pmatrix} \nu_1 \\ \nu_2 \\ \nu_3 \end{pmatrix}$$

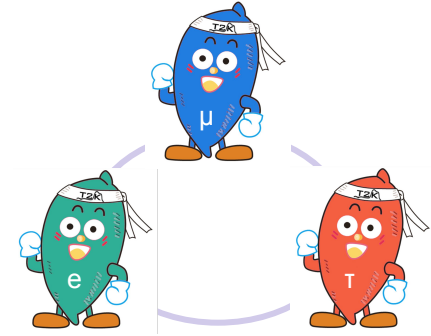
$$s_{ij} = \sin \theta_{ij}, \quad c_{ij} = \cos \theta_{ij}$$

**Two flavour neutrino oscillation probability** where:

$$P(\nu_\alpha \rightarrow \nu_\beta) = \sin^2(2\theta) \sin^2\left(1.27 \Delta m^2 \frac{L(km)}{E(GeV)}\right)$$

- ❑  $L$  is the distance to the source
- ❑  $E$  is the neutrino energy
- ❑  $\Delta m^2 = m_k^2 - m_l^2$  with  $m$  neutrino mass eigenstate  $k, l = 1, 2, 3$

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$$s_{ij} = \sin \theta_{ij}, \quad c_{ij} = \cos \theta_{ij}$$

## Atmospherics and LBL

$$\theta_{23} \sim 47.8^\circ$$

$$|\Delta m_{32}^2| \sim 2.5 \times 10^{-3} \text{ eV}^2$$

## Reactors and LBL

$$\theta_{13} \sim 8.5^\circ$$

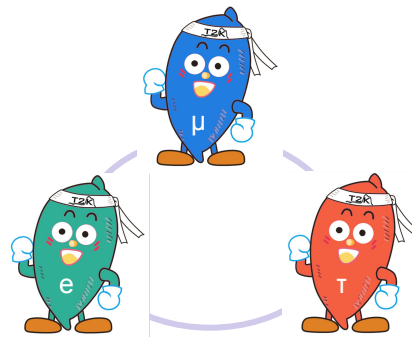
$$\delta_{CP} \text{ unknown}$$

## Solar and Reactors

$$\theta_{12} \sim 33.6^\circ$$

$$\Delta m_{12}^2 \sim 7.5 \times 10^{-5} \text{ eV}^2$$

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Reactors and LBL

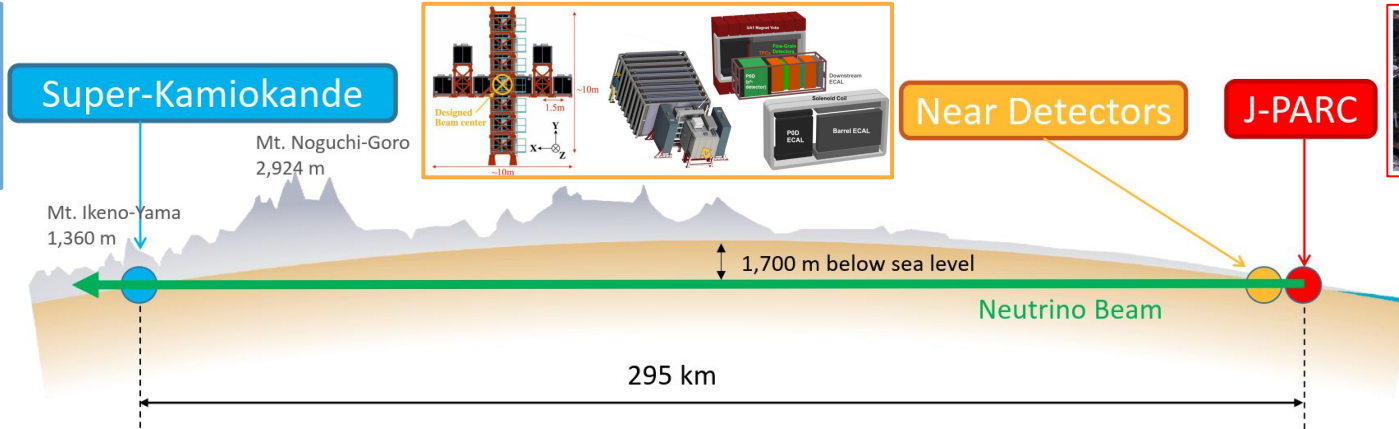
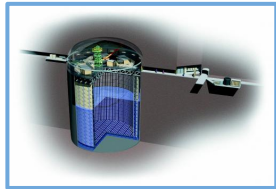
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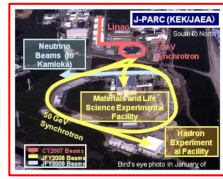
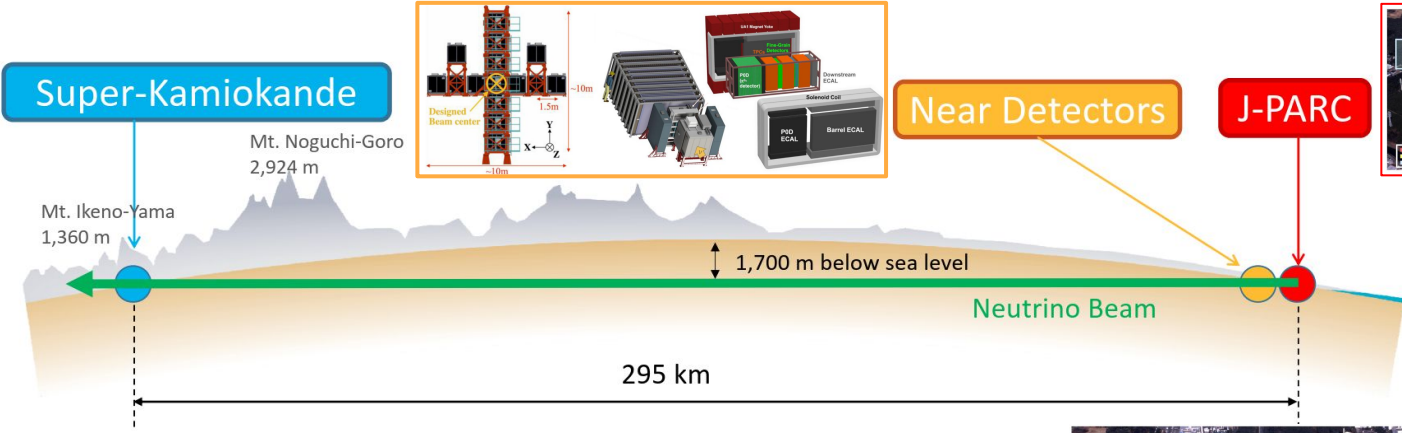
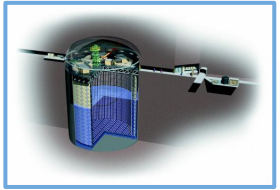
Solar and Reactors

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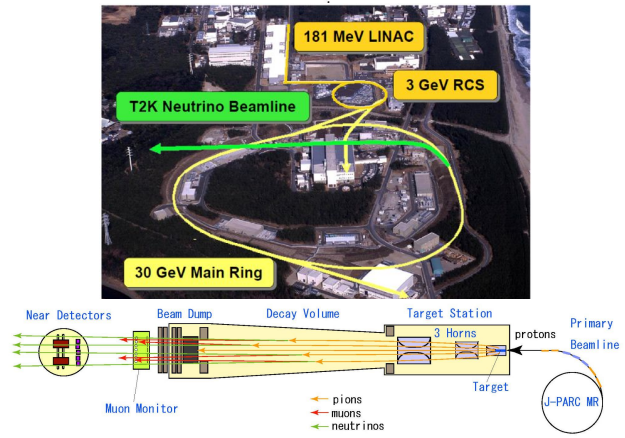
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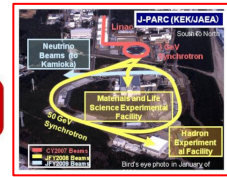
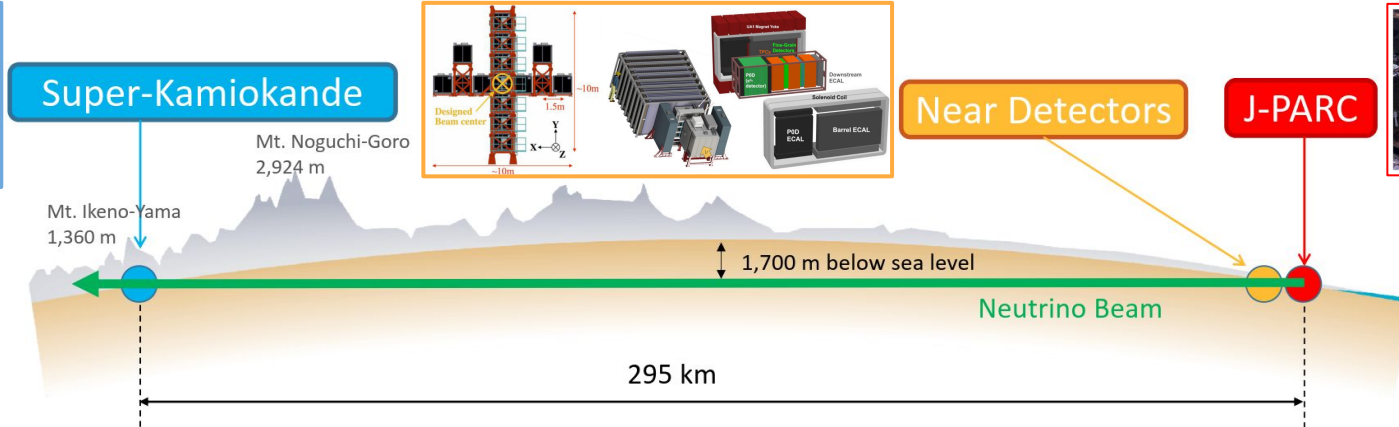
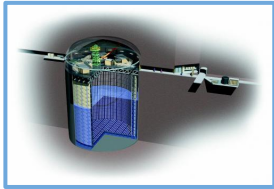
- ❑ Long-baseline neutrino experiment
- ❑ Neutrino oscillation parameters measurement:  $\delta_{CP}$ ,  $\theta_{23}$ ,  $\Delta m_{23}^2$ ,  $\theta_{13}$
- ❑ Accelerator  $\nu_{\mu}$  produced at J-PARC
- ❑  $\nu$  detected in two facilities:
  - ❑ Three near detectors (INGRID, WAGASCI-BabyMIND, ND280)
  - ❑ Super-Kamiokande far detector (2.5° off-axis)



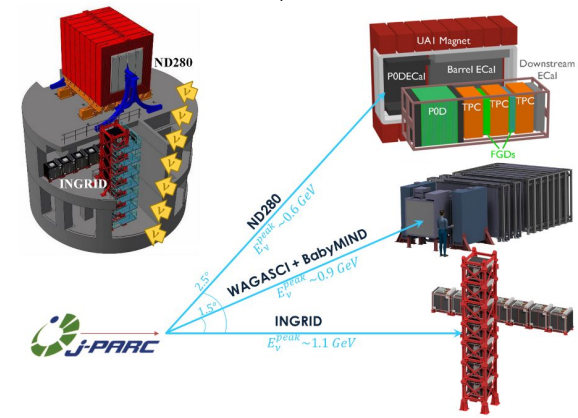
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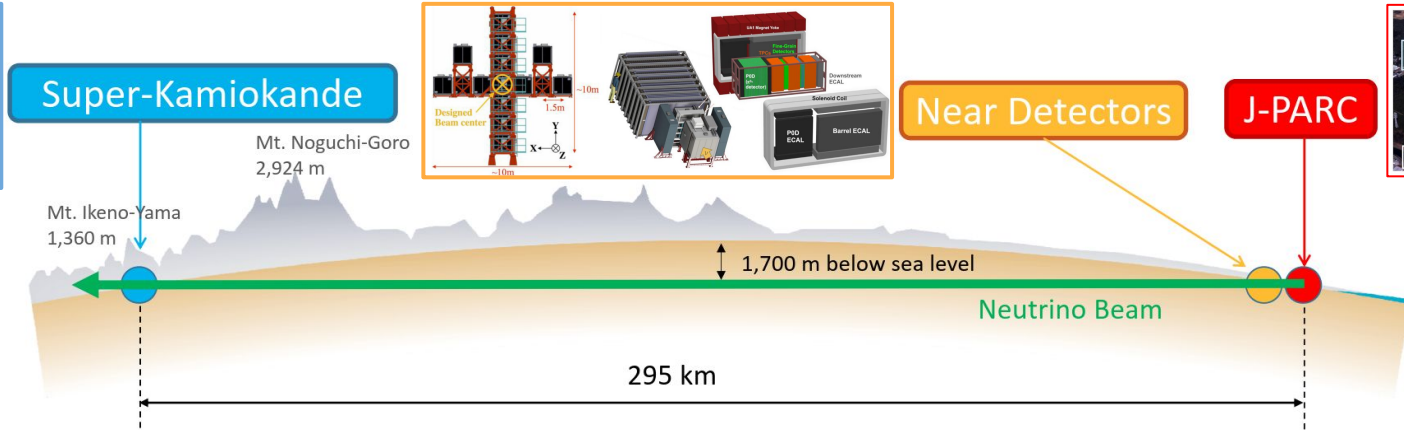
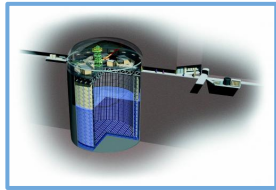




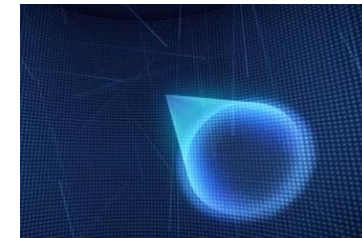
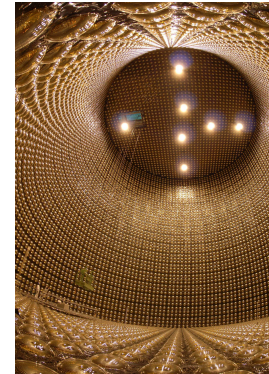


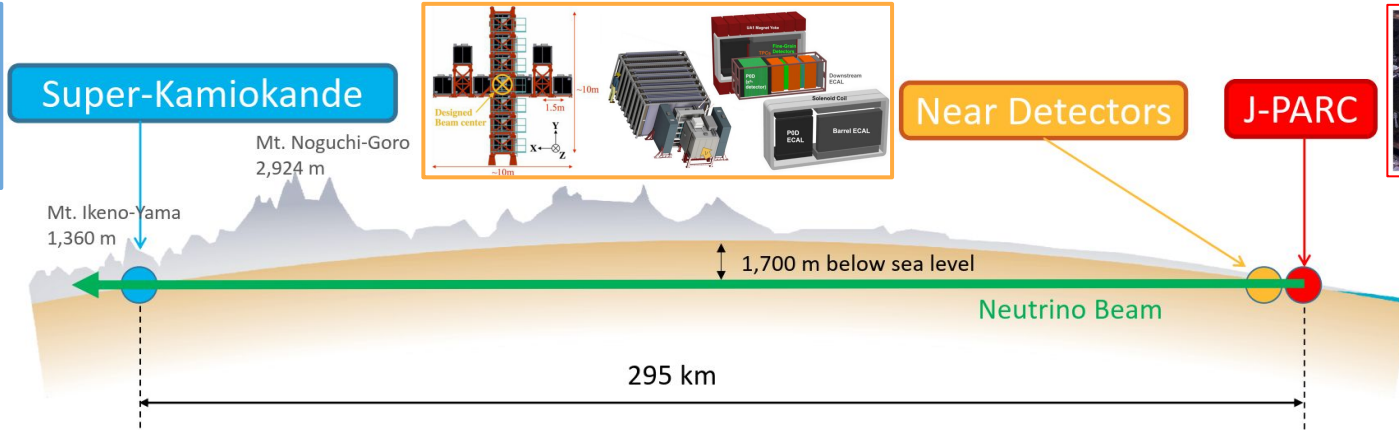
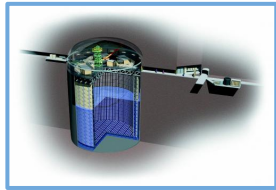
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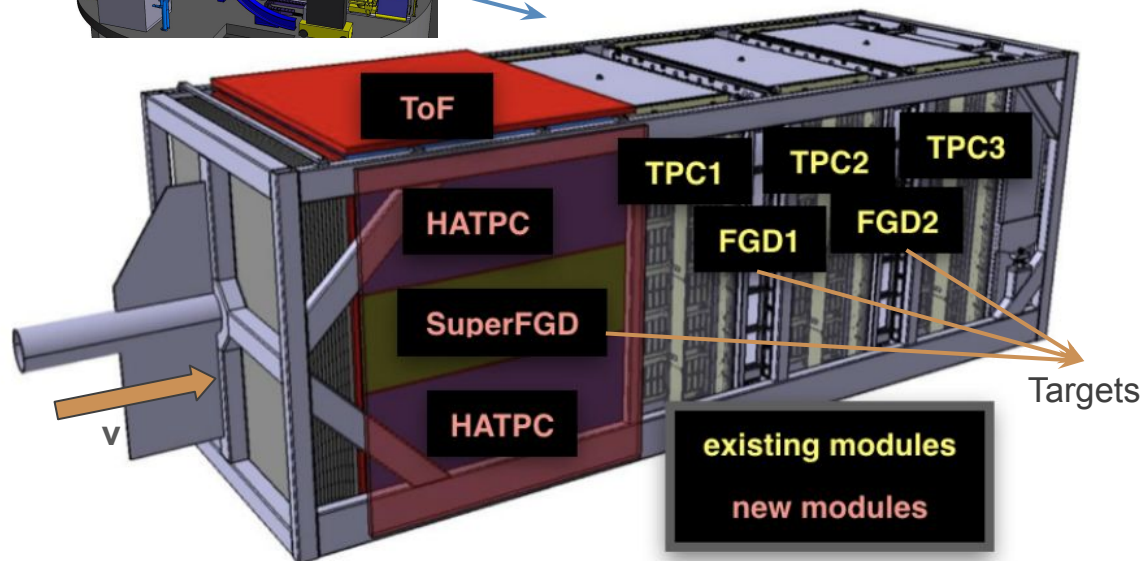
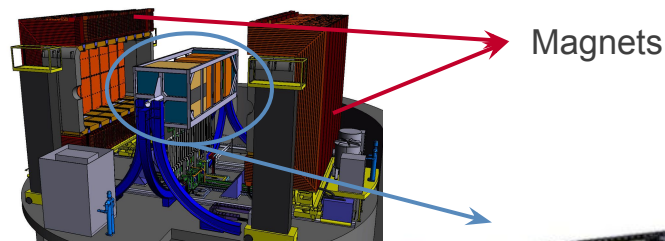


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- ❑ 2.5° off-axis
- ❑ Flux and detector error
- ❑ **ND280 Upgrade**

Three new detectors:

- ❑ **Time of Flight (ToF)**
  - ❑ 150 ps time resolution
- ❑ **High-Angle TPC (HATPC)**
  - ❑ full angular coverage
- ❑ **Super Fine-Grained Detector (FGD)**
  - ❑ novel detector concept



- ❑ **> 2M independent 1 cm<sup>3</sup> cubes** of plastic scintillator
- ❑ **WLS fibers** crossing the cube along three orthogonal directions
- ❑ **Photosensor** at one end of each WLS fibers

High granularity

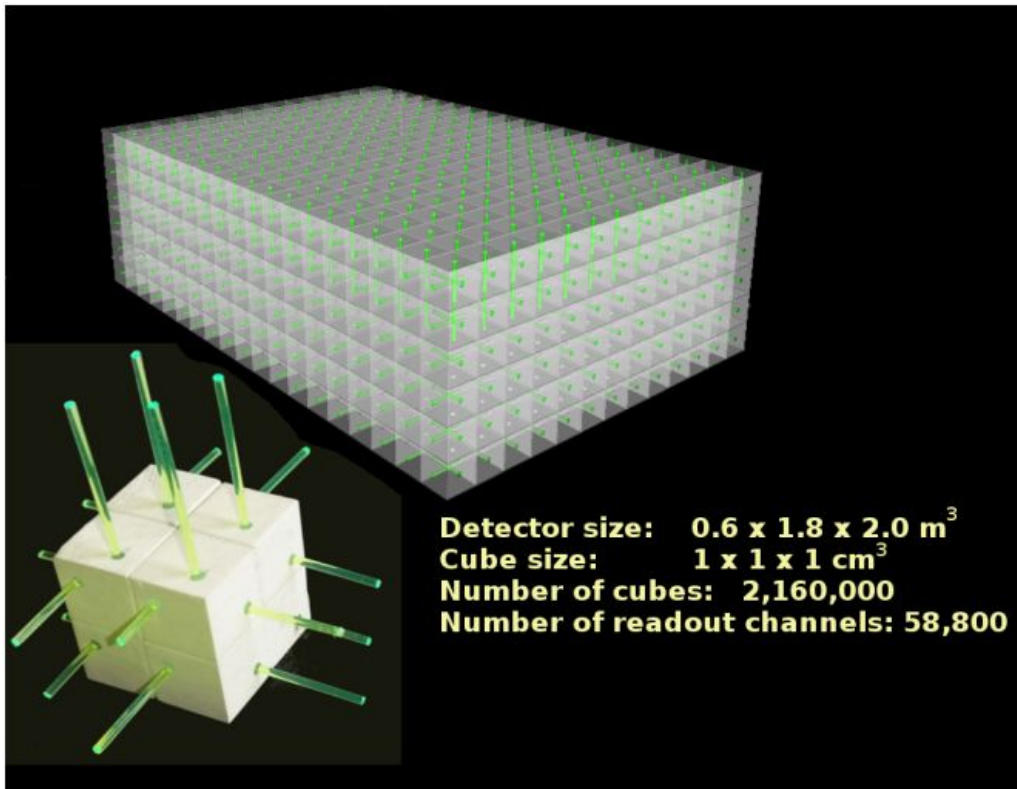
3D reconstruction



PID improvement

Low energy threshold

Neutron reconstruction



**Detector size:** 0.6 x 1.8 x 2.0 m<sup>3</sup>  
**Cube size:** 1 x 1 x 1 cm<sup>3</sup>  
**Number of cubes:** 2,160,000  
**Number of readout channels:** 58,800

- ❑ **> 2M independent 1 cm<sup>3</sup> cubes** of plastic scintillator
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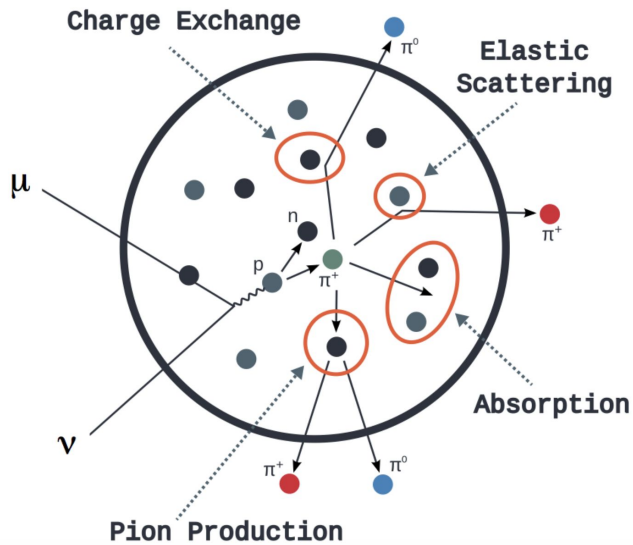


Neutrino energy reconstruction from lepton kinematics

**Nuclear effects** are not considered



**Bias in the neutrino energy reconstruction**



- ❑ **Final State Interactions (FSI)** make interaction identification challenging
- ❑ Topologies identification is the way
- ❑ Neutrons come in many of these topologies: (**CC0 $\pi$ 1n**, **CC0 $\pi$ 0p1n**, **CC0 $\pi$ 0pNn**)

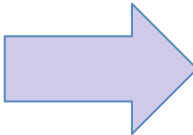
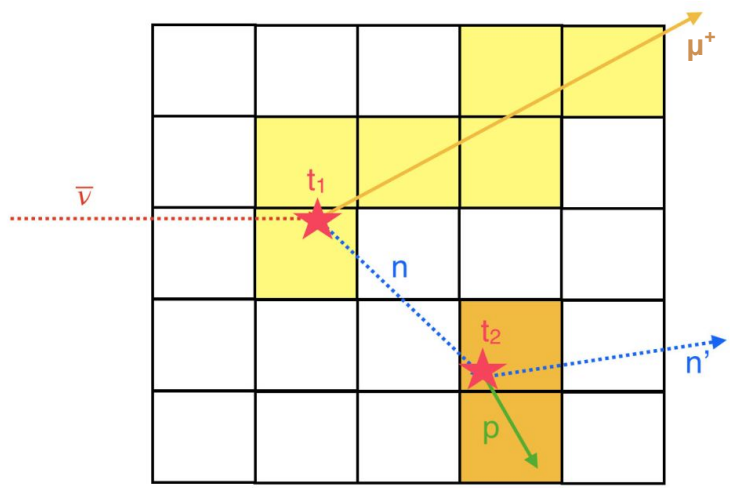


**Neutron reconstruction can recover the bias**

Neutrino energy reconstruction from lepton kinematics  
**Nuclear effects** are not considered



**Bias in the neutrino energy reconstruction**



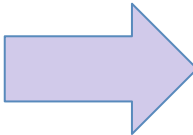
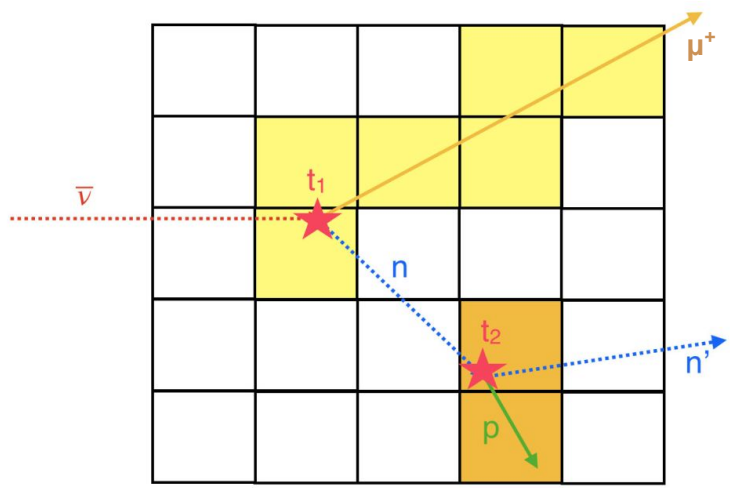
**Neutron energy reconstruction from time of flight t2 - t1 measurement**



Neutrino energy reconstruction from lepton kinematics  
**Nuclear effects** are not considered



**Bias in the neutrino energy reconstruction**



Neutron energy reconstruction from  
 time of flight  $t_2 - t_1$  measurement

**Timing characterization of the  
 detector is crucial!**

**Time calibration:  
time offsets + time walk**

**Time resolution**

**Goal:** evaluate

- ❑ **Channel time offsets** (bias in time ~ns given by electronic units)
- ❑ **Time walk** (effect ~10 ns depending on signal amplitude)

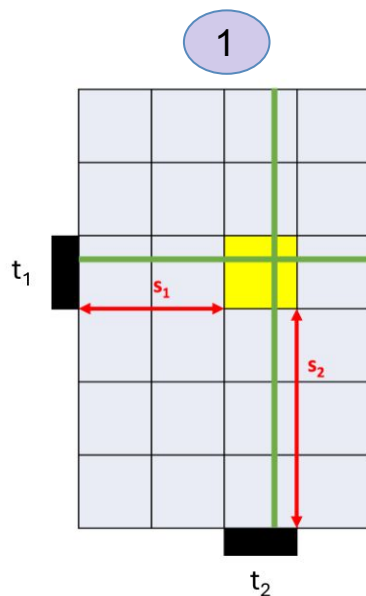
Important for:

- ❑ **Neutron analysis**
- ❑ **Track direction**

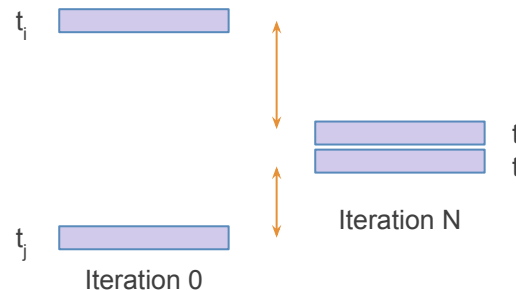
Exploit the **granularity** of the detector



- 1 Get a set of **matched hits**
- 2 Apply **iterative process** to evaluate channel offsets

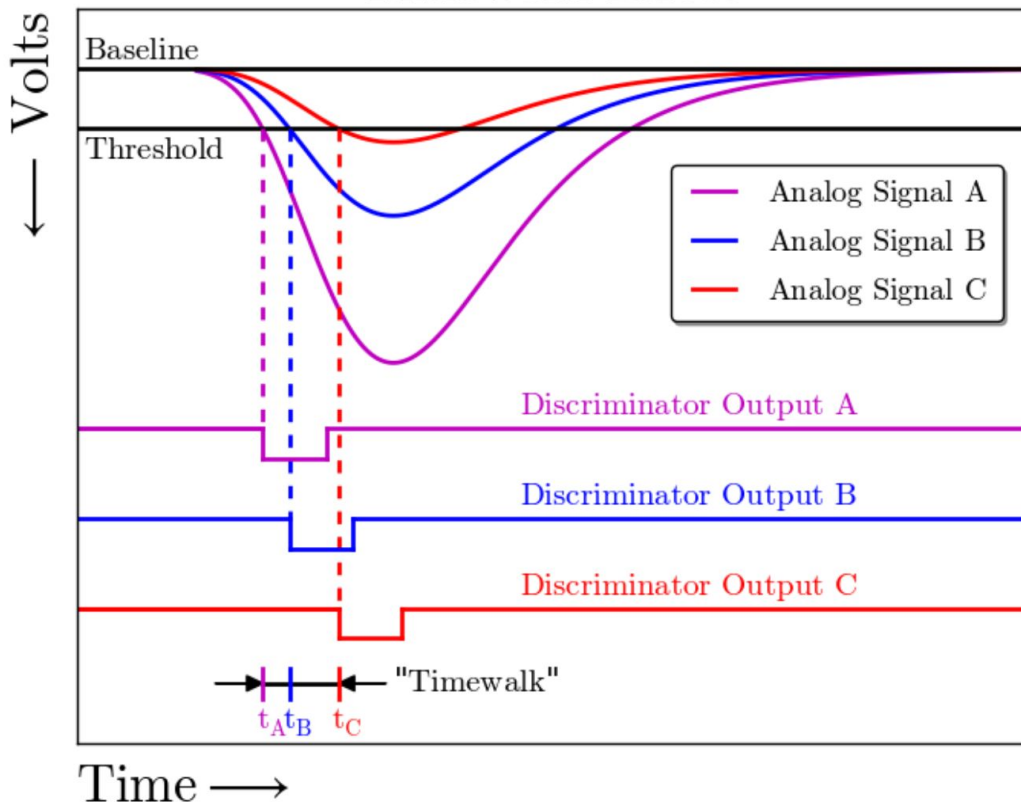


$$\text{Minimize: } \Delta t_{ij} = \frac{(t_i - t_j) - (s_i - s_j) / v}{2}$$



**Channel time offsets**

## Timewalk Effect



- ❑ The threshold is needed to cut the noise out
- ❑ Registered time depends on the time when the signal passes the threshold
- ❑ Very different amplitude means high time walk contribution
- ❑ By comparing similar amplitude the time walk is little

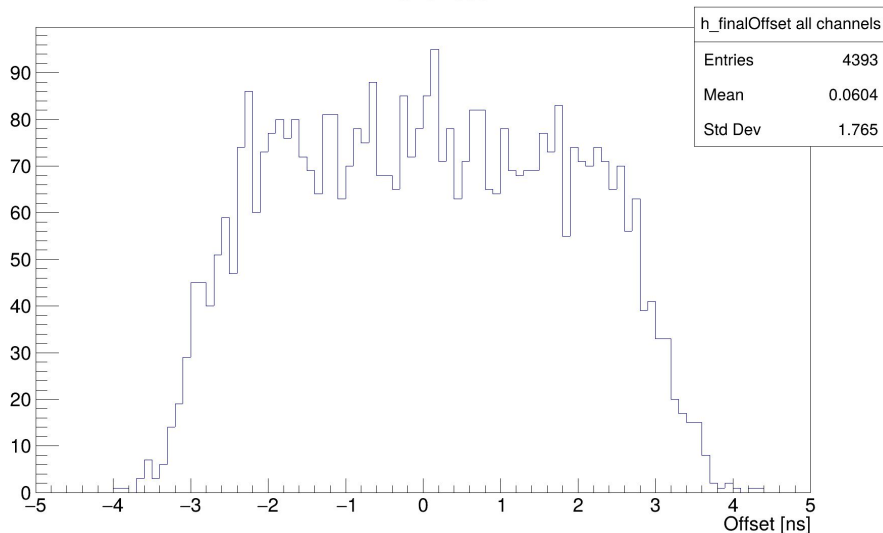


Channel offsets evaluated with matched hits with  $0.9 < q_1/q_2 < 1.11$  requirement

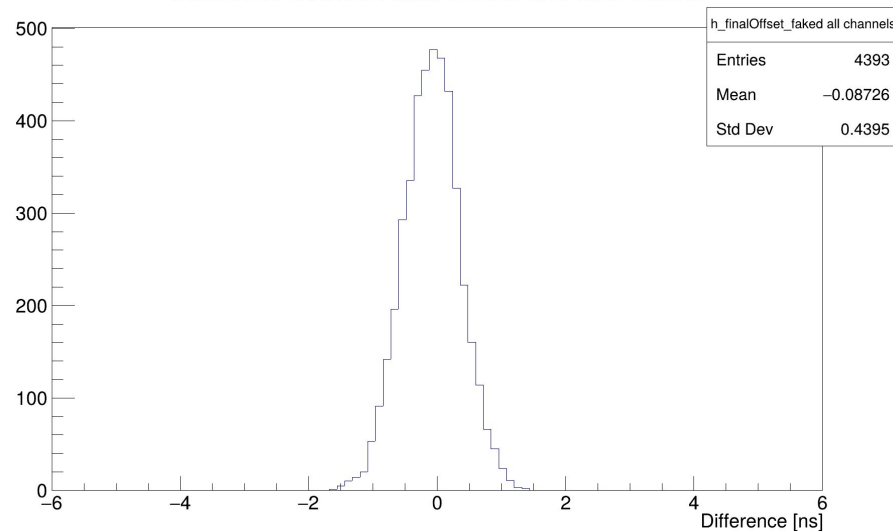
Run on 11k cosmic muons mc

Applied a **fake offset** between -3 and 3 ns

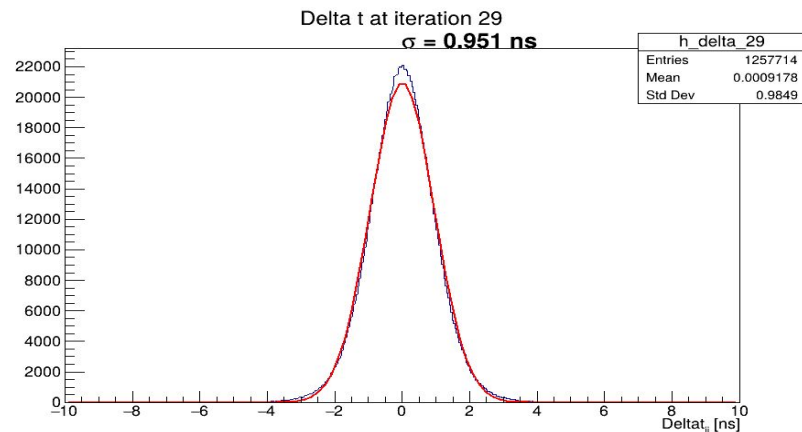
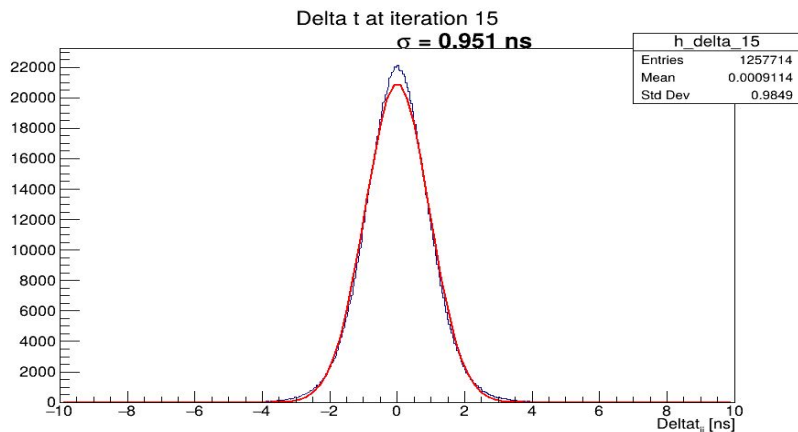
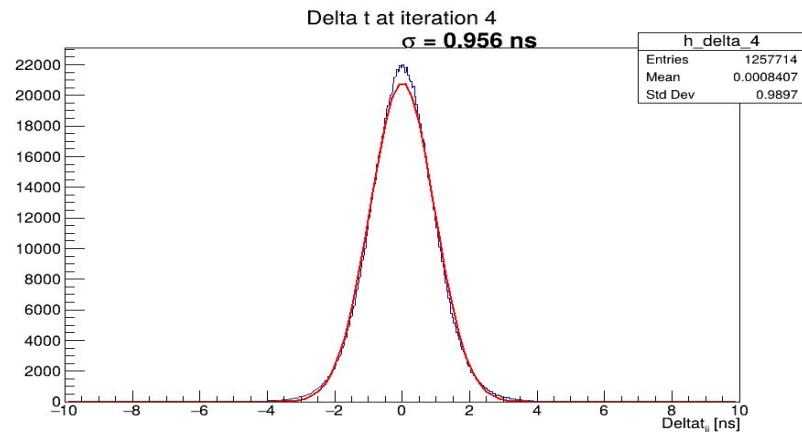
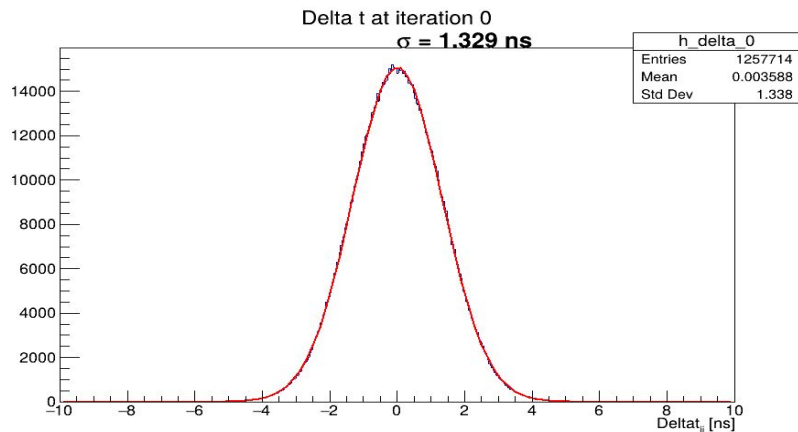
Final offset



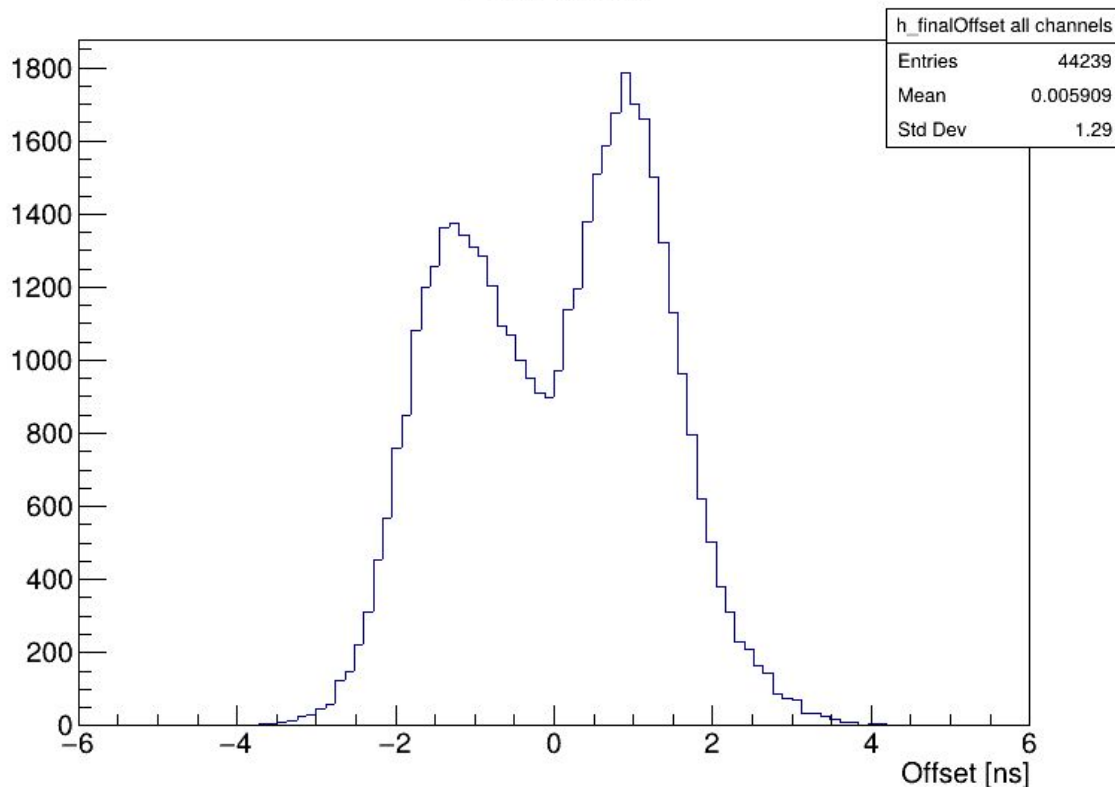
Difference between final offsets and fake offsets



**Algorithm works pretty well in MC**

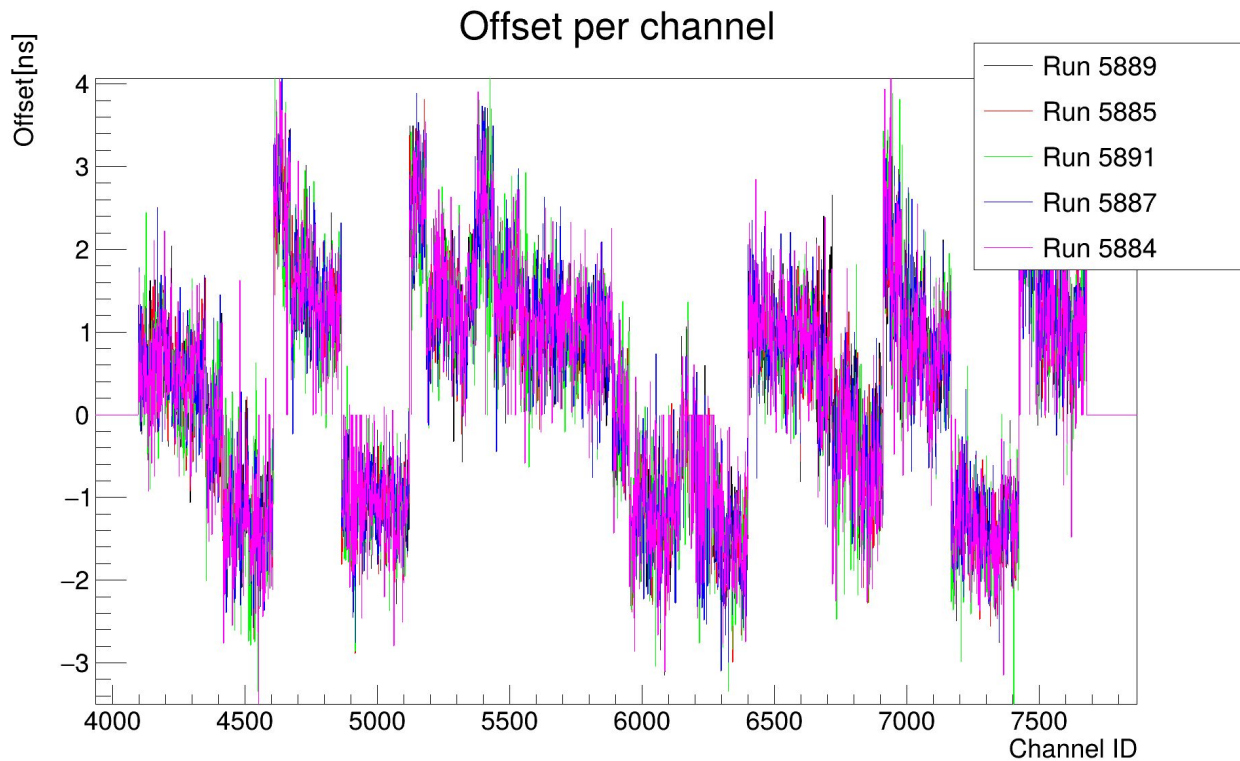


Final offset



- ❑ Only channels with more than 20 hits are saved
- ❑ 44k channels evaluated out of 58k (depend on stats)





**Good stability in data as well**

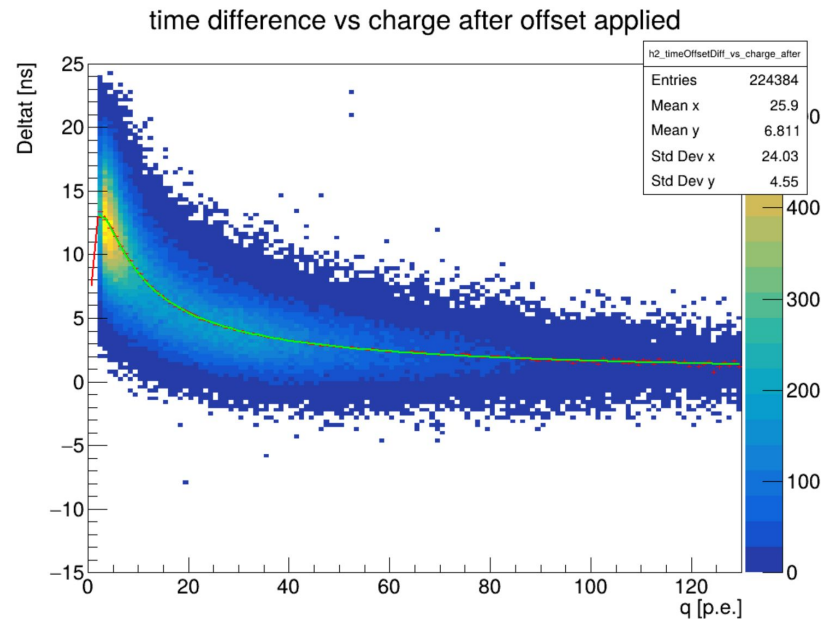
**Time walk correction**

- ❑ Apply the offset correction
- ❑ Check the difference between the expected times vs charge

$$\Delta t_{ij} = (t_i - t_j) - (s_i - s_j) / v$$

Fit function: 
$$\frac{A}{\log(B * q + C)} - \frac{D}{q + E}$$

Fit is not perfect but work in progress



**Time resolution**

For each pair of neighbouring cubes:



- ❑  $\Delta t_{reco}$ : reconstructed cube time difference
- ❑  $\Delta t_{calc}$ : calculated cube time difference from position
- ❑ Obtain time resolution  $\sigma_t$

$$t_{i+1} - t_i$$

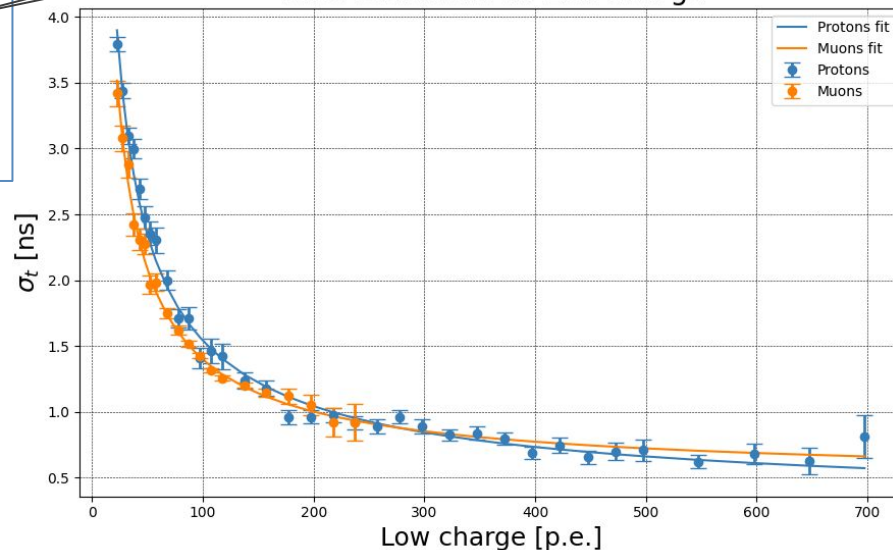
$$\text{Dist}(i+1, i)/\text{speed}$$

$$\sigma_t = \frac{\sigma_{st} (\Delta t_{reco} - \Delta t_{calc})}{\sqrt{2}}$$

For each pair of neighbouring cubes:

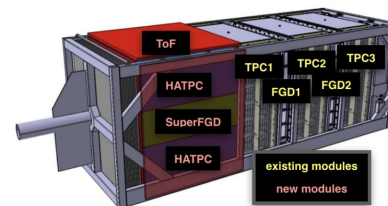
- ❑ Build  $(\Delta t_{reco} - \Delta t_{calc})$  distribution
- ❑ **Low charge** from hit charges  $q_i$  and  $q_{i+1}$
- ❑ Scan  $\sigma_t$  through low charge

Time resolution vs Low charge

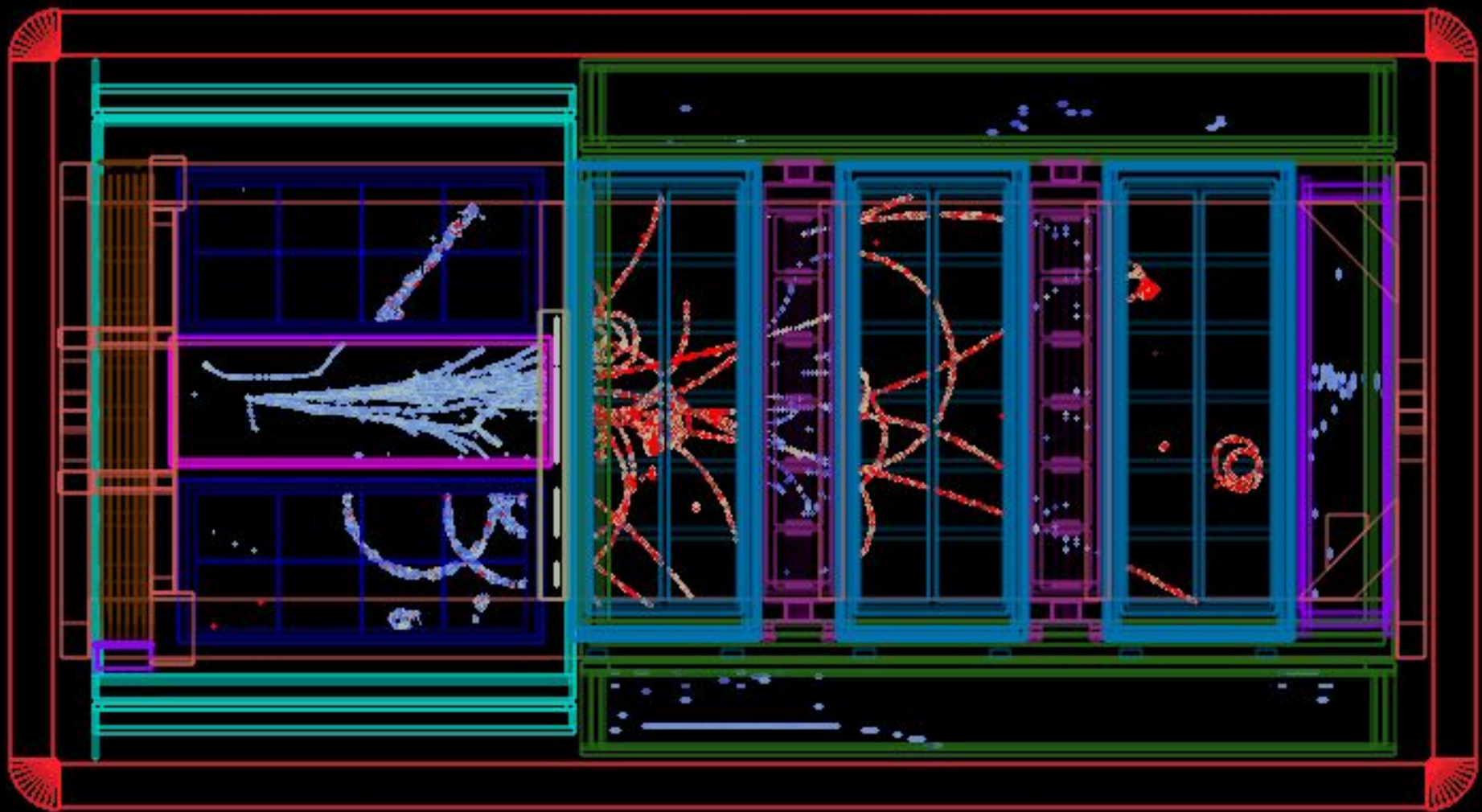


Fit function:  $\sigma(A) = p_0/A^{p_1} + p_2$

- ❑ **ND280 Upgrade** for systematic errors reduction
  - ❑ Neutrons will play a key role in systematic errors
  - ❑ Neutron reconstruction relies on time of flight of neutrons
- 
- ❑ **Time calibration** on going
  - ❑ **Time resolution** only tested in MC
  - ❑ Given a methodology for **time calibration for high granularity detector**



**Thanks for your attention!**



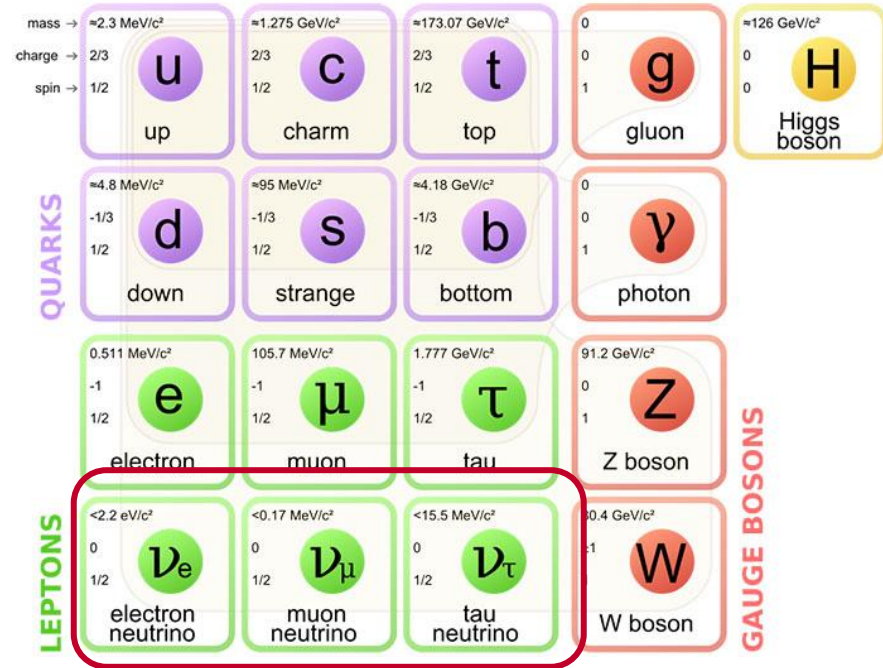
**Backup**



- ❑ Neutrinos  $\nu$  belong to the Standard Model of particle physics
- ❑ Chargeless leptons
- ❑ Treated as **massless particles** (wrongly)



Homestake first (of many) experiment to observe **neutrino oscillations** proves this wrong



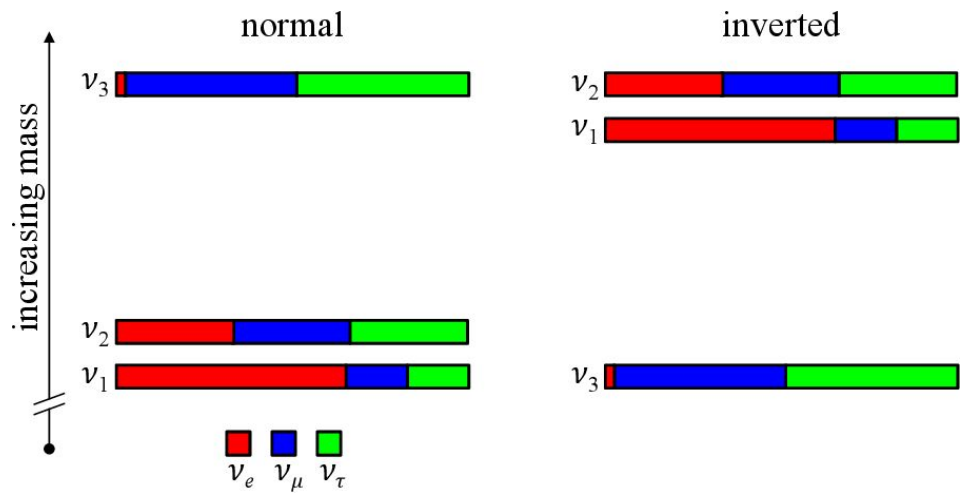
$$\begin{pmatrix} \nu_e \\ \nu_\mu \\ \nu_\tau \end{pmatrix} = \begin{pmatrix} \text{PMNS} \\ \text{matrix} \end{pmatrix} \begin{pmatrix} \nu_1 \\ \nu_2 \\ \nu_3 \end{pmatrix}$$

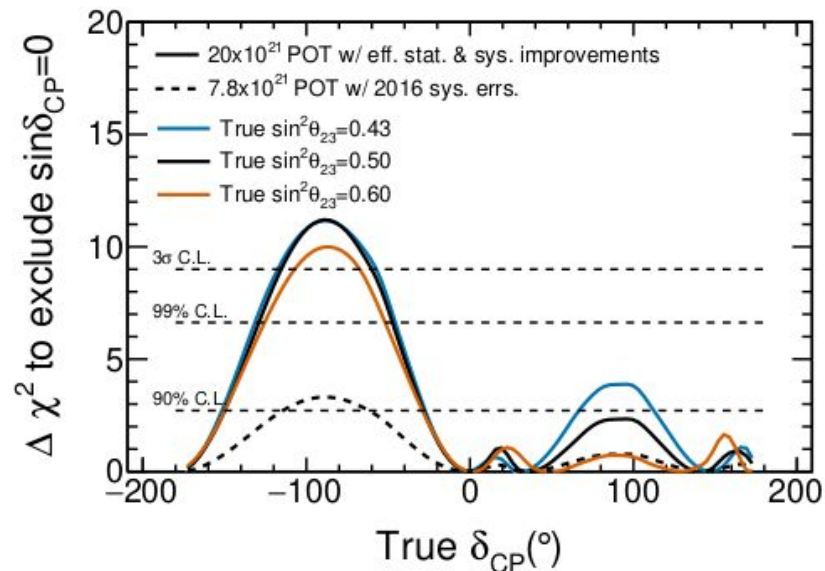
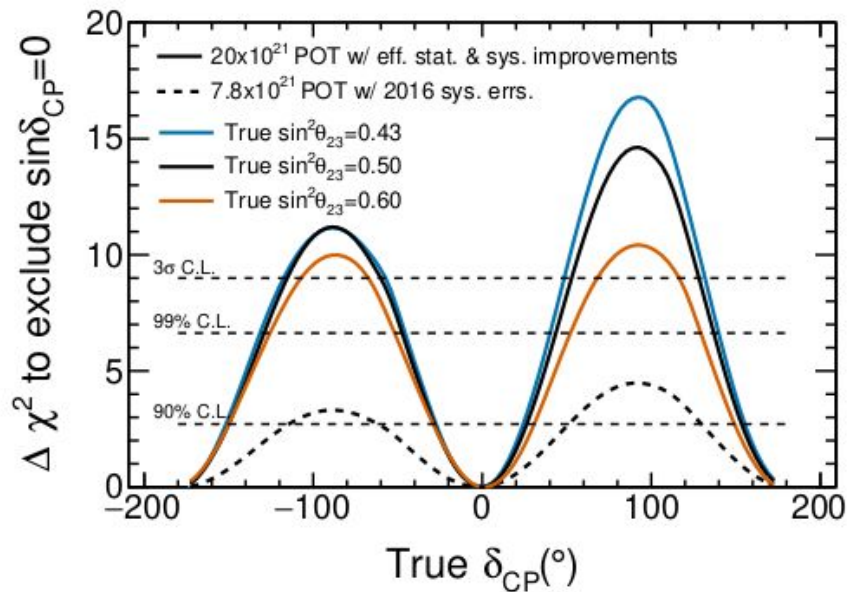
$$U = \begin{pmatrix} 1 & 0 & 0 \\ 0 & c_{23} & s_{23} \\ 0 & -s_{23} & c_{23} \end{pmatrix} \cdot \begin{pmatrix} c_{13} & 0 & s_{13}e^{-i\delta_{CP}} \\ 0 & 1 & 0 \\ -s_{13}e^{i\delta_{CP}} & 0 & c_{13} \end{pmatrix} \cdot \begin{pmatrix} c_{12} & s_{12} & 0 \\ -s_{12} & c_{12} & 0 \\ 0 & 0 & 1 \end{pmatrix}$$

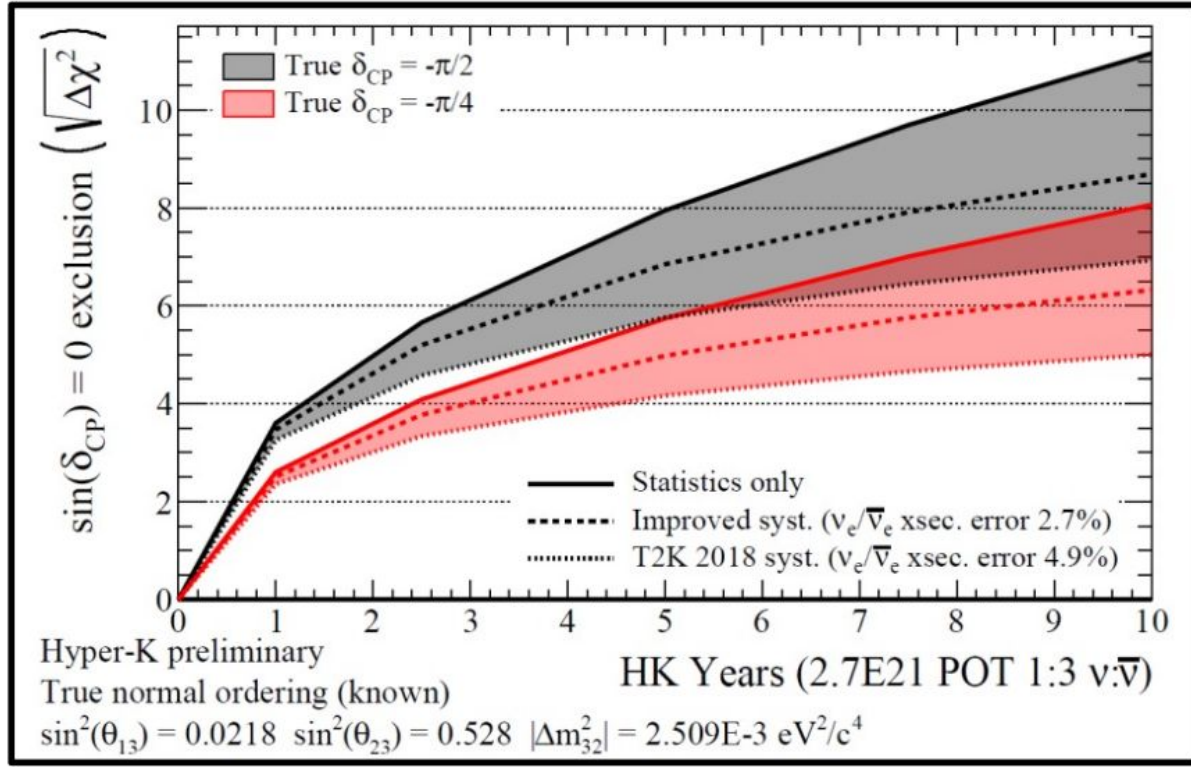
$$= \begin{pmatrix} c_{12}c_{13} & s_{12}c_{13} & s_{13}e^{-i\delta_{CP}} \\ -s_{12}c_{23} - c_{12}s_{23}s_{13}e^{i\delta_{CP}} & c_{12}c_{23} - s_{12}s_{23}s_{13}e^{i\delta_{CP}} & s_{23}c_{13} \\ s_{12}s_{23} - c_{12}c_{23}s_{13}e^{i\delta_{CP}} & -c_{12}s_{23} - s_{12}c_{23}s_{13}e^{i\delta_{CP}} & c_{23}c_{13} \end{pmatrix}$$

$$P_{\alpha \rightarrow \beta, \alpha \neq \beta} = \sin^2(2\theta) \sin^2 \left( 1.27 \frac{\Delta m^2 L}{E} \frac{[\text{eV}^2] [\text{km}]}{[\text{GeV}]} \right)$$

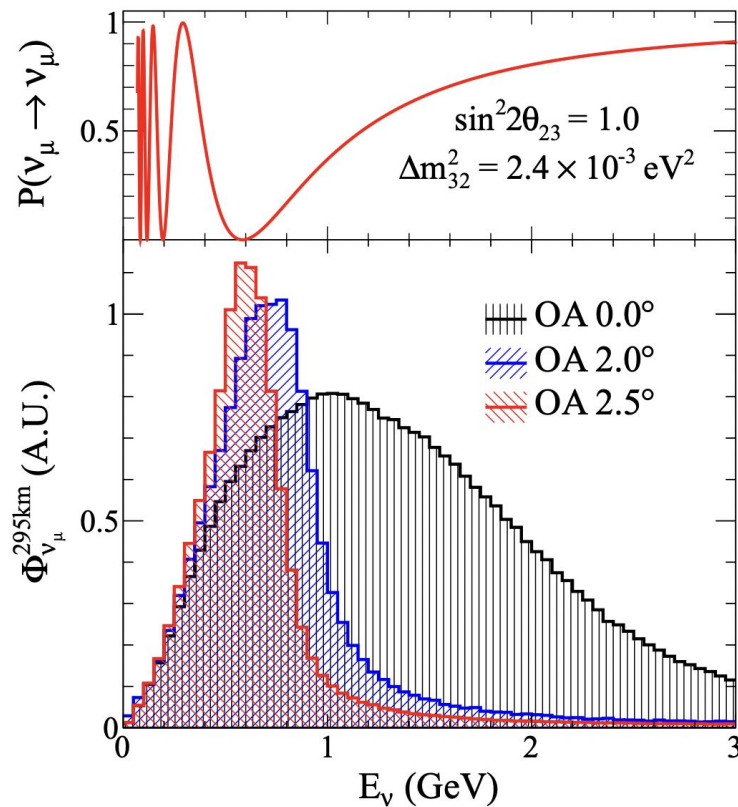
- Open questions**
- CP violation phase
  - Neutrino Mass ordering
  - $\theta_{23}$  octant





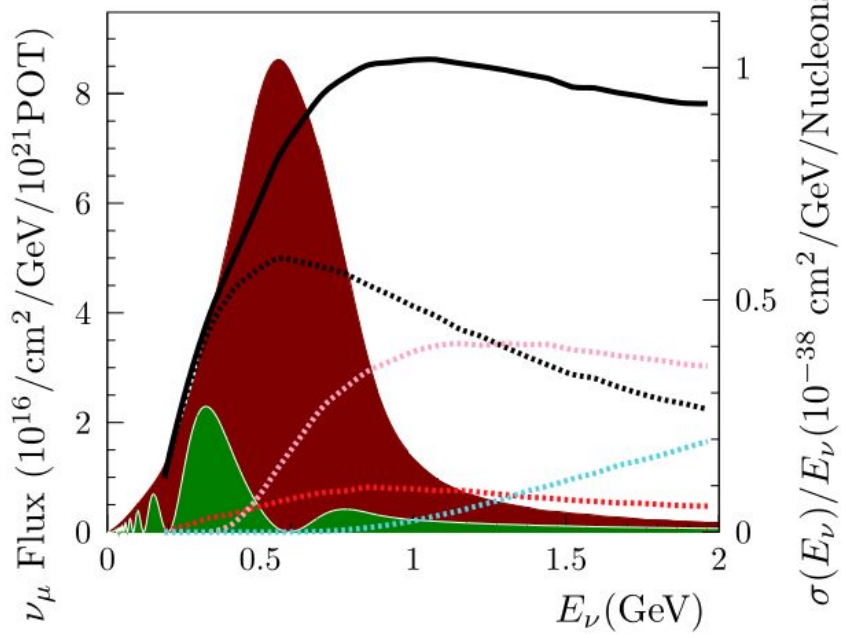


|   | Normal Ordering (best fit)      |                               | Inverted Ordering ( $\Delta\chi^2 = 7.1$ ) |                               |
|---|---------------------------------|-------------------------------|--|-------------------------------|
|   | bfp $\pm 1\sigma$               | $3\sigma$ range               | bfp $\pm 1\sigma$                          | $3\sigma$ range               |
| $\sin^2 \theta_{12}$                              | $0.304^{+0.012}_{-0.012}$       | $0.269 \rightarrow 0.343$     | $0.304^{+0.013}_{-0.012}$                  | $0.269 \rightarrow 0.343$     |
| $\theta_{12}/^\circ$                              | $33.44^{+0.77}_{-0.74}$         | $31.27 \rightarrow 35.86$     | $33.45^{+0.78}_{-0.75}$                    | $31.27 \rightarrow 35.87$     |
| $\sin^2 \theta_{23}$                              | $0.573^{+0.016}_{-0.020}$       | $0.415 \rightarrow 0.616$     | $0.575^{+0.016}_{-0.019}$                  | $0.419 \rightarrow 0.617$     |
| $\theta_{23}/^\circ$                              | $49.2^{+0.9}_{-1.2}$            | $40.1 \rightarrow 51.7$       | $49.3^{+0.9}_{-1.1}$                       | $40.3 \rightarrow 51.8$       |
| $\sin^2 \theta_{13}$                              | $0.02219^{+0.00062}_{-0.00063}$ | $0.02032 \rightarrow 0.02410$ | $0.02238^{+0.00063}_{-0.00062}$            | $0.02052 \rightarrow 0.02428$ |
| $\theta_{13}/^\circ$                              | $8.57^{+0.12}_{-0.12}$          | $8.20 \rightarrow 8.93$       | $8.60^{+0.12}_{-0.12}$                     | $8.24 \rightarrow 8.96$       |
| $\delta_{CP}/^\circ$                              | $197^{+27}_{-24}$               | $120 \rightarrow 369$         | $282^{+26}_{-30}$                          | $193 \rightarrow 352$         |
| $\frac{\Delta m_{21}^2}{10^{-5} \text{ eV}^2}$    | $7.42^{+0.21}_{-0.20}$          | $6.82 \rightarrow 8.04$       | $7.42^{+0.21}_{-0.20}$                     | $6.82 \rightarrow 8.04$       |
| $\frac{\Delta m_{3\ell}^2}{10^{-3} \text{ eV}^2}$ | $+2.517^{+0.026}_{-0.028}$      | $+2.435 \rightarrow +2.598$   | $-2.498^{+0.028}_{-0.028}$                 | $-2.581 \rightarrow -2.414$   |

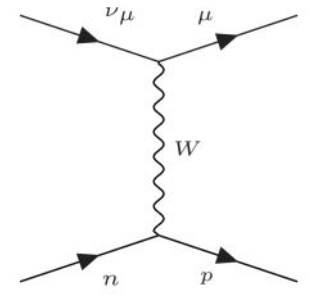


$\nu$ -mode beam  
■ ND280  
■ SK Osc. ( $\times 10^6$ )

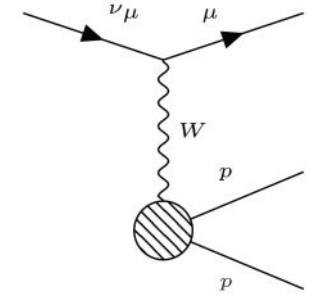
NEUT,  $\nu_\mu - {}^{12}\text{C}$   
 — CC-INC    ..... CC-1p1h  
 ..... CC-2p2h    ..... CC-SPP  
 ..... CC-DIS



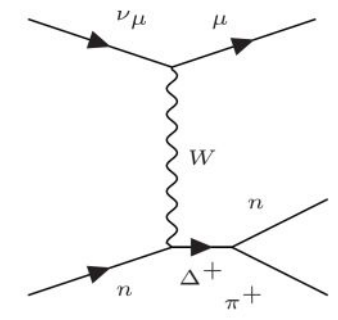
**CC-QE**  
(Charged-Current Quasi-Elastic)



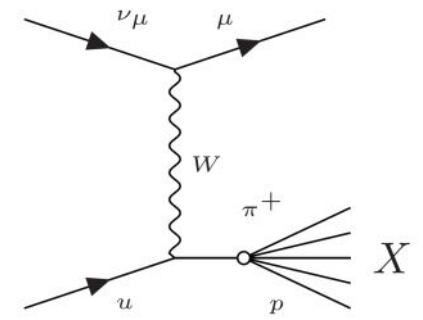
**CC-2p2h**  
(Two-Particle-Two-Hole)



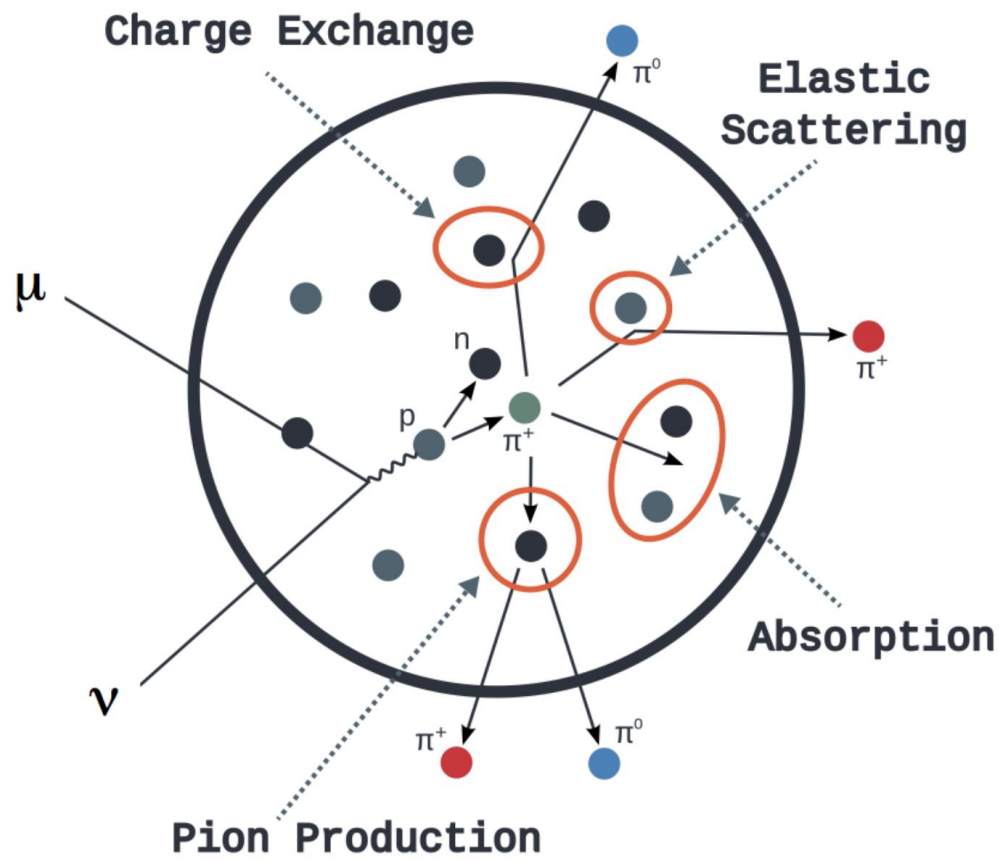
**CC-SPP**  
(Single Pion Production)



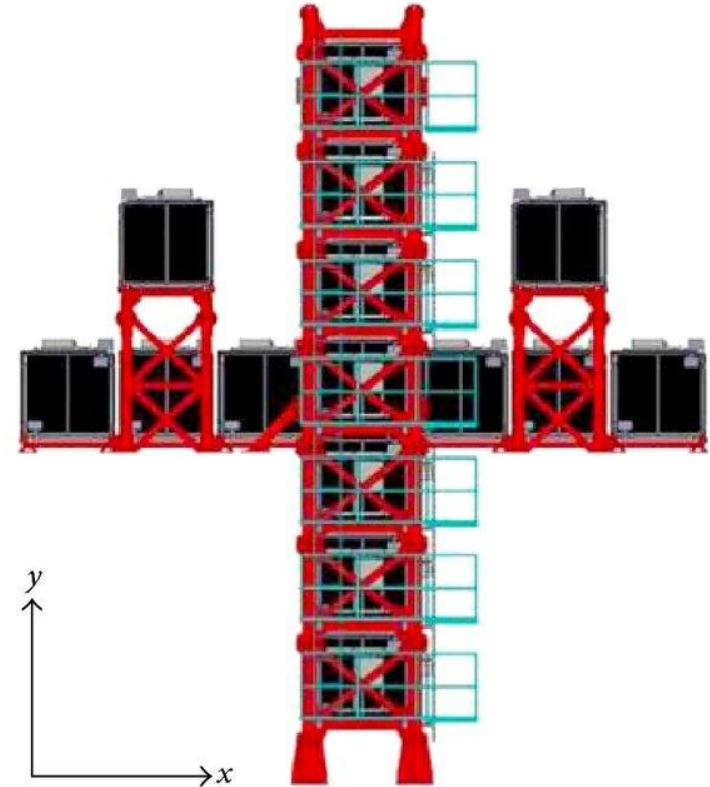
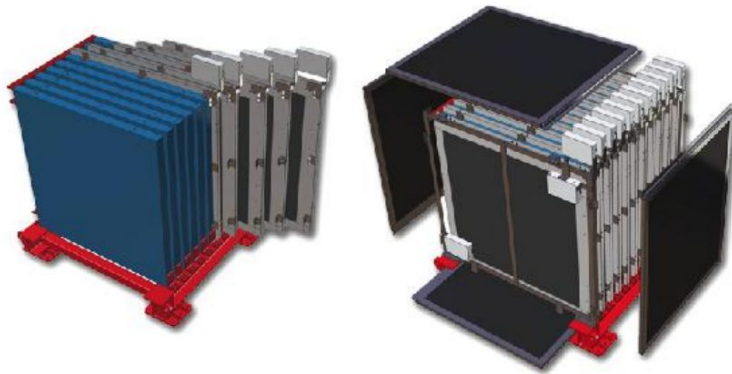
**CC-DIS**  
(Deep Inelastic Scattering)

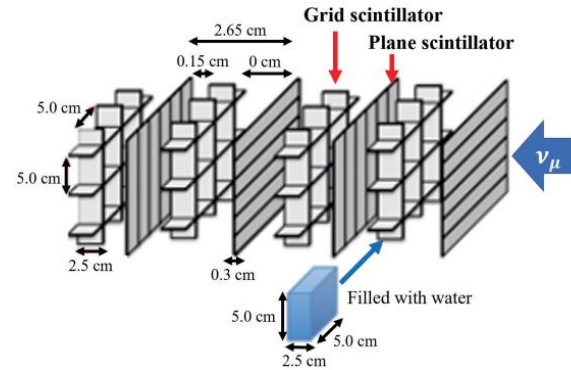






- ❑ On-axis
- ❑ 16 modules
- ❑ 9 iron target plates + 11 tracking scintillator planes
- ❑ Surrounded by scintillator veto planes
- ❑ Flux direction monitoring





- ❑ 1.5° off-axis
- ❑ Reduce systematic errors

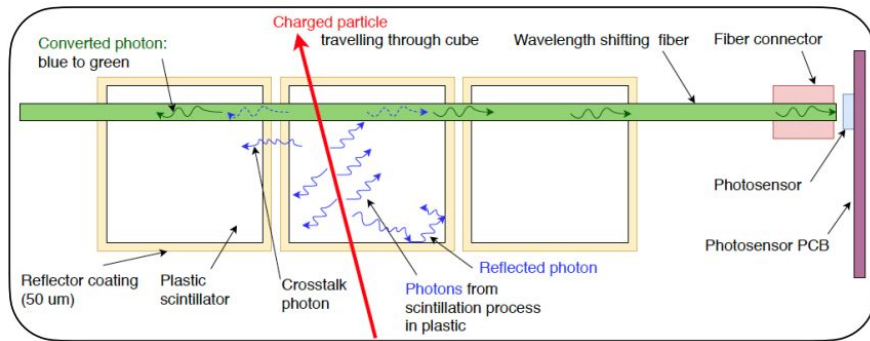
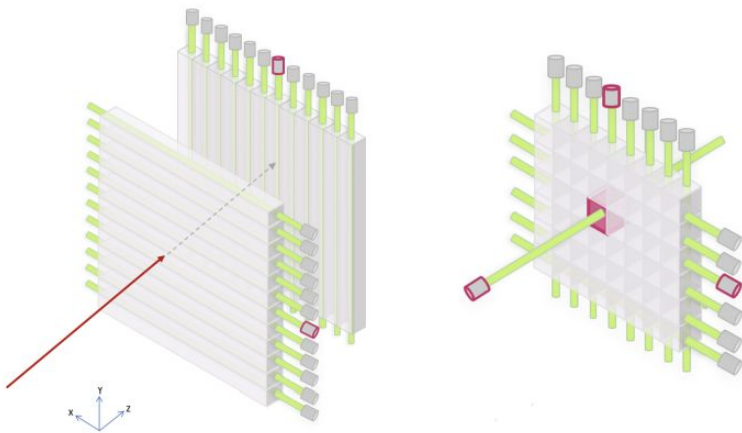
## WAGASCI

- ❑ Main target: pure water
- ❑ Cross section ratio scintillator-water

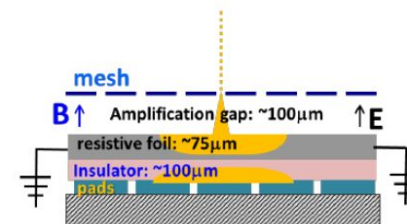
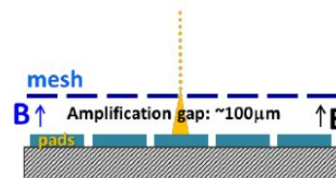
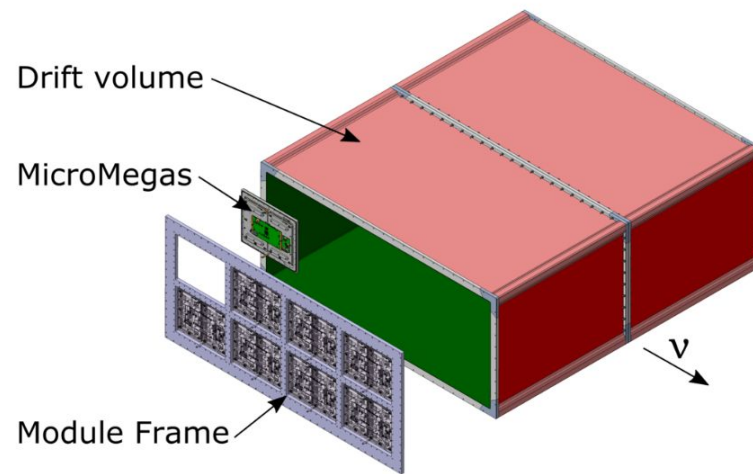
## Baby MIND

- ❑ Magnetized iron neutrino detector
- ❑ Charge and momentum of outgoing muon

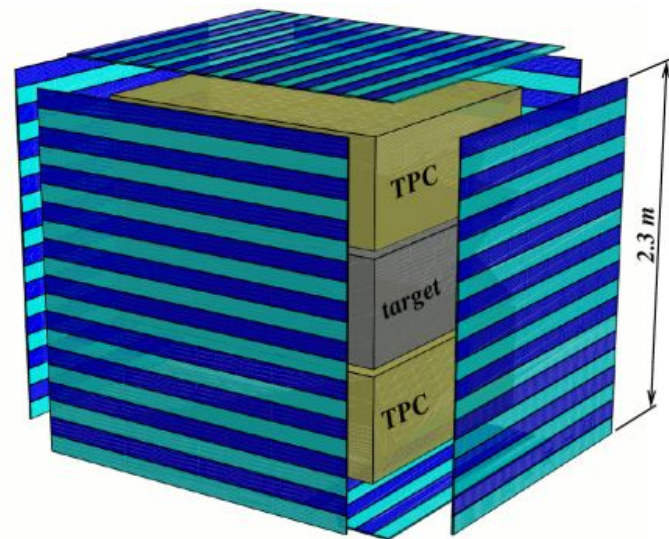
- ❑ Polystyrene doped with 1.5% of paraterphenyl (PTP) and 0.01% of POPOP
- ❑ Hamamatsu MPPC S13081-050C
- ❑ L.Y. around 40 p.e. for MIP



- ❑ Similar to old TPCs
- ❑ Resistive MicroMegas
- ❑ Thin field cage
- ❑ Better spatial resolution
- ❑ Reduced risk of sparks



- ❑ 6 modules of plastic scintillator layers
- ❑ 150 ps time resolution
- ❑ Background discrimination
- ❑ Ingoing and outgoing particle identification



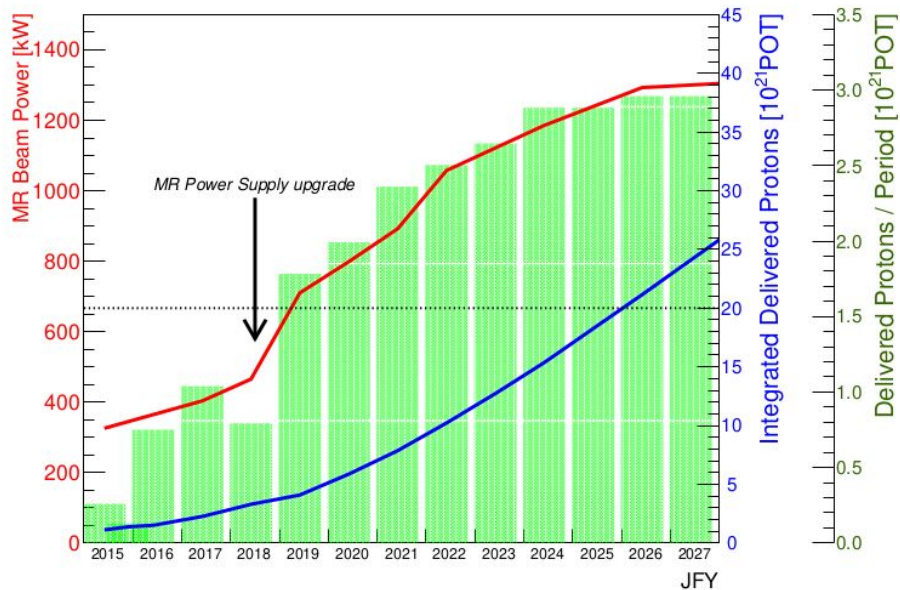
**Aim:** CP violation with  $> 3\sigma$  significance exclusion

- ❑ Extended T2K running time (Hyper-Kamiokande)
- ❑ Collect  $20 \times 10^{21}$  Protons-on-Target (PoT)
- ❑ J-PARC **accelerator upgrade reaching 1.3 MW** in few years (485 kW currently achieved)

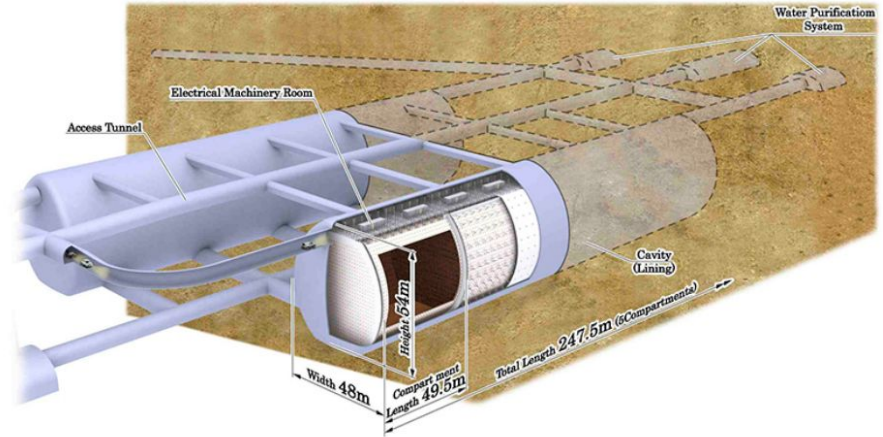
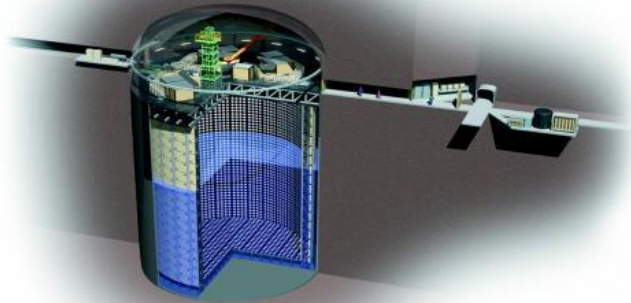
**Better understanding of detector uncertainties is crucial!**



**ND280 UPGRADE**





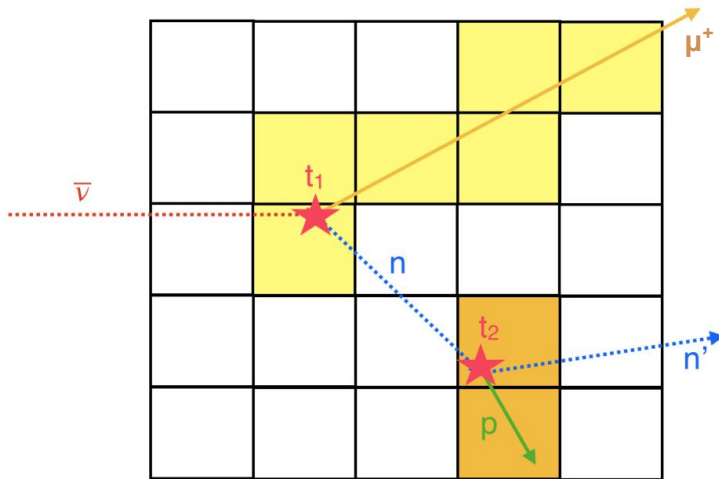


- ❑ 41.4 m tall, 39.3 m diameter
- ❑ 50k metric tons
- ❑ 11'000 PMTs

- ❑ 48 x 54 x 250 m<sup>3</sup>
- ❑ 1 million metric tons
- ❑ 99'000 PMTs

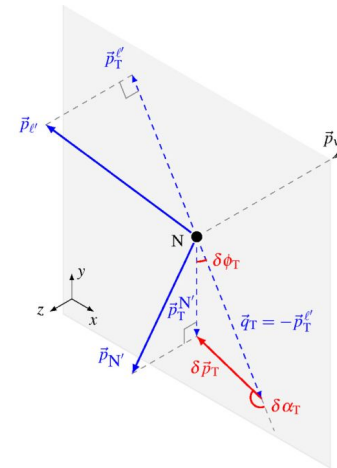
Neutrino energy reconstruction from lepton kinematics  
**Nuclear effects** (FSI) are not considered

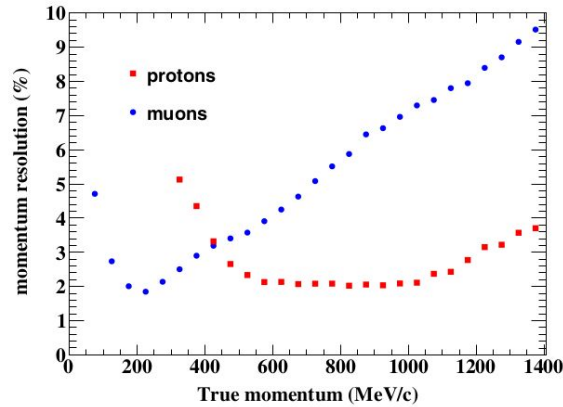
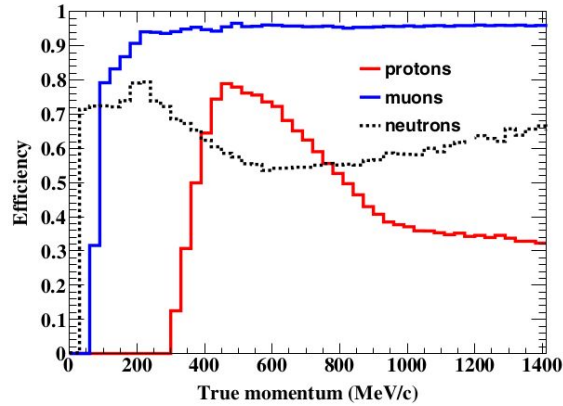
Transverse momentum imbalance  $\delta p_T$



$$E_\nu = \frac{m_n^2 - m_p^2 - m_\mu^2 + 2m_p E_\mu}{2(m_p - E_\mu + p_\mu \cos \theta_\mu)}$$

$$\bar{\nu}_\mu + p \rightarrow \mu^+ + n \quad \delta p_T = |\vec{p}_T^l + \vec{p}_T^n|$$





SFGD time calibration doesn't take into account **channel offsets** and **time walk correction**

**Goal:** time offsets and time walk evaluation

Important for:

- ❑ **Neutron analysis**
- ❑ **Track direction**

We define:

- ❑ **fiber hit:** a THit with time, charge and MPPC position
- ❑ **matchedHit:** a match between 2 fiber hits
- ❑ **t<sub>i</sub>:** the time recorded by the i<sub>th</sub> channel
- ❑ **s<sub>i</sub>:** distance from the MPPC to the reconstructed cube in one of the 2 hits matched
- ❑ **v:** speed of light in the fiber

Offset ~ns given by the electronic components

Effect ~10ns depending on signal amplitude

Exploit the **granularity** of the detector



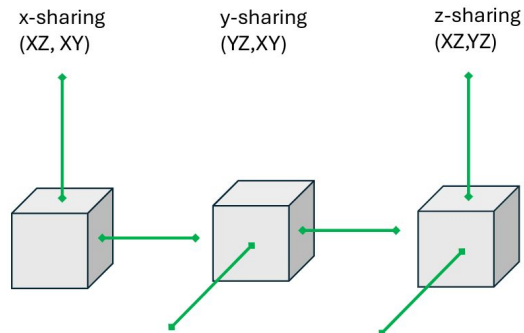
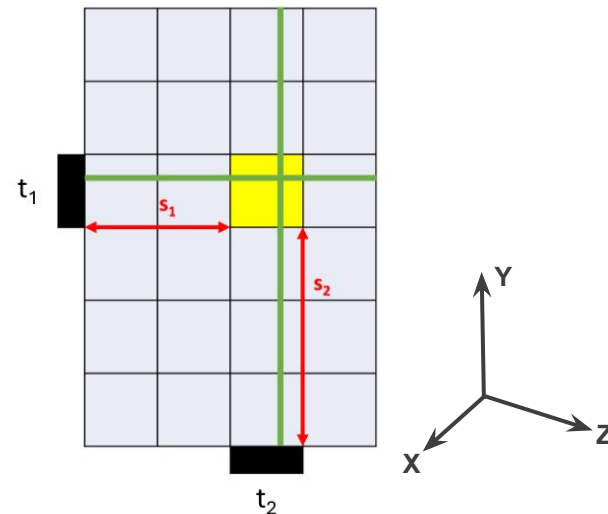
- ❑ Get a set of **matched hits**
- ❑ Apply **iterative process** to evaluate channel offsets

Using **cosmics** ( $\mu$ ) in both MonteCarlo and Data:

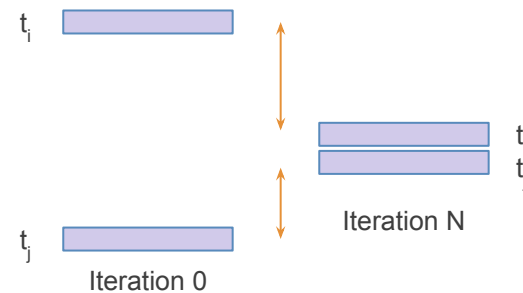
- ❑ Each event has a single track
- ❑ Each track is reconstructed as a set of **3D cube hits**
- ❑ Each cube hit has **3 fiber hits**, meaning **3 matched hits**
- ❑ **Noise cut:** 10 photoelectrons (p.e.)



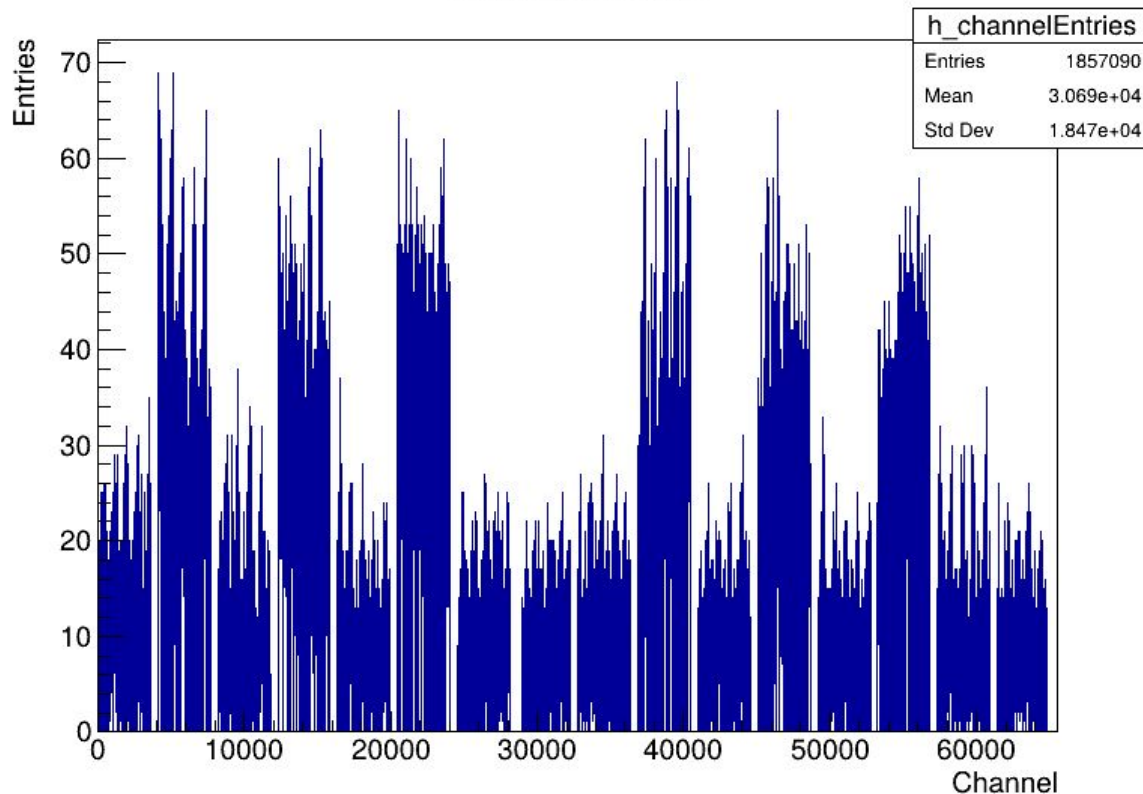
End up with a set of vectors of:  $[(t,s,ch,q)_1, (t,s,ch,q)_2]$



- ❑ Loop through matched hits
- ❑ For each matched hit:
  - ❑ Evaluate:  $\Delta t_{ij} = \frac{(t_i - t_j) - (s_i - s_j)/v}{2}$  dividing by 2 necessary to distribute the correction equally to each channel of the matched hit
  - ❑ Evaluate:  $\Delta t_i = \sum_j \frac{\Delta t_{ij}}{N}$  which is the **correction on the offset**
  - ❑ Apply the correction:  $t_i^{n+1} = t_i^n \pm \Delta t_i$
- ❑ Repeat loop until convergence: maximum number of iteration
- ❑ Final channel offset is the sum of all the corrections at each iteration
- ❑ In the end the two times should be as close as possible

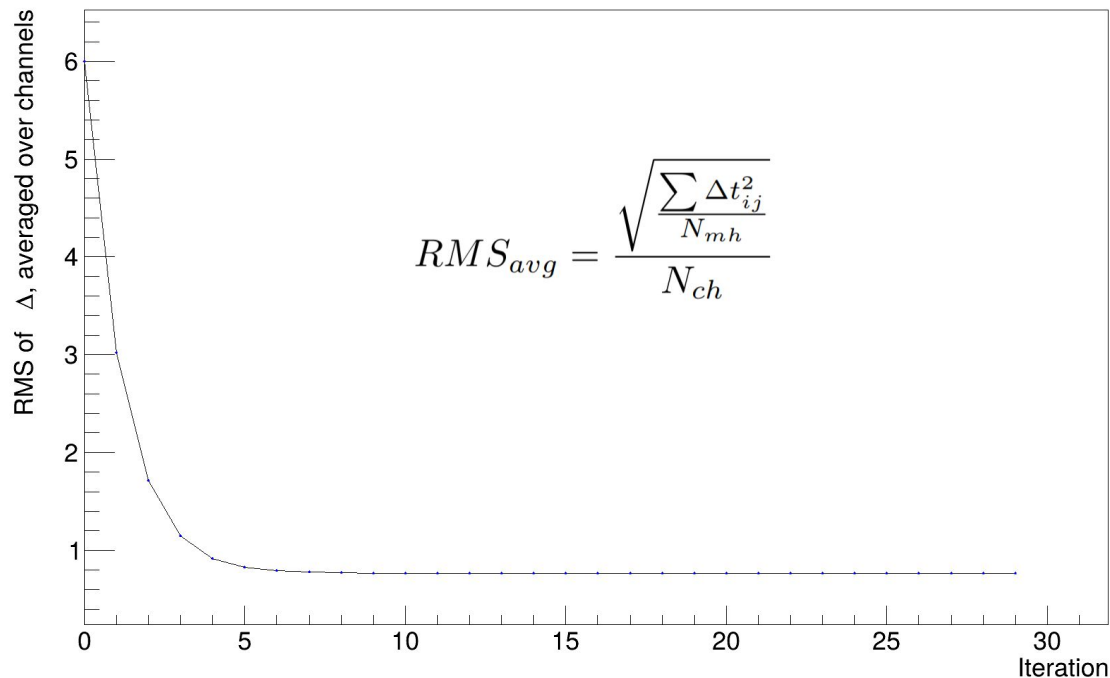


## Channel entries

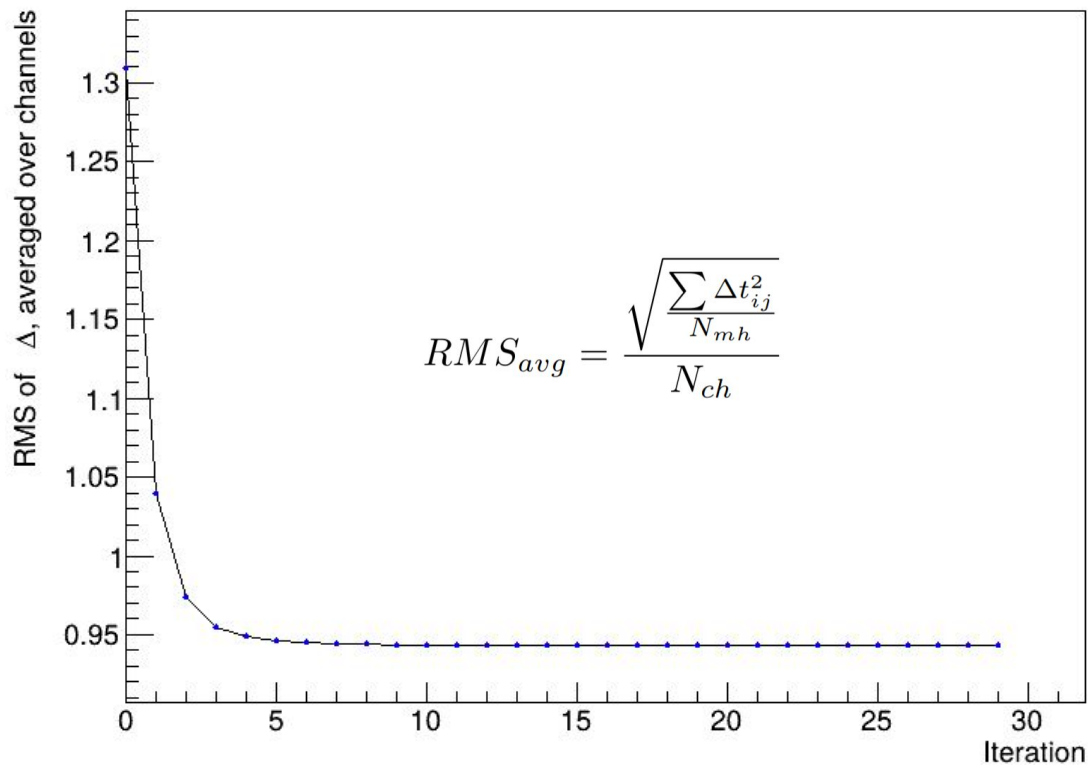


Run on 11k cosmic mc

$RMS_{avg}$  is a metric to the convergence of the algorithm

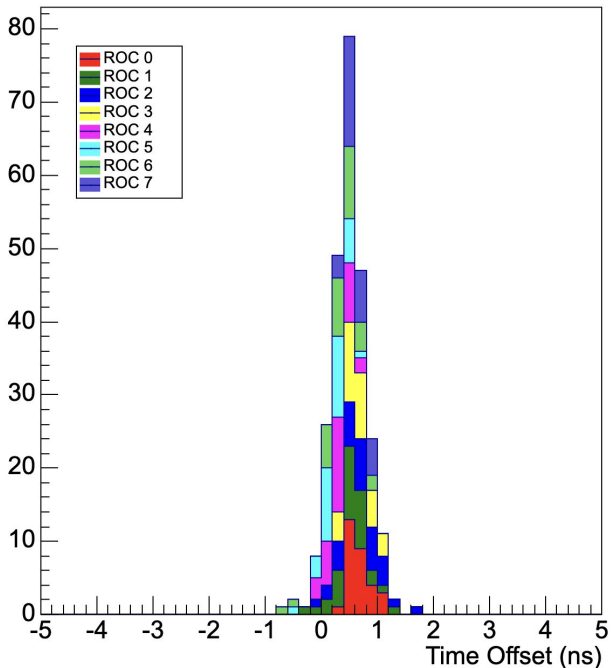






For the “unseen” channels we could use the citiroc’s average offset

FEB 16 Offset Distribution per ROC



ROC 1

