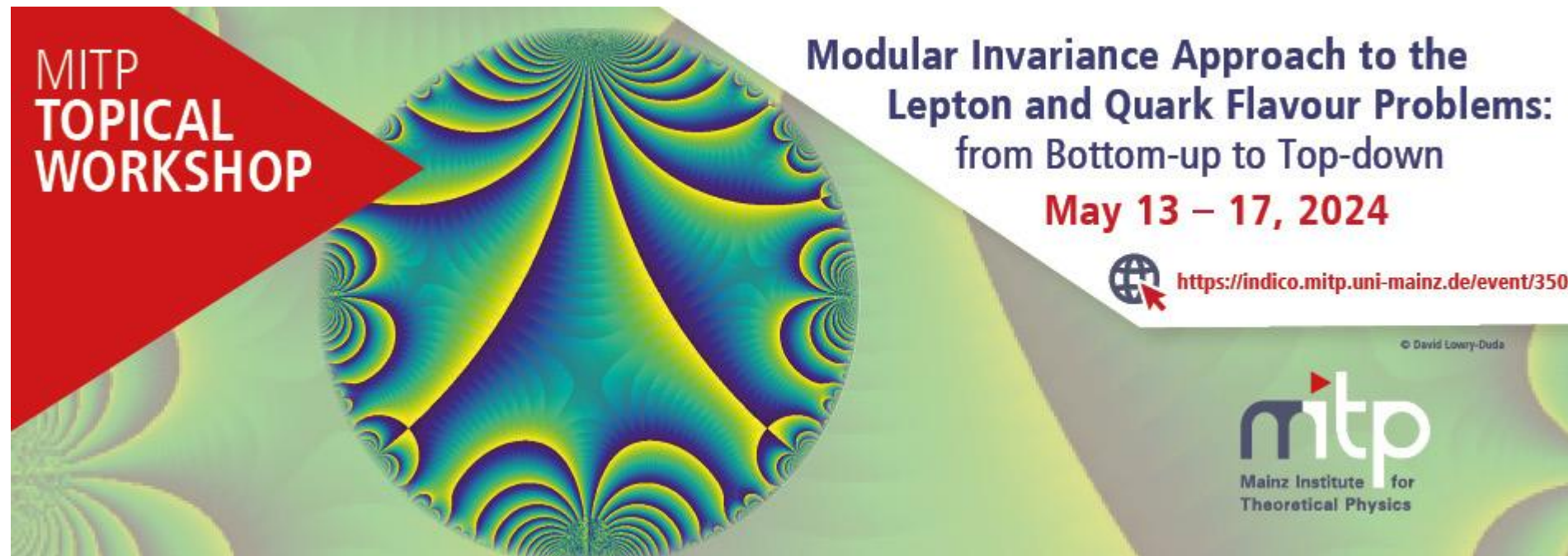


Bottom-up modular invariance approach to flavour I

Gui-Jun Ding


University of Science and Technology of China




MITP
TOPICAL
WORKSHOP

Modular Invariance Approach to the
Lepton and Quark Flavour Problems:
from Bottom-up to Top-down

May 13 – 17, 2024

 <https://indico.mitp.uni-mainz.de/event/350>

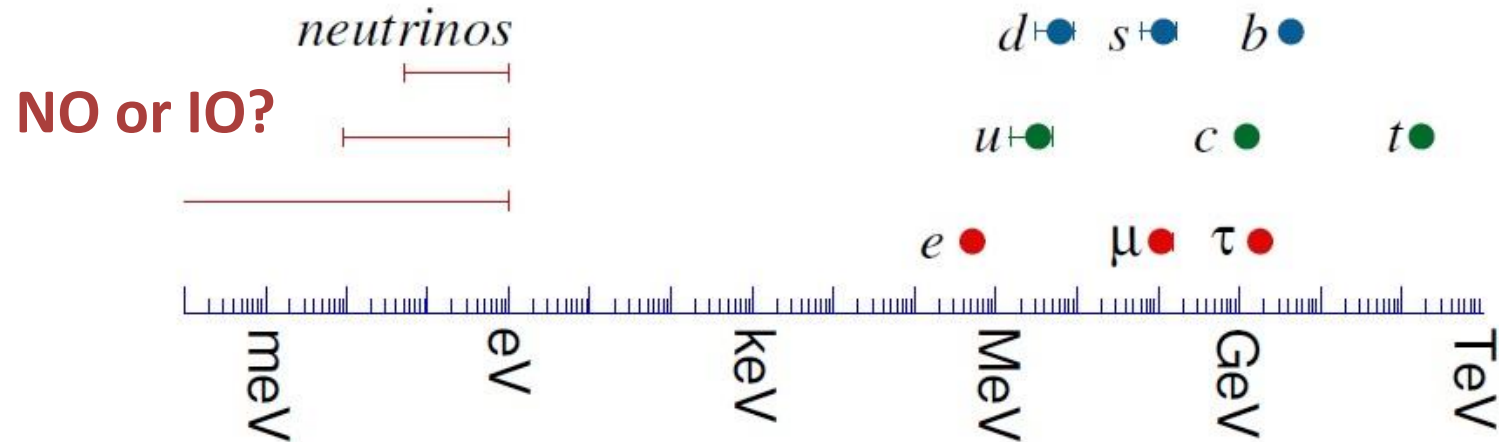
© David Lowery-Duda


mitp
Mainz Institute for
Theoretical Physics

Flavor puzzles in SM

talk by Stephen King

- Hierarchical masses, e.g. $\frac{m_t}{m_e} = \mathcal{O}(10^5)$



- Quark mixing vs lepton mixing

Quark mixings are small

CKM

$$|V| = \begin{matrix} & d & s & b \\ \begin{matrix} u \\ c \\ t \end{matrix} & \begin{bmatrix} \text{large} & \text{small} & \cdot \\ \text{small} & \text{large} & \text{small} \\ \cdot & \text{small} & \text{large} \end{bmatrix} \end{matrix}$$

Lepton mixings are large

MNS

$$|U| = \begin{matrix} & 1 & 2 & 3 \\ \begin{matrix} e \\ \mu \\ \tau \end{matrix} & \begin{bmatrix} \text{large} & \text{large} & \text{small} \\ \text{large} & \text{large} & \text{large} \\ \text{small} & \text{large} & \text{large} \end{bmatrix} \end{matrix}$$

Quark and lepton mixing matrices have distinctive structures!

Flavor symmetry to flavor puzzle

The fundamental principle underlying the fermion masses and flavor mixing structure is unknown so far. Symmetry can help to reduce the number of free parameters in the Yukawa coupling.

➤ **Flavor symmetry:** relate three families

1st-family



2nd-family



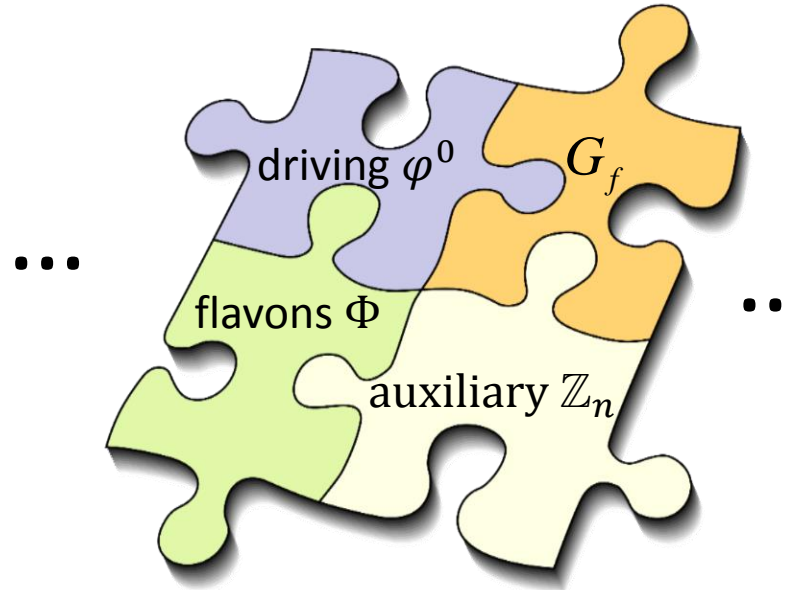
3rd-family

Flavor symmetry(horizontal)

Gauge symmetry(vertical)

Quarks	u	c	t	
	up	charm	top	
	d	s	b	
	down	strange	bottom	
	Leptons	ν_e	ν_μ	ν_τ
		e-neutrino	μ -neutrino	τ -neutrino
e		μ	τ	
	electron	muon	tau	

talks by Stephen King, Myriam Mondragon



[reviews: Altarelli, Feruglio, 1002.0211; Ishimori, Kobayashi et al, 1003.3552; King, Luhn, 1301.1340; King, Merle et al, 1402.4271; King, 1701.04413; Xing, 1909.09610; Feruglio, Romanino, 1912.06028; Almumin, Chen et al, 2204.08668; Ding, Valle, 2402.16963]

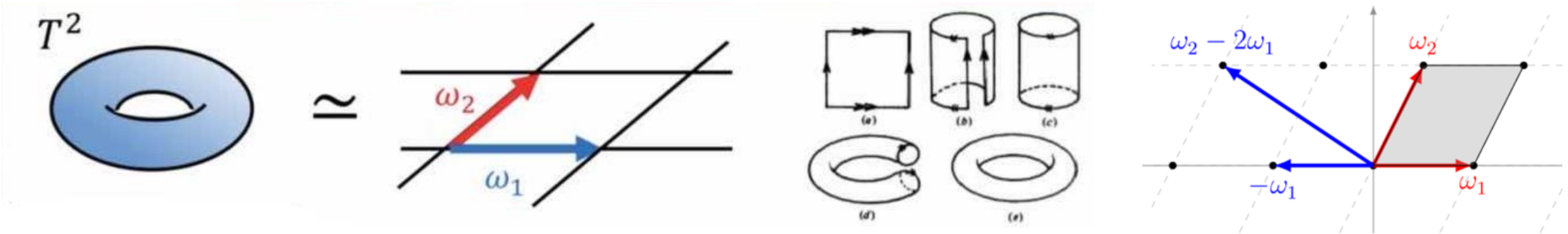
- To make the Lagrangian invariant under flavor symmetry, Higgs-like fields “flavons” Φ_e, Φ_ν are needed
- Structure of Yukawa couplings arises from **the vacuum alignment of flavons**
- Flavons and the vacuum alignment make the models **quite complicated**

Modular symmetry

Modular symmetry $SL(2, \mathbb{Z})$ is the geometrical symmetry of the torus T^2

Torus $T^2 \cong \mathbb{C}/\Lambda_{(\omega_1, \omega_2)}$, lattice $\Lambda_{(\omega_1, \omega_2)} = n_1\omega_1 + n_2\omega_2$

[Feruglio, Romanino, 1912.06028]



Up to rotations and dilations, the shape of a torus is parameterized by $\tau = \omega_2/\omega_1, \text{Im}\tau > 0$

The lattice (torus) is left **invariant** by $SL(2, \mathbb{Z})$ modular transformation

$$\begin{pmatrix} \omega'_2 \\ \omega'_1 \end{pmatrix} = \begin{pmatrix} a & b \\ c & d \end{pmatrix} \begin{pmatrix} \omega_2 \\ \omega_1 \end{pmatrix} \implies \tau \rightarrow \gamma\tau = \frac{a\tau + b}{c\tau + d}$$

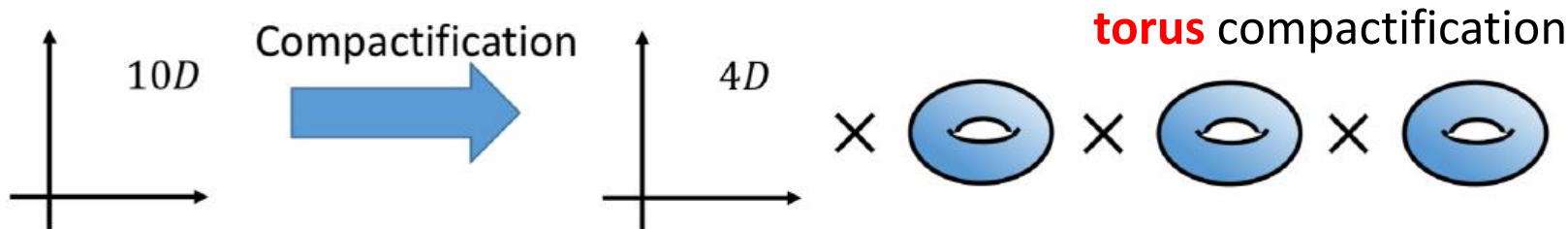
$$SL(2, \mathbb{Z}) = \left\{ \gamma = \begin{pmatrix} a & b \\ c & d \end{pmatrix} \middle| a, b, c, d \in \mathbb{Z}, ad - bc = 1 \right\}$$

modular generators:

$$S: \tau \rightarrow -\frac{1}{\tau}$$

$$T: \tau \rightarrow \tau + 1$$

Effective 4D theories are modular invariant:



$$S = \int d^4x d^6y \mathcal{L}_{10D} \implies \int d^4x \mathcal{L}_{\text{eff}}(\varphi, \tau_i)$$

Modular invariant SUSY theory

➤ The field transformation

[Lauer, Mas, Nilles, 1989; Ferrara, Lust et al, 1989; Feruglio, 1706.08749]

$$\psi \rightarrow (c\tau + d)^{-k} \rho(\gamma) \psi$$

weight $k \in \mathbb{Z}$ ρ is a unitary representation of Γ_N or Γ'_N

➤ Superpotential

$$\mathcal{W} = \sum Y_{I_1 I_2 \dots I_n}(\tau) \psi_{I_1} \psi_{I_2} \dots \psi_{I_n}$$

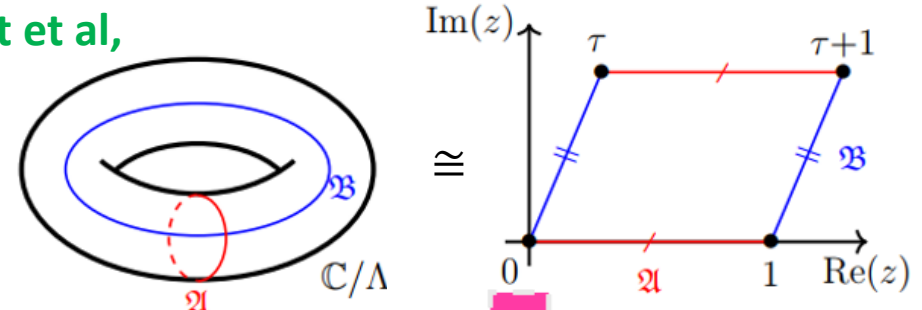
Modular invariance requires

$$Y_{I_1 I_2 \dots I_n}(\tau) \rightarrow Y_{I_1 I_2 \dots I_n}(\gamma\tau) = (c\tau + d)^{k_Y} \rho_Y(\gamma) Y_{I_1 I_2 \dots I_n}(\tau)$$

$$k_Y = k_{I_1} + k_{I_2} + \dots + k_{I_n}, \quad \rho_Y \otimes \rho_{I_1} \otimes \dots \otimes \rho_{I_n} \supset 1$$

- Yukawa couplings are modular forms $Y_{I_1 I_2 \dots I_n}(\tau)$
- Γ_N and Γ'_N play the role of flavor symmetry groups

N	2	3	4	5
Γ_N	S_3	A_4	S_4	A_5
Γ'_N	S_3	$A'_4 = T'$	$S'_4 \cong SL(2, \mathbb{Z}_4)$	$A'_5 \cong SL(2, \mathbb{Z}_5)$



$SL(2, \mathbb{Z})$ on torus T^2

finite modular groups

$$\begin{cases} \Gamma_N \equiv SL(2, \mathbb{Z}) / \pm\Gamma(N) \\ \Gamma'_N \equiv SL(2, \mathbb{Z}) / \Gamma(N) \end{cases}$$

➤ **Advantages:**

- ① No flavons (except τ) and complexity of vacuum alignment;
- ② No corrections from higher dimensional operators in SUSY limit;
- ③ **significant reduction** in the number of parameters

Modular invariant flavor models

Are neutrino masses modular forms?

Ferruccio Feruglio (INFN, Padua and Padua U.) (Jun 27, 2017)

e-Print: [1706.08749](https://arxiv.org/abs/1706.08749) [hep-ph]

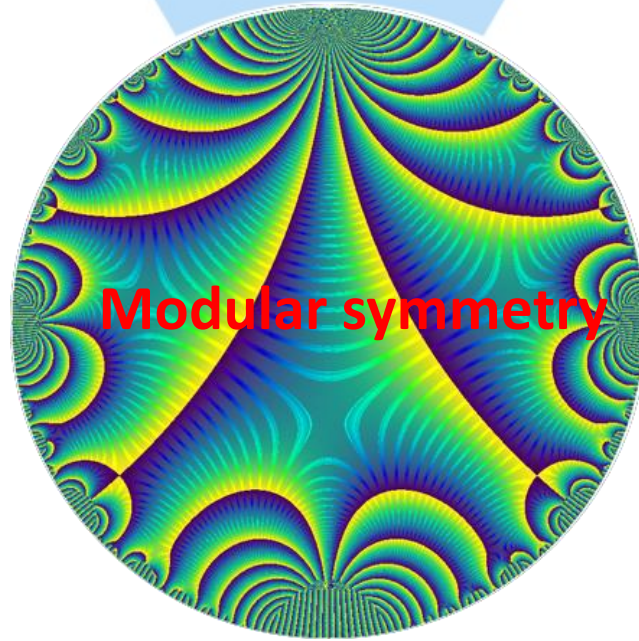
[pdf](#)
[DOI](#)
[cite](#)
[claim](#)
[reference search](#)
[244 citations](#)



➤ Bottom-up models for lepton and quark [reviews: Kobayashi, Tanimoto, 2307.03384; Ding, King, arXiv:2311.09282]

	Γ_N/Γ'_N	leptons alone	leptons & quarks	$SU(5)$	$SO(10)$
$N = 2$	S_3	Kobayashi et al, 1803.10391...	—	Kobayashi et al, 1906.10341...	—
$N = 3$	A_4	Feruglio, 1706.08749 , 1807.01125 ; Kobayashi, Tanaka, et al, 1803.10391 ; Kobayashi, Omoto, et al, 1808.03012...	Okada,Tanimoto, 1905.13421 ; King, King, 2002.00969 ; Yao, Lu, Ding, 2012.13390...	Anda, King,Perdomo, 1812.05620 ; Chen, Ding, King, 2101.12724...	Ding, King,Lu, 2108.09655
	T'	Liu, Ding, 1907.01488...	Lu, Liu, Ding, 1912.07573...	—	—
$N = 4$	S_4	Penedo,Petcov, 1806.11040 ; Novichkov, Penedo et al, 1811.04933...	Qu, Liu et al, 2106.11659	Zhao, Zhang, 2101.02266 ; Ding, King, Yao, 2103.16311...	—
	S'_4	Novichkov,Penedo,Petcov, 2006.03058...	Liu, Yao, Ding, 2006.10722...	—	—
$N = 5$	A_5	Novichkov, Penedo et al, 1812.02158 ; Ding, King, Liu, 1903.12588...	—	—	—
	A'_5	Wang, Yu, Zhou, 2010.10159 ...	Yao, Liu, Ding, 2011.03501	—	—
$N = 6$	Γ_6	—	—	Abe,Higaki et al, 2307.01419	—
	Γ'_6	Li,Liu,Ding, 2108.02181	—	—	—
$N = 7$	Γ_7	Ding, King et al, 2004.12662	—	—	—
	Γ'_7	—	—	—	—

General finite
modular groups



General finite modular groups: $SL(2, \mathbb{Z})/\text{normal subgroups}$

Normal subgroups $\ker(\rho)$			Finite modular groups $\Gamma/\ker(\rho) \cong \text{Im}(\rho)$	
Index	Label	Additional relators	Group structure	GAP Id
6	$\Gamma(2)$	T^2	S_3	[6, 1]
12	—	S^2T^2	$Z_3 \rtimes Z_4 \cong 2D_3$	[12, 1]
	$\pm\Gamma(3)$	S^2, T^3	A_4	[12, 3]
18	—	$ST^{-2}ST^2$	$S_3 \times Z_3$	[18, 3]
24	$\Gamma(3)$	T^3	T'	[24, 3]
	—	S^2T^3		
	$\pm\Gamma(4)$	S^2, T^4	S_4	[24, 12]
	—	$S^2, (ST^{-1}ST)^2$	$A_4 \times Z_2$	[24, 13]
36	—	$S^3T^{-2}ST^2$	$(Z_3 \rtimes Z_4) \times Z_3$	[36, 6]
42	—	$T^6, (ST^{-1}S)^2TST^{-1}ST^2$	$Z_7 \rtimes Z_6$	[42, 1]
	—	$T^6, ST^{-1}ST(ST^{-1}S)^2T^2$		
48	—	S^2T^4	$2O$	[48, 28]
	—	T^8, ST^4ST^{-4}	$GL(2, 3)$	[48, 29]
	$\Gamma(4)$	T^4	$A_4 \rtimes Z_4 \cong S'_4$	[48, 30]
	—	$(ST^{-1}ST)^2$	$A_4 \times Z_4$	[48, 31]
	—	$S^2(ST^{-1}ST)^2$	$T' \times Z_2$	[48, 32]
54	—	T^{12}, ST^3ST^{-3}	$((Z_4 \times Z_2) \rtimes Z_2) \rtimes Z_3$	[48, 33]
	—	$T^6, (ST^{-1}ST)^3$		
60	$\pm\Gamma(5)$	S^2, T^5	A_5	[60, 5]
72	—	T^{12}, ST^4ST^{-4}	$S_4 \times Z_3$	[72, 42]
	$\pm\Gamma(6)$	$S^2, T^6, (ST^{-1}STST^{-1}S)^2T^2$	$A_4 \times S_3$	[72, 44]

[Liu,Ding, 2112.14761;
Ding, Liu, Lu, Weng,
2307.14926; Arriaga-
Osante, Liu, Ramos-
Sanchez, 2311.10136]

The flavor symmetry can be general finite modular groups rather than Γ_N and Γ'_N .

Minimal modular lepton model

Modular symmetry allows to construct quite predictive lepton models. The modular flavor symmetry is modular binary octahedral group **2O** which is the Shur double cover of S_4

	L	$E_D^c = (e^c, \mu^c)$	τ^c	N^c	$H_{u,d}$
$2O$	3	$\widehat{\mathbf{2}}'$	1'	3	1
k_I	-1	6	5	1	0

[Ding, Liu, Lu, Weng, 2307.14926]

➤ Charged leptons

$$\mathcal{W}_E = \alpha \left(E_D^c L Y_{\widehat{\mathbf{2}}'}^{(5)} \right)_1 H_d + \beta \left(E_D^c L Y_{\widehat{\mathbf{4}}}^{(5)} \right)_1 H_d + \gamma \left(\tau^c L Y_{\mathbf{3}'}^{(4)} \right)_1 H_d$$

$$\longrightarrow M_E = \begin{pmatrix} -\alpha Y_{\widehat{\mathbf{2}}',2}^{(5)} - \sqrt{2}\beta Y_{\widehat{\mathbf{4}},3}^{(5)} & \sqrt{3}\beta Y_{\widehat{\mathbf{4}},1}^{(5)} & \sqrt{2}\alpha Y_{\widehat{\mathbf{2}}',1}^{(5)} + \beta Y_{\widehat{\mathbf{4}},4}^{(5)} \\ -\alpha Y_{\widehat{\mathbf{2}}',1}^{(5)} + \sqrt{2}\beta Y_{\widehat{\mathbf{4}},4}^{(5)} & -\sqrt{2}\alpha Y_{\widehat{\mathbf{2}}',2}^{(5)} + \beta Y_{\widehat{\mathbf{4}},3}^{(5)} & -\sqrt{3}\beta Y_{\widehat{\mathbf{4}},2}^{(5)} \\ \gamma Y_{\mathbf{3}',1}^{(4)} & \gamma Y_{\mathbf{3}',3}^{(4)} & \gamma Y_{\mathbf{3}',2}^{(4)} \end{pmatrix} v_d$$

➤ Neutrino mass : seesaw mechanism

$$\mathcal{W}_\nu = g H_u (N^c L)_1 + \Lambda \left(N^c N^c Y_{\mathbf{2}}^{(2)} \right)_1$$

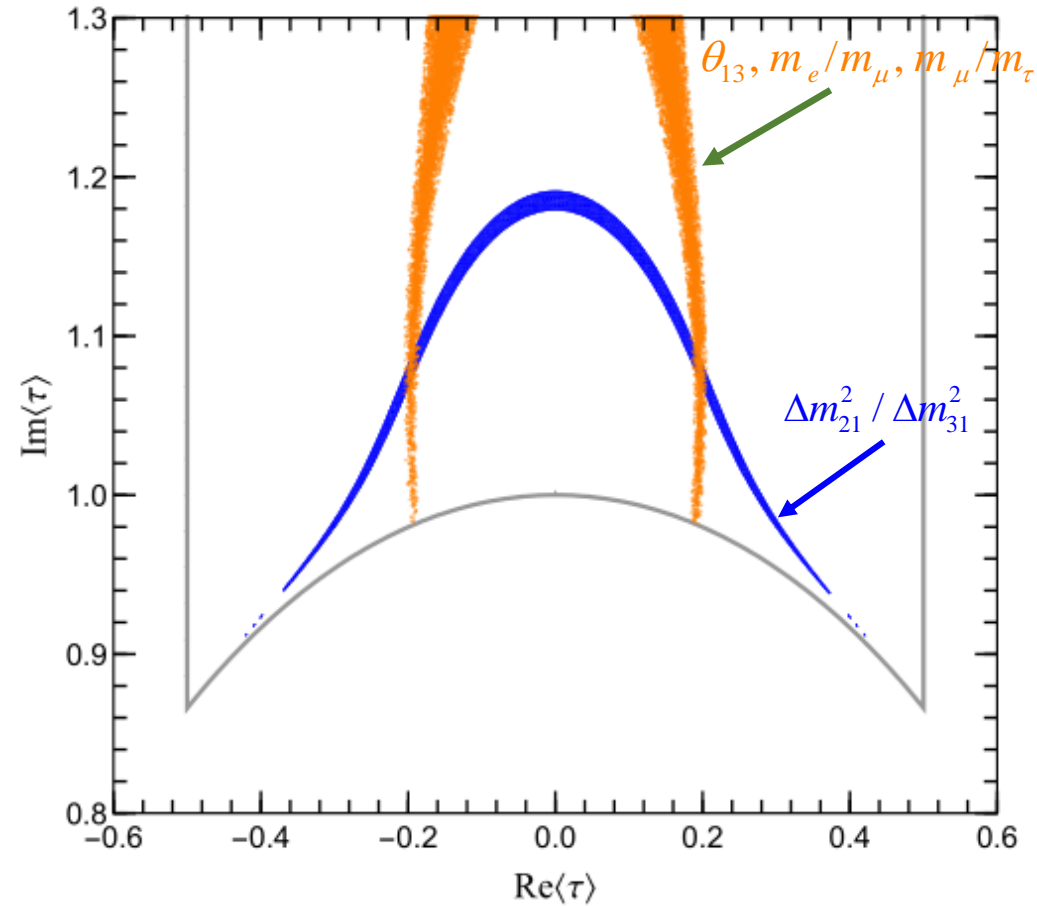
Minimal #p: $\alpha, \beta, \gamma, g^2/\Lambda$

$$\longrightarrow M_D = g \begin{pmatrix} 1 & 0 & 0 \\ 0 & 0 & 1 \\ 0 & 1 & 0 \end{pmatrix} v_u, \quad M_N = \begin{pmatrix} -2Y_{\mathbf{2},1}^{(2)} & 0 & 0 \\ 0 & \sqrt{3}Y_{\mathbf{2},2}^{(2)} & Y_{\mathbf{2},1}^{(2)} \\ 0 & Y_{\mathbf{2},1}^{(2)} & \sqrt{3}Y_{\mathbf{2},2}^{(2)} \end{pmatrix} \Lambda$$

Light neutrino mass

$$m_1 = \frac{1}{|2Y_{2,1}^{(2)}|} \frac{g^2 v_u^2}{\Lambda}, \quad m_2 = \frac{1}{|Y_{2,1}^{(2)} - \sqrt{3}Y_{2,2}^{(2)}|} \frac{g^2 v_u^2}{\Lambda}, \quad m_3 = \frac{1}{|Y_{2,1}^{(2)} + \sqrt{3}Y_{2,2}^{(2)}|} \frac{g^2 v_u^2}{\Lambda}$$

only depends on modulus τ up to overall scale



Neutrino mass spectrum is normal ordering

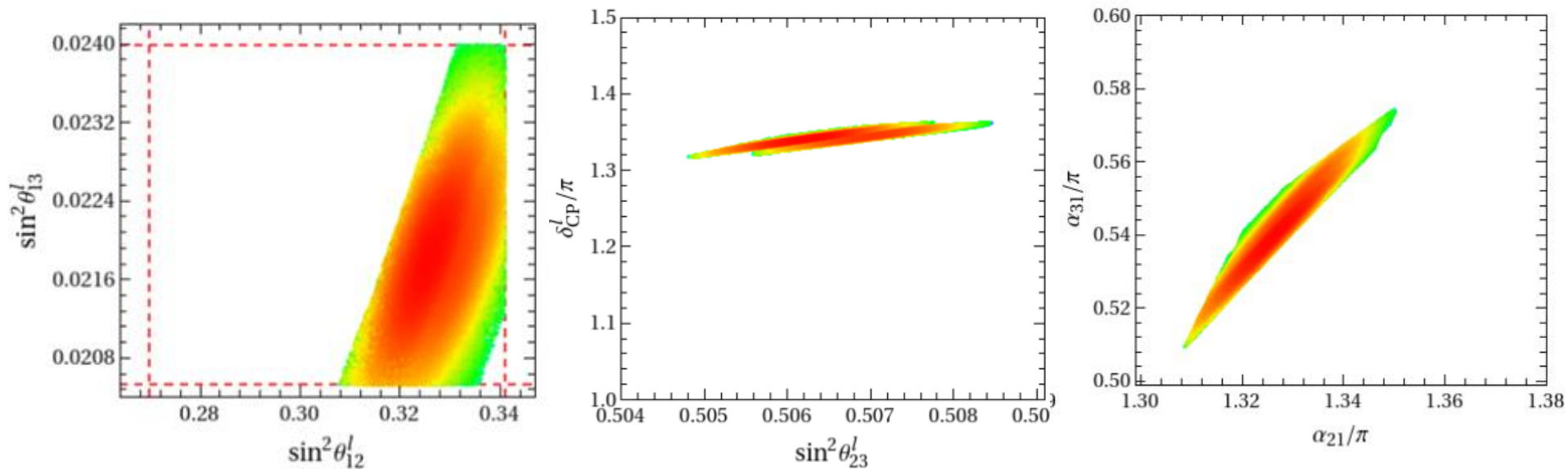
Minimal: only 4 real couplings plus modulus τ can explain **12 observables**

$$\langle \tau \rangle = -0.1921 + 1.0854i, \quad \beta / \alpha = 0.7159, \quad \gamma / \alpha = 87.4471,$$

$$\alpha v_d = 0.02881 \text{ MeV}, \quad g^2 v_u^2 / \Lambda = 71.8888 \text{ meV}$$

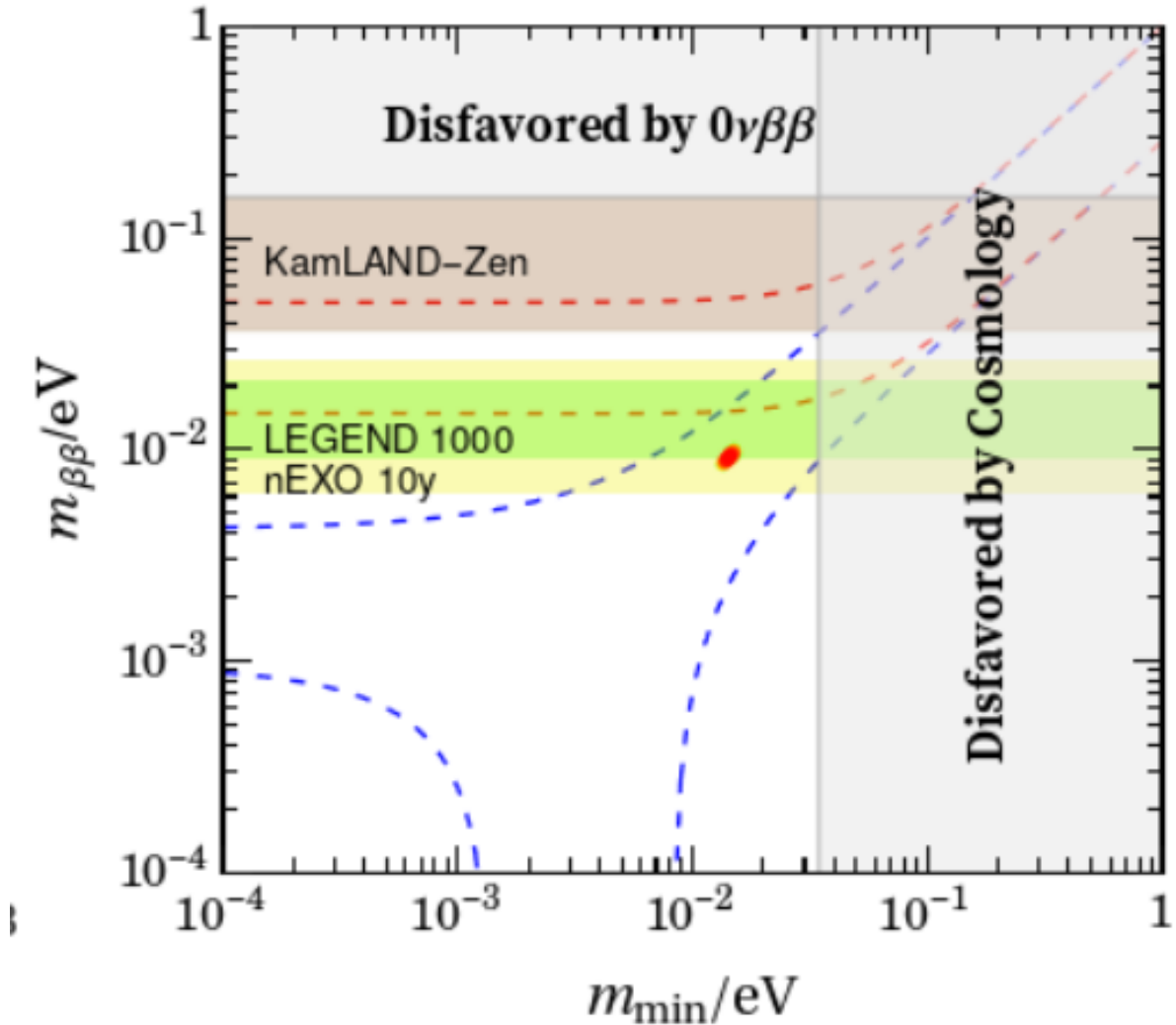
τ is the unique source breaking both modular and CP symmetries. All observables are within the 3σ regions

$$\sin^2 \theta_{12} = 0.3261, \quad \sin^2 \theta_{13} = 0.02182, \quad \sin^2 \theta_{23} = 0.5063, \quad \delta_{CP} = 1.34\pi,$$
$$\alpha_{21} = 1.3268\pi, \quad \alpha_{31} = 0.5401\pi, \quad m_e / m_\mu = 0.004737, \quad m_\mu / m_\tau = 0.05876,$$
$$m_1 = 14.27 \text{ meV}, \quad m_2 = 16.67 \text{ meV}, \quad m_3 = 51.64 \text{ meV}$$



The effective neutrino masses:


$$m_\beta = 16.76 \text{ meV}, \quad m_{\beta\beta} = 9.17 \text{ meV}$$



within the sensitivity of future $0\nu\beta\beta$ experiments

Extension to quark sector

	$Q_D = (Q_1, Q_2)$	Q_3	$U_D^c = (u^c, c^c)$	t^c	$D_D^c = (d^c, s^c)$	b^c
$2O$	2	1'	$\widehat{2}'$	1'	2	1'
k_I	k_{Q_D}	k_{Q_D}	$3 - k_{Q_D}$	$6 - k_{Q_D}$	$6 - k_{Q_D}$	$-k_{Q_D}$



$$M_u = \begin{pmatrix} \alpha_u Y_{\widehat{4},3}^{(3)} & -\alpha_u Y_{\widehat{4},2}^{(3)} & \boxed{0} \\ \alpha_u Y_{\widehat{4},4}^{(3)} & \alpha_u Y_{\widehat{4},1}^{(3)} & \boxed{0} \\ -\beta_u Y_{2,2}^{(6)} & \beta_u Y_{2,1}^{(6)} & \gamma_u Y_1^{(6)} \end{pmatrix} v_u,$$

$$M_d = \begin{pmatrix} \alpha_d Y_1^{(6)} - \gamma_d Y_{2,1}^{(6)} & \beta_d Y_{1'}^{(6)} + \gamma_d Y_{2,2}^{(6)} & -\delta_d Y_{2,2}^{(6)} \\ \gamma_d Y_{2,2}^{(6)} - \beta_d Y_{1'}^{(6)} & \alpha_d Y_1^{(6)} + \gamma_d Y_{2,1}^{(6)} & \delta_d Y_{2,1}^{(6)} \\ \boxed{0} & \boxed{0} & \varepsilon_d \end{pmatrix} v_d$$

The complex modulus τ is common in both quark and lepton sectors, and its value is fixed by the lepton parameters

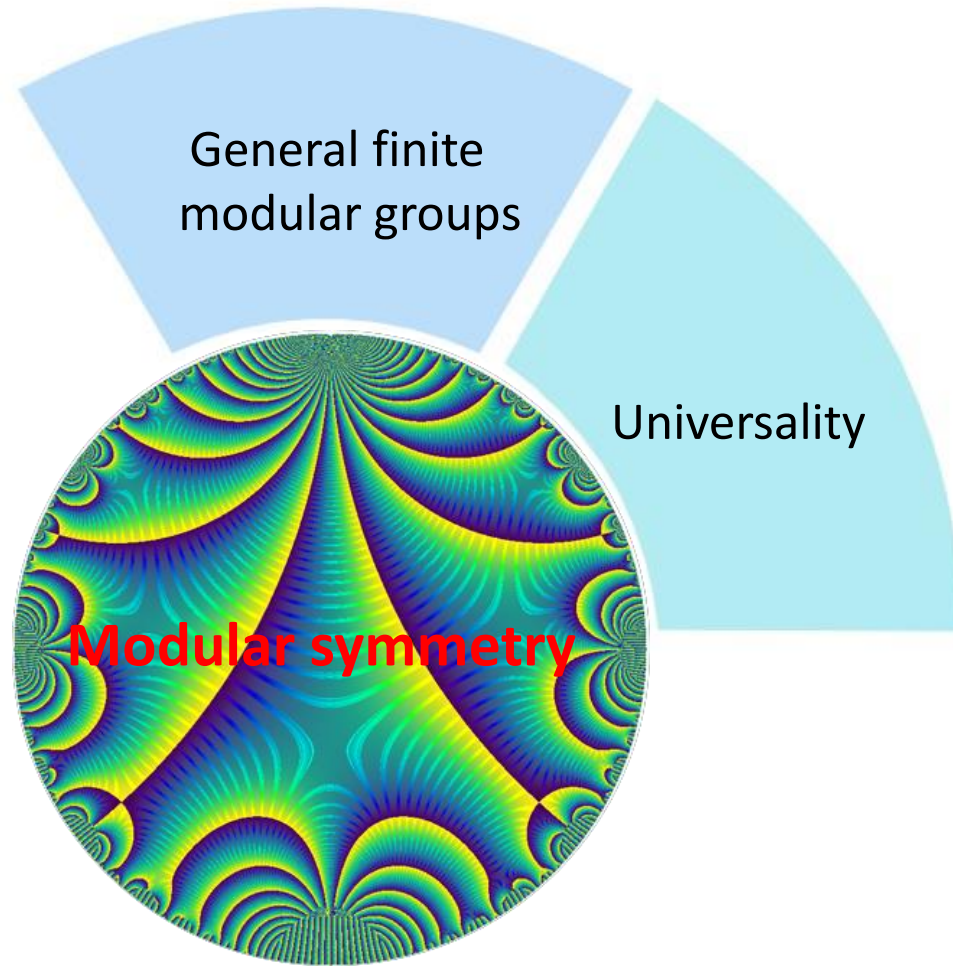
$$\langle \tau \rangle = -0.1946 + 1.0799i$$

The quark masses and CKM mixing parameters can be well accommodated with $\chi_q^2 = 6.4$:

$$\theta_{12}^q = 0.229, \quad \theta_{13}^q = 0.00393, \quad \theta_{23}^q = 0.0388, \quad \delta_{CP}^q = 61.27^\circ,$$

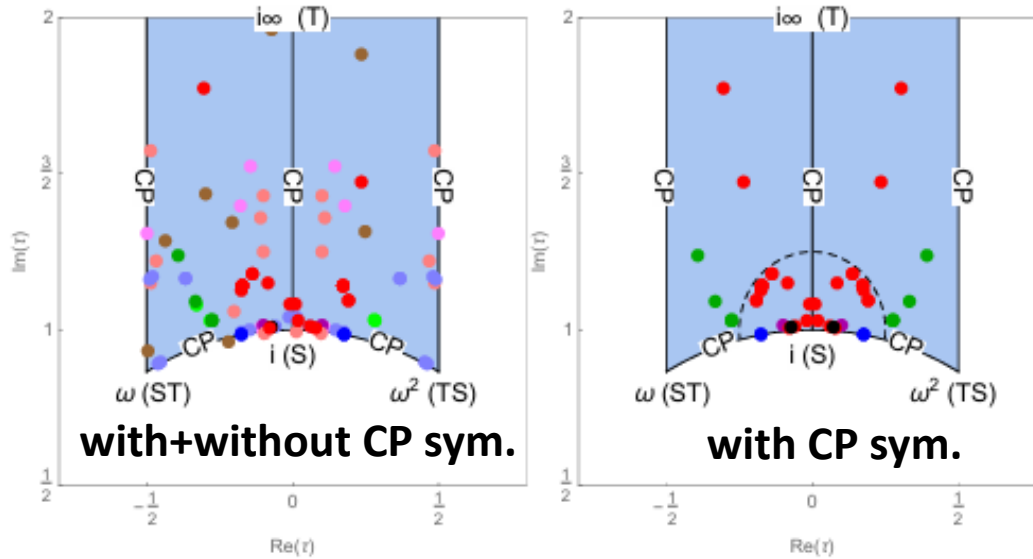
$$m_u / m_c = 0.00243, \quad m_c / m_t = 0.00245, \quad m_d / m_s = 0.0510, \quad m_s / m_b = 0.0234$$

- The model uses **14** parameters to describe the masses and mixing of both quark and lepton sectors: **12** masses+**6** mixing angles+**4** CP phases.



Universality of modular invariant models around fixed points

Many modular invariant bottom-up models differ in the level N , the representation ρ_I and weight k_I assignments of matter fields.



[Feruglio, 2211.00659,2302.11580]

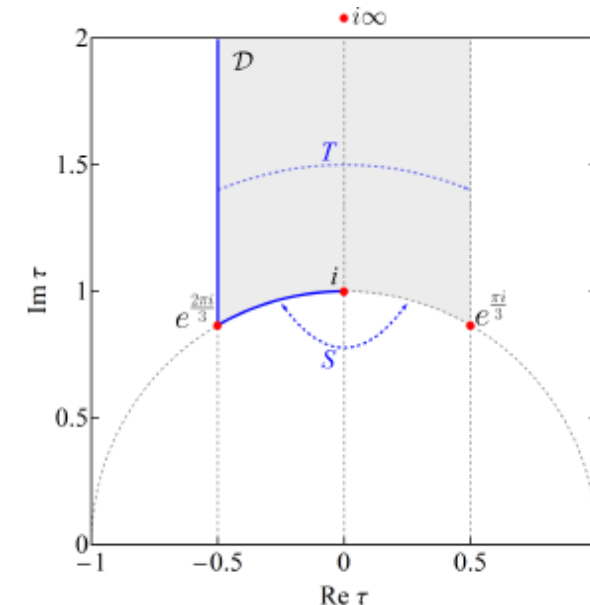
- At certain values of τ , modular symmetry is partially broken and some residual symmetry remains.

➤ There are only 3 inequivalent modular fixed points in the fundamental domain

$$\gamma_0 \tau_0 = \tau_0 \quad \longrightarrow \quad \tau_0 = i, e^{2\pi i/3}, i\infty$$

Fixed points of modular symmetry:

τ_0	γ_0	invariant under	Residual sym. G_0
i	S	$\tau \xrightarrow{S} -\frac{1}{\tau}$	\mathbb{Z}_4^S
$e^{2\pi i/3}$	ST	$\tau \xrightarrow{ST} -\frac{1}{\tau+1}$	\mathbb{Z}_3^{ST}
$i\infty$	T	$\tau \xrightarrow{T} \tau + 1$	\mathbb{Z}_N^T



$$\Omega(\gamma_0) \equiv (c_0\tau_0 + d_0)^{-k} \rho(\gamma_0)$$

$$\left. \begin{aligned} \tau_0 = i, e^{2\pi i/3}: u = \frac{\tau - \tau_0}{\tau - \tau_0^*}, \Phi = (1 - u)^k \psi \\ \tau_0 = i\infty: u = e^{\frac{2\pi i\tau}{N}}, \Phi = \psi \end{aligned} \right\} \longrightarrow \begin{cases} u \xrightarrow{\gamma_0} \frac{a_0 - c_0\tau_0}{a_0 - c_0\tau_0^*} u, & u \xrightarrow{CP} u^* \\ \Phi \xrightarrow{\gamma_0} \Omega(\gamma_0)\Phi, & \Phi \xrightarrow{CP} \Phi^* \end{cases}$$

The residual symmetry acts as traditional flavor abelian symmetry broken by the VEV of u

➤ Fermion mass terms in terms of new variable **see talk by Joao Penedo**

$$\mathcal{L}_m = \Phi_i^c m_{ij}(u, u^*) \Phi_j \longrightarrow \begin{cases} m(\gamma_0 u, \gamma_0 u^*) = \Omega^{c*} m(u, u^*) \Omega \\ m(u^*, u) = m^*(u, u^*) \end{cases} \quad \text{modular invariance under residual } \gamma_0 \text{ and CP}$$

Expanding $m_{ij}(\tau)$ in powers of u in the vicinity of τ_0

$$m_{ij}(u, u^*) = m_0 [x_{ij}^0 + x_{ij}^{10} u + x_{ij}^{01} u^* + x_{ij}^{20} u^2 + x_{ij}^{11} uu^* + x_{ij}^{02} u^{*2} + \dots]$$

- The coefficients $x_{ij}^0, x_{ij}^{10}, x_{ij}^{01}, \dots$ are assumed to be of order one, normalization of modular form is relevant. **[Petcov,2311.04185]**
- Invariance under the residual symmetry γ_0 allows to fix the pattern of mass matrix
- Only need $\Omega(\gamma_0)$ of LH leptons which are assumed to be irreducible triplet of Γ'_N
- **Universality:** independent of level, the modular weight of matter fields and the kinetic terms

➤ The vicinity of $\tau = i$ is preferred by neutrino oscillation [Feruglio, 2302.11580]

τ	mass ordering	$m_\nu(0,0)$	regular	NO/IO	$\frac{\Delta m_{sol}^2}{\Delta m_{atm}^2}$	$\sin^2 \theta_{12}$	$\sin^2 \theta_{13}$	$\sin^2 \theta_{23}$
$\approx i$	k_S even	$m_\nu(0,0)$	regular	NO/IO	$\mathcal{O}(1)$	$\mathcal{O}(x^2)$	$\mathcal{O}(x^2)$	$\mathcal{O}(1)$
$\approx i$	k_S odd	$m_\nu(0,0)$	regular	IO	$\mathcal{O}(x)$	$\frac{1}{2}(1 + \mathcal{O}(x))$	$\mathcal{O}(x^2)$	$\mathcal{O}(1)$
$\approx i$	k_S odd	$m_\nu(0,0)$	singular	NO	$\mathcal{O}(x^3)$	$\frac{1}{2}(1 + \mathcal{O}(x))$	$\mathcal{O}(x^2)$	$\mathcal{O}(1)$
$\approx \omega$				NO/IO	$\mathcal{O}(x)$	$\frac{1}{2}(1 + \mathcal{O}(x))$	$\mathcal{O}(x^2)$	$\mathcal{O}(x^2)$

$$x = |u|$$

$$m_\nu^{-1} = m_{0\nu}^{-1} \begin{pmatrix} x_{11} & x & x_{12}^0 & x_{13}^0 \\ \cdot & x_{22} & x & x_{23} & x \\ \cdot & \cdot & x_{33} & x \end{pmatrix} + \mathcal{O}(x^2)$$

[Ding, Feruglio, Liu, 2402.14915]

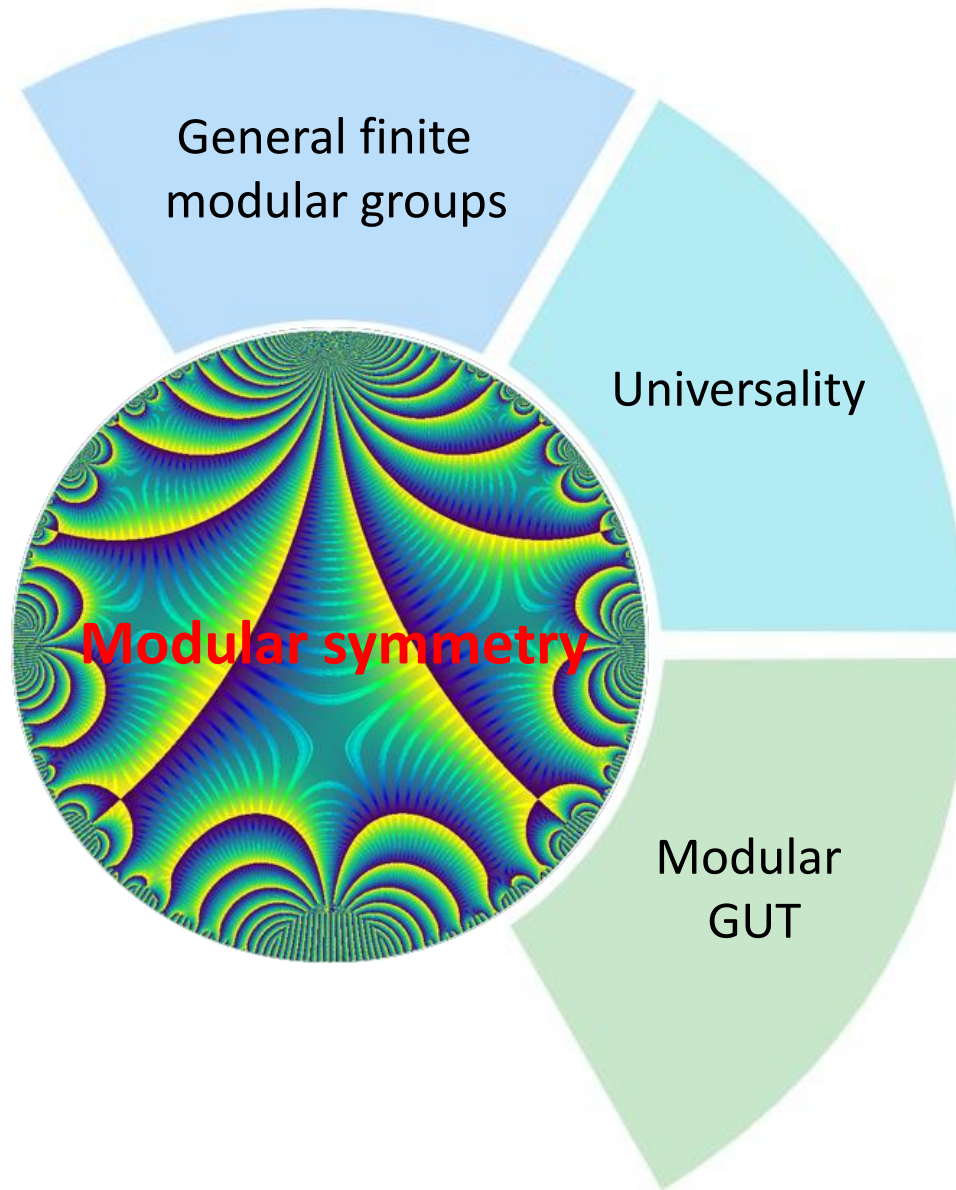
The same universal behavior in Symplectic modular invariant theories near the fixed points.

- Charged lepton mass hierarchies require tuning the order one coefficients for $\tau \approx i$, they can be naturally generated for $\tau \approx e^{2\pi i/3}, i\infty$, large lepton angles necessitates singlet assignment of LH leptons, but predictivity is reduced somewhat. [Novichkov, Penedo, Petcov, 2102.07488]

• Predictive model for both charged lepton mass hierarchies and neutrino mixing without fine-tuning? both quark and lepton sectors without fine-tuning?

• Critical behavior similar to phase transition or cosmological evolution?

[Feruglio, 2302.11580]

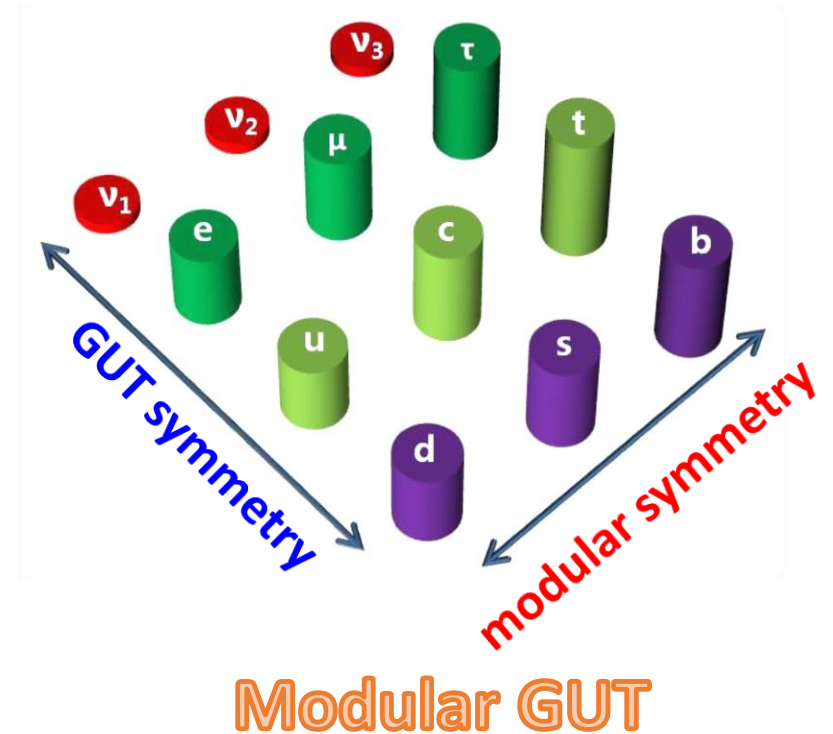
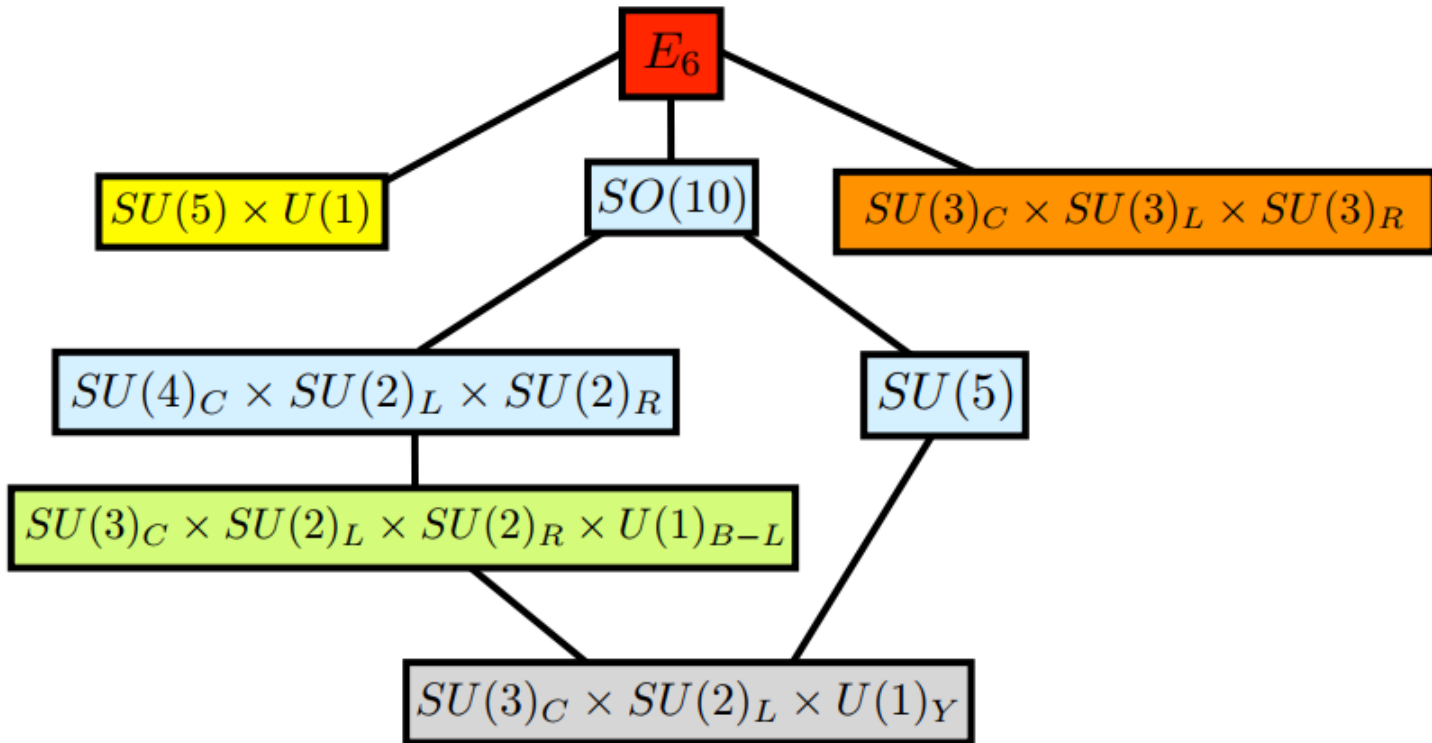


Modular GUTs

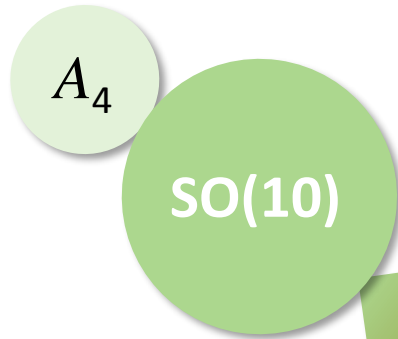
The symmetry group of modular GUT is $G_{\text{GUT}} \times \Gamma_N$ or $G_{\text{GUT}} \times \Gamma'_N$

- **GUTs:** connecting quarks and leptons leptons \longleftrightarrow quarks
- **Modular symmetry:** relating three families and Yukawa couplings are modular forms

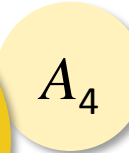
➤ Candidates of GUT gauge group G_{GUT}



Ding, King, Lu, 2108.09655;
Ding, King, Lu, Qu, 2206.14675



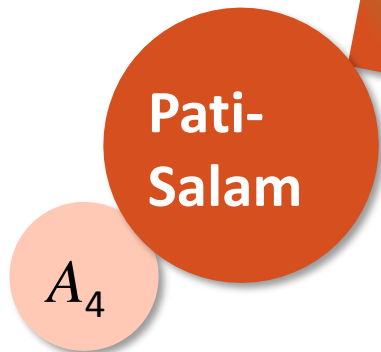
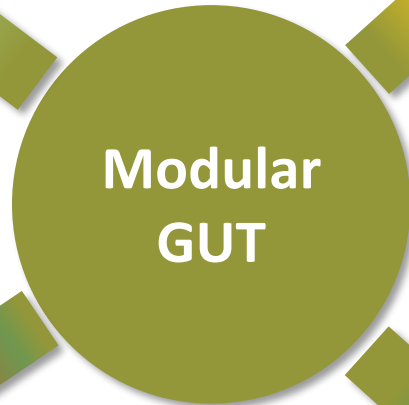
Kobayashi, Shimizu, Takagi, Tanimoto,
Tatsuishi, 1906.10341;
Du, Wang, 2012.01397



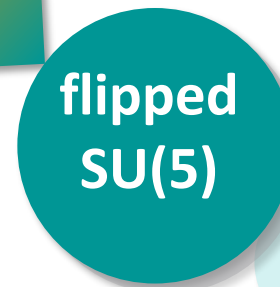
Anda, King, Perdomo, 1812.05620;
Chen, Ding, King, 2101.12724



Zhao, Zhang, 2101.02266;
King, Zhou, 2103.02633;
Ding, King, Yao, 2103.16311;
Varzielas, King, Levy, 2309.15901



Ding, Jiang, King, Lu, Qu, 2404.06520

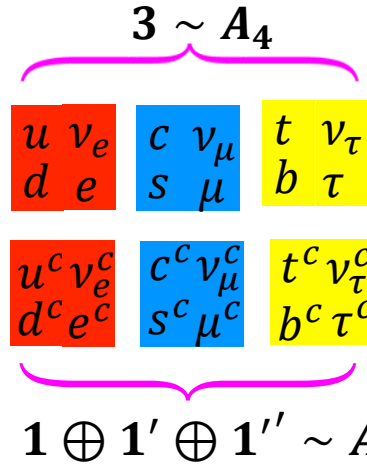


Charalampous, King, Leontaris, Zhou, 2109.11379;
Du, Wang, 2209.08796

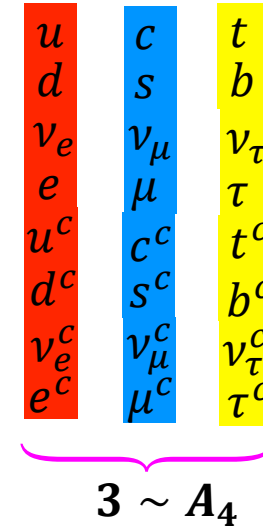
- GUTs unify the quark and leptons fields into few multiplets, the possible transformations of matter fields under modular symmetry is quite limited.

u	ν_e	c	ν_μ	t	ν_τ
d	e	s	μ	b	τ
u^c	ν_e^c	c^c	ν_μ^c	t^c	ν_τ^c
d^c	e^c	s^c	μ^c	b^c	τ^c

SM: $3 \times 6 = 18$ multiplets



Pati-Salam: $3 \times 2 = 6$ multiplets



SO(10): $3 \times 1 = 3$ multiplets

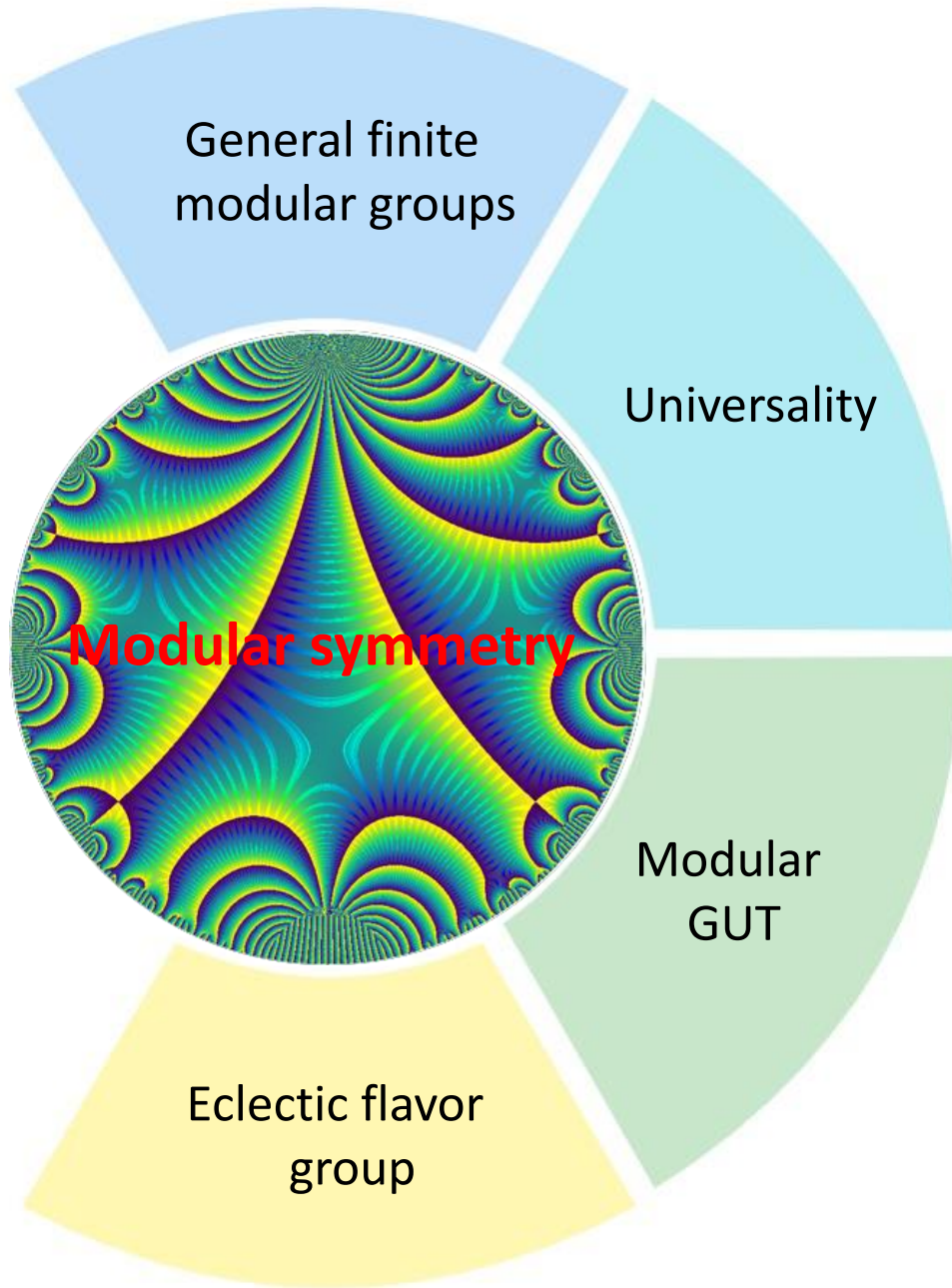
- The modular symmetry can **strongly constrain** the GUT Yukawa couplings and thus enhance the predictive power, but the complexity of GUT breaking to SM remains

SO(10) fermion mass matrices:

$$\begin{aligned}
 M_u &= (\mathcal{Y}^{10} + r_2 \mathcal{Y}^{\overline{126}} + r_3 \mathcal{Y}^{120}) v_u, & M_d &= r_1 (\mathcal{Y}^{10} + \mathcal{Y}^{\overline{126}} + \mathcal{Y}^{120}) v_d, \\
 M_\ell &= r_1 (\mathcal{Y}^{10} - 3\mathcal{Y}^{\overline{126}} + c_e \mathcal{Y}^{120}) v_d, & M_{\nu_D} &= (\mathcal{Y}^{10} - 3r_2 \mathcal{Y}^{\overline{126}} + c_\nu \mathcal{Y}^{120}) v_u, \\
 M_{\nu_R} &= v_R \mathcal{Y}^{\overline{126}}, & M_{\nu_L} &= v_L \mathcal{Y}^{\overline{126}}.
 \end{aligned}$$

$r_{1,2,3}$ and $c_{e,\mu}$: mixture of SO(10) Higgs multiplets into SM Higgs

- Modular GUT for flavor puzzle so far, the interplay of modular symmetry with proton decay, GW ?



General finite
modular groups

Universality

Modular symmetry

Modular
GUT

Eclectic flavor
group

Kähler potential problem in modular symmetry

- Kähler potential not fixed by modular flavor symmetry

$$\mathcal{K} = (-i\tau + i\bar{\tau})^{-k_\psi} (\psi^\dagger \psi)_1 + \sum_{n, r_1, r_2} c^{(n, r_1, r_2)} (-i\tau + i\bar{\tau})^{-k_\psi + n} (\psi^\dagger Y_{r_1}^{(n)\dagger} Y_{r_2}^{(n)} \psi)_1$$

minimal Kähler potential

non-canonical terms

many non-canonical terms on the same footing as the minimal Kähler potential

- Modifying Kähler metric and kinetic terms [\[Chen, Ramon-Sanchez, Ratz, 1909.06910\]](#)

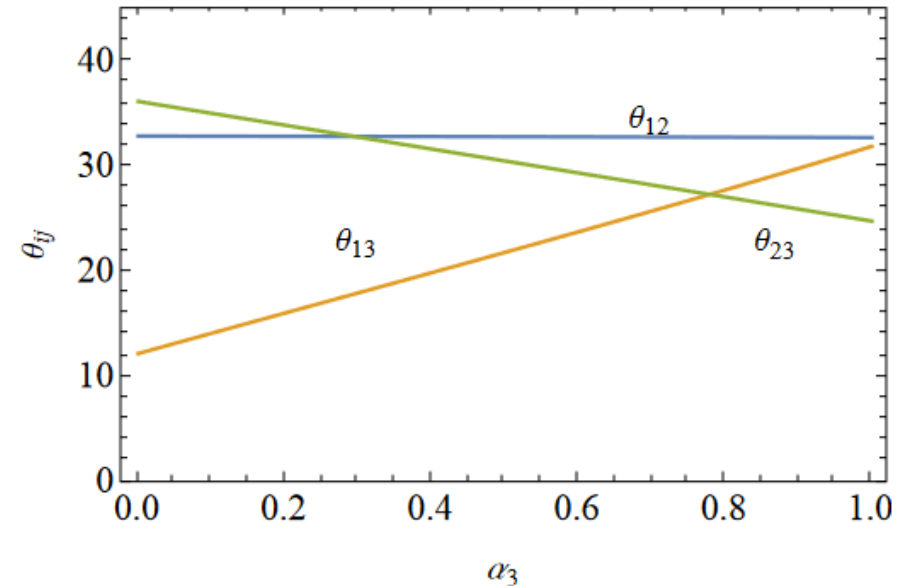
$$\mathcal{K}_\psi^{ij} = \frac{\partial^2 \mathcal{K}}{\partial \psi_i^\dagger \partial \psi_j} = \underbrace{\langle -i\tau + i\bar{\tau} \rangle^{-k} \delta^{ij}}_{\text{flavor universal}} + \underbrace{\Delta \mathcal{K}_\psi^{ij}}_{\text{flavor off-diagonal}}$$

- Go to canonical basis

$$\mathcal{K}_\psi(\tau, \tau^*) \mathcal{K} = Z_\psi^{-1\dagger}(\tau, \tau^*) Z_\psi^{-1}(\tau, \tau^*)$$

$$\mathcal{K}_{\psi^c}(\tau, \tau^*) \mathcal{K} = Z_{\psi^c}^{-1\dagger}(\tau, \tau^*) Z_{\psi^c}^{-1}(\tau, \tau^*)$$

$$m(\tau) \rightarrow Z_{\psi^c}^T(\tau, \tau^*) m(\tau) Z_\psi(\tau, \tau^*) \rightarrow \text{sizeable corrections to mixing parameters}$$



Eclectic flavor group as a solution to Kähler potential problem

- Modular flavor symmetries from top-down approach (orbifold string compactification) gives
 - Normal symmetries of extra dimensions → traditional flavor symmetries
 - String duality transformations → modular flavor symmetries
 - The multiplicative closure of these groups is defined as the eclectic flavor group

[Nilles, Ramos-Sanchez, Vaudrevange, 2001.01736; 2004.05200]

$$G_{\text{eclectic}} = G_{\text{traditional}} \cup G_{\text{modular}}$$

see talk by Saul Ramos-Sanchez

- Traditional flavor symmetry vs. modular symmetry transformations

modular: $\tau \xrightarrow{\gamma} \gamma\tau = \frac{a\tau + b}{c\tau + d}, \psi \xrightarrow{\gamma} (c\tau + d)^{-k} \rho(\gamma)\psi, \gamma \in SL(2, \mathbb{Z})$

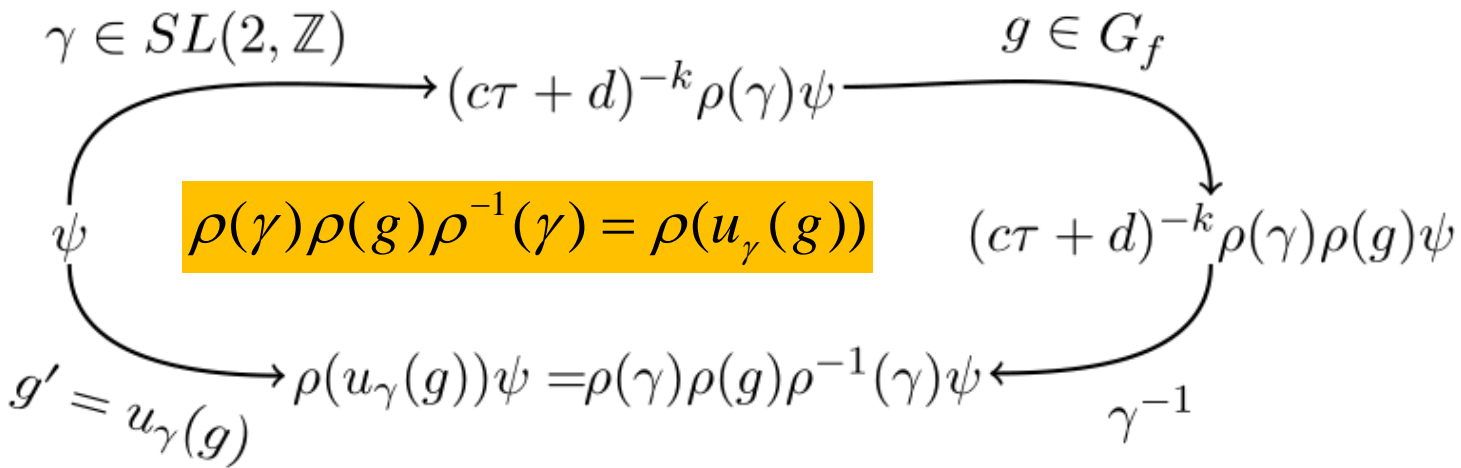
flavor: $\tau \xrightarrow{g} \tau, \psi \xrightarrow{g} \rho(g)\psi, g \in G_f$

τ distinguishes flavor symmetry from modular symmetry

	Traditional flavor symmetry	Modular symmetry	Eclectic flavor Group
Kähler potential			
Superpotential			
Vacuum problem			

- The matter fields transform nontrivially under both flavor and modular symmetries, the Kähler potential is constrained
- Enhanced flavor symmetry at the modular symmetry fixed points $\tau = i, e^{2\pi i/3}, i\infty$
- The notorious flavons reoccur to break flavor symmetry

➤ Consistency condition [Nilles, Ramos-Sanchez, Vaudrevange, 2001.01736]



$$\begin{aligned}
 \gamma &\rightarrow u_\gamma \\
 S &\rightarrow u_S \\
 T &\rightarrow u_T \\
 S^{N_s} = 1 &\rightarrow (u_S)^{N_s} = 1 \\
 T^N = 1 &\rightarrow (u_T)^{N_T} = 1 \\
 (ST)^3 = 1 &\rightarrow (u_S \circ u_T)^3 = 1 \\
 S^2 T = T S^2 &\rightarrow (u_S)^2 \circ u_T = u_T \circ (u_S)^2
 \end{aligned}$$

$$N_s = 4 \text{ (2) for } \Gamma'_N \text{ (}\Gamma_N\text{)}$$

Finite modular group Γ'_N (Γ_N) must be a subgroup of the outer automorphism group of G_f

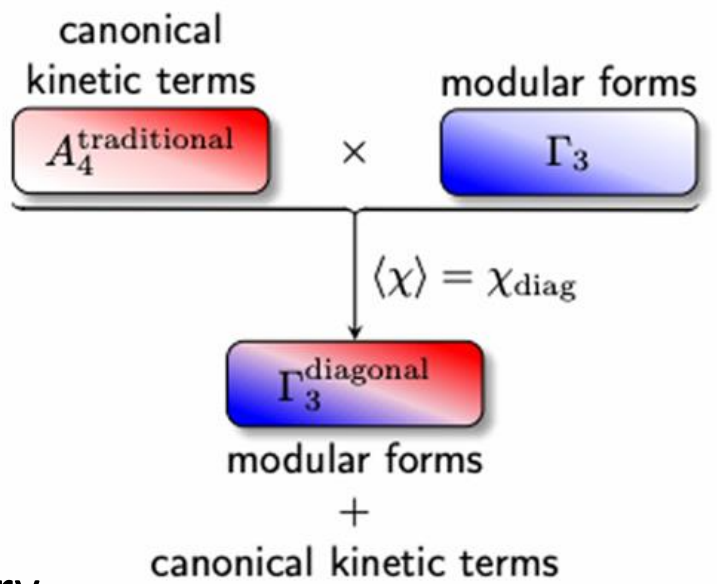
$$G_{\text{eclectic}} \cong G_f \rtimes \Gamma'_N \text{ (} G_f \rtimes \Gamma_N \text{)}$$

➤ Quasi-eclectic flavor symmetry: $u_\gamma(g) = g$

[Chen, Perez, Hamud, Sanchez, Ratz, Shukla, 2108.02240]

$$G_{\text{Quasi-eclectic}} \cong G_f \times \Gamma'_N \text{ (} G_f \times \Gamma_N \text{)}$$

- Flavor symmetry and modular symmetry are commutable
- No constraint on the choice of G_f and Γ'_N (Γ_N)
- Non-canonical Kähler potentials are suppressed by flavor symmetry



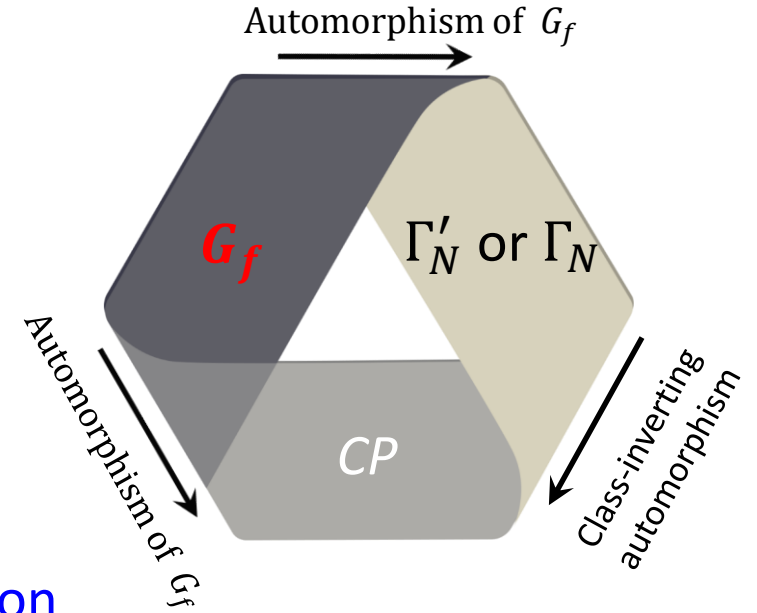
➤ Eclectic flavor group can combine with CP: **unification of flavor, CP and modular symmetries**

CP transformation: $\psi(x) \xrightarrow{CP} X_r \bar{\psi}(x_P)$

$$X_r \rho_r^*(g) X_r^{-1} = \rho_r(g')$$

$$X_r \rho_r^*(S) X_r^{-1} = \rho_r(g_1 S^{-1}),$$

$$X_r \rho_r^*(T) X_r^{-1} = \rho_r(g_2 T^{-1}), \quad g', g_1, g_2 \in G_f$$



➤ Possible eclectic flavor groups: a few G_f suitable to eclectic extension

flavor group G_f	GAP ID	Aut(G_f)	finite modular groups		eclectic flavor group
Q_8	[8, 4]	S_4	without CP	S_3	GL(2, 3)
			with CP	-	-
$\mathbb{Z}_3 \times \mathbb{Z}_3$	[9, 2]	GL(2, 3)	without CP	S_3	$\Delta(54)$
			with CP	$S_3 \times \mathbb{Z}_2$	[108, 17]
A_4	[12, 3]	S_4	without CP	S_3	S_4
			with CP	S_4	S_4
T'	[24, 3]	S_4	without CP	-	-
			with CP	-	-
$\Delta(27)$	[27, 3]	[432, 734]	without CP	S_3	$\Delta(54)$
			with CP	T'	$\Omega(1)$
$\Delta(54)$	[54, 8]	[432, 734]	without CP	$S_3 \times \mathbb{Z}_2$	[108, 17]
			with CP	GL(2, 3)	[1296, 2891]
$\Delta(54)$	[54, 8]	[432, 734]	without CP	T'	$\Omega(1)$
			with CP	GL(2, 3)	[1296, 2891]

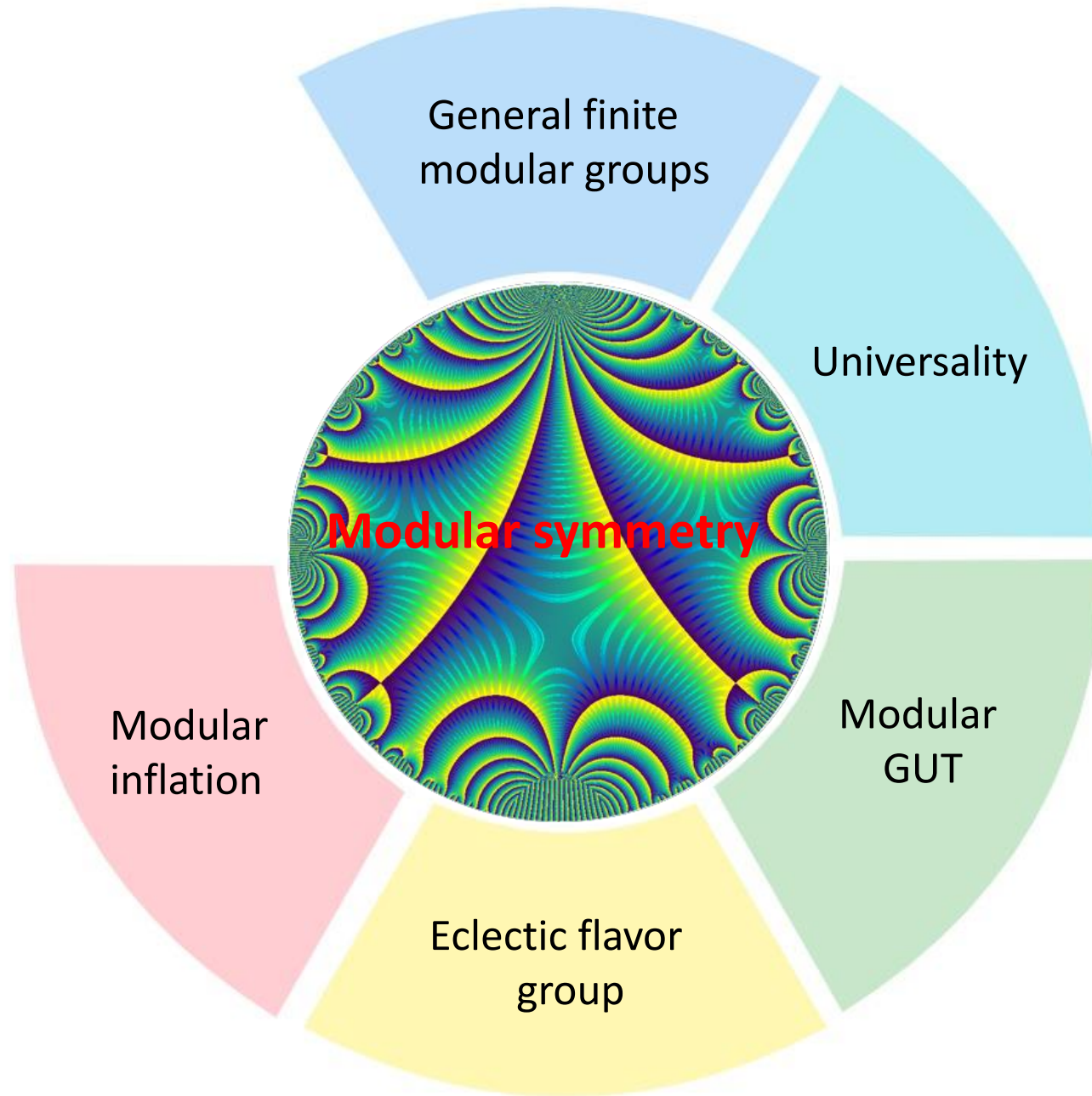
[Nilles,Ramos-Sanchez, Vaudrevange, et al, 2001.01736]

[Ding, Li, Lu,2405.xxxxx]

[Ding, Li, 2308.16901]

[Ding,King, Li, Liu,Lu, 2303.02071]

[Baur, Nilles et al,2207.10677]



Modular inflation

Accelerated expansion of the early universe favored by cosmic microwave background (CMB) observations

$$\ln(10^{10} A_s) = 3.044 \pm 0.014 \text{ (68\%CL)},$$

$$n_s = 0.9649 \pm 0.0042 \text{ (68\%CL)},$$

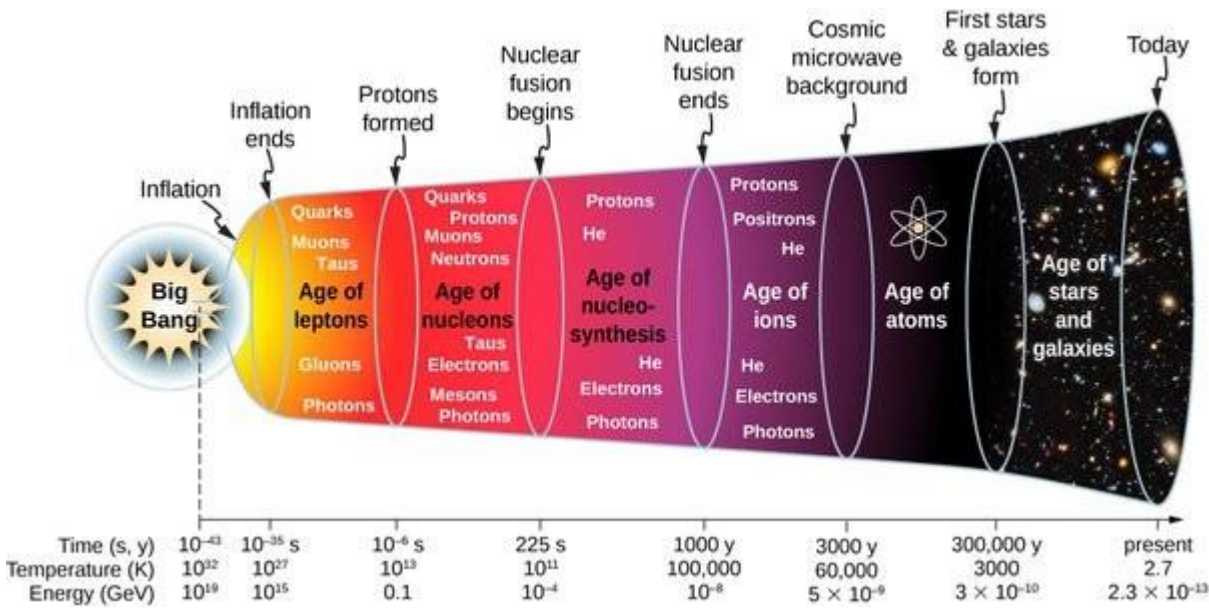
$$r < 0.036 \text{ (95\%CL)}.$$

A_s : scalar amplitude

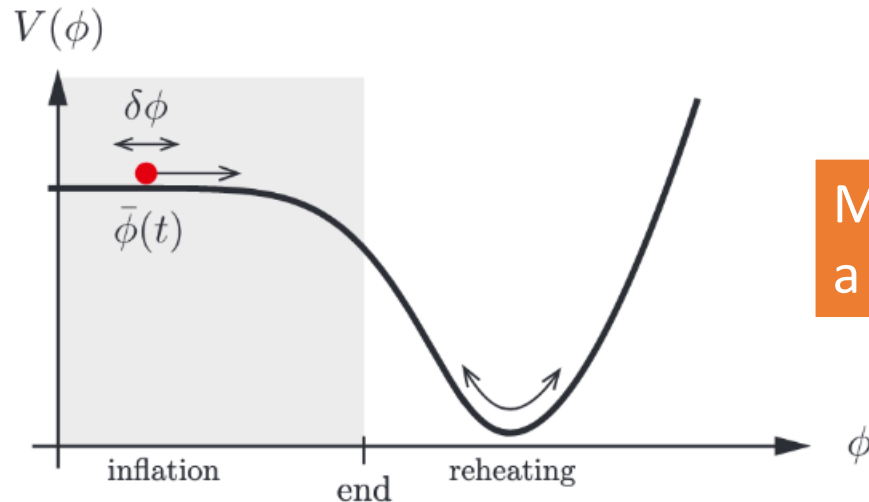
n_s : spectral index

r : tensor-to-scalar ratio

[Planck Collaboration, 1807.06211]



- Inflation can be realized by a slow-rolling scalar field (inflaton)



SM cannot explain inflation!

Modular inflation: modulus τ plays the role of inflaton, a plateau in the scalar potential is necessary

Modular invariant scalar potential of τ

- The modular invariant effective action in SUGRA for $\tau + S$ (dilaton)

[Cvetic et al., Nucl. Phys. B 361 (1991); Leedom, Righi, Westphal, 2212.03876]

$$\mathcal{G}(\tau, \bar{\tau}, S, \bar{S}) = \kappa^2 \mathcal{K}(\tau, \bar{\tau}, S, \bar{S}) + \ln |\kappa^3 \mathcal{W}(\tau, S)|^2$$

Kähler potential: $\kappa^2 \mathcal{K}(\tau, \bar{\tau}, S, \bar{S}) = K(S, \bar{S}) - 3 \ln [-i(\tau - \bar{\tau})],$
 $K(S, \bar{S}) = -\ln(S + \bar{S}) + \delta k(S, \bar{S})$

$$\kappa = \sqrt{8\pi G_N}$$

$j(\tau)$: j -invariant function

Superpotential: $\mathcal{W}(S, \tau) = \Lambda_W^3 \frac{\Omega(S) H(\tau)}{\eta^6(\tau)}, \quad H(\tau) = (j(\tau) - 1728)^{m/2} j(\tau)^{n/3} \mathcal{P}(j(\tau))$ $m, n = 0, 1, 2, \dots$

$\Omega(S)$ is technically arbitrary, $\mathcal{P}(j)$ is an arbitrary polynomial function of $j(\tau)$

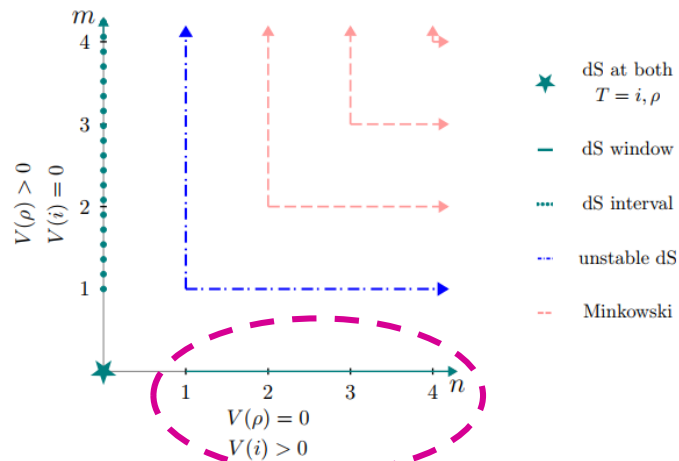


$$V(\tau, S) = \Lambda^4 e^{K(S, \bar{S})} Z(\tau, \bar{\tau}) |\Omega(S)|^2 \left[(A(S, \bar{S}) - 3) |H(\tau)|^2 + \widehat{V}(\tau, \bar{\tau}) \right]$$

- Vacuum structure at fixed points $\tau = i, e^{2\pi i/3} \equiv \rho$

see talk by Nicole Righi

[Leedom, Righi, Westphal, 2212.03876]



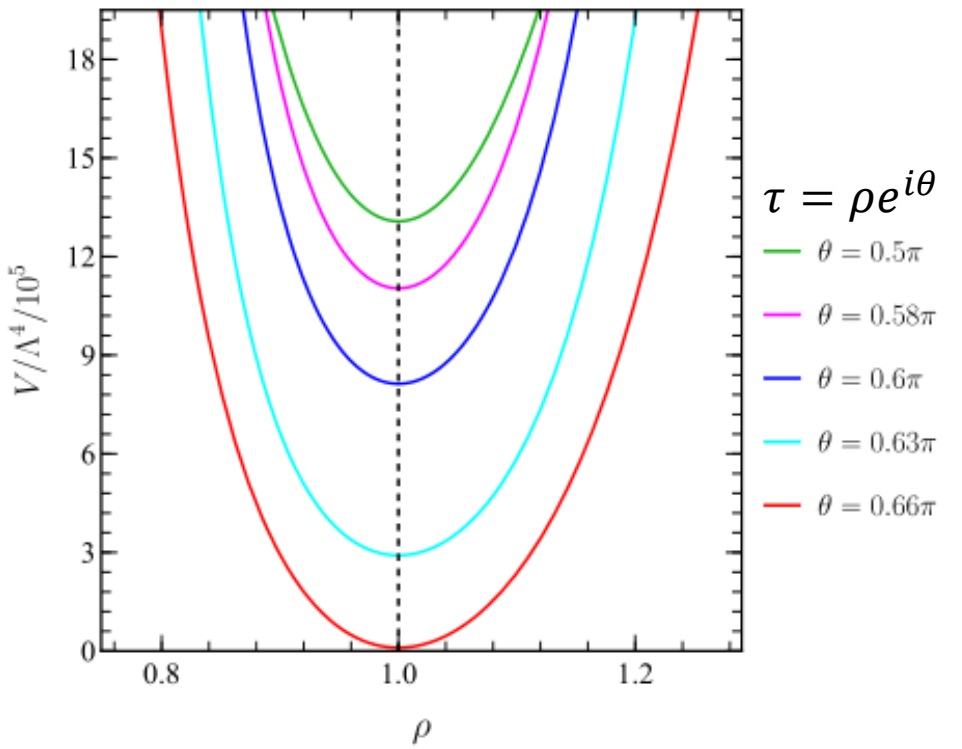
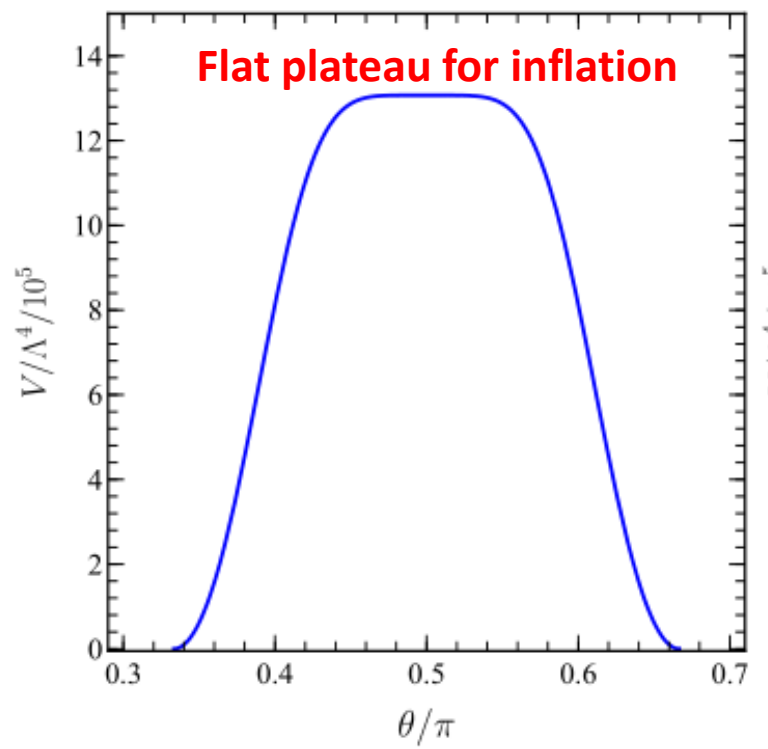
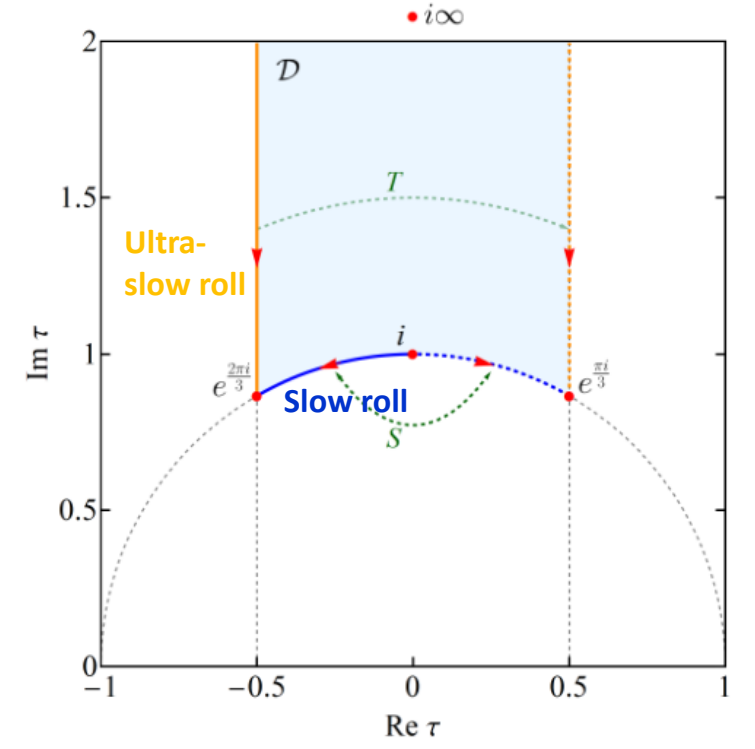
τ rolls toward the fixed points

➤ Slow roll along the unit arc $\tau = e^{i\theta}$ from $\tau = i$ to $\tau = e^{2\pi i/3}$

Hilltop-like inflation: $V(\phi) = V_0 [1 - C_2\phi^2 - C_4\phi^4 - C_6\phi^6 + \dots]$

Invariant under: $\phi \xrightarrow{S} -\phi$ $\phi = \sqrt{3/2}M_{\text{Pl}} \ln(\tan(\theta/2))$

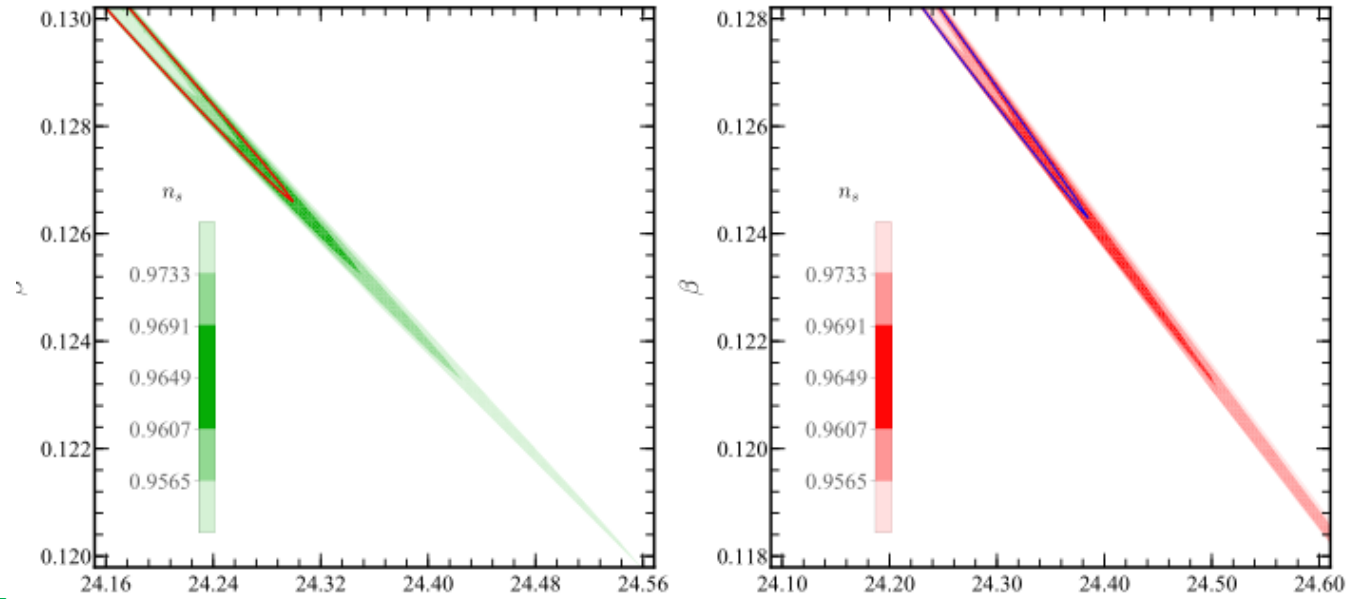
- For the power exponents $m = 0, n \geq 2, V(i) > 0, V(e^{2\pi i/3}) = 0$, the scalar potential reaches minimum at the unit arc along the radial direction, can only roll along angular direction



[Ding, Jiang, Zhao, 2405.06497; King, Wang, 2405.xxxxx]

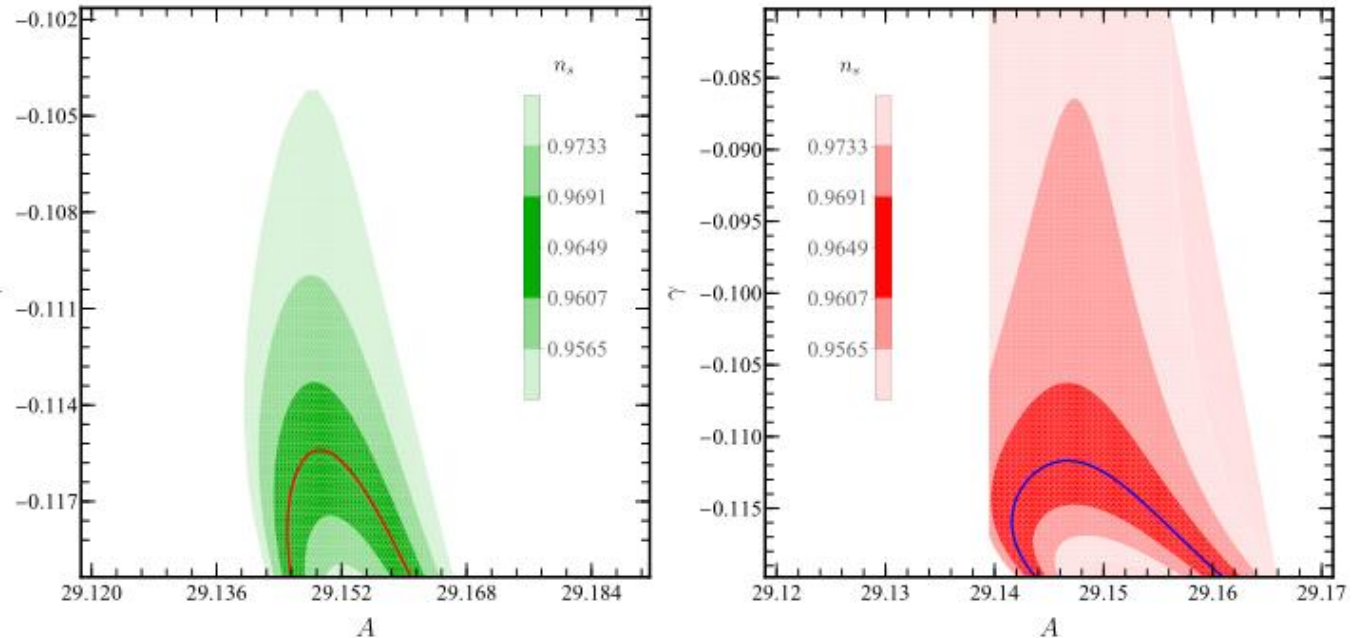
➤ Successful inflation can be reproduced, the tensor-to-scalar ratio $r < 10^{-6}$

(a) $\mathcal{P}(j) = 1 + \beta(1 - j/1728)$

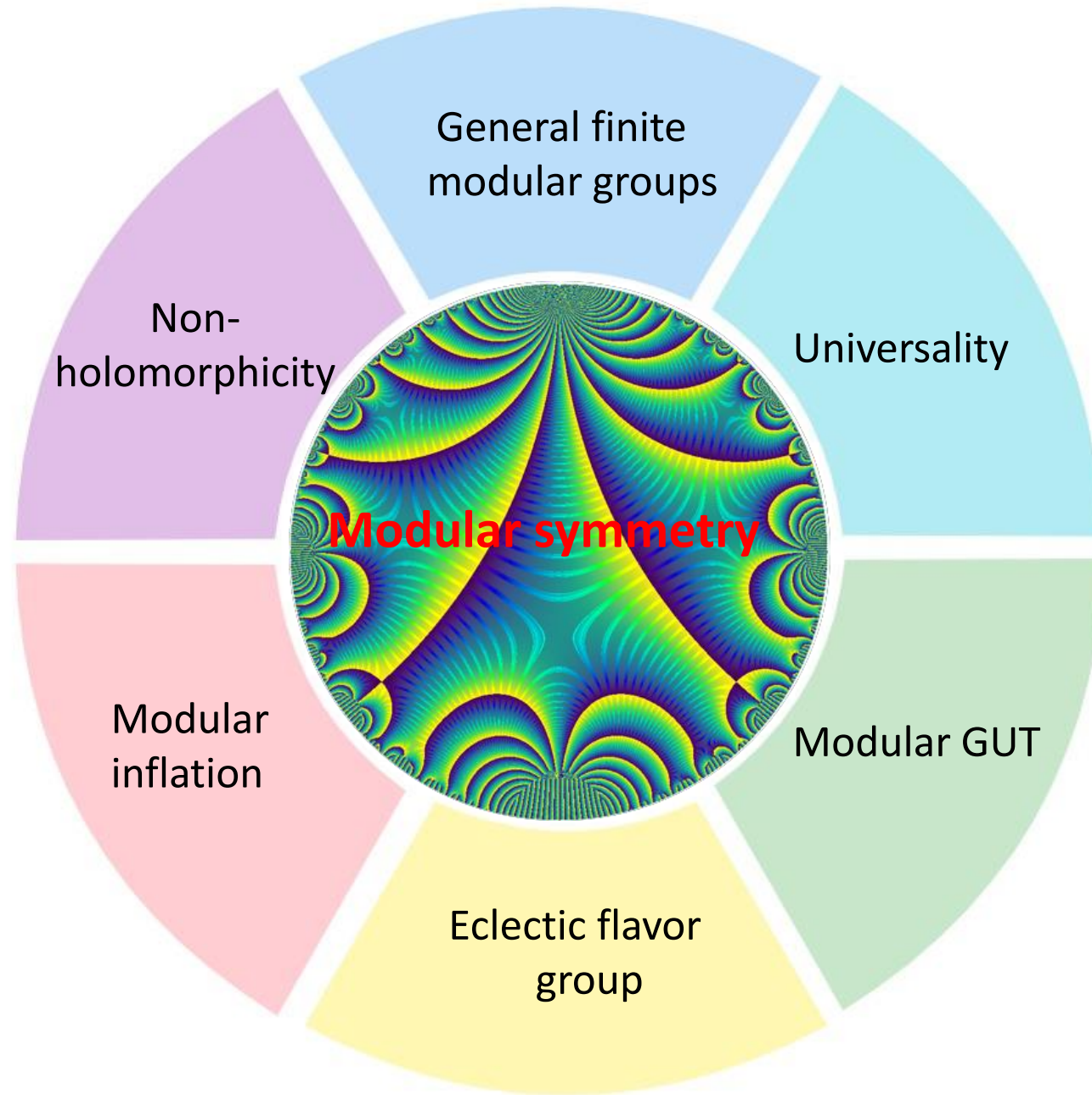


[Ding, Jiang, Zhao, 2405.06497;
King, Wang, 2405.xxxxx]

(b) $\mathcal{P}(j) = 1 + \gamma(1 - j/1728)^2$



[other scenarios of
modular inflation, see
Gunji, Ishiwata,
Yoshida, 2208.10086;
Abe, Higaki, Kaneko,
Kobayashi, Otsuka,
2303.02947...]



Non-holomorphic modular flavor symmetry

➤ Modular symmetry requires SUSY to protect holomorphicity of modular form

SUSY



holomorphicity



$Y(\tau)$: Yukawa couplings=modular form of level N and weight k_Y

ATLAS SUSY Searches* - 95% CL Lower Limits
August 2023

ATLAS Preliminary
 $\sqrt{s} = 13$ TeV

Model	Signature	$\mathcal{L}_{\text{int}} [\text{fb}^{-1}]$	Mass limit	Reference				
Inclusive Searches	$0, \mu, \mu$	2-8 jets	$A_{\text{eff}}^{\text{mono}} \rightarrow 140$	140	1.0, 1.85	$m(\tilde{g}) > 1000 \text{ GeV}$	2018.14293	
	$0, \mu, \mu$	1-3 jets	$A_{\text{eff}}^{\text{mono}} \rightarrow 140$	140	0.3	$m(\tilde{g}) > 1000 \text{ GeV}$	2103.10074	
	$0, \mu, \mu$	2-6 jets	$A_{\text{eff}}^{\text{mono}} \rightarrow 140$	140	2.2	$m(\tilde{g}) > 1000 \text{ GeV}$	2018.14293	
	$0, \mu, \mu$	2-6 jets	$A_{\text{eff}}^{\text{mono}} \rightarrow 140$	140	1.15-1.95	$m(\tilde{g}) > 1000 \text{ GeV}$	2018.14293	
	$1, \mu, \mu$	2-6 jets	$A_{\text{eff}}^{\text{mono}} \rightarrow 140$	140	2.2	$m(\tilde{g}) > 600 \text{ GeV}$	2101.01029	
	$0, \mu, \mu$	2 jets	$A_{\text{eff}}^{\text{mono}} \rightarrow 140$	140	2.2	$m(\tilde{g}) > 700 \text{ GeV}$	2004.13072	
	$0, \mu, \mu$	7-11 jets	$A_{\text{eff}}^{\text{mono}} \rightarrow 140$	140	1.97	$m(\tilde{g}) > 600 \text{ GeV}$	2008.06032	
	$0, \mu, \mu$	6 jets	$A_{\text{eff}}^{\text{mono}} \rightarrow 140$	140	1.15	$m(\tilde{g}) > 200 \text{ GeV}$	2007.01094	
	$0, \mu, \mu$	3-6 jets	$A_{\text{eff}}^{\text{mono}} \rightarrow 140$	140	2.45	$m(\tilde{g}) > 500 \text{ GeV}$	2011.06026	
	$0, \mu, \mu$	6 jets	$A_{\text{eff}}^{\text{mono}} \rightarrow 140$	140	1.25	$m(\tilde{g}) > 300 \text{ GeV}$	1909.06407	
3 ℓ jet states discrimination	$0, \mu, \mu$	2-6 jets	$A_{\text{eff}}^{\text{mono}} \rightarrow 140$	140	0.68, 1.25	$m(\tilde{g}) > 400 \text{ GeV}$	2101.12527	
	$0, \mu, \mu$	6-8 jets	$A_{\text{eff}}^{\text{mono}} \rightarrow 140$	140	0.25-1.25	$10 \text{ GeV} < m(\tilde{g}) < 200 \text{ GeV}$	2101.12527	
	$2, \tau, \tau$	2-6 jets	$A_{\text{eff}}^{\text{mono}} \rightarrow 140$	140	Forbidden	$m(\tilde{g}) > 100 \text{ GeV}$	1908.02122	
	$0, \mu, \mu$	2-6 jets	$A_{\text{eff}}^{\text{mono}} \rightarrow 140$	140	0.13-0.85	$m(\tilde{g}) > 100 \text{ GeV}$	2103.10109	
	$0, 1, \mu, \mu$	≥ 1 jet	$A_{\text{eff}}^{\text{mono}} \rightarrow 140$	140	1.25	$m(\tilde{g}) > 100 \text{ GeV}$	2004.14053, 2010.03799	
	$1, \mu, \mu$	3 jets/1 b	$A_{\text{eff}}^{\text{mono}} \rightarrow 140$	140	Forbidden	$m(\tilde{g}) > 500 \text{ GeV}$	2012.02799, ATLAS-COHE-2023-043	
	$1, 2, \tau, \tau$	2 jets/1 b	$A_{\text{eff}}^{\text{mono}} \rightarrow 140$	140	Forbidden	$m(\tilde{g}) > 100 \text{ GeV}$	2103.07065	
	$0, \mu, \mu$	2 τ	$A_{\text{eff}}^{\text{mono}} \rightarrow 140$	140	0.55, 0.85	$m(\tilde{g}) > 100 \text{ GeV}$	1805.01649	
	$0, \mu, \mu$	mono jet	$A_{\text{eff}}^{\text{mono}} \rightarrow 140$	140	0.55, 0.85	$m(\tilde{g}) > 100 \text{ GeV}$	2102.10074	
	$0, \mu, \mu$	1-4 b	$A_{\text{eff}}^{\text{mono}} \rightarrow 140$	140	0.067-1.18	$m(\tilde{g}) > 500 \text{ GeV}$	2006.05980	
EW direct	$1, 1, 2, \mu, \mu$	Multiple ℓ jets	$A_{\text{eff}}^{\text{mono}} \rightarrow 140$	140	0.205, 0.36	$m(\tilde{g}) > 0, \text{ wino-bino}$	2105.01676, 2108.07595	
	$2, \mu, \mu$	≥ 1 jet	$A_{\text{eff}}^{\text{mono}} \rightarrow 140$	140	0.42	$m(\tilde{g}) > 0, \text{ wino-bino}$	1911.12006	
	$1, 1, 2, \mu, \mu$	Multiple ℓ jets	$A_{\text{eff}}^{\text{mono}} \rightarrow 140$	140	Forbidden	$m(\tilde{g}) > 70 \text{ GeV, wino-bino}$	1908.02115	
	$1, 1, 2, \mu, \mu$	Multiple ℓ jets	$A_{\text{eff}}^{\text{mono}} \rightarrow 140$	140	0.34, 0.40	$m(\tilde{g}) > 70 \text{ GeV, wino-bino}$	2004.10094, 2108.07595	
	$1, 1, 2, \mu, \mu$	Multiple ℓ jets	$A_{\text{eff}}^{\text{mono}} \rightarrow 140$	140	0.34, 0.40	$m(\tilde{g}) > 0, \text{ wino-bino}$	1908.02115	
	$1, 1, 2, \mu, \mu$	Multiple ℓ jets	$A_{\text{eff}}^{\text{mono}} \rightarrow 140$	140	0.26	$m(\tilde{g}) > 0, \text{ wino-bino}$	ATLAS-COHE-2023-029	
	$0, \mu, \mu$	0 jets	$A_{\text{eff}}^{\text{mono}} \rightarrow 140$	140	0.26	$m(\tilde{g}) > 0$	1908.02115	
	$0, \mu, \mu$	≥ 1 jet	$A_{\text{eff}}^{\text{mono}} \rightarrow 140$	140	0.26	$m(\tilde{g}) > 10 \text{ GeV}$	1911.12006	
	$0, \mu, \mu$	≥ 1 b	$A_{\text{eff}}^{\text{mono}} \rightarrow 140$	140	0.35, 0.94	$m(\tilde{g}) > 100 \text{ GeV}$	1911.12006	
	$0, \mu, \mu$	0 jets	$A_{\text{eff}}^{\text{mono}} \rightarrow 140$	140	0.35, 0.94	$m(\tilde{g}) > 100 \text{ GeV}$	2102.11004	
Long-lived particles	Direct \tilde{g}, \tilde{g} prod., long-lived \tilde{g}	Disapp. bk	1 jet	$A_{\text{eff}}^{\text{mono}} \rightarrow 140$	140	0.21, 0.86	Pump Wino Pump Higgsino	2001.02472
	Stable g R hadron	pinel dE/dx	$A_{\text{eff}}^{\text{mono}} \rightarrow 140$	140	2.05	$m(\tilde{g}) > 100 \text{ GeV}$	2005.06013	
	Metastable g R hadron, $g \rightarrow \text{qq}$	pinel dE/dx	$A_{\text{eff}}^{\text{mono}} \rightarrow 140$	140	2.2	$m(\tilde{g}) > 100 \text{ GeV}$	2005.06013	
	$\tilde{g}, \tilde{g} \rightarrow \text{qq}$	Disapp. bk	$A_{\text{eff}}^{\text{mono}} \rightarrow 140$	140	0.34, 0.7	$m(\tilde{g}) > 0.1 \text{ ms}$	2011.07812	
	$\tilde{g}, \tilde{g} \rightarrow \text{qq}$	pinel dE/dx	$A_{\text{eff}}^{\text{mono}} \rightarrow 140$	140	0.34, 0.36	$m(\tilde{g}) > 0.1 \text{ ms}$	2005.06013	
	$\tilde{g}, \tilde{g} \rightarrow \text{qq}$	pinel dE/dx	$A_{\text{eff}}^{\text{mono}} \rightarrow 140$	140	0.34, 0.36	$m(\tilde{g}) > 10 \text{ ms}$	2005.06013	
	$\tilde{g}, \tilde{g} \rightarrow \text{qq}$	pinel dE/dx	$A_{\text{eff}}^{\text{mono}} \rightarrow 140$	140	0.34, 0.36	$m(\tilde{g}) > 10 \text{ ms}$	2005.06013	
	$\tilde{g}, \tilde{g} \rightarrow \text{qq}$	pinel dE/dx	$A_{\text{eff}}^{\text{mono}} \rightarrow 140$	140	0.34, 0.36	$m(\tilde{g}) > 10 \text{ ms}$	2005.06013	
	$\tilde{g}, \tilde{g} \rightarrow \text{qq}$	pinel dE/dx	$A_{\text{eff}}^{\text{mono}} \rightarrow 140$	140	0.34, 0.36	$m(\tilde{g}) > 10 \text{ ms}$	2005.06013	
	$\tilde{g}, \tilde{g} \rightarrow \text{qq}$	pinel dE/dx	$A_{\text{eff}}^{\text{mono}} \rightarrow 140$	140	0.34, 0.36	$m(\tilde{g}) > 10 \text{ ms}$	2005.06013	
RPV	$\tilde{g}, \tilde{g} \rightarrow \text{qq}$	0 jets	$A_{\text{eff}}^{\text{mono}} \rightarrow 140$	140	0.625, 1.05	$m(\tilde{g}) > 200 \text{ GeV}$	2011.10543	
	$\tilde{g}, \tilde{g} \rightarrow \text{qq}$	0 jets	$A_{\text{eff}}^{\text{mono}} \rightarrow 140$	140	0.35, 1.55	$m(\tilde{g}) > 200 \text{ GeV}$	2103.11004	
	$\tilde{g}, \tilde{g} \rightarrow \text{qq}$	0 jets	$A_{\text{eff}}^{\text{mono}} \rightarrow 140$	140	0.35, 1.55	$m(\tilde{g}) > 200 \text{ GeV}$	2103.11004	
	$\tilde{g}, \tilde{g} \rightarrow \text{qq}$	0 jets	$A_{\text{eff}}^{\text{mono}} \rightarrow 140$	140	0.35, 1.55	$m(\tilde{g}) > 200 \text{ GeV}$	2103.11004	
	$\tilde{g}, \tilde{g} \rightarrow \text{qq}$	0 jets	$A_{\text{eff}}^{\text{mono}} \rightarrow 140$	140	0.35, 1.55	$m(\tilde{g}) > 200 \text{ GeV}$	2103.11004	
	$\tilde{g}, \tilde{g} \rightarrow \text{qq}$	0 jets	$A_{\text{eff}}^{\text{mono}} \rightarrow 140$	140	0.35, 1.55	$m(\tilde{g}) > 200 \text{ GeV}$	2103.11004	
	$\tilde{g}, \tilde{g} \rightarrow \text{qq}$	0 jets	$A_{\text{eff}}^{\text{mono}} \rightarrow 140$	140	0.35, 1.55	$m(\tilde{g}) > 200 \text{ GeV}$	2103.11004	
	$\tilde{g}, \tilde{g} \rightarrow \text{qq}$	0 jets	$A_{\text{eff}}^{\text{mono}} \rightarrow 140$	140	0.35, 1.55	$m(\tilde{g}) > 200 \text{ GeV}$	2103.11004	
	$\tilde{g}, \tilde{g} \rightarrow \text{qq}$	0 jets	$A_{\text{eff}}^{\text{mono}} \rightarrow 140$	140	0.35, 1.55	$m(\tilde{g}) > 200 \text{ GeV}$	2103.11004	
	$\tilde{g}, \tilde{g} \rightarrow \text{qq}$	0 jets	$A_{\text{eff}}^{\text{mono}} \rightarrow 140$	140	0.35, 1.55	$m(\tilde{g}) > 200 \text{ GeV}$	2103.11004	
$\tilde{g}, \tilde{g} \rightarrow \text{qq}$	0 jets	$A_{\text{eff}}^{\text{mono}} \rightarrow 140$	140	0.35, 1.55	$m(\tilde{g}) > 200 \text{ GeV}$	2103.11004		

*Only a selection of the available mass limits on new states or phenomena is shown. Many of the limits are based on simplified models, c.f. refs. for the assumptions made.

[PDG,2023]

low energy SUSY is not observed so far!

Non-SUSY



???



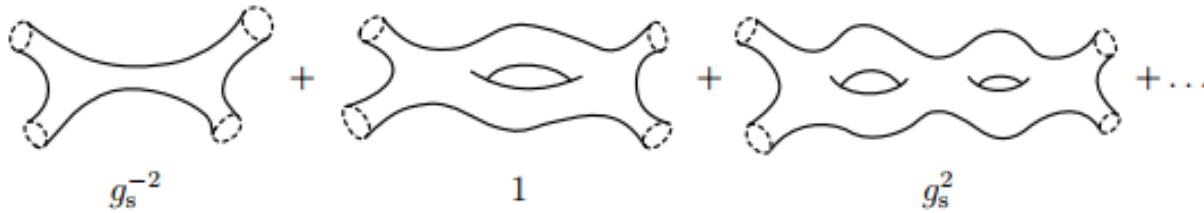
$Y(\tau, \bar{\tau})$: non-holomorphic Yukawa couplings



A finite choice of $Y(\tau, \bar{\tau})$ to have prediction power

Harmonic (Laplacian) condition replacing holomorphicity

- From top-down, the effective interactions in the low energy expansion of the four-graviton amplitude are **non-holomorphic automorphic** functions satisfying Laplace eigenvalue equations.



[Green, Gutperle, hep-th/9701093; Green, Russo, Vanhove, 1001.2535; Hoker, Kaidi, 2208.07242]

- The bottom-up approach based on automorphic forms [Ding, Feruglio, Liu, 2010.07952]
- Automorphic forms coincide with the harmonic Maass forms for single modulus τ

- Harmonic Maass forms of level N and weight k

$$\left\{ \begin{array}{l} \textcircled{1} \text{ modularity: } Y(\gamma\tau) = (c\tau + d)^k Y(\tau), \quad \gamma \in \Gamma(N) \\ \textcircled{2} \text{ harmonic condition: } \Delta_k Y(\tau) = 0, \quad \Delta_k \equiv -4y^2 \frac{\partial}{\partial\tau} \frac{\partial}{\partial\bar{\tau}} + 2iky \frac{\partial}{\partial\bar{\tau}} \\ \textcircled{3} \text{ growth condition: } Y(\tau) = \mathcal{O}(y^\alpha), \quad y \rightarrow +\infty \end{array} \right.$$

- Fourier expansion:

$$Y(\tau) = \underbrace{\sum_{n \in \frac{1}{N}\mathbb{Z}^+} c_n^+ q^n}_{\text{holomorphic}} + \underbrace{c_0^- y^{1-k} + \sum_{n \in \frac{1}{N}\mathbb{Z}^-} c_n^- \Gamma(1-k, -4\pi ny) q^n}_{\text{non-holomorphic}}$$

[Qu, Ding, in progress]

- Transformation under modular group: $Y_i(\gamma\tau) = (c\tau + d)^k \rho_{ij}(\gamma) Y_j(\tau), \quad \gamma \in SL(2, \mathbb{Z})$

$\rho(\gamma)$ is unitary representation of $\Gamma'_N = \Gamma/\Gamma(N)$ for odd k and $\Gamma_N = \Gamma/\pm\Gamma(N)$ for even k

Harmonic Maass forms of level 3

- Weight k harmonic Maass forms $Y_r^{(k)}(\tau)$ related with modular forms $Y_r^{(2-k)}(\tau)$ of weight $2 - k$

$$2iy^k \frac{\partial}{\partial \tau} \overline{Y_r^{(k)}(\tau)} \propto Y_r^{(2-k)}(\tau), \quad \left(\frac{1}{2\pi i} \frac{\partial}{\partial \tau} \right)^{1-k} Y_r^{(k)}(\tau) \propto Y_r^{(2-k)}(\tau)$$

- weight 2 modular forms transforming as an A_4 triplet 3 [Feruglio, 1706.08749; Liu, Ding, 1907.01488]

$$Y_3^{(2)}(\tau) \equiv \begin{pmatrix} \varepsilon^2(\tau) \\ \sqrt{2} \vartheta(\tau) \varepsilon(\tau) \\ -\vartheta^2(\tau) \end{pmatrix} = \begin{pmatrix} 1 + 12q + 36q^2 + \dots \\ -6q^{1/3} (1 + 7q + 8q^2 + \dots) \\ -18q^{2/3} (1 + 2q + 5q^2 + \dots) \end{pmatrix}$$

$$\vartheta(\tau) = 3\sqrt{2} \frac{\eta^3(3\tau)}{\eta(\tau)},$$

$$\varepsilon(\tau) = -\frac{3\eta^3(3\tau) + \eta^3(\tau/3)}{\eta(\tau)}$$

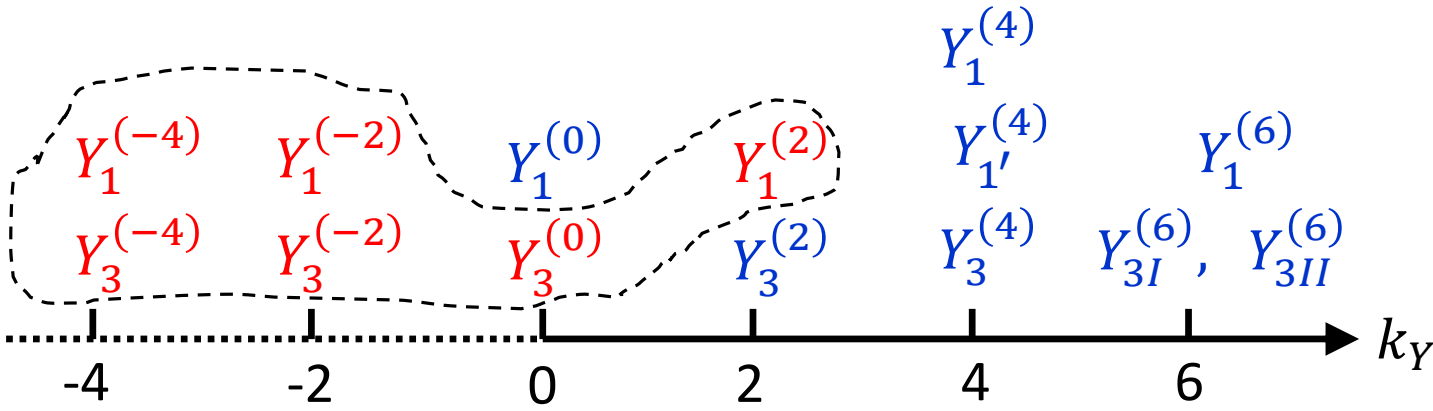
Lifted to weight zero harmonic Maass forms



$$Y_3^{(0)}(\tau) = \begin{pmatrix} -\left(0.7868 + \frac{3q}{\pi} + \frac{9q^2}{2\pi} + \dots\right) + \left(y - \frac{3e^{-4\pi y}}{\pi q} - \frac{9e^{-8\pi y}}{2\pi q^2} + \dots\right) \\ \frac{9q^{1/3}}{2\pi} \left(1 + \frac{7q}{4} + \frac{8q^2}{7} + \dots\right) + \frac{27q^{1/3} e^{\pi y/3}}{\pi} \left(\frac{e^{-3\pi y}}{4q} + \frac{e^{-7\pi y}}{5q^2} + \dots\right) \\ \frac{27q^{2/3}}{\pi} \left(\frac{1}{4} + \frac{q}{5} + \frac{5q^2}{16} + \dots\right) + \frac{9q^{2/3} e^{2\pi y/3}}{2\pi} \left(\frac{e^{-2\pi y}}{q} + \frac{7e^{-6\pi y}}{4q^2} + \dots\right) \end{pmatrix}$$

[Qu, Ding, in progress]

➤ Harmonic Maass forms coincide with modular forms for weight $k > 3$, negative weight Maass forms and phenomenological implications in mass hierarchies and strong CP problem?



➤ **General comments and questions**

- Origin of harmonic condition and the possible symmetry related?
- Possible top-down connection?
- Are harmonic Maass forms applicable to flavor puzzle?
-

An example model based on A_4 modular symmetry

➤ field content and assignment

[Qu, Ding, in progress]

	L	e^c	μ^c	τ^c	H
$SU(2)_L \times U(1)_Y$	$(\mathbf{2}, -1/2)$		$(\mathbf{1}, 1)$		$(\mathbf{2}, 1/2)$
A_4	$\mathbf{3}$	$\mathbf{1}$	$\mathbf{1}''$	$\mathbf{1}'$	$\mathbf{1}$
k_I	-2	0	2	2	0

➤ Neutrino masses arise from Weinberg operator (two-component formalism for fermions)

$$\begin{aligned}
 -\mathcal{L}_e &= \alpha(H^\dagger e^c L Y_{\mathbf{3}}^{(-2)})_{\mathbf{1}} + \beta(H^\dagger \mu^c L Y_{\mathbf{3}}^{(0)})_{\mathbf{1}} + \gamma(H^\dagger \tau^c L Y_{\mathbf{3}}^{(0)})_{\mathbf{1}} + \text{h.c.}, \\
 -\mathcal{L}_\nu &= \frac{1}{2\Lambda}(LLH^2 Y_{\mathbf{3}}^{(-4)})_{\mathbf{1}} + \frac{g}{2\Lambda}(LLH^2 Y_{\mathbf{1}}^{(-4)})_{\mathbf{1}} + \text{h.c.},
 \end{aligned}$$

Lepton mass matrices

$$M_e = \begin{pmatrix} \alpha Y_{\mathbf{3},1}^{(-2)} & \alpha Y_{\mathbf{3},3}^{(-2)} & \alpha Y_{\mathbf{3},2}^{(-2)} \\ \beta Y_{\mathbf{3},2}^{(0)} & \beta Y_{\mathbf{3},1}^{(0)} & \beta Y_{\mathbf{3},3}^{(0)} \\ \gamma Y_{\mathbf{3},3}^{(0)} & \gamma Y_{\mathbf{3},2}^{(0)} & \gamma Y_{\mathbf{3},1}^{(0)} \end{pmatrix} v, \quad M_\nu = \begin{pmatrix} 2Y_{\mathbf{3},1}^{(-4)} + gY_{\mathbf{1}}^{(-4)} & -Y_{\mathbf{3},3}^{(-4)} & -Y_{\mathbf{3},2}^{(-4)} \\ -Y_{\mathbf{3},3}^{(-4)} & 2Y_{\mathbf{3},2}^{(-4)} & -Y_{\mathbf{3},1}^{(-4)} + gY_{\mathbf{1}}^{(-4)} \\ -Y_{\mathbf{3},2}^{(-4)} & -Y_{\mathbf{3},1}^{(-4)} + gY_{\mathbf{1}}^{(-4)} & 2Y_{\mathbf{3},3}^{(-4)} \end{pmatrix} \frac{v^2}{\Lambda}$$

Adjusting α, β, γ for electron, muon and tau masses

Neutrino masses and mixing described by three parameters: $g, \frac{v^2}{\Lambda}, \tau$

➤ Best agreement with data is achieved at

$$\beta/\alpha = 125.743, \quad \gamma/\alpha = 1574.04, \quad \alpha v = 3.62033 \text{ MeV},$$

$$g = 0.753046 + 0.385855i, \quad \langle \tau \rangle = 0.369080 + 0.961090i, \quad v^2/\Lambda = 97.3898 \text{ meV}$$

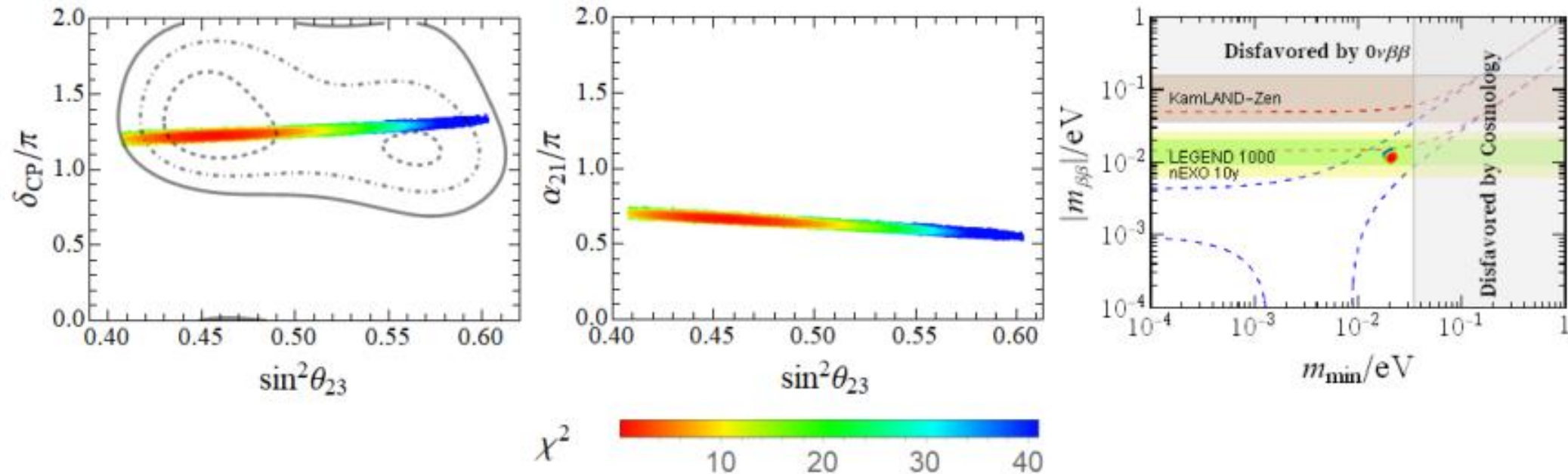
Predictions for lepton masses and mixing parameters

$$m_e/m_\mu = 0.004737, \quad m_\mu/m_\tau = 0.05868, \quad m_e = 0.46965 \text{ MeV},$$

$$m_1 = 21.0918 \text{ meV}, \quad m_2 = 22.7808 \text{ meV}, \quad m_3 = 54.3128 \text{ meV},$$

$$\sin^2 \theta_{12} = 0.3070, \quad \sin^2 \theta_{13} = 0.02224, \quad \sin^2 \theta_{23} = 0.4541,$$

$$\delta_{CP} = 1.2252\pi, \quad \alpha_{21} = 0.6667\pi, \quad \alpha_{31} = 1.3855\pi, \quad \chi^2_{\min} \approx 0$$



Summary

Modular flavor symmetry is a very interesting approach to address flavor puzzle of SM, closely related to top-down. It was proposed for flavor, but not only for flavor. There are still many aspects needing clarification and development toward theory of flavor.

**LET'S
DO
IT
TOGETHER!**

