



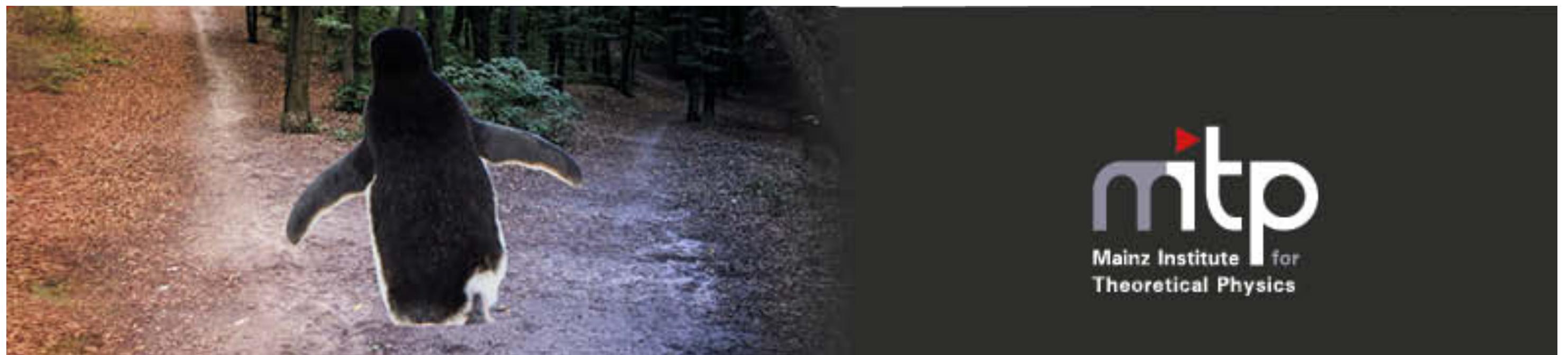
Universität  
Zürich<sup>UZH</sup>

# Renormalization Group Evolved 4321 and intriguing behaviors in the UV

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MITP Scientific Program, Flavor at the Crossroads, 27 April 2022

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University of Zurich (UZH)



**mitp**  
Mainz Institute for  
Theoretical Physics

Based on

*R. Houtz, J. Pagès and S. Trifinopoulos, arXiv:2204.06440*

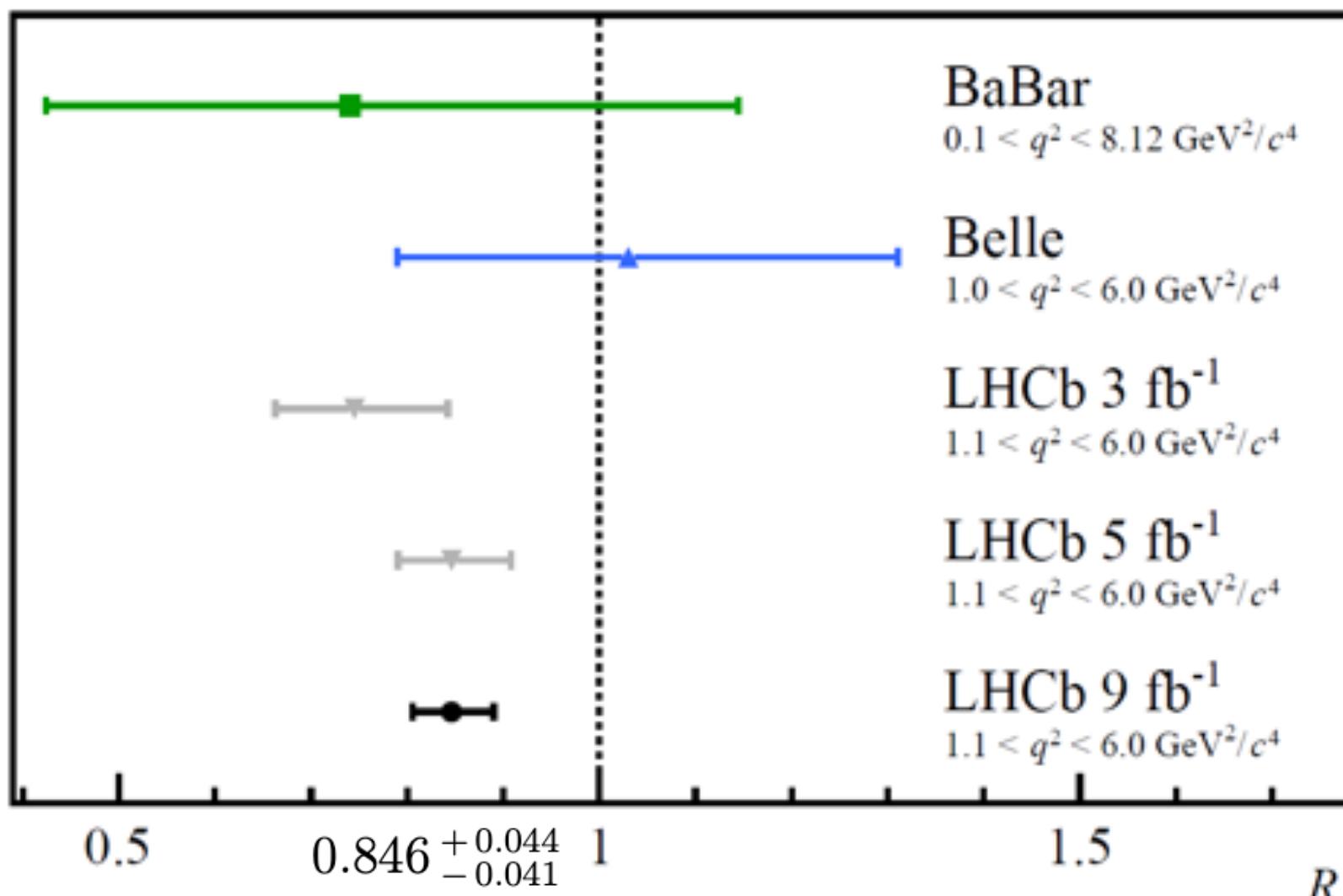


# B anomalies

## Neutral currents

$b \rightarrow s \ell^+ \ell^-$  : universality in  $\mu$  vs.  $e$

$$R_K = \frac{\Gamma(B \rightarrow K \mu^+ \mu^-)}{\Gamma(B \rightarrow K e^+ e^-)} \quad 3.1 \sigma$$



+ other observables:

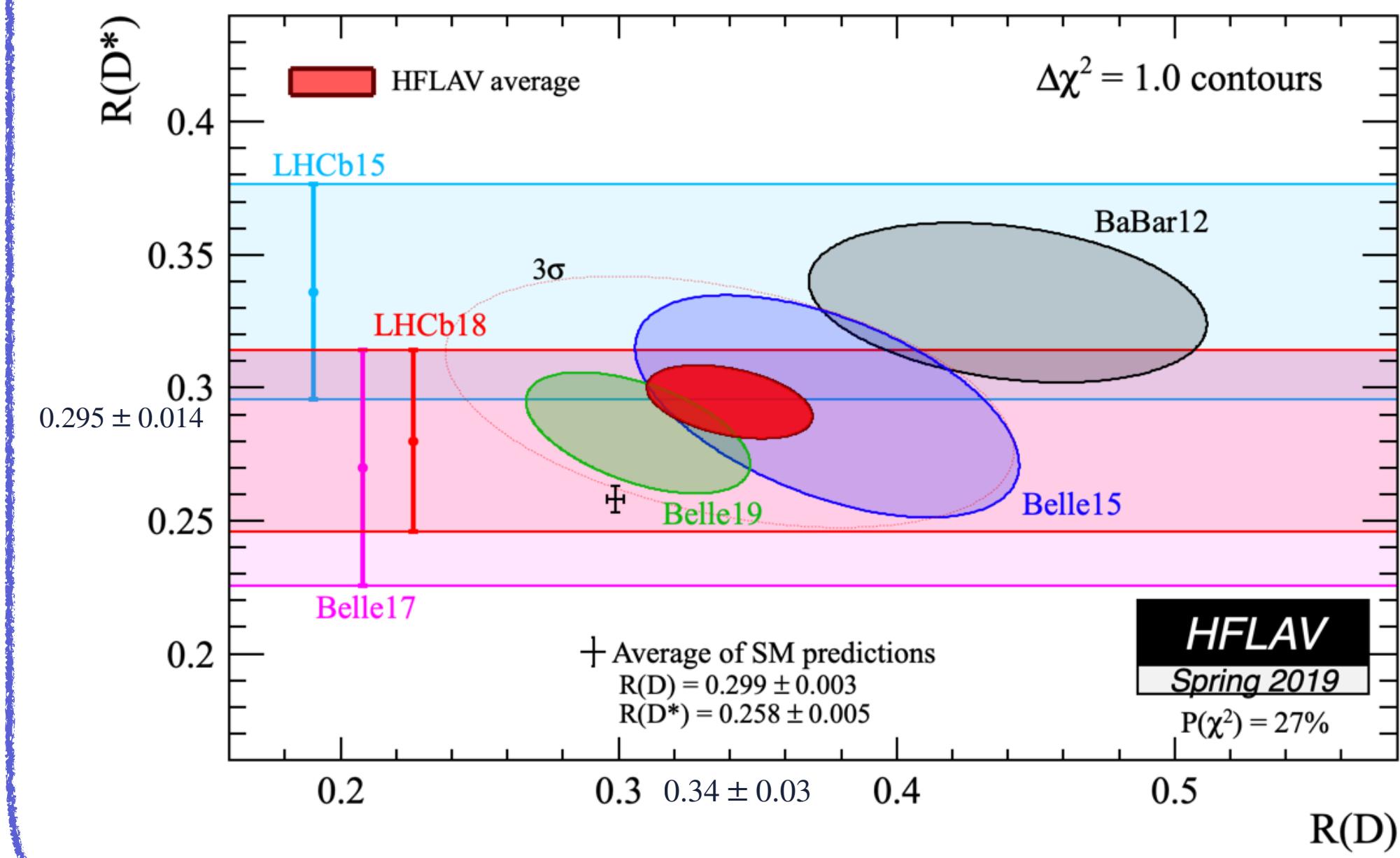
$$R_{K^{(*)}}, P'_5, B \rightarrow K^{(*)} \mu^+ \mu^-, B_s \rightarrow \mu^+ \mu^-, B_s \rightarrow \phi \mu^+ \mu^-, \dots$$

> 4  $\sigma$

## Charged currents

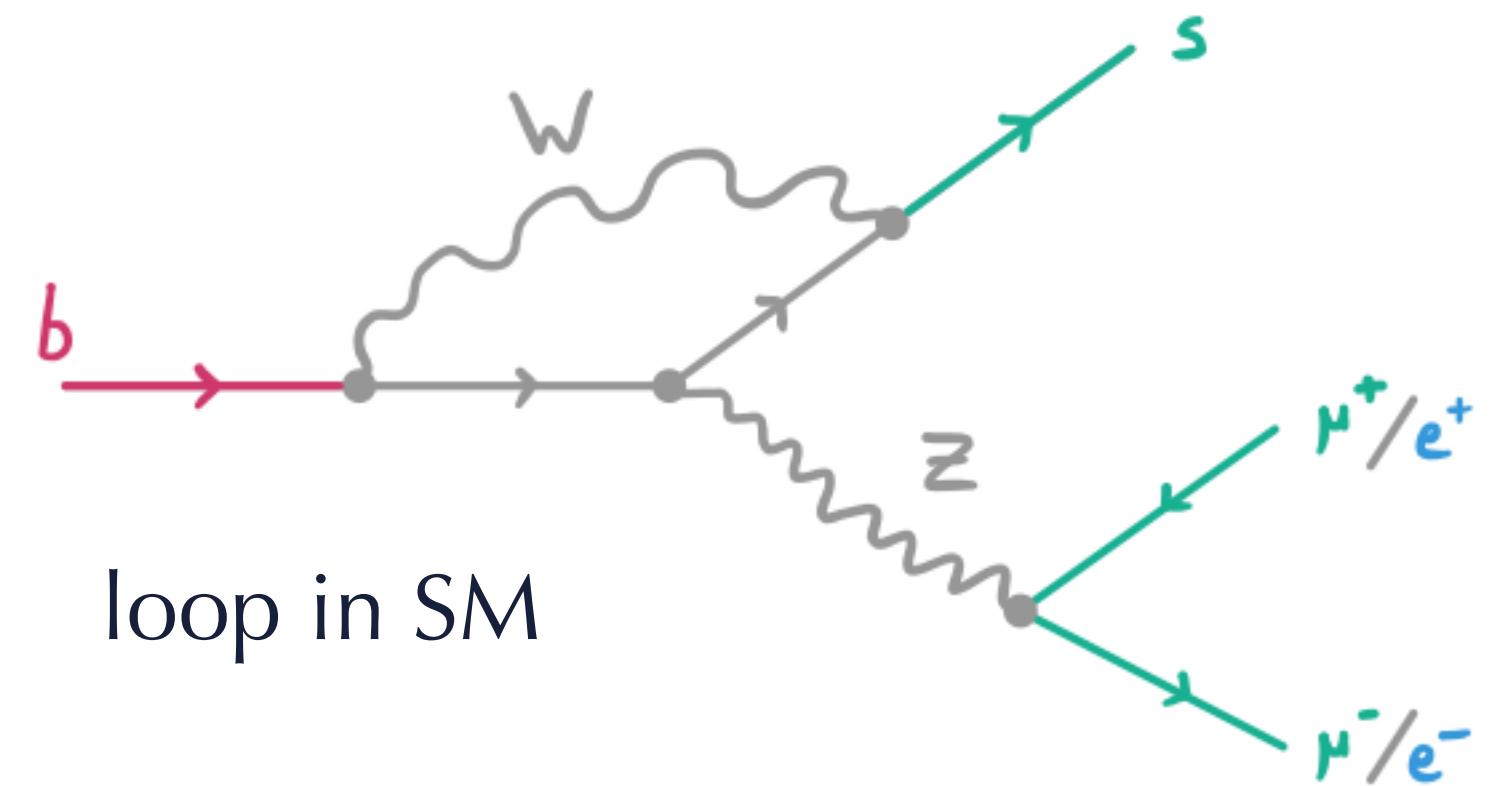
$b \rightarrow c \ell \nu$  : universality in  $\tau$  vs.  $\mu, e$

$$R_{D^{(*)}} = \frac{\Gamma(B \rightarrow D^{(*)} \tau \nu)}{\Gamma(B \rightarrow D^{(*)} \ell \nu)} \quad 3.4 \sigma$$



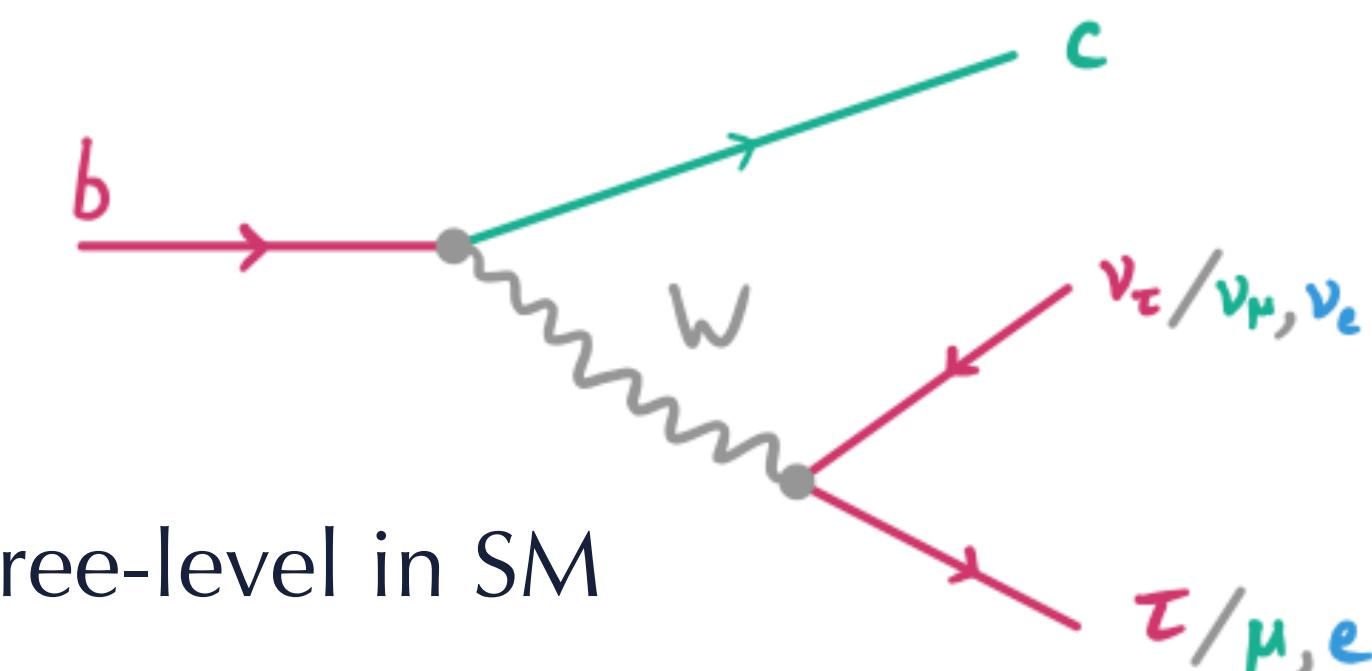
# Vector LQ mediator for the B anomalies

Neutral currents:  $b \rightarrow s\ell^+\ell^-$



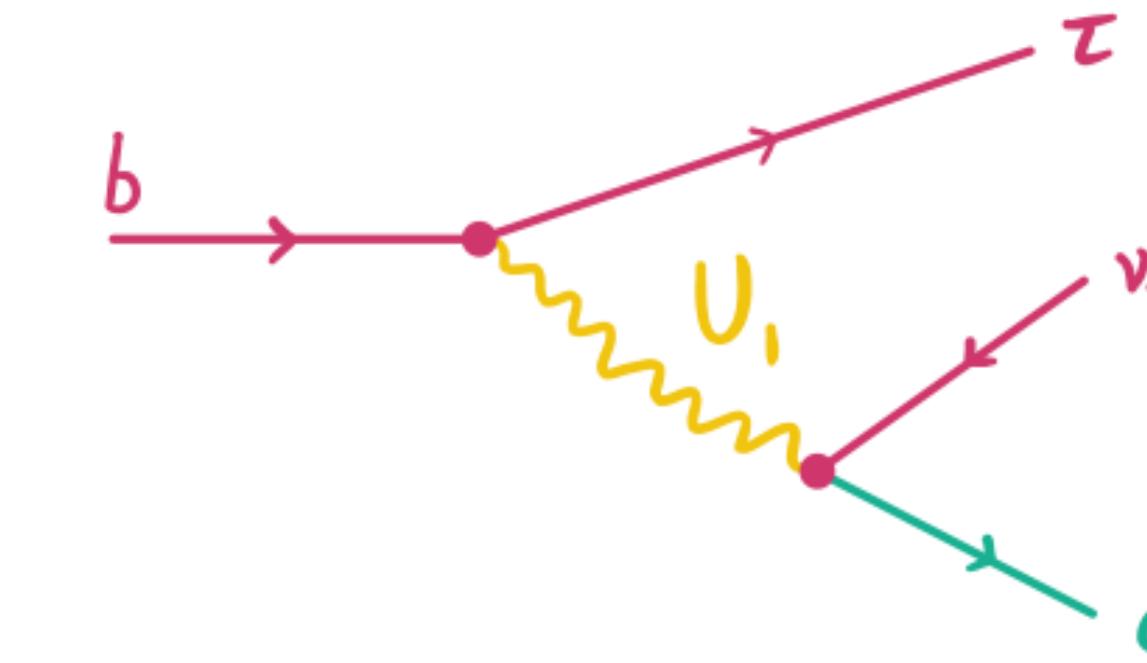
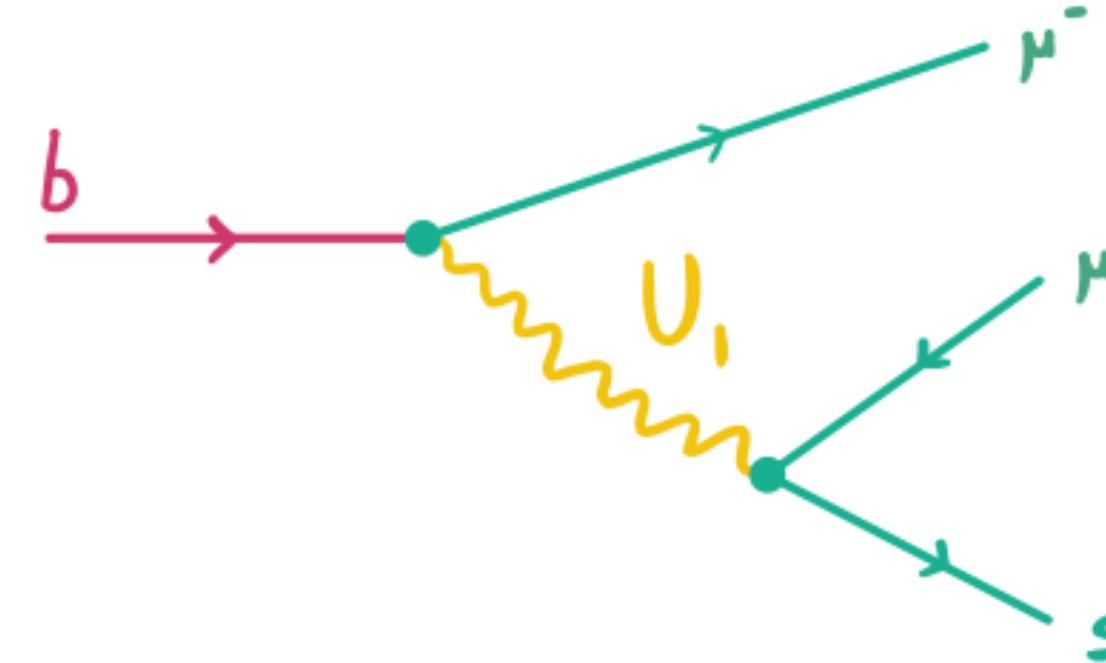
loop in SM

Charged currents:  $b \rightarrow c\ell\nu$

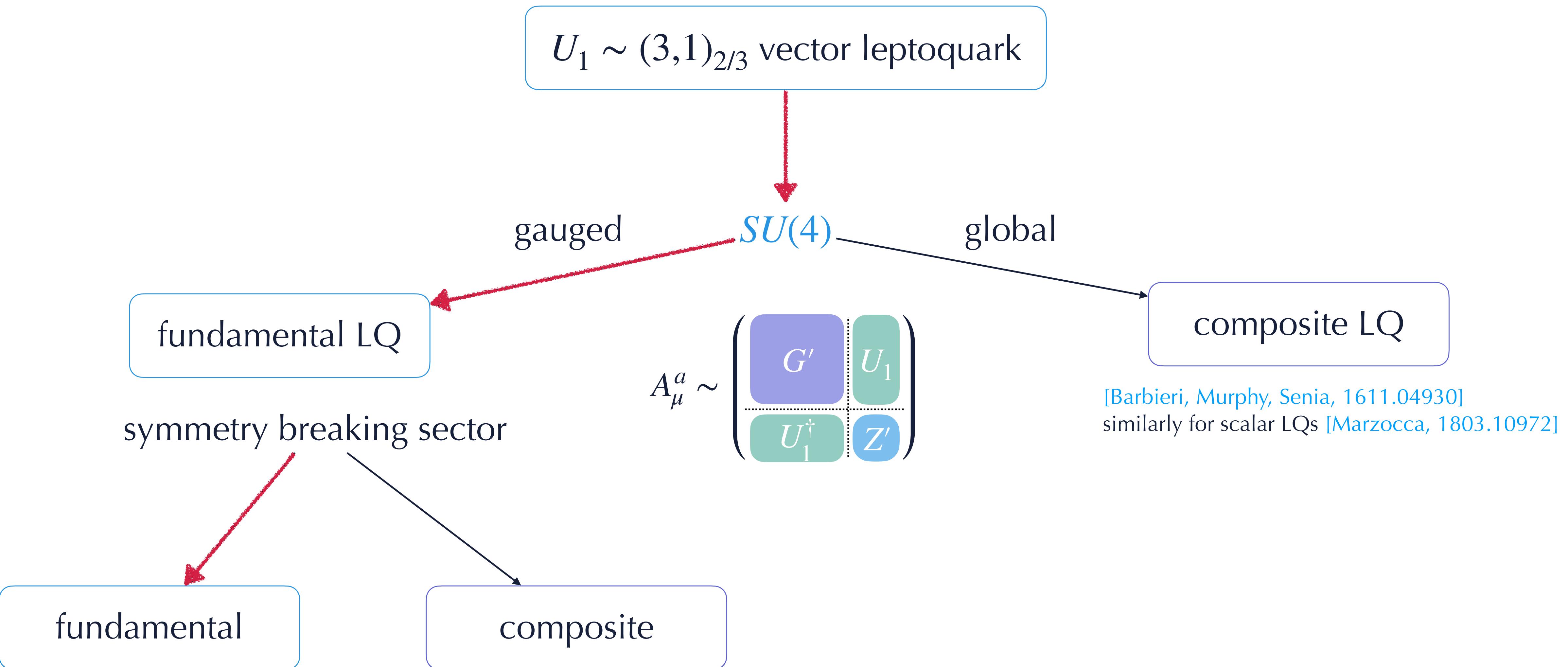


tree-level in SM

Combined single mediator solution: the  $U_1 \sim (3,1)_{2/3}$  vector leptoquark



# UV origin of the $U_1$ leptoquark



[Diaz, Schmaltz, Zhong, 1706.05033]

[Di Luzio, Greljo, Nardecchia, 1708.08450]

[Fuentes-Martín, Stangl, 2004.11376]

[Fuentes-Martin, Isidori, Lizana, Selimovic, Stefanek, 2203.01952]

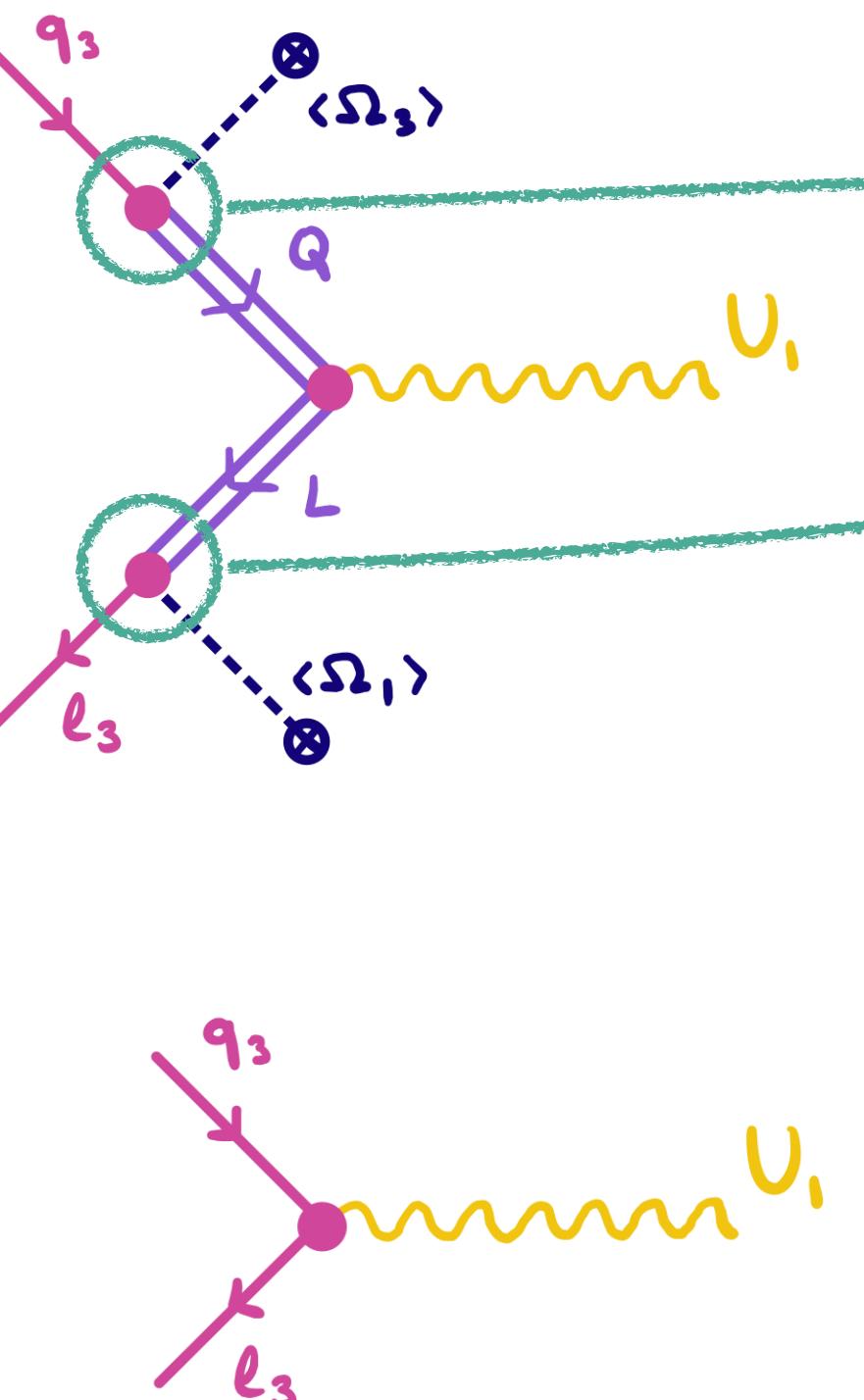
# 4321 model

- Flavor universal

[Di Luzio, Greljo, Nardecchia, 1708.08450]

[Di Luzio, Fuentes-Martín, Greljo, Nardecchia, Renner, 1808.00942]

- large mixing between 3<sup>rd</sup> generation SM fields and vector-like partners  
 ↳ Landau poles  $\lesssim 100$  TeV



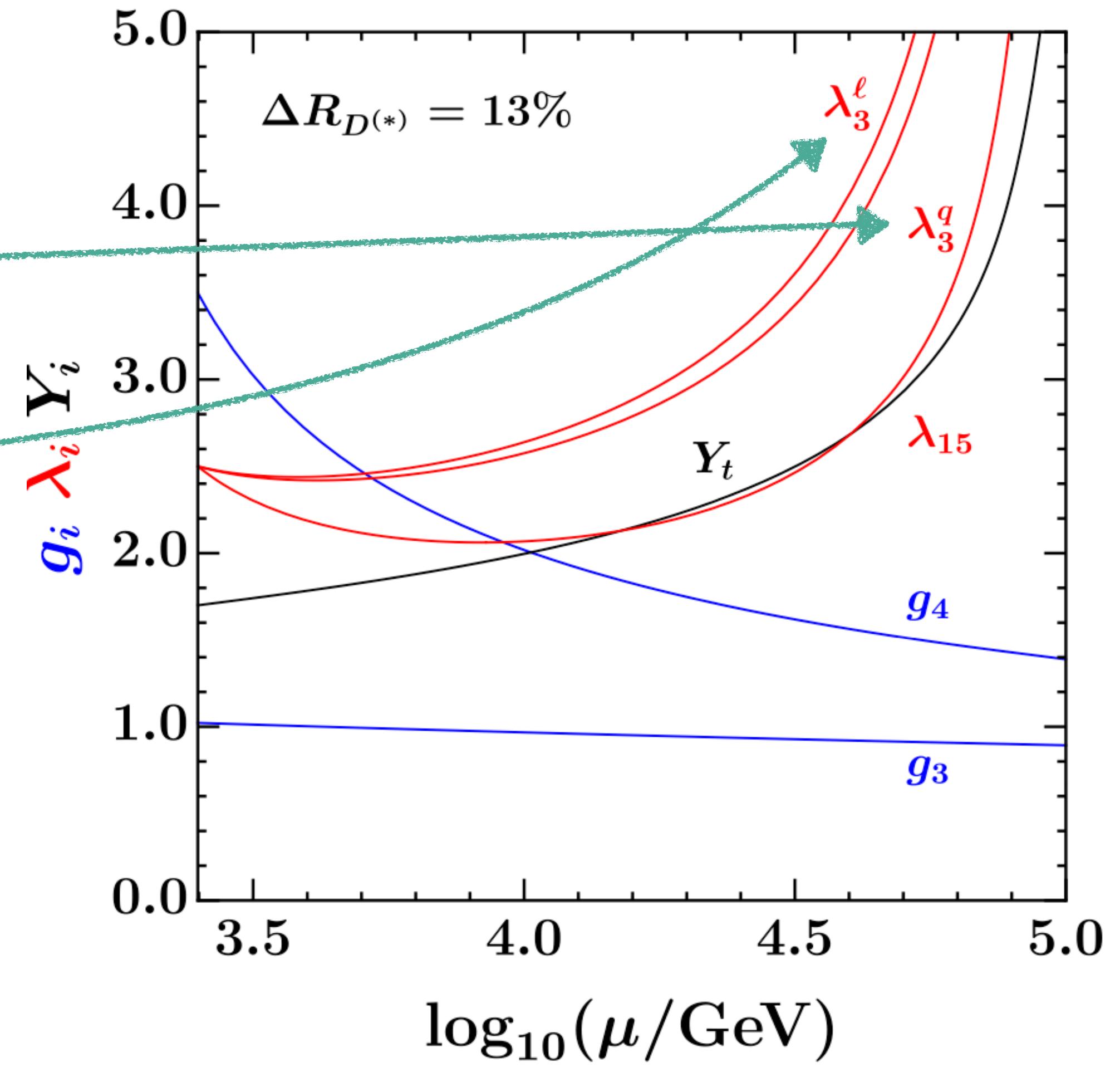
- Flavor non-universal

[Greljo, Stefanek, 1802.04274]

[Bordone, Cornella, Fuentes-Martín, Isidori, 1805.09328]

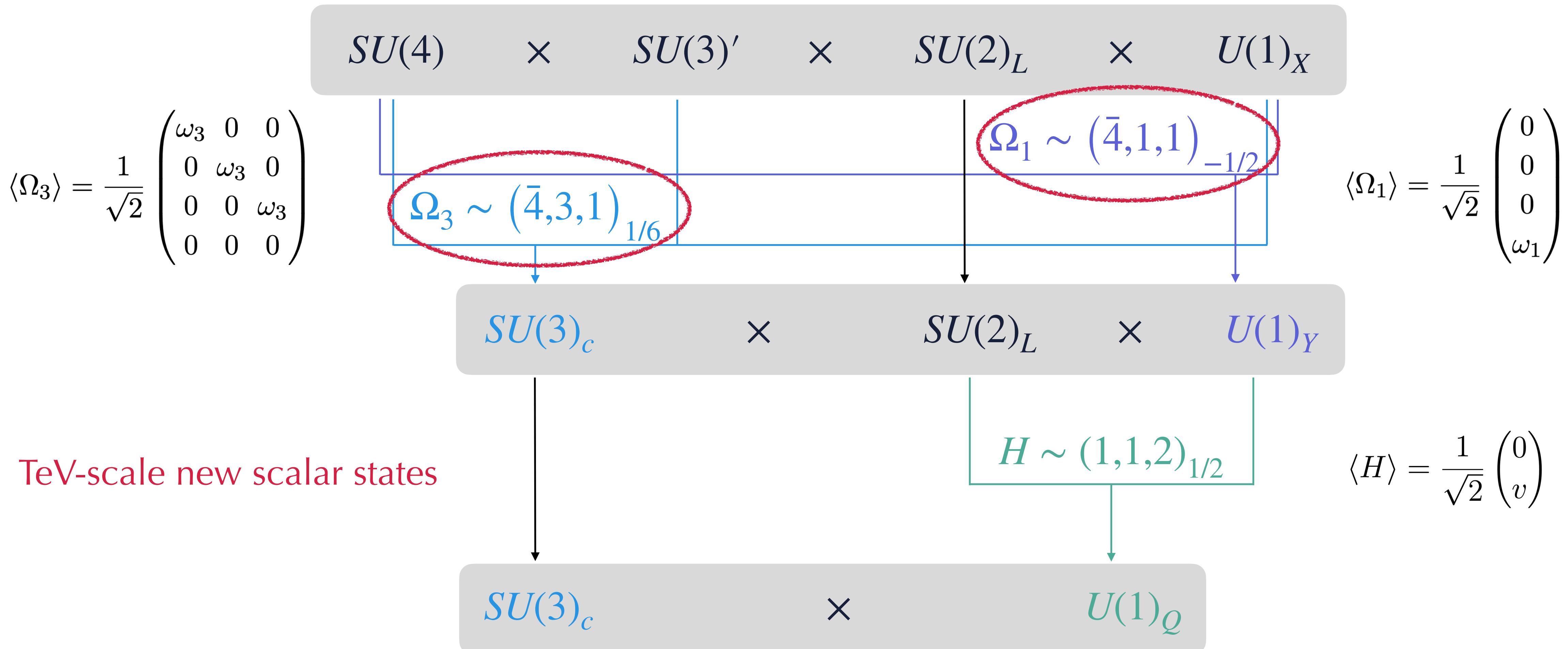
[Cornella, Fuentes-Martin, Isidori, 1903.11517]

see Claudia's talk



[Di Luzio, Fuentes-Martín, Greljo, Nardecchia, Renner, 1808.00942]

# Spontaneous symmetry breaking of the 4321 model



# Scalar potential in the 4321 model

Omega potential

with quartics  $\rho_i$

$$\begin{aligned} V_\Omega = & m_{\Omega_3}^2 \text{Tr}[\Omega_3^\dagger \Omega_3] + m_{\Omega_1}^2 \Omega_1^\dagger \Omega_1 + \frac{\rho_1}{2} (\Omega_1^\dagger \Omega_1)^2 + \frac{\rho_2}{2} \text{Tr}[\Omega_3^\dagger \Omega_3]^2 + \frac{\rho'_2}{2} \text{Tr}[\Omega_3^\dagger \Omega_3 \Omega_3^\dagger \Omega_3] \\ & + \rho_3 \text{Tr}[\Omega_3^\dagger \Omega_3] \Omega_1^\dagger \Omega_1 + \rho_4 \Omega_1^\dagger \Omega_3 \Omega_3^\dagger \Omega_1 + \frac{\rho_5}{3!} (\epsilon_{\alpha\beta\gamma\delta} \epsilon^{abc} (\Omega_3)_a^\alpha (\Omega_3)_b^\beta (\Omega_3)_c^\gamma (\Omega_1)^\delta + \text{h.c.}) \end{aligned}$$

Higgs potential

with quartic  $\lambda$

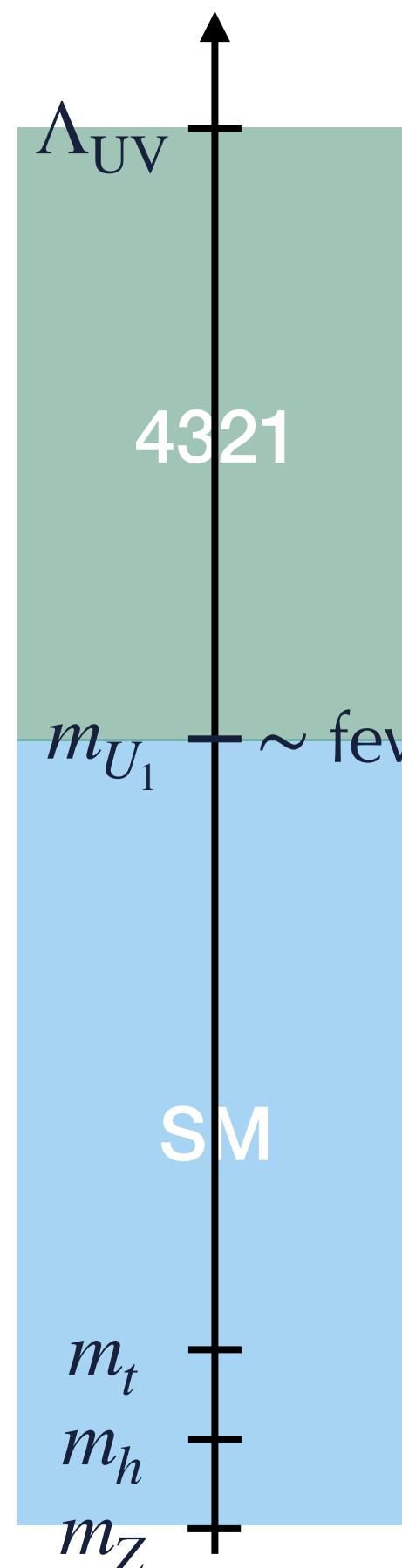
$$V_H = \mu_H^2 H^\dagger H + \frac{\lambda}{2} (H^\dagger H)^2$$

$H - \Omega_{1,3}$  mixing

with quartics  $\eta_i$

$$V_{\Omega H} = \eta_1 H^\dagger H \Omega_1^\dagger \Omega_1 + \eta_3 H^\dagger H \text{Tr}[\Omega_3^\dagger \Omega_3]$$

# Renormalization Group Evolution of the 4321

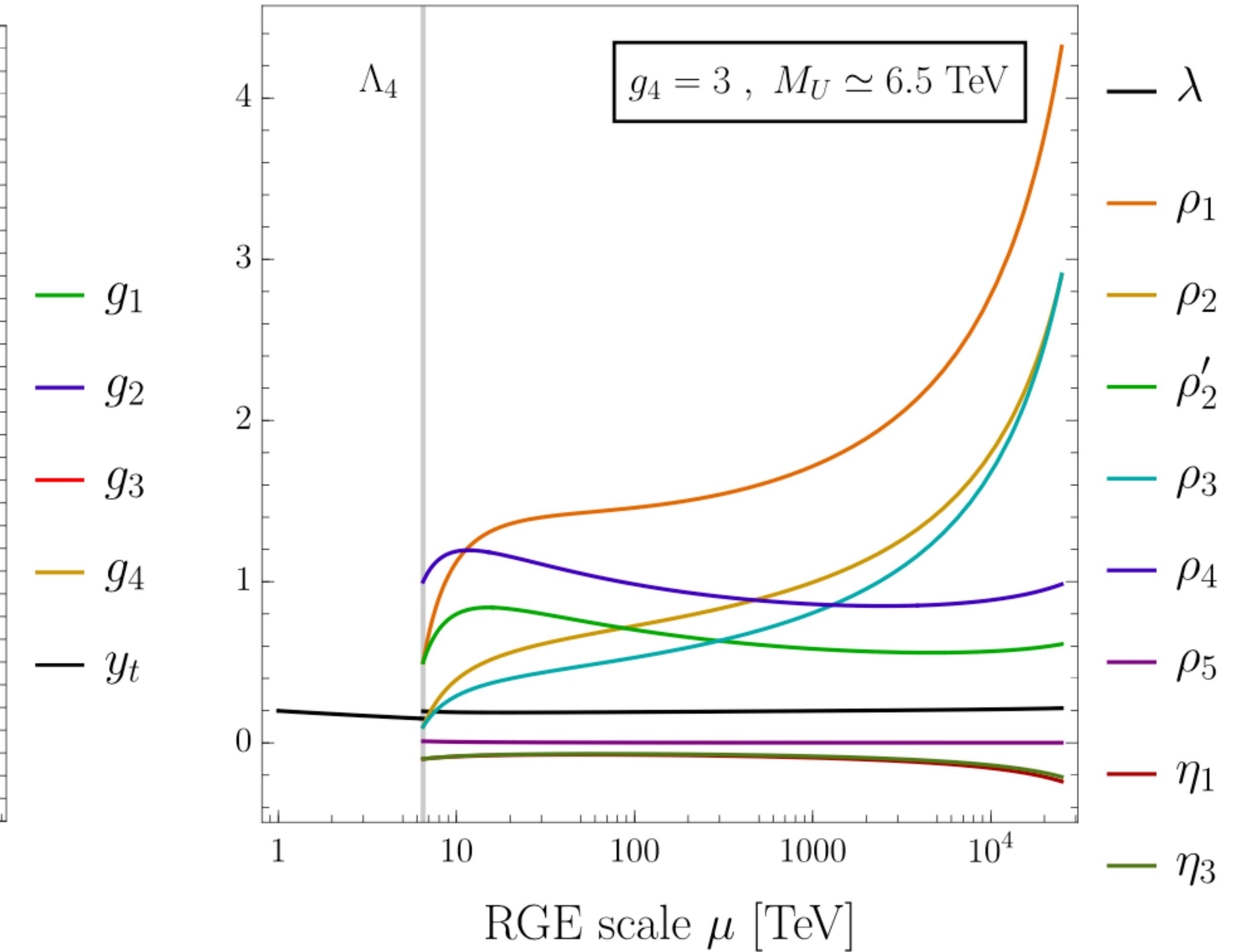
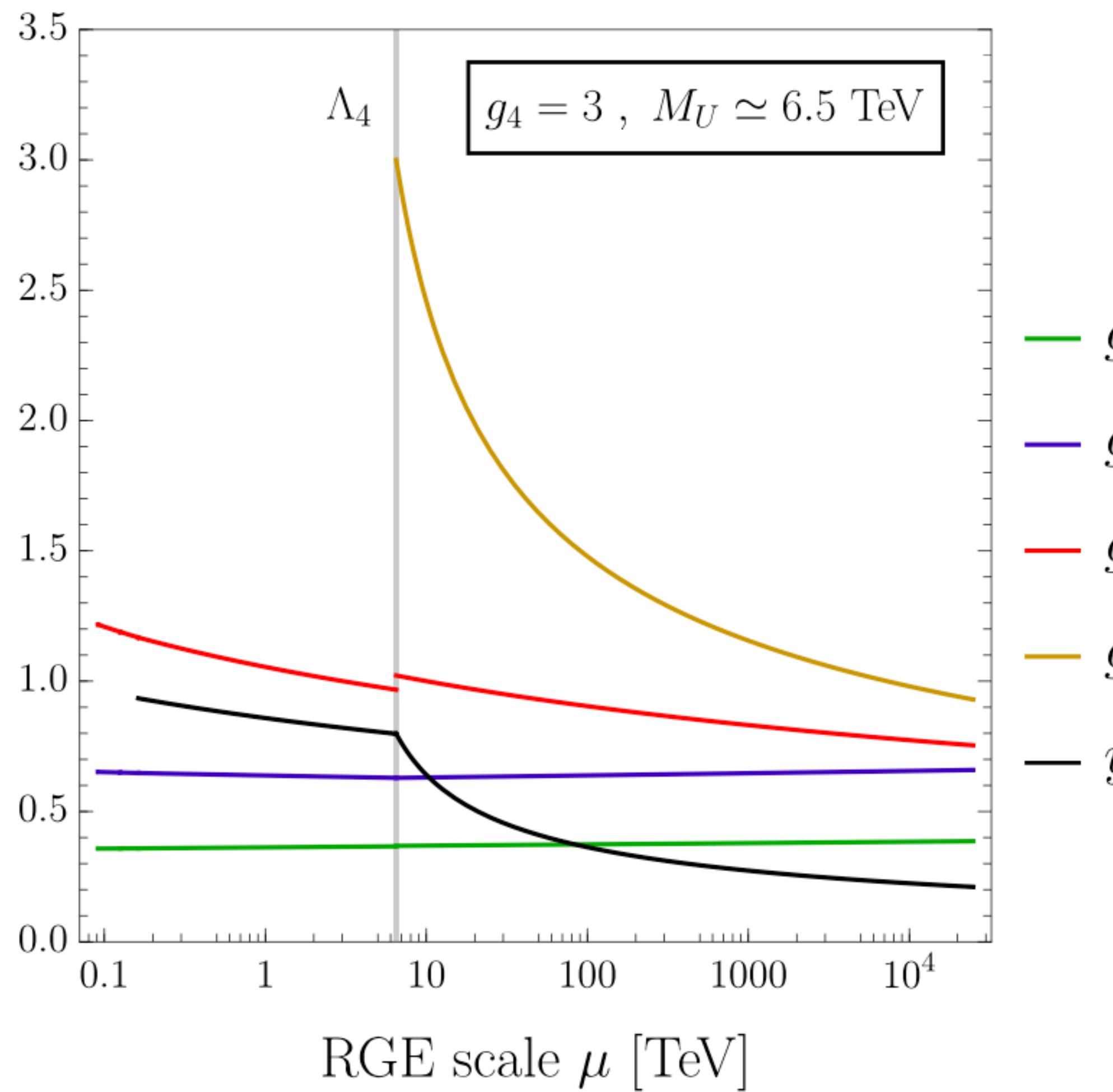


$$V = V_\Omega + V_H + V_{\Omega H}$$
$$\{\rho_i\}, m_{\Omega_{1,3}} \quad \lambda, \mu_H \quad \{\eta_i\}$$

RGE study:

- 1) Fix boundary conditions from phenomenology constraints  
    ↪ since under-constrained, pick a benchmark satisfying the constraints
- 2) Derive all  $\beta$ -functions for each theory/energy range at one-loop  
    crosschecked/computed with RGBeta [Thomsen, 2101.08265]
- 3) Run, Match (at tree-level), Run, etc...

# 1. Landau poles



“Asymptotic freedom” in gauge couplings

but

Landau poles develop in the  $\Omega$  potential quartics

# Masses of the radials

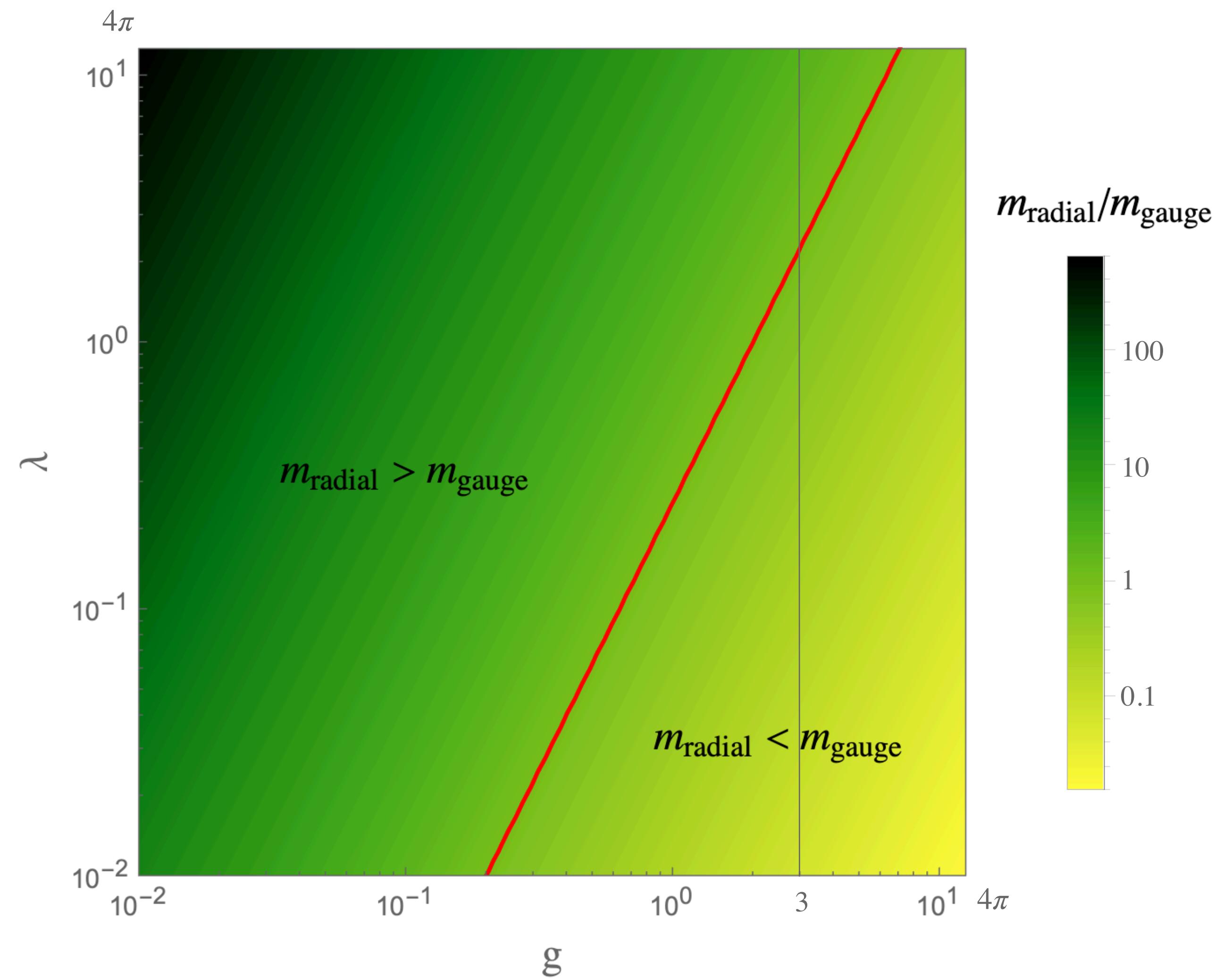
Decoupling of radial modes  
is not easily realised with  $g_4 = 3$

$$m_{\text{radial}} = \sqrt{\lambda} v$$

$$m_{\text{gauge}} = \frac{g v}{2}$$

$$\frac{m_{\text{radial}}}{m_{\text{gauge}}} = 2 \frac{\sqrt{\lambda}}{g} \lesssim 2.4 \left( \frac{3}{g} \right)$$

Decoupling limit corresponds to  $\lambda \gg 2$



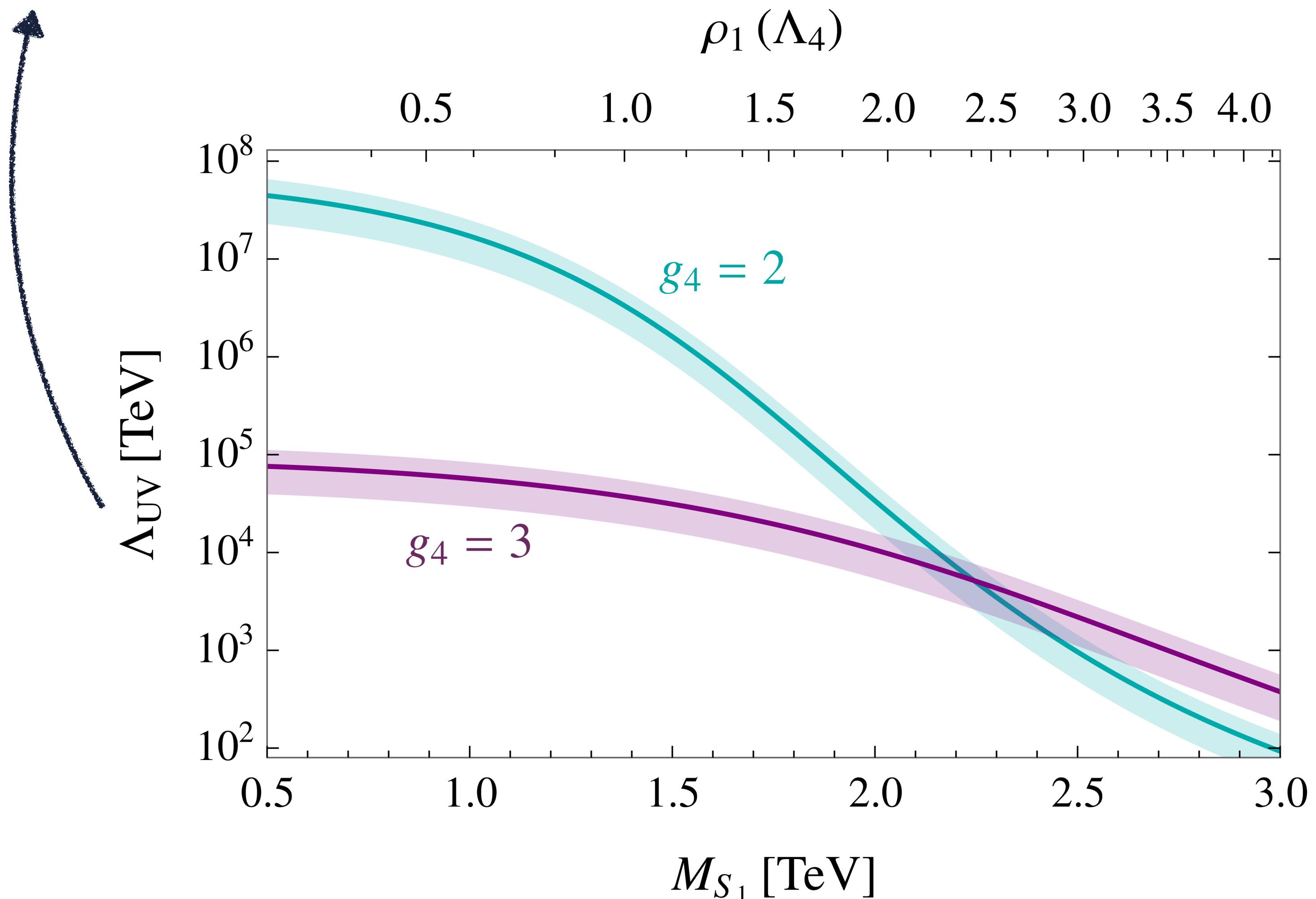
# Landau poles and masses of the radials

Running of the quartic  $\rho_1$

$$\begin{aligned}\beta_{\rho_1} = & \left( 16\rho_1 - 3g_1^2 - \frac{45}{2}g_4^2 \right) \rho_1 \\ & + 24\rho_3^2 + 12\rho_3\rho_4 + 6\rho_4^2 + 4\eta_1^2 \\ & + \frac{3}{4}g_1^4 + \frac{9}{4}g_1^2g_4^2 + \frac{99}{16}g_4^4\end{aligned}$$

Large quartic couplings  
=  
heavier radials  
=  
lower cut-off scale

Cut-off scale where  $\rho_1(\Lambda_{\text{UV}}) \geq 4\pi$

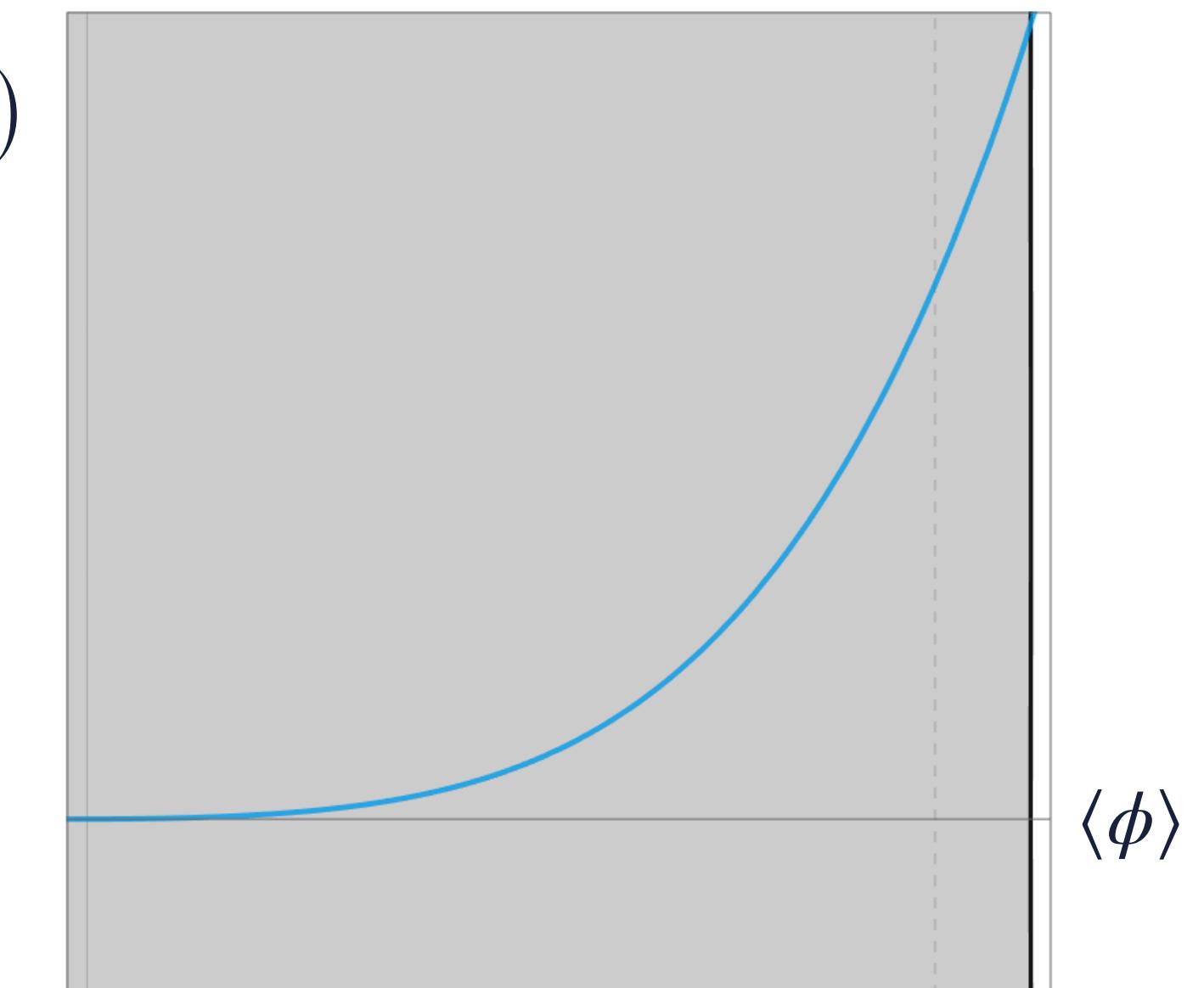


## 2. Radiative Electroweak symmetry breaking

Radiative electroweak symmetry breaking:

Electroweak symmetry is conserved at the classical level, but loop corrections to the mass parameter of the Higgs boson trigger its spontaneous breaking.

⇒ A positive Higgs mass parameter at high field value can turn negative at lower scale via the renormalization group flow.



Examples:

[Babu, Gogoladze, Khan, 1512.05185]

- Standard Model
- Type-I seesaw model
- Scalar singlet dark matter model

$$\begin{aligned}\beta_H^{\text{SM}} &\propto m_H^2 & \times \\ \beta_H^{\text{SS1}} &= \beta_H^{\text{SM}} - 4 |y_\nu|^2 |m_R|^2 & \times \\ \beta_H^{\text{sDM}} &= \beta_H^{\text{SM}} + \lambda_3 m_s^2 & \checkmark\end{aligned}$$

Necessary ingredients:

- new states  
positive contribution  
⇒ **TeV-scale new scalars**

# Radiative Electroweak Symmetry Breaking

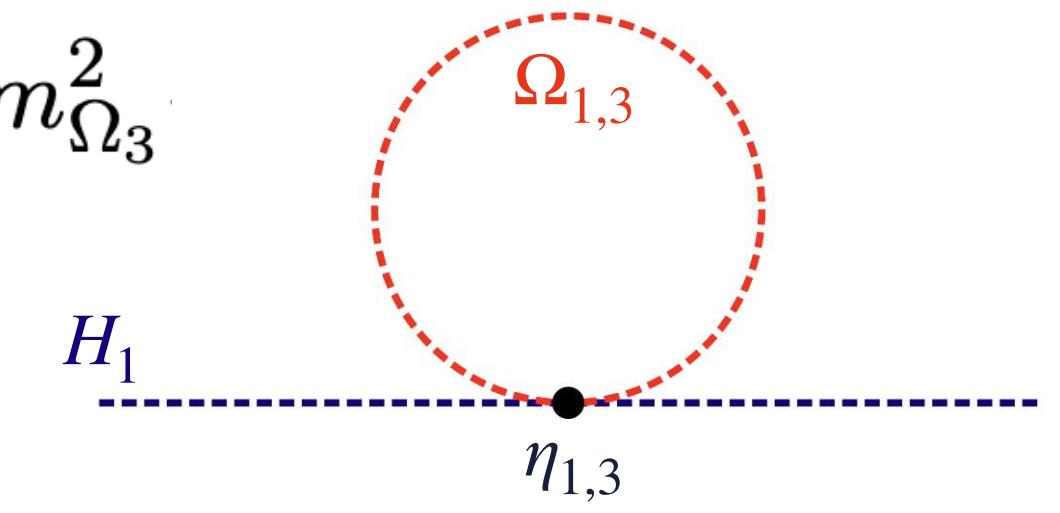
All ingredients are here!

| Field      | $SU(4)$ | $SU(3)'$ | $SU(2)_L$ | $U(1)_X$ |
|------------|---------|----------|-----------|----------|
| $\Omega_1$ | 4       | 1        | 1         | -1/2     |
| $\Omega_3$ | 4       | 3        | 1         | 1/6      |

TeV-scale new scalar states

$$\beta_{\mu_H^2} = \left( 6\lambda + 8y_t^2 - \frac{3}{2}g_1^2 - \frac{9}{2}g_2^2 \right) \mu_H^2$$

$$+ 8\eta_1 m_{\Omega_1}^2 + 24\eta_3 m_{\Omega_3}^2$$



positive contribution for  $\eta_{1,3} < 0$

Diagonalization of the Hessian (equivalent to integrating out  $\Omega_{1,3}$ ) gives the effective SM Higgs mass:

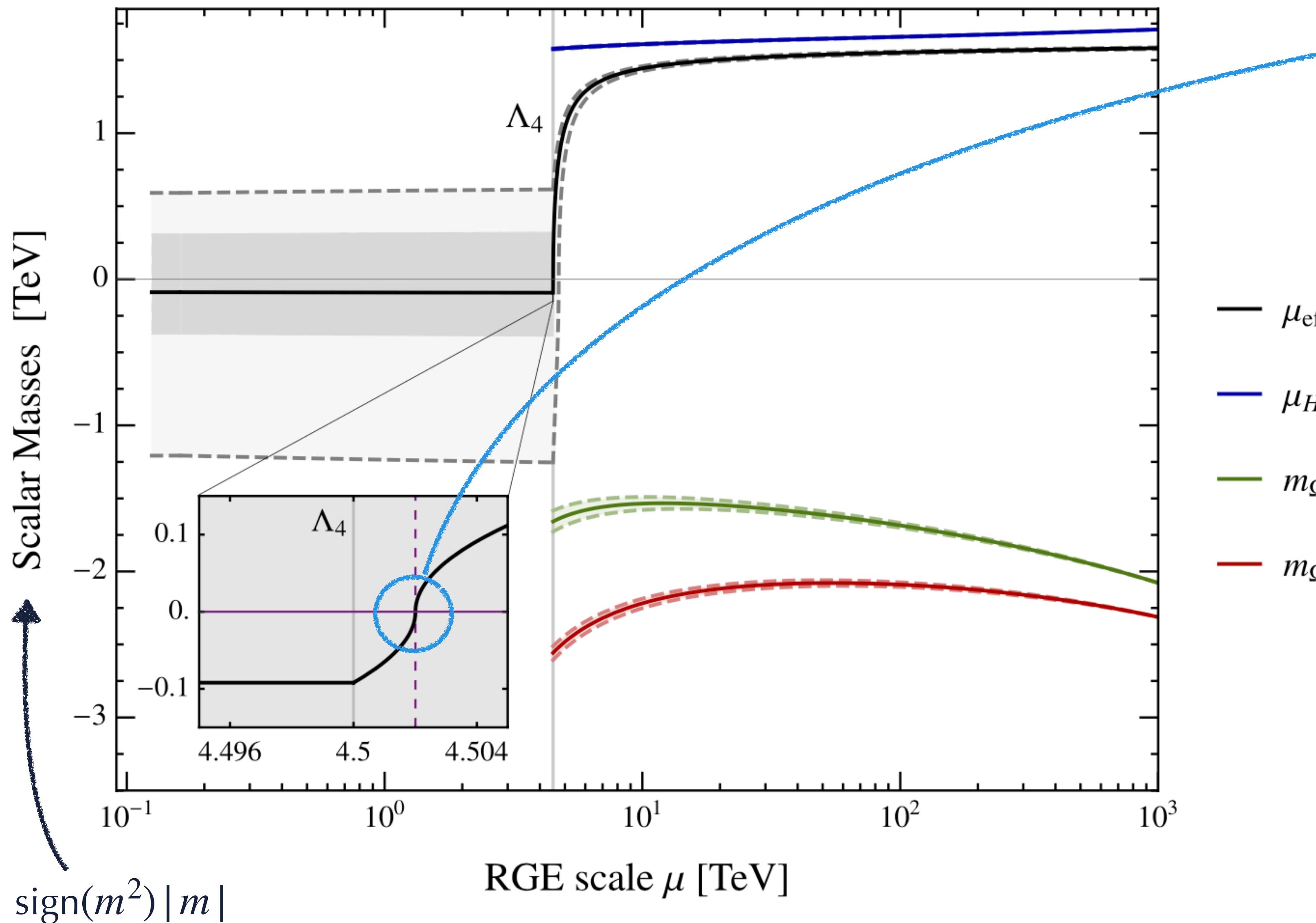
$$\mu_{\text{eff}}^2 = -\frac{\lambda_{\text{eff}}}{2}v^2 = \mu_H^2 - \frac{\eta_1 D_1 - 3\eta_3 \rho_3}{D_{12}} m_{\Omega_1}^2 - 3 \frac{\eta_3 \rho_1 - \eta_1 \rho_3}{D_{12}} m_{\Omega_3}^2$$

$$D_1 = 3\rho_2 + \rho'_2 > 0$$

$$D_{12} = \rho_1 D_1 - 3\rho_3^2 > 0$$

$$D_{123} = \lambda D_{12} - (3\eta_3^2 \rho_1 - 6\eta_1 \eta_3 \rho_3 + \eta_1^2 D_1) > 0$$

# Radiative Electroweak Symmetry Breaking



Radiative EWSB realised  
when  $\mu_{\text{eff}}$  flips sign

Checked:

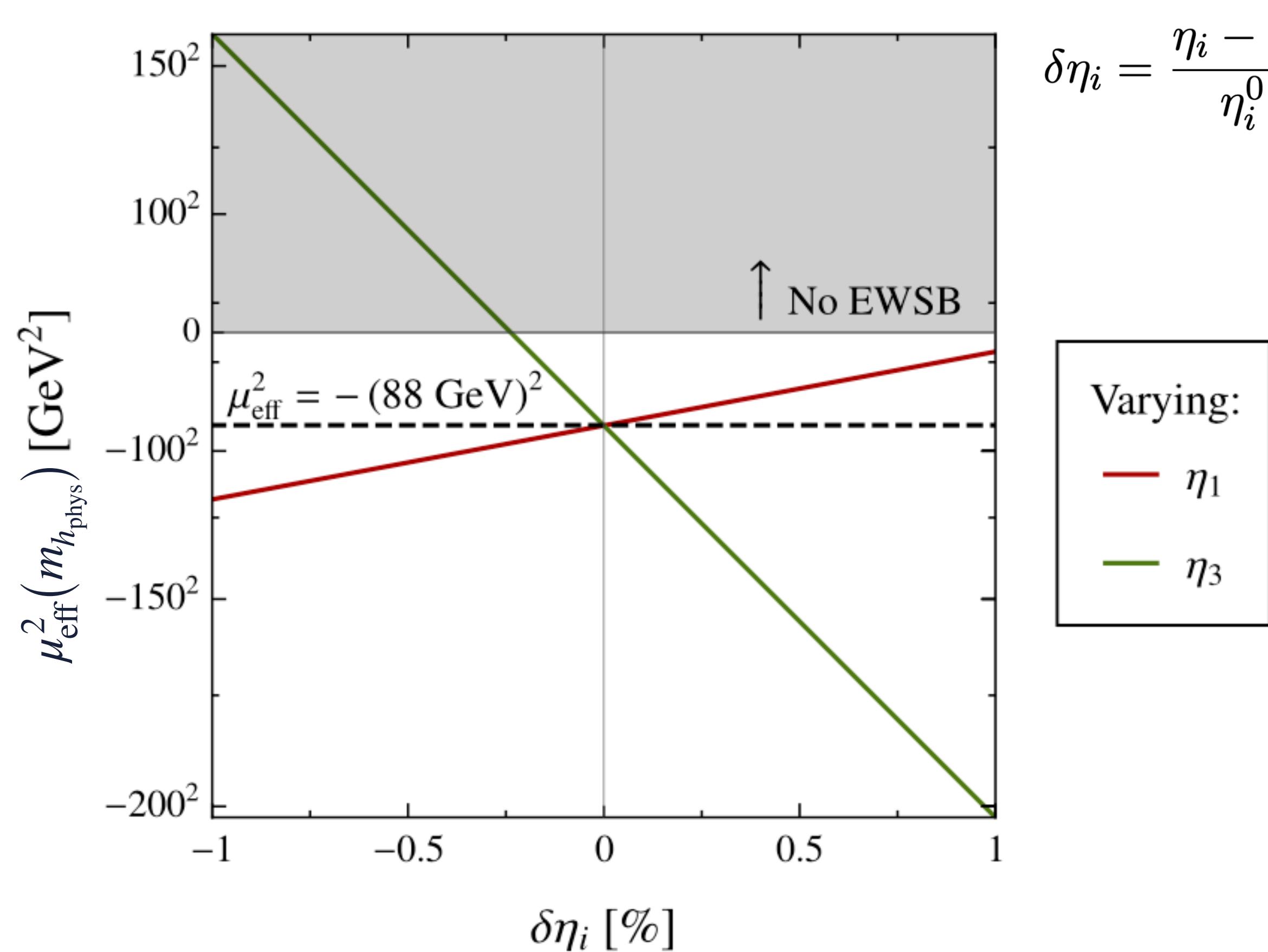
- ✓ perturbativity
- ✓ bounded from below conditions

Some observations:

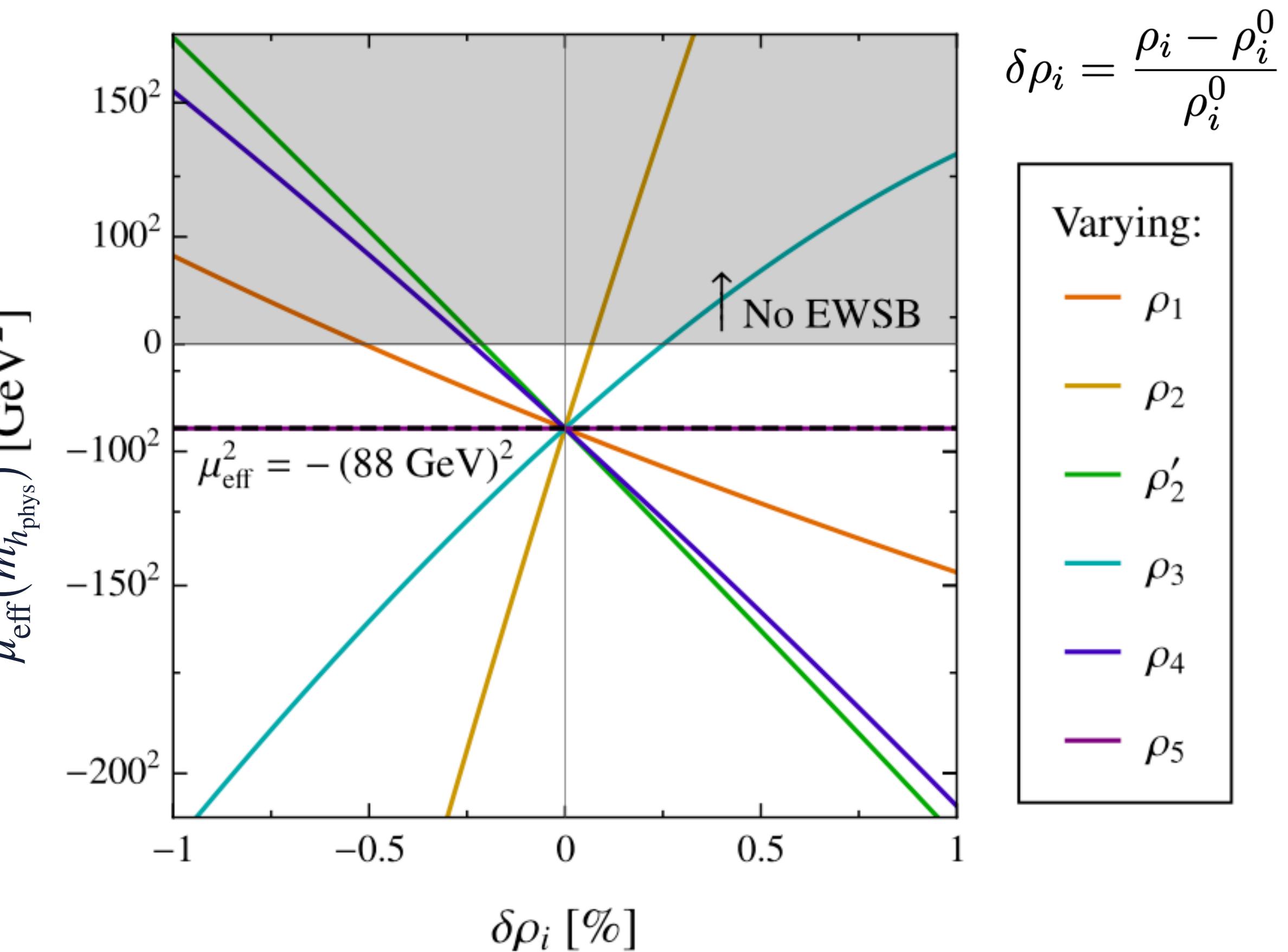
- ✓ undeniably an RG effect
- ↪ without running, positive eigenvalues of the hessian stay positive.
- ✗ unavoidable fine-tuning
- ↪ strong acceleration near the 4321 breaking scale

# Fine-tuning

Quantifying the fine-tuning and identifying sensitivity of  $m_{h_{\text{phys}}}$  to variation of the quartic by  $\pm 1\%$



$$\delta\eta_i = \frac{\eta_i - \eta_i^0}{\eta_i^0}$$



$$\delta\rho_i = \frac{\rho_i - \rho_i^0}{\rho_i^0}$$

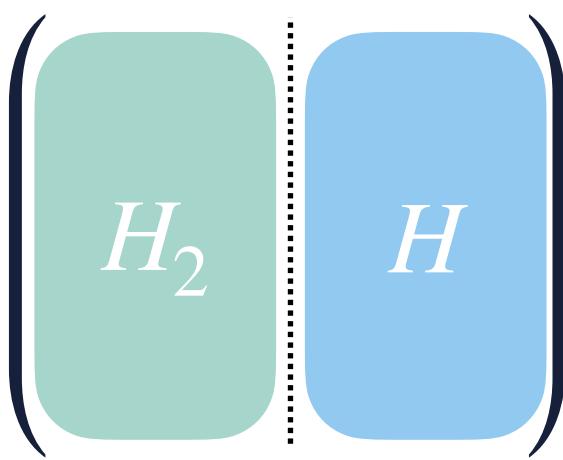
### 3. Unification in the UV

In some UV completion of 4321 see Ben's talk 

[Fuentes-Martín, Isidori, Lizana, Selimovic, Stefanek, 2203.01952]

[Bordone, Cornella, Fuentes-Martín, Isidori, 1712.01368]

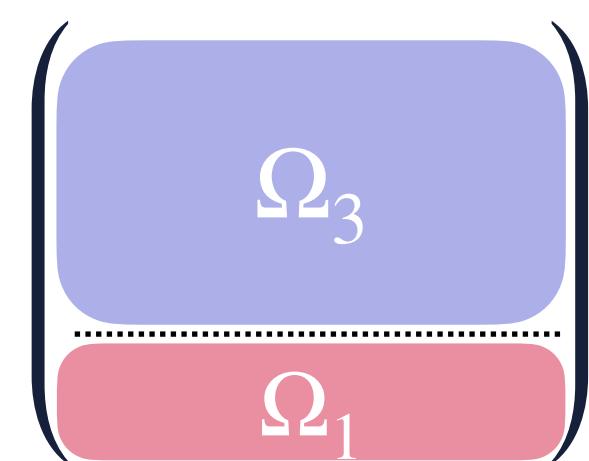
- the doublet Higgs comes from a bi-doublet field under  $SU(2)_L$  and  $SU(2)_R$   $\Phi \sim$



leading to a two-Higgs-doublet-model after 4321 breaking scale with relations:

$$\mu_H = \mu_2 \quad \text{and} \quad \lambda = \lambda_2 = \lambda_3$$

- the two  $\Omega_{1,3}$  come from the same “bi-quadruplet” under  $SU(4)_{\text{light}}$  and  $SU(4)_3$   $\Omega \sim$



leading to the relations after SSB:

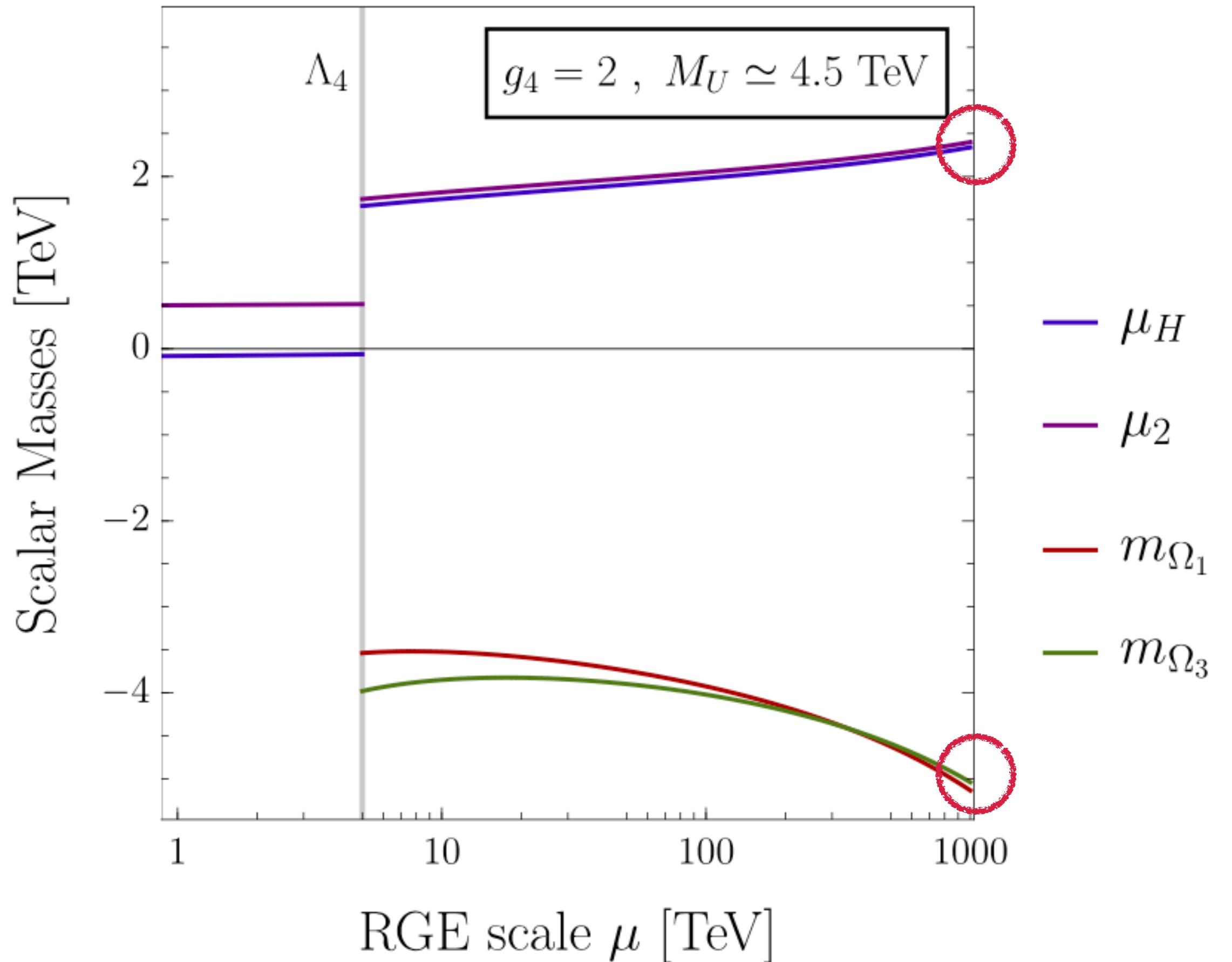
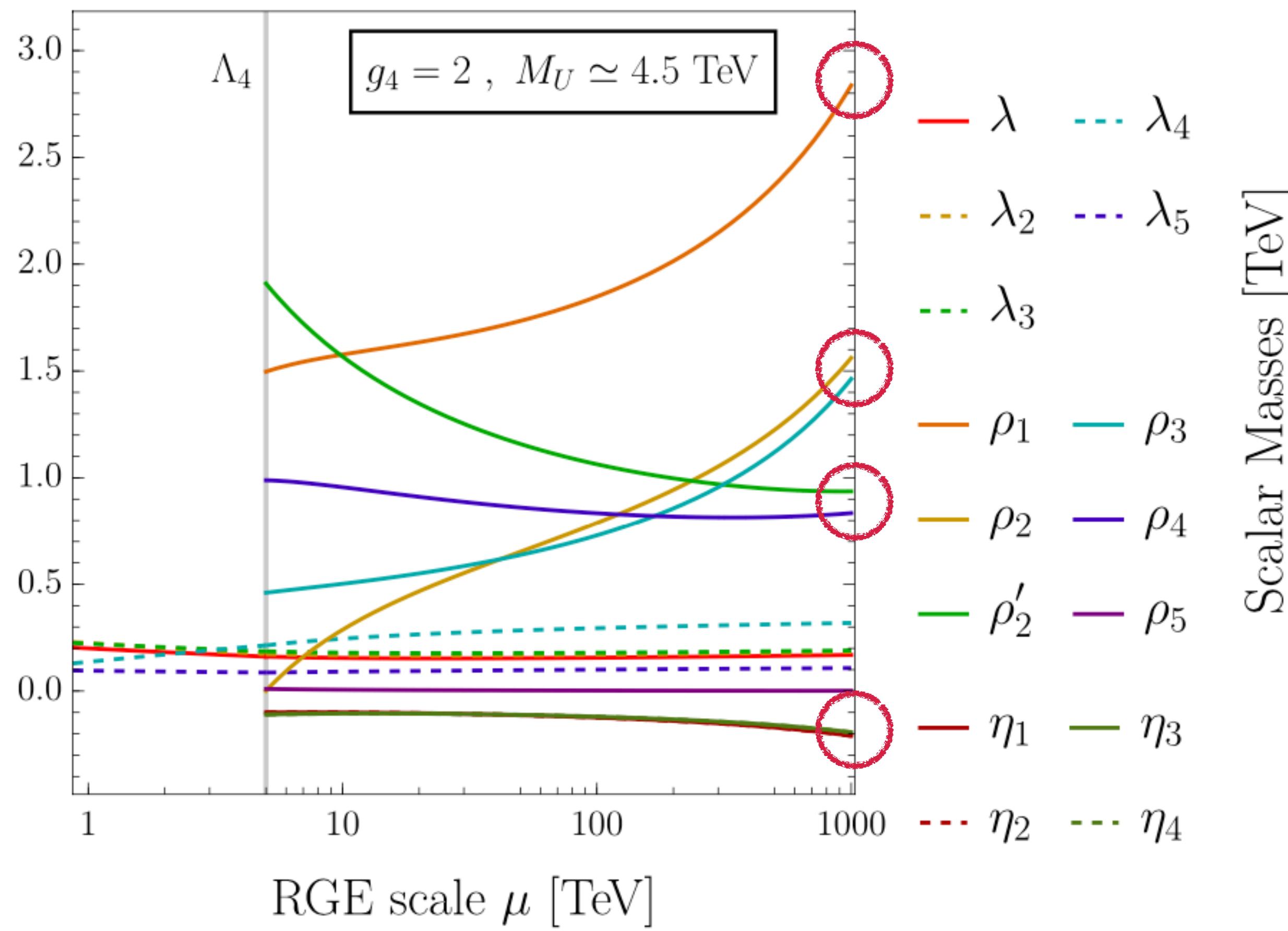
$$m_{\Omega_1} = m_{\Omega_3}, \quad \rho_1 = \rho_2 + \rho'_2, \quad \rho_2 = \rho_3 \quad \text{and} \quad \rho'_2 = \rho_4$$

$$\eta_1 = \eta_2 = \eta_3 = \eta_4$$

# UV unification

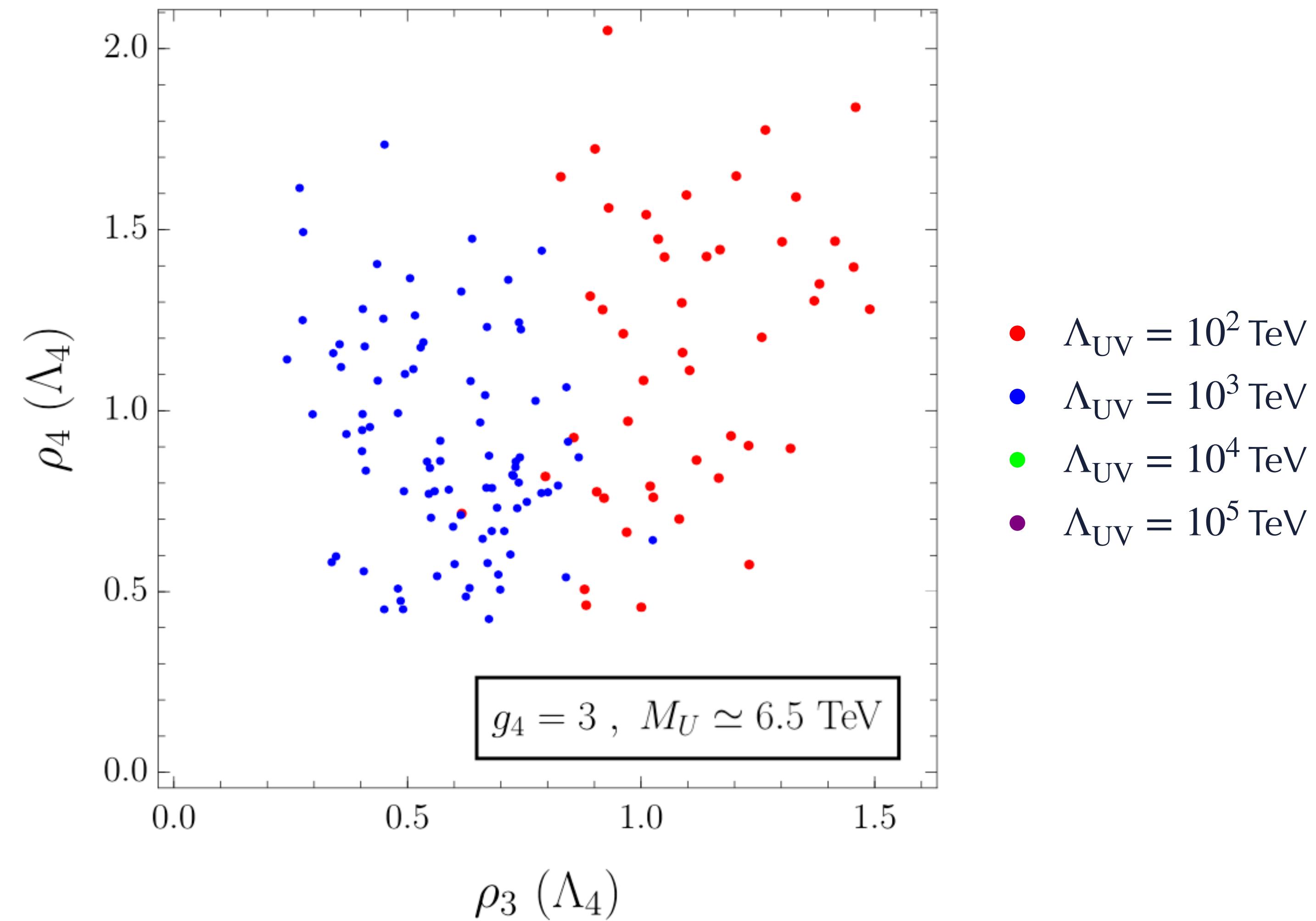
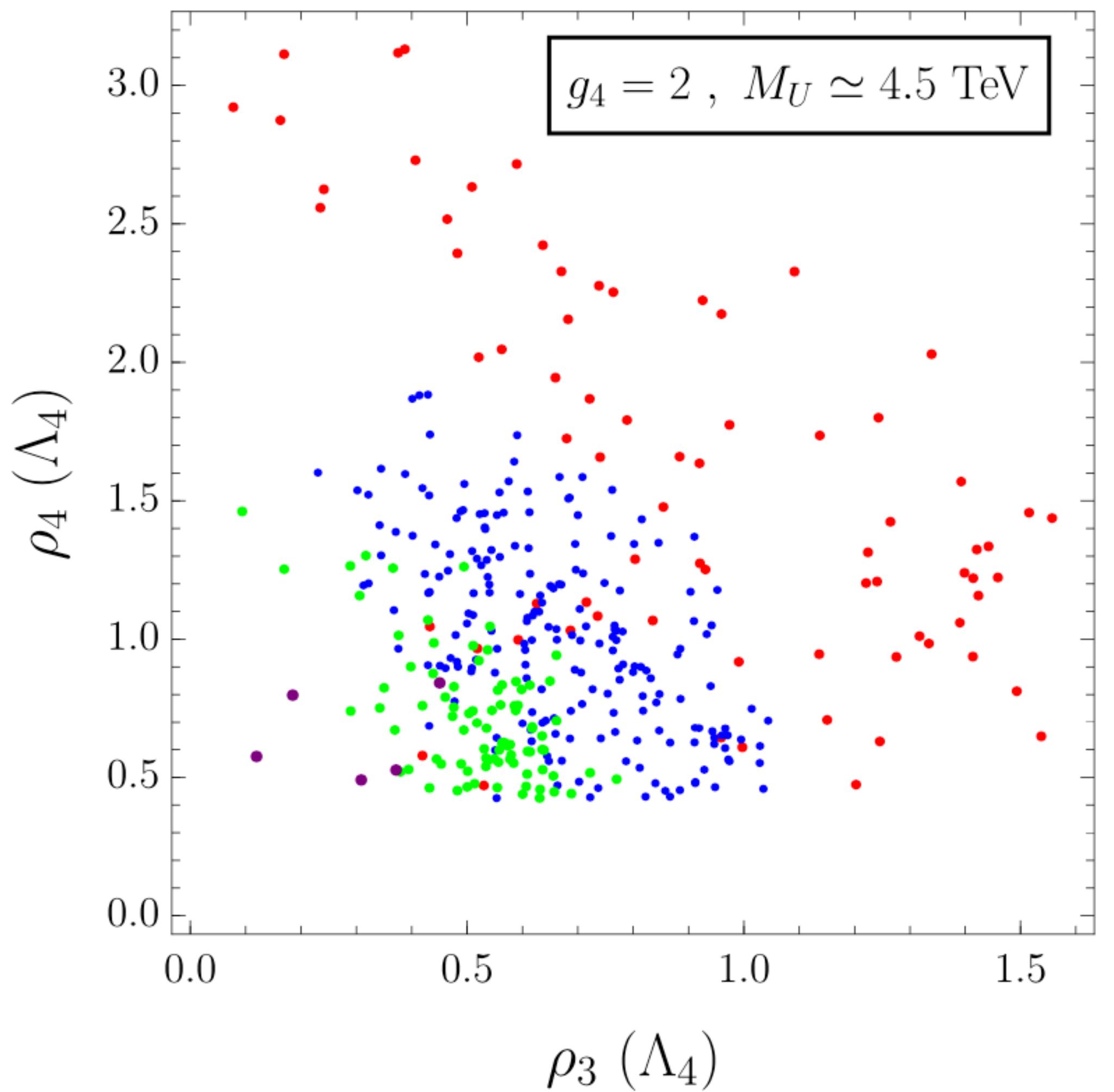
Unification in the UV can be realised allowing for 10 % threshold corrections at  $\sim 1000$  TeV

$$\begin{aligned} \lambda &= \lambda_2 = \lambda_3 \\ \eta_1 &= \eta_2 = \eta_3 = \eta_4 \end{aligned} \quad \begin{aligned} \rho_1 &= \rho_2 + \rho'_2 \\ \rho_2 &= \rho_3 \\ \rho'_2 &= \rho_4 \end{aligned} \quad \begin{aligned} \mu_H &= \mu_2 \\ m_{\Omega_1} &= m_{\Omega_3} \end{aligned}$$



# UV unification scale

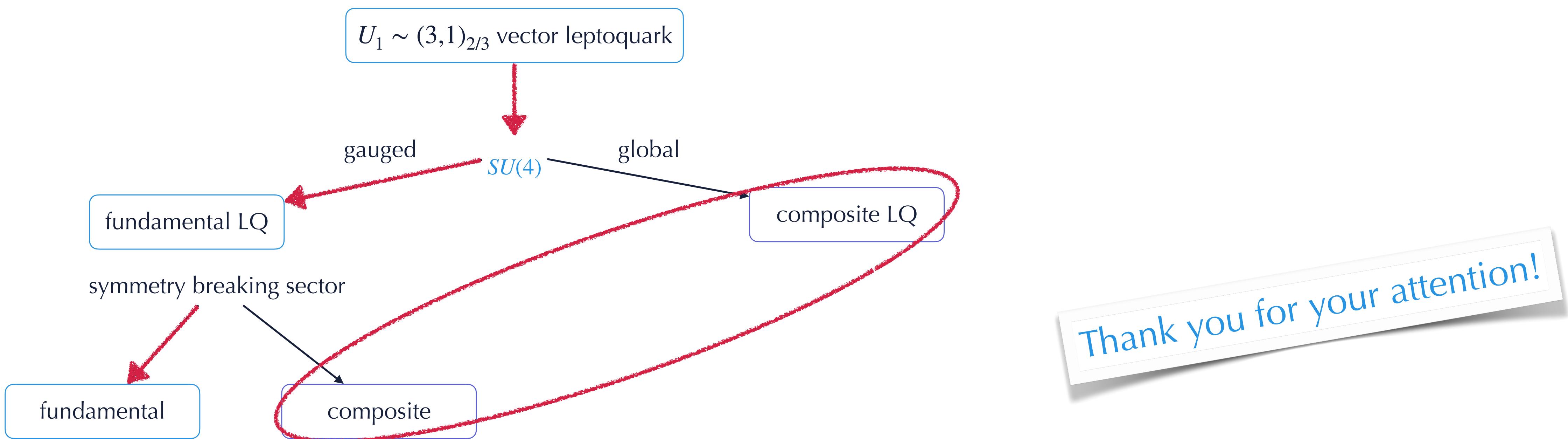
Number of benchmark points satisfying the UV conditions per energy scale  
↔ unification scale between  $10^2 - 10^4$  TeV



# Conclusion

The RGE of the 4321 scalar sector revealed *intriguing* features:

- ▶ Laudau poles appear as early as 100 TeV for heavy scalar radial modes
- ▶ Radiative EWSB can happen but fine-tuning seems ineluctable
- ▶ Couplings can unify in the UV between  $10^2 - 10^4$  TeV



# Back-up

# $U_1$ , $G'$ and $Z'$ masses and couplings

Gauge bosons masses

$$M_U = \frac{g_4}{2} \sqrt{\omega_1^2 + \omega_3^2}$$

$$M_{G'} = \sqrt{\frac{g_4^2 + g_3^2}{2}} \omega_3$$

$$M_{Z'} = \frac{1}{2\sqrt{6}} \sqrt{(3g_4^2 + 2g_1^2) (3\omega_1^2 + \omega_3^2)}$$

Gauge bosons couplings

$$g_U = g_4$$

$$g_Y = \frac{g_1 g_4}{\sqrt{g_4^2 + \frac{2}{3}g_1^2}}$$

$$g_{G'} = \sqrt{g_4^2 - g_s^2}$$

$$g_s = \frac{g_4 g_3}{\sqrt{g_4^2 + g_3^2}}$$

$$g_{Z'} = \frac{1}{2\sqrt{6}} \sqrt{g_4^2 - \frac{2}{3}g_Y^2}$$

# Radial modes spectrum

$\Omega_{1,3}$  radials

$$\Omega_1^\dagger = \begin{pmatrix} \tilde{T}_1 \\ \frac{\omega_1}{\sqrt{2}} + \tilde{S}_1^* \end{pmatrix} \quad \Omega_3^\dagger = \begin{pmatrix} \left( \frac{\omega_3}{\sqrt{2}} + \frac{\tilde{S}_3^*}{\sqrt{3}} \right) \mathbb{1}_3 + \tilde{O}_3^{a*} t^a \\ \tilde{T}_3^\dagger \end{pmatrix}$$

- $M_{O_R}^2 = \omega_3(\rho'_2\omega_3 - \rho_5\omega_1)$  with  $\rho'_2\omega_3 > \rho_5\omega_1$
- $M_{T_R}^2 = \frac{1}{2} \left( \rho_4 - \rho_5 \frac{\omega_3}{\omega_1} \right) (\omega_1^2 + \omega_3^2)$  with  $\rho_4\omega_1 > \rho_5\omega_3$
- $M_{S_0}^2 = \frac{\rho_5}{2} \frac{\omega_3}{\omega_1} (3\omega_1^2 + \omega_3^2)$  with  $\rho_5 > 0$
- $M_{S_1}^2 = \frac{1}{2} \left( \rho_1\omega_1^2 + (3\rho_2 + \rho'_2)\omega_3^2 + \frac{\rho_5}{2} \frac{\omega_3}{\omega_1} (\omega_1^2 - \omega_3^2) \right) - \frac{u}{2\omega_1}$
- $M_{S_2}^2 = \frac{1}{2} \left( \rho_1\omega_1^2 + (3\rho_2 + \rho'_2)\omega_3^2 + \frac{\rho_5}{2} \frac{\omega_3}{\omega_1} (\omega_1^2 - \omega_3^2) \right) + \frac{u}{2\omega_1}$

$$u^2 = \left[ \rho_1\omega_1^3 + (3\rho_2 + \rho'_2)\omega_1\omega_3^2 + \frac{\rho_5}{2}\omega_3(\omega_1^2 - \omega_3^2) \right]^2 - 4\omega_1\omega_3 \left[ \rho_1\omega_1^3 \left( \frac{\rho_5}{2}\omega_1 + (3\rho_2 + \rho'_2)\omega_3 \right) - 3\omega_3 \left( \rho_3^2\omega_1^3 + \rho_3\rho_5\omega_3\omega_1^2 + \frac{\rho_5^2}{3}\omega_3^2\omega_1 + \frac{\rho_5}{6}(3\rho_2 + \rho'_2)\omega_3^3 \right) \right].$$

$H, H_2$  radials

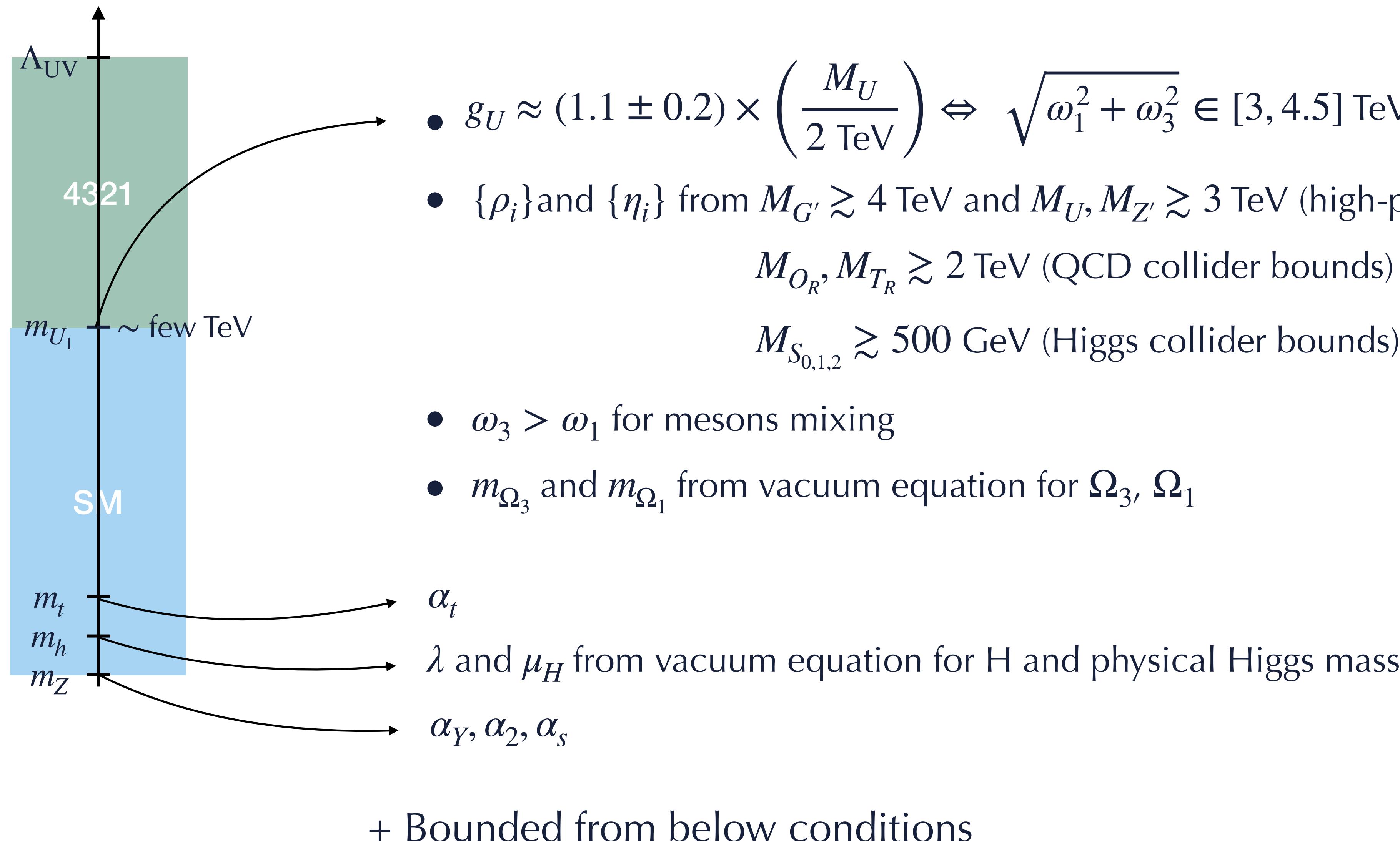
$$H = \begin{pmatrix} \eta_W^+ \\ \frac{v+h}{\sqrt{2}} + i\eta_Z \end{pmatrix} \quad H_2 = \begin{pmatrix} h^+ \\ \frac{1}{\sqrt{2}}(h_R + i h_I) \end{pmatrix}$$

$$\begin{aligned} m_h^2 &= \lambda v^2, \\ m_\pm^2 &= \mu_2^2 + \frac{\lambda_3}{2} v^2, \\ m_R^2 &= \mu_2^2 + \frac{\lambda_3 + \lambda_4 + \lambda_5}{2} v^2 \\ m_I^2 &= \mu_2^2 + \frac{\lambda_3 + \lambda_4 - \lambda_5}{2} v^2 \end{aligned}$$

$$\begin{aligned} V_H &\supset \mu_2^2 H_2^\dagger H_2 + \frac{\lambda_2}{2} (H_2^\dagger H_2)^2 + \lambda_3 (H^\dagger H)(H_2^\dagger H_2) \\ &\quad + \lambda_4 (H^\dagger H_2)(H_2^\dagger H) + \frac{\lambda_5}{2} ((H^\dagger H_2)^2 + \text{h.c}) \end{aligned}$$

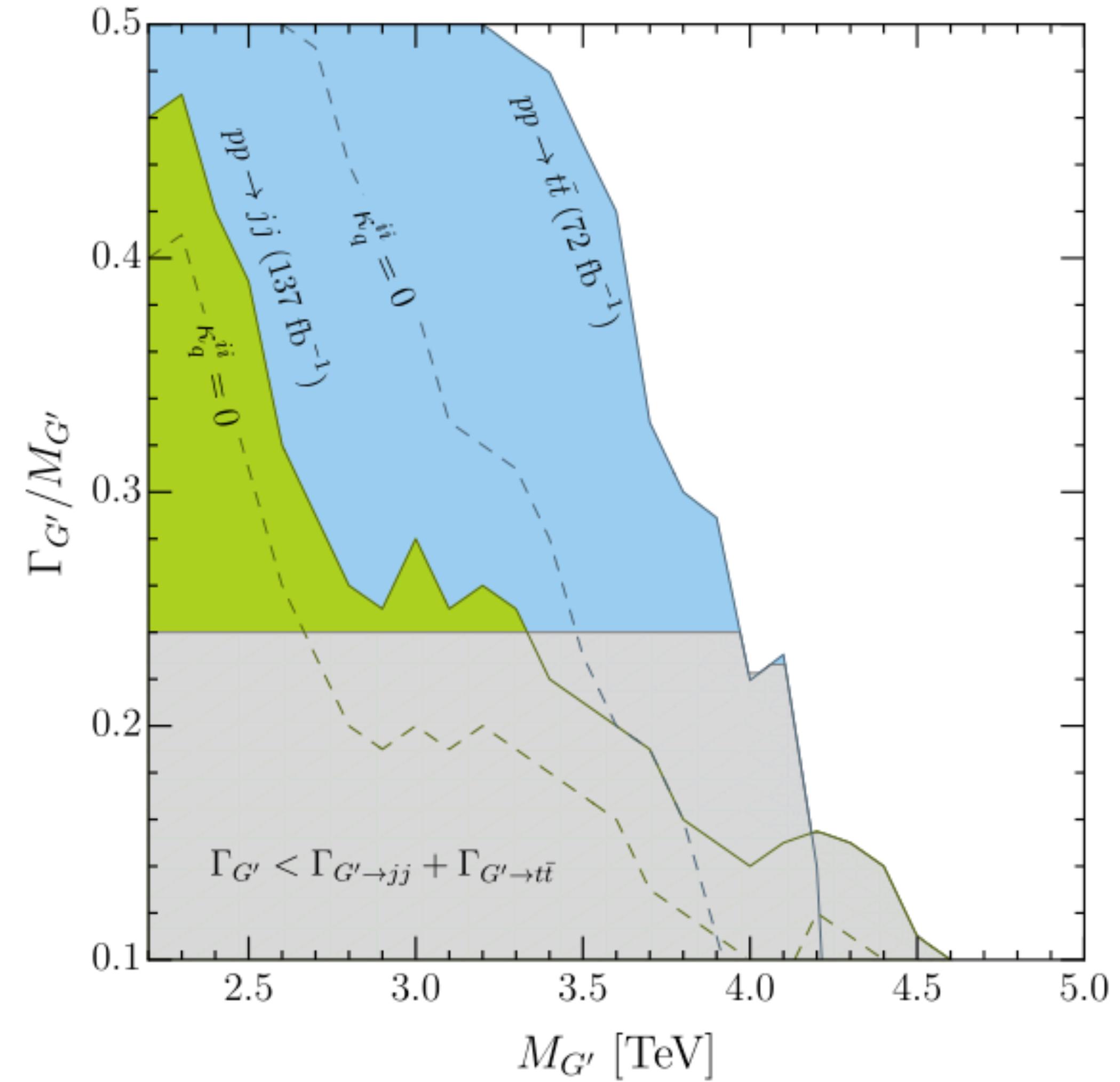
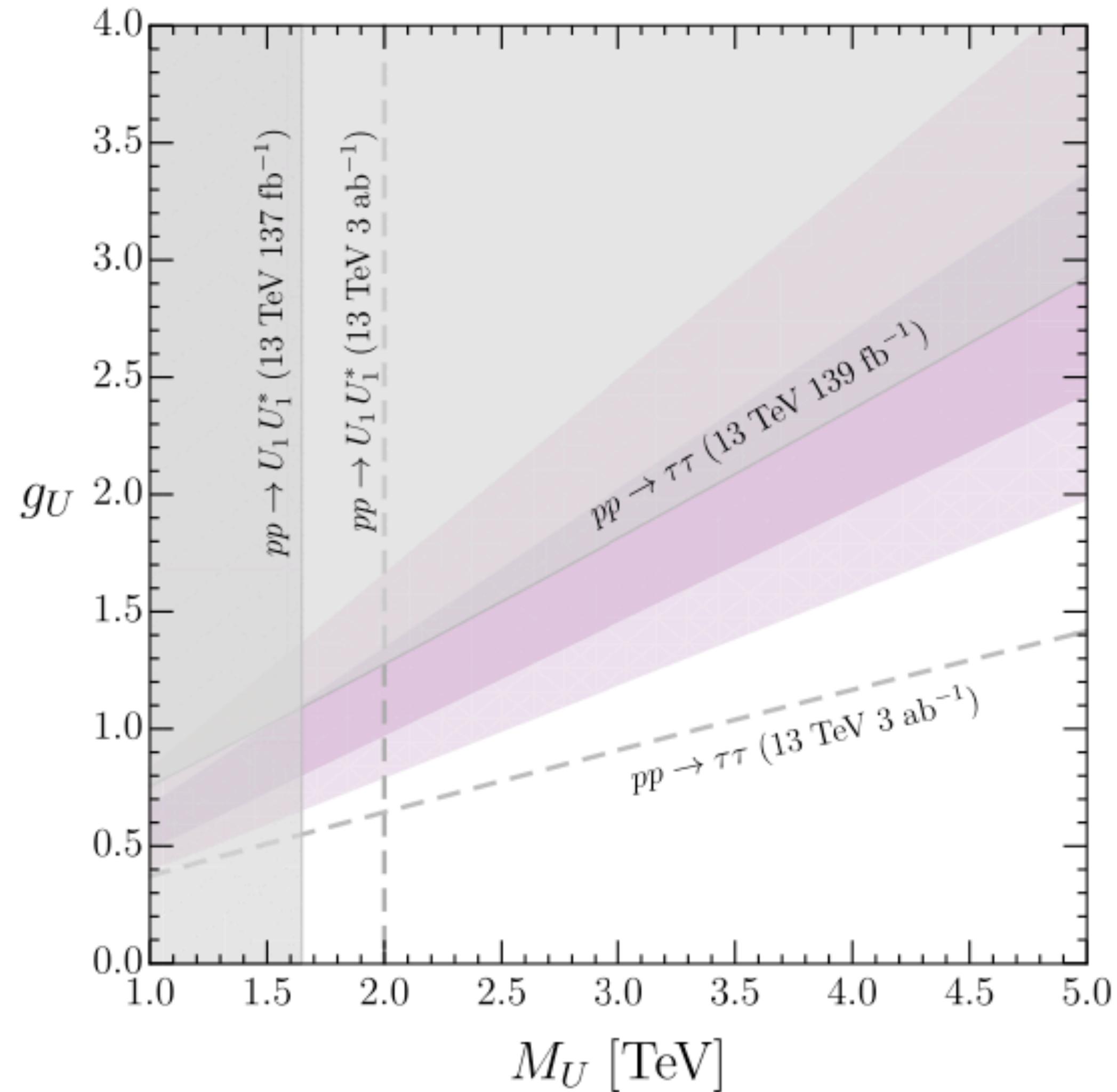
$$V_{\Omega H} \supset \eta_2 H_2^\dagger H_2 \Omega_1^\dagger \Omega_1 + \eta_4 H_2^\dagger H_2 \text{Tr}[\Omega_3^\dagger \Omega_3]$$

# Boundary conditions



# $U_1, G'$ searches

[Cornella, Faroughy, Fuentes-Martín, Isidori, Neubert, 2103.16558]



# Benchmark points

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| $\Omega$ VEVs    | values                 | scalar couplings | values |
|------------------|------------------------|------------------|--------|
| $\omega_1$       | 1.5 TeV                | $\lambda$        | 0.2    |
| $\omega_3$       | 4 TeV                  | $\rho_1$         | 0.5    |
| mass parameters  |                        | $\rho_2$         | 0.1    |
| $\mu_H^2$        | $(1.6 \text{ TeV})^2$  | $\rho'_2$        | 0.5    |
| $m_{\Omega_1}^2$ | $-(1.8 \text{ TeV})^2$ | $\rho_3$         | 0.1    |
| $m_{\Omega_3}^2$ | $-(2.5 \text{ TeV})^2$ | $\rho_4$         | 1      |
| gauge coupling   |                        | $\rho_5$         | 0.01   |
| $g_4$            | 3                      | $\eta_1$         | -0.1   |
|                  |                        | $\eta_3$         | -0.1   |

# Beta-functions - Gauge and Higgs potential

Gauge couplings:

$$\mu \frac{d\alpha_i}{d\mu} \equiv \frac{1}{16\pi^2} \beta_{\alpha_i} = -\frac{\alpha_i^2}{2\pi} b_i$$

$$U(1)': b_1 = -\frac{115}{18}, \quad SU(2)_L: b_2 = \frac{19}{6} - \frac{8}{3}n_{VL}$$

$$SU(3)_{1+2}: b_3 = \frac{23}{3}, \quad SU(4)_3: b_4 = \frac{38}{3} - \frac{4}{3}n_{VL}$$

Top yukawa:

$$\beta_{y_t} = \left( \frac{11}{2}y_t^2 - \frac{9}{4}g_2^2 - \frac{45}{4}g_4^2 - \frac{3}{4}g_1^2 \right) y_t$$

Higgs quartics:

$$\begin{aligned} \beta_\lambda &= (12\lambda + 16y_t^2 - 3g_1^2 - 9g_2^2) \lambda + 8\eta_1^2 + 24\eta_3^2 - 16y_t^4 + 4\lambda_3^2 + 4\lambda_3\lambda_4 + 2\lambda_4^2 + 2\lambda_5^2 \\ &\quad + \frac{3}{4}g_1^4 + \frac{3}{2}g_1^2g_2^2 + \frac{9}{4}g_2^4 \\ \beta_{\lambda_2} &= (12\lambda_2 + 16y_t^2 - 3g_1^2 - 9g_2^2) \lambda_2 + 8\eta_2^2 + 24\eta_4^2 - 16y_t^4 + 4\lambda_3^2 + 4\lambda_3\lambda_4 + 2\lambda_4^2 + 2\lambda_5^2 \\ &\quad + \frac{3}{4}g_1^4 + \frac{3}{2}g_1^2g_2^2 + \frac{9}{4}g_2^4 \\ \beta_{\lambda_3} &= (4\lambda_3 + 6\lambda + 6\lambda_2 - 3g_1^2 - 9g_2^2 + 16y_t^2) \lambda_3 + 2\lambda_4(\lambda + \lambda_2) + 2\lambda_4^2 + 2\lambda_5^2 - 16y_t^4 \\ &\quad + 8\eta_1\eta_2 + 24\eta_3\eta_4 + \frac{3}{4}g_1^4 - \frac{3}{2}g_1^2g_2^2 + \frac{9}{4}g_2^4 \\ \beta_{\lambda_4} &= (4\lambda_4 + 2\lambda + 2\lambda_2 + 8\lambda_3 - 3g_1^2 - 9g_2^2 + 16y_t^2) \lambda_4 + 8\lambda_5^2 + 3g_1^2g_2^2 + 16y_t^4 \\ \beta_{\lambda_5} &= (2\lambda + 2\lambda_2 + 8\lambda_3 + 12\lambda_4 + 16y_t^2 - 9g_2^2 - 3g_1^2) \lambda_5 \end{aligned}$$

Higgs masses:

$$\beta_{\mu_H^2} = \left( 6\lambda + 8y_t^2 - \frac{3}{2}g_1^2 - \frac{9}{2}g_2^2 \right) \mu_H^2 + 8\eta_1 m_{\Omega_1}^2 + 24\eta_3 m_{\Omega_3}^2 + (4\lambda_3 + 2\lambda_4) \mu_2^2$$

$$\beta_{\mu_2^2} = \left( 6\lambda_2 + 8y_t^2 - \frac{3}{2}g_1^2 - \frac{9}{2}g_2^2 \right) \mu_2^2 + (4\lambda_3 + 2\lambda_4) \mu_H^2 + 8\eta_2 m_{\Omega_1}^2 + 24\eta_4 m_{\Omega_3}^2$$

# Beta-functions - $\Omega$ potential

$\Omega$  quartics:

$$\begin{aligned}\beta_{\rho_1} &= \left(16\rho_1 - 3g_1^2 - \frac{45}{2}g_4^2\right)\rho_1 + 24\rho_3^2 + 12\rho_3\rho_4 + 6\rho_4^2 + 4\eta_1^2 + 4\eta_2^2 \\ &\quad + \frac{3}{4}g_1^4 + \frac{9}{4}g_1^2g_4^2 + \frac{99}{16}g_4^4 \\ \beta_{\rho_2} &= \left(32\rho_2 + 28\rho'_2 - \frac{1}{3}g_1^2 - 16g_3^2 - \frac{45}{2}g_4^2\right)\rho_2 + 6\rho'^2_2 + 8\rho_3^2 + 4\rho_3\rho_4 + 4\rho_5^2 \\ &\quad + 4\eta_3^2 + 4\eta_4^2 + \frac{1}{108}g_1^4 + \frac{11}{6}g_3^4 + \frac{27}{16}g_4^4 - \frac{1}{9}g_1^2g_3^2 - \frac{1}{12}g_1^2g_4^4 + \frac{13}{2}g_3^2g_4^2 \\ \beta_{\rho'_2} &= \left(14\rho'_2 + 12\rho_2 - \frac{1}{3}g_1^2 - 16g_3^2 - \frac{45}{2}g_4^2\right)\rho'_2 + 2\rho_4^2 - 4\rho_5^2 \\ &\quad + \frac{5}{2}g_3^4 + \frac{9}{2}g_4^4 + \frac{1}{3}g_1^2g_3^2 + \frac{1}{3}g_1^2g_4^2 - \frac{7}{2}g_3^2g_4^2 \\ \beta_{\rho_3} &= \left(4\rho_3 + 10\rho_1 + 26\rho_2 + 14\rho'_2 - \frac{5}{3}g_1^2 - 8g_3^2 - \frac{45}{2}g_4^2\right)\rho_3 + 2\rho_4^2 + 4\rho_5^2 \\ &\quad + 2(\rho_1 + 3\rho_2 + \rho'_2)\rho_4 + 4\eta_1\eta_3 + 4\eta_2\eta_4 + \frac{1}{12}g_1^4 + \frac{1}{4}g_1^2g_4^2 + \frac{27}{16}g_4^4 \\ \beta_{\rho_4} &= \left(8\rho_4 + 2\rho_1 + 2\rho_2 + 6\rho'_2 + 8\rho_3 - \frac{5}{3}g_1^2 - 8g_3^2 - \frac{45}{2}g_4^2\right)\rho_4 - 4\rho_5^2 \\ &\quad - g_1^2g_4^2 + \frac{9}{2}g_4^4 \\ \beta_{\rho_5} &= \left[6(\rho_2 - \rho'_2 + \rho_3 - \rho_4) - g_1^2 - 12g_3^2 - \frac{45}{2}g_4^2\right]\rho_5\end{aligned}$$

$\Omega$  masses:

$$\begin{aligned}\beta_{m_{\Omega_1}^2} &= \left(10\rho_1 - \frac{3}{2}g_1^2 - \frac{45}{4}g_4^2\right)m_{\Omega_1}^2 + (24\rho_3 + 6\rho_4)m_{\Omega_3}^2 + 4\eta_1\mu_H^2 + 4\eta_2\mu_2^2 \\ \beta_{m_{\Omega_3}^2} &= \left(26\rho_2 + 14\rho'_2 - \frac{1}{6}g_1^2 - 8g_3^2 - \frac{45}{4}g_4^2\right)m_{\Omega_3}^2 + 4\eta_3\mu_H^2 + (8\rho_3 + 2\rho_4)m_{\Omega_1}^2 + 4\eta_4\mu_2^2\end{aligned}$$

Mixing quartics:

$$\begin{aligned}\beta_{\eta_1} &= \left(4\eta_1 + 6\lambda + 10\rho_1 - 3g_1^2 - \frac{9}{2}g_2^2 - \frac{45}{4}g_4^2 + 8y_t^2\right)\eta_1 \\ &\quad + 4\lambda_3\eta_2 + 2\lambda_4\eta_2 + 24\rho_3\eta_3 + 6\rho_4\eta_3 + \frac{3}{4}g_1^4 \\ \beta_{\eta_2} &= \left(4\eta_2 + 6\lambda_2 + 10\rho_1 - 3g_1^2 - \frac{9}{2}g_2^2 - \frac{45}{4}g_4^2 + 8y_t^2\right)\eta_2 \\ &\quad + 4\lambda_3\eta_1 + 2\lambda_4\eta_1 + 24\rho_3\eta_4 + 6\rho_4\eta_4 + \frac{3}{4}g_1^4 \\ \beta_{\eta_3} &= \left(6\lambda + 26\rho_2 + 14\rho'_2 - \frac{5}{3}g_1^2 - \frac{9}{2}g_2^2 - 8g_3^2 - \frac{45}{4}g_4^2 + 8y_t^2\right)\eta_3 \\ &\quad + (4\lambda_3 + 2\lambda_4)\eta_4 + (8\rho_3 + 2\rho_4)\eta_1 + 4\eta_3^2 + \frac{1}{12}g_1^4 \\ \beta_{\eta_4} &= 6\lambda_2\eta_4 + (4\lambda_3 + 2\lambda_4)\eta_3 + 2(13\rho_2 + 7\rho'_2)\eta_4 + (8\rho_3 + 2\rho_4)\eta_2 + 4\eta_4^2 \\ &\quad - \left(\frac{5}{3}g_1^2 + \frac{9}{2}g_2^2 + 8g_3^2 + \frac{45}{4}g_4^2\right)\eta_4 + \frac{1}{12}g_1^4 + 8\eta_4 Y_{12}\end{aligned}$$

# Bounded from below conditions

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$V_\Omega$  conditions

$$\begin{aligned}\rho_1 &> 0 , \\ \rho_2 + \rho'_2 &> 0 , \quad 3\rho_2 + \rho'_2 > 0 , \\ \rho_3 &> -\sqrt{\rho_1(\rho_2 + \rho'_2)} , \quad \rho_3 + \rho_4 > -\sqrt{\rho_1(\rho_2 + \rho'_2)} \\ |\rho_5| &< \frac{1}{4}(\rho_1 + 3(3\rho_2 + \rho'_2) + 6\rho_3) ,\end{aligned}$$

$V_{\Omega H}$  conditions

$$\begin{aligned}\eta_1 &> -\sqrt{\lambda\rho_1} , \\ \eta_3 &> -\min \left[ \sqrt{\lambda(\rho_2 + \rho'_2)}, \sqrt{\lambda(3\rho_2 + \rho'_2)} \right] \\ \eta_2 &> -\sqrt{\lambda_2\rho_1} \\ \eta_4 &> -\min \left[ \sqrt{\lambda_2(\rho_2 + \rho'_2)}, \sqrt{\lambda_2(3\rho_2 + \rho'_2)} \right]\end{aligned}$$

$V_{2\text{HDM}}$  conditions

$$\lambda > 0 , \quad \lambda_2 > 0 , \quad \lambda_3 > -\sqrt{\lambda\lambda_2} , \quad \lambda_3 + \lambda_4 - |\lambda_5| > -\sqrt{\lambda\lambda_2}$$

# Effective quantities for EWSB

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Potential minimization

$$\frac{\partial V(s_1, s_3, h)}{\partial h_i} \Big|_{\langle h \rangle, \langle s_1 \rangle, \langle s_3 \rangle} = 0 \quad \text{for } h_i = h, s_1, s_3$$

$$[M_{s_3, s_1, h}^2]_{ij} = \frac{\partial^2 V}{\partial h_i \partial h_j} \text{ with } h_i \in (s_3, s_1, h)$$

positive-definite:

$$M_{s_3, s_1, h}^2 = \begin{pmatrix} 2(3\rho_2 + \rho'_2)\omega_3^2 & 2\sqrt{3}\rho_3\omega_1\omega_3 & \sqrt{6}\eta_3 v \omega_3 \\ 2\sqrt{3}\rho_3\omega_1\omega_3 & 2\rho_1\omega_1^2 & \sqrt{2}\eta_1 v \omega_1 \\ \sqrt{6}\eta_3 v \omega_3 & \sqrt{2}\eta_1 v \omega_1 & \lambda v^2 \end{pmatrix}$$

$$D_1 = 3\rho_2 + \rho'_2 > 0$$

$$D_{12} = \rho_1 D_1 - 3\rho_3^2 > 0$$

$$D_{123} = \lambda D_{12} - (3\eta_3^2 \rho_1 - 6\eta_1 \eta_3 \rho_3 + \eta_1^2 D_1) > 0$$

Effective mass and quartic coupling:

$$\mu_{\text{eff}}^2 = -\frac{\lambda_{\text{eff}}}{2}v^2 = \mu_H^2 - \frac{\eta_1 D_1 - 3\eta_3 \rho_3}{D_{12}} m_{\Omega_1}^2 - 3 \frac{\eta_3 \rho_1 - \eta_1 \rho_3}{D_{12}} m_{\Omega_3}^2$$

$$m_{h_{\text{phys}}}^2 = \lambda_{\text{eff}} v^2 = \frac{D_{123}}{D_{12}} v^2 + O\left(\frac{v^2}{\omega_{1,3}^2}\right)$$