

Inclusive $|V_{ub}|$ from SIMBA

Zoltan Ligeti

[On behalf of F. Bernlochner, H. Lacker, ZL, I. Stewart, F. Tackmann, K. Tackmann]

- Model independent shape function treatment
- Fits for $B \rightarrow X_s \gamma$
- Fits for $B \rightarrow X_u \ell \bar{\nu}$
- Future progress
- Conclusions



V_{ub} — the beginning

CLEO, PRL **64** (1990) 16, Received 8 Nov 1989, (212+101)/pb

ARGUS, PLB **234** (1990) 409, Received 28 Nov 1989, (201+69)/pb

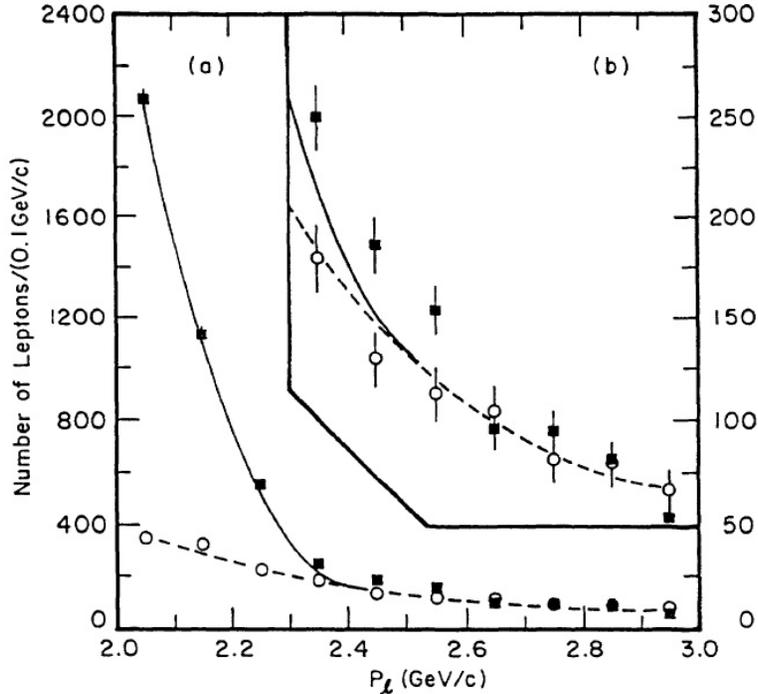


FIG. 1. Sum of the e and μ momentum spectra for ON data (filled squares), scaled OFF data (open circles), the fit to the OFF data (dashed line), and the fit to the OFF data plus the $b \rightarrow cl\nu$ yield (solid line). Note the different vertical scales in (a) and (b).

“ $|V_{ub}/V_{cb}|$... is approximately 0.1; it is sensitive to the theoretical model.”

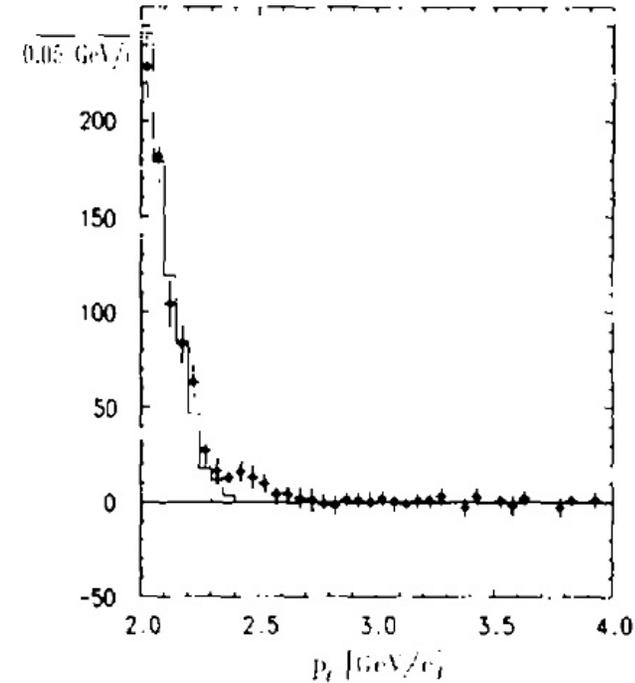


Fig. 5. Combined lepton momentum spectrum for direct $\Upsilon(4S)$ decays: the histogram is a $b \rightarrow c$ contribution normalized in the region 2.0–2.3 GeV/c.

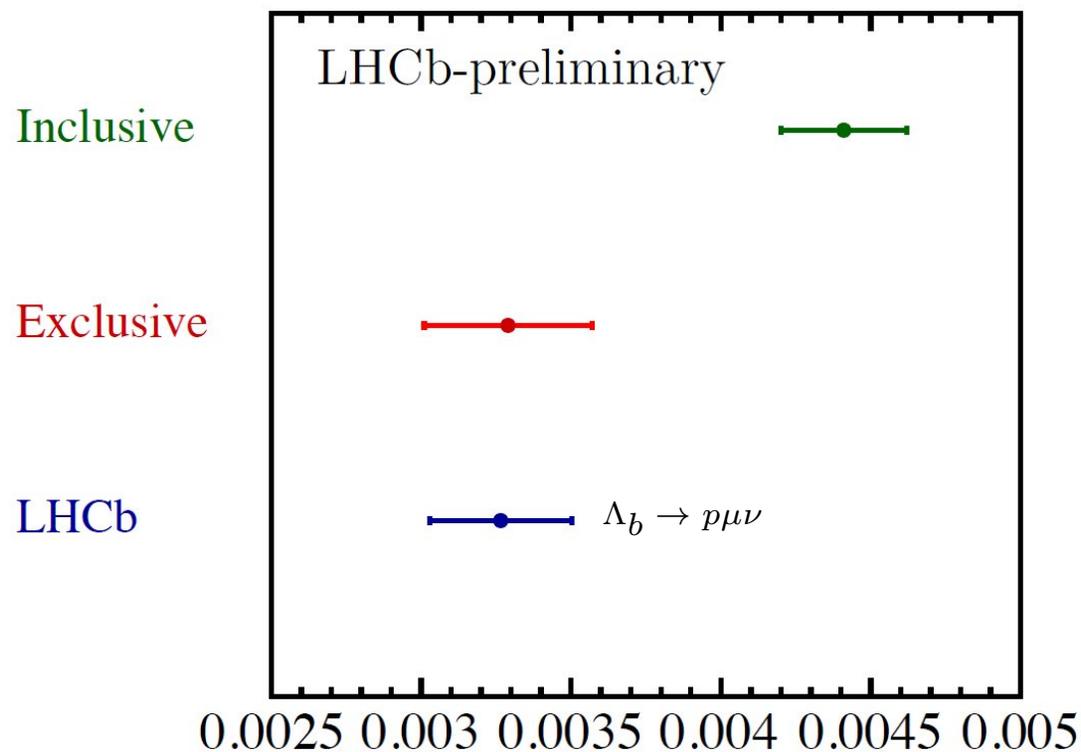
“If interpreted as a signal of $b \rightarrow u$ coupling ... $|V_{ub}/V_{cb}|$ of about 10%.”

25 years later — situation still confusing

- By now, 5000 times more data, persistent tensions — I think the jury is still out:

HFAG:

$ V_{ub} _{\text{incl-BLL}} = (4.62 \pm 0.35) \times 10^{-3}$	$ V_{ub} _{\pi\ell\bar{\nu}\text{-LQCD}} = (3.4 \pm 0.4) \times 10^{-3}$
$ V_{ub} _{\text{incl-BLNP}} = (4.45 \pm 0.27) \times 10^{-3}$	$ V_{ub} _{\tau\nu} = ?$
$ V_{ub} _{\text{incl-GGOU}} = (4.51 \pm 0.22) \times 10^{-3}$	SM fit: $(3.6 \pm 0.2) \times 10^{-3}$

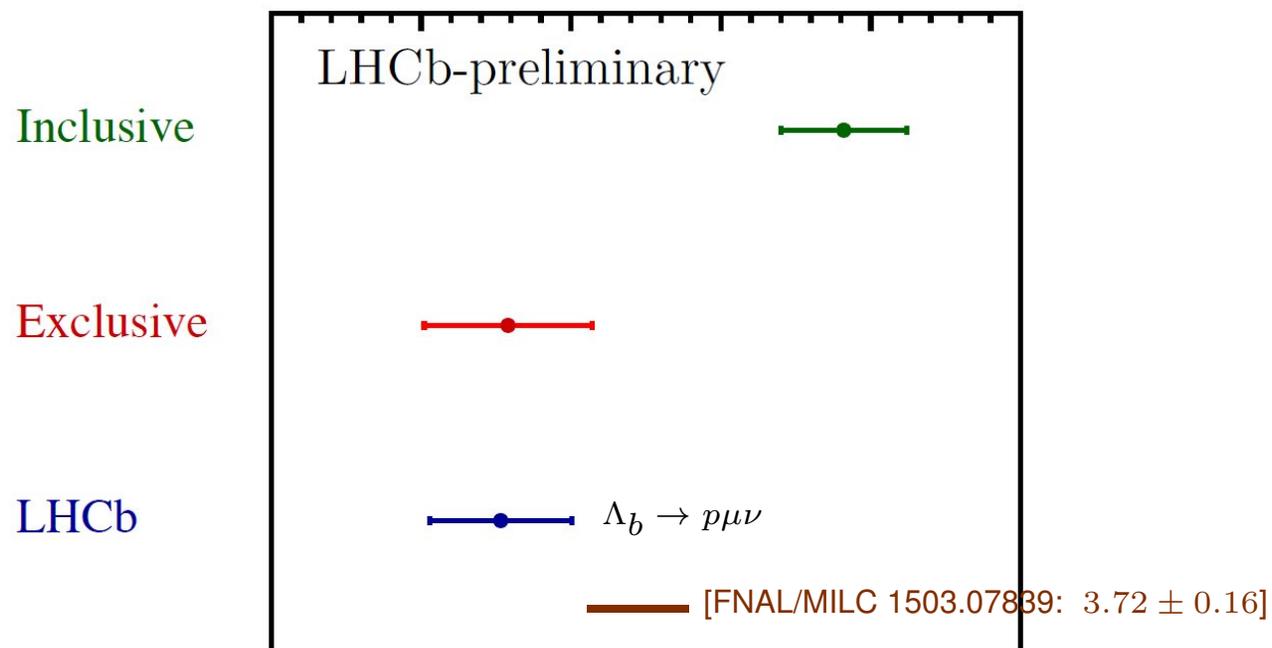


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- What would it take to conclude that there is unambiguous evidence for NP?

Tensions in $|V_{ub}|$ determinations

- $\sim 3\sigma$ tension among $|V_{ub}|$ measurements

Tim Gershon @ FPCP 2014: “Understanding this will involve a great deal of effort, but is essential for continued progress in the field”

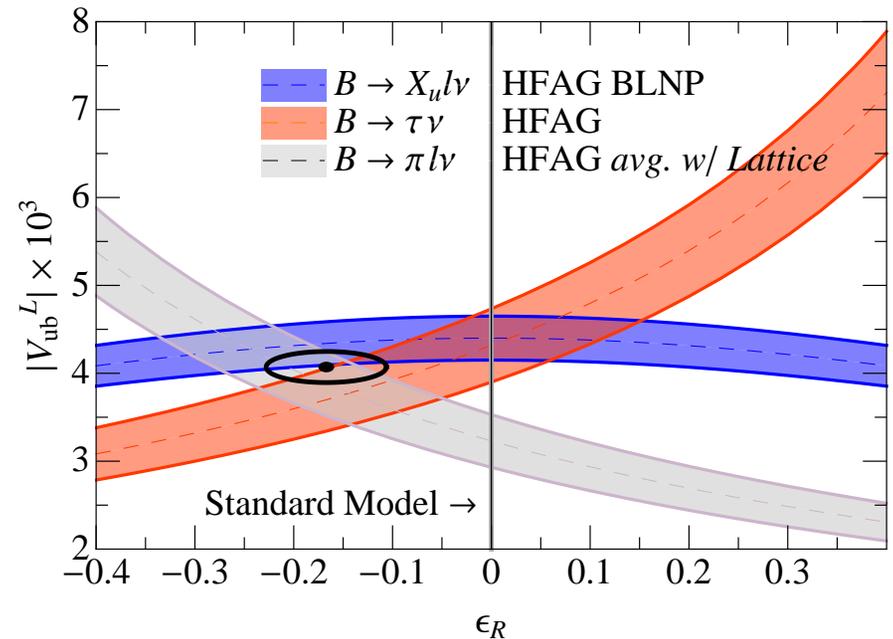
- Too early to conclude:
 - Inclusive determination can improve
 - Exclusive measured better with full reco
 - Lattice QCD results will improve

- A BSM possibility:

$$\mathcal{L} = -\frac{4G_F}{\sqrt{2}} V_{ub}^L (\bar{u}\gamma_\mu P_L b + \epsilon_R \bar{u}\gamma_\mu P_R b) (\bar{\nu}_\ell \gamma^\mu P_L \ell)$$

Can we construct observables which give “more vertical” constraints?

- NB: Cleanest $|V_{ub}|$ I know, only isospin, $\mathcal{B}(B_u \rightarrow \ell \bar{\nu}) / \mathcal{B}(B_d \rightarrow \mu^+ \mu^-)$ [Grinstein, CKM'06 (?)]



[Bernlochner, ZL, Turczyk, 1408.2516]

Decay	$ V_{ub} \times 10^4$	adm.
$B \rightarrow \pi \ell \bar{\nu}_\ell$	3.23 ± 0.30	$(1 + \epsilon_R)$
$B \rightarrow X_u \ell \bar{\nu}_\ell$	4.39 ± 0.21	$\sqrt{1 + \epsilon_R^2}$
$B \rightarrow \tau \bar{\nu}_\tau$	4.32 ± 0.42	$(1 - \epsilon_R)$

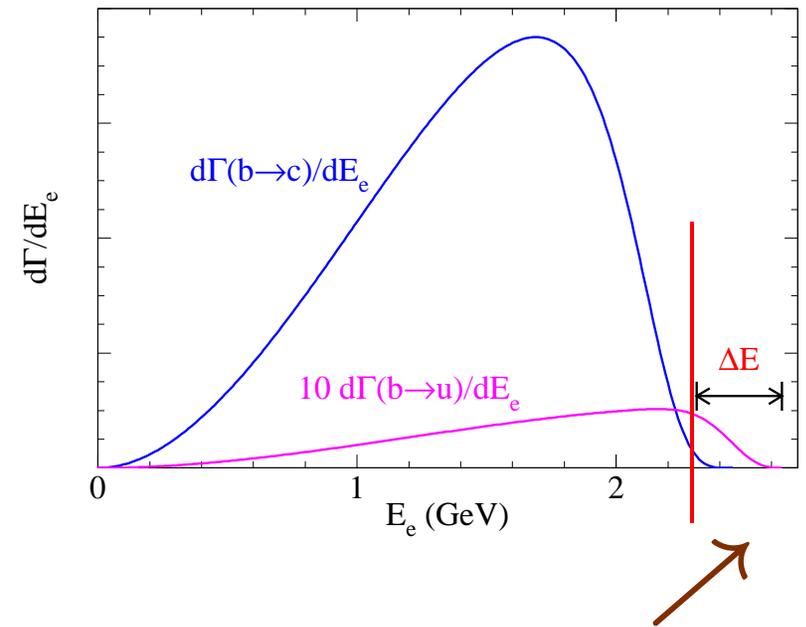
Features of SIMBA

- Optimally combine all information on $B \rightarrow X_u \ell \bar{\nu}$ & $B \rightarrow X_s \gamma$
Consistently treat uncertainties and their correlations (exp, theo, parameters)
- Simultaneously determine:
 - Overall normalization: $\mathcal{B}(B \rightarrow X_s \gamma)$, $|V_{ub}|$
 - Parameters: m_b , shape function(s)
- Utilize all measurements:
 - Different $B \rightarrow X_s \gamma$ spectra, or partial rates
 - Different $B \rightarrow X_u \ell \bar{\nu}$ spectra, or partial rates
 - Include other constraints on m_b , λ_1 , etc.
 - Eventually use or predict $B \rightarrow X_s \ell^+ \ell^-$
- Same strategy as for inclusive $|V_{cb}|$, just a lot more complicated...

Shape function

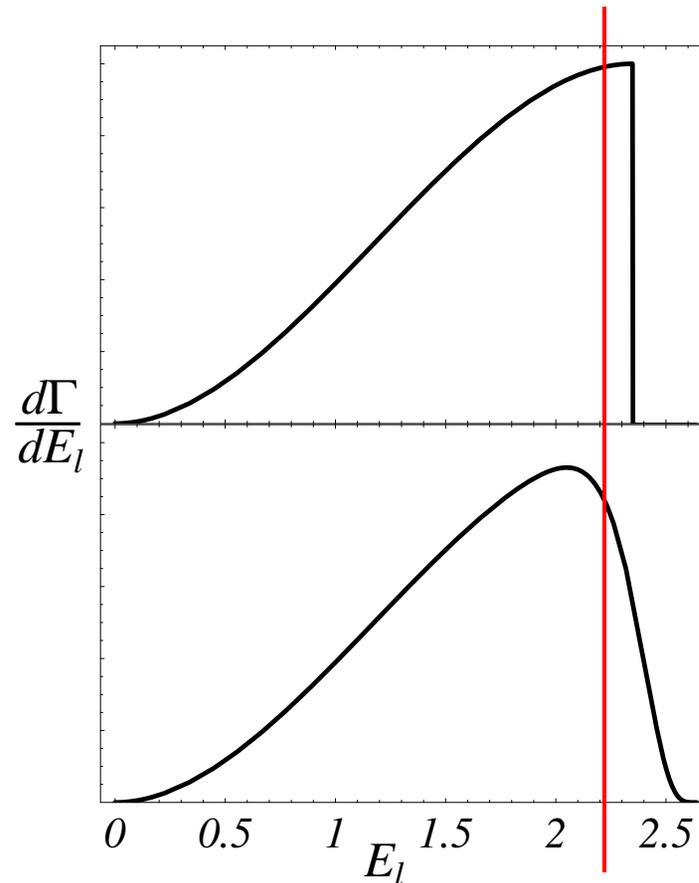
The challenge of inclusive $|V_{ub}|$ measurements

- Total rate calculable with $\sim 4\%$ uncertainty, similar to $\mathcal{B}(B \rightarrow X_c \ell \bar{\nu})$
- To remove the huge charm background ($|V_{cb}/V_{ub}|^2 \sim 100$), need phase space cuts
Phase space cuts can enhance perturbative and nonperturbative corrections drastically
- Hadronic parameters are functions (like PDFs)
Leading order: universal & related to $B \rightarrow X_s \gamma$;
 $\mathcal{O}(\Lambda_{\text{QCD}}/m_b)$: several new unknown functions
- Nonperturbative effects shift endpoint $\frac{1}{2} m_b \rightarrow \frac{1}{2} m_B$ & determine its shape
- Shape in the endpoint region is determined by b quark PDF in B [“shape function”]
Related to $B \rightarrow X_s \gamma$ photon spectrum at lowest order [Bigi, Shifman, Uraltsev, Vainshtein; Neubert]



Shape function: lepton endpoint vs. $B \rightarrow X_s \gamma$

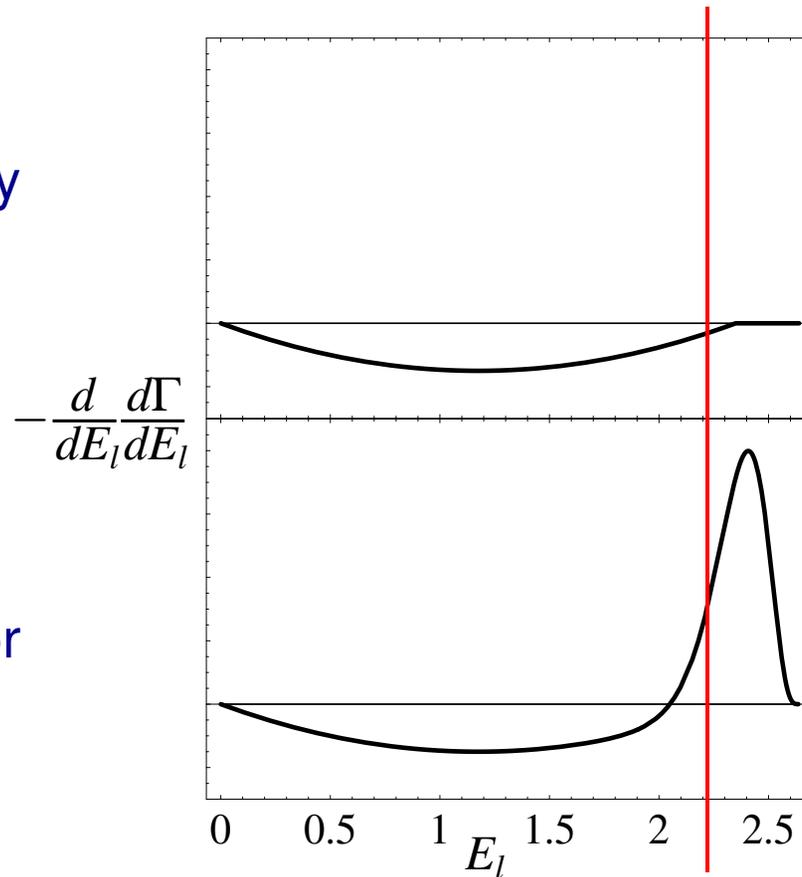
b quark decay
spectrum



with a model for
 b quark PDF

Shape function: lepton endpoint vs. $B \rightarrow X_s \gamma$

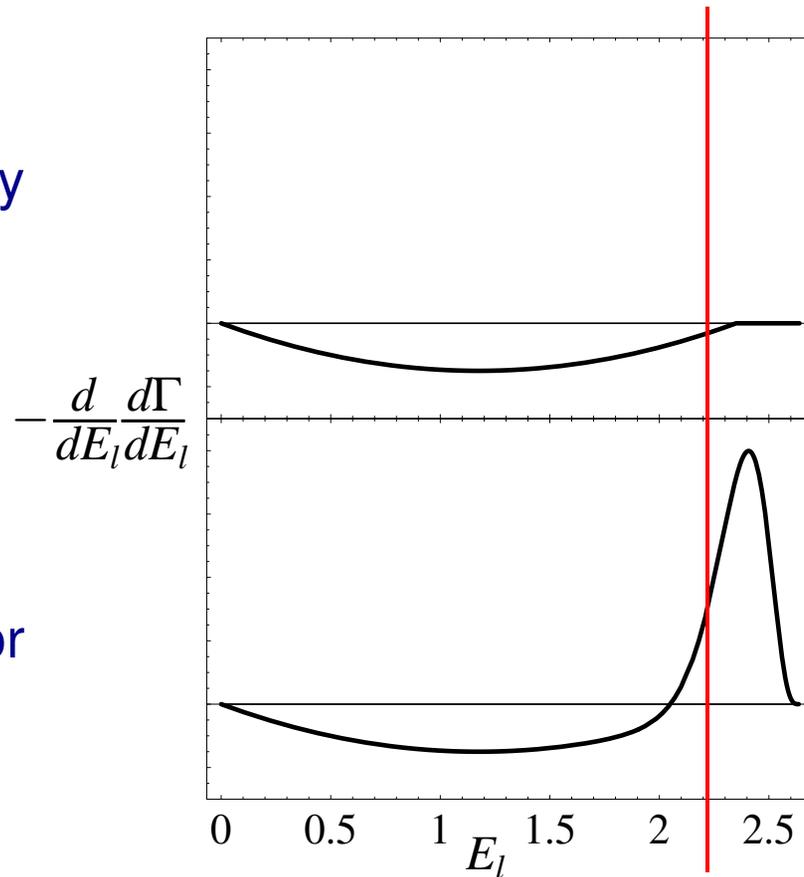
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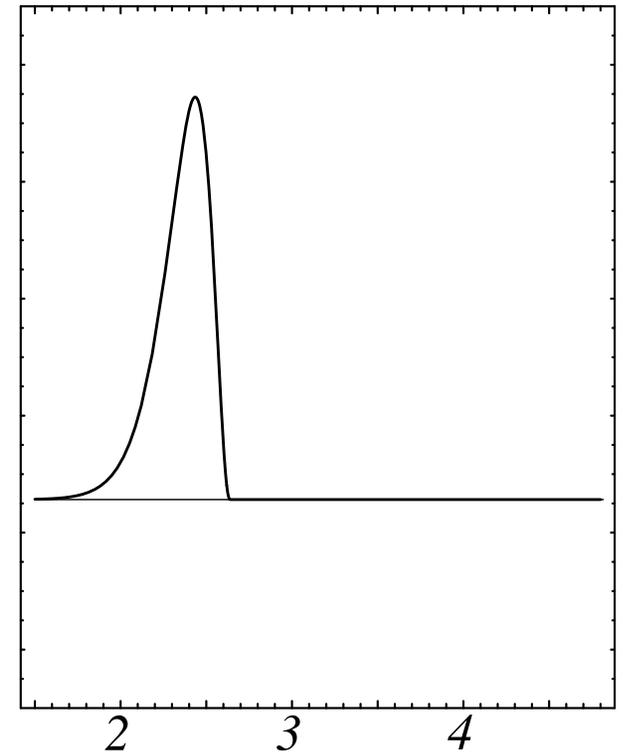
Shape function: lepton endpoint vs. $B \rightarrow X_s \gamma$

b quark decay spectrum



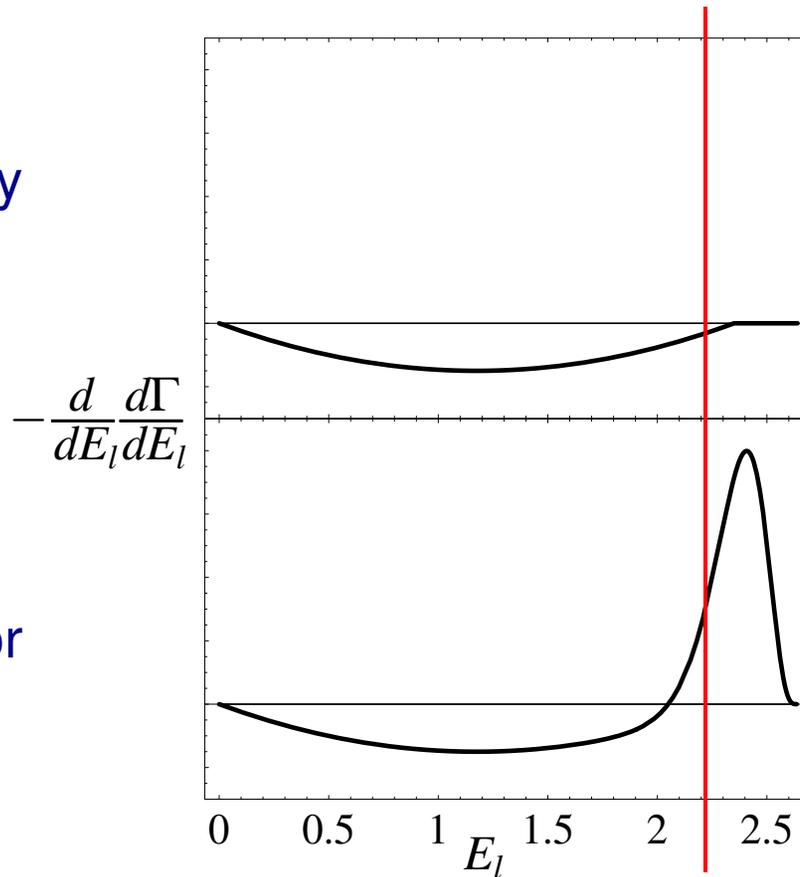
with a model for b quark PDF

difference:



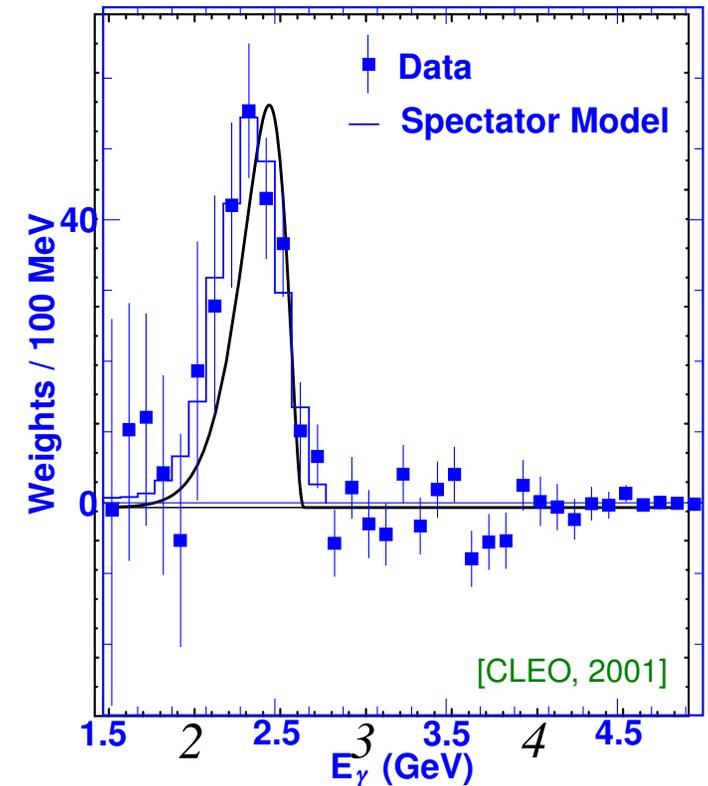
Shape function: lepton endpoint vs. $B \rightarrow X_s \gamma$

b quark decay spectrum



with a model for b quark PDF

difference:



- Both spectra determined at lowest order by the b quark PDF in B meson

Start with $B \rightarrow X_s \gamma$

Regions of $B \rightarrow X_s \gamma$ photon spectrum

- Important both for $|V_{ub}|$ and constraining NP

- Peak around $E_\gamma \sim 2.3 \text{ GeV}$ ($m_B - 2E_\gamma \sim 0.8 \text{ GeV}$)

Three cases: 1) $\Lambda_{\text{QCD}} \sim m_B - 2E_\gamma \ll m_B$ ["SCET"]

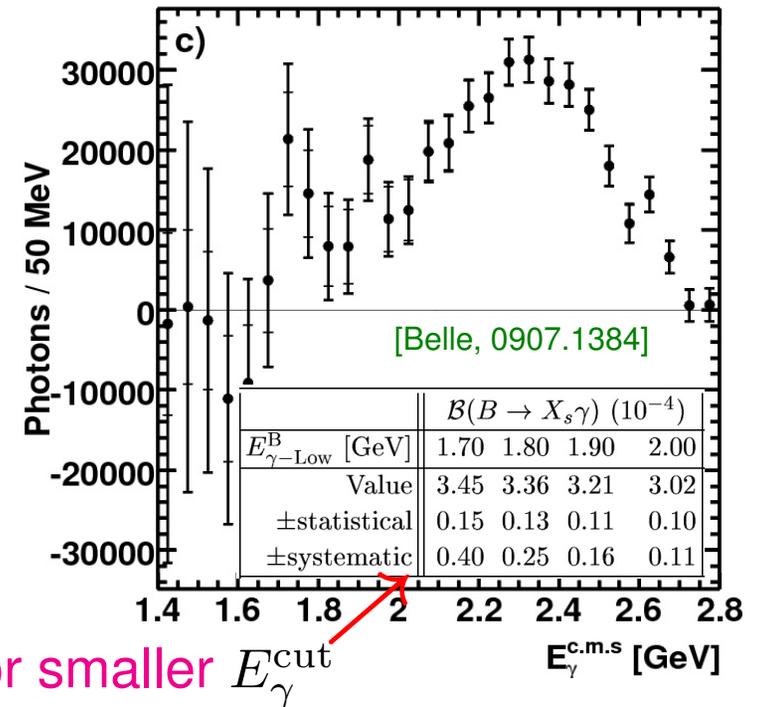
2) $\Lambda_{\text{QCD}} \ll m_B - 2E_\gamma \ll m_B$ ["MSOPE"]

3) $\Lambda_{\text{QCD}} \ll m_B - 2E_\gamma \sim m_B$

Expansions and theory uncertainties differ in the 3 regions

Neither 1) nor 2) is fully appropriate

- Experimental systematic error rapidly increases for smaller E_γ^{cut}



- Current practice: Compare rate extrapolated to 1.6 GeV with theoretical prediction

Con: (i) extrapolation uses theory, so comparison of theory and data is effectively done at the measured values; (ii) best use of the most precise measurements?

The shape function (b quark PDF in B)

- The shape function $S(\omega, \mu)$ contains nonperturbative physics and obeys a RGE

Even if $S(\omega, \mu_\Lambda)$ has exponentially small tail, RGE running gives long tail and divergent moments

$$S(\omega, \mu_i) = \int d\omega' U_S(\omega - \omega', \mu_i, \mu_\Lambda) S(\omega', \mu_\Lambda)$$

[Balzereit, Mannel, Kilian]

Constraint: moments (OPE) + $B \rightarrow X_s \gamma$ shape

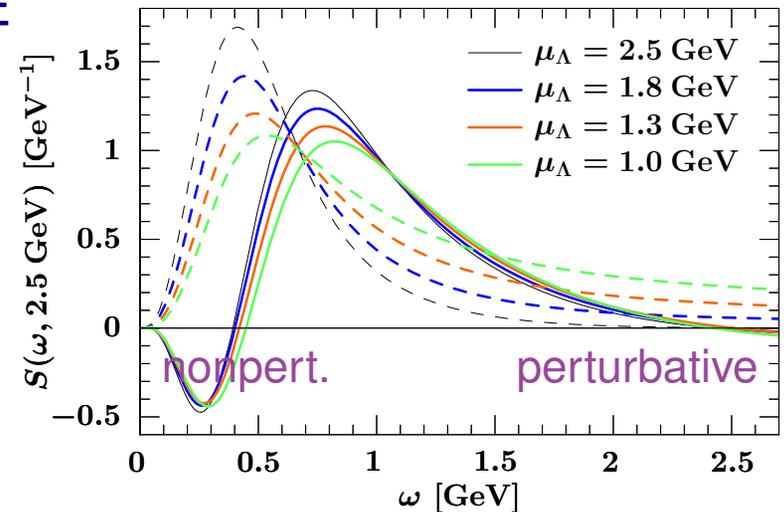
- Derive: $S(\omega, \mu_\Lambda) = \int dk C_0(\omega - k, \mu_\Lambda) F(k)$
- [ZL, Stewart, Tackmann, 0807.1926]

– Can use any (mass) scheme, work to any order

– Stable results for varying μ_Λ (SF modeling scale, part of uncertainty, often ignored)

– Similar to how all matrix elements are defined [e.g., $B_K(\mu) = \hat{B}_K \times [\alpha_s(\mu)]^{2/9}(1 + \dots)$]

- Consistent to impose moment constraints on $F(k)$, but not on $S(\omega, \mu_\Lambda)$ w/o cutoff



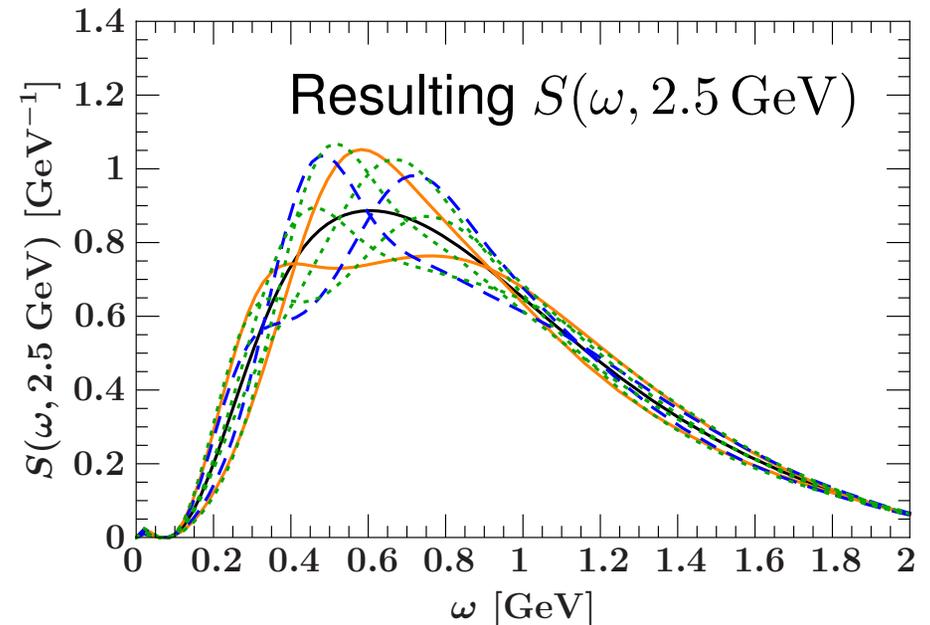
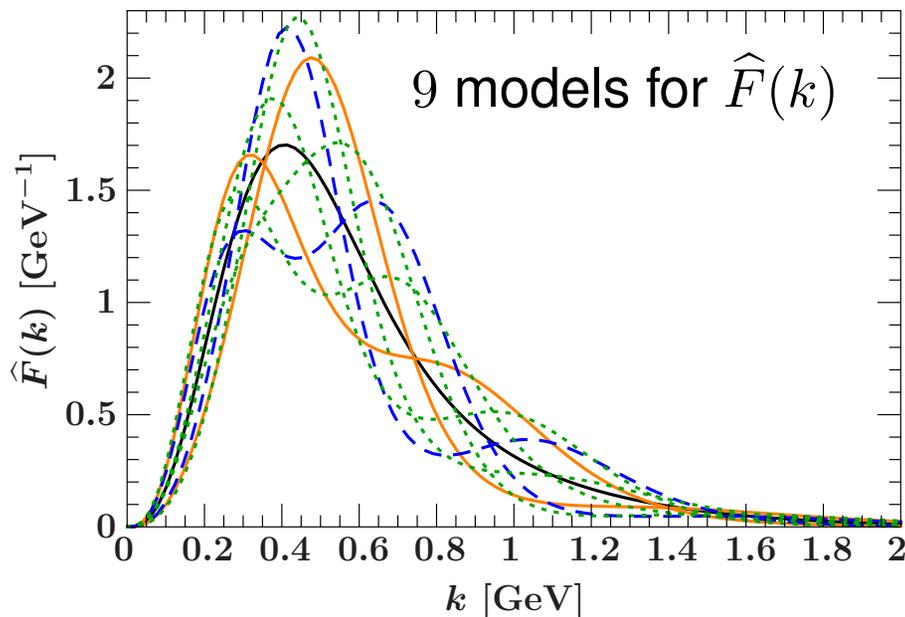
Model $\begin{cases} S & \text{(dash)} \\ F & \text{(solid)} \end{cases}$ run to 2.5 GeV

Shape function: the bottom line

$$S(\omega, \mu_\Lambda) = \int dk \hat{F}(k) \hat{C}_0(\omega - k, \mu_\Lambda)$$

\hat{F} : nonperturbative
 determines peak region
 well-defined moments
 fit from data

\hat{C}_0 : perturbative
 generates tail consistent with RGE
 divergent moments
 calculable



Designer orthonormal functions

- Devise suitable orthonormal basis functions (avoid: fit parameters of model functions to data)

$$\widehat{F}(\lambda x) = \frac{1}{\lambda} \left[\sum c_n f_n(x) \right]^2, \quad n \text{th moment} \sim \Lambda_{\text{QCD}}^n$$

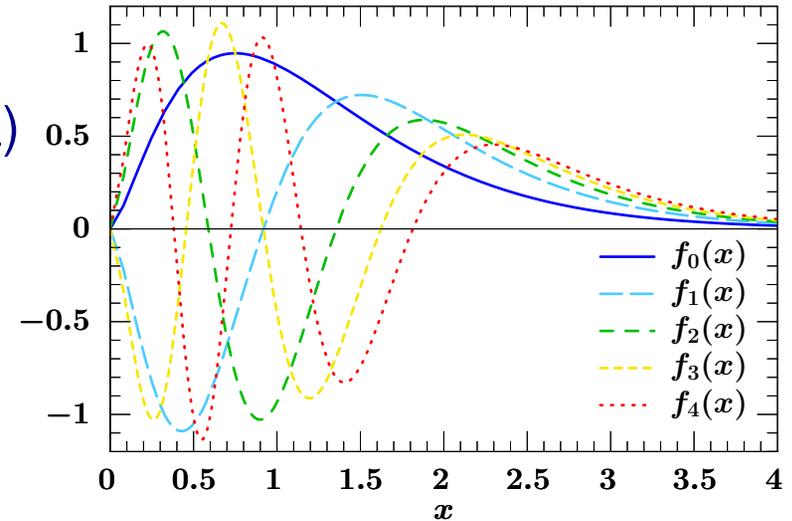
$$f_n(x) \sim P_n[y(x)] \quad \leftarrow \text{Legendre polynomials}$$

- Approximating a model shape function

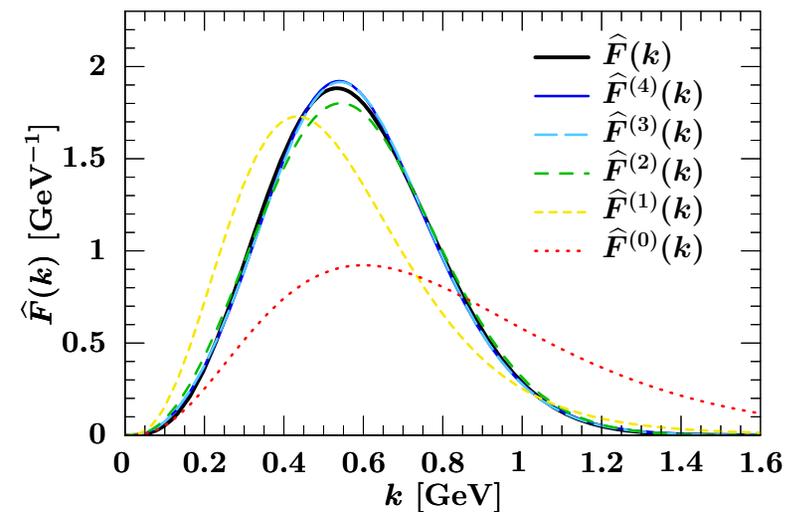
Better to add a new term in an orthonormal basis than a new parameter to a model:

- less parameter correlations
- uncertainties easier to quantify

“With four parameters I can fit an elephant, and with five I can make him wiggle his trunk.” (John von Neumann)



[ZL, Stewart, Tackmann, 0807.1926]



Details of fitting the data

- $\widehat{F}(k)$ enters the spectra linearly
 \Rightarrow can calculate independently the contribution of $f_m f_n$ in the expansion of $\widehat{F}(k)$:

$$d\Gamma = \sum \underbrace{c_m c_n}_{\text{fit}} \underbrace{d\Gamma_{mn}}_{\text{compute}}$$

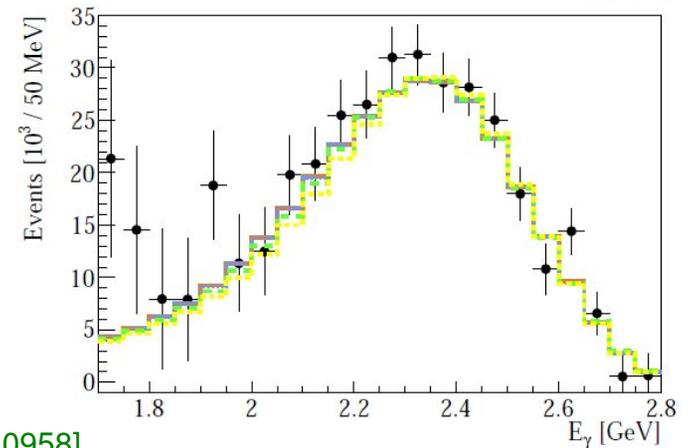
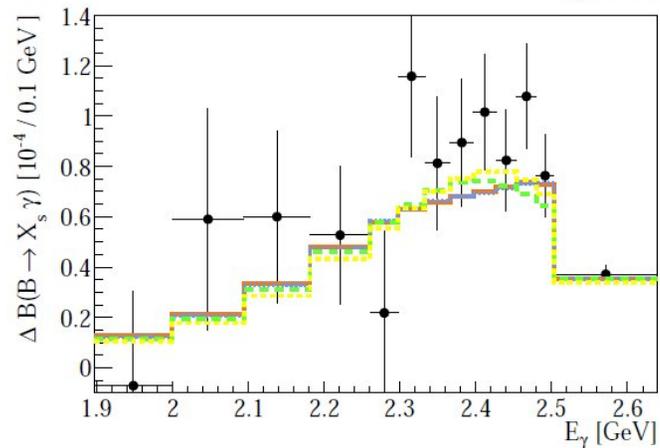
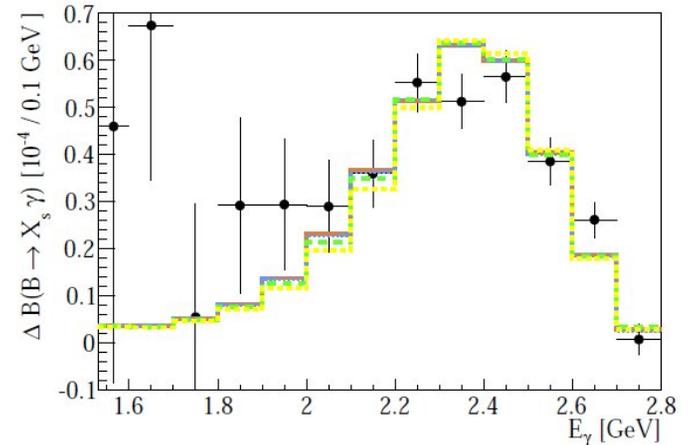
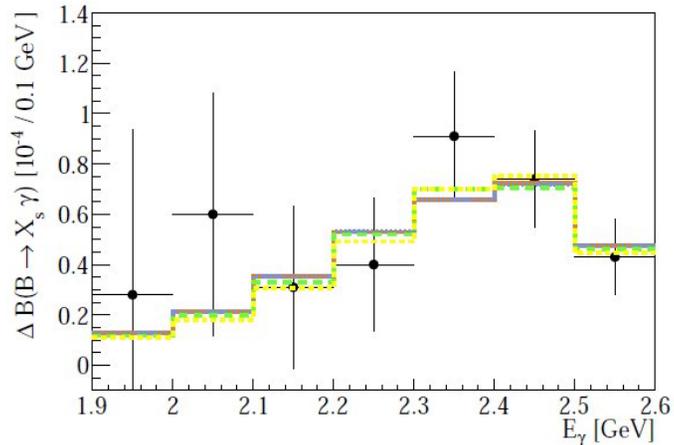
$$d\Gamma_{mn} = \Gamma_0 H(p_X^\pm) \int_0^{p_X^\pm} dk \frac{\widehat{P}(p^-, k)}{\lambda} \underbrace{f_m\left(\frac{p_X^+ - k}{\lambda}\right) f_n\left(\frac{p_X^+ - k}{\lambda}\right)}_{\text{basis functions}}$$

Fit the c_i coefficients from all measured (binned) spectra (similar to $|V_{cb}|$ fit)

- SIMBA includes:
 - Simultaneous fit using all available information
 - Correlations in data, propagation of SF uncertainties
 - Validate the fits with pseudo-experiments
 - Check model independence by varying number of basis functions in fit (up to 5)



Fit results for $B \rightarrow X_s \gamma$



[SIMBA, 1303.0958]

- Fit with $\lambda = 0.5$ GeV with 2 (yellow), 3 (green), 4 (blue), and 5 (orange) coefficients

Can change the basis by varying λ — find consistent results

Comments on uncertainties

- Theoretical inputs:

- Scale variations: μ_i and profiles

- Subleading SF: tree level C_7^2 terms absorbed in C_7^{incl}
estimate uncertainty due to $\mathcal{O}(\alpha_s \Lambda_{\text{QCD}}/m_b)$ shape functions

- Non- C_7^2 subleading SF (4-quark) less important than sometimes claimed,
since $\mathcal{O}(\alpha_s \Lambda_{\text{QCD}}/m_b)$ in peak region, which dominates the fit

- λ_2, ρ_2, m_c — mild sensitivity

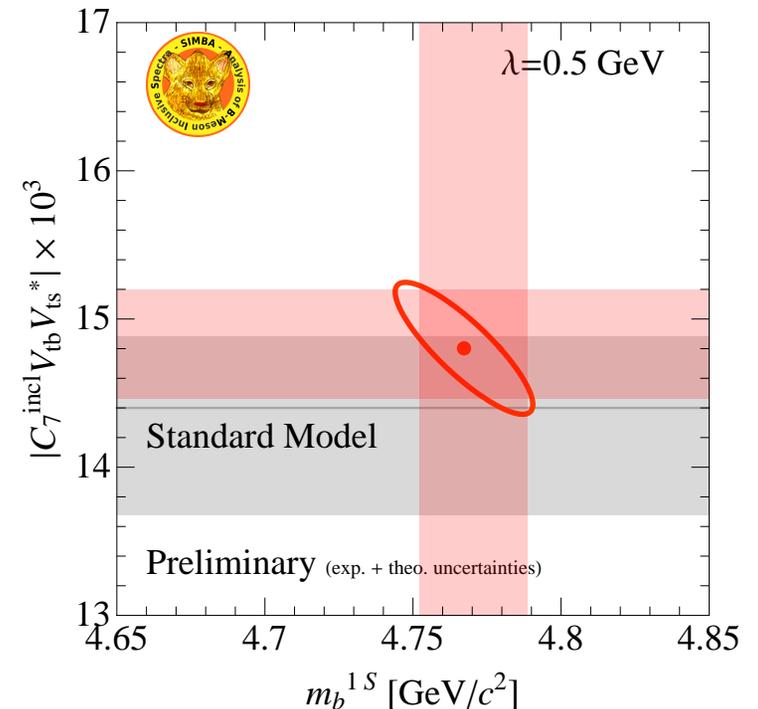
- Fit procedure and validation:

- Measurements with all available correlations

- Shape function basis, number of terms in fit, test with toys

Fit results for $B \rightarrow X_s \gamma$ (2)

- Have complete NNLO + NNLL' (2-loop matching & running, 3-loop cusp) [1303.0958]
- $\chi^2/\text{ndf} = 41.7/48$
- SM prediction: $|C_7^{\text{incl}}|^{\text{SM}} = 0.354^{+0.011}_{-0.012}$
 $|V_{tb}V_{ts}| = (40.4 \pm 0.1) \times 10^{-3}$
- Fit: $|C_7^{\text{incl}} V_{tb}V_{ts}| = (14.83 \pm 0.53_{[\text{exp}]} \pm 0.37_{[\text{th}]}) \times 10^{-3}$
- Data slightly above SM prediction, as in HFAG combination vs. Misiak *et al.*



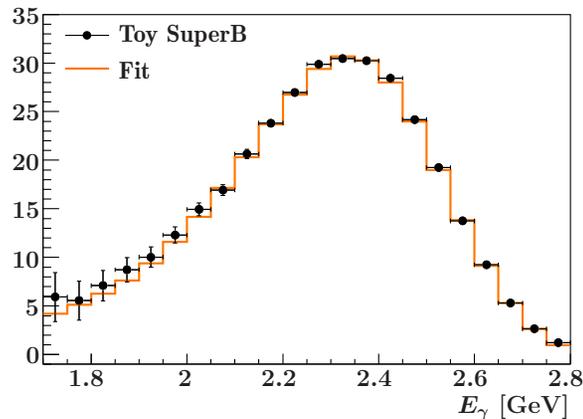
Future of $B \rightarrow X_s \gamma$

- Toy fits few years ago for 75/ab:

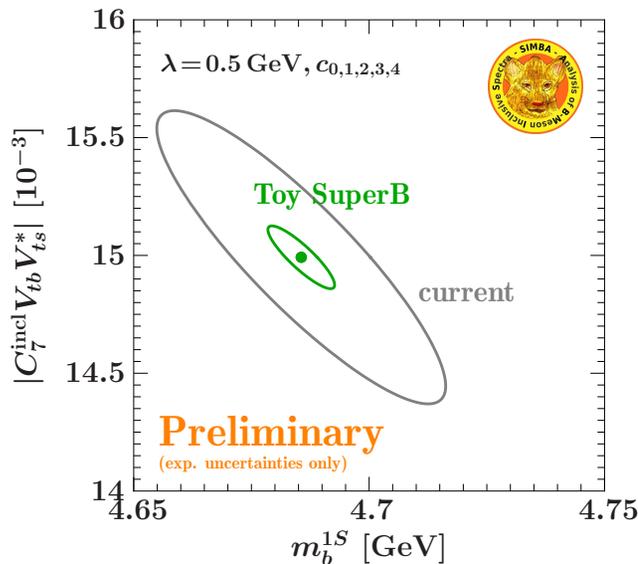
5 coefficients

$\lambda = 0.5 \text{ GeV}$

Theory uncert.
will dominate



[BELLE2-NOTE-0021]	Statistical	Systematic (reducible, irreducible)	Total Exp
$B(B \rightarrow X_s \gamma)$ inclusive (untagged)			
605 fb ⁻¹	4.2	(10.3, 5.3)	12.3
5 ab ⁻¹	1.5	(3.6, 5.3)	6.6
50 ab ⁻¹	0.5	(1.1, 5.3)	5.4
$B(B \rightarrow X_s \gamma)$ inclusive (hadron tagged)			
210 fb ^{-1†}	23.2	(15.7, 4.8)	28.4
5 ab ⁻¹	4.8	(3.2, 4.8)	7.5
50 ab ⁻¹	1.5	(1.0, 4.8)	5.1



We assumed factor of 3 reduction in systematic uncertainty, slightly (but not wastly) optimistic

High precision data can be used to fit with more coefficients and constrain subleading effects

$$B \rightarrow X_u l \bar{\nu}$$

$B \rightarrow X_u \ell \bar{\nu}$ is more complicated

- “Natural” kinematic variables: $p_X^\pm = E_X \mp |\vec{p}_X|$ (ratio is “jettiness” of hadrons)

$B \rightarrow X_s \gamma$: $p_X^+ = m_B - 2E_\gamma$ & $p_X^- \equiv m_B$ — independent variables in $B \rightarrow X_u \ell \bar{\nu}$

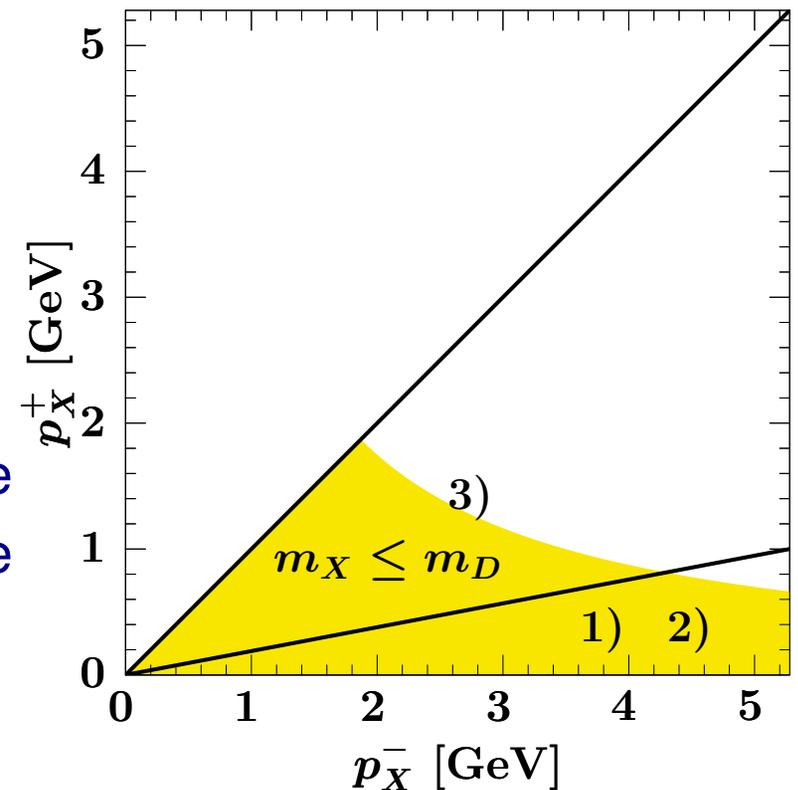
- Three cases:

1) $\Lambda \sim p_X^+ \ll p_X^-$	}	Shape Fn region
2) $\Lambda \ll p_X^+ \ll p_X^-$		
3) $\Lambda \ll p_X^+ \sim p_X^-$		

Want to make no assumptions how p_X^- compares to m_B

- $B \rightarrow X_u \ell \bar{\nu}$: 3-body final state, appreciable rate in region 3), where hadronic final state not jet-like

E.g., $m_X^2 < m_D^2$ does not imply $p_X^+ \ll p_X^-$



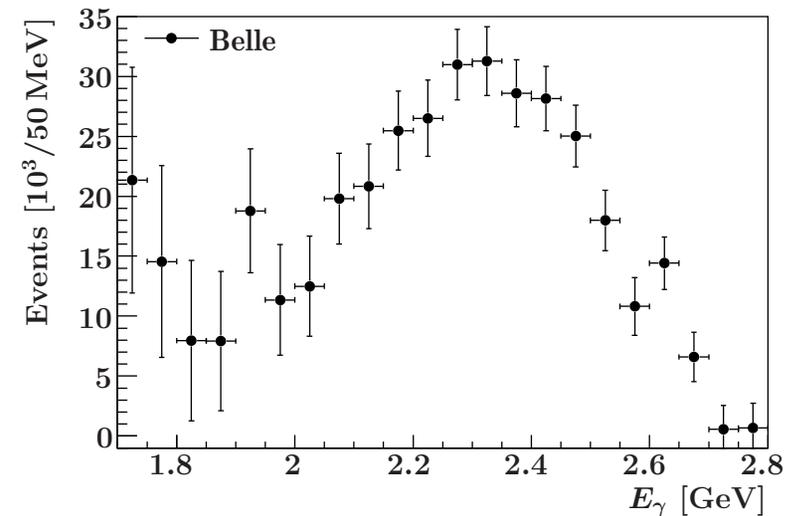
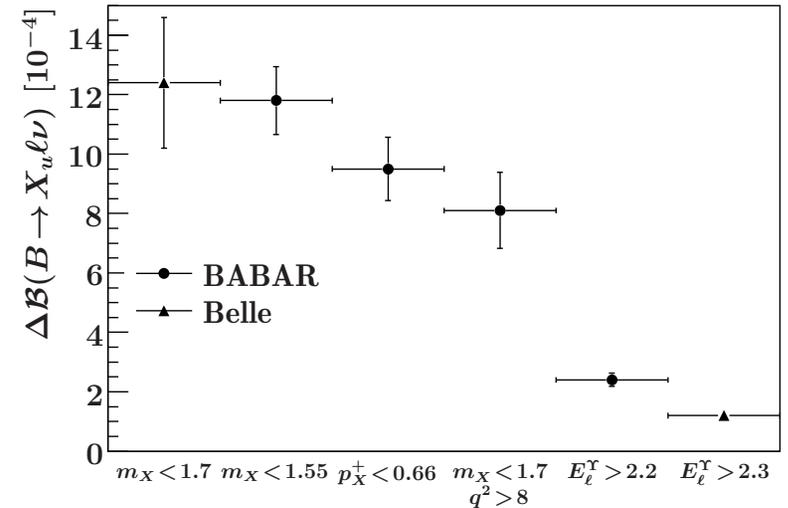
- Existing results based on theory in one region, extrapolated / modeled to rest

Other approaches

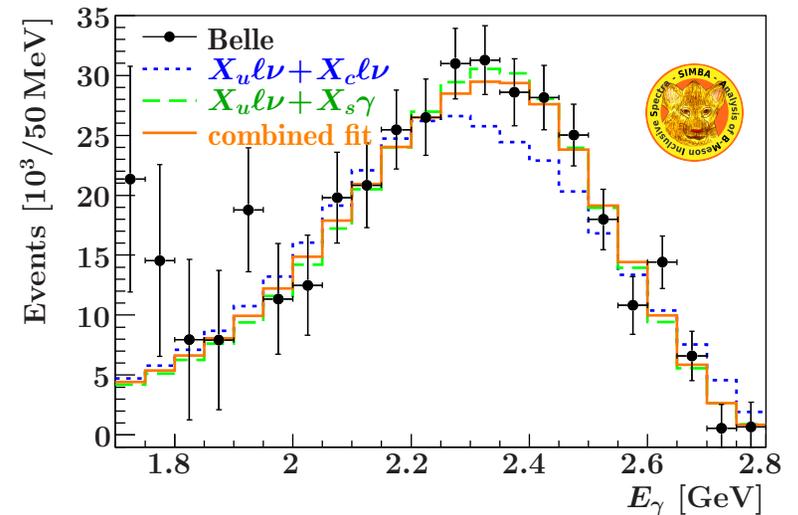
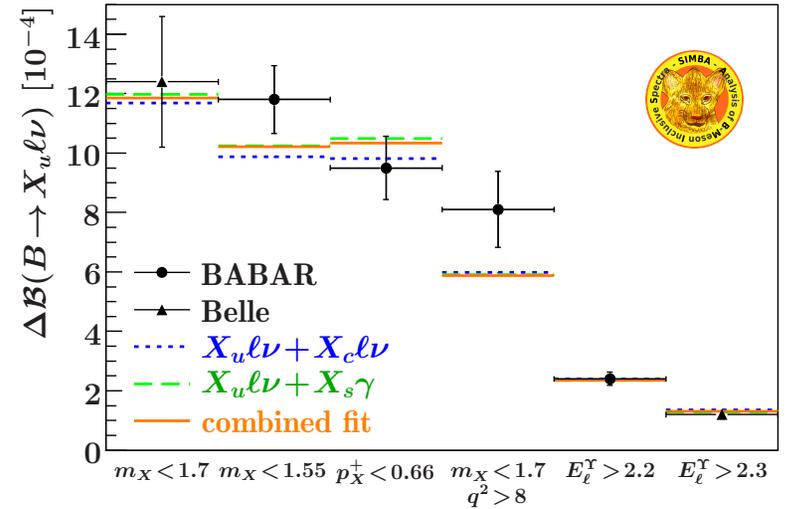
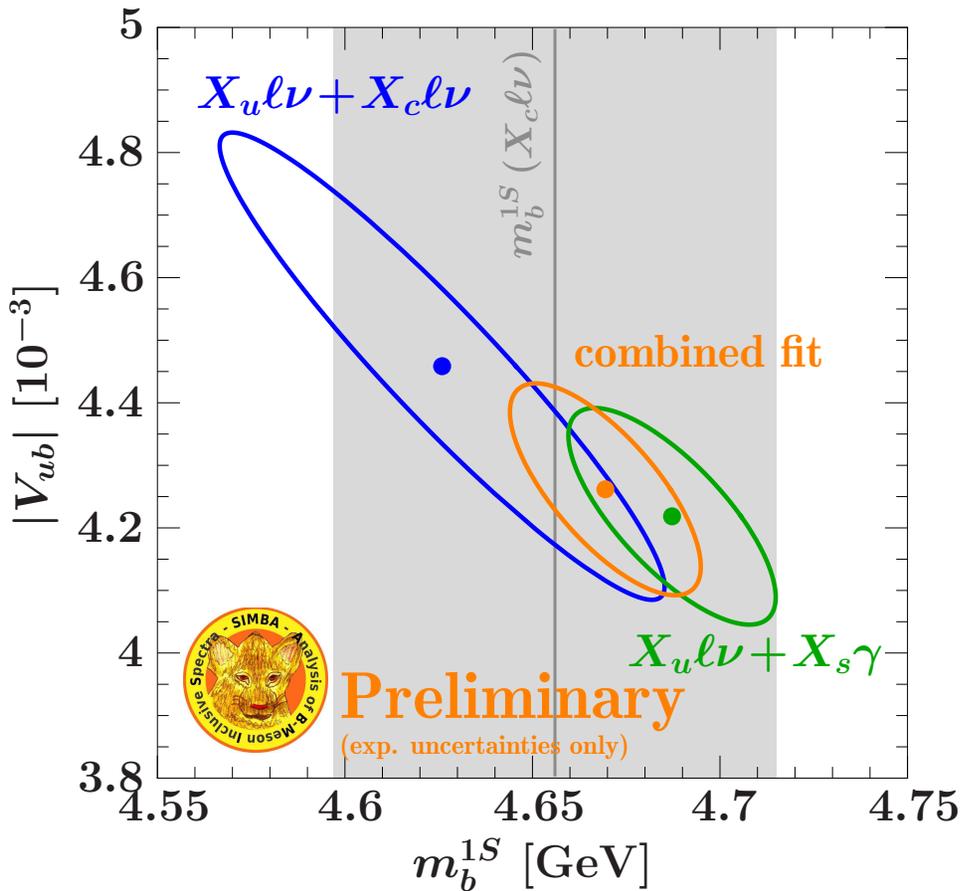
- **BLNP** [Bosch *et al.*] — based on **SCET** region, tied to “shape function scheme”
- **DGE** [Andersen & Gardi] — based on **SCET** region + perturbative model for the SF
- **GGOU** [Gambino *et al.*] — based on **local OPE** region + SF smearing
- **BLL** [Bauer, ZL, Luke] — based on **local OPE** at large q^2 (but expansion scale is smaller)
 - combine q^2 and m_X cuts, such that SF effect is kept small
- **Shape function independent relations** [Leibovich, Low, Rothstein; Hoang, ZL, Luke; Lange, Neubert, Paz; Lange]
 - beautiful at leading order, less so when $\mathcal{O}(\Lambda_{\text{QCD}}/m_b)$ included

Exploratory: $|V_{ub}|$ w/ NLO + NLL' only

- $B \rightarrow X_u \ell \bar{\nu}$ hadronic tag
 - BaBar $m_X, m_X - q^2, p_X^+$
 - Belle m_X
- $B \rightarrow X_u \ell \bar{\nu}$ lepton endpoint
 - BaBar $E_\ell^\Upsilon > 2.2$ GeV
 - Belle $E_\ell^\Upsilon > 2.3$ GeV
- $B \rightarrow X_s \gamma$ spectra
 - Belle latest result (shown)
 - BaBar sum over exclusive + hadronic tag
- m_b^{1S}, λ_1 from $B \rightarrow X_c \ell \bar{\nu}$ fit
 - $m_b^{1S} = (4.66 \pm 0.05)$ GeV
 - $\lambda_1 = (-0.34 \pm 0.05)$ GeV²



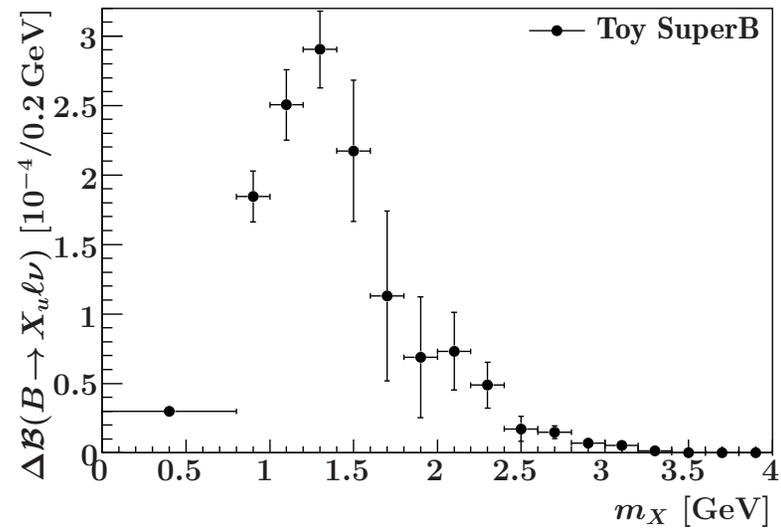
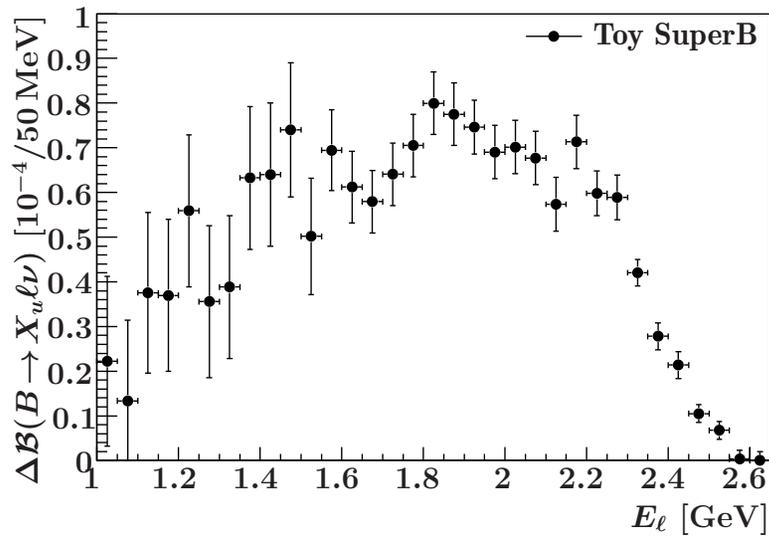
Exploratory: $|V_{ub}|$ w/ NLO + NLL' only



- E_γ spectrum is off without $B \rightarrow X_s \gamma$ in the fit
- Including it, favors lower values of $|V_{ub}|$

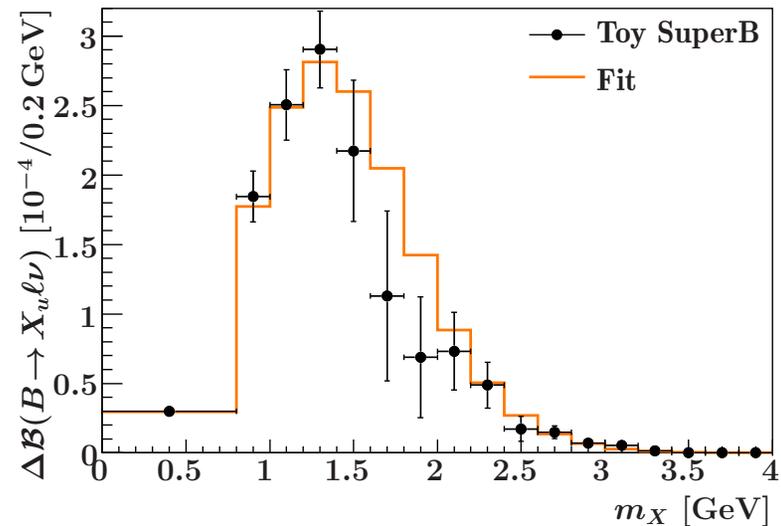
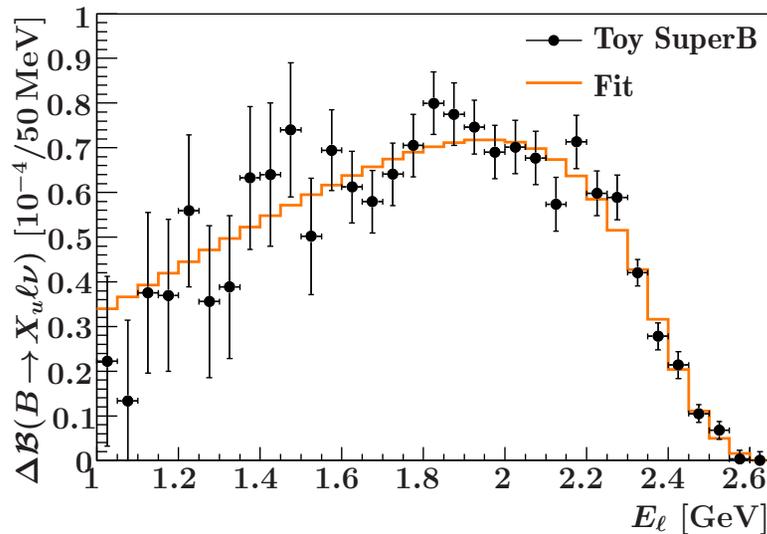
Future of $B \rightarrow X_u \ell \bar{\nu}$

- Spectra generated with $\lambda = 0.6 \text{ GeV}$ and $c_0 = 1$ (Assumed uncertainties & correlations similar to BaBar full reco analysis, 1112.0702 — by now Belle hadronic tagging efficiency is better)



Future of $B \rightarrow X_u \ell \bar{\nu}$

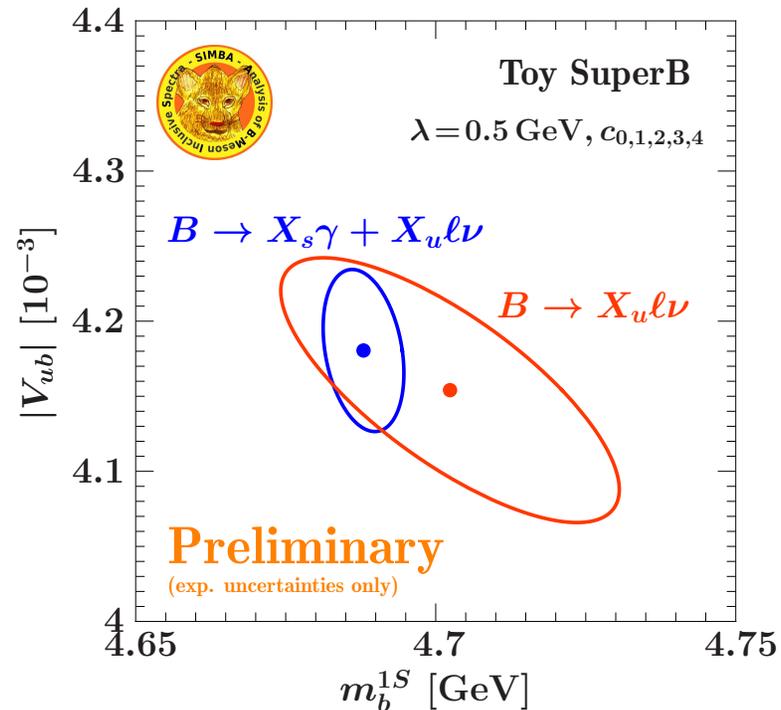
- Spectra generated with $\lambda = 0.6 \text{ GeV}$ and $c_0 = 1$ (Assumed uncertainties & correlations similar to BaBar full reco analysis, 1112.0702 — by now Belle hadronic tagging efficiency is better)



- Measure spectra — the rate with low E_ℓ or high m_X cut cannot give optimal $|V_{ub}|$
 - Uncertainties grow, as for $d\Gamma(B \rightarrow X_s \gamma)/dE_\gamma$
 - Experimental analysis needs input on shape in any case
- Large data sets will push analysis to the limits, constrain subleading SF effects

Future of $B \rightarrow X_u \ell \bar{\nu}$ (2)

- Toy fit with 5 coefficients for 75/ab:



- With Belle II data sets:

- Combination with $B \rightarrow X_s \gamma$ will be essential for ultimate sensitivity
- Combination with $B \rightarrow X_c \ell \bar{\nu}$ (moments, shapes?) also possible

Final comments

Some comments on $|V_{ub}|$ inclusive

- Is {in/ex}clusive tension a nuisance or tip of an iceberg? (right-handed currents?)
- Recently $\Gamma(D_s \rightarrow X\ell\bar{\nu})$ gave some indication of what the resolution is not
- Qualitatively better analyses are possible than those implemented so far
 - Fitting $F(k)$ instead of modeling $S(\omega, \mu)$
 - Designer orthonormal functions — reduce role of shape function modeling
 - Fully consistent combination of all phase space regions
 - Decouple SF shape variation from m_b variation
- Inclusive $|V_{cb}|$ uses a combined fit; clearly the right method for $|V_{ub}|$ as well
Combine all $B \rightarrow X_s\gamma, X_u\ell\bar{\nu}, X_c\ell\bar{\nu}$ data to constrain short distance physics & SFs
Need all available spectra and correlations
- $|V_{ub}|$ is tricky: to draw conclusions about new physics, we'll want ≥ 2 extractions with different uncertainties to agree well (inclusive, exclusive, leptonic)

Conclusions

- Current status of $|V_{ub}|$ unsettled — improvement crucial to better constrain NP
Hope to see measurements w/ different uncertainties agree (incl., excl., leptonic)
- Qualitatively better inclusive $|V_{ub}|$ analysis possible than those implemented so far
Full hadronic reconstruction B sample at Belle II is crucial for this
- Measure all possible **spectra** with best possible precision
- **SIMBA allows:** eliminate shape function modeling
constrain subleading shape functions
consistently combine all $B \rightarrow X_u \ell \bar{\nu}$ measurements
consistently include $B \rightarrow X_s \gamma$ measurements
consistently include $B \rightarrow X_c \ell \bar{\nu}$ constraints
eventually include/predict $B \rightarrow X_s \ell^+ \ell^-$



Backup slides

Derivation of the magic formula (1)

- The shape function is the matrix element of a nonlocal operator:

$$S(\omega, \mu) = \langle B | \underbrace{\bar{b}_v \delta(iD_+ - \delta + \omega) b_v}_{O_0(\omega, \mu)} | B \rangle, \quad \delta = m_B - m_b$$

Integrated over a large enough region, $0 \leq \omega \leq \Lambda$, one can expand O_0 as

$$O_0(\omega, \mu) = \sum C_n(\omega, \mu) \underbrace{\bar{b}_v (iD_+ - \delta)^n b_v}_{Q_n} + \dots = \sum C_n(\omega - \delta, \mu) \underbrace{\bar{b}_v (iD_+)^n b_v}_{\tilde{Q}_n} + \dots$$

The C_n are the same for Q_n and \tilde{Q}_n (since O_0 only depends on $\omega - \delta$)

- Matching:** $\langle b_v | O_0(\omega + \delta, \mu) | b_v \rangle = \sum C_n(\omega, \mu) \langle b_v | \tilde{Q}_n | b_v \rangle = C_0(\omega, \mu), \quad \langle b_v | \tilde{Q}_n | b_v \rangle = \delta_{0n}$

$$\langle b_v(k_+) | O_0(\omega + \delta, \mu) | b_v(k_+) \rangle = C_0(\omega + k_+, \mu) = \sum \frac{k_+^n}{n!} \frac{d^n C_0(\omega, \mu)}{d\omega^n}$$

$$\langle b_v(k_+) | O_0(\omega + \delta, \mu) | b_v(k_+) \rangle = \sum C_n(\omega, \mu) \langle b_v | \tilde{Q}_n | b_v \rangle = \sum C_n(\omega, \mu) k_+^n$$

- Comparing last two lines:** $C_n(\omega, \mu) = \frac{1}{n!} \frac{d^n C_0(\omega, \mu)}{d\omega^n}$

[Bauer & Manohar]

Derivation of the magic formula (2)

- Define the nonperturbative function $F(k)$ by: [ZL, Stewart, Tackmann; Lee, ZL, Stewart, Tackmann]

$$S(\omega, \mu_\Lambda) = \int dk C_0(\omega - k, \mu_\Lambda) F(k), \quad C_0(\omega, \mu) = \langle b_v | O_0(\omega + \delta, \mu) | b_v \rangle$$

uniquely defines $F(k)$: $\tilde{F}(y) = \tilde{S}(y, \mu) / \tilde{C}_0(y, \mu)$

- Expand in k : $S(\omega, \mu) = \sum_n \frac{1}{n!} \frac{d^n C_0(\omega, \mu)}{d\omega^n} \int dk (-k)^n F(k)$

Compare with previous page $\Rightarrow \int dk k^n F(k) = (-1)^n \langle B | Q_n | B \rangle$

$$\langle B | Q_0 | B \rangle = 1, \quad \langle B | Q_1 | B \rangle = -\delta, \quad \langle B | Q_2 | B \rangle = -\frac{\lambda_1}{3} + \delta^2$$

More complicated situation for higher moments, so stop here

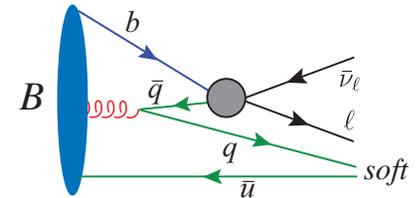
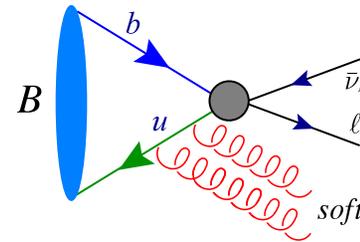
- This treatment is fully consistent with the OPE

Weak annihilation

- **Hard to estimate:** $(16\pi^2) (\Lambda_{\text{QCD}}^3/m_b^3) \varepsilon$, centered near $q^2 = m_B^2$ and $E_\ell = m_B/2$

$$\langle B | (\bar{b}\gamma^\mu P_L u) (\bar{u}\gamma_\mu P_L b) | B \rangle = \frac{f_B^2 m_B}{8} B_1$$

$$\langle B | (\bar{b}P_L u) (\bar{u}P_L b) | B \rangle = \frac{f_B^2 m_B}{8} B_2$$



Overall shift vs. splitting between B^\pm and B^0

Factorization + vacuum saturation: $B_{1,2} = \begin{cases} 1, & B^\pm \\ 0, & B^0 \end{cases}$ assume $\varepsilon \equiv B_1 - B_2 \sim 0.1$

Rate: $\Gamma_{\text{WA}} = \frac{G_F^2 m_b^2 |V_{ub}|^2}{12\pi} f_B^2 m_B (B_2 - B_1) \sim 3\%$ of $\Gamma(B \rightarrow X_u \ell \bar{\nu})$ [Voloshin, hep-ph/0106040]

- **Enters all $|V_{ub}|$ measurements, enhanced by $(m_b/m_c)^3 \sim 30$ in $D_{u,d,s}$ decays**

$\Gamma(D^0 \rightarrow X \ell \bar{\nu}) \approx \Gamma(D^\pm \rightarrow X \ell \bar{\nu})$ to $\lesssim 3\%$, recently $\Gamma(D_s \rightarrow X \ell \bar{\nu})$ [CLEO-c, arXiv:0912.4232]

No evidence that WA is bigger when light quark in operator = spectator flavor

- **Probably a smaller effect in the determination of $|V_{ub}|$ than typically assumed**

[ZL, Luke, Manohar, arXiv:1003.1351; Gambino & Kamenik, arXiv:1004.0114]