

# Status of the DMRadio Program

Nicholas Rapidis

Stanford University – Irwin Group

DMRadio Collaboration

Shoot for the Stars, Aim for the Axions – October 6, 2022

# DMRadio Collaboration

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*California State University, East Bay*

B. R. Safdi  
*Department of Physics*  
*University of California Berkeley*



Berkeley  
UNIVERSITY OF CALIFORNIA



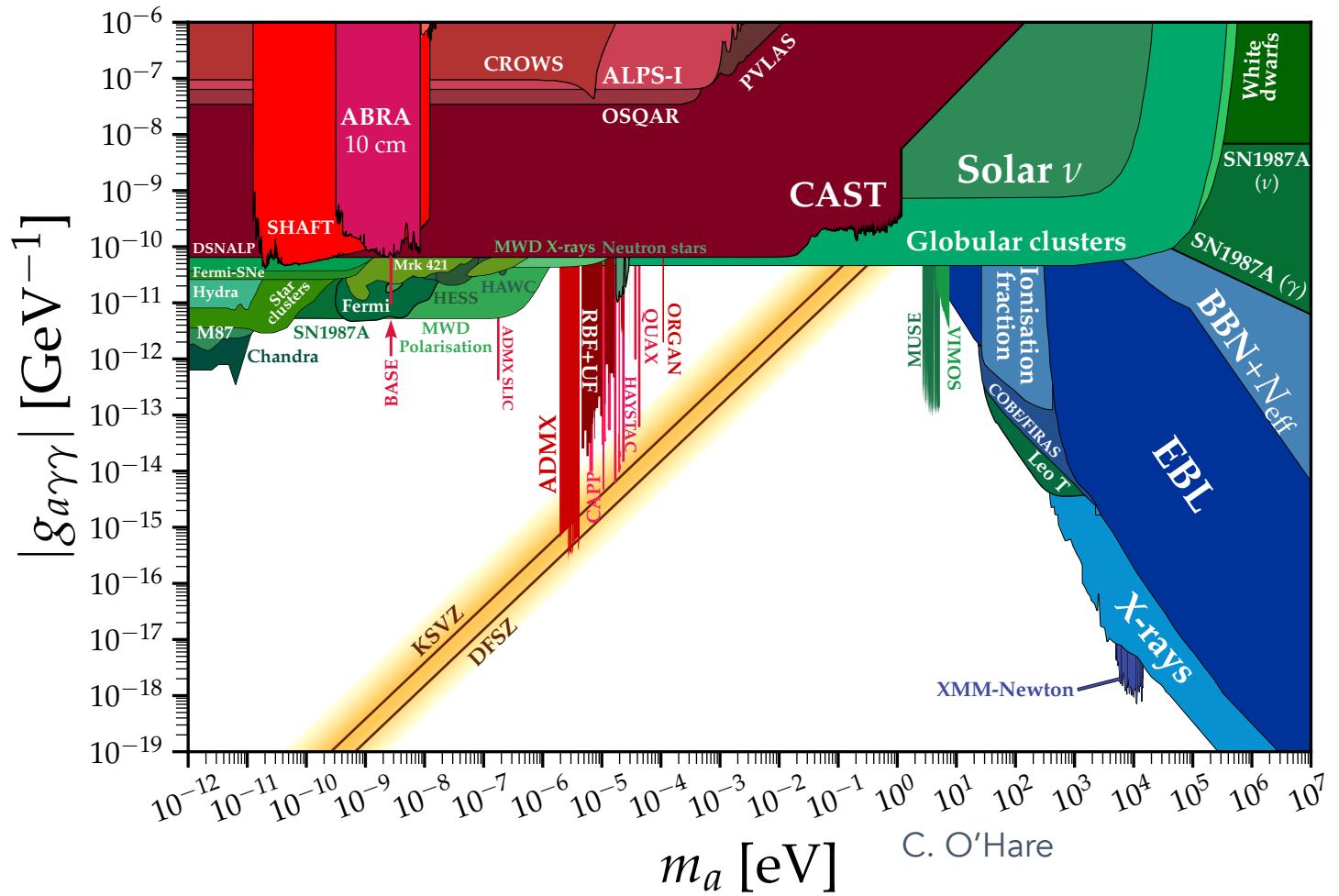
# Outline

1. Axions and low frequency haloscopes
2. DMRadio 50L
  1. Overview
  2. Status
3. DMRadio m<sup>3</sup>
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4. Outlook, future, & collaboration

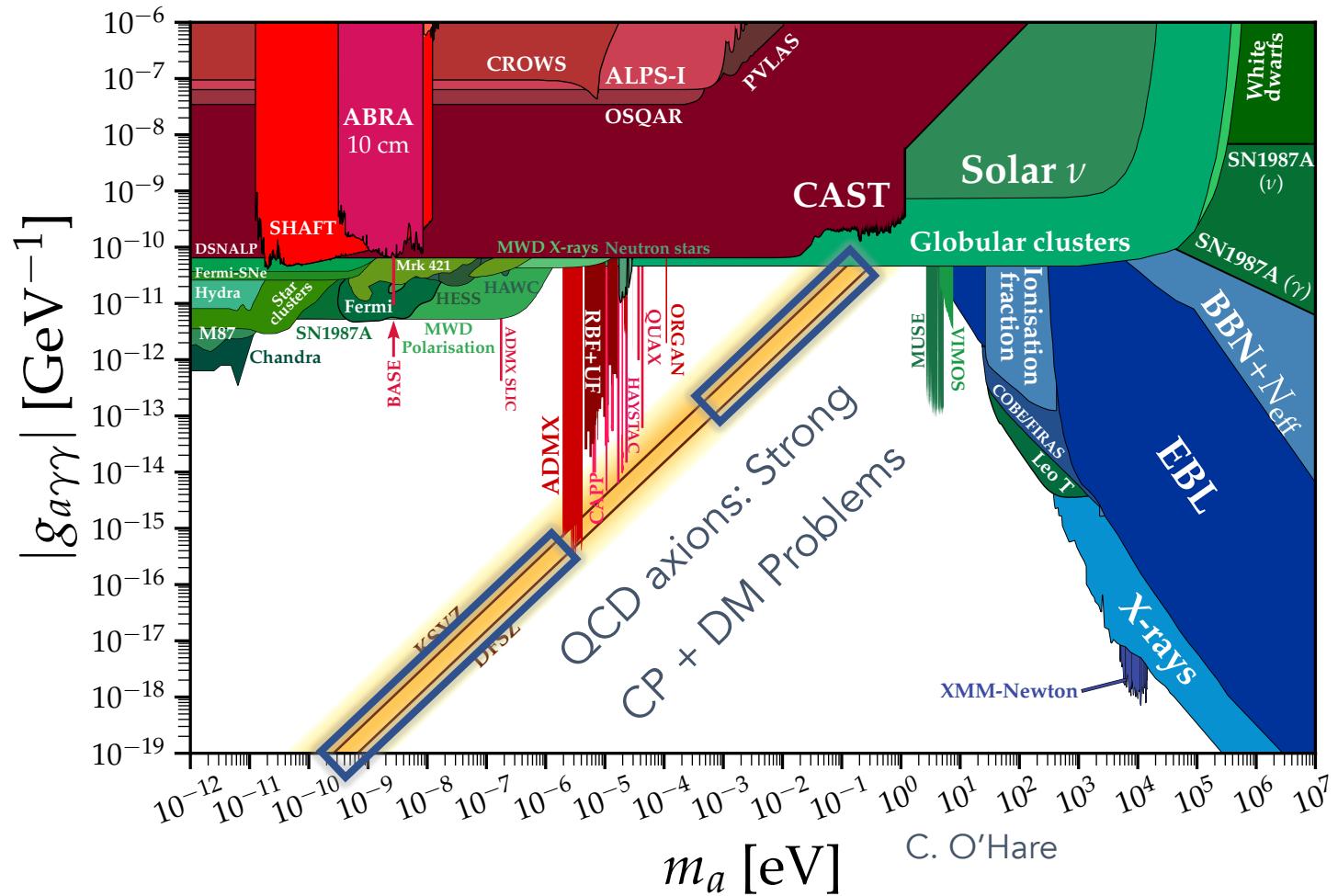
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# Axions



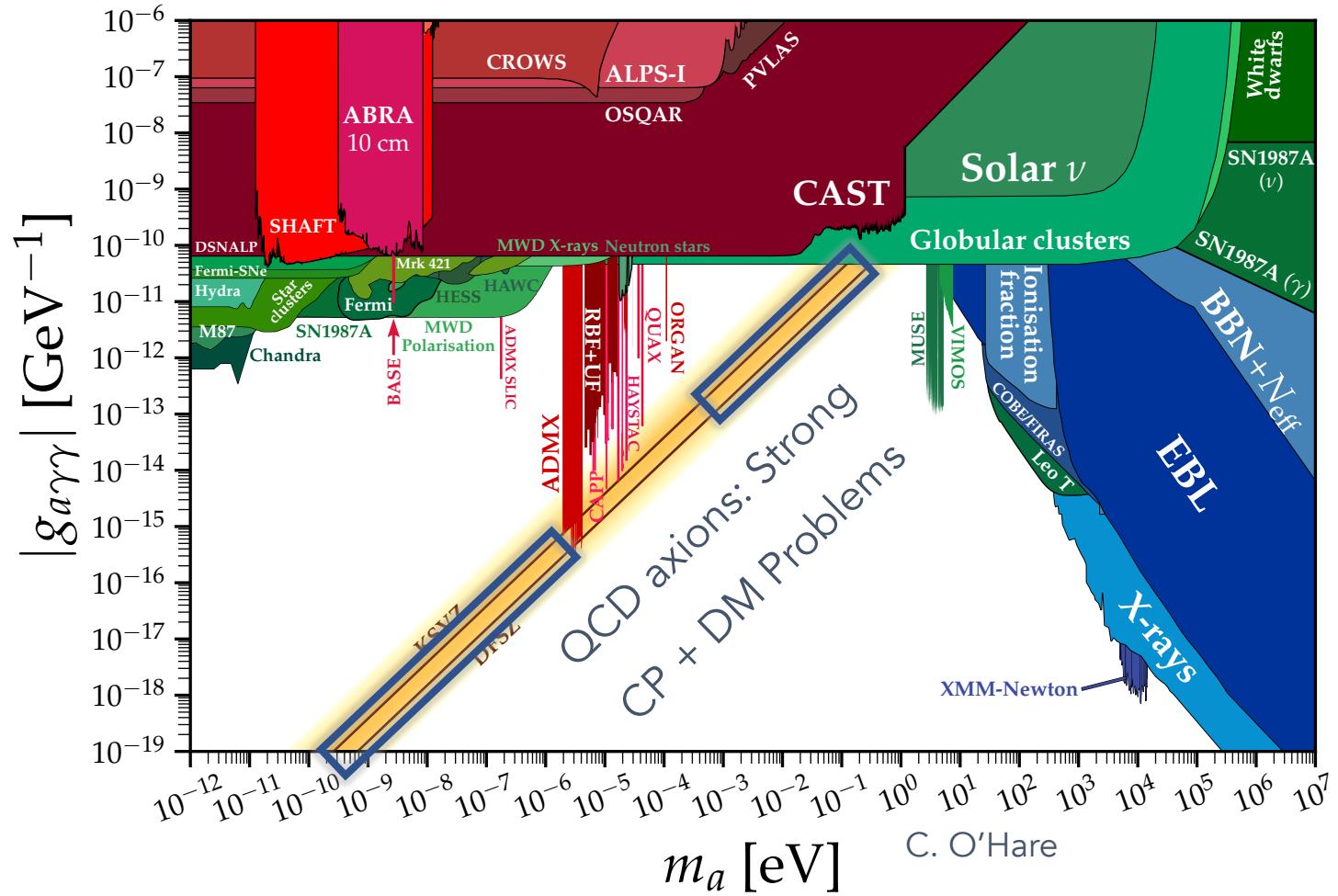
# Axions



# Axions

Axion to photon conversion:

$$\nu_a = \frac{m_a c^2}{h}$$



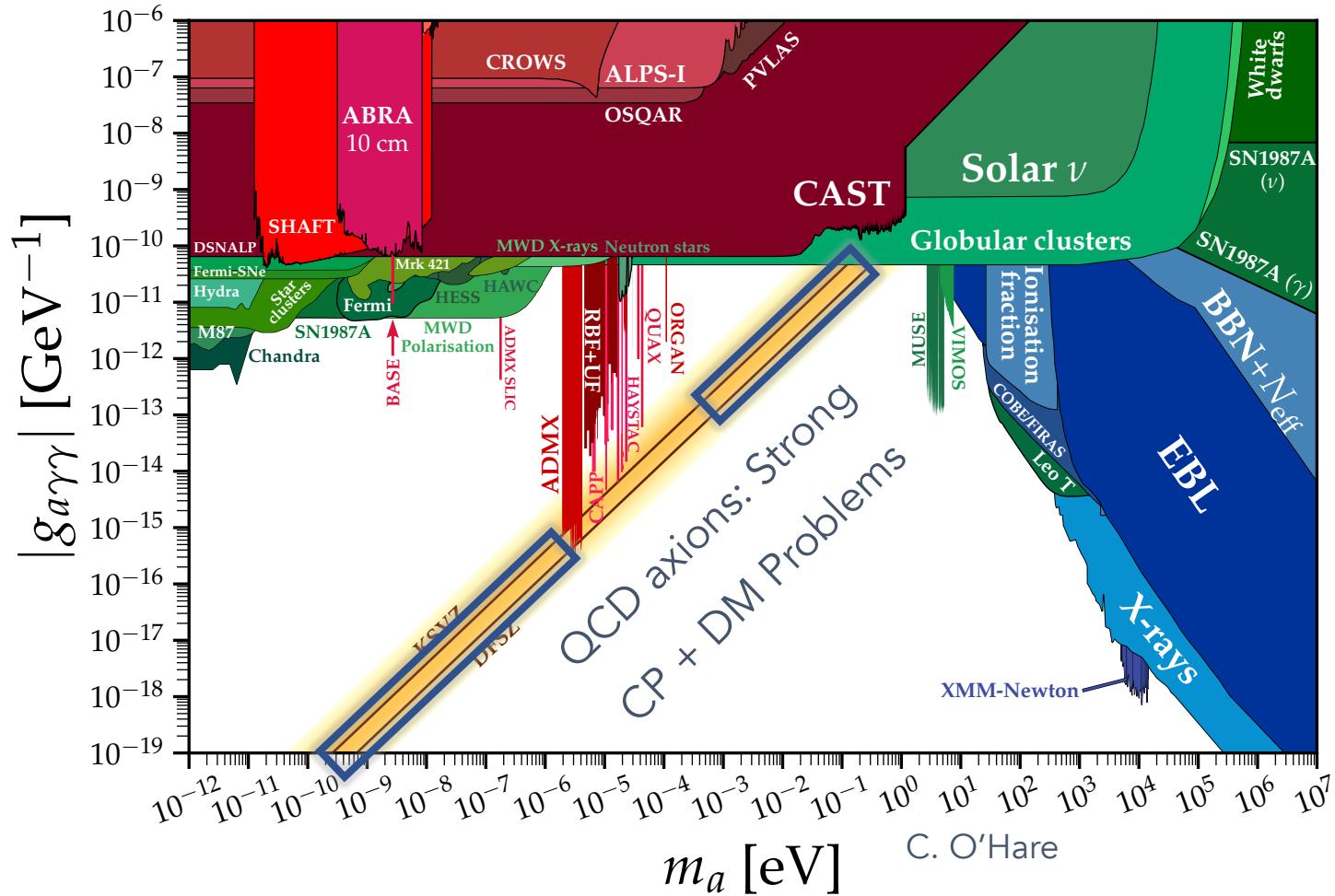
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QCD axion symmetry  
breaking scale:

$$f_a \sim g_a \gamma\gamma^{-1}$$



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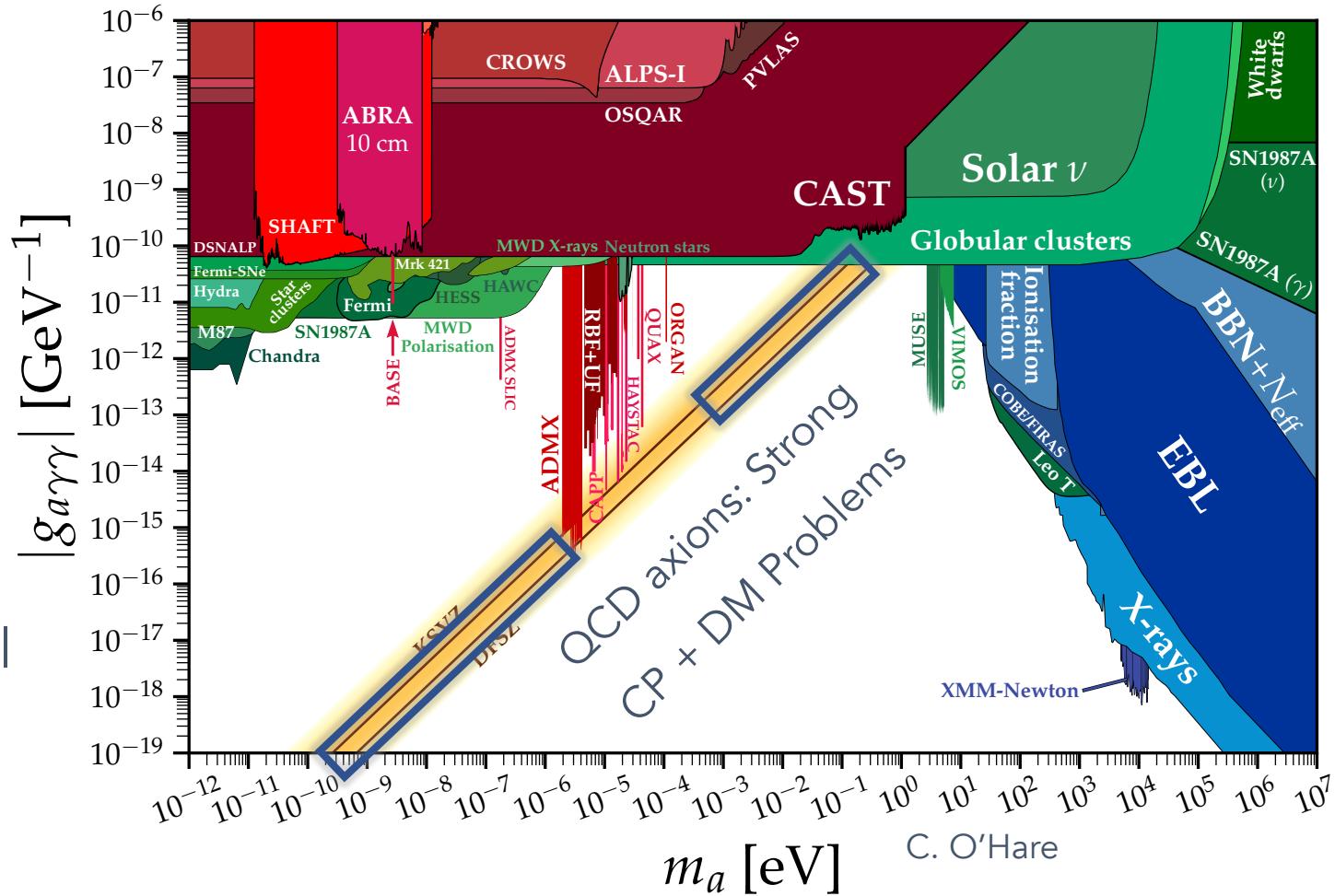
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Axion parameter space is well  
motivated for:

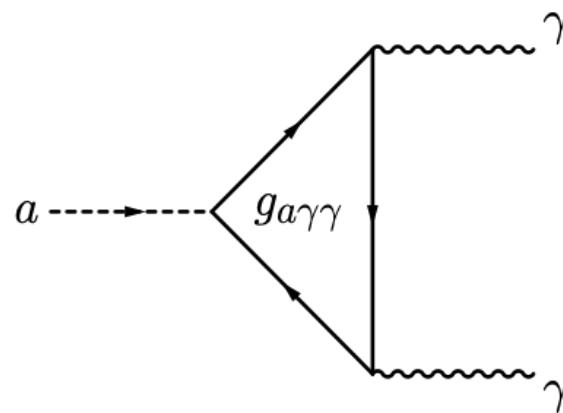
$$\Lambda_{\text{inflation}} < f_a \lesssim m_{\text{Planck}}$$



# Axions and Axion DM

Axion-photon coupling:

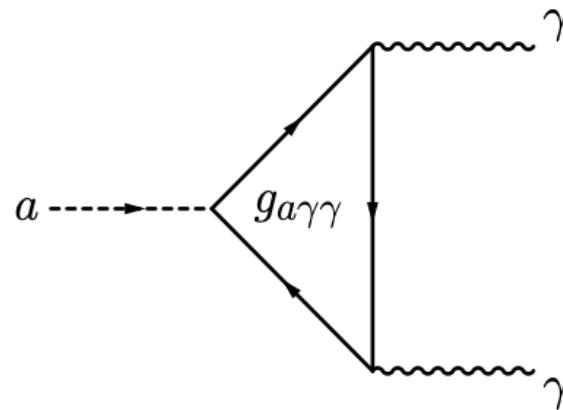
$$\mathcal{L} \supset g_{a\gamma\gamma} a F \tilde{F} \sim g_{a\gamma\gamma} a \vec{E} \cdot \vec{B}$$



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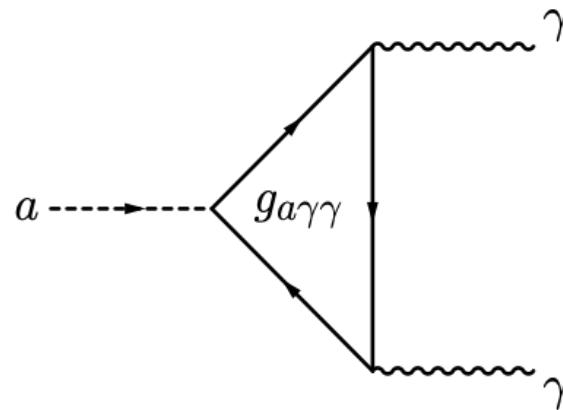
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Axion + magnetic field photon  
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Expected axion number density (for neV axion)

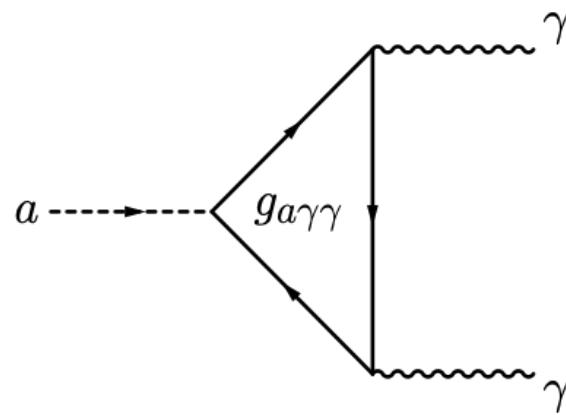
$$n_a \sim 10^{17} \text{ cm}^{-3} \gg 1 \text{ per quantum state}$$

Axion can be modeled as classical wave

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$$\nabla \cdot \vec{E} = -g_{a\gamma\gamma} \vec{B} \cdot \nabla a$$

$$\nabla \cdot \vec{B} = 0$$

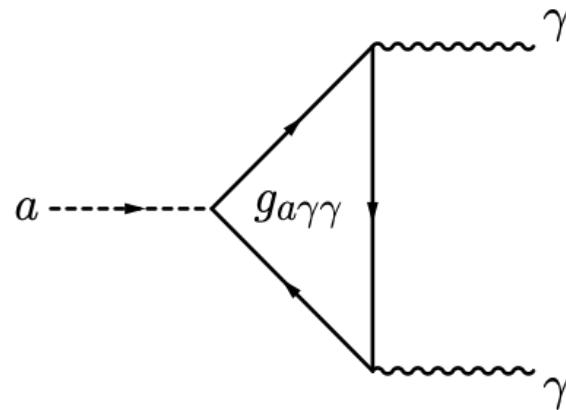
$$\nabla \times \vec{E} = -\partial_t \vec{B}$$

$$\nabla \times \vec{B} = \partial_t \vec{E} - g_{a\gamma\gamma} (\vec{E} \times \nabla a - \partial_t a \vec{B})$$

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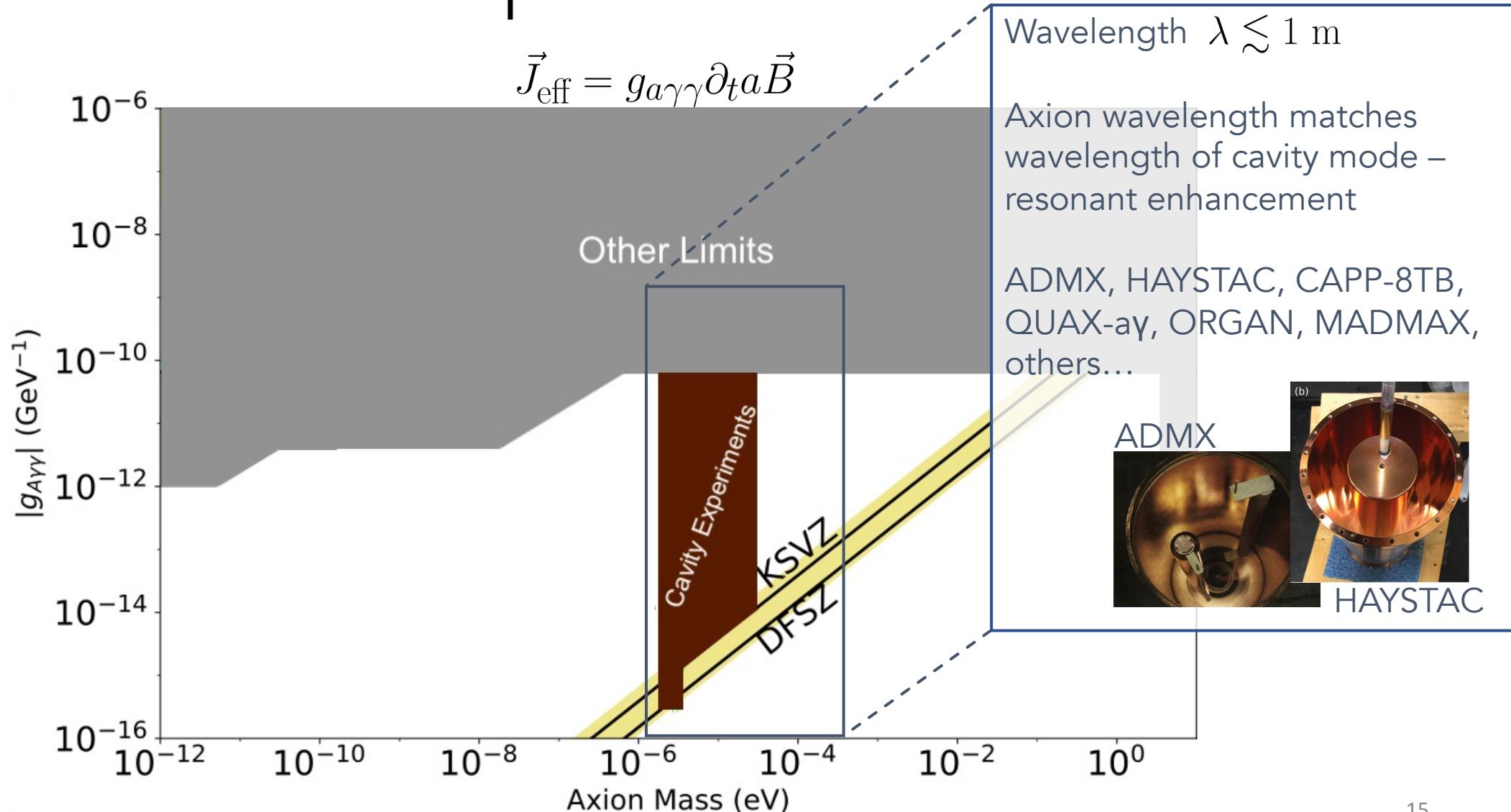
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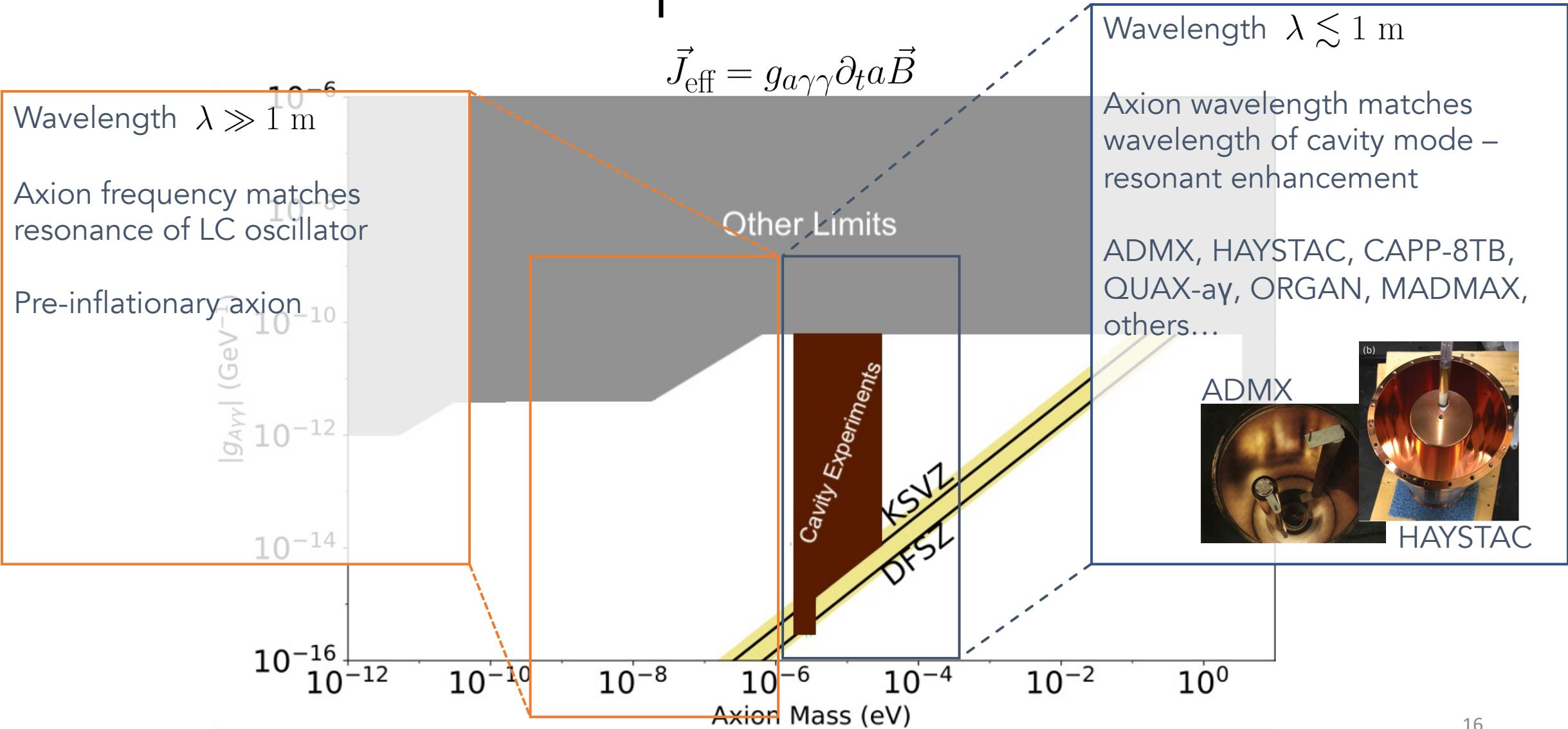
$$\nabla \times \vec{B} = \partial_t \vec{E} - g_{a\gamma\gamma} (\vec{E} \times \nabla a - \boxed{\partial_t a \vec{B}})$$

$$\vec{J}_{\text{eff}} = g_{a\gamma\gamma} \partial_t a \vec{B}$$

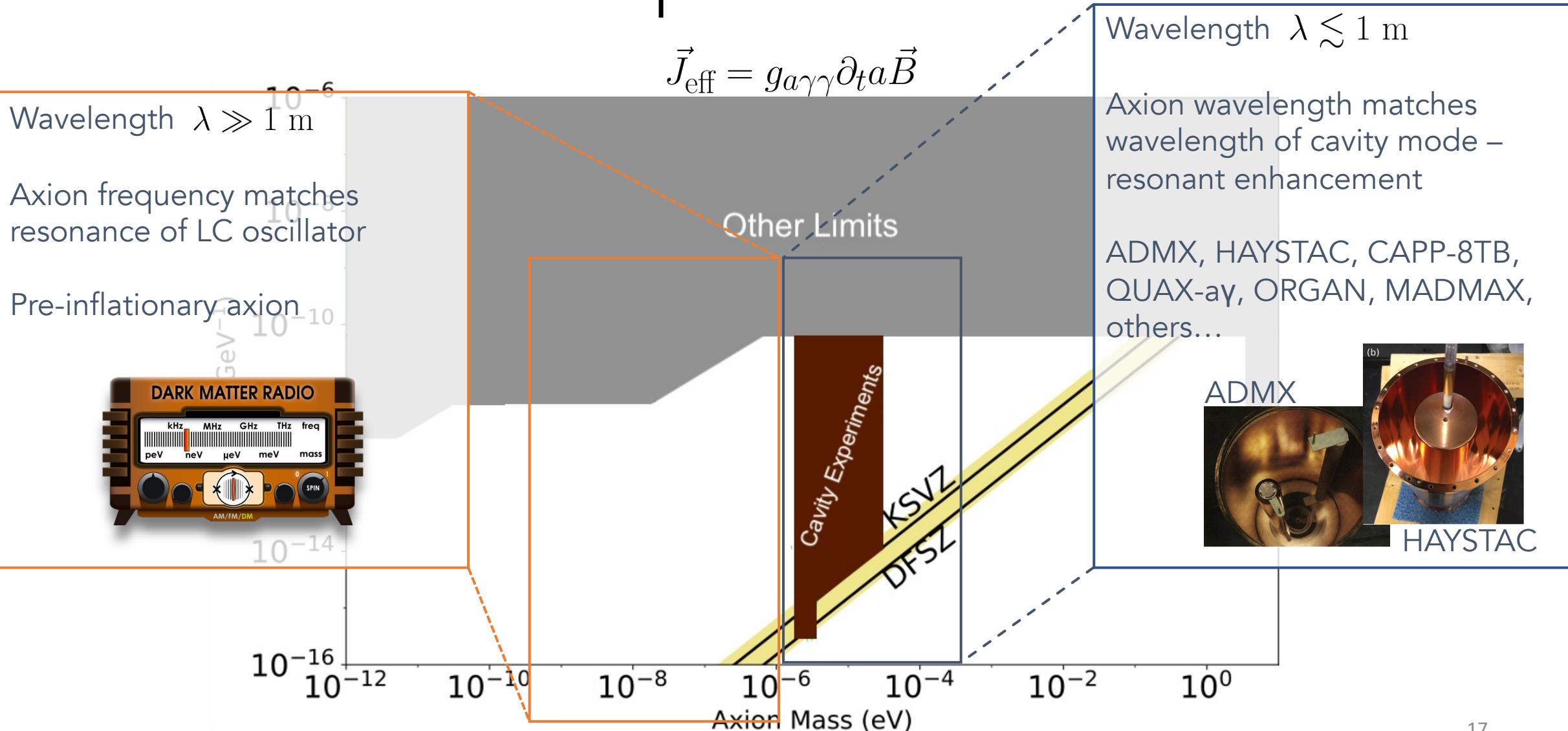
# Axion Parameter Space



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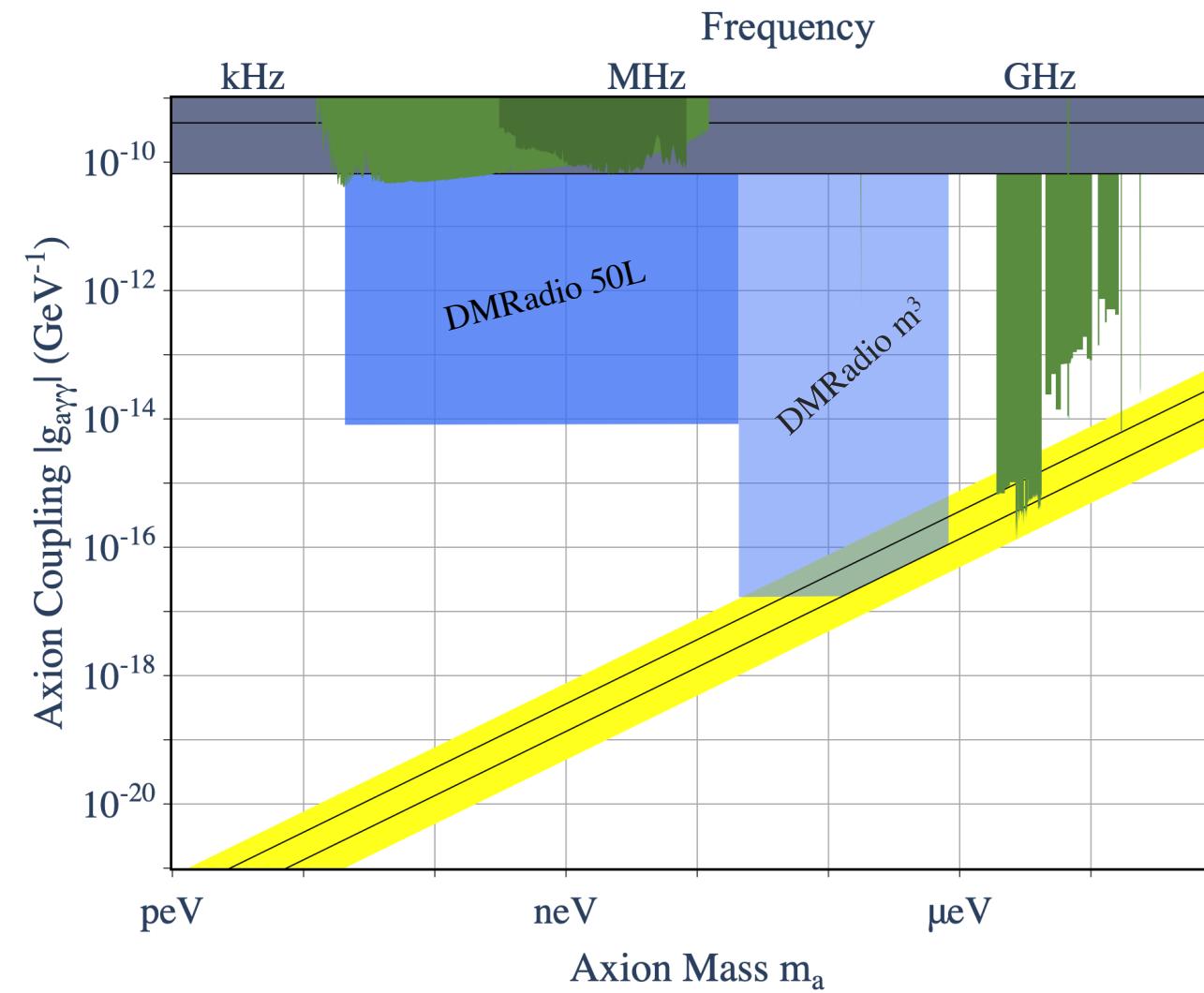


# Axion Parameter Space

Wavelength  $\lambda \gg 1$  m

Axion frequency matches resonance of LC oscillator

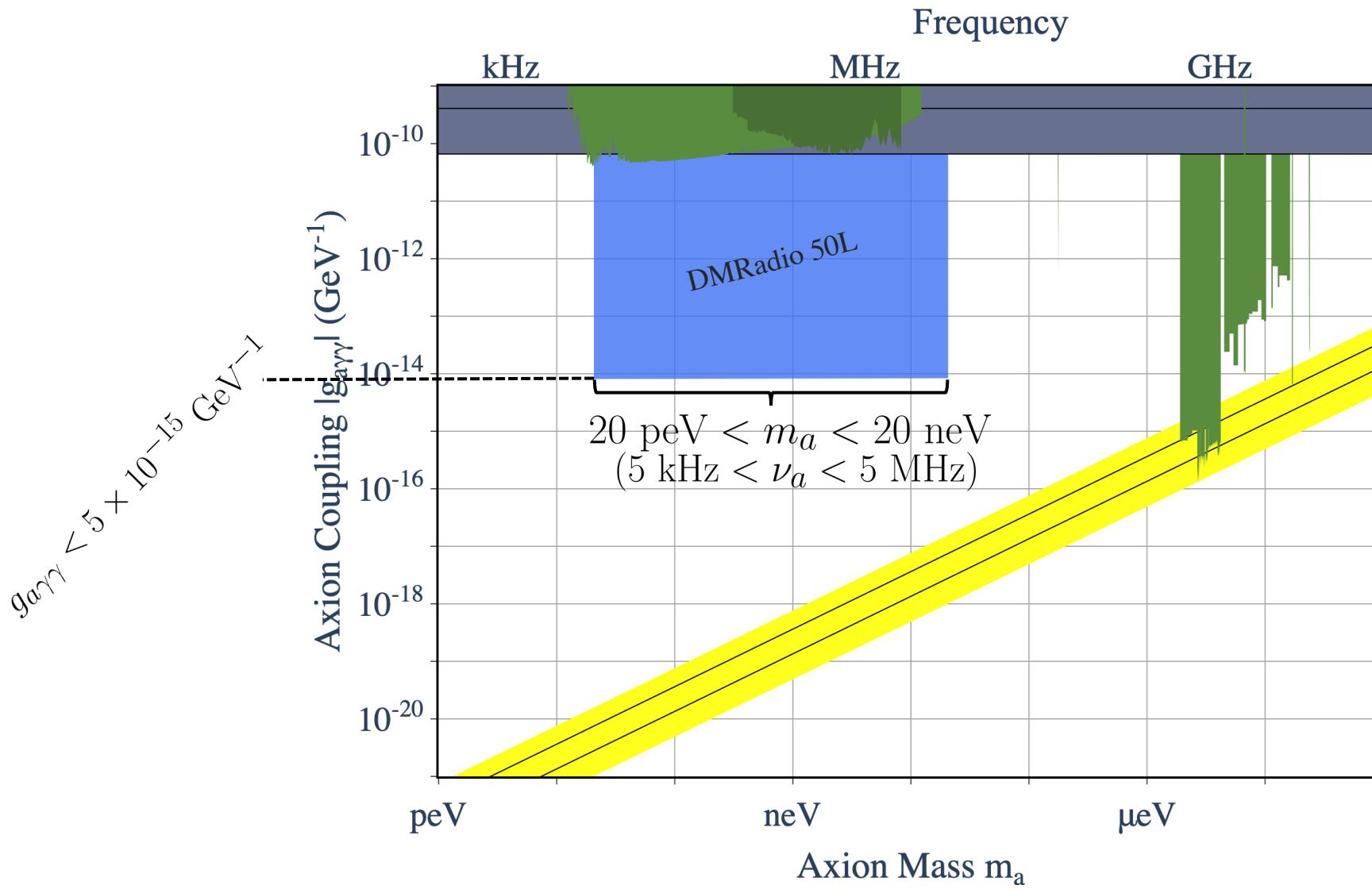
Pre-inflationary axion



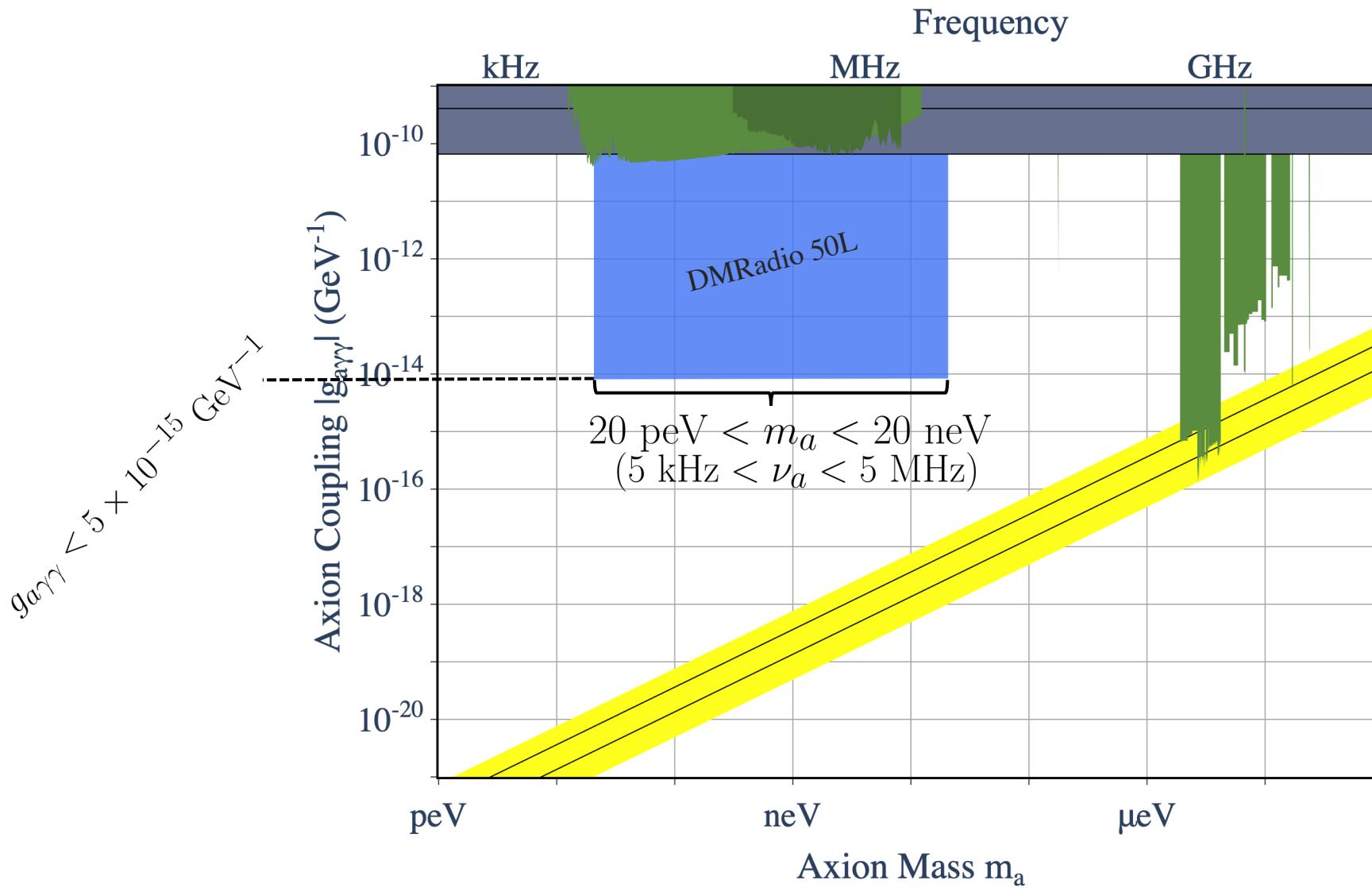
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# 50L Goals

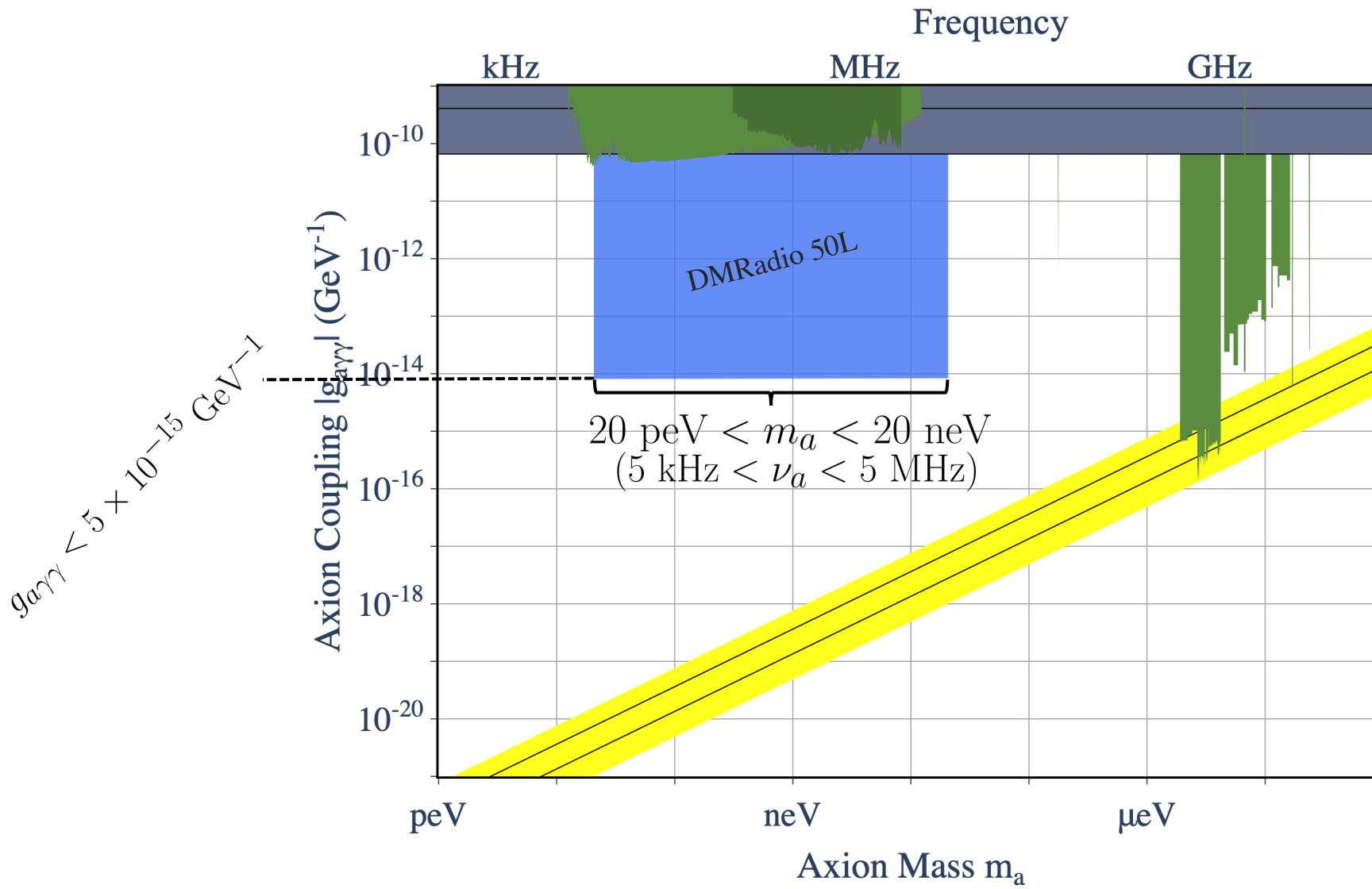


# 50L Goals



Demonstration of LC  
resonator + magnet

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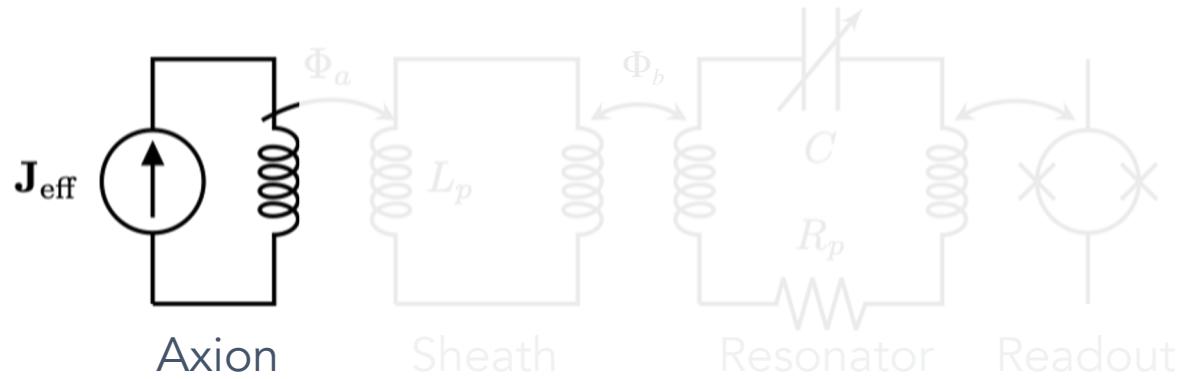


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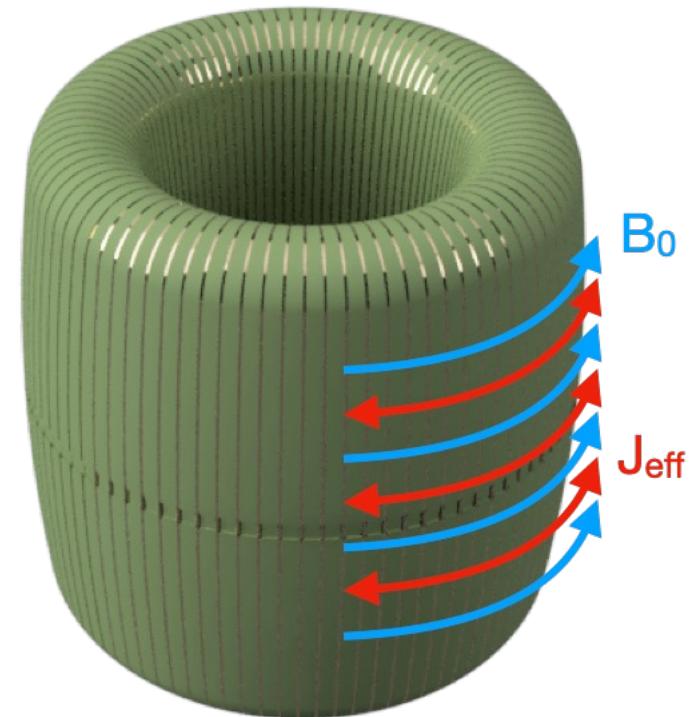
Testbed for novel  
quantum devices

# 50L Circuit & Design

$$\vec{J}_{\text{eff}} = g_a \gamma \gamma \partial_t a \vec{B}$$

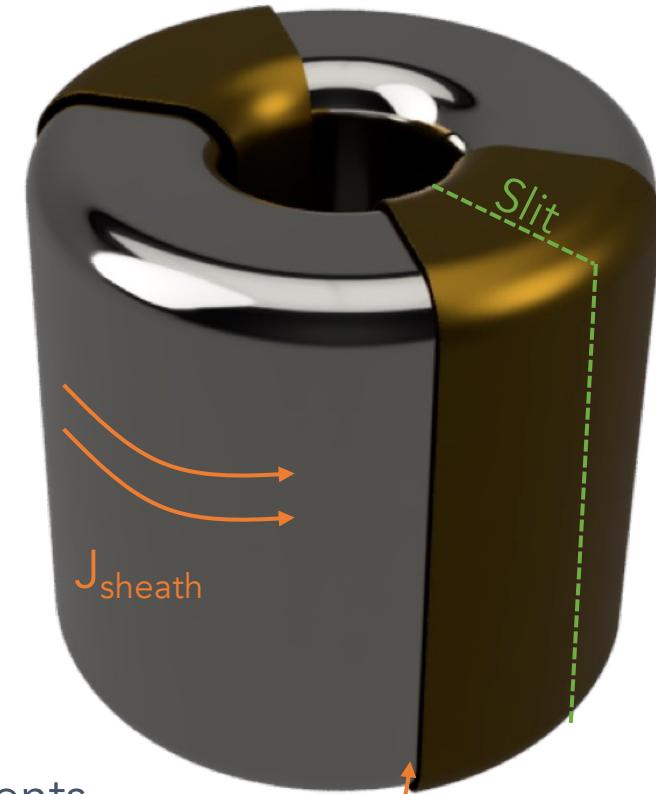
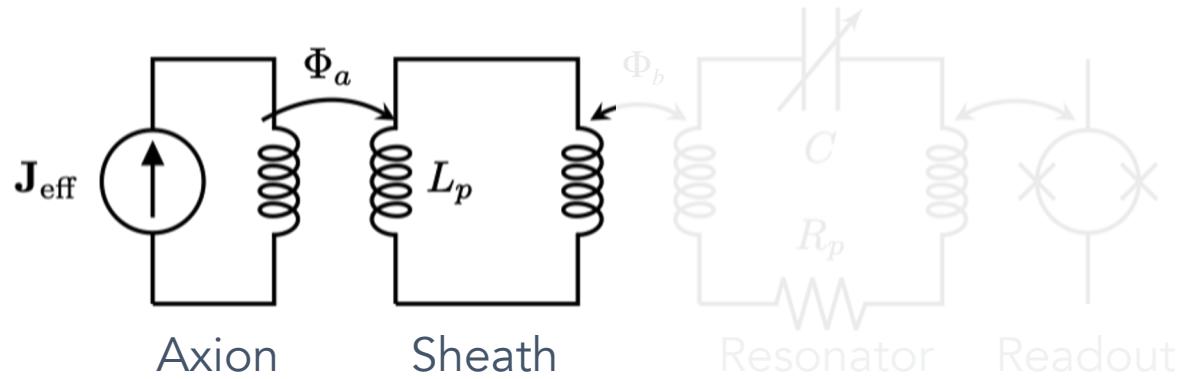


$\sim 1 \text{ T}$  magnet



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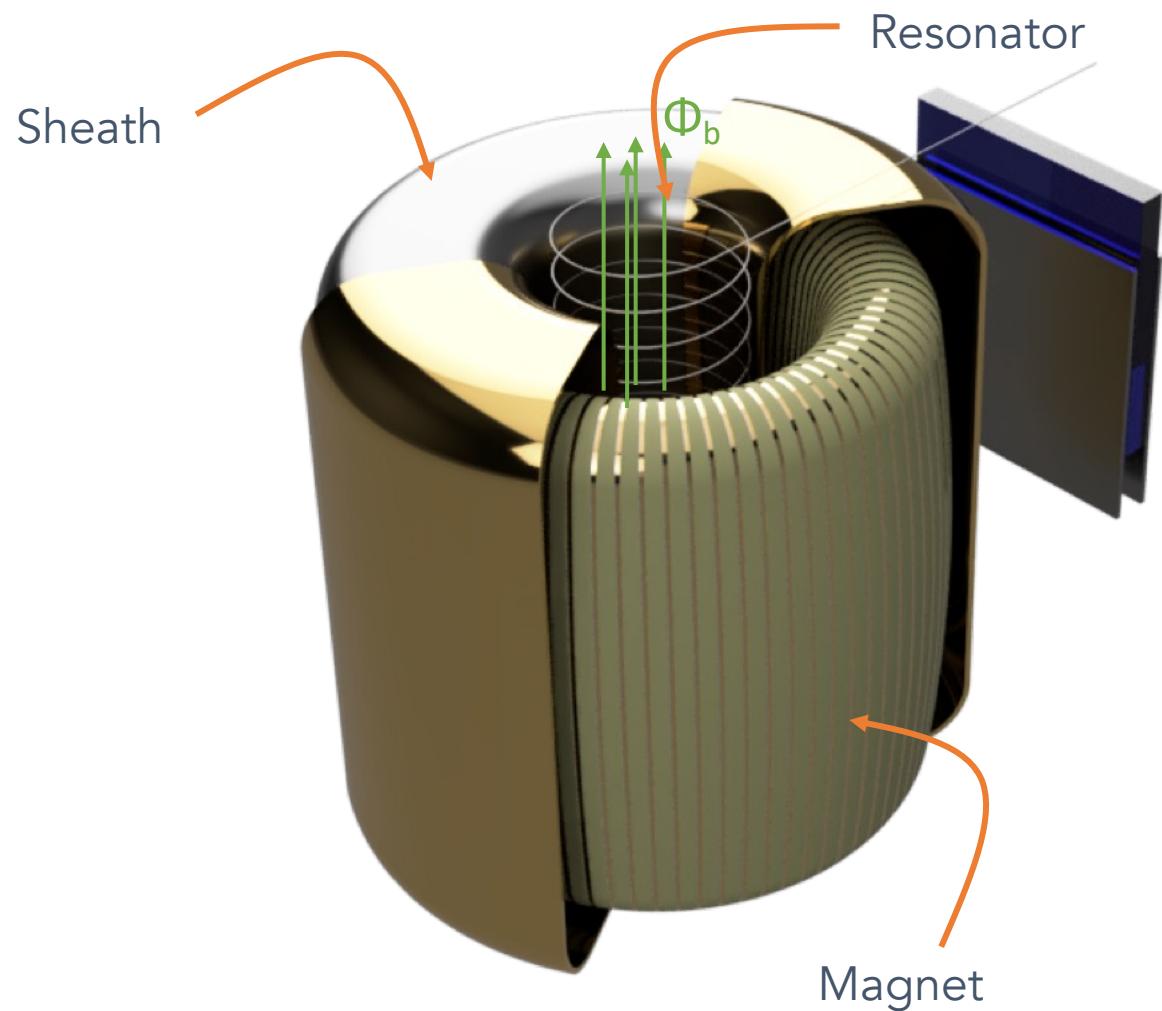
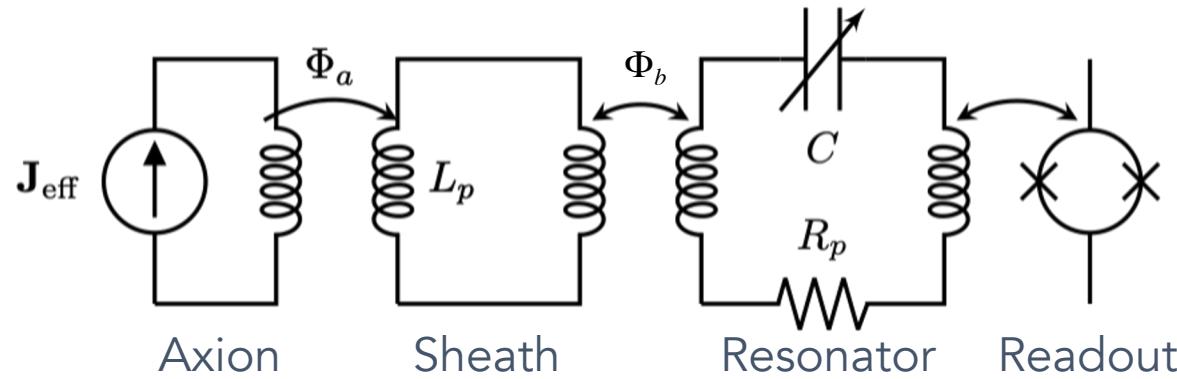


Nb sheath shields lossy magnet elements

Slit (+ sleeve) allows currents to flow outside

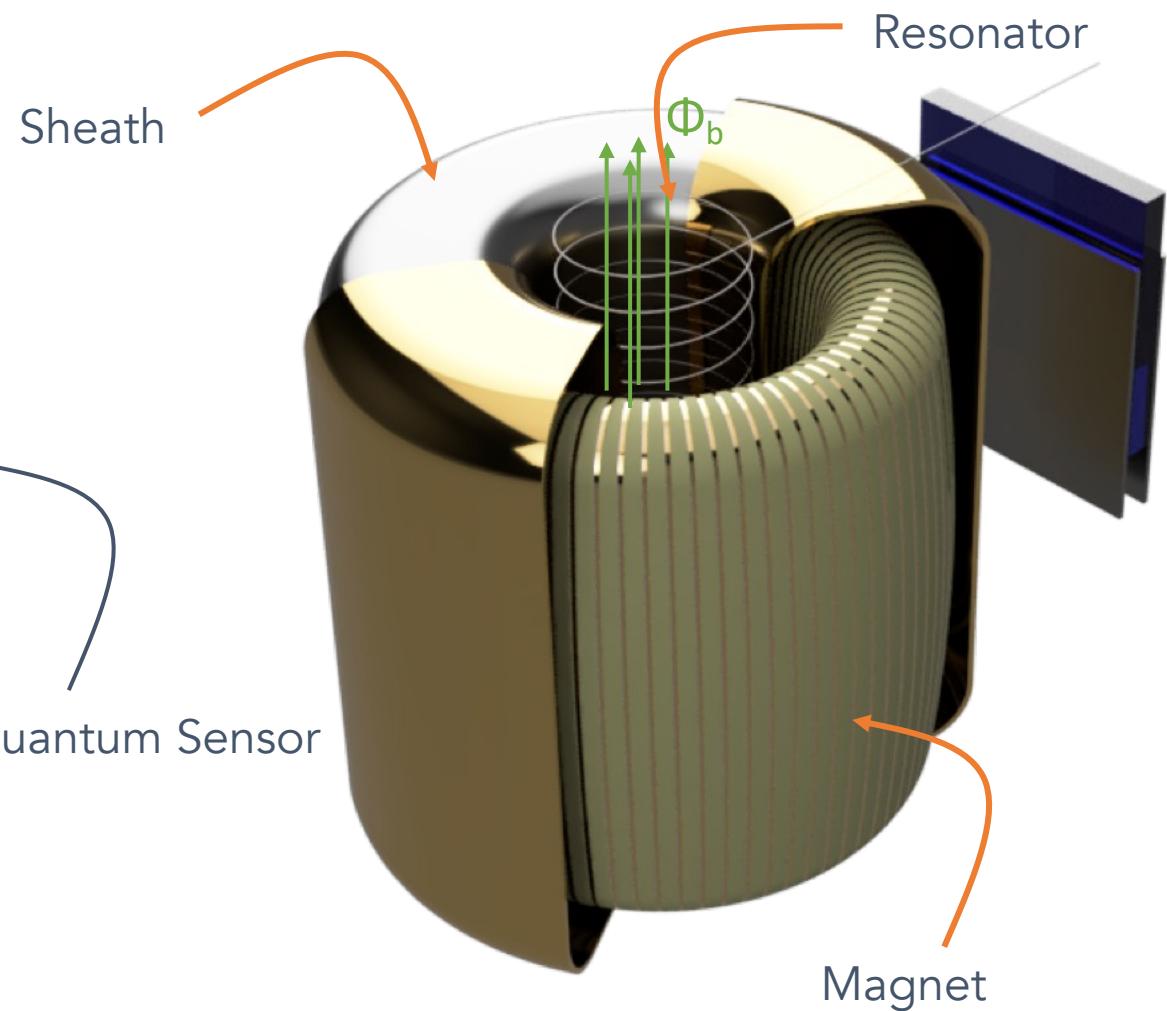
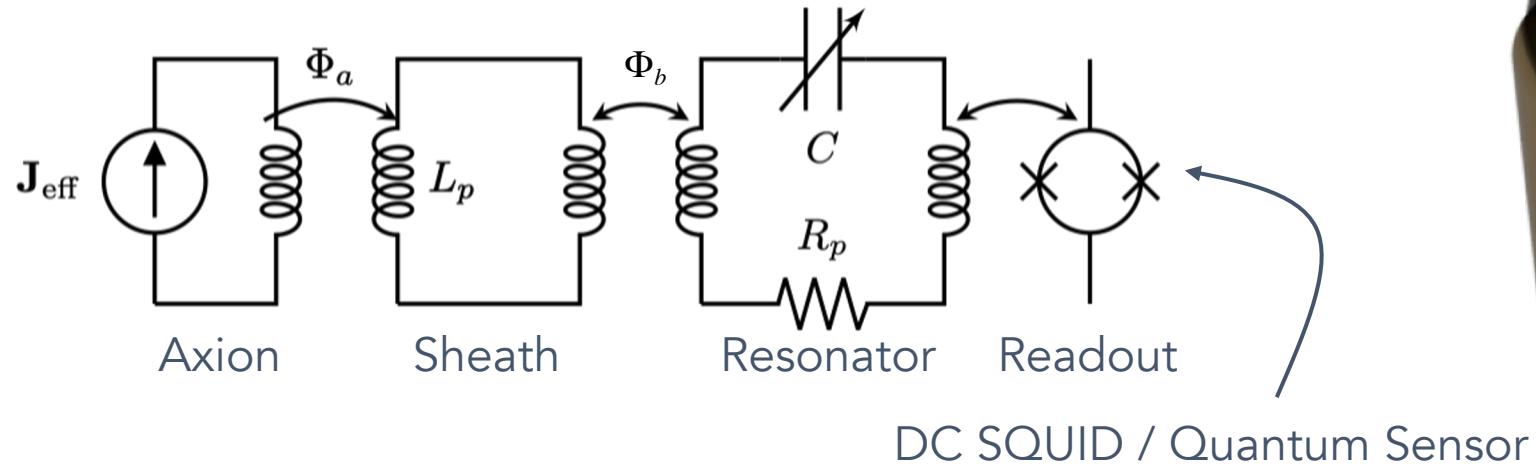
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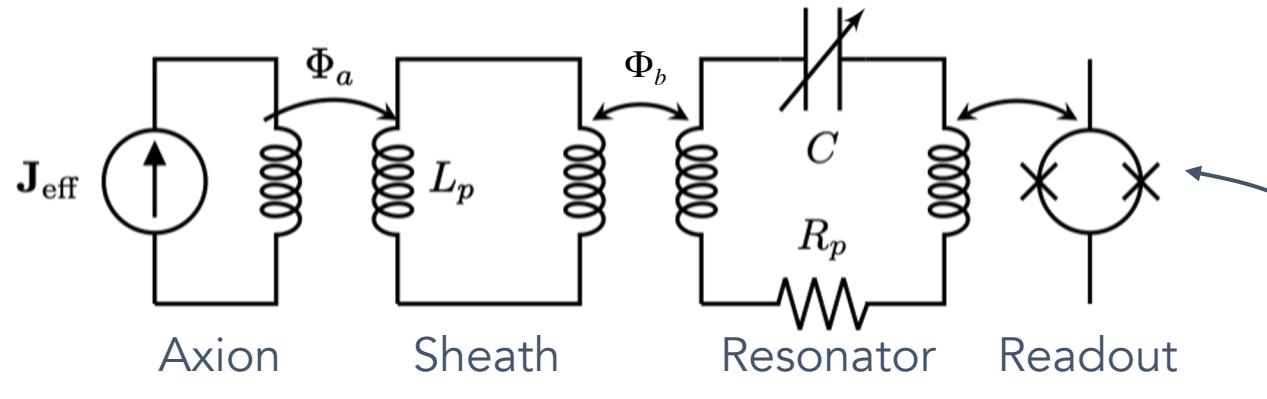
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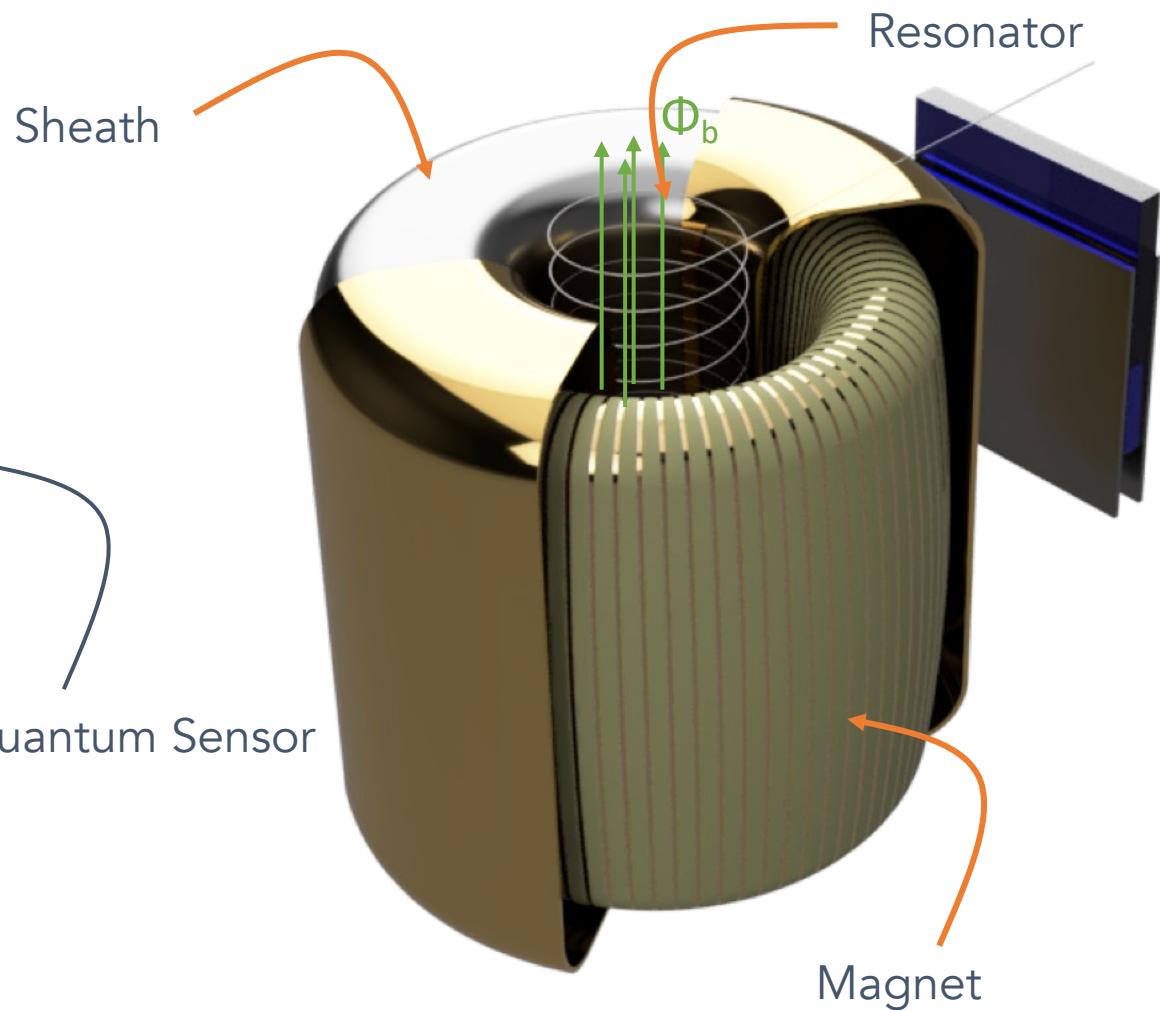
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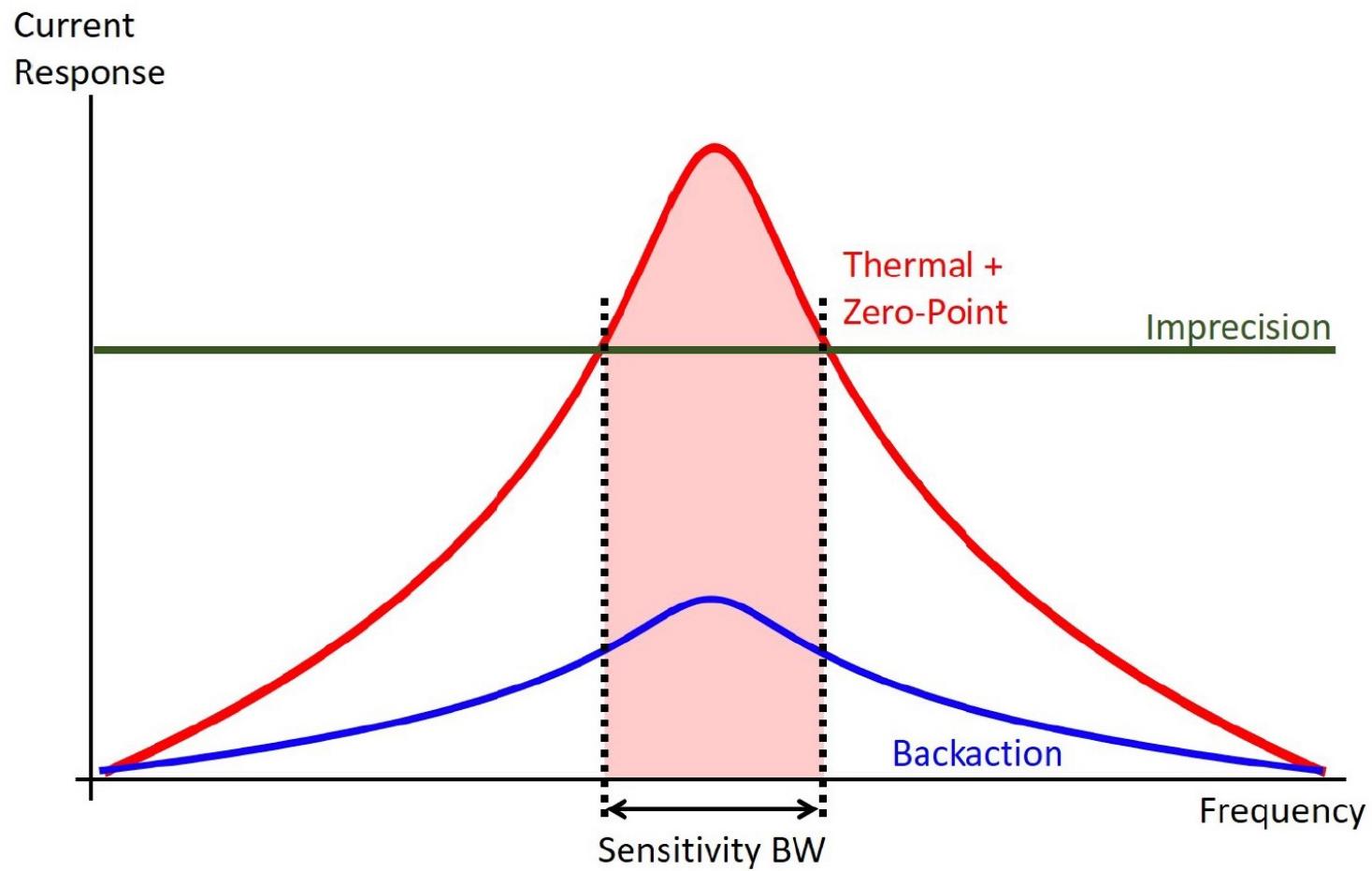
Quality factor of circuit  $\sim 10^6$

Peak B-field  $\sim 1$  T

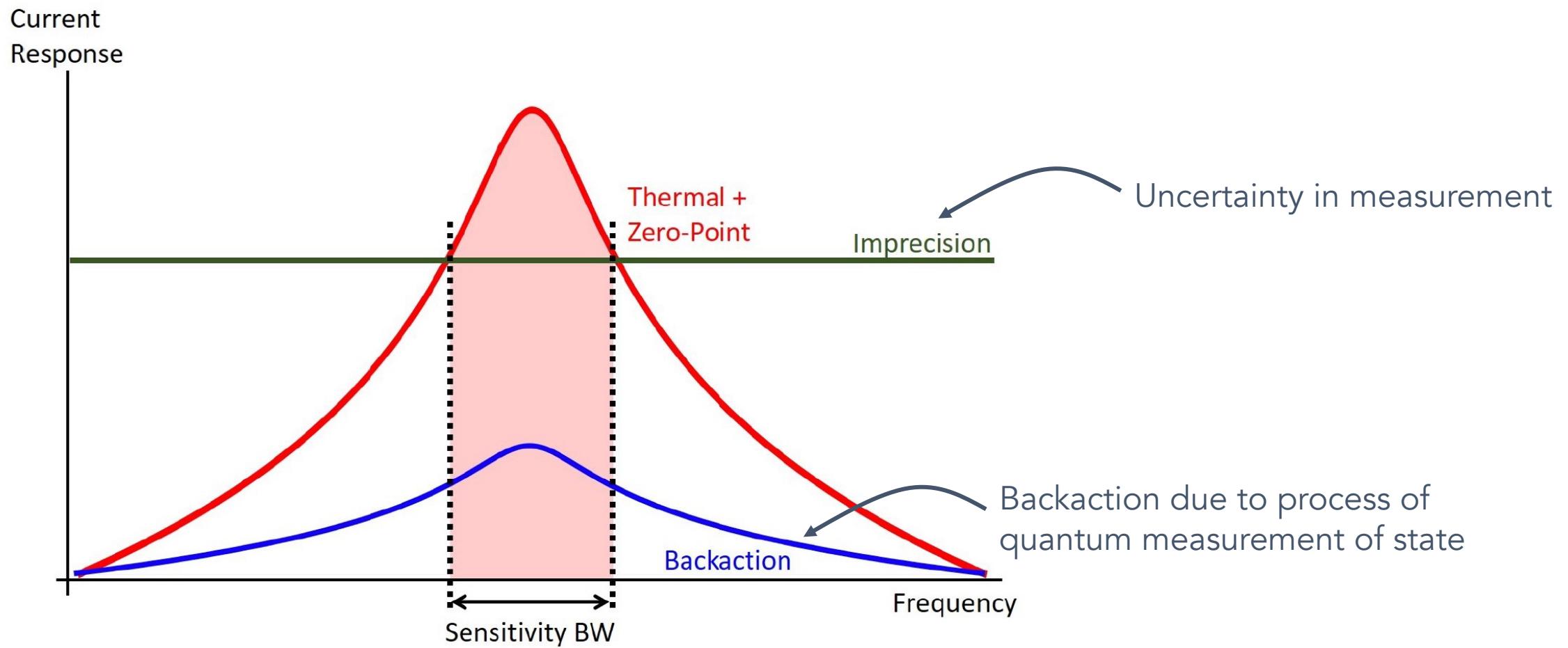
Physical temperature of resonator  $< 100$  mK



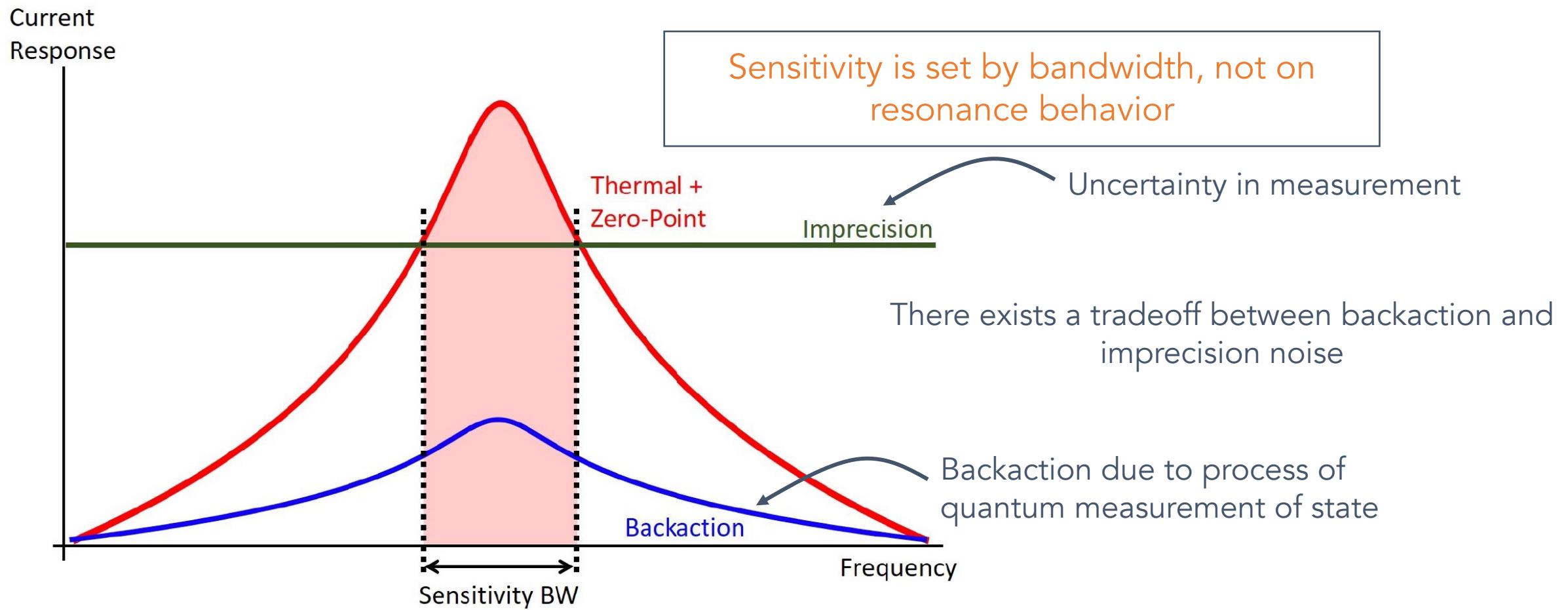
# Integrated Sensitivity and Quantum Sensors



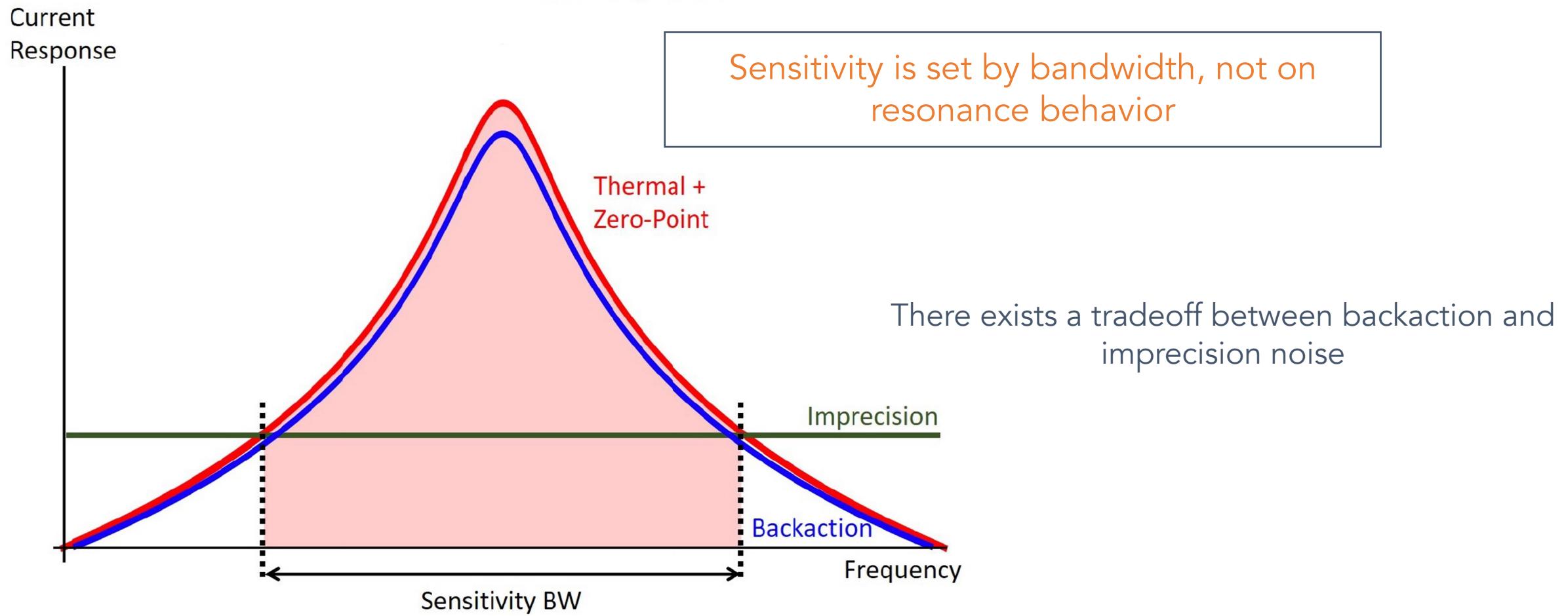
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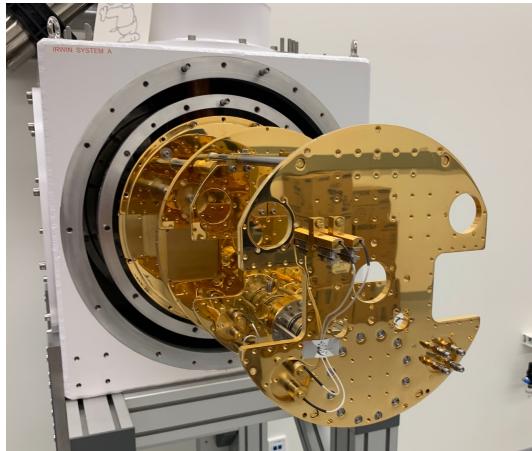


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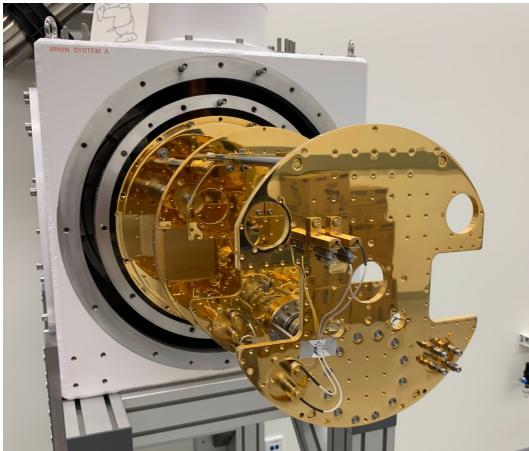
# 50L Status

BlueFors Dil Fridge at Stanford

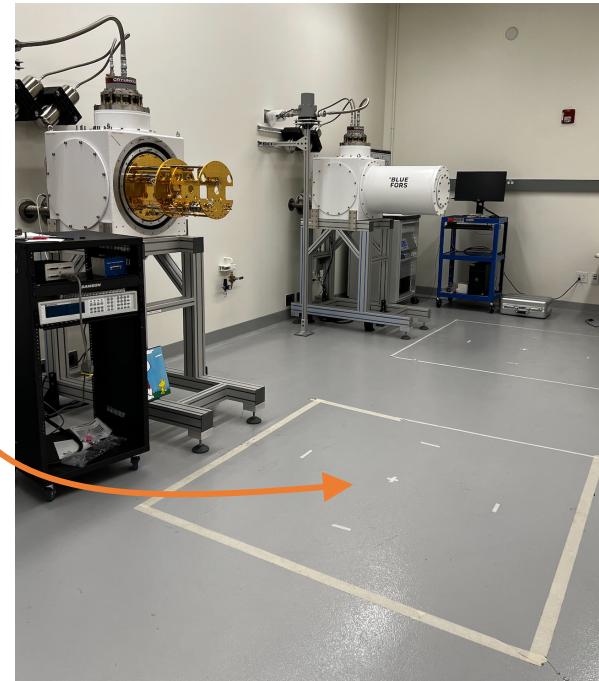


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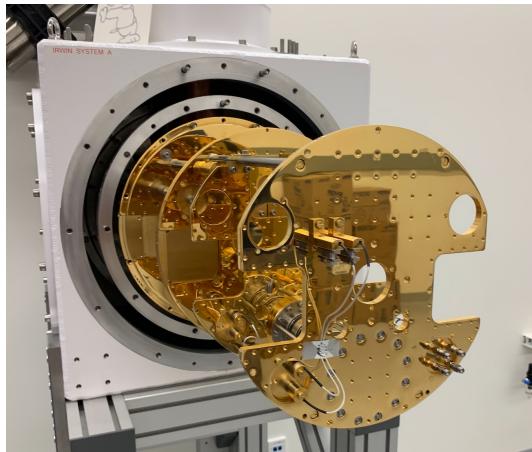


Cryostat currently being manufactured

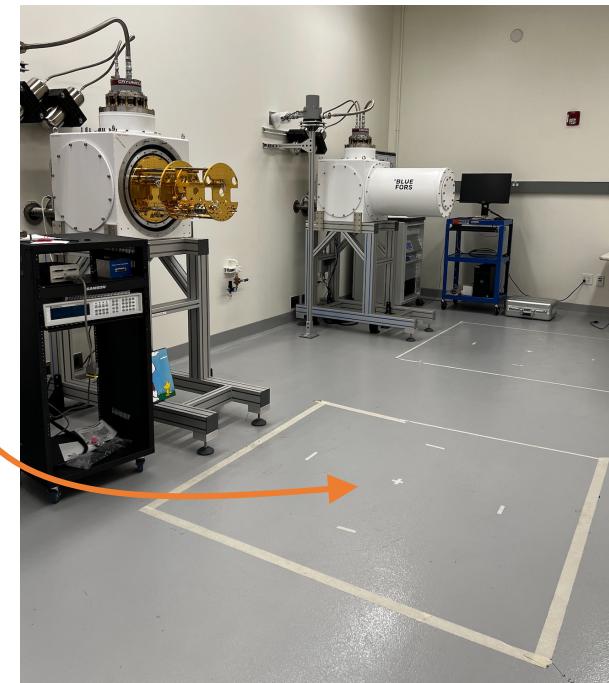


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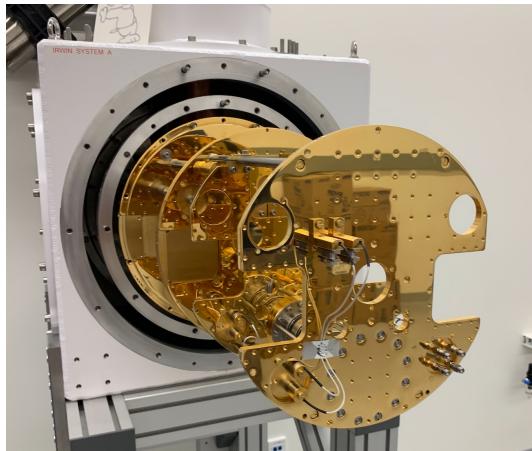


Magnet being manufactured by SSI

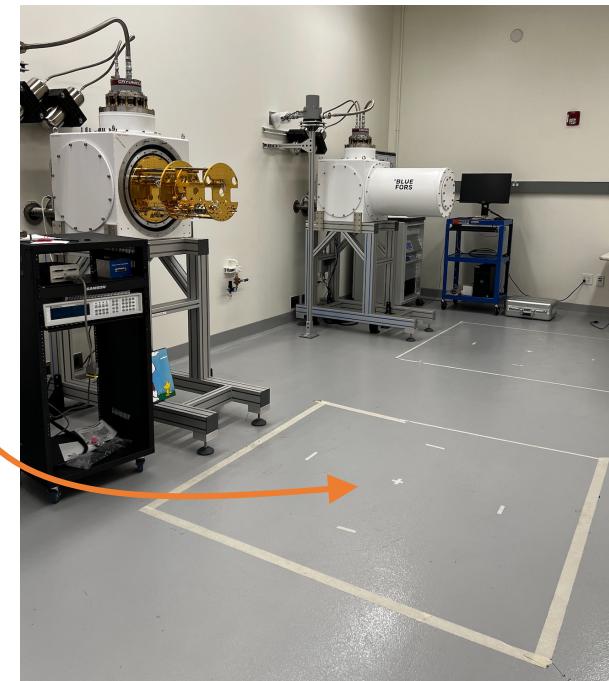


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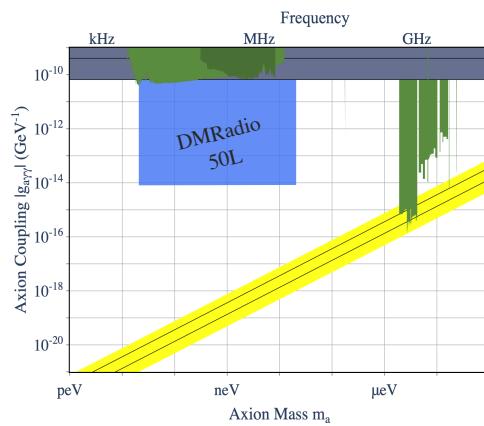
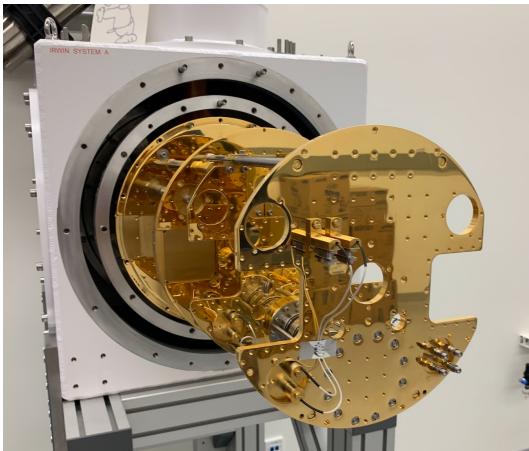


Sheath design being finalized

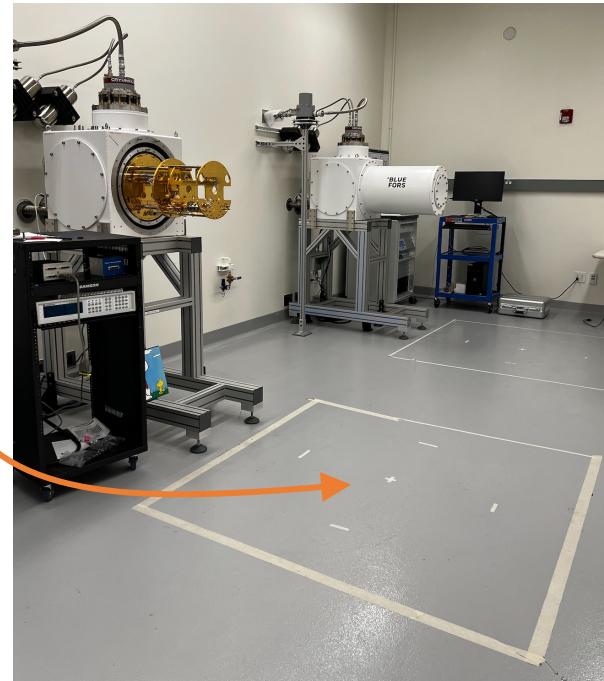


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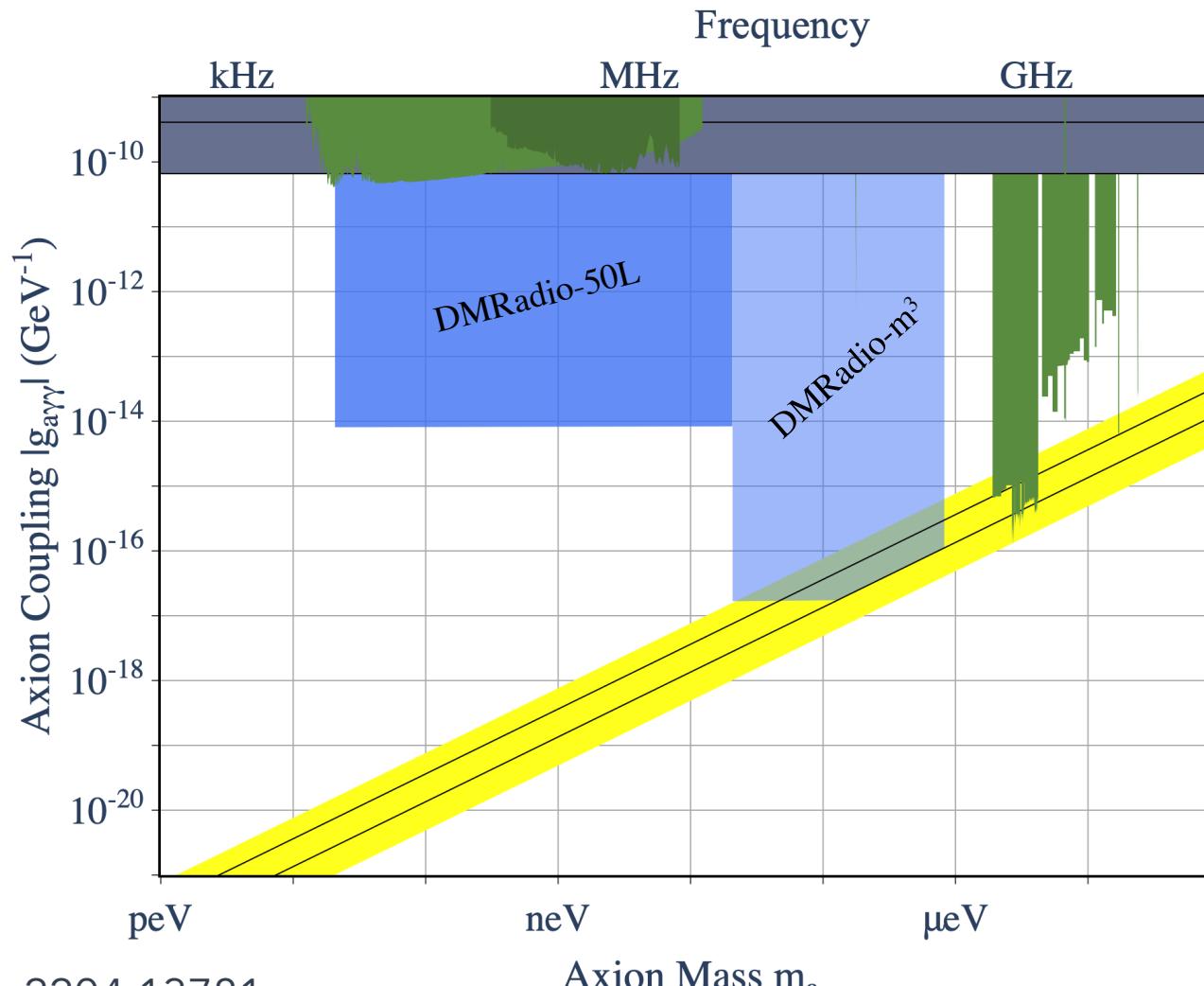


Science starting in 2023

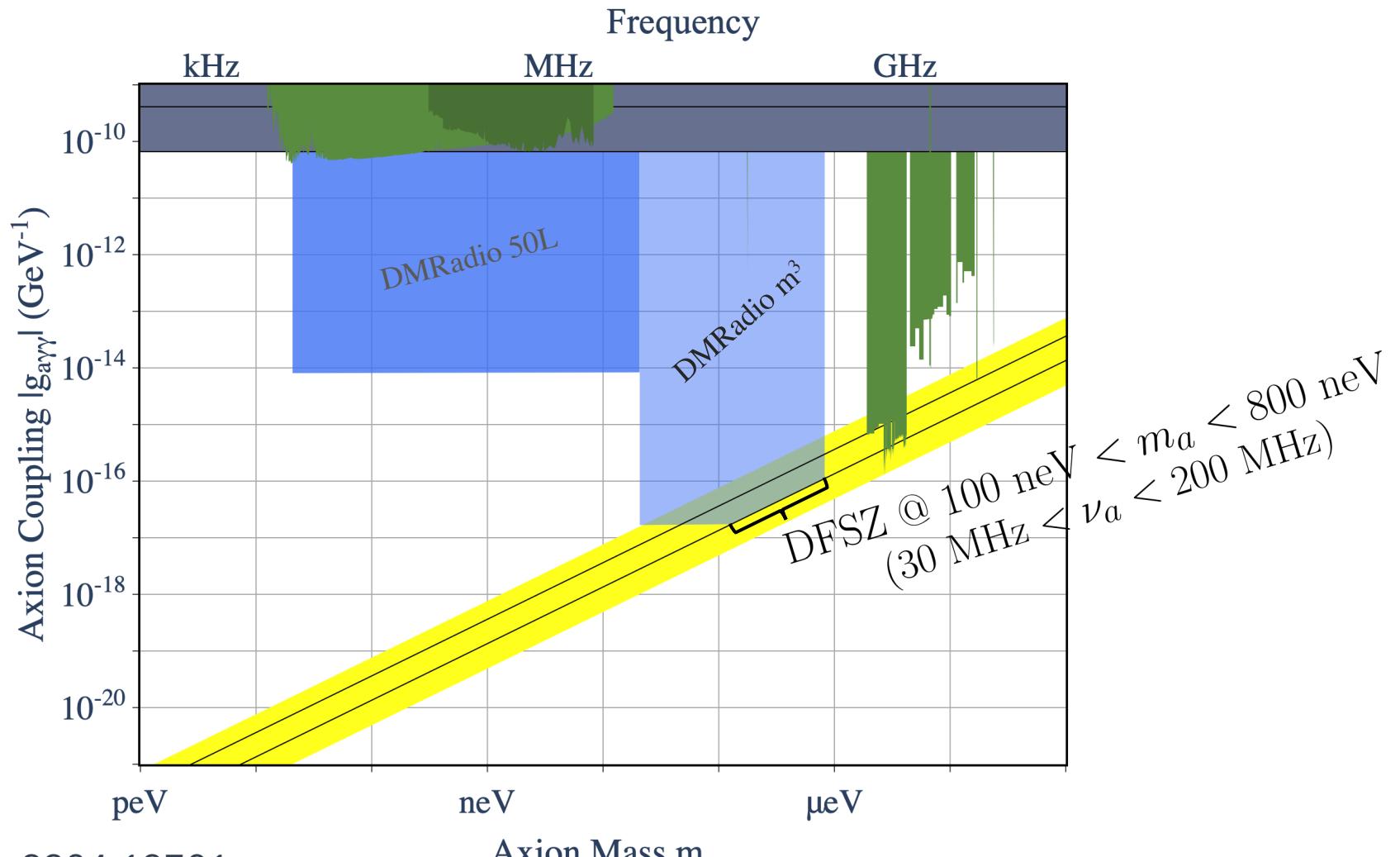
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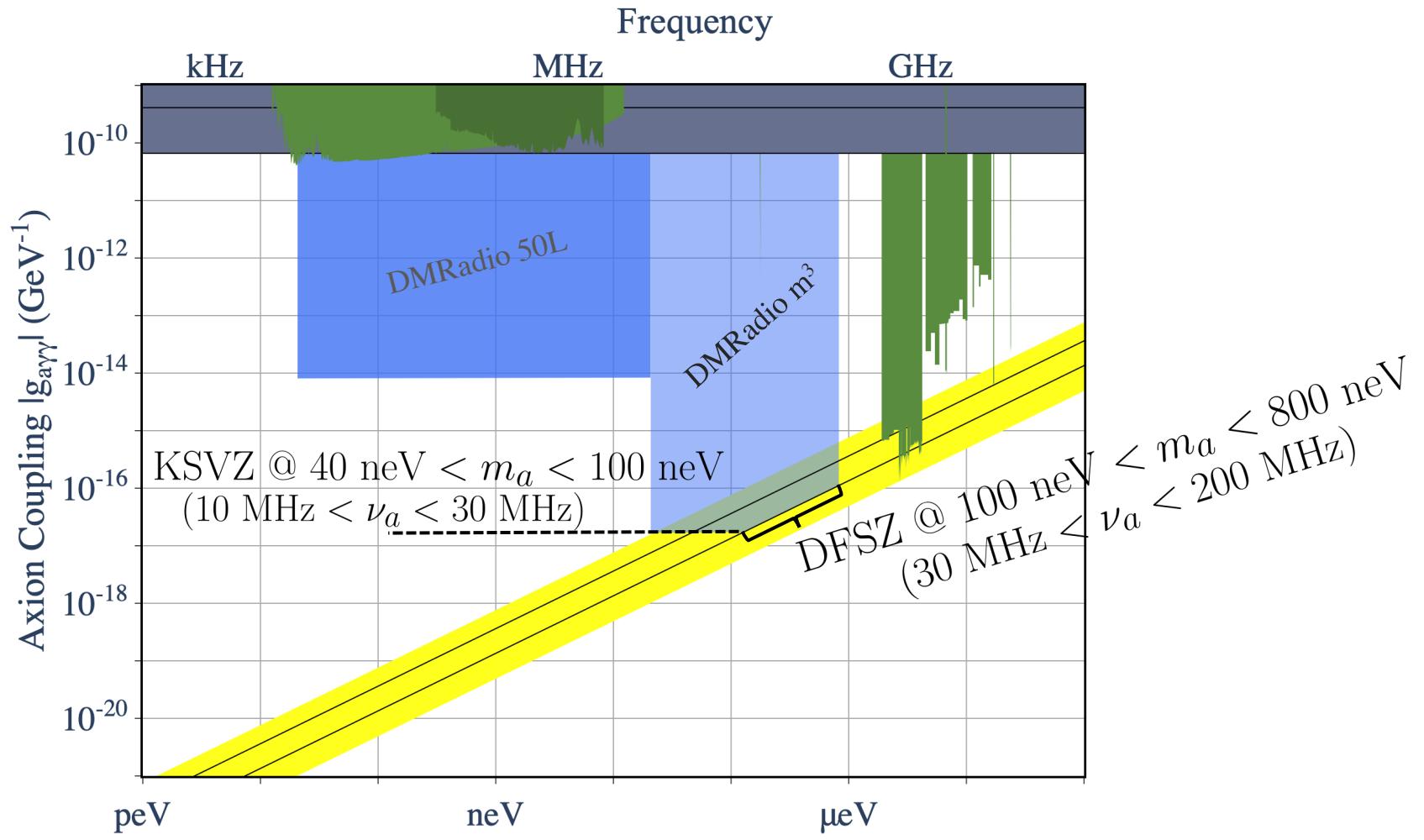
# DMRadio m<sup>3</sup> Goals



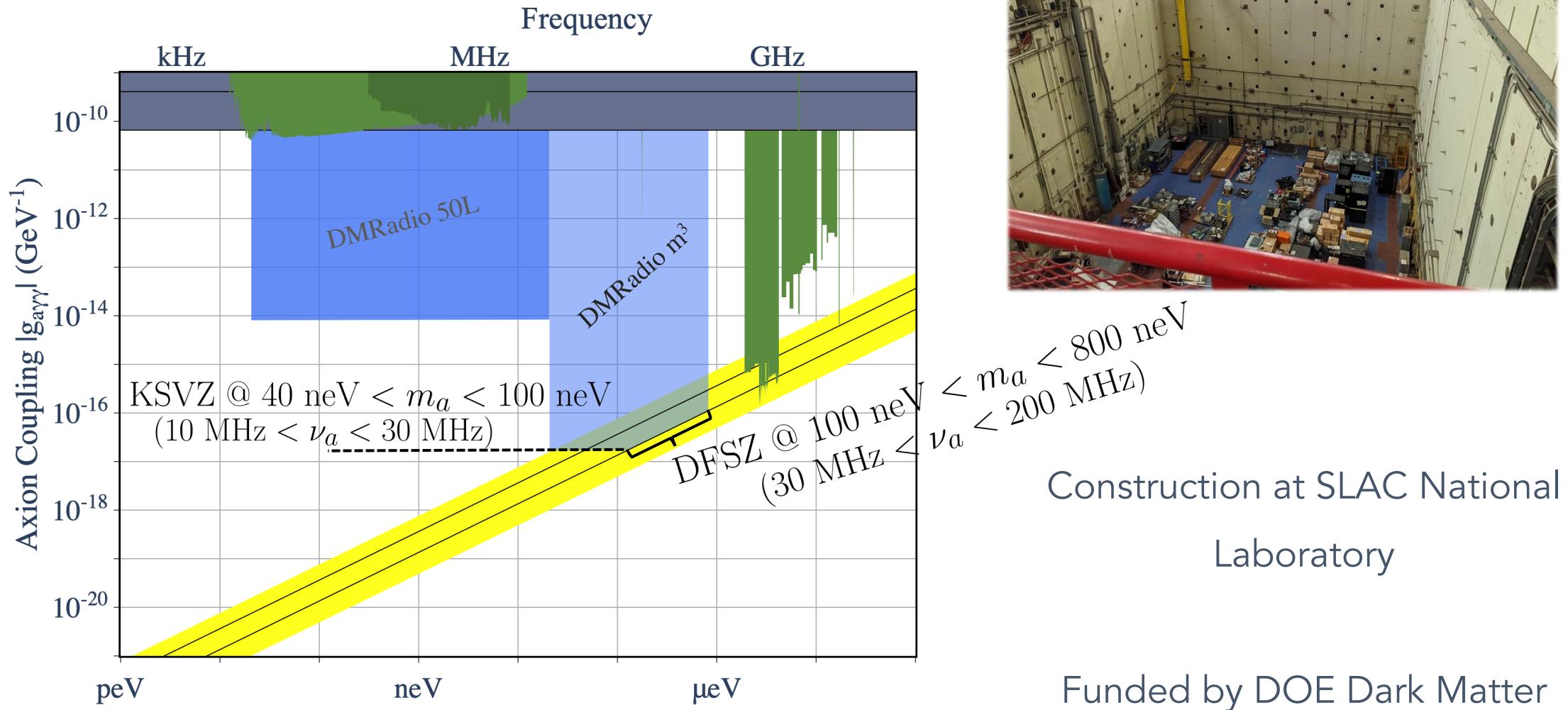
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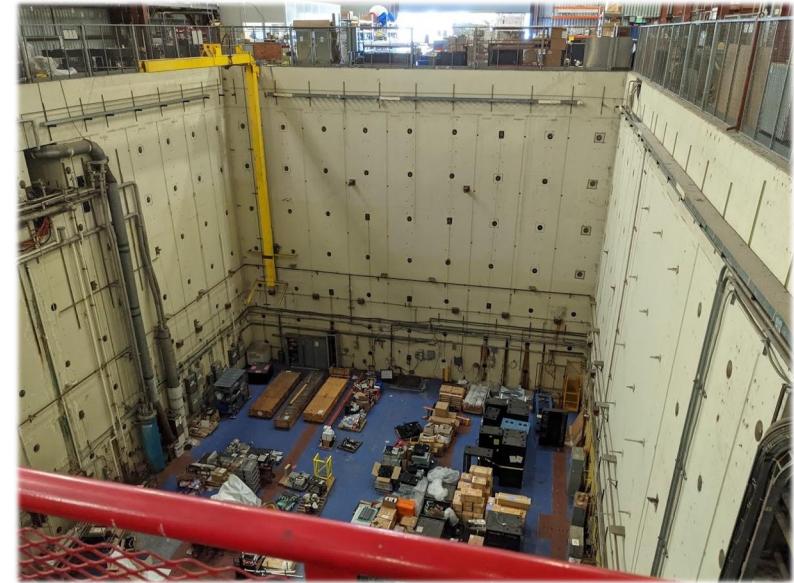
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arXiv: 2204.13781



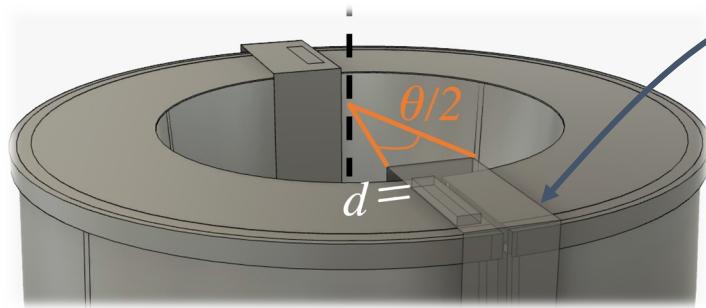
Construction at SLAC National  
Laboratory

Funded by DOE Dark Matter  
New Initiatives

# DMRadio m<sup>3</sup> Design

50L-like geometry non optimal – problems  
at high frequency

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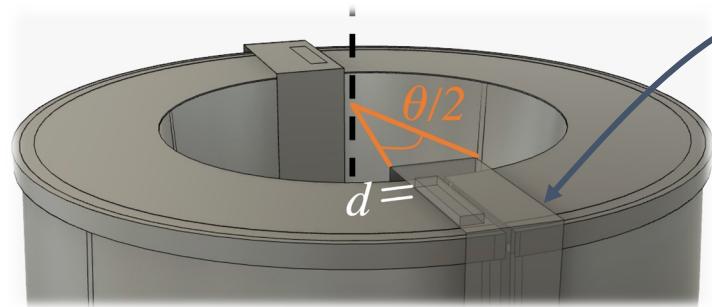


DMRadio 50L

Parasitic capacitive coupling shorts out signal!

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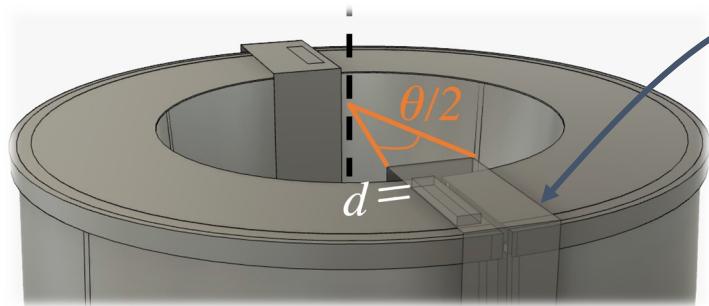
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$$f_{\text{paras}} = 48.5 \text{ MHz} \left( \frac{0.05 \text{ m}^3}{V_{\text{sheath}}} \right)^{1/2} \left( \frac{1 \text{ rad}}{\theta} \right)^{1/2} \left( \frac{d}{1 \text{ cm}} \right)^{1/2}$$

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50L-like geometry non optimal – problems  
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Axion signal will destructively  
interfere with **cavity modes** of this  
coaxial structure (as  $\lambda \sim l$ )

# m<sup>3</sup> design

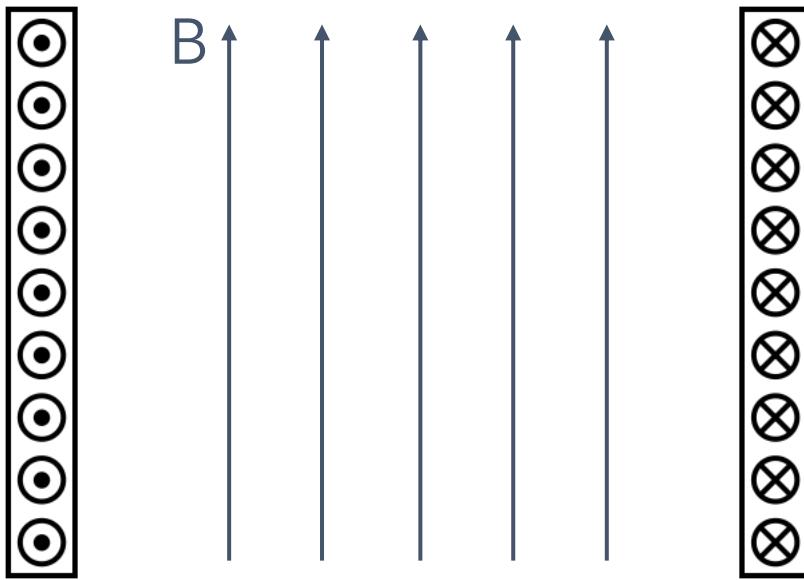
m<sup>3</sup> uses a solenoidal magnet + coaxial copper pickup

No parasitic capacitance

# $m^3$ design

$m^3$  uses a solenoidal magnet + coaxial copper pickup

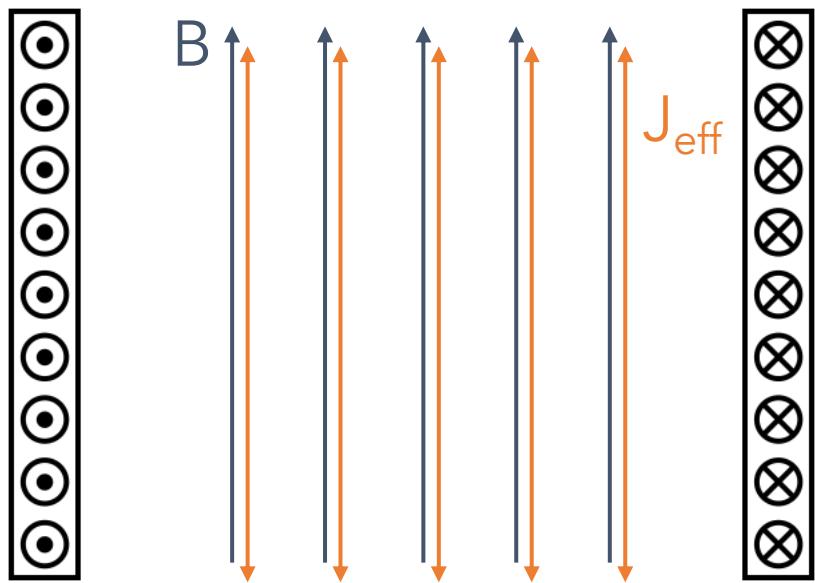
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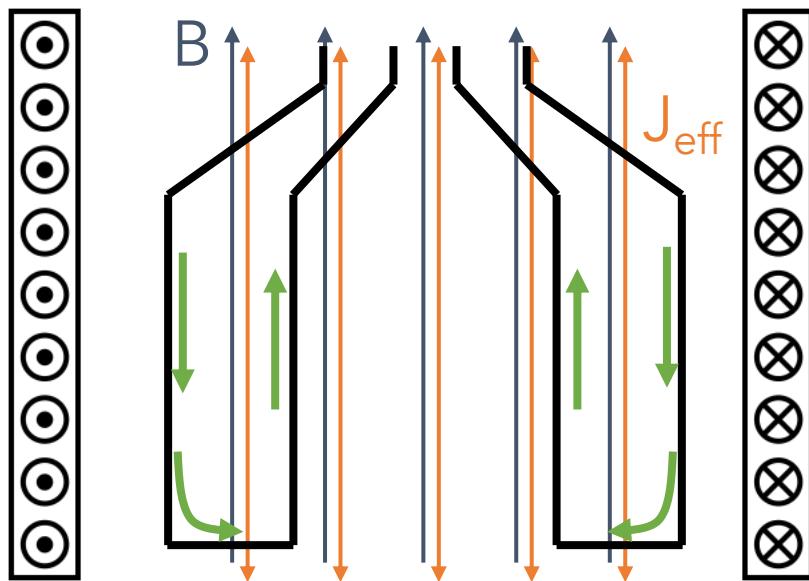
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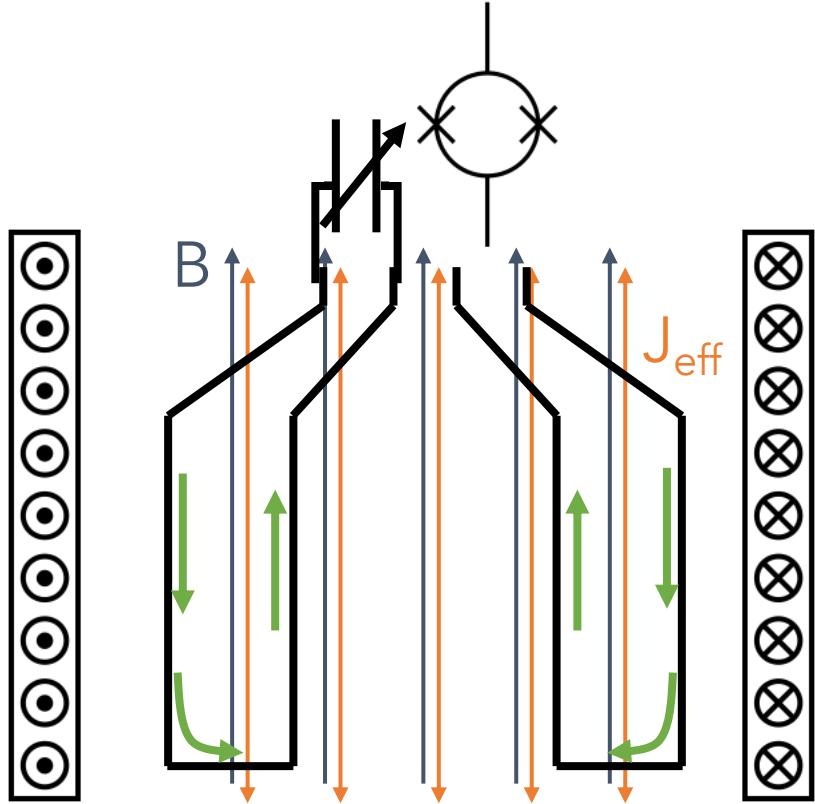
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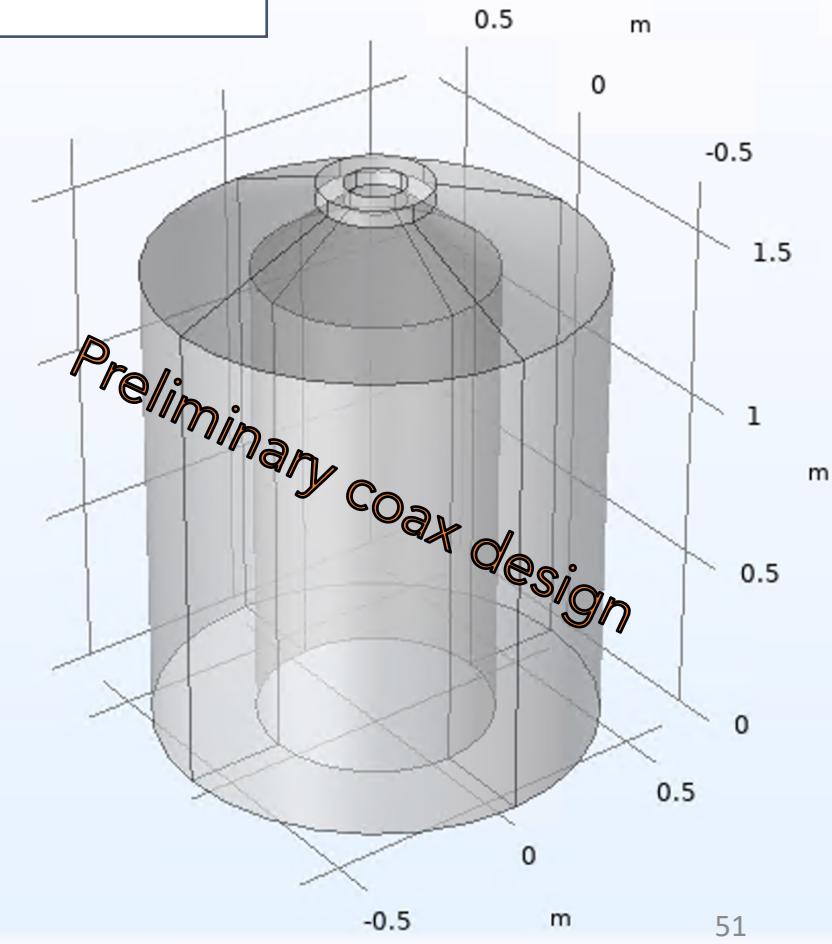
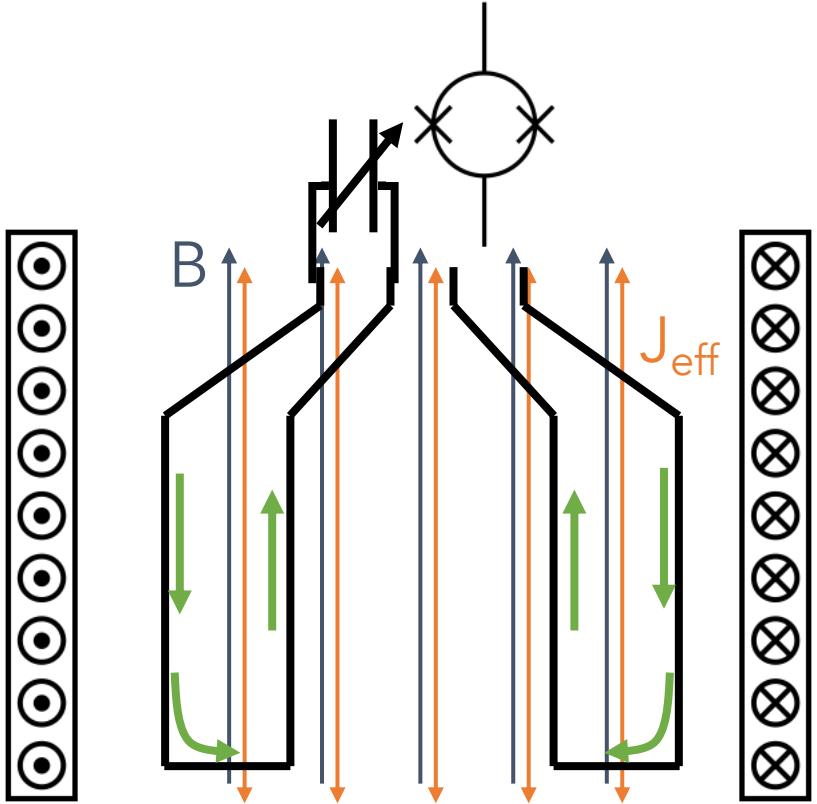
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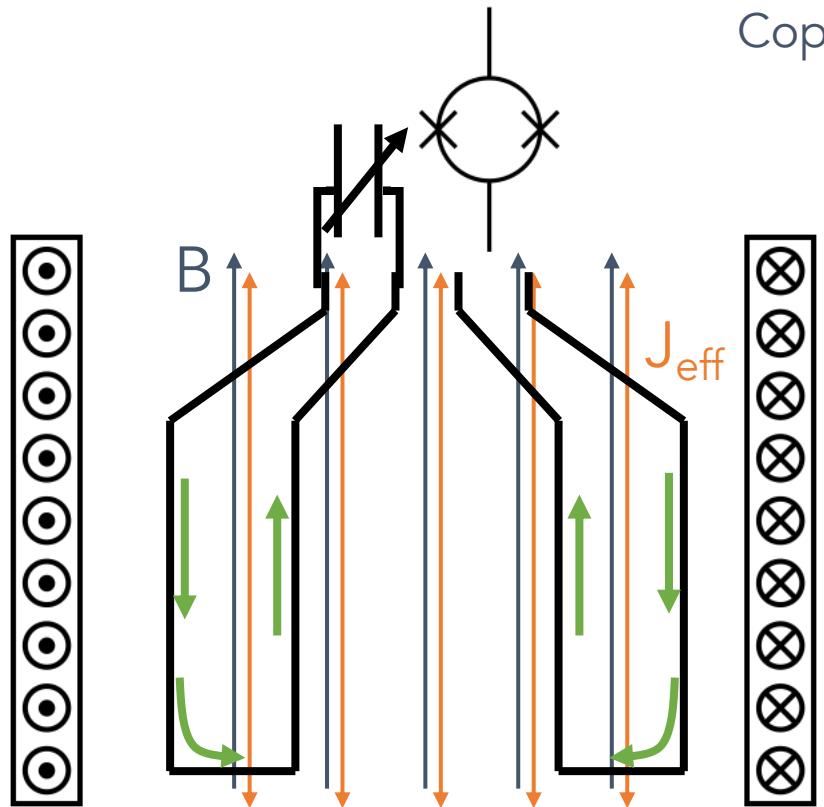
$m^3$  uses a solenoidal magnet + coaxial copper pickup

No parasitic capacitance



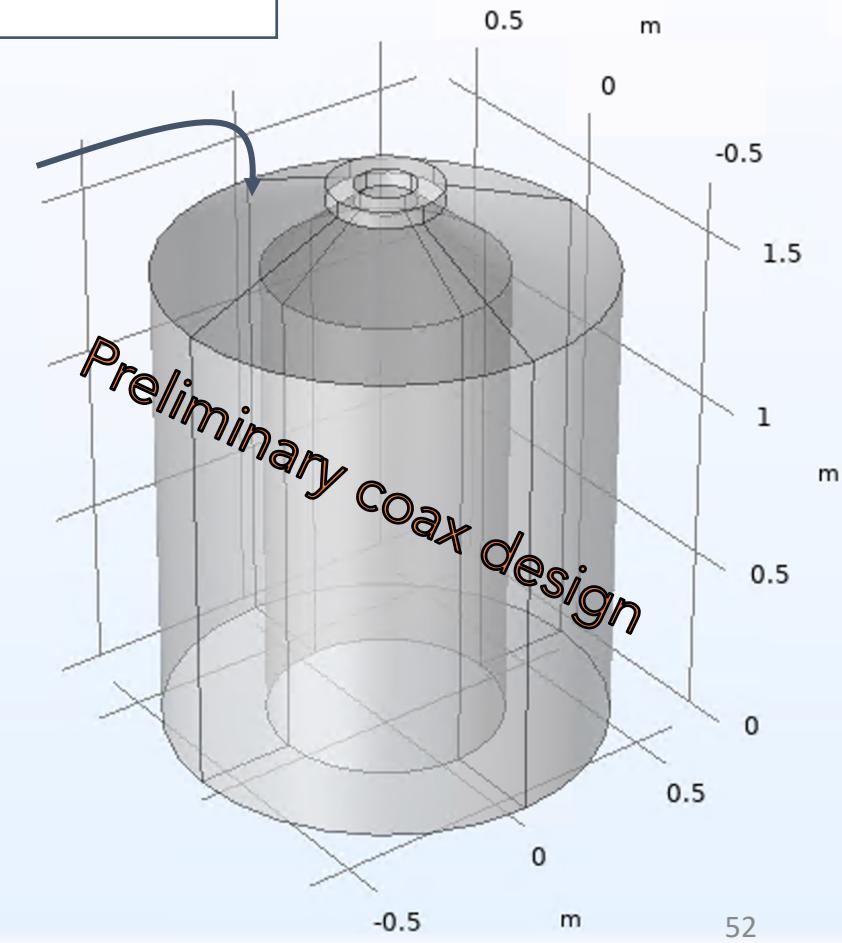
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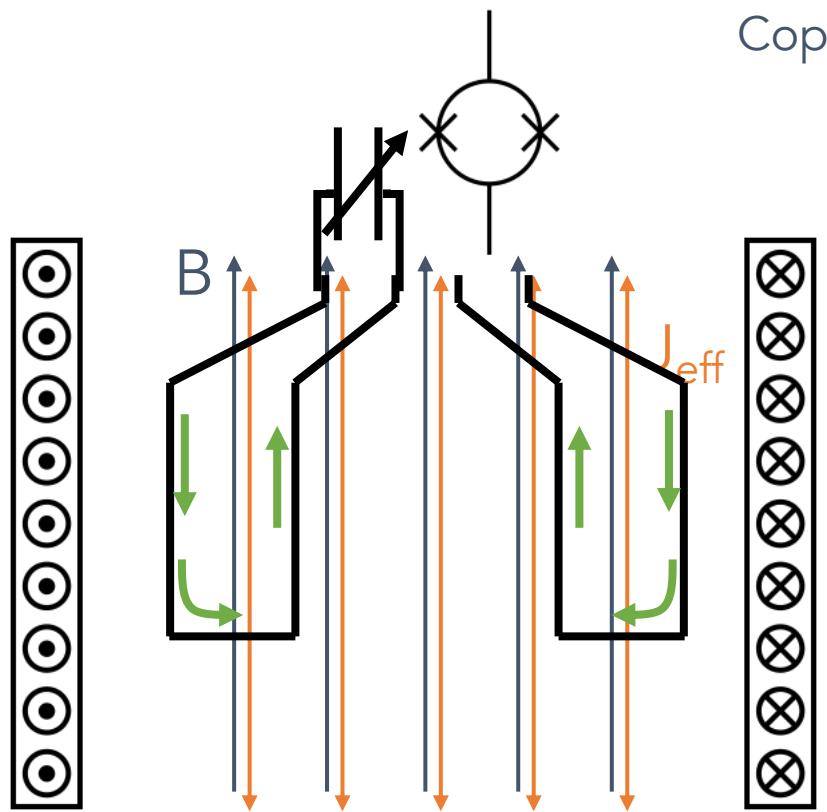
Copper pickup in high B-field region

Coax design favorable  
for higher frequencies—  
need careful treatment  
nevertheless



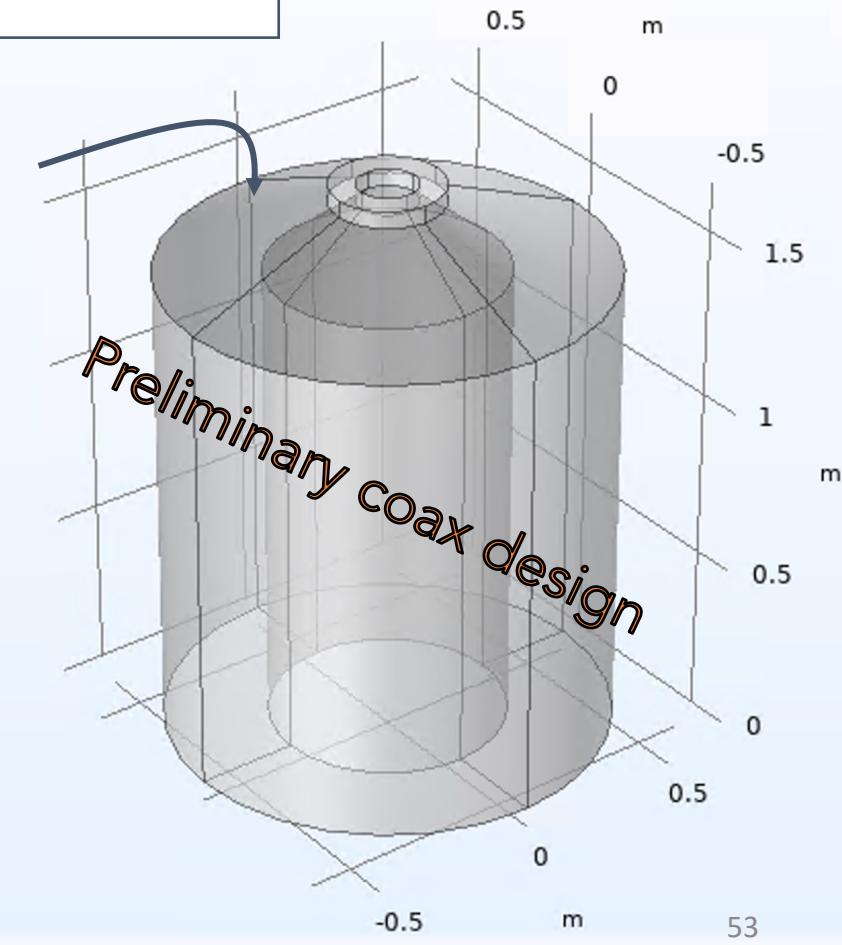
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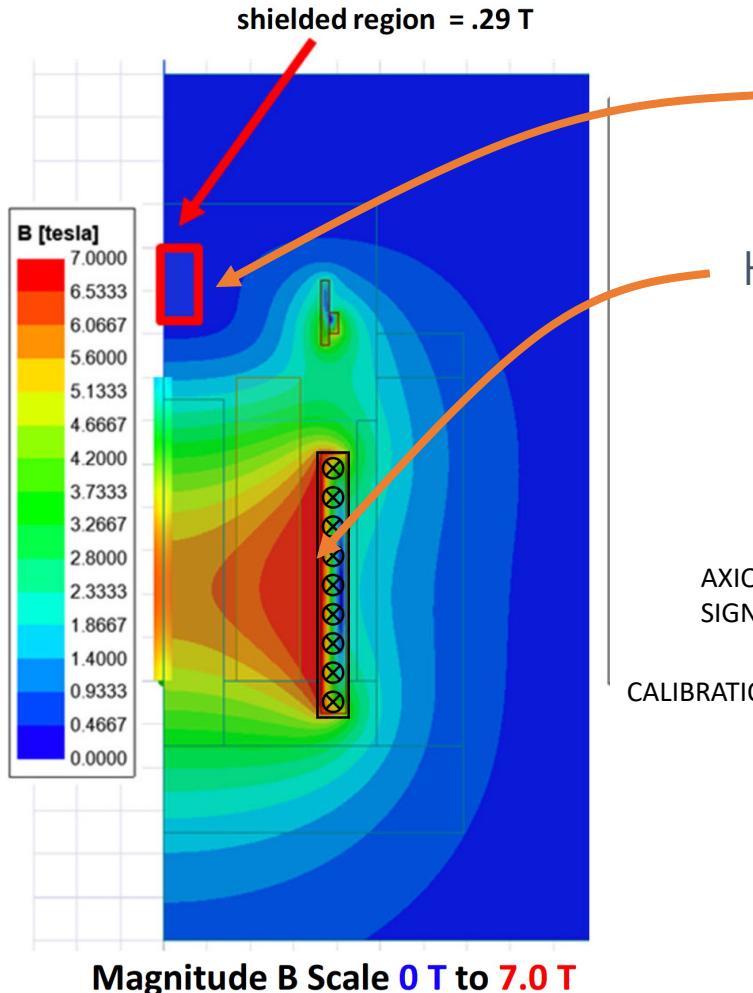


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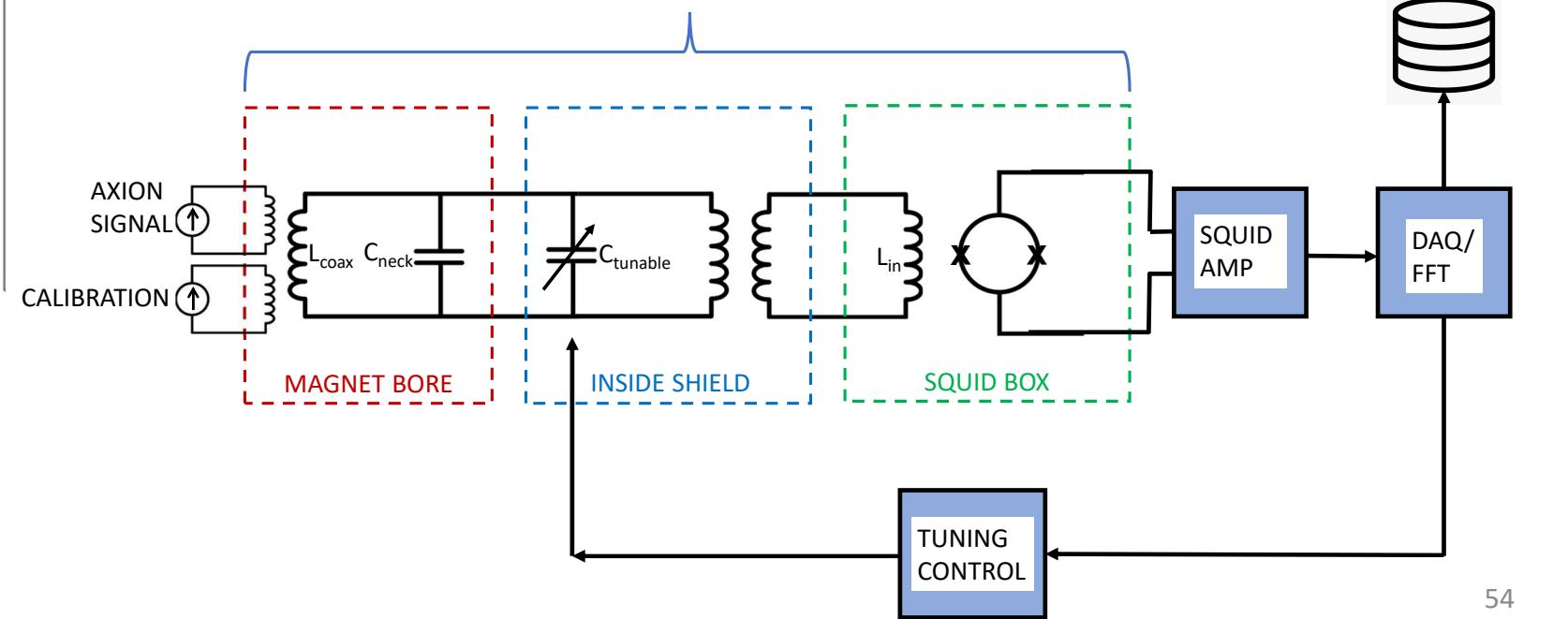
# $m^3$ design



Low B-field region allows for superconducting electronics (less loss)

High B-field region produces larger  $J_{\text{eff}}$

COLD COMPONENTS (Magnet at 4.5 K, LC resonator and SQUID at 20 mK)

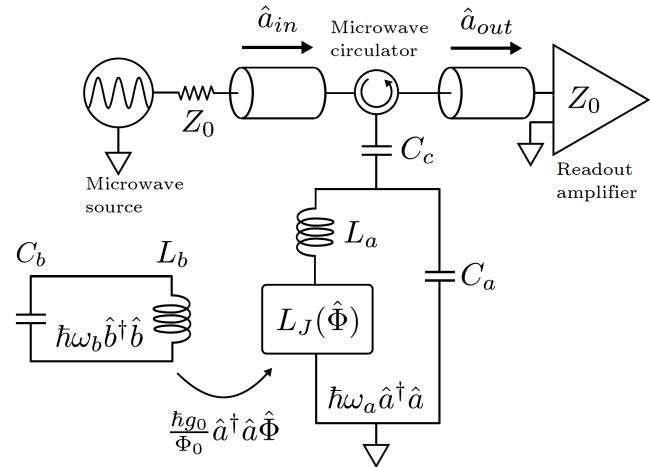


# Outline

1. Axions and low frequency haloscopes
2. DMRadio 50L
  1. Overview
  2. Status
3. DMRadio m<sup>3</sup>
  1. Overview
  2. Status
4. Outlook, future, & collaboration

# DMRadio GUT

RF Quantum Upconverters:



Magnet:

Low vs High  $T_C$  still under consideration – significant cryogenic constraints

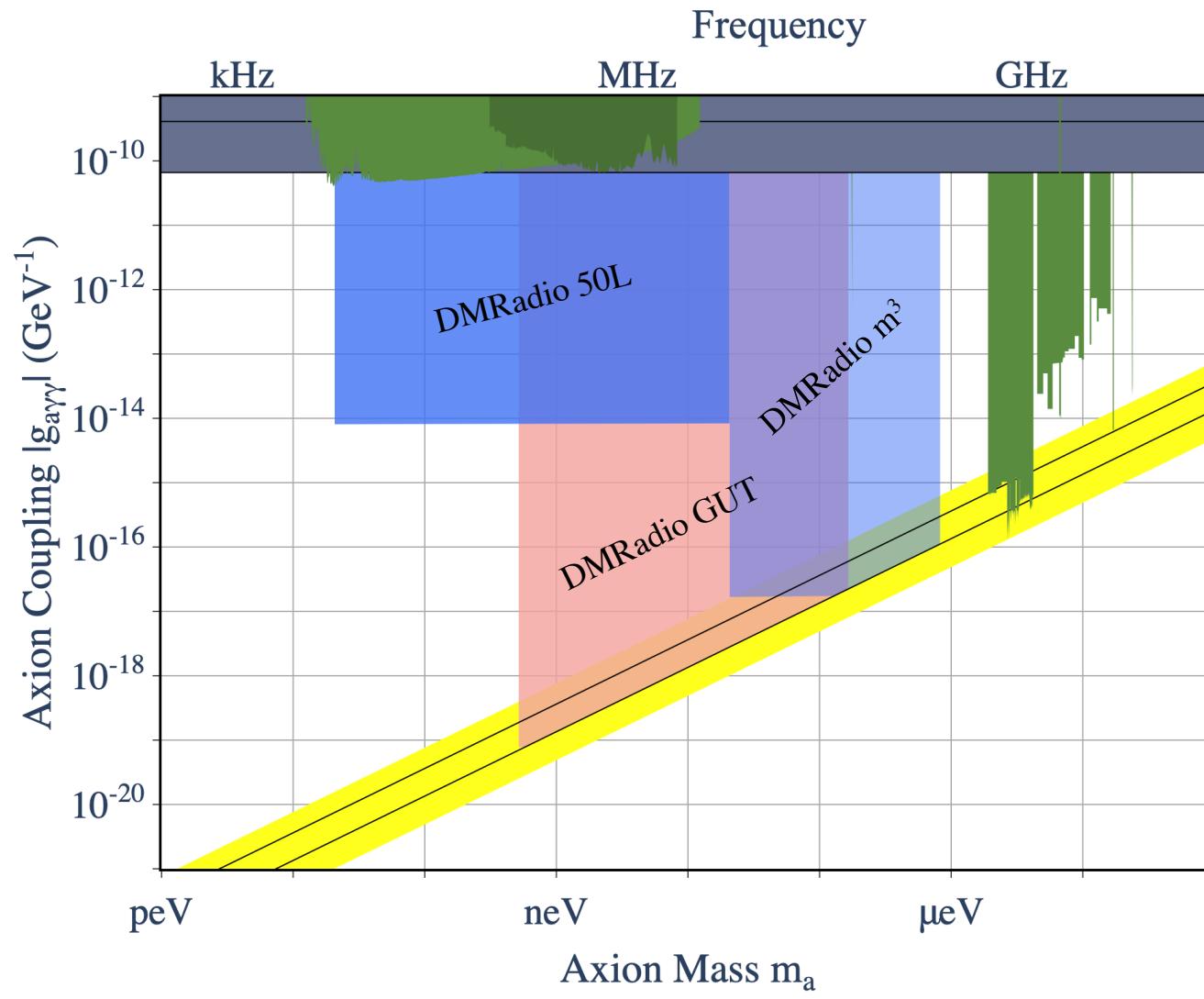
16 T, 10 m<sup>3</sup>

Scan Rate:

$$\begin{aligned} \frac{d\nu_r}{dt} \approx & 41 \frac{\text{kHz}}{\text{year}} \left( \frac{3}{\text{SNR}} \right)^2 \left( \frac{g_{a\gamma\gamma}}{10^{-19} \text{ GeV}^{-1}} \right)^4 \left( \frac{\rho_{\text{DM}}}{0.45 \text{ GeV/cm}^3} \right)^2 \\ & \times \left( \frac{\nu_r}{100 \text{ kHz}} \right) \left( \frac{c_{PU}}{0.1} \right)^4 \left( \frac{B_0}{16 \text{ T}} \right)^4 \left( \frac{V}{10 \text{ m}^3} \right)^{10/3} \left( \frac{Q}{2 \times 10^7} \right) \left( \frac{10 \text{ mK}}{T} \right) \left( \frac{0.1}{\eta_A} \right). \end{aligned}$$

$Q \sim 20 \times 10^6$

# DMRadio Outlook



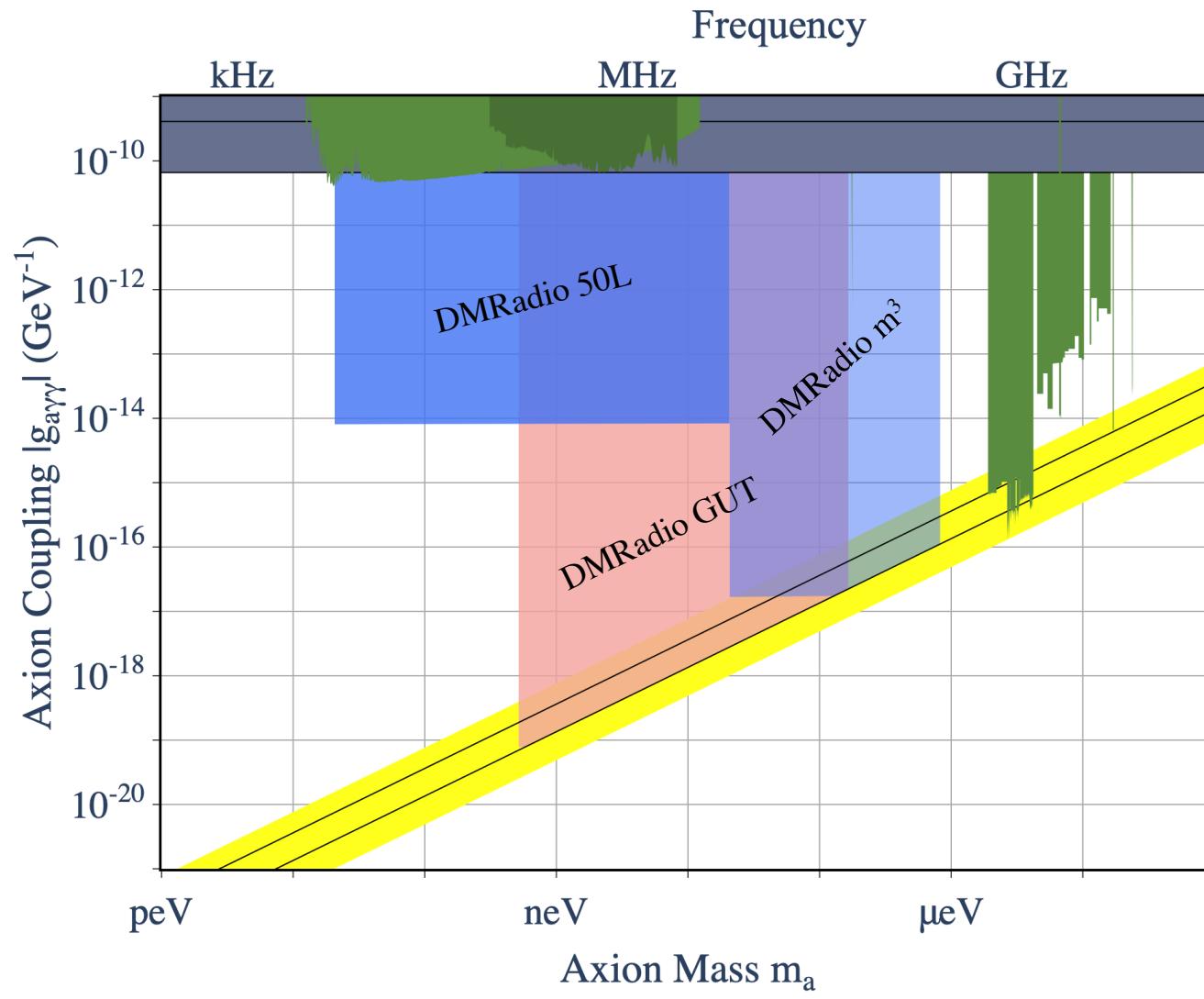
DFSZ @  $0.4 \text{ neV} < m_a < 800 \text{ neV}$

Projects developed in parallel

$f_a < 10^{19} \text{ GeV}$  for QCD axion models

Testbed for quantum devices

# DMRadio Outlook



DFSZ @  $0.4 \text{ neV} < m_a < 800 \text{ neV}$

Projects developed in parallel

$f_a < 10^{19} \text{ GeV}$  for QCD axion models

Testbed for quantum devices

## Publications

50L: coming soon

$m^3$ : arXiv: 2204.13781

GUT: arXiv: 2203.11246

RF Quantum Upconverters:

arXiv: 2210.xxxxxx

# DMRadio Collaboration

H.M. Cho, W. Craddock, D. Li, C. P. Salemi, W. J. Wisniewski  
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*University of California Berkeley*

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*Princeton University*

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*California State University, East Bay*

B. R. Safdi  
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*University of California Berkeley*

## Publications

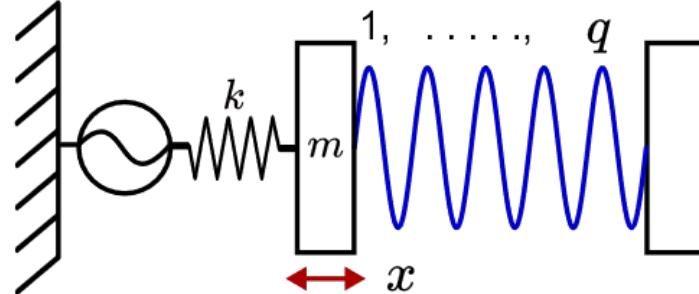
m<sup>3</sup>: arXiv: 2204.13781

GUT: arXiv: 2203.11246

# Backup Slides

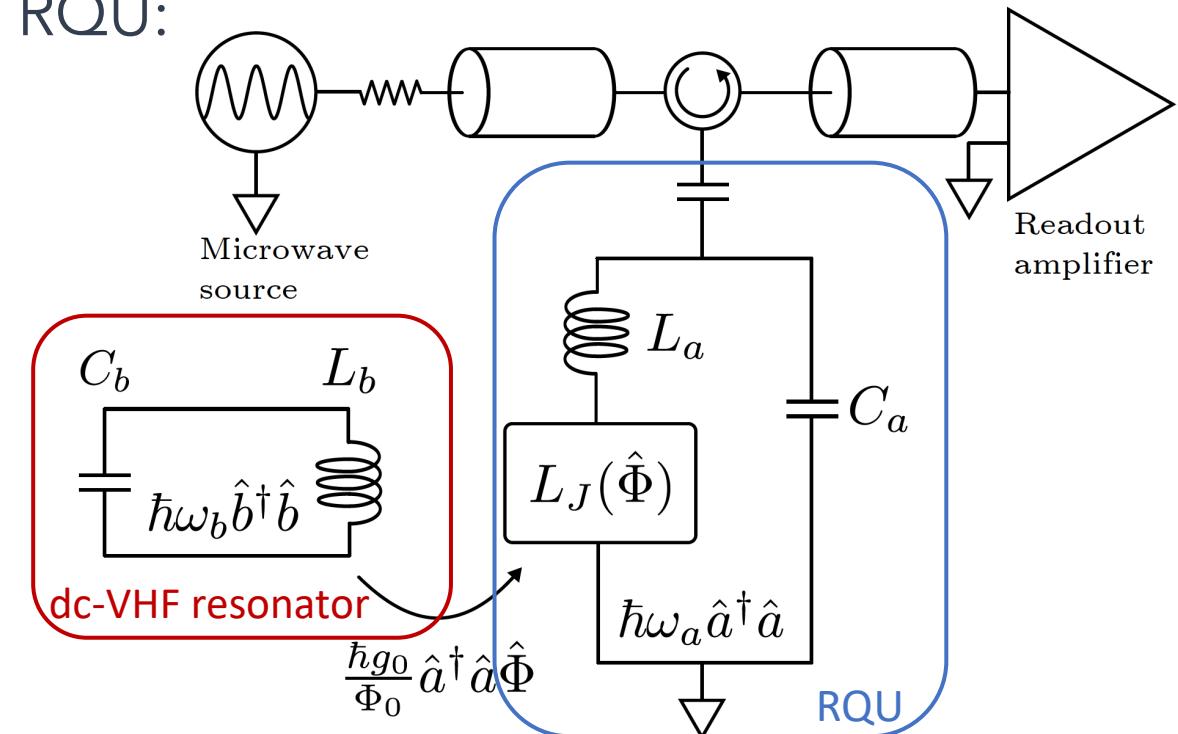
# Radiofrequency Quantum Upconverters

LIGO:



$$\omega_B = \sqrt{\frac{k}{m}} \quad \omega_A = \frac{2\pi qc}{l(x)}$$

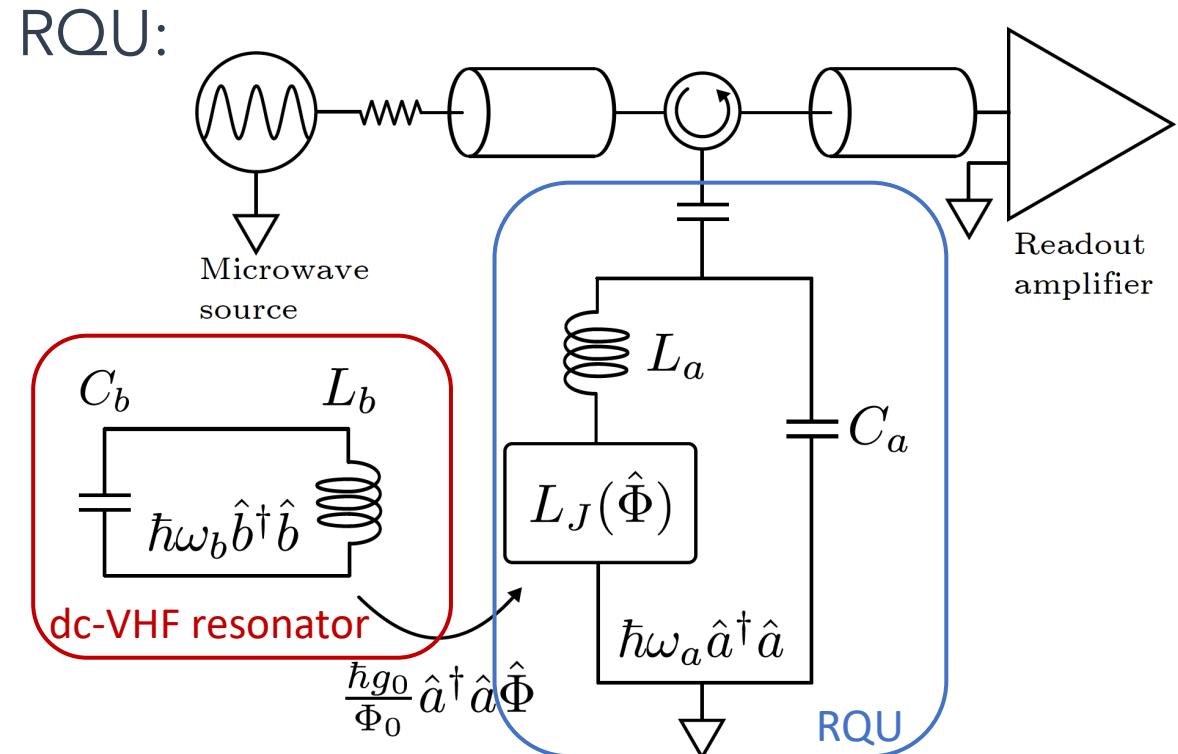
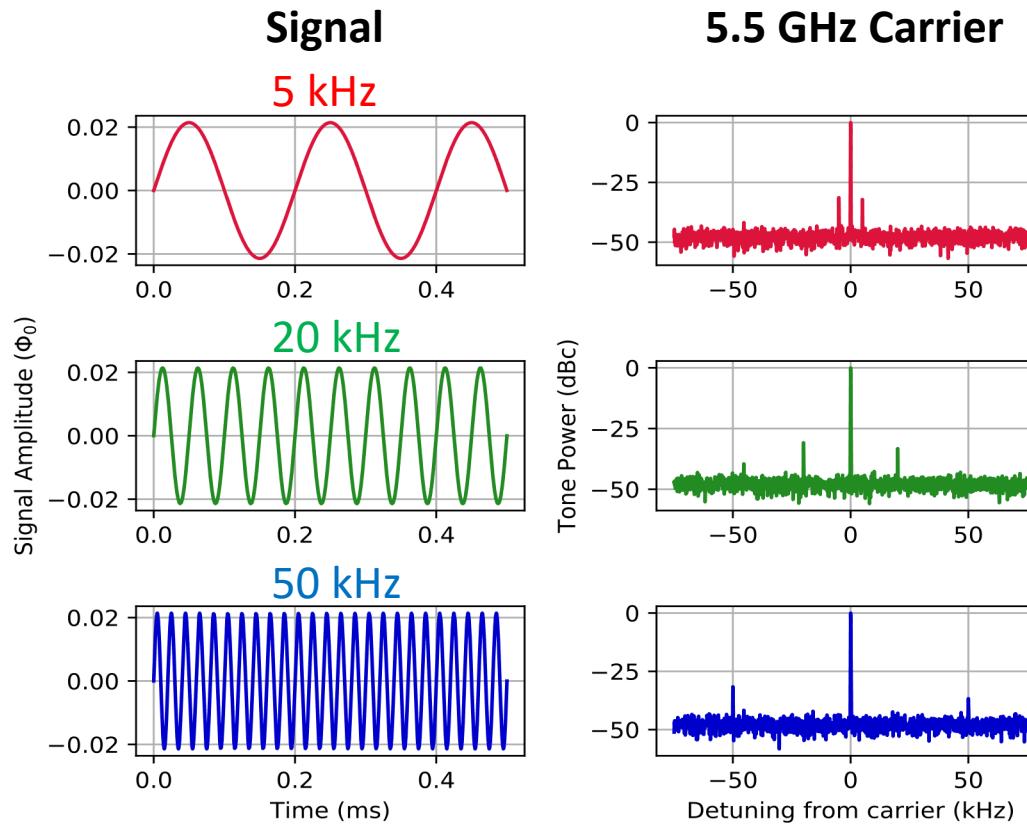
RQU:



Identical Hamiltonian Formalism:

$$H = \hbar\omega_b \left( b^\dagger b + \frac{1}{2} \right) + \hbar\omega_a \left( a^\dagger a + \frac{1}{2} \right) + H_{\text{int}} \quad H_{\text{int}} = -\frac{\hbar}{2} F b^\dagger b (a^\dagger + a)$$

# Radiofrequency Quantum Upconverters

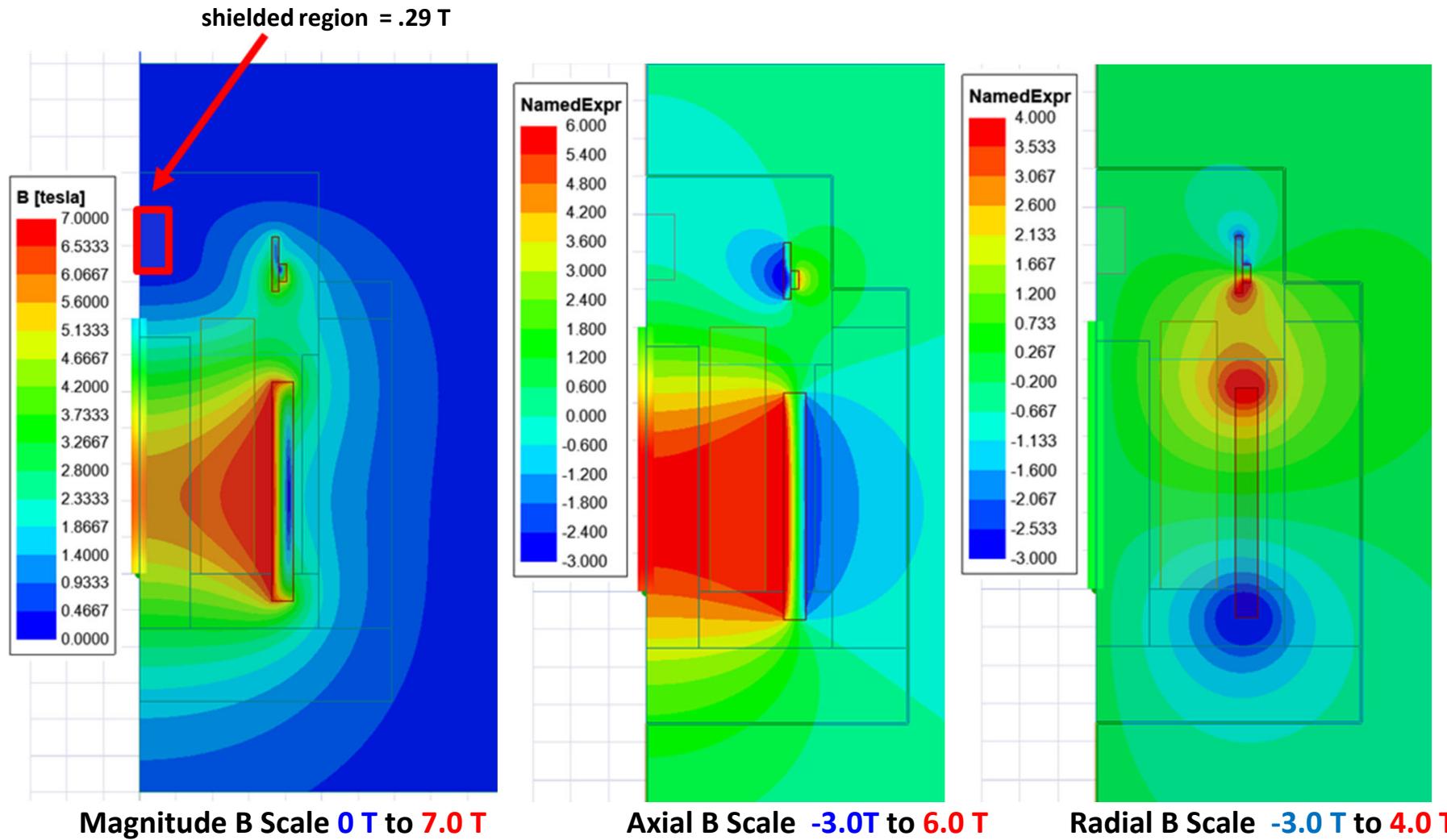


Identical Hamiltonian Formalism:

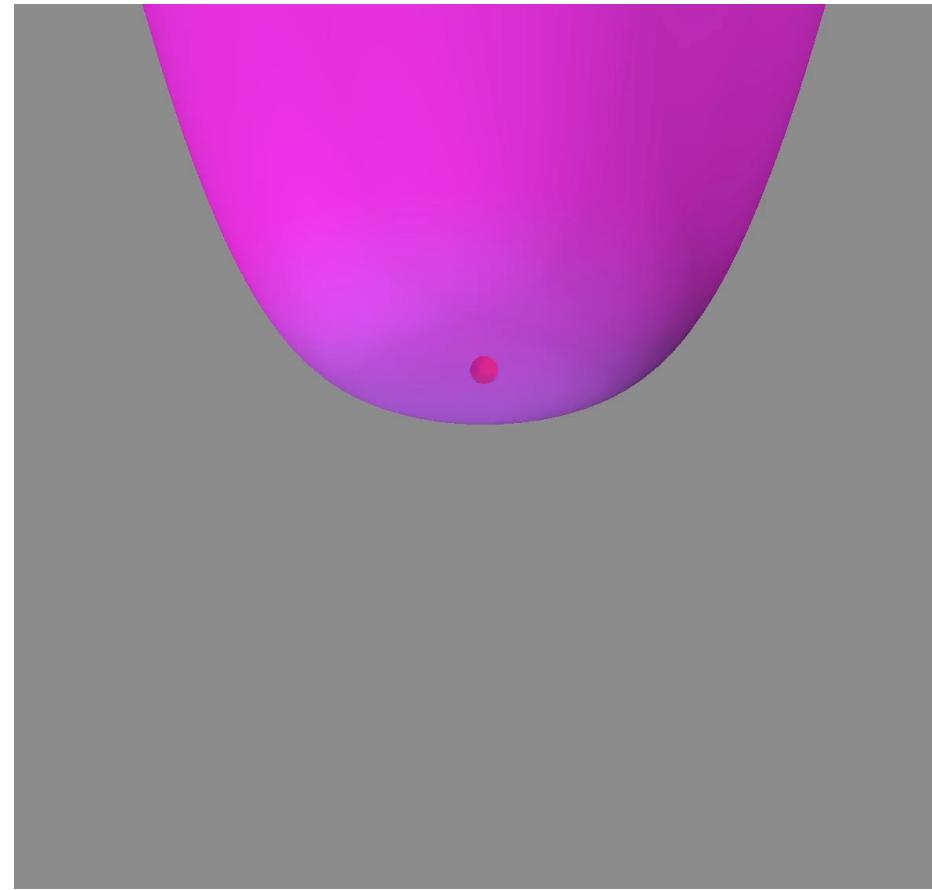
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$$H_{\text{int}} = -\frac{\hbar}{2} F b^\dagger b (a^\dagger + a)$$

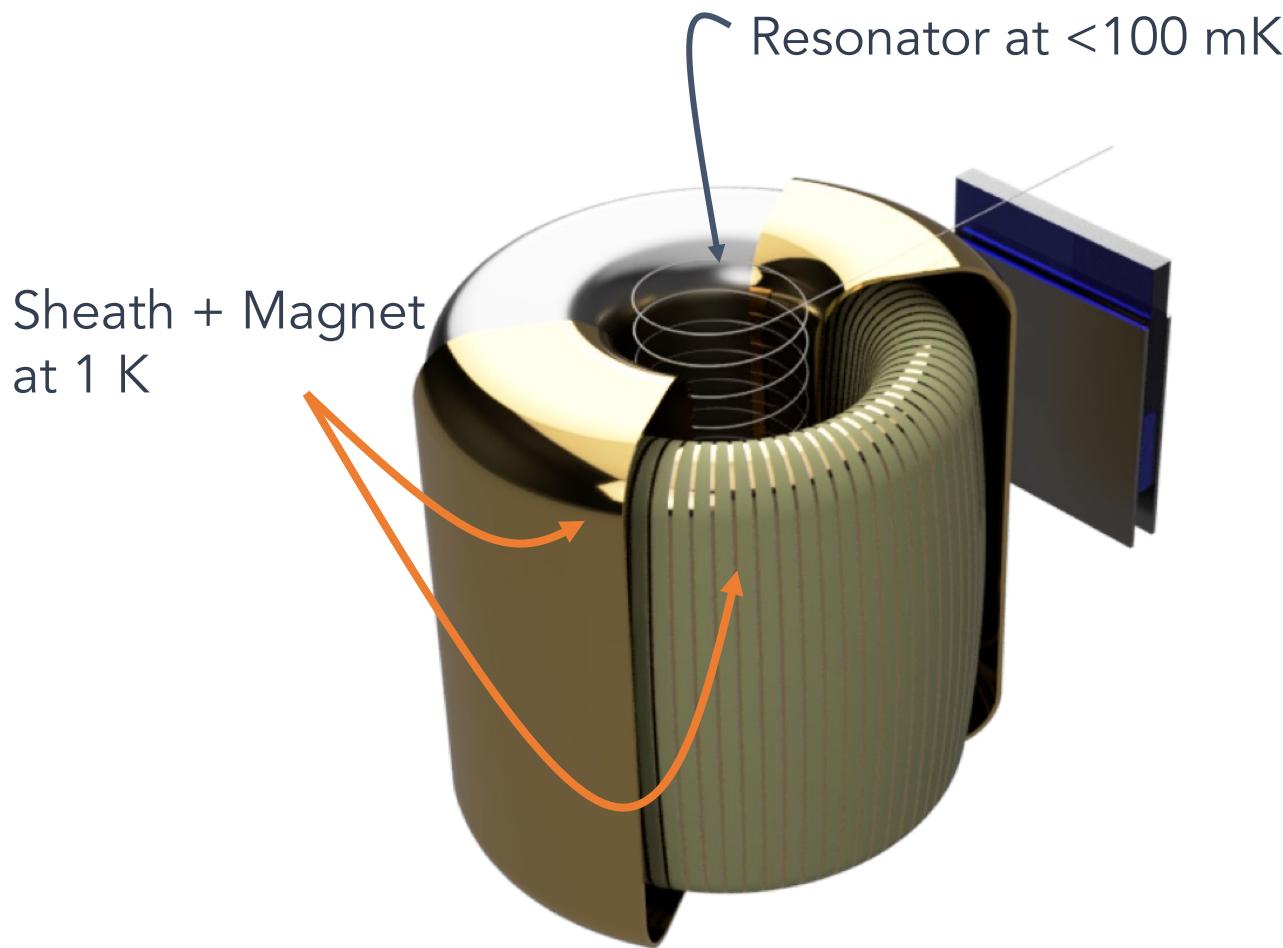
# $m^3$ magnet



# Misalignment mechanism



# 50L details



~Coupled energy in resonator vs  
energy in axion field

$$\frac{1}{g_{a\gamma\gamma,\min}} = \frac{c_{\text{PU}} B_0 V^{5/6} Q^{1/4} \tau_{\text{int}}}{T^{1/4} \eta^{1/4}}$$

Integration time

Amplifier noise temperature,  
normalized to  $\frac{1}{2}$  photon SQL

Physical Temperature