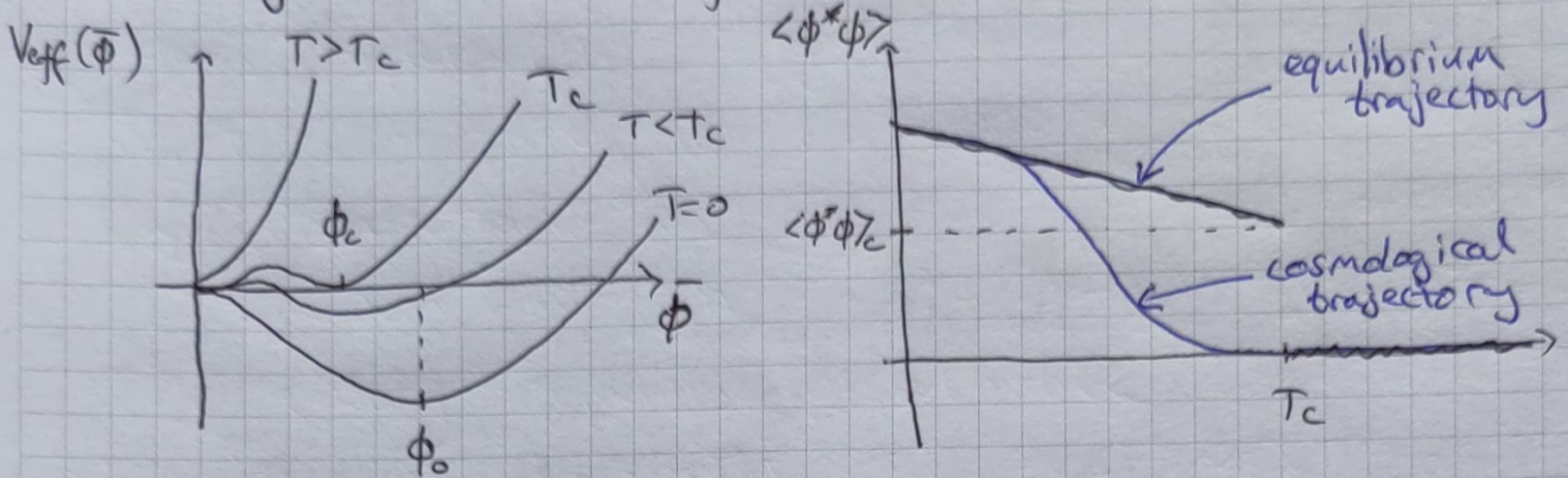


Phase transitions III

Cosmological evolution of a 1st order phase transition



Transition mechanism:

- Transition for volume L^3 suppressed by $e^{-V_{\text{eff}}(\bar{\phi}_{\text{barrier}}) L^3}$
- ⇒ must be local
- Thermal fluctuation over barrier
- Quantum tunnelling through barrier

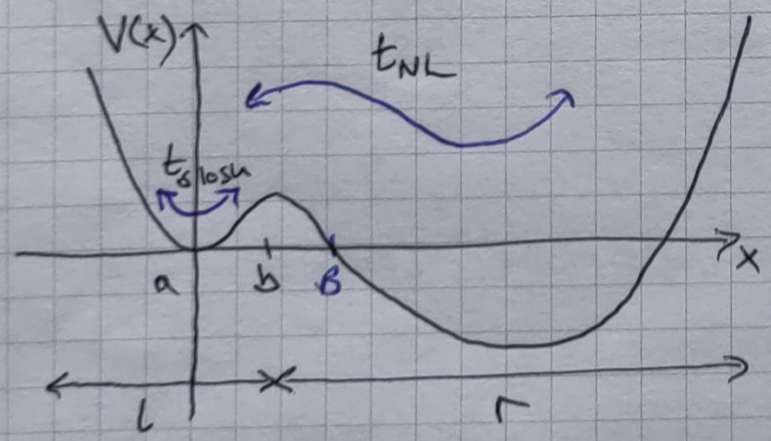
For cosmology, the transition rate must be faster than Hubble

$\Gamma \gtrsim H^4 \Rightarrow \gtrsim \text{once per Hubble patch}$

 (Transition rate per unit volume)

Need to understand these rates, Γ .

Particle escape rate



Start at 'a' in metastable well. How long until $x > b$?

Assume, $\rho(x > b, p) = 0$
 $\rho(x = b, p < 0) = 0$

$t_{\text{slash}} \ll \Gamma^{-1} \ll t_{\text{NL}}$

 $\rho(x < b, p) \propto e^{-\beta H}$

Classical, thermal

$H = \frac{p^2}{2m} + V(x)$

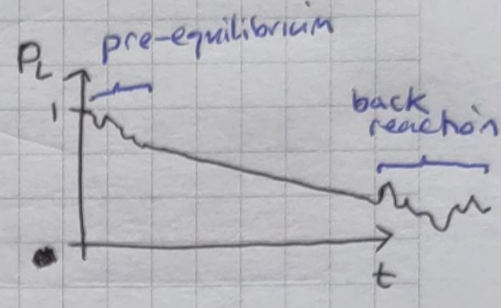
Liouville's equation for probability density:

$$\frac{\partial \rho}{\partial t} + \frac{\partial}{\partial x} \left(\frac{p}{m} \rho \right) + \frac{\partial}{\partial p} \left(-V'(x) \rho \right) = \frac{\partial \rho}{\partial t} + \nabla \cdot \mathbf{J} = 0$$

$J_x = \dot{x} \rho$ $J_p = \dot{p} \rho$

Define rate as:

$$\Gamma \equiv \lim_{\substack{t/t_{\text{slush}} \rightarrow \infty \\ t/t_{NL} \rightarrow 0}} \frac{1}{P_L(t)} \frac{dP_L(t)}{dt}$$



And as $P_r = 1 - P_L$,
 $\Rightarrow \dot{P}_L = -\Gamma P_L \Rightarrow P_L(t) = e^{-\Gamma t}$

$$\begin{aligned} \frac{dP_r}{dt} &= \frac{d}{dt} \int_r dx dp \rho(x,p) = - \int_r dx dp \nabla \cdot \mathbf{J} \\ &= \int_{x=b} J \cdot \hat{x} = \int dp \cdot \frac{p}{m} N e^{-\beta H(b,p)} \theta(p) \\ &= \left[-\Gamma N e^{-\beta \left(\frac{p^2}{2m} + V(b) \right)} \right]_{p=0}^{p=\infty} = \Gamma N e^{-\beta V(b)} \\ P_L &= \int dx dp N e^{-\beta H} \approx \int dx dp N e^{-\beta \left(\frac{p^2}{2m} + \frac{1}{2} m \omega_a^2 x^2 \right)} \\ &= \sqrt{2\pi m T} \sqrt{\frac{2\pi T}{m \omega_a^2}} N = \frac{2\pi T N}{\omega_a} \end{aligned}$$

$$\Rightarrow \Gamma = \underbrace{\frac{\omega_a}{2\pi}}_{\text{frequency of 'attempts'}} \underbrace{e^{-\beta V(b)}}_{\text{Boltzmann suppression}}$$

This assumes Hamiltonian evolution. For Langevin evolution (with noise and damping) the result is similar, but the 'frequency of attempts' is reduced by the damping, γ ; Kramers (1940) Physica 7 284.

$$\Gamma = \left(\frac{\sqrt{\omega_b^2 + \gamma^2/4} - \gamma/2}{\omega_b} \right) \cdot \frac{\omega_a}{2\pi} e^{-\beta V(b)}$$

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Now let's try the quantum case ($T=0$).

$$P_R(t) = \int_R dx_f |\langle x_f, t | \alpha, 0 \rangle|^2$$

initial state
final state

$D_F(\alpha, 0; x_f, t)$

To transfer easily to field theory, we develop a path integral approach,

$$D_F(\alpha, 0; x_f, t) = \int_{\substack{x(0)=\alpha \\ x(t)=x_f}} Dx e^{iS[x]}$$

When $t \ll t_{NL}$ we don't expect the evolution for $x > B$ to affect the rate, so let's factor it out

$$D_F(\alpha, 0; x_f, t) = \int_0^t dt_0 \int_{\substack{x(0)=\alpha \\ x(t_0)=B}} Dx \delta(t_B[x] - t_0) e^{iS[x]} \int_{\substack{x(t_0)=B \\ x(t)=x_f}} Dx e^{iS[x]}$$

$\bar{D}_F(\alpha, 0; B, t_0)$ $D_F(B, t_0; x_f, t)$

Rather than writing everything out longhand, I'm going to use:
 $\longrightarrow \equiv D_F$, $\dashrightarrow \equiv \bar{D}_F$

$$P_R(t) = \int_R dx_f \int_{\alpha} \int_{x_f} \dots$$

(Integrals over intermediate times implied.)

$$\approx \int_R dx_f \langle B, t_0 | x_f, t \rangle \langle x_f, t | B, t_0' \rangle$$

$$\approx \langle B, t_0 | B, t_0' \rangle$$

(4)

$$\Gamma \equiv \lim_{t \rightarrow \infty} \frac{\overline{D_F(a,0;B,t)} D_F^*(a,0;B,t) + c.c.}{\int dx |D_F(a,0;x,t)|^2}$$

This can be simplified by Wick rotation ($\tau \rightarrow it$), noting that:

- D_F and D_F^* rotate with opposite phases

- The $\delta(t_B[x] - t_0)$ factors pick up 'i'.

- The two terms in the numerator can be

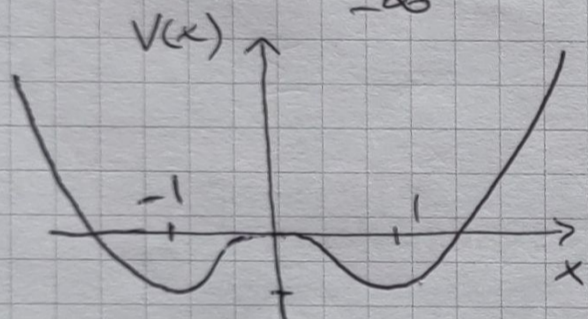
combined into one going from $-\tau$ to τ ; $x(\pm\tau) = a$.

$$\Gamma = \lim_{\tau \rightarrow \infty} 2 \operatorname{Im} \left(\frac{\int dx e^{-S_E[x]} \delta(t_B[x])}{\int dx e^{-S_E[x]}} \right)$$

Imaginary parts from 'real' integrals?

$$I \equiv \int_{-\infty}^{\infty} dx e^{-\beta V(x)}$$

$$V(x) = -\frac{x^2}{2} + \frac{x^4}{4}, \quad \beta \gg 1.$$



Steepest descent method:

$$\begin{aligned} \text{solve extrema} & \begin{cases} V'(x) = -x(1-x^2) \stackrel{!}{=} 0 \\ \Rightarrow x = 0, \pm 1 \end{cases} \\ \text{quadratic fluct}^{\text{ns}} & \begin{cases} V''(x) = -1 + 3x^2 \end{cases} \end{aligned}$$

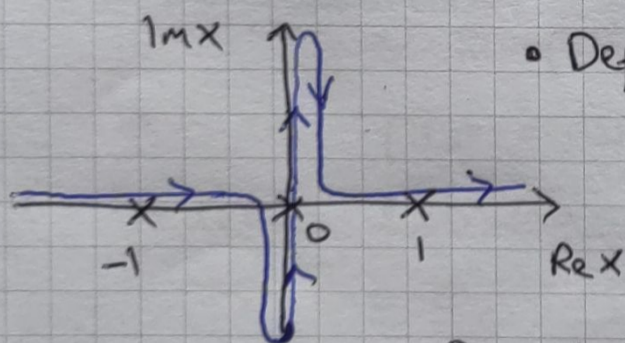
$$I = I_1 + I_2 + I_3$$

- Deform contour in complex plane to

pass through each extremum so

that the real part diminishes

fastestst away from the extremum.



$$I_1 = I_3 \approx \int_{-\infty}^{\infty} dx e^{\beta(\frac{1}{4} - \frac{3}{2}x^2)} = \sqrt{\frac{2\pi}{3\beta}} e^{\beta/4}$$

$$I_2 \approx \int_{-i\infty}^{i\infty} dx e^{\beta x^2/2} = -i \int_{-\infty}^{\infty} dy e^{-\beta y^2} = -i \sqrt{\frac{2\pi}{\beta}}$$

$$\Rightarrow I \approx \sqrt{\frac{2\pi}{3\beta}} e^{\beta/4} - i \sqrt{\frac{2\pi}{\beta}} + \sqrt{\frac{2\pi}{3\beta}} e^{\beta/4}$$

↑ exponentially subdominant

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- I_2 should be dropped at this order. I is real, and all imaginary parts must cancel between the saddles in a complete calculation.
- In Γ , the $\delta(\tau_B[x])$ removes the dominant saddles, leaving us with a genuine imaginary part.

Decay rate in QFT

$$\Gamma = \lim_{\tau \rightarrow \infty} 2 \operatorname{Im} \left(\frac{\int D\phi e^{-S_E} \delta(\tau_\Sigma[\phi])}{V \int D\phi e^{-S_E}} \right)$$

$\Sigma \equiv$ co-dimension 1 surface in field configuration space separating metastable and stable minima.

Steepest descent:

solve extrema: $\left. \frac{\delta S_E}{\delta \phi} \right|_{\tilde{\phi}} = 0$, $\tilde{\phi}(\pm\infty) = \phi_a \leftarrow$ metastable minimum
 $\tilde{\phi}(\tau_0=0) \in \Sigma$ in numerator.

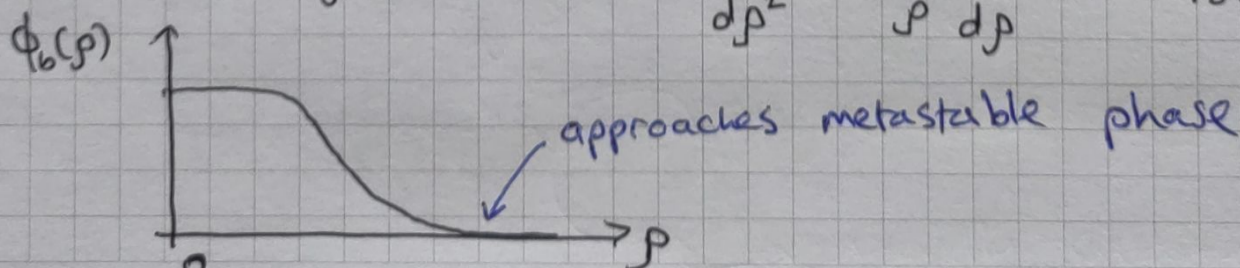
quadratic fluctuations: $-\partial_\mu \partial_\mu + V''(\phi_b)$ $\frac{\delta^2 S_E}{\delta \phi \delta \phi} \Big|_{\tilde{\phi}} f_\alpha = \lambda_\alpha f_\alpha$

If there is a single negative eigenvalue, $\lambda_- < 0$
 \Rightarrow steepest descent contour gives 'c'.

For a first-order phase transition there are two relevant extrema:

- $\tilde{\phi} = \phi_a$ relevant for denominator
- $\tilde{\phi} = \phi_{\text{bounce}}(\rho)$ relevant for numerator

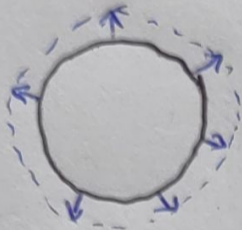
$O(4)$ symmetric solⁿ: $\frac{d^2 \phi_b}{d\rho^2} + \frac{3}{\rho} \frac{d\phi_b}{d\rho} - V'(\phi_b) = 0$.



$\phi_b(\rho) \in \Sigma \iff \exists$ small perturbations which lower action towards both metastable and stable phases = negative eigenvalue ⑥

Negative eigenmode is growing/shrinking, roughly.

$$\phi_b(\rho) \rightarrow \phi_b(\rho) + \epsilon \rho \underbrace{\frac{\partial \phi_b(\rho)}{\partial \rho}}_{\text{approximate negative eigenmode}}$$



In addition there are zero modes corresponding to symmetries broken by the solⁿ (Goldstone modes), most notably translations.

$$\partial_\nu (\partial_\mu \partial_\mu \phi_b - V'(\phi_b)) = 0.$$

$$\Rightarrow (\partial_\mu \partial_\mu - V''(\phi_b)) \partial_\nu \phi_b = 0.$$

$$\Rightarrow \partial_\nu \phi_b \text{ are zero eigenmodes.}$$

In total, write,

$$\phi = \phi_b(x - \bar{x}) + \sum_n' C_n f_n(x - \bar{x}),$$

want to integrate over \bar{x}_0 rather than $\partial_\nu \phi_b$.

$$\int_x f_n f_m = \delta_{nm}$$

$$\|\delta\phi\|^2 = \int_x (\phi_b(x - \delta\bar{x}) + \sum_n' C_n f_n(x - \delta\bar{x}) - \phi_b(x))^2$$

$$= \int_x (\delta\bar{x}_\mu \partial_\mu \phi_b + \sum_n' C_n f_n)^2$$

$$= \delta\bar{x}_\mu \delta\bar{x}_\nu \int_x \partial_\mu \phi_b \partial_\nu \phi_b + \sum_n' C_n^2$$

$$= \delta\bar{x}_\mu \delta\bar{x}_\nu \underbrace{\frac{1}{d} \int_x \partial_\mu \phi_b \partial_\nu \phi_b}_{J^2} + \sum_n' C_n^2.$$

$$\Rightarrow \int D\phi = \left(\prod_n \frac{dC_n}{\sqrt{2\pi}} \right) \cdot J^d \frac{d^d x}{(2\pi)^{d/2}}$$

$$S_E[\phi] \approx S_E[\phi_b] + \frac{1}{2} \sum_n' \lambda_n c_n^2, \quad (7)$$

And now putting it all together,

$$\begin{aligned} & \int D\phi e^{-\frac{S_E[\phi]}{\hbar}} \delta(\tau_E[\phi]) \\ & \approx \int \left(\prod_n' \frac{dc_n}{\sqrt{2\pi}} \right) \int \frac{d^4x}{(2\pi)^4} \exp\left(-\frac{S_E[\phi_b]}{\hbar} + \frac{i}{2} |\lambda-1| c^2 - \frac{i}{2} \sum_n' \lambda_n c_n^2\right) \delta(\tau_E) \\ & = e^{-\frac{S_E[\phi_b]}{\hbar}} \cdot \frac{\int^4}{(2\pi)^4} \cdot \nu \left(\prod_n' \left(\frac{1}{\lambda_n} \right)^{\frac{1}{2}} \right) \cdot \frac{1}{2} \frac{i}{|\lambda-1|} \end{aligned}$$

And the denominator,

$$\nu \int D\phi e^{-\frac{S_E[\phi]}{\hbar}} = \nu \prod_n \left(\frac{1}{\lambda_n} \right)^{\frac{1}{2}}$$

So that,

$$\Gamma = \left(\frac{\int}{\sqrt{2\pi}} \right)^4 \cdot \sqrt{\frac{\det(-\partial_\mu \partial_\mu + V''(\phi_a))}{|\det'(-\partial_\mu \partial_\mu + V''(\phi_b))|}} e^{-\frac{S_E[\phi_b]}{\hbar}}$$

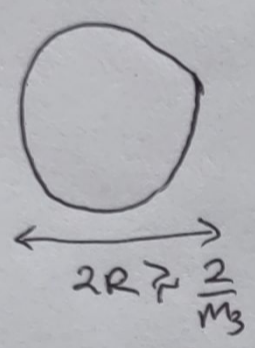
Adding nonzero temperature

Start from a thermal initial state in metastable phase $\hat{\rho}_{\text{meta}}$, rather than a pure state $|\phi_a\rangle$.
 General structure expected to follow from $\mathbb{R}^4 \rightarrow \mathbb{R}^3 \times S^1$:

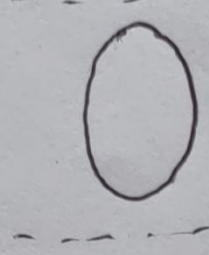
$$\Gamma_T \sim 2 \text{Im} \frac{\int \mathcal{D}\phi e^{-S_E[\phi]} \delta(\tau_\pm[\phi])}{\int \mathcal{D}\phi e^{-S_E[\phi]}} \quad \left| \quad \tau = \frac{1}{T}, \text{ periodic in } \tau \right.$$

Bounce solutions:

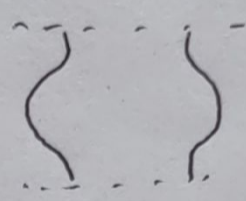
$T=0, O(4)$



$T < \frac{1}{2R}, O(3)$



$T \sim \frac{1}{2R}, O(3)$



$T > \frac{1}{2R}, O(3) \times O(1)$



$T \gg m_3$
 \Rightarrow this case

High temperature bounce:

$$\frac{d^2 \phi_b}{dr^2} + \frac{2}{r} \frac{d\phi_b}{dr} - V'_{\text{EFT}}(\phi_b) = 0$$

potential at tree-level in 3d

First do dimensional reduction, then compute bounce to find, like 3d vacuum decay rate

$$\Gamma_T = A_{\text{dyn}} \left(\frac{J_{\text{EFT}}}{\sqrt{2\pi}} \right)^3 \sqrt{\frac{\det(-\partial_i \partial_i + V''_{\text{EFT}}(\phi_a))}{|\det'(-\partial_i \partial_i + V''_{\text{EFT}}(\phi_b))|}} e^{-S_{\text{EFT}}[\phi_b]}$$

\leftarrow real-time part associated with $\delta(\tau_\pm)$ direction
 \leftarrow modes $< \Lambda$
 \leftarrow modes $> \Lambda$ enter through parameters of EFT

$A_{\text{dyn}} \approx$ (initial growth rate of bubble)

$$\approx \frac{1}{2\pi} \lambda^{-1/2}$$

} again, damping modifies this
 Langer, Annals of Physics 54: 258-275 (1969)