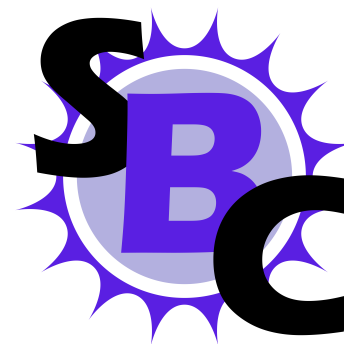


Coherent Elastic Neutrino-Nucleus Scattering in Reactors Using a Liquid Argon Scintillating Bubble Chamber



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Instituto de Física, UNAM

Precision Tests with Neutral-Current Coherent Interactions with Nuclei
Mainz, Germany, May 27th, 2022

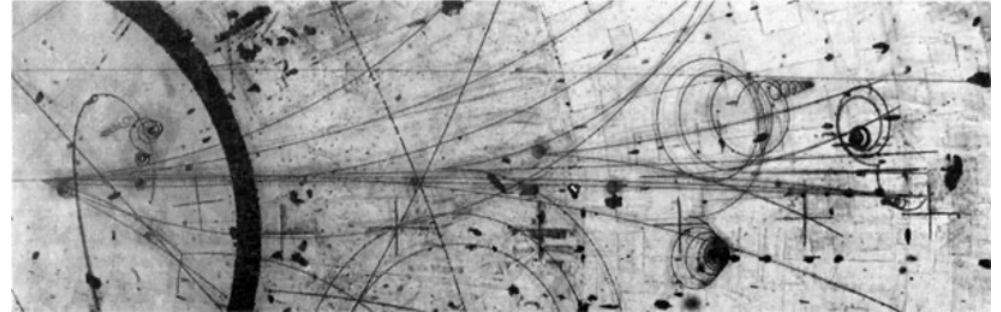
Outline

- Bubble chambers
- SBC: A 10 kg LAr bubble chamber for dark matter and $\text{CE}\nu\text{NS}$
- Physics reach for reactor neutrinos
- Final remarks

Physics with bubble chambers

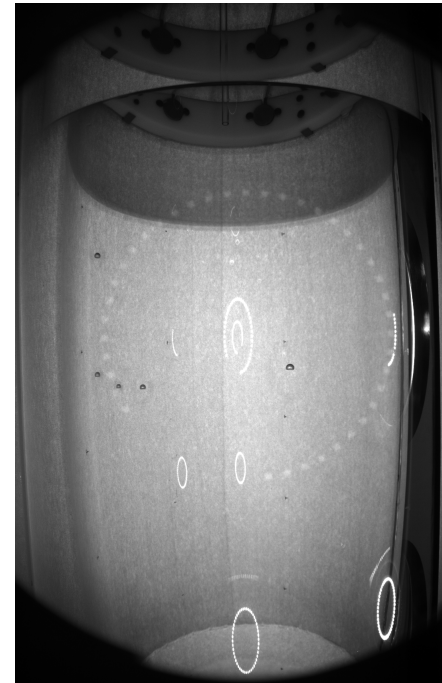
1970s: Neutrino Beam Physics

- Sensitive to MIPs
- Particle tracks visible
- Threshold $\ll 1$ keV
- Multi-ton chambers, multiple fluids



2000-today: Nuclear Recoil Detectors

- Dark matter searches with fluorocarbon bubble chambers
- Electron recoil blind
- Nuclear recoil threshold ~ 3 keV
- Scalable at modest cost

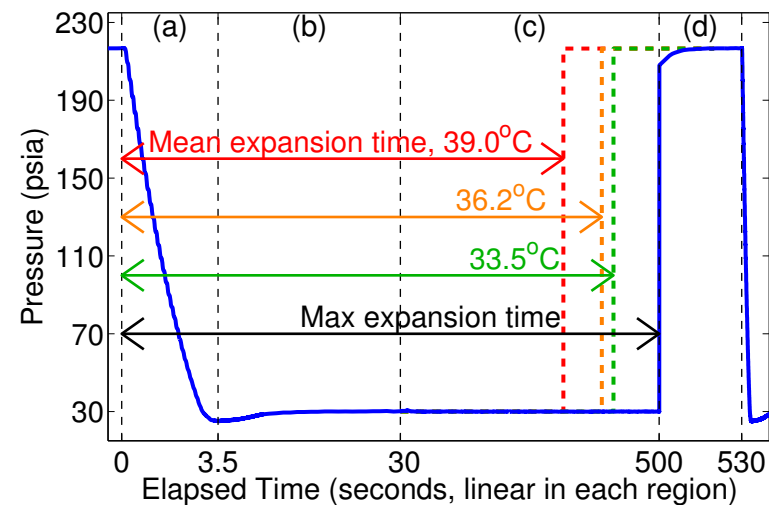
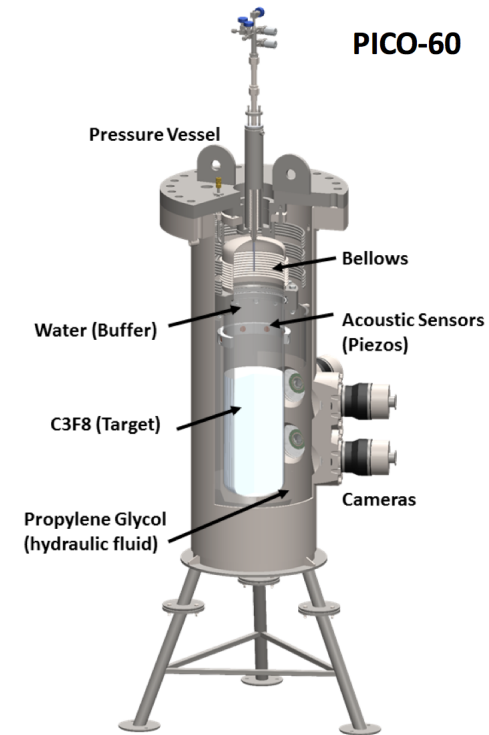


PICO bubble chambers

- Target material:
superheated CF_3I ,
 C_3F_8 , C_4F_{10}
spin-dependent/independent

Could make a
dark matter bubble
chamber with any liquid!

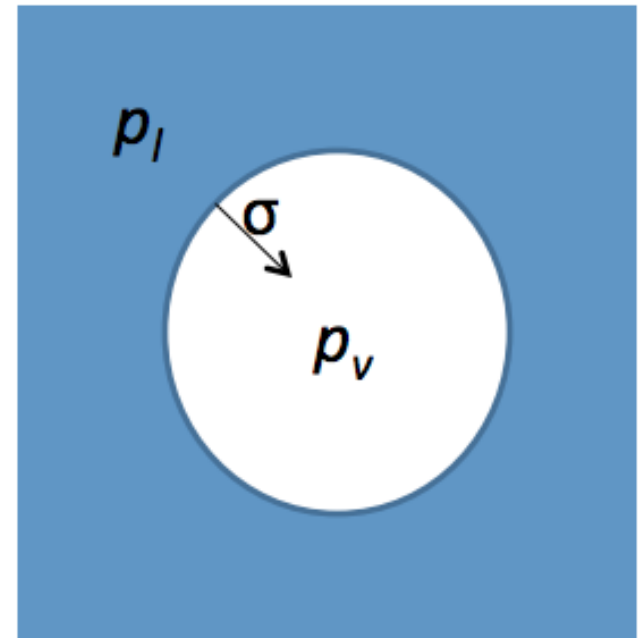
- Particles interacting
evaporate a small
amount of material:
bubble nucleation
- Four Cameras record bubbles
- Eight piezo-electric acoustic
sensors detect sound
- Recompression after
each event



Bubble chambers: Physics

- In a superheated fluid, energy deposition greater than E_{th} in a radius less than r_c will result in a bubble large enough to overcome surface tension (Seitz "Hot-Spike" Model)
- Low E or dE/dx result in smaller bubbles that immediately collapse
- Classical Thermodynamics:

$$p_v - p_l = \frac{2\sigma}{r_c}$$
$$E_{th} = \underbrace{4\pi r_c^2 \left(\sigma - T \frac{\partial \sigma}{\partial T} \right)}_{\text{Surface energy}} + \underbrace{\frac{4}{3}\pi r_c^3 \rho_v h}_{\text{Latent heat}}$$

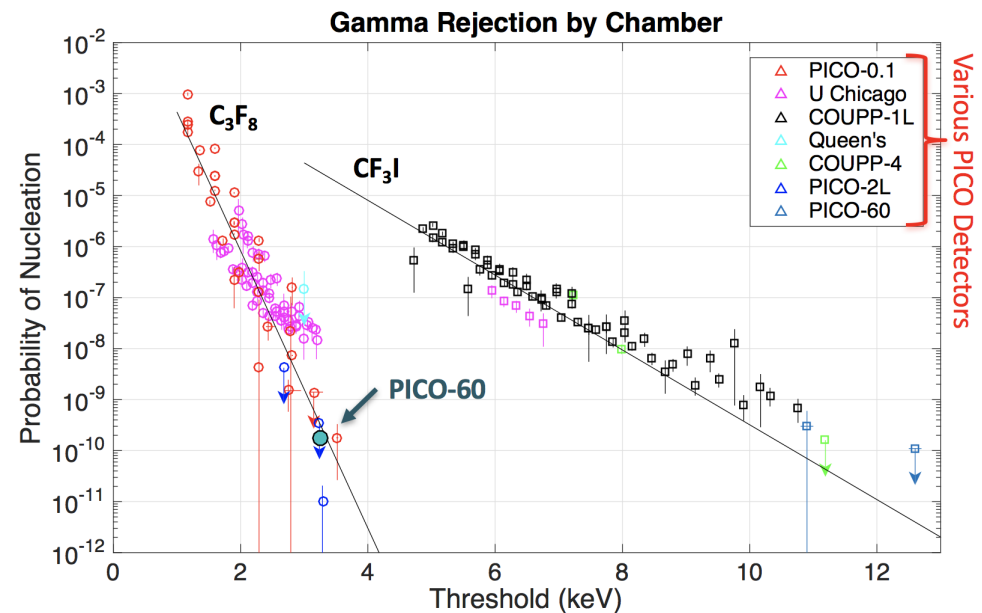
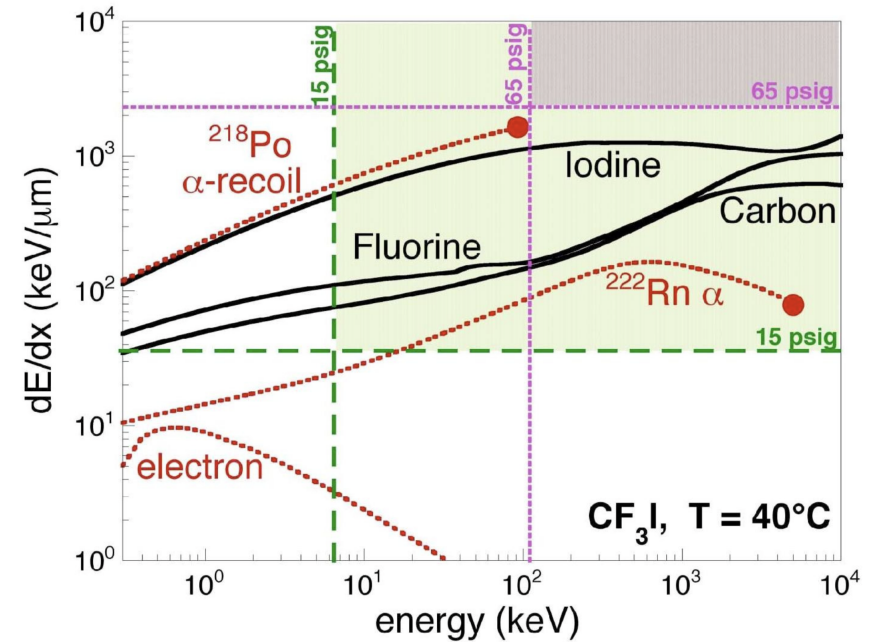


Bubble nucleation

Dependence of bubble nucleation on the total deposited energy and dE/dx

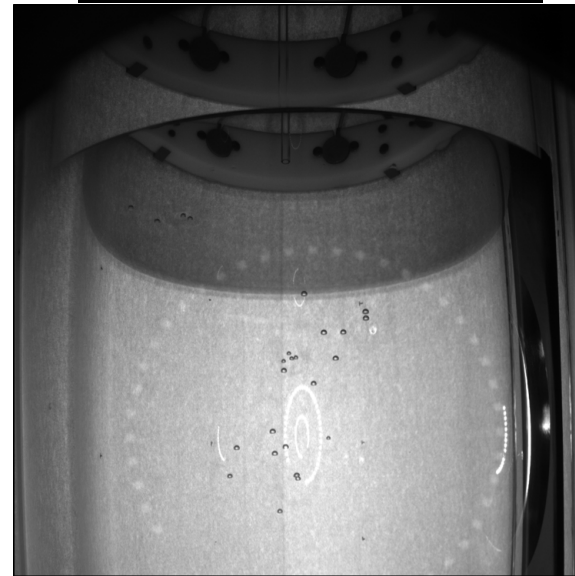
- Region of bubble nucleation at 15 psig
- Backgrounds: electrons, ^{218}Po , ^{222}Rn
- Signal processes of Iodine, Fluorine and Carbon nuclear recoils

insensitive to electrons and gammas



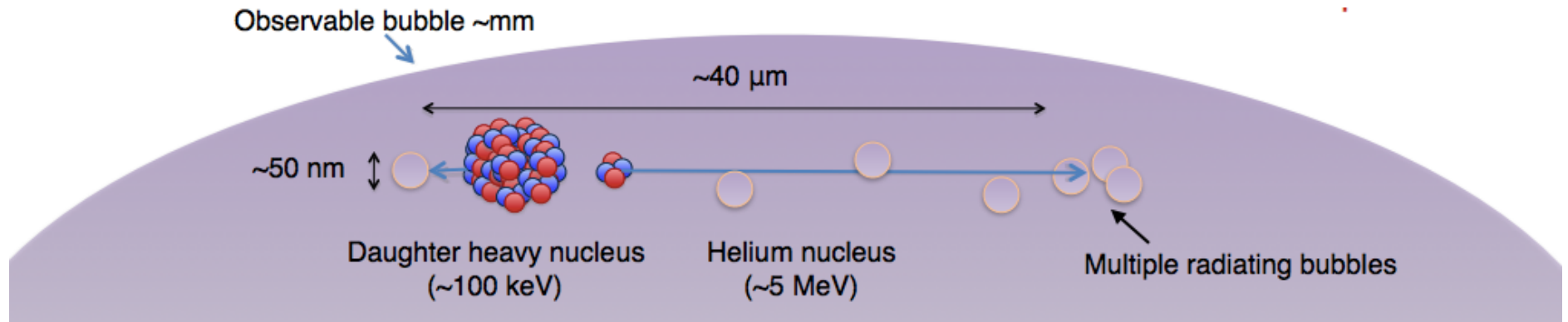
Bubble chambers: signal

- Alpha decays:
Nuclear recoil and
40 μm alpha track
1 bubble
- Neutrons:
Nuclear recoils
mean free path ~ 20 cm
3:1 multiple-single ratio
in PICO-60
- Neutrinos or WIMPs:
Nuclear recoil
mean free path $> 10^{10}$ cm
1 bubble

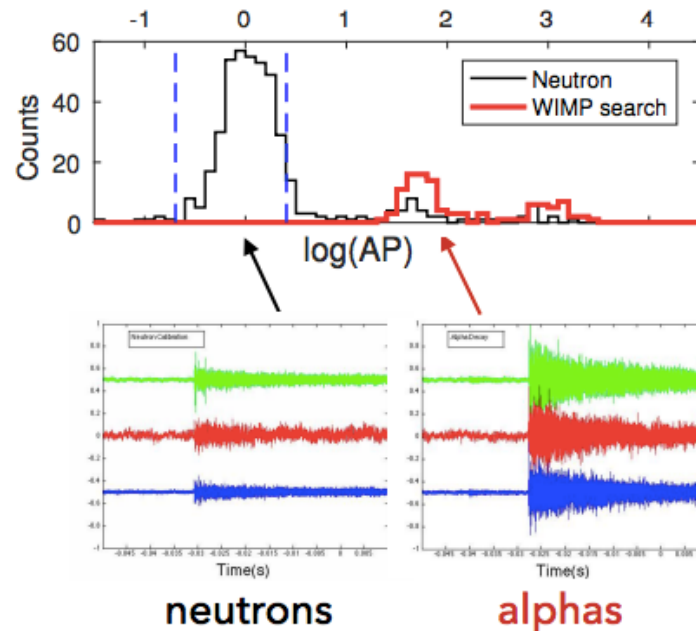


Bubble chambers: Acoustics

- Alphas are ~ 4 times louder than nuclear recoil bubbles



- $> 99.4\%$ discrimination against alpha events demonstrated
- Discovered by the PICASSO collaboration



Why bubble chambers?

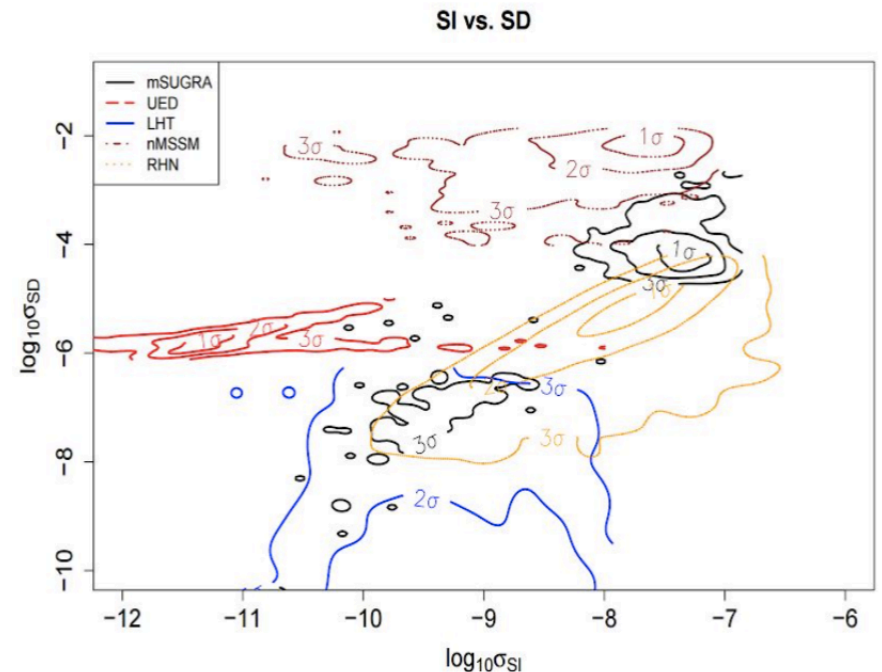
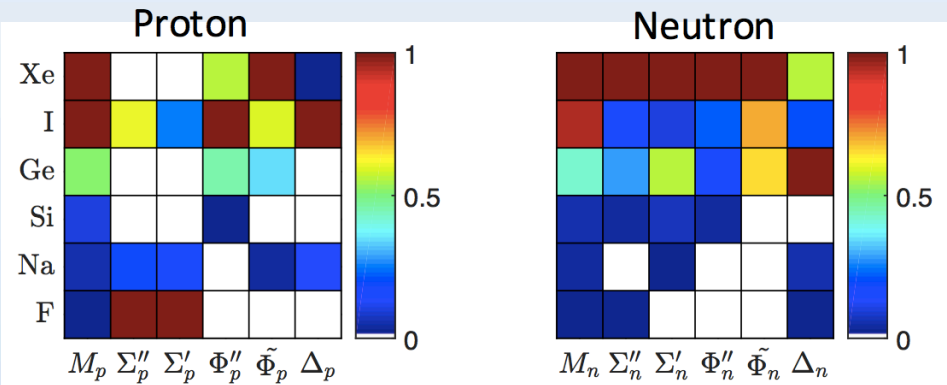
- Zero background
- Large target mass
- Low energy threshold (a few keV, and down to eV for some fluids)
- Multiple target nuclei
test expected cross section dependences on
atomic number and nuclear spin
(Fluorine, Iodine, Chlorine, Xenon, Argon, Bromine, Hydrogen...)
- Measure nuclear recoil energies (by varying threshold)
- No measure of nuclear recoil direction.

EFT and SI vs SD

Capability to instrument a wide range of target nuclei with sensitivity to diverse WIMP-nucleon couplings.
Unknown how WIMPs couple to matter

- Fluorine: Best sensitivity to spin dependent interactions.
- Iodine, Bromine, Xenon, Argon: High A targets to exploit A^2 dependence of spin-independent cross section.
- Hydrogen: Enhanced sensitivity to low mass particles.

Fitzpatrick, Haxton et al. Effective Field Theory Couplings



V. Barger, W-Y Keung and G. Shaughnessy, Phys. Rev. D78 (2008) 056007

SBC Collaboration



- **Eric Dahl**
- Rocco Coppejans
- Zhiheng Sheng
- Aaron Brandon
- David Velasco
- Ari Sloss
- Maheebub Khatri
- Dishen Wang
- Shishir Bandapalli



- **Ken Clark**
- Hector Hawley
- Patrick Hatch
- Austin De St Croix



- Marie-Cécile Piro
- Carsten Krauss
- Daniel Durnford
- Sumanta Pal
- Youngtak Ko
- Mitchel Baker



- Pietro Giampa
- Eric Poulin



- Mathieu Laurin



- Orin Harris



- Chris Jackson



- Eric Vázquez-Jáuregui
- Ernesto Alfonso-Pita
- Ariel Zuniga-Reyes
- Daniel Lámbarri



- Russell Neilson
- Matt Bressler
- Noah Lamb
- Stephen Windle



- Ilan Levine
- Ed Behnke
- Cody Cripe



- Hugh Lippincott
- TJ Whitis
- Runze Zhang



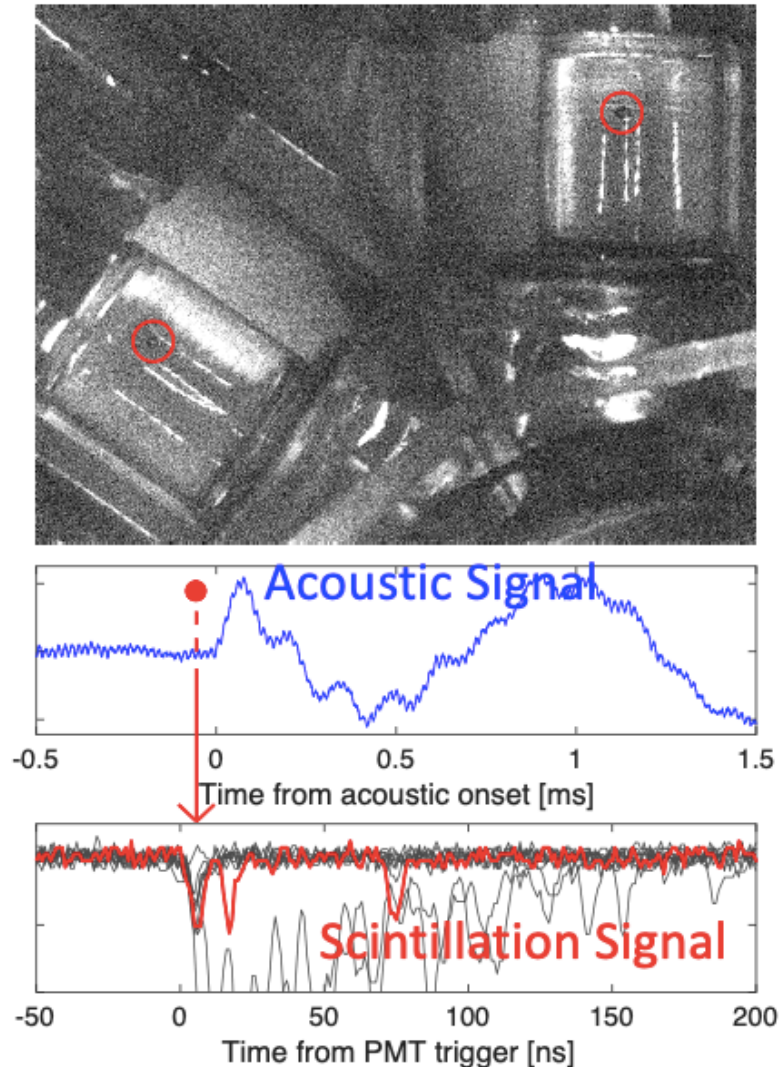
- Mike Crisler



First demonstration of SBC

Phys Rev Lett 118, 231301

A nuclear recoil:

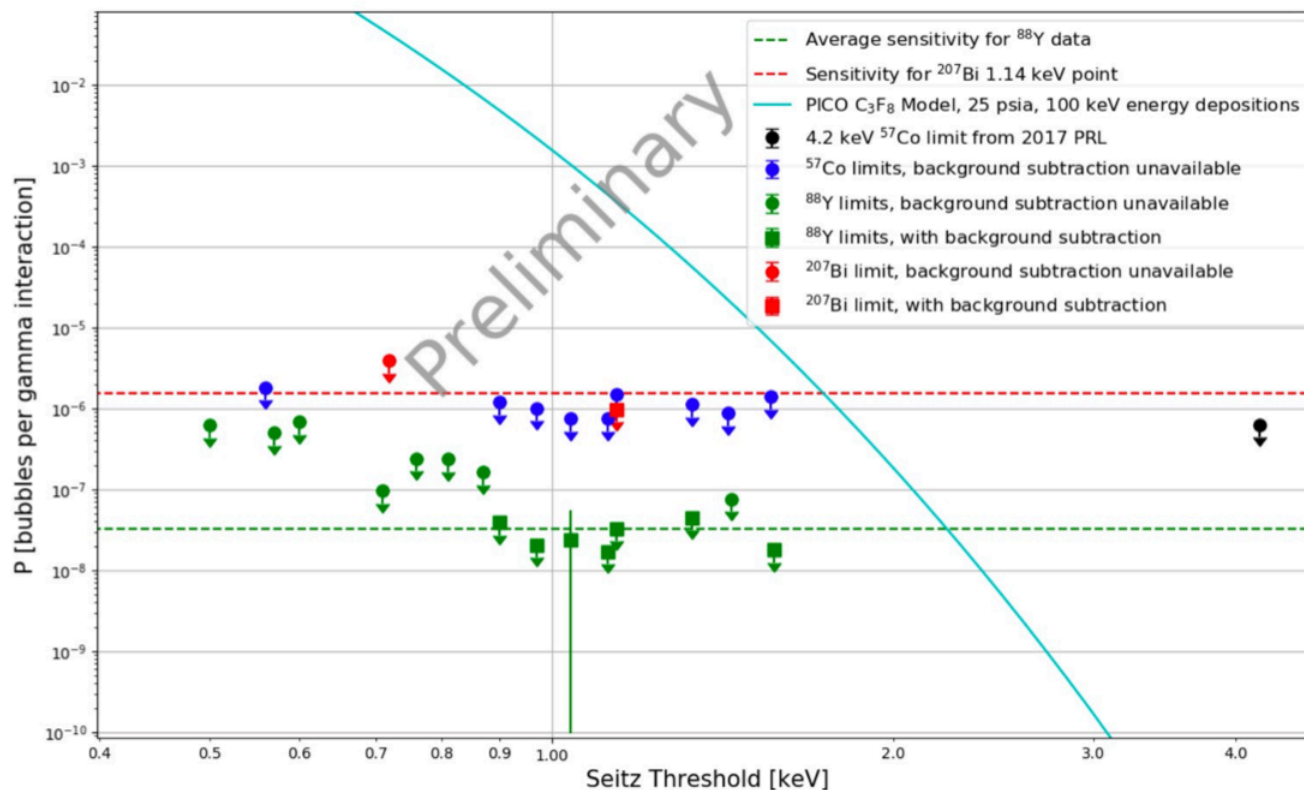


- Demonstrated (NU):
 - Xenon at 500 eV threshold
 - 30-gram target
 - 0.3% photon-detection efficiency
- Argon down to 40 eV threshold (1 bubble/ton-year from thermal fluctuations)
 - 10-kg target
 - 5% photon-detection efficiency (1 phd @ 2 keVr)Events with zero photons are signal

Xenon bubble chamber

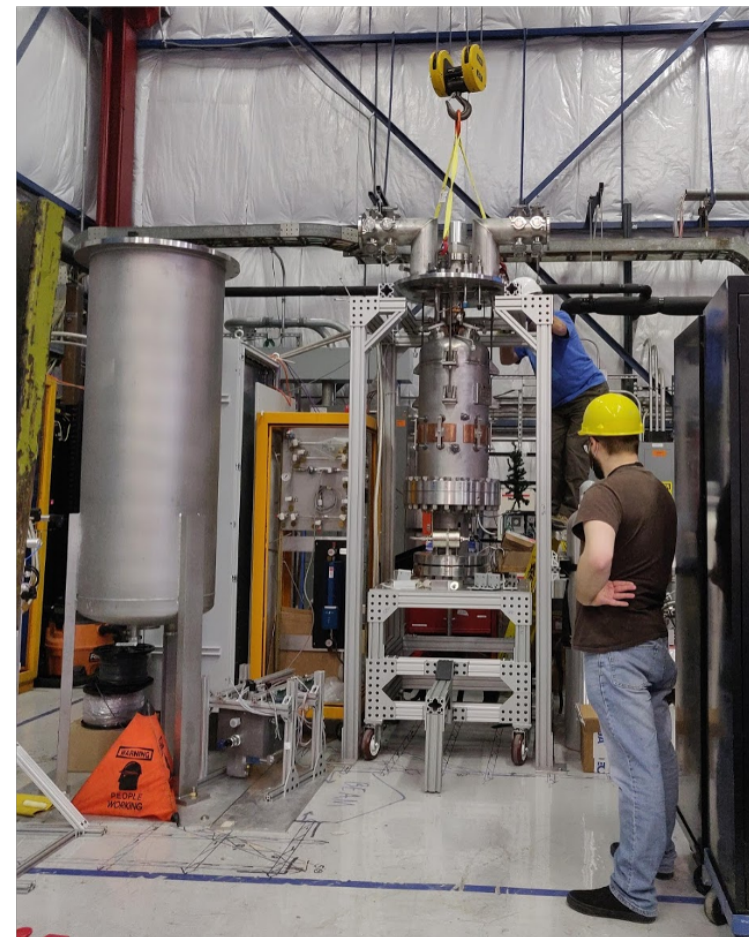
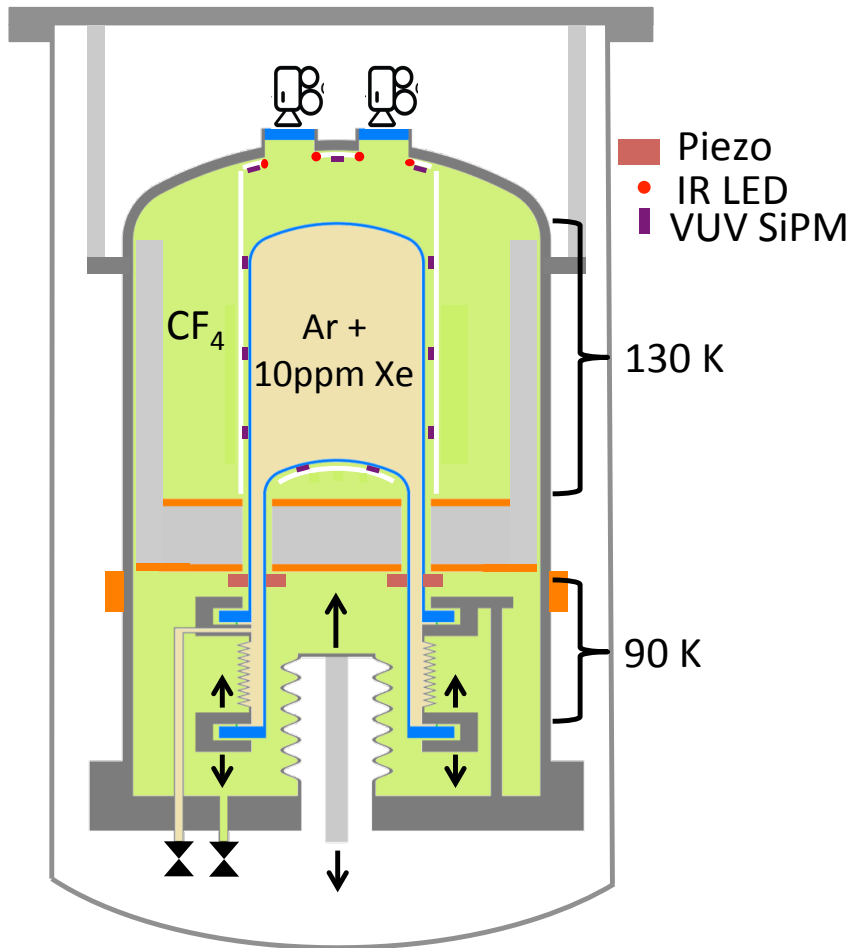
- Xenon measured to have outstanding ER discrimination
- Thresholds explored down to 500 eV
- No gamma induced ER observed
- Xe bubble chambers don't work for tracks (J.L. Brown, D.A. Glaser and M.L. Perl, Phys Rev 102, 1956), “solved” by adding 2% ethylene.

30g of LXe, 30% Overall Light Collection Efficiency



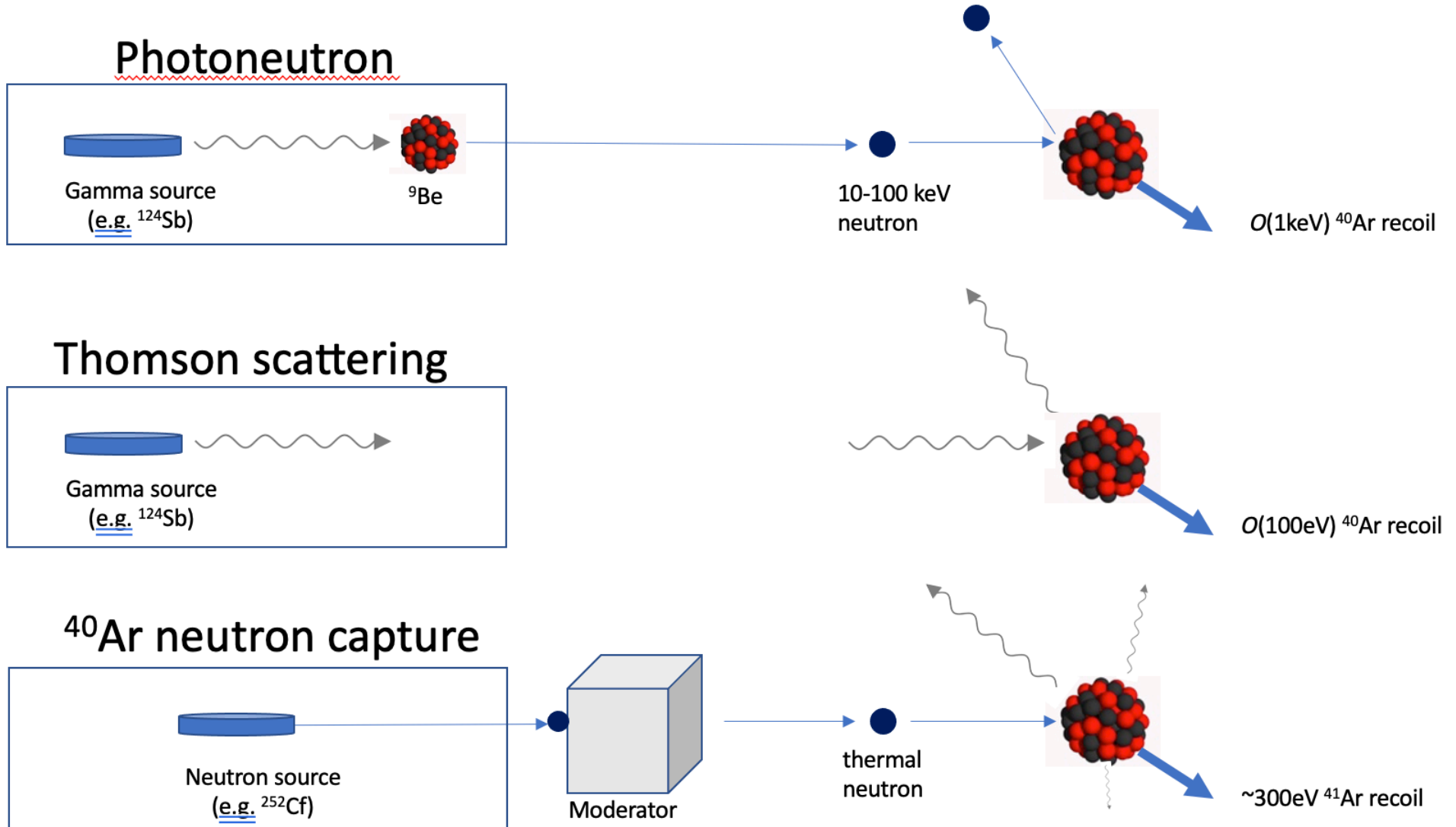
10 kg liquid Argon bubble chamber: 100 eV threshold

- Ar + 10-100 ppm Xe target, 178 nm scintillation
- SiPMs immersed in hydraulic fluid (CF₄ at 130K)
- 20-360 psia (\sim 1-25 bar) cycles
- Single-fluid, “right-side-up” geometry used by PICO-40L



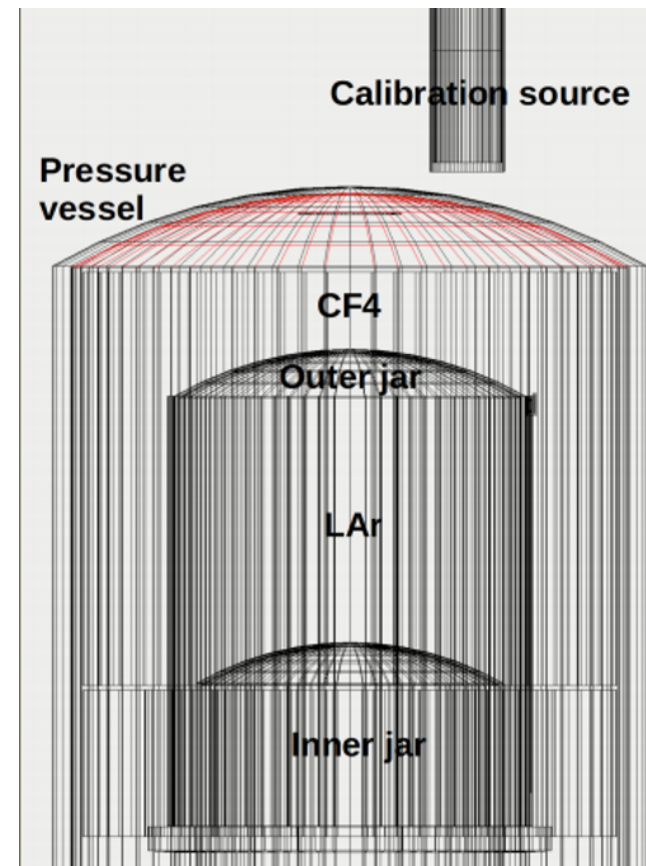
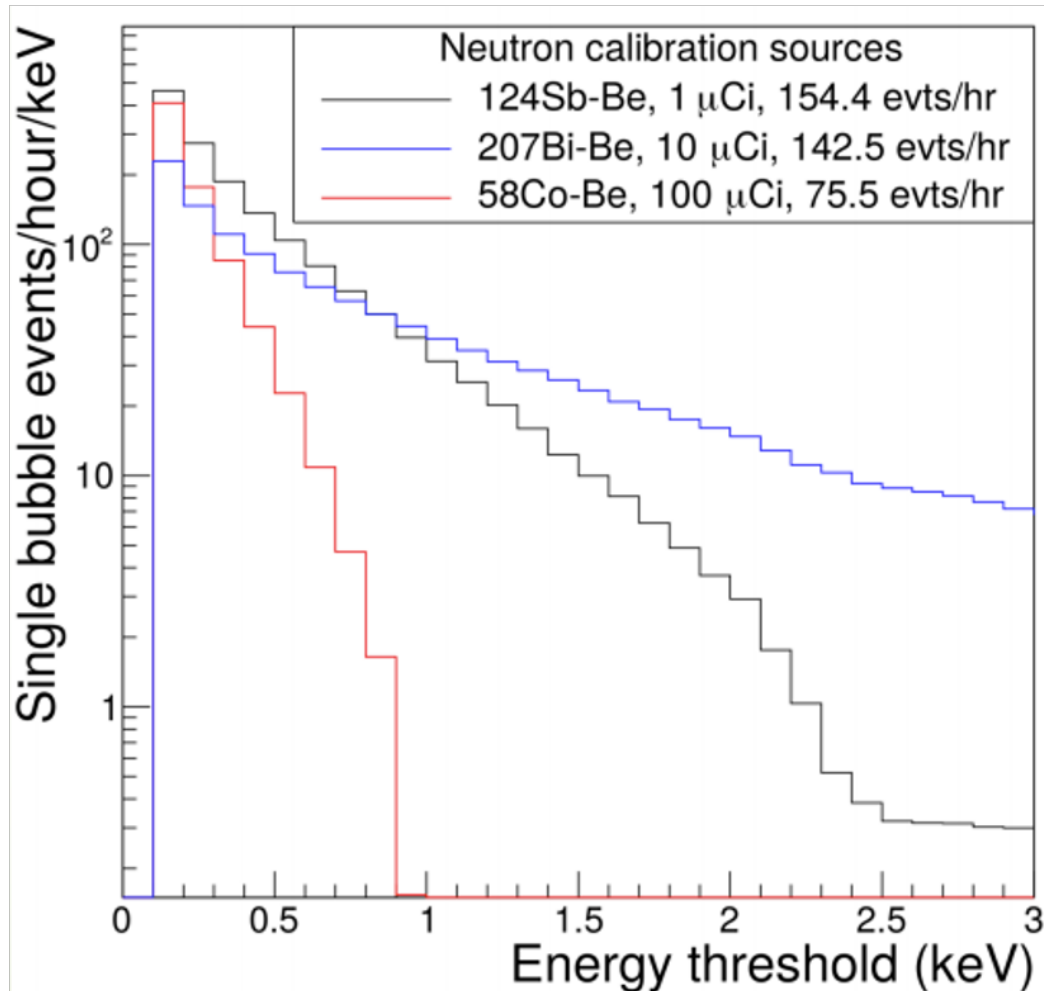
Calibration

- Different nuclear recoil calibration techniques

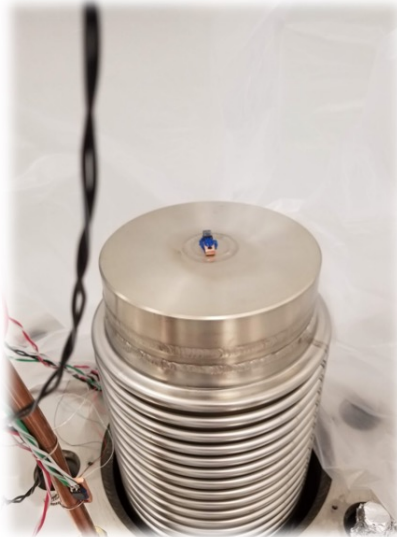
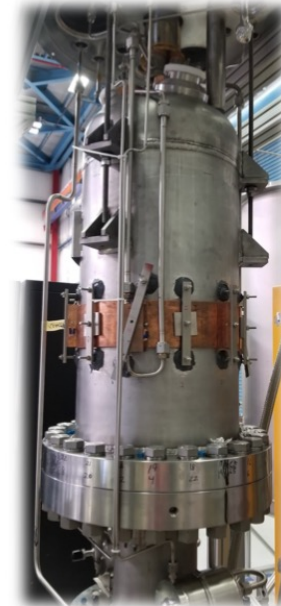


Calibration: photo-neutron sources

- $^9\text{Be}(\gamma, n)$, $Q = 1664$, keV gives nearly monoenergetic neutrons.
 ^{207}Bi : 90keV, ^{124}Sb : 23keV, ^{58}Co : 9keV

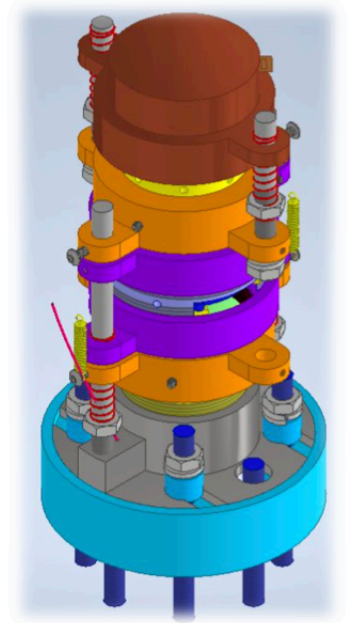
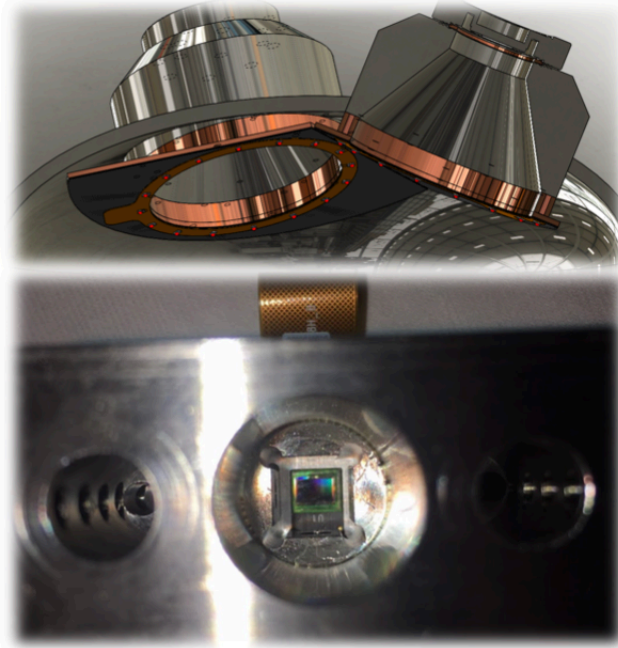
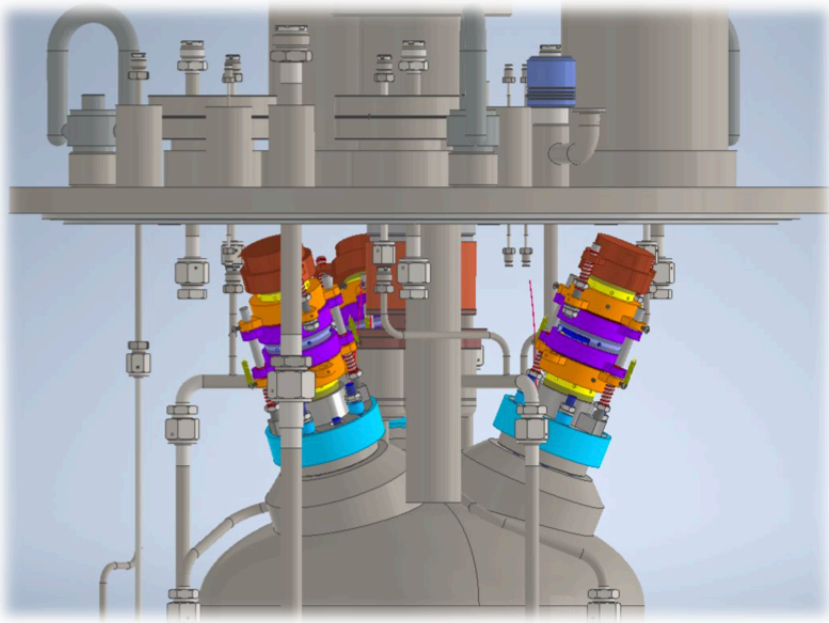


10 kg liquid Argon bubble chamber



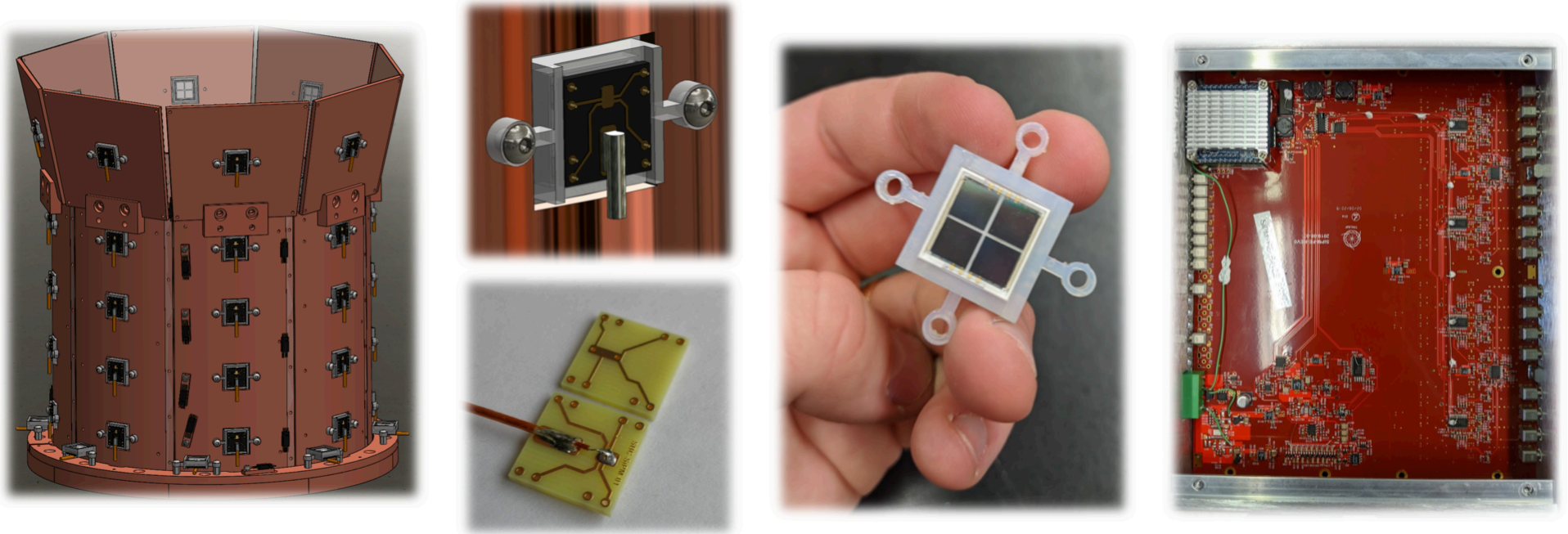
SBC-10kg: Readout systems

- Three Raspberry-Pi Controlled Camera System
- Three LED Rings To Provide Illumination



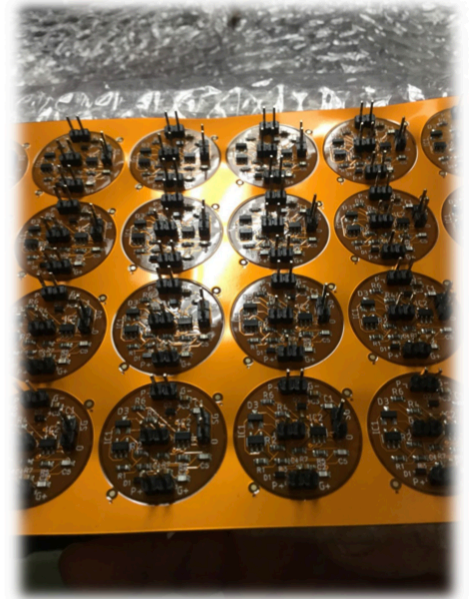
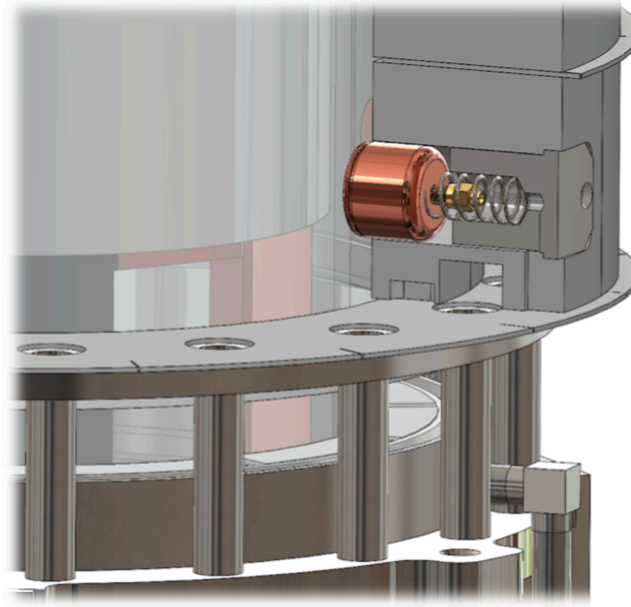
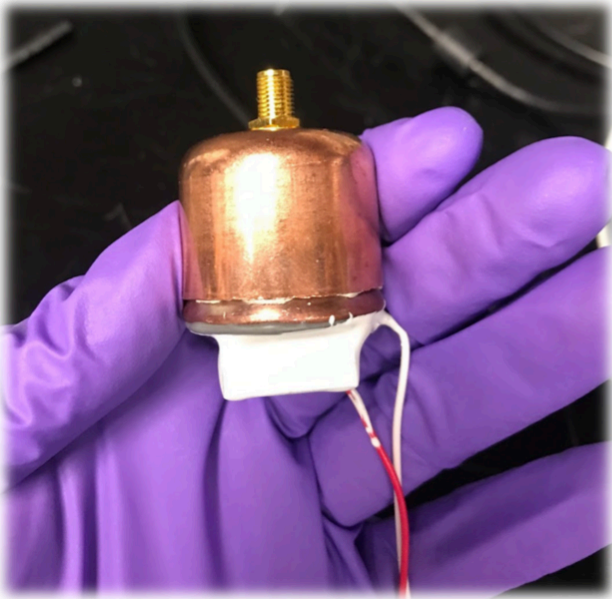
SBC-10kg: Readout systems

- 32 Hamamatsu VUV4 Quads
- Measure Scintillation Light In The Target Fluid.



SBC-10kg: Readout systems

- Eight Piezo Acoustic Sensors
- Monitor The Nucleation Process

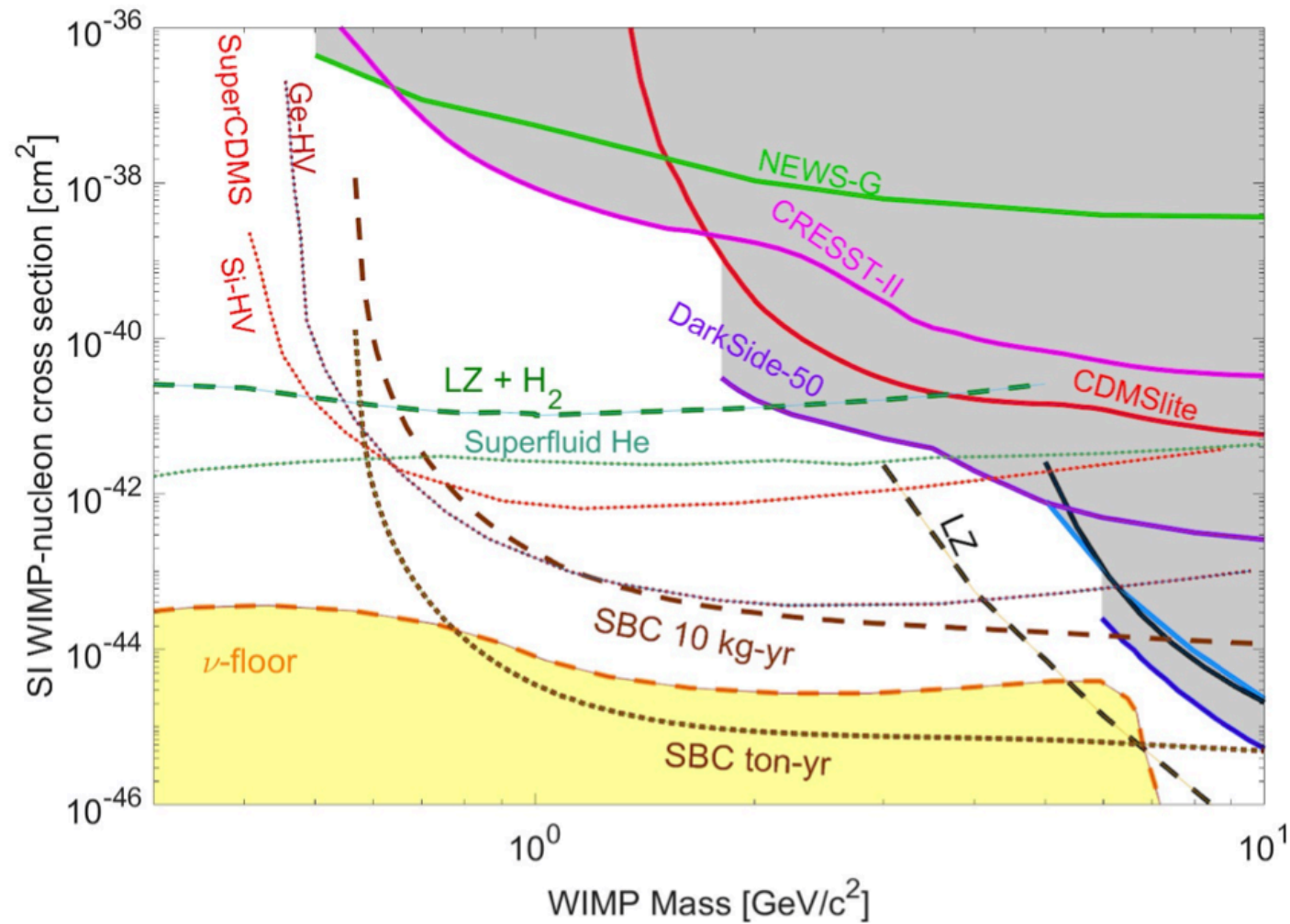


SBC: possible strategy

- SBC-Fermilab:
Build and commission detector
Calibrate NR and ER
- SBC-SNOLAB:
Build and install 2nd detector
Low mass dark matter searches
- SBC-CE ν NS:
Upgrade SBC-Fermilab detector
Install at a reactor site for CE ν NS



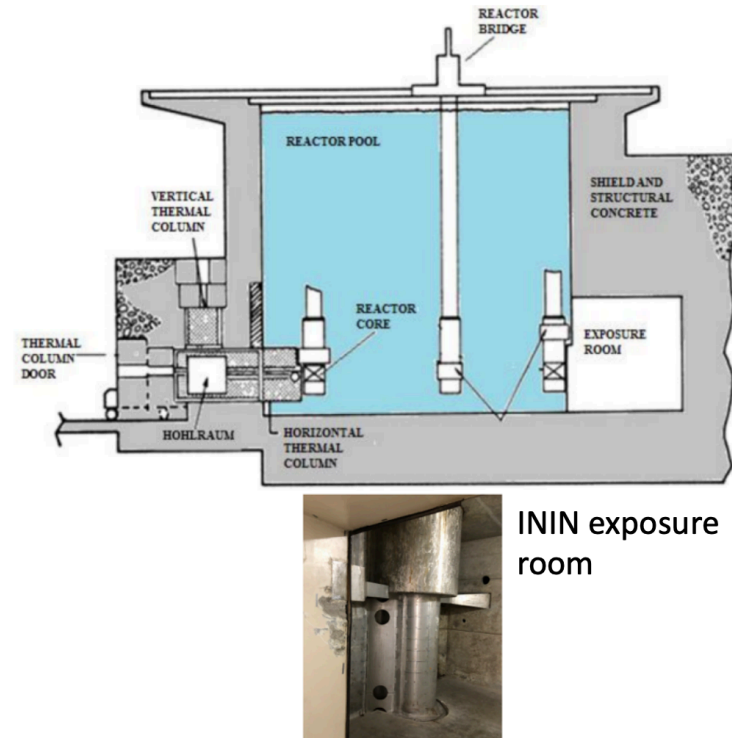
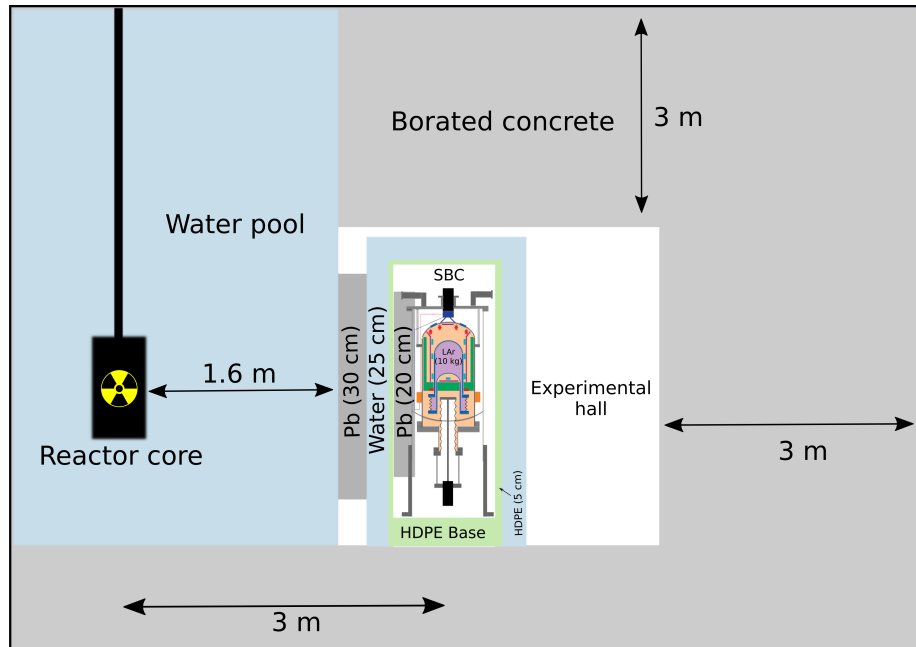
SBC Physics: Dark Matter



- WIMP searches down to solar floor (0.7-7 GeV)

SBC CE ν NS: physics reach

ININ 1MW Triga Mark-III reactor in Mexico



- Two sites explored: ININ and Laguna Verde

| Setup | LAr mass (kg) | Power (MW_{th}) | Distance (m) | Anti- ν flux uncertainty (%) | Threshold uncertainty (%) |
|--------|---------------|---------------------|--------------|----------------------------------|---------------------------|
| A | 10 | 1 | 3 | 2.4 | 5 |
| B | 100 | 2000 | 30 | 2.4 | 5 |
| B(1.5) | 100 | 2000 | 30 | 1.5 | 2 |

SBC Physics: $\text{CE}\nu\text{NS}$ reach

- Setup A:

$\sim 8 \text{ CE}\nu\text{NS/day}$ at 100 eV

0.25 evts/day - reactor backgrounds

0.85 evts/day - cosmogenic

Shielding = 0.3m Pb, 0.25m H₂O,

0.5m Polyethene, 0.2m Pb

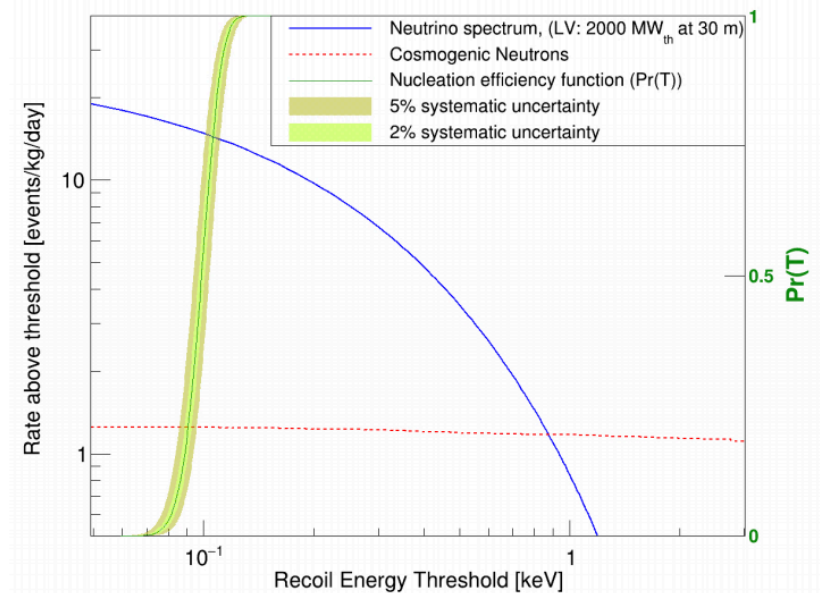
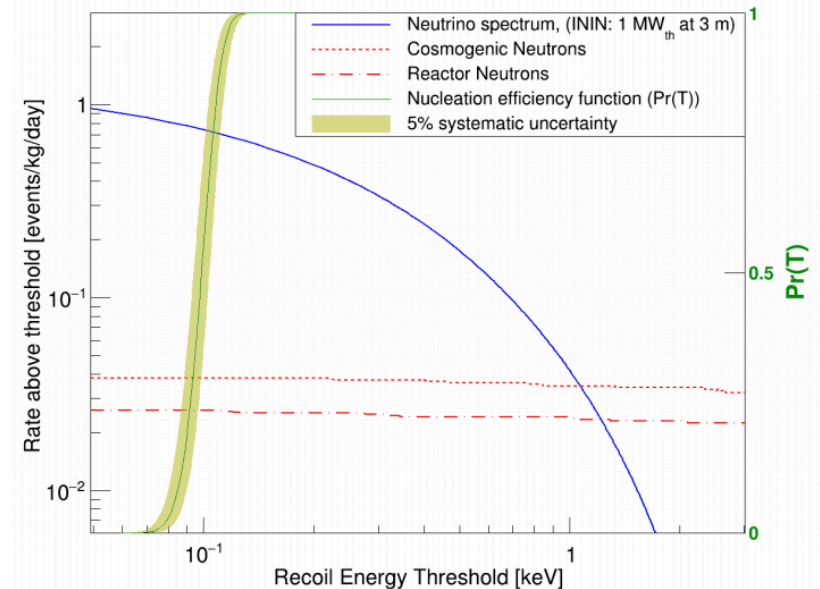
- Setup B:

$\sim 1570 \text{ CE}\nu\text{NS/day}$ at 100 eV

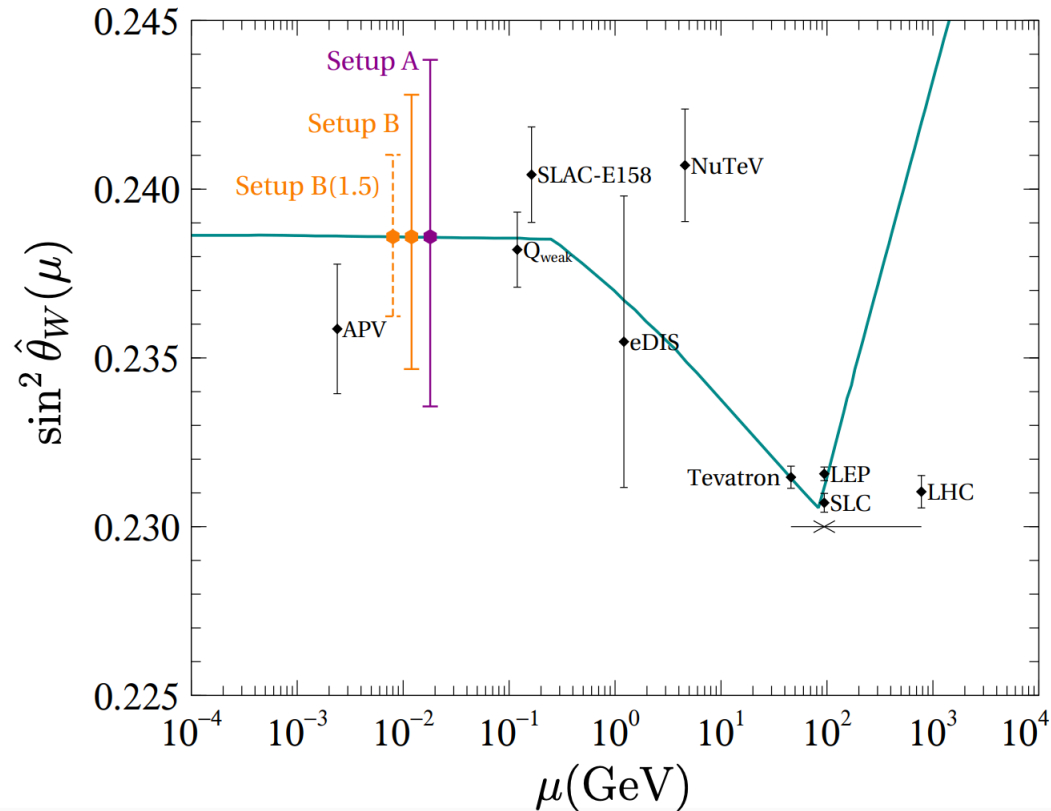
negligible reactor backgrounds
(30m + shielding)

180 evts/day - cosmogenic

Shielding = 3m H₂O, 0.5m Polyethen



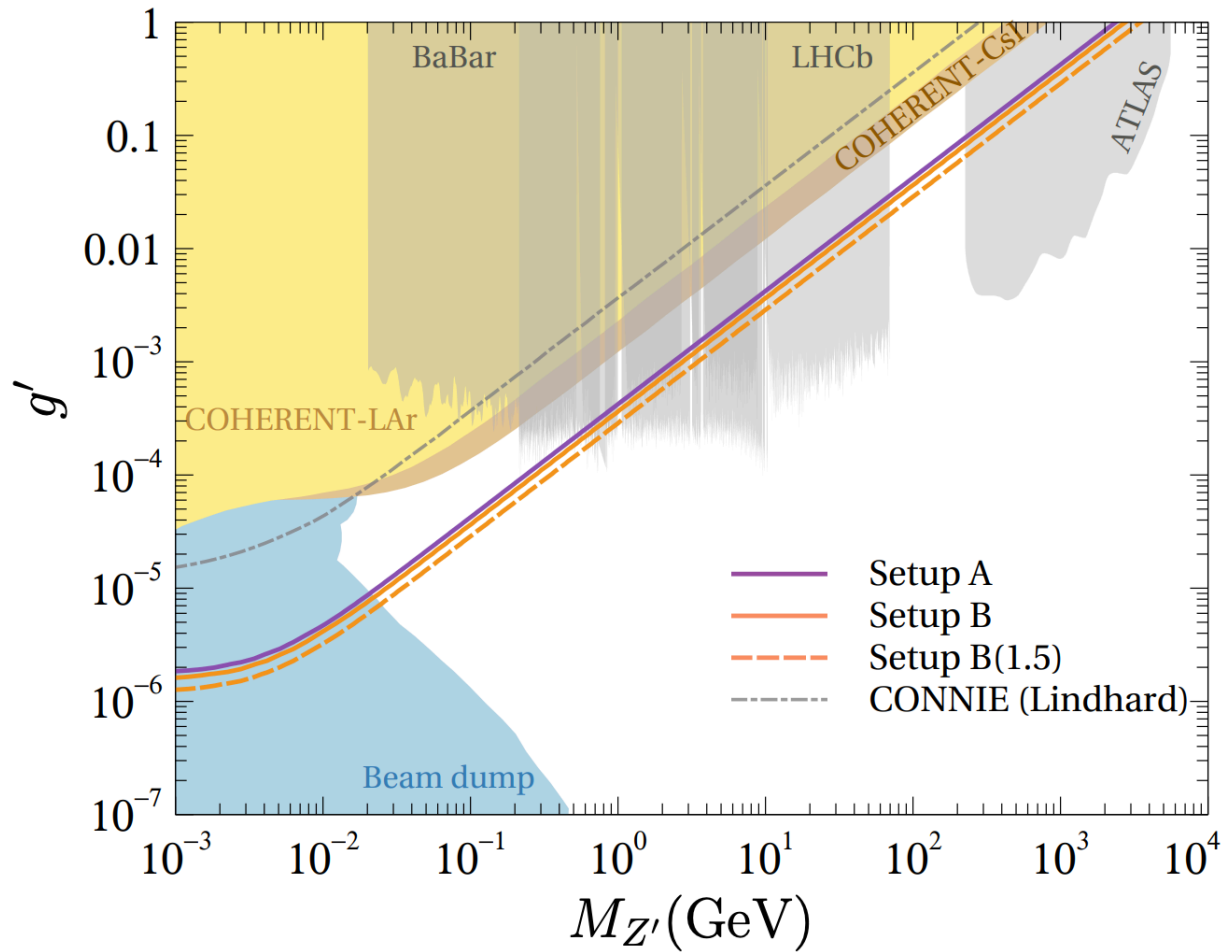
SBC CE ν NS Physics: weak mixing angle



- Precision as good as 1% in the weak mixing angle, similar to APV.
- Conservative: one year exposure, 2.4% flux uncertainty, 5% threshold uncertainty (A: ININ 10 kg, B: Laguna Verde 100 kg)
- Aggressive 1.5% flux, 2% threshold (B(1.5): Laguna Verde 100 kg)

$$\frac{d\sigma}{dT} = \frac{G_F^2}{2\pi} M_N Q_w^2 \left(2 - \frac{M_N T}{E_\nu^2} \right) F^2(q^2)$$

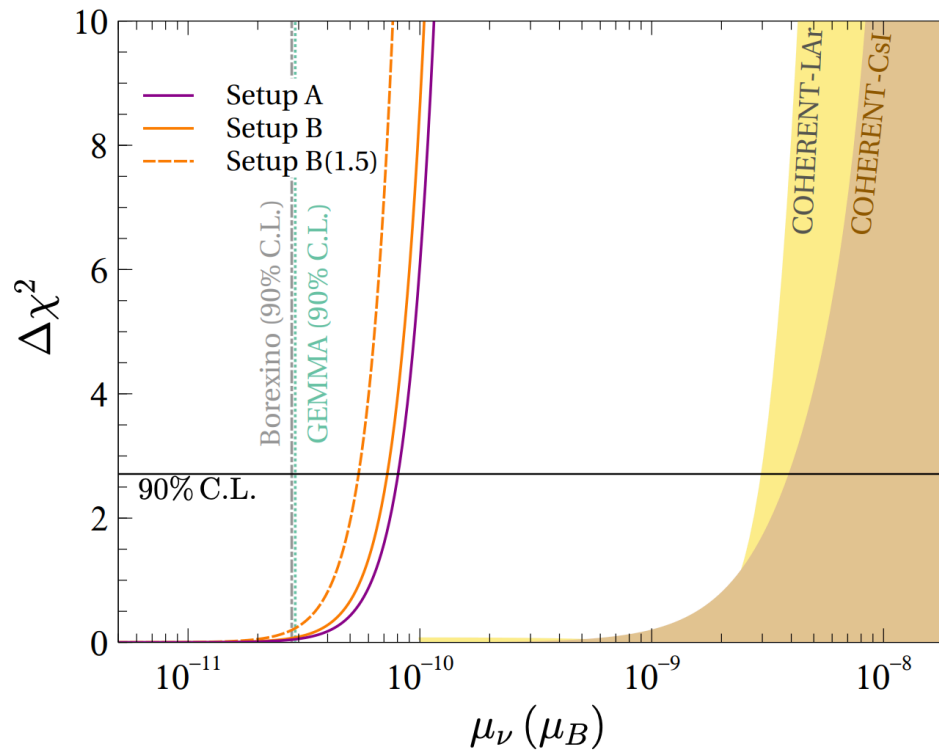
SBC CE ν NS Physics: Z' boson



- Most stringent bounds for new gauge vector bosons (20 MeV - 1 GeV and 70 - 230 GeV).

$$\mathcal{L}_{\text{eff}} = -\frac{g'^2 Q_l Q_q}{q^2 + M_{Z'}^2} \left[\sum_{\alpha} \bar{\nu}_{\alpha} \gamma^{\mu} P_L \nu_{\alpha} \right] \left[\sum_q \bar{q} \gamma_{\mu} q \right]$$

SBC CE ν NS Physics: ν magnetic moment



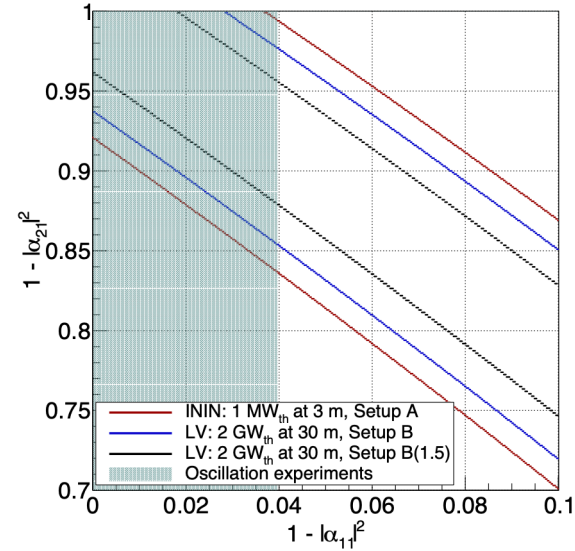
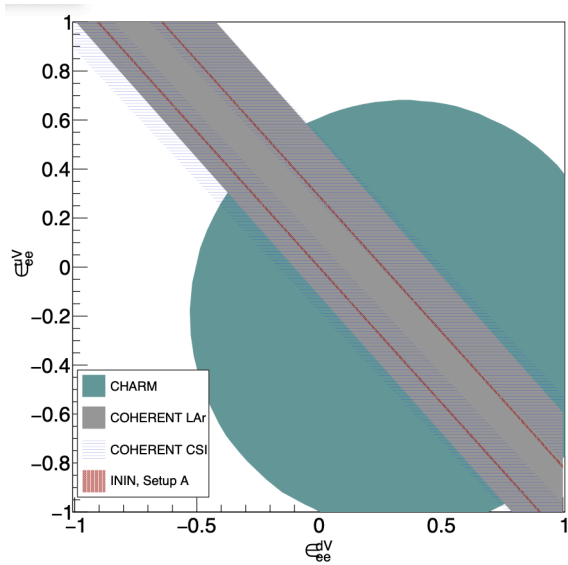
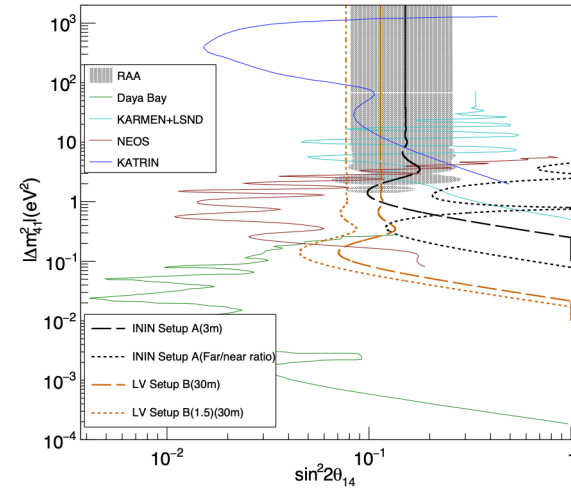
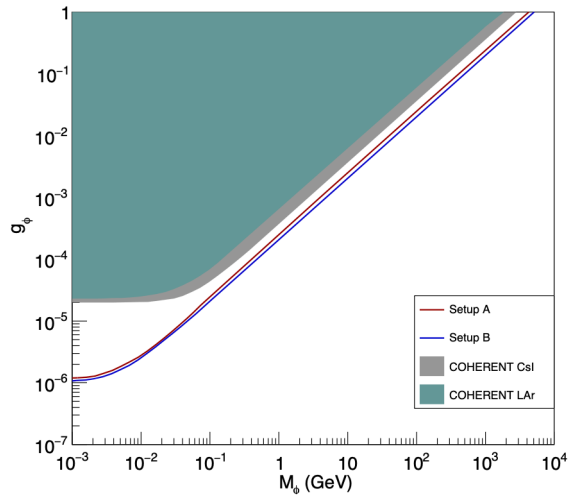
- $\mu_\nu = 5.4 \times 10^{-11} \mu_B$ (90% C.L.), similar to GEMMA and Borexino.

$$\frac{d\sigma}{dT} = \pi \frac{\alpha_{\text{EM}}^2 Z^2 \mu_\nu^2}{m_e^2} \left(\frac{1}{T} - \frac{1}{E_\nu} + \frac{T}{4E_\nu^2} \right) F^2(q^2),$$

Physics reach of a low threshold scintillating argon bubble chamber
in coherent elastic neutrino-nucleus scattering reactor experiments

[Phys. Rev. D 103, L091301 \(2021\)](#)

SBC CE ν NS: New Physics



New Physics searches in a low threshold scintillating argon bubble chamber measuring coherent elastic neutrino-nucleus scattering in reactors

<https://arxiv.org/abs/2203.05982>

Final remarks

- SBC is a 10 kg LAr bubble chamber:
unique potential for reactor CE ν NS measurement
with low backgrounds
 - 100 eV nuclear recoil detection
 - Dark matter searches for masses 0.7-7 GeV
 - Rich CE ν NS physics programme:
weak mixing angle, Z' boson, neutrino magnetic moment,
sterile neutrinos, NSI, unitarity violation