# Light-ray OPE of EEC and Collinear Limit

Hao Chen Zhejiang University

Energy Correlators at the Collider Frontier July 18, 2024

## **Collider Physics**



Examples of observables: thrust, C-parameter, energy-energy correlator,...

Probing microscopic physics from high-energy collision, through measurement at "infinity".

this workshop

## Holography



### AdS/CFT correspondence

AdS scattering = CFT correlation functions

Various theoretical interest in understanding bulk physics from the boundary of spacetime.



flat space

less clear about <u>operators(?)</u> on the flat space boundary

Also see Sruthi's talk



## **Correlators on the Celestial Sphere**

The experiment observables can serve as correlators on the boundary of flat space.

[Basham, Brown, Ellis and Love, 1978] **Example:** energy-energy correlator (EEC)



a point on the celestial sphere  $\rightarrow$ 

calorimeter



a light-ray at null infinity

light-ray operator

## Energy Flow Operator $\mathcal{E}(\vec{n})$

• Building block of EEC, with which we can immediately generalize 2-point EEC to multi-point EECs.

energy E, e.g.  $\frac{d^{3}p}{(2\pi)^{3}2E_{\vec{n}}}$ 



- Integrate t to get the total received energy
- Detector is effectively located at infinity

- "Perturbative" vs "Non-perturbative"
- A calorimeter only detects particles flowing along direction  $\vec{n}$ , and weight with its

$$E_{\vec{p}} a_{\vec{p}}^{\dagger} a_{\vec{p}} \delta^{(2)} (\vec{n} - \hat{p}) \longrightarrow \text{Amplitude Method}$$

Non-perturbative definition via energymomentum tensor

$$\mathcal{E}(\vec{n}) = \lim_{R \to \infty} R^2 \int_0^\infty dt \, n_i T^{0i}(t, R\vec{n}) \longrightarrow \mathbb{N}$$

[Sveshnikov, Tkachov, 1996; Hofman, Maldacena, 2008]







## Amplitudes v.s. Correlation Functions

### Energy Correlation Correlators Functions Que of the simplest observances from the theorem cal perspective is the Energy-Energy Correlator (EEC), defined as (1.1)Here $\mathcal{E}_i$ and $E_j$ are the energies (of ) $\overline{\mathbf{f}_i}$ state parton $\hat{\mathbf{y}}$ and j in the center-of-mass frame, and AneroIntgelsinspleatacionerisables drove the phoquetical the spective in the Danser Frank view Sparselatore (FFGsudefinedeas FC dan also be defined in terms of correlation function of ANEC operators [4–7] $\frac{d\sigma}{dz} = \sum_{i,j} \int d\sigma \frac{E_i E_j}{Q^2} \delta\left(z - \frac{1 - \cos\chi_{ij}}{2}\right).$ (1.1) The energies of final-state partons *i* and *j* in the center-of-mass frame, (1.1) EEC in M=4and their angular separation is $\chi_{ii}$ . $d\sigma$ is the product of the squared matrix element and the whase it is esized we. The EEC can also be defined in terms of correlation function of ANEC $\psi_{ij}$ and $\psi_{ij}$ on mirror curve 4 transmission respectively rowichts 1-4 and 2-3 respectively rowichts 1-4 and 2-3 respectively rowichts and 2-3 respectively rowichts 1-4 and 2-3 respectively rowichts 1-(1.3)Sokatchev, Zhiboedov, 2014] woving the study of event shapes to profit from is vice prover the first of the stand of the stand of the poly department of the conversely, the EEC provide a finantappogaes functions the EEC (1.3) a tenned from partity batten the out and integrative bation theory and integrabil-for generic another the differences for all constructions proceeded and the state les, the EEC has been computed at next-to-Bothen setween and the setwart of the setween and the setween and the setween and the setwart of the A AMESY APPRATOR allowing builds to be a set of the second and the standard of the standard standard shares to a second at the second standard standard standard standard standard and second at the second standard stand recent developments in the study of ANEC appended of an enversely the study of ANEC appended a provide a provide a concrete situation for studying the behavior of antifice situations for tarlying the behavior of ANEC operators of the EEC, which occur as





## **End-point Regions in EEC**



in this workshop...

TMD Factorization and Resummation [Moult, Zhu, 2018]

Recently achieved N<sup>4</sup>LL accuracy



Thankfully, Dingyu gave a back-to-back story!





### **Collinear Limit Result from CMS** [Meng Xiao et al, 2023]

### **Previous exploration from CMS Open Data**



### See illustration on

https://cms.cern/news/jets-elucidate-how-partons-evolve-hadrons



Enables precision measurement of  $\alpha_s$  in jet substructure





### Common interest on collinear limit from different perspectives (May 3, 2019):

### **CFT** Perspective

	nformal collid
Murat Kologlu (Caltech), Petr K 2019)	ravchuk (CERN),
Published in: JHEP 01 (2021) 1	28 • e-Print: 19
」pdf & DOI ⊡ cite	🗟 claim
Energy correlations in the	end-point re
Energy correlations in the G.P. Korchemsky (IPhT, Saclay)	<b>e end-point re</b> (May 4, 2019)
Energy correlations in the G.P. Korchemsky (IPhT, Saclay) Published in: <i>JHEP</i> 01 (2020) 0	e end-point re (May 4, 2019) 008 • e-Print: 190

### **QCD** Perspective

### Collinear limit of the energy-energy correlator

2019)

Published in: *Phys.Rev.D* 100 (2019) 1, 014009 • e-Print: 1905.01310 [hep-ph]

🔎 pdf ∂ DOI [ → cite 🗟 claim









### Universality in the Collinear Limit

Exhibits collinear universality in pQCD, described by the factorization formula



### Light-ray Operators and OPE

## Light-ray Operators

### light transform of local operators

Energy flow operator  $\mathcal{E}(\vec{n}) = \lim_{r \to \infty} r^2$ 

Energy radiation obeys inverse square law,  $r^2$  compensates this effect to be non-vanishing.

### Generalization to other local operators

$$\mathbb{O}(\vec{n}) = \lim_{r \to \infty} r^{\text{twist}} \int_0^\infty dt \ O^{\mu_1 \dots \mu_J} (t, r\vec{n}) \bar{n}_{\mu_1} \dots \bar{n}_{\mu_J}$$

In other contexts of physics, light-ray operators are not necessarily at null infinity—they can live on any light-ray.

$$\mathbf{L}[\mathcal{O}](\boldsymbol{x},\boldsymbol{n}) = \int_{-\infty}^{\infty} d\alpha (-\alpha)^{-\Delta - J} \mathcal{O} \left( \mathbf{x},\boldsymbol{n} \right) = \int_{-\infty}^{\infty} d\alpha (-\alpha)^{-\Delta - J} \mathcal{O} \left( \mathbf{x},\boldsymbol{n} \right) = \int_{-\infty}^{\infty} d\alpha (-\alpha)^{-\Delta - J} \mathcal{O} \left( \mathbf{x},\boldsymbol{n} \right) = \int_{-\infty}^{\infty} d\alpha (-\alpha)^{-\Delta - J} \mathcal{O} \left( \mathbf{x},\boldsymbol{n} \right) = \int_{-\infty}^{\infty} d\alpha (-\alpha)^{-\Delta - J} \mathcal{O} \left( \mathbf{x},\boldsymbol{n} \right) = \int_{-\infty}^{\infty} d\alpha (-\alpha)^{-\Delta - J} \mathcal{O} \left( \mathbf{x},\boldsymbol{n} \right) = \int_{-\infty}^{\infty} d\alpha (-\alpha)^{-\Delta - J} \mathcal{O} \left( \mathbf{x},\boldsymbol{n} \right) = \int_{-\infty}^{\infty} d\alpha (-\alpha)^{-\Delta - J} \mathcal{O} \left( \mathbf{x},\boldsymbol{n} \right) = \int_{-\infty}^{\infty} d\alpha (-\alpha)^{-\Delta - J} \mathcal{O} \left( \mathbf{x},\boldsymbol{n} \right) = \int_{-\infty}^{\infty} d\alpha (-\alpha)^{-\Delta - J} \mathcal{O} \left( \mathbf{x},\boldsymbol{n} \right) = \int_{-\infty}^{\infty} d\alpha (-\alpha)^{-\Delta - J} \mathcal{O} \left( \mathbf{x},\boldsymbol{n} \right) = \int_{-\infty}^{\infty} d\alpha (-\alpha)^{-\Delta - J} \mathcal{O} \left( \mathbf{x},\boldsymbol{n} \right) = \int_{-\infty}^{\infty} d\alpha (-\alpha)^{-\Delta - J} \mathcal{O} \left( \mathbf{x},\boldsymbol{n} \right) = \int_{-\infty}^{\infty} d\alpha (-\alpha)^{-\Delta - J} \mathcal{O} \left( \mathbf{x},\boldsymbol{n} \right) = \int_{-\infty}^{\infty} d\alpha (-\alpha)^{-\Delta - J} \mathcal{O} \left( \mathbf{x},\boldsymbol{n} \right) = \int_{-\infty}^{\infty} d\alpha (-\alpha)^{-\Delta - J} \mathcal{O} \left( \mathbf{x},\boldsymbol{n} \right) = \int_{-\infty}^{\infty} d\alpha (-\alpha)^{-\Delta - J} \mathcal{O} \left( \mathbf{x},\boldsymbol{n} \right) = \int_{-\infty}^{\infty} d\alpha (-\alpha)^{-\Delta - J} \mathcal{O} \left( \mathbf{x},\boldsymbol{n} \right) = \int_{-\infty}^{\infty} d\alpha (-\alpha)^{-\Delta - J} \mathcal{O} \left( \mathbf{x},\boldsymbol{n} \right) = \int_{-\infty}^{\infty} d\alpha (-\alpha)^{-\Delta - J} \mathcal{O} \left( \mathbf{x},\boldsymbol{n} \right) = \int_{-\infty}^{\infty} d\alpha (-\alpha)^{-\Delta - J} \mathcal{O} \left( \mathbf{x},\boldsymbol{n} \right) = \int_{-\infty}^{\infty} d\alpha (-\alpha)^{-\Delta - J} \mathcal{O} \left( \mathbf{x},\boldsymbol{n} \right) = \int_{-\infty}^{\infty} d\alpha (-\alpha)^{-\Delta - J} \mathcal{O} \left( \mathbf{x},\boldsymbol{n} \right) = \int_{-\infty}^{\infty} d\alpha (-\alpha)^{-\Delta - J} \mathcal{O} \left( \mathbf{x},\boldsymbol{n} \right) = \int_{-\infty}^{\infty} d\alpha (-\alpha)^{-\Delta - J} \mathcal{O} \left( \mathbf{x},\boldsymbol{n} \right) = \int_{-\infty}^{\infty} d\alpha (-\alpha)^{-\Delta - J} \mathcal{O} \left( \mathbf{x},\boldsymbol{n} \right) = \int_{-\infty}^{\infty} d\alpha (-\alpha)^{-\Delta - J} \mathcal{O} \left( \mathbf{x},\boldsymbol{n} \right) = \int_{-\infty}^{\infty} d\alpha (-\alpha)^{-\Delta - J} \mathcal{O} \left( \mathbf{x},\boldsymbol{n} \right) = \int_{-\infty}^{\infty} d\alpha (-\alpha)^{-\Delta - J} \mathcal{O} \left( \mathbf{x},\boldsymbol{n} \right) = \int_{-\infty}^{\infty} d\alpha (-\alpha)^{-\Delta - J} \mathcal{O} \left( \mathbf{x},\boldsymbol{n} \right) = \int_{-\infty}^{\infty} d\alpha (-\alpha)^{-\Delta - J} \mathcal{O} \left( \mathbf{x},\boldsymbol{n} \right) = \int_{-\infty}^{\infty} d\alpha (-\alpha)^{-\Delta - J} \mathcal{O} \left( \mathbf{x},\boldsymbol{n} \right) = \int_{-\infty}^{\infty} d\alpha (-\alpha)^{-\Delta - J} \mathcal{O} \left( \mathbf{x},\boldsymbol{n} \right) = \int_{-\infty}^{\infty} d\alpha (-\alpha)^{-\Delta - J} \mathcal{O} \left( \mathbf{x},\boldsymbol{n} \right) = \int_{-\infty}^{\infty} d\alpha (-\alpha)^{-\Delta - J} \mathcal{O} \left( \mathbf{x},\boldsymbol{n} \right) = \int_{-\infty}^{\infty} d\alpha (-\alpha)^{-\Delta - J} \mathcal{O} \left( \mathbf{x},\boldsymbol{n} \right) = \int_{-\infty}^{\infty} d\alpha (-\alpha)^{-\Delta - J} \mathcal{O} \left( \mathbf{x},\boldsymbol{n} \right) = \int_{-\infty}^{\infty} d\alpha (-\alpha)^{-\Delta - J} \mathcal{O} \left( \mathbf{x},\boldsymbol{n} \right) = \int_{-\infty}^{\infty} d\alpha (-\alpha)^{-\Delta - J} \mathcal{O} \left( \mathbf{x},\boldsymbol{n} \right) = \int_{-\infty}^{\infty} d\alpha (-\alpha)^{-\Delta - J} \mathcal{$$

$$\int_0^\infty dt \ \vec{n}_i T^{0i}(t,r\vec{n})$$

( n )x - -, n $\mathcal{X}$ Duffin, 2018] 12

Examples of more general light-ray operators, see [Chang, Kologlu, Kravchuk, Simmons-Duffin, Zhiboedov, 2020; Caron-Huot, Kologlu, Kravchuk, Meltzer, Simmons-Duffin, 2022;...]





## **Spacetime Symmetry**

# Little group that fixes a light-ray in future null infinity consists of

**Poincare group part** 

 $\mathbb{O}(\vec{n}) =$ • Dimension = J - 1

 $\lim_{\bar{n}\cdot x\to\infty} (x)$ • Collinear Spin =  $1 - \Delta$ **Boost quantum number** 

Boost along  $\vec{n}$   $n^{\mu} \to \lambda n^{\mu}$ ,  $\bar{n}^{\mu} \to \lambda^{-1} \bar{n}^{\mu}$ 

- Transverse Spin = transverse spin of the local operator
- Momentum = 0 (invariant under translations)

- translations, collinear boost, transverse rotations, dilatation

$$= \lim_{r \to \infty} r^{\Delta - J} \int_0^\infty dt \, O^{\mu_1 \dots \mu_J}(t, r\vec{n}) \bar{n}_{\mu_1} \dots \bar{n}_{\mu_J}$$
$$-(\Delta - J) \quad -1 \qquad +\Delta$$

$$(\bar{n}\cdot x)^{\Delta-J}\int_{-\infty}^{\infty} d(n\cdot x) \ O^{\mu_1\dots\mu_J}(x)\bar{n}_{\mu_1}\dots\bar{n}_{\mu_J}$$

 $\mathbb{O}(\vec{n}) \to \lambda^{1-\Delta} \mathbb{O}(\vec{n})$ 

13

## Interesting Aspects of Light-ray Operators

- Analyticity in spin
   [Caron-Huot, 2017; Kravchuk, Simmons-Duffin, 2018]
- Level crossing near Regge intercept [Caron-Huot, Kologlu, Kravchuk, Meltzer, Simmons-Duffin, 2022]
- ANEC and new perspective on *a*-theorem and *c*-anomaly [Hartman, Mathys, 2023 and 2024]
- Applications in collider physics





## Light-ray OPE



### Short distance scaling behavior is determined by local Operator Product Expansion (OPE).



[Hofman, Maldacena, 2008]



 $\mathcal{O}(x_1)$ 



Small angle behavior is controlled by the OPE of these light-ray operators.



## Light-ray OPE in CFT



Small angle scaling is dominated by the leading twist operators. still true in pQCD

Conformal symmetry constrains the operators to have spin=3. broken by running coupling

Light-ray OPE in CFT is rigorous and convergent. [Kologlu, Kravchuk, Simmons-Duffin, Zhiboedov, 2019]









## Leading Twist Operators in QCD



The analytic continuation of even spin branch is



### For unpolarized cases, there are only two kinds of twist-2 operators in QCD

See HuaXing's talk for polarized case

$$\frac{d^3 p}{(2\pi)^3 2E} \delta^{(2)}(\hat{p} - \vec{n}) E^{J-1}(b^{\dagger}_{\vec{p},s}b_{\vec{p},s} + d^{\dagger}_{\vec{p},s}d_{\vec{p},s}) \frac{d^3 p}{(2\pi)^3 2E} \delta^{(2)}(\hat{p} - \vec{n}) E^{J-1} a^{\dagger}_{\vec{p},\lambda,c} a_{\vec{p},\lambda,c}$$

## LP Light-ray OPE in QCD at LO

QCD in 4d is classically conformal

$$\mathcal{E}(\hat{n}_1)\mathcal{E}(\hat{n}_2) = -\frac{1}{2\pi}\frac{2}{\theta^2}\vec{\mathcal{J}}\left[\hat{C}_{\phi}(\mathbf{r}_2)\right]$$
$$\vec{\mathcal{J}} = (1, 1, 0, 0)$$
$$\hat{C}_{\phi}(\mathbf{r}_2)$$



### [HC, Moult, Zhu, 2020]

 $(2) - \widehat{C}_{\phi}(3) \left| \vec{\mathbb{O}}^{[3]}(\hat{n}_1) + \text{higher twist} \right|$ 

### **PE coefficient matrix**





at LL accuracy

$$(\mu,\mu) = \int_0^1 dx \, x^2 \vec{J}(\log \frac{\zeta x^2 Q^2}{\mu^2},\mu) \cdot \vec{H}(x,\frac{Q^2}{\mu^2},\mu)$$

[Dixon, Moult, Zhu, 2019]



### "Bootstrapping" Factorization Formula

[2311.00350]

### **Beyond Conformal Symmetry Renormalization of Light-ray Operators**

- In perturbation theory, the light-ray operators have divergences.
  - require renormalization [Caron-Huot, Kologlu, Kravchuk, Meltzer, Simmons-Duffin, 2022]

$$\mathbb{O}_{a;\text{bare}}^{[J]}(n) = \lim_{\bar{n}\cdot x \to \infty} \left(\frac{\bar{n}\cdot x}{2}\right)^2 \int_{-\infty}^{\infty} d(n\cdot x) \mathcal{O}_{a;\text{bare}}^{[J]}(x;\bar{n})$$

Only  $\mathbb{O}_a^{[J]}$  and  $\mathbb{O}_a^{[J]}$  can mix, as long as J is large enough. **RG equation:**  $\frac{d}{d \ln \mu^2} \mathbb{O}_{a;\text{ren}}^{[J]}(n;\mu) = \gamma_{ab}^T(J;\alpha_s(\mu)) \mathbb{O}_{b;\text{ren}}^{[J]}(n;\mu)$ 

Introducing a renormalization factor and define renormalized light-ray operators

 $\mathbb{O}_{a:\text{bare}}^{[J]} = \mathbb{Z}_{ab}^{[J]} \mathbb{O}_{b;\text{ren}}^{[J]}$ 

[IR behaviors of detectors]

contains scale dependence

Time-like anomalous dimension different from the space-like anomalous dimension for the local operators. But they are related by **reciprocity relation**.

bare twist-2 local operators

## Generalizing Conformal Light-ray OPE

Recall the conformal case:  $\mathcal{E}(n)\mathcal{E}(n)$ 

In non-conformal theories, dimension is not a good quantum number.

However, QCD in 4d is classically conformal and the running coupling effect is a higher order effect - We expect the breaking is a small perturbation

Ansatz: 
$$\mathcal{E}(n)\mathcal{E}(n') = \sum_{k=0}^{\infty} C_a^{(k)}(z;\mu) \left[ \frac{\partial_J^k \mathbb{O}_{a;\text{ren}}^{[J]}(n;\mu)}{J=3} \right]_{J=3} + \text{higher twists}$$

Goal: constrain the ansatz from general principles

$$n') \sim \sum_{i} c_i \theta^{\tau_i - 4} \mathbb{O}_i^{[J=3]}(n)$$

To balance the dimension

assume **analyticity** in perturbation theory

$$\mathcal{E}(n)\mathcal{E}(n') = \sum_{k=0}^{\infty} C_a^{(k)}(z;\mu) \left[ \partial_J^k \mathbb{O}_{a;\mathrm{ren}}^{[J]}(n;\mu) \right] \Big|_{J=3} + \text{higher twists}$$

$$(1,0,0,1) \quad (1,\sin\theta\cos\phi,\sin\theta\sin\phi,\cos\theta)$$

Act boost generator on both sides of the ansatz collinear approx  $-2(3+z\partial_z)$  $= \frac{(-3(1+\cos\theta)-\sin\theta\partial_{\theta})\mathcal{E}(n)\mathcal{E}(n'(\theta,\phi))}{-(J+1)\partial_{J}^{k}\mathbb{O}_{a;\mathrm{ren}}^{[J]}(n;\mu)-k\partial_{J}^{k-1}\mathbb{O}_{a;\mathrm{ren}}^{[J]}(n;\mu)},$ 

L.H.S. : 
$$-i[\vec{n} \cdot \vec{\mathbf{K}}, \mathcal{E}(n)\mathcal{E}(n'(\theta, \phi))] =$$
  
R.H.S. :  $-i[\vec{n} \cdot \vec{\mathbf{K}}, \partial_J^k \mathbb{O}_{a;ren}^{[J]}(n; \mu)] =$ 

**Recursion Relation:** 

$$\left(\frac{\partial}{\partial \ln z} + 1\right) C_a^{(k)}(z;\mu) = \frac{k+1}{2} C_a^{(k+1)}(z;\mu)$$

Solution:

$$C_a^{(k)}(z;\mu) = \frac{1}{z} \frac{2^k}{k!} \left(\frac{\partial}{\partial \ln z}\right)^k \widetilde{C}_a(z;\mu)$$

classical scaling behavior

## Lorentz Symmetry

 Only single undetermined function Contains logarithms in perturbation theory



## **RG** invariance

### The RG evolution of coefficients are determined by the RG of operators



### They are compatible with the Lorentz symmetry constraint

$$\begin{split} & \left[ \mathbb{O}_{a;\text{ren}}^{[J]}(n;\mu) \right] \Big|_{J=3} + \text{higher twists} \\ & \textbf{known RG behavior} \\ & \frac{d}{d \ln \mu^2} \mathbb{O}_{a;\text{ren}}^{[J]}(n;\mu) = \gamma_{ab}^T(J;\alpha_s(\mu)) \mathbb{O}_{b;\text{ren}}^{[J]}(n;\mu) \\ & \text{efficients} \end{split}$$

$$\frac{k+m)!}{k!m!}C_b^{(k+m)}(z;\mu)\partial_J^m\gamma_{ba}^T(J;\alpha_s(\mu))$$

$$\frac{d}{d\ln\mu^2} \widetilde{C}_a(z;\mu) = -\sum_{m=0}^{\infty} \frac{2^m}{m!} \left(\frac{\partial}{\partial\ln z}\right)^m \widetilde{C}_b(z;\mu) \left[\partial_J^m \gamma_{ba}^T(J;\alpha_s(\mu))\right] \Big|_{J=3}$$

### **Constraint from Physical Observables**

Consider EEC with the center of mass energy Q. In perturbative massless QCD, we expect the functional from

$$\langle \mathcal{E}(n)\mathcal{E}(n')\rangle_Q = Q^2 f(z, \ln \frac{Q^2}{\mu^2}, \alpha_s(\mu))$$

But the light-ray OPE does not manifest this property

$$\langle \mathcal{E}(n)\mathcal{E}(n')\rangle_Q = \frac{1}{z}\sum_{k=0}^{\infty}\frac{2^k}{k!}\left[\left(\frac{\partial}{\partial\ln z}\right)\right]$$

$$\begin{array}{c} \text{Rearrange} \quad Q^2 \\ \hline z \\ m = 0 \end{array} \overset{\infty}{\longrightarrow} \frac{2^m}{m!} \left[ \partial_J^m h_a(J, \ln \frac{Q^2}{\mu^2}; \alpha_s(\mu)) \right] \Big|_{J=3} \left( \frac{\partial}{\partial \ln z} \right)^m \sum_{k=0}^{\infty} \frac{2^k}{k!} (\ln Q)^k \left( \frac{\partial}{\partial \ln z} \right)^k \widetilde{C}_a(z; \mu) = \widetilde{C}_a(zQ)^k \left( \frac{\partial}{\partial \ln z} \right)^k \left( \frac{\partial}{\partial \ln z}$$

**Constraint:** 

$$\widetilde{C}_a(z;\mu) = \widetilde{C}_a(\ln \frac{z}{\mu^2};\alpha_s(\mu)$$

 $Q^{J-1}h_a(J,\ln\frac{Q^2}{\mu^2};\alpha_s(\mu))$  $\frac{1}{z} \overset{\kappa}{\widetilde{C}}_{a}(z;\boldsymbol{\mu}) \left[ \partial_{J}^{k} \langle \mathbb{O}_{a;\mathrm{ren}}^{[J]}(n;\boldsymbol{\mu}) \rangle_{Q} \right] \Big|_{J=3}$ 





## **Deriving Factorization from Light-ray OPE**

$$\frac{\langle \mathcal{E}(n)\mathcal{E}(n')\rangle_Q}{Q^2} = \frac{1}{z} \sum_{m=0}^{\infty} \frac{1}{m!} \left[ \frac{\partial_J^m h_a(J, \ln \frac{Q^2}{\mu^2}; \alpha_s(\mu))}{\mu^2} \right] \Big|_{J=3} \times 2^m \left( \frac{\partial}{\partial \ln z} \right)^m \widetilde{C}_a(\ln \frac{zQ^2}{\mu^2}; \alpha_s(\mu))$$

**Mellin moment**  $h_a(J, \ln \frac{Q^2}{\mu^2}; \alpha_s(\mu))$ 

$$= \frac{1}{z} \int_{0}^{1} dx \, x^{2} \widetilde{h}_{a}(x, \ln \frac{Q^{2}}{\mu^{2}}; \alpha_{s}(\mu)) \sum_{m=0}^{\infty} \underbrace{\frac{2^{m}}{m!} (\ln x)^{m}}_{m!} \left(\frac{\partial}{\partial \ln z}\right)^{m}}_{\widetilde{C}_{a}(\ln \frac{zQ^{2}}{\mu^{2}}; \alpha_{s}(\mu))} \widetilde{C}_{a}(\ln \frac{zx^{2}Q^{2}}{\mu^{2}}; \alpha_{s}(\mu)),$$

$$= \frac{1}{z} \int_{0}^{1} dx \, x^{2} \widetilde{h}_{a}(x, \ln \frac{Q^{2}}{\mu^{2}}; \alpha_{s}(\mu)) \widetilde{C}_{a}(\ln \frac{zx^{2}Q^{2}}{\mu^{2}}; \alpha_{s}(\mu)),$$
hard function jet function QCD factorization formula for EEC

$$\mu)) = \int_0^1 dx \, x^{J-1} \widetilde{h}_a(x, \ln \frac{Q^2}{\mu^2}; \alpha_s(\mu))$$

### **Back to Conformal Case**

For simplicity, let's assume there is no degeneracy and mixing

**RG equation**  $\frac{d}{d \ln \mu^2} \widetilde{C}(\ln \frac{z}{\mu^2}; \alpha_s(\mu)) = -\int_{\eta}^{\eta} \frac{d}{d \ln \mu^2} \widetilde{C}(\ln \frac{z}{\mu^2}; \alpha_s(\mu)) = -\int_{\eta}^{\eta} \frac{d}{d \ln \mu^2} \frac{d}{d \ln \mu^2} + \frac{\beta(\alpha_s)}{2} \frac{d}{d \ln \mu^2} \frac{$ 

Diagonalizing the evolution equation

$$0 = (\nu - \gamma^{T}(J + 2\nu; \alpha_{s}))\mathcal{C}(\nu; \alpha_{s}) \xrightarrow{\text{solution}} \mathcal{C}(\nu; \alpha_{s}) \propto \delta(\nu - \gamma^{T}(J + 2\nu; \alpha_{s})) \propto \delta(\nu - \gamma^{S})$$

$$\gamma^{S}(J) - \gamma^{T}(J + 2\gamma^{S}(J))$$

Exact scaling behavior  $C(\ln z; \alpha_s) \propto z^{\gamma^{\sim}(3; \alpha_s)}$ 

$$-\sum_{m=0}^{\infty} \frac{2^m}{m!} \left(\frac{\partial}{\partial \ln z}\right)^m \widetilde{C}(\ln \frac{z}{\mu^2}; \alpha_s(\mu)) \left[\partial_J^m \gamma^T(J; \alpha_s(\mu))\right]$$

$$\frac{(\alpha_s)}{2} \frac{\partial}{\partial \alpha_s} \longrightarrow \frac{\partial}{\partial \ln \mu^2} = -\frac{\partial}{\partial \ln z}$$

**n with** 
$$\widetilde{C}(\ln z; \alpha_s) = \int \frac{d\nu}{2\pi i} z^{\nu} \mathcal{C}(\nu; \alpha_s)$$

**Reciprocity Relation** 



## Light-ray OPE and Regge Trajectory

### Light-ray operators are expected to be the analytic continuation of local operators. [Kravchuk, Simmons-Duffin, 2018]











## **Evolution of Leading Power Correction**

### Assume power corrections = (hadron minus parton) results from Monte Carlo generator. **Hadron Collider Lepton Collider**



### See HuaXing's talk





## Strong Coupling Limit

Strong coupling limit in generic QFTs are hard. For holographic CFTs, we can access this limit through AdS/CFT correspondence.

Example: N=4 SYM and Type II B superstring theory on  $AdS_5 \times S^5$ 

Expansion in  $1/\lambda$ : Stringy Correction Classical GR < — — large  $\lambda$  and  $N_c$  limit: Expansion in  $1/N_c$ : QG Correction

$$\langle \mathcal{E}(n_1)\mathcal{E}(n_2)\rangle = \left(\frac{p^0}{4\pi}\right)^2 \left[1 + \frac{1}{\lambda}4\pi^2(1 - 6z)\right]$$

Leading QG Correction:  $-384\left(\frac{p^0}{4\pi}\right)^2 \frac{\log c_T}{c_T}\log z \left(1-36z\right)^2$ 

't Hooft coupling:  $\lambda = g_{YM}^2 N_c$ string length:  $\ell_s = \lambda^{-1/4} R_{AdS}$  $G_N \sim N_c^{-2} \sim 1/c_T$ 

EEC in the strong coupling limit: [Hofman, Maldacena, 2008]  $(\mathcal{E})^2$  [Hofman, Maldacena, 2008]  $(z + 6z^2) + ...]$  Also see Riccardo's and Matthew's talks Energy detectors correspond to shockwaves in AdS

$$+216z^2 - 400z^3 + 225z^4 + \cdots$$
 [HC, Karlsson Zhiboedov, 2024]

Enhancement property compared to the local correlator:  $\langle \mathcal{O}^{\dagger}T_{\mu\nu}T_{\rho\sigma}\mathcal{O}\rangle_{c}$  = Tree-level Sugra +  $\frac{1}{\lambda^{3/2}}$  Stringy +  $\frac{1}{c_{\sigma}}$  One-loop Sugra + ...







### Summary

• Light-ray operators play an important role in collider physics.

• Light-ray OPE, organized as twist expansion, governs the small angle scaling behavior.

the QCD factorization formula.

applications, including hadronization effects.

• By generalizing light-ray OPE to non-conformal theory, we are able to derive

• Light-ray OPE is an interesting formalism that has fruitful phenomenological