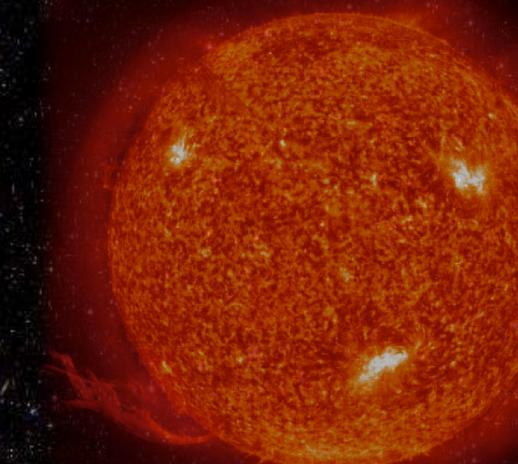


LUNA400 and LUNA-MV: present and future of Nuclear Astrophysics at Gran Sasso

C. Gustavino
INFN-Roma

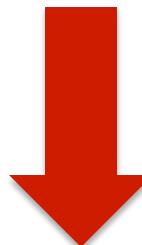


LUNA400 (2015-2018):

| | |
|--------------------------|-------------------|
| $D(p,\gamma)^3He$ | BBN |
| $^{13}C(\alpha,n)^{16}O$ | S-process |
| $^{12}C(p,\gamma)^{13}N$ | Stellar evolution |
| $^{13}C(p,\gamma)^{14}N$ | Stellar evolution |
| $^6Li(p,\gamma)^7Be$ | Nuclear Physics |

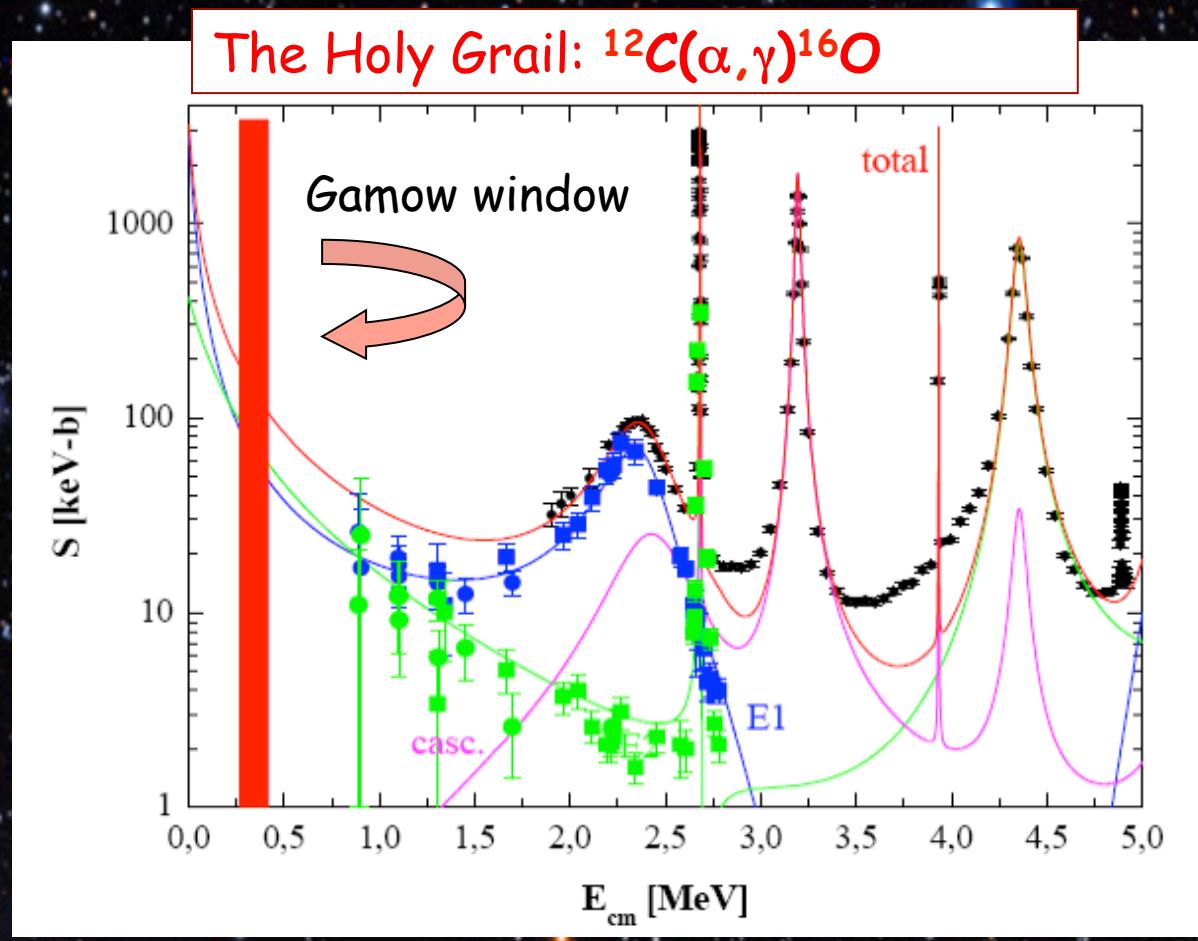
LUNA-MV:

| | |
|-------------------------------|------------------|
| $^3He(^4He,\gamma)^7Be$ | pp chain and BBN |
| $^{12}C(\alpha,\gamma)^{16}O$ | He Burning |
| $^{13}C(\alpha,n)^{16}O$ | S-process |
| $^{22}Ne(\alpha,n)^{25}Mg$ | S-process |
| $^{12}C+^{12}C$ reactions | Carbon burning |
| $D(\alpha,\gamma)^6Li$ | BBN |

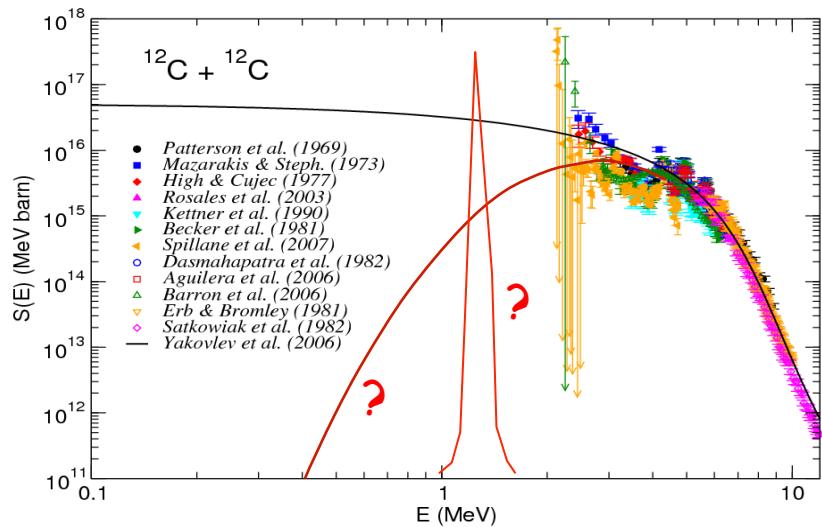
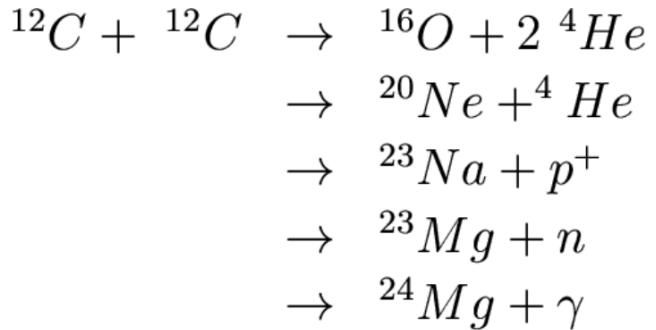
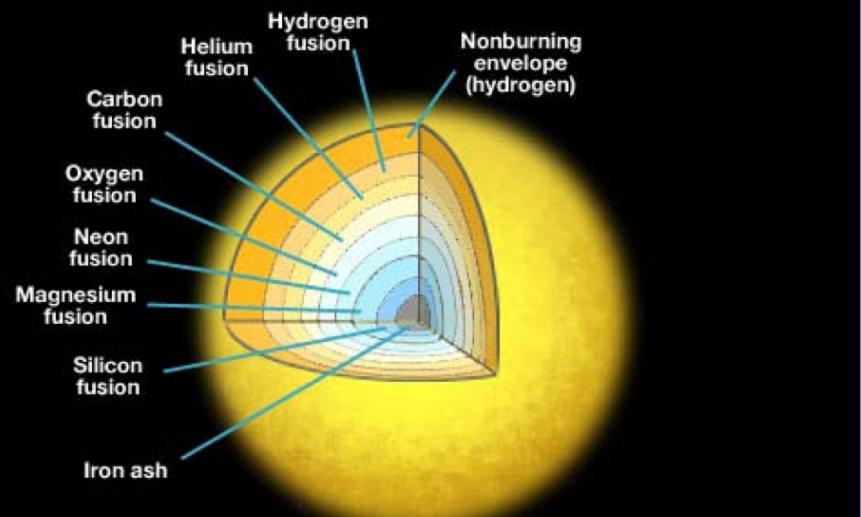


- Solar Neutrino mixing parameters
- Temperature and metallicity of the Sun
- Age of Universe
- Evolution of stars
- Isotopic abundances
- Cosmology and Particle physics

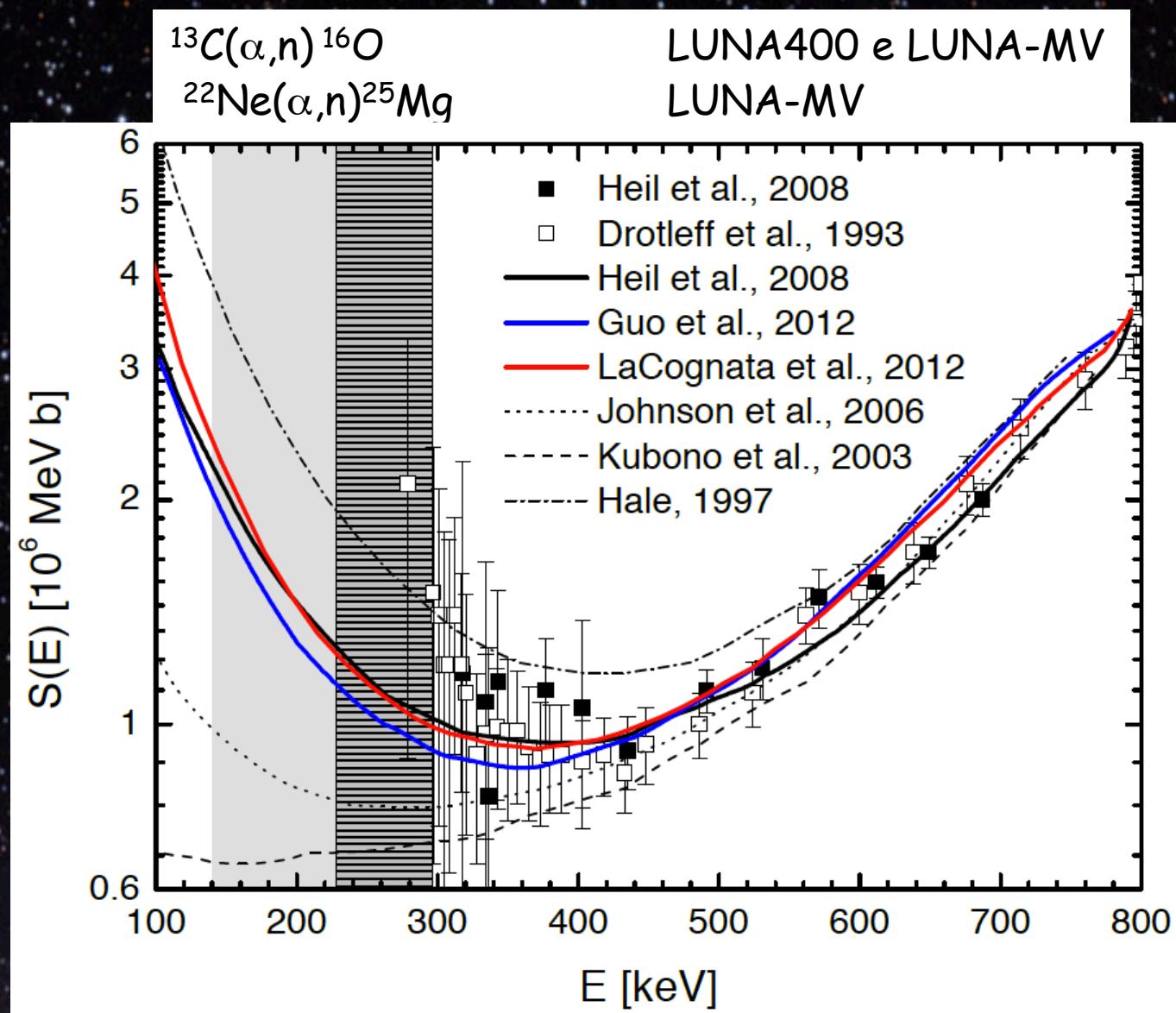
Evolution of stars: Helium Burning



Carbon burning & type Ia supernovae



S-process



Big Bang Nucleosynthesis: The D(p,γ) ^3He reactions

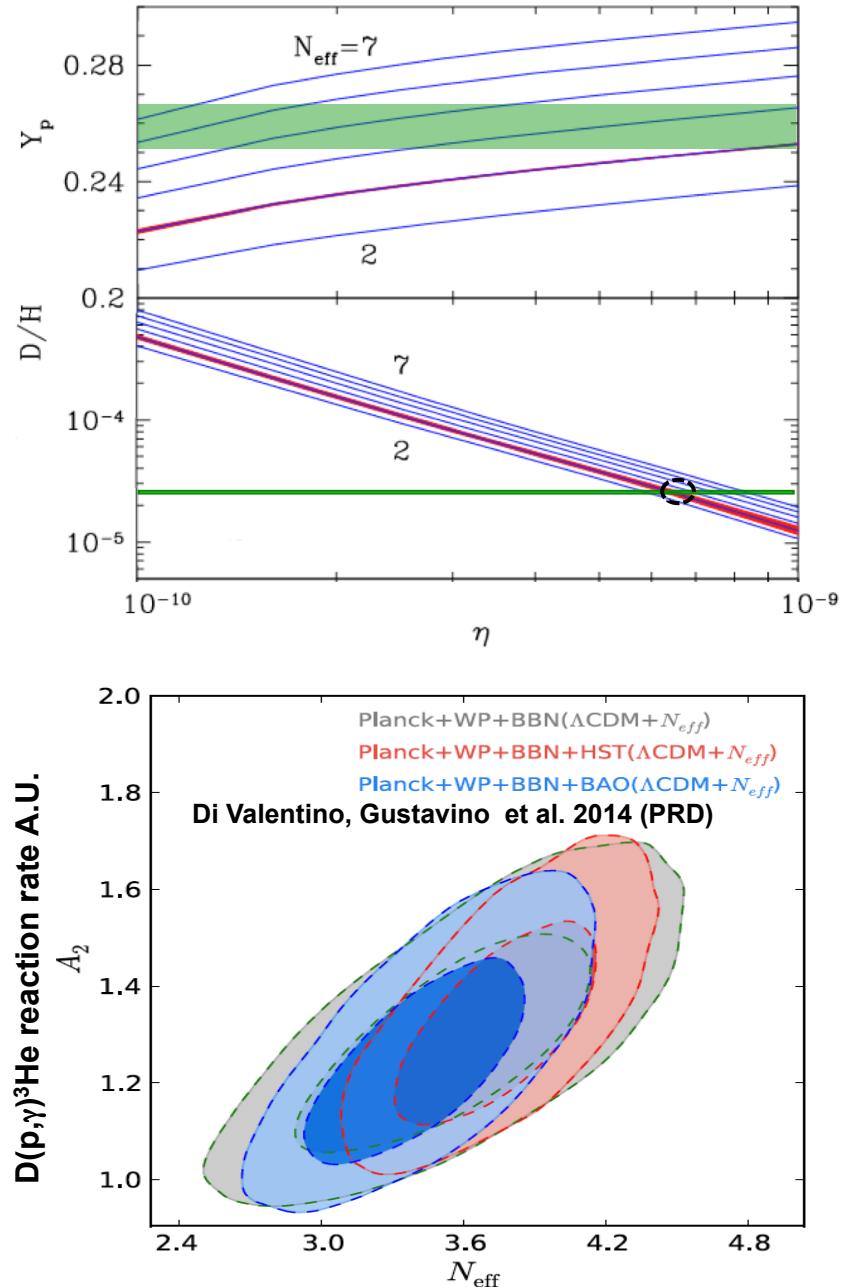
The error of computed deuterium abundance (D/H)_{BBN} is mainly due to the D(p,γ) ^3He reaction, because of the paucity of data in the BBN energy region

| Reaction | Rate Symbol | $\sigma_2 \text{H}/\text{H} \cdot 10^5$ |
|--------------------------|-------------|---|
| $p(n,\gamma)^2\text{H}$ | R_1 | ± 0.002 |
| $d(p,\gamma)^3\text{He}$ | R_2 | ± 0.062 |
| $d(d,n)^3\text{He}$ | R_3 | ± 0.020 |
| $d(d,p)^3\text{H}$ | R_4 | ± 0.013 |

Physics:

- 1) Cosmology: measurement of $\Omega_b h^2$.
- 2) Neutrino physics: measurement of N_{eff} .
- 3) Nuclear physics: comparison of data with theoretical "ab initio" predictions.

See also D. Trezzi talk



Other reactions?

Future LUNA program is in progress.
Suggestions are Welcome!

Example of a chain

