

WIMP scattering off Xenon: nuclear structure insights

Javier Menéndez

JSPS Fellow, The University of Tokyo

with P. Klos, A. Schwenk, L. Vietze (Darmstadt), D. Gazit (HUJI) and W. Haxton (Berkeley)

“Effective Theories and Dark Matter”

Mainz Institute for Theoretical Physics (MITP)
18th March 2015

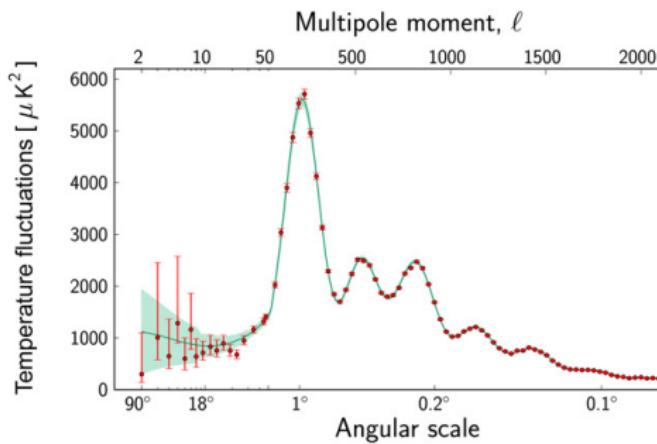
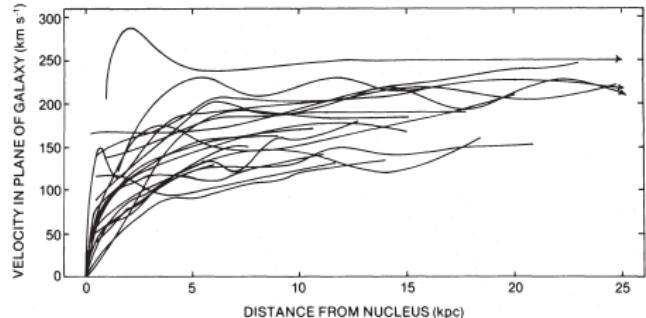


日本学術振興会
Japan Society for the Promotion of Science



東京大学
THE UNIVERSITY OF TOKYO

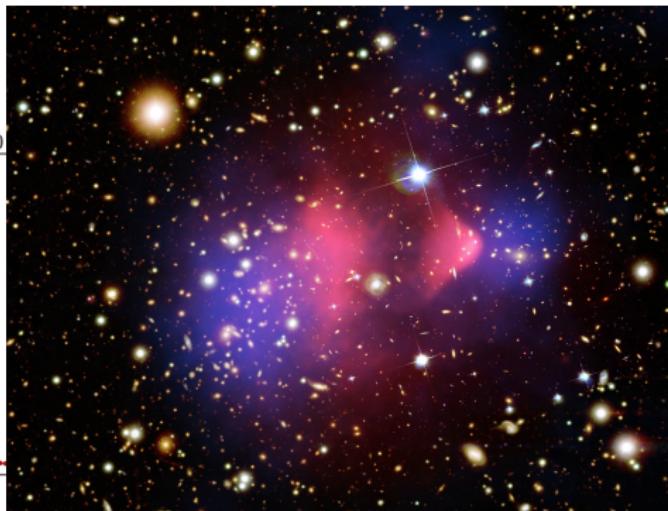
Dark Matter: evidence



Solid evidence of Dark Matter
in very different observations:

Rotation curves, Lensing, CMB...

Zwicky 1930's, Rubin 1970's,..., Planck (2013)

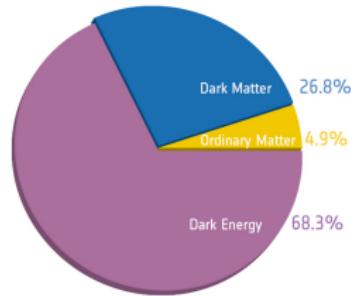


What is Dark Matter made of?

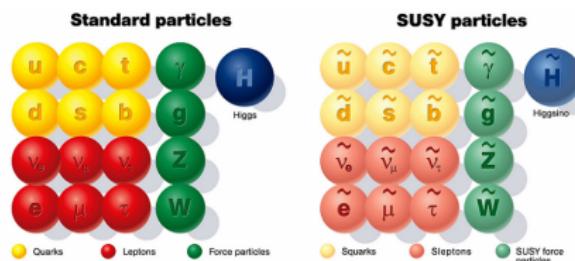
The composition of Dark Matter is unknown

New particles: To be detected

- Weakly interacting massive particles (WIMPs)
- Sterile neutrinos
- Axions
- Gravitons
- ...



Lightest supersymmetric particles (usually neutralino) predicted in SUSY extensions of the Standard Model



Expected WIMP-density naturally accounts for observed Dark Matter density

WIMP scattering off nuclei

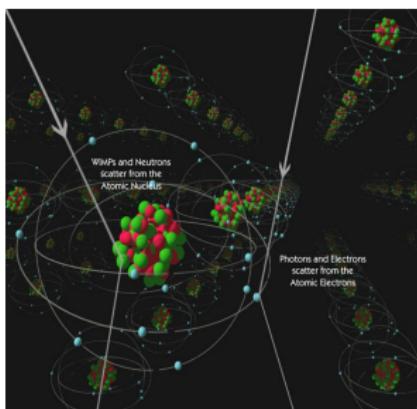
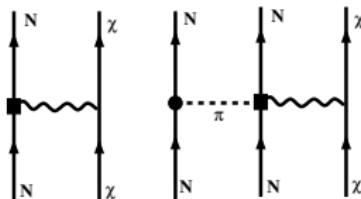
WIMPs interact with quarks \Rightarrow nuclei

Direct detection experiments: XENON100, LUX
nuclear recoil from WIMP scattering off nuclei
sensitive to Dark Matter masses $\gtrsim 1$ GeV

Assume spin 1/2 non-relativistic WIMPs
couple to nuclear density or spin

$$\langle \text{Initial} | \int dx j^\mu(x) J_\mu(x) | \text{Final} \rangle$$

- WIMP-nucleus interaction:
 - isoscalar spin-independent
 - isoscalar spin-dependent
 - isovector spin-dependent:
correspond to axial weak neutral currents
1b + 2b currents predicted by chiral EFT
- Nuclear structure calculation:
initial and final states



CDMS Collaboration

Outline

1 Nuclear Structure of Xenon Isotopes

2 Spin-Independent WIMP scattering

3 Spin-Dependent WIMP scattering

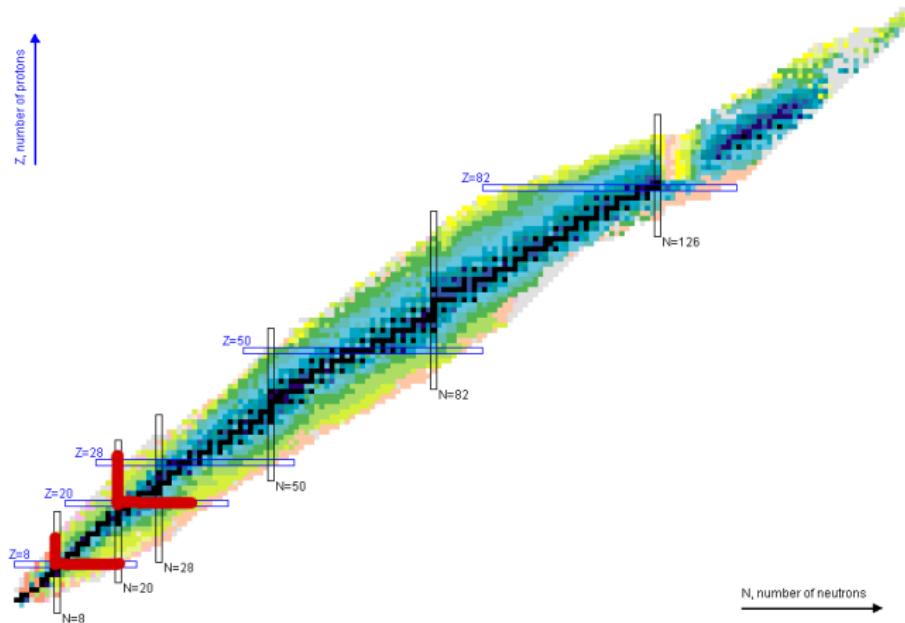
Outline

1 Nuclear Structure of Xenon Isotopes

2 Spin-Independent WIMP scattering

3 Spin-Dependent WIMP scattering

Nuclear landscape



Big variety of nuclei in the nuclear chart,
 $A \sim 2\ldots 300$

Hard many-body problem: approximate methods suited for different regions

Ab initio methods remarkable success in selected medium-mass nuclei:
Coupled-Cluster (CC), In-medium Similarly Renormalization Group (IMSRG),
Self-Consistent Green's Function (SCGF)...

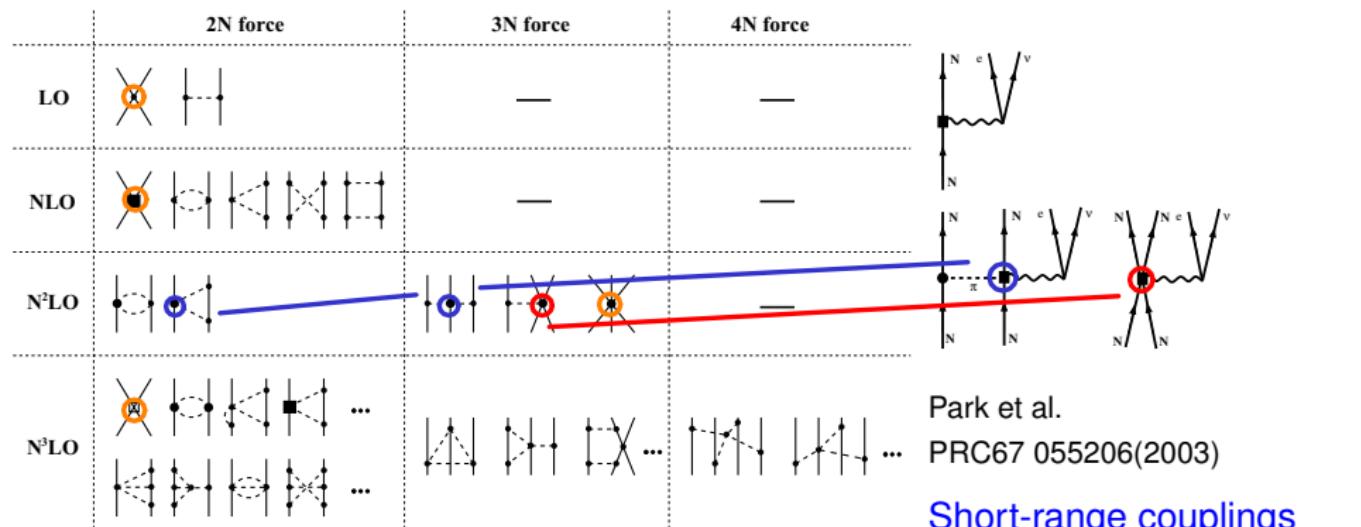
Microscopic shell model: reduced degrees of freedom (valence space),
cover wider range range of medium-mass nuclei

Chiral effective field theory

Chiral EFT: low energy approach to QCD, nuclear structure energies

Approximate chiral symmetry: pion exchanges, contact interactions

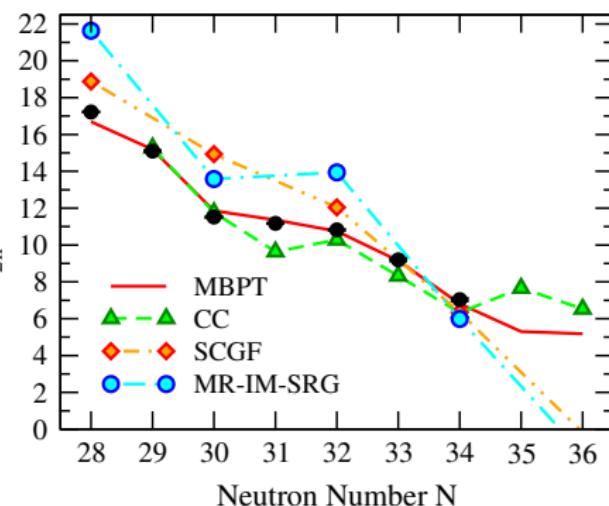
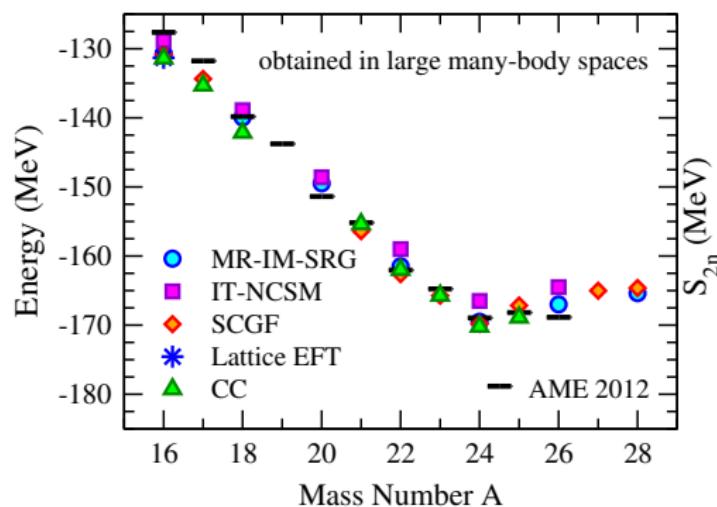
Systematic expansion: nuclear forces and electroweak currents



Weinberg, van Kolck, Kaplan, Savage, Weise, Epelbaum, Meißner...

Nuclear structure with chiral EFT forces

Great success prediction of oxygen dripline and calcium separation energies



Hergert et al. PRL110 242501 (2013)

Cipollone et al. PRL111 062501 (2013)

Jansen et al. PRL113 142502 (2014)

Gallant et al. PRL 109 032506 (2012)

Wienholtz et al. Nature 498 346 (2013)

Hagen et al. PRL 109 032502 (2012)

Somà et al. PRC 89 061301 (2014)

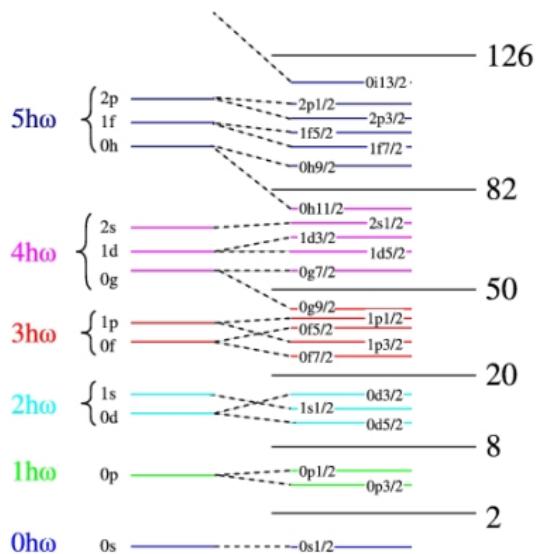
Hergert et al. PRC 90 041302 (2014) ↗ ↘ ↙

Stable Xenon isotopes

Xenon isotopes beyond present efforts with ab-initio calculations

Seven stable isotopes: $^{128,129,130,131,132,134,136}\text{Xe}$, $Z = 54$, $N = 74 - 82$

Solve with phenomenological shell model



Configuration space divided into

- Inner core: filled orbits up to $Z = 50$, $N = 50$ (^{100}Sn)
- Valence space: active orbits $0g_{7/2}$, $1d_{3/2}$, $1d_{5/2}$, $2s_{1/2}$ and $0h_{11/2}$
- Outer orbits: empty orbits

$$\text{Dim} \sim \binom{(p+1)(p+2)_\nu}{N} \binom{(p+1)(p+2)_\pi}{Z}$$

Solving the nuclear many-body problem

Solve the nuclear many-body problem in the valence space:

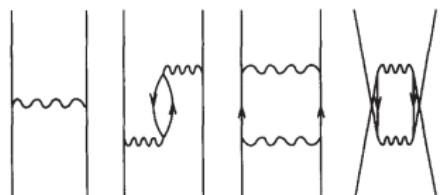
$$H |\Psi\rangle = E |\Psi\rangle \rightarrow H_{\text{eff}} |\Psi\rangle_{\text{eff}} = E |\Psi\rangle_{\text{eff}}$$

with H_{eff} based on NN forces,

many-body perturbation theory

include effects of core and high-energy orbits

phenomenological modifications needed to
compensate for absence of 3N forces



Many body states are linear combination of Slater Determinants

$$|\phi_\alpha\rangle = a_{i1}^+ a_{i2}^+ \dots a_{iA}^+ |0\rangle \quad |\Psi\rangle_{\text{eff}} = \sum_\alpha c_\alpha |\phi_\alpha\rangle$$

The ISM code Antoine diagonalizes up to 10^{10} Slater determinants

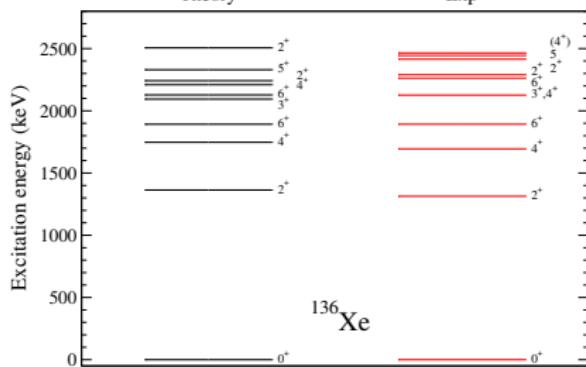
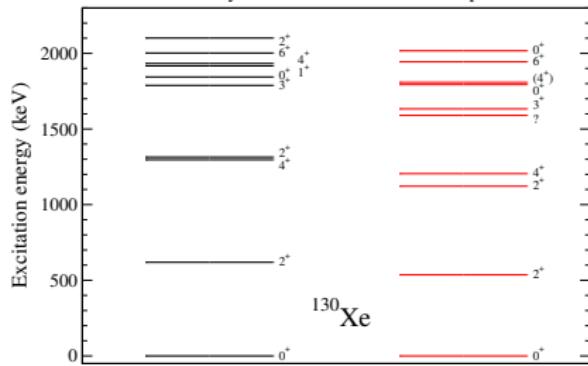
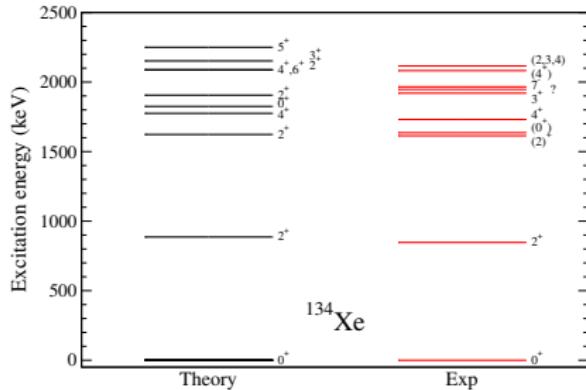
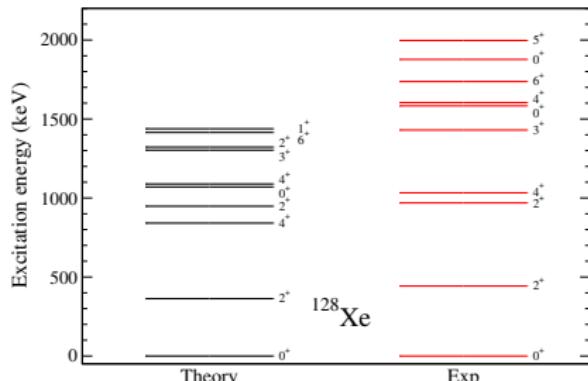
Caurier *et al.* RMP 77 (2005)

Valence space and interaction tested in nuclear structure, β , $\beta\beta$ decay studies

Even-mass Xe spectra

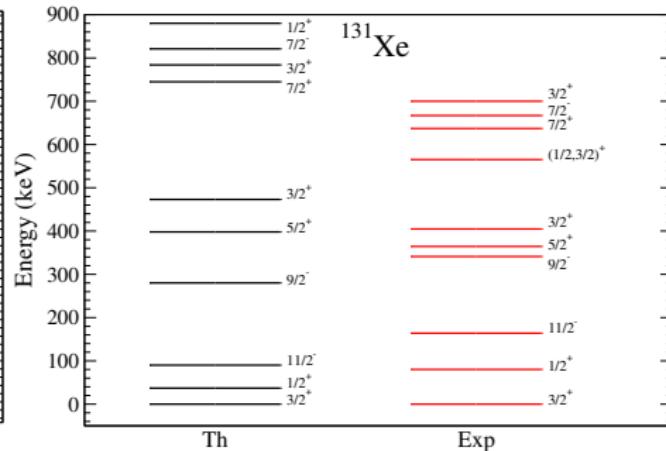
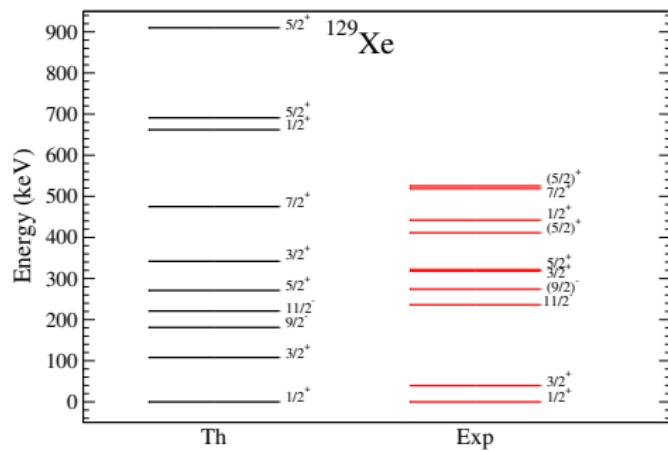
Low-lying excitation spectra in good agreement to experiment in all cases

Vietze, Klos, JM, Haxton, Schwenk PRD91 043520 (2015)



$^{129,131}\text{Xe}$ spectra

For odd-mass Xenon isotopes also very good agreement with experiment



JM, Gazit, Schwenk PRD86 103511(2012)

Very low-lying first-excited states $\sim 40, 80$

If WIMPs have enough kinetic energy,
inelastic scattering allowed

$$p_{\pm} = \mu v_i \left(1 \pm \sqrt{1 - \frac{2E^*}{\mu v_i^2}} \right)$$

See talk by P. Klos on Tuesday 24!

Outline

1 Nuclear Structure of Xenon Isotopes

2 Spin-Independent WIMP scattering

3 Spin-Dependent WIMP scattering

Spin-Independent scattering

Spin-Independent (SI) interaction: WIMPs couple to the nuclear density

The interaction Lagrangian at low-energies is given by the scalar leptonic $j(\mathbf{r})$ and scalar hadronic $S(\mathbf{r})$ currents:

$$\mathcal{L}_\chi^{\text{SI}} = \frac{G_F}{\sqrt{2}} \int d^3\mathbf{r} j(\mathbf{r}) S(\mathbf{r})$$

The scattering cross-section is proportional to the structure factor $S_S(q)$:

$$\frac{d\sigma}{dq^2} = \frac{2}{(2J_i + 1)\pi v^2} \sum_{s_f, s_i} \sum_{M_f, M_i} |\langle f | \mathcal{L}_\chi^{\text{SI}} | i \rangle|^2 = \frac{8G_F^2}{(2J_i + 1)v^2} S_S(q)$$

Only 1 body currents considered in SI scattering: $S(\mathbf{r}) = c_0 \sum_{i=1}^A \delta^{(3)}(\mathbf{r} - \mathbf{r}_i)$

Their impact is under discussion, but could be relevant

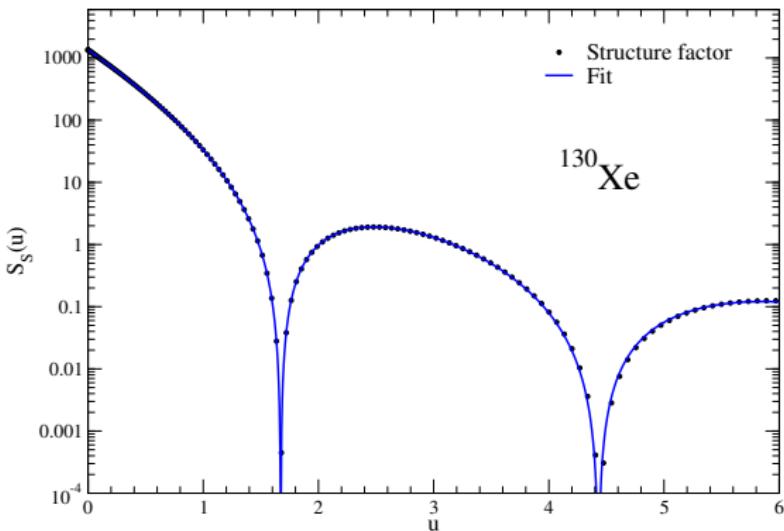
Prezeau et al. PRL91 231301 (2003), Cirigliano et al. JHEP10 025 (2012),

Beane et al. PRD89 074505 (2014)

SI structure factors: coherence

Structure factor decomposed in multipole expansion

$$S_S(q) = \sum_{L=0,2,4,\dots} \left| \langle J \parallel c_0 \sum_{i=1}^A j_L(qr_i) Y_{LM}(\mathbf{r}_i) \parallel J \rangle \right|^2,$$



Plot as function of dimensionless $u = p^2 b^2 / 2$

Coherent response at $p = 0$:

$$\frac{S_S(0)}{2J+1} = A^2 \frac{c_0^2}{4\pi}$$

Lost at $p > 0$

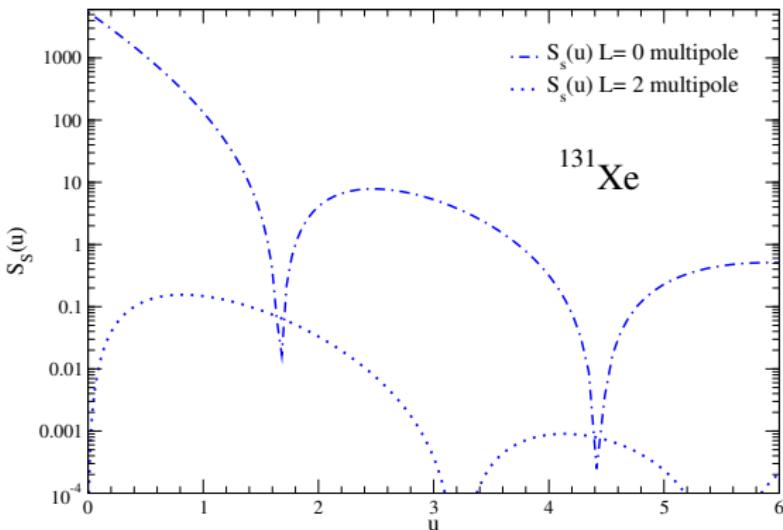
With similar couplings for different responses, SI scattering always dominant

Vietze, Klos, JM, Haxton, Schwenk PRD91 043520 (2015)

SI structure factors: multipole decomposition

Structure factor decomposed in multipole expansion

$$S_S(q) = \sum_{L=0,2,4,\dots} \left| \langle J \| c_0 \sum_{i=1}^A j_L(qr_i) Y_{LM}(\mathbf{r}_i) \| J \rangle \right|^2,$$



Plot as function of
dimensionless $u = p^2 b^2 / 2$

For ^{131}Xe , $J = 3/2$,
 $L = 0, 2$ multipoles allowed

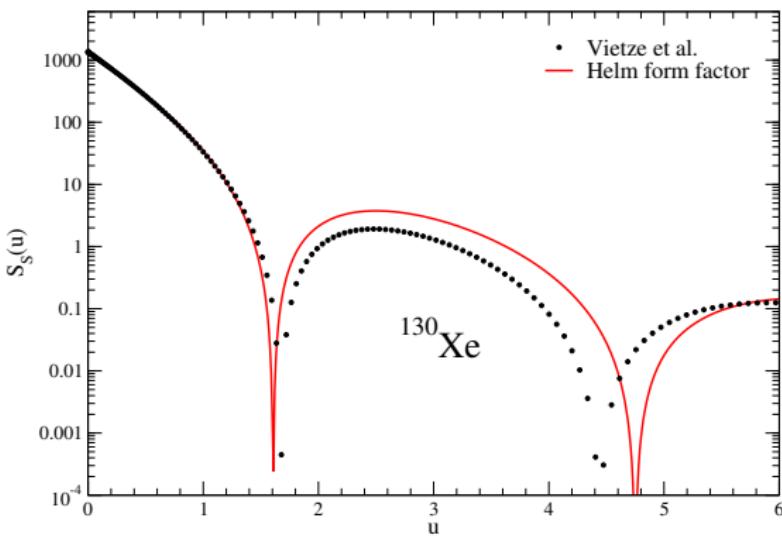
SI scattering completely
dominated by $L = 0$,
except at its minima

Vietze, Klos, JM, Haxton, Schwenk PRD91 043520 (2015)

Comparison to Helm form factor

Phenomenological Helm form factors,
based on a model of constant nuclear density with Gaussian surface,
are used in experimental data analysis of SI scattering

$$S_S^{\text{Helm}}(q) = S_S(0) \left(\frac{3 j_1(qr)}{qr} \right)^2 e^{-(qs)^2}$$



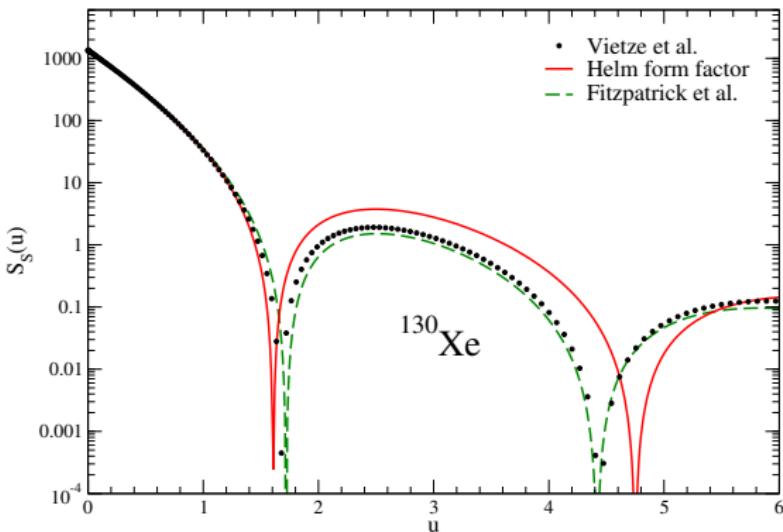
At low momentum transfers
good agreement between
state-of-the-art calculation
and Helm form factor

Small differences in amplitude
and location of minima appear
at higher momentum transfers

SI scattering not very sensitive
to nuclear structure details

Comparison to Fitzpatrick et al.

Fitzpatrick et al. JCAP02 004 (2013) also calculated Xe with the shell model but using older shell model calculations and larger valence space truncations



Shell model calculations by Vietze et al., Fitzpatrick et al. give almost identical SI structure factors

SI scattering not very sensitive to nuclear structure details

Vietze, Klos, JM, Haxton, Schwenk PRD91 043520 (2015)

Outline

1 Nuclear Structure of Xenon Isotopes

2 Spin-Independent WIMP scattering

3 Spin-Dependent WIMP scattering

Spin-Dependent scattering

Spin-Dependent (SD) interaction: WIMPs couple to the nuclear spin

The interaction Lagrangian at low-energies is given by the axial-vector leptonic $j_\mu(\mathbf{r})$ and axial-vector hadronic $J_\mu^A(\mathbf{r})$ currents:

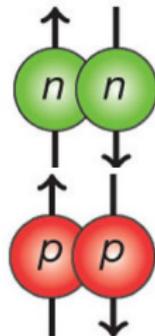
$$\mathcal{L}_\chi^{\text{SD}} = \frac{G_F}{\sqrt{2}} \int d^3\mathbf{r} j_\mu(\mathbf{r}) J_\mu^A(\mathbf{r})$$

Pairing interaction: pairs of spins couple to $S = 0$

No coherence for SD scattering

Only stable nuclei with odd neutrons/protons relevant for experiment searches:
for Xenon ^{129}Xe and ^{131}Xe

Specially sensitive to nuclear structure,
distribution of spin among nucleons

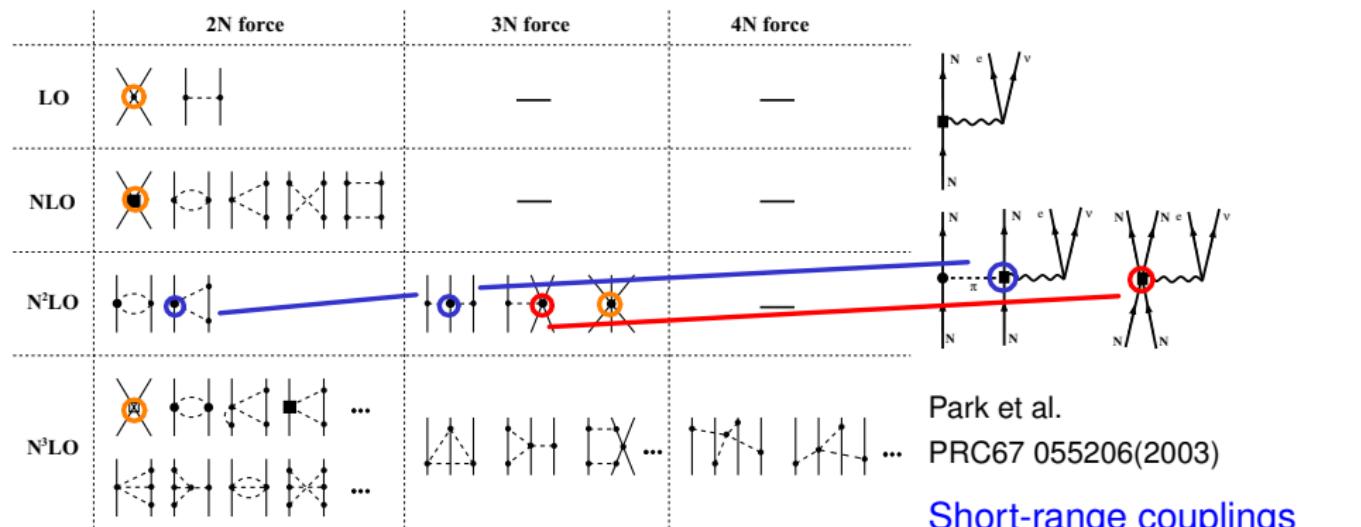


Chiral effective field theory

Chiral EFT: low energy approach to QCD, nuclear structure energies

Approximate chiral symmetry: pion exchanges, contact interactions

Systematic expansion: nuclear forces and electroweak currents



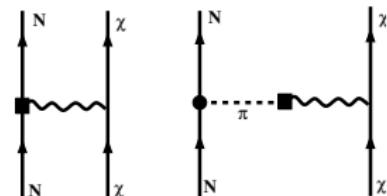
Weinberg, van Kolck, Kaplan, Savage, Weise, Epelbaum, Meißner...

SD scattering: 1b currents

At lowest orders in chiral EFT, 1b current agrees with phenomenological derivation Engel et al. IJMPE1 1(1992)

$$Q^0 : \sum_{i=1}^A \mathbf{J}_{i,1b} = \sum_{i=1}^A \frac{1}{2} \left[a_0 \sigma_i + a_1 \tau_i^3 \sigma_i \right],$$

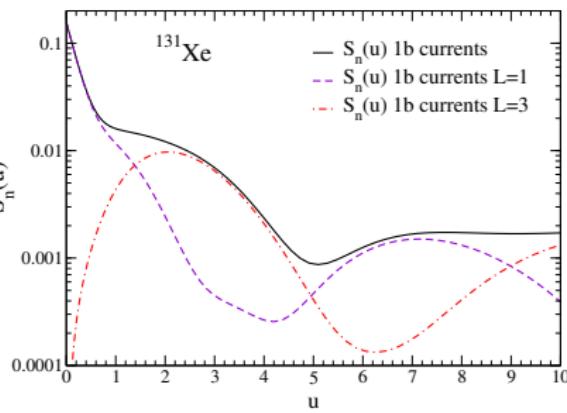
$$Q^2 : \sum_{i=1}^A \mathbf{J}_{i,1b} = \sum_{i=1}^A \frac{1}{2} \left[a_0 \sigma_i + a_1 \tau_i^3 \left(\frac{g_A(p^2)}{g_A} \sigma_i - \frac{g_P(p^2)}{2m g_A} (\mathbf{p} \cdot \sigma_i) \mathbf{p} \right) \right],$$



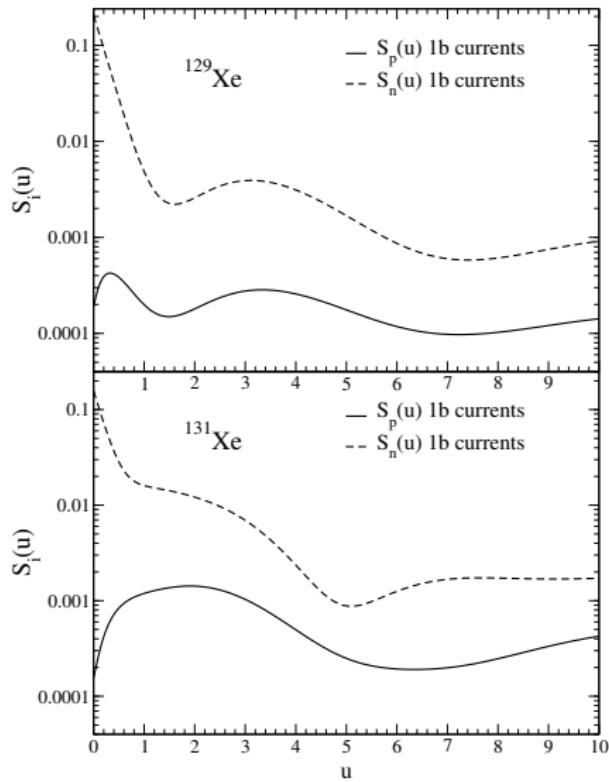
Multipole decomposition of structure factor:

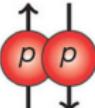
$$S_A(p) = \sum_{L \geq 0} \left| \langle J_f | \mathcal{L}_L^5 | J_i \rangle \right|^2 + \sum_{L \geq 1} \left| \langle J_f | \mathcal{T}_L^{e15} | J_i \rangle \right|^2 \quad S_A(u)$$

$L = 1$ multipole dominates at low p , then various multipoles contribute



SD Structure Factors with 1b currents




 $\text{In } ^{129,131}_{54}\text{Xe } \langle \mathbf{S}_n \rangle \gg \langle \mathbf{S}_p \rangle$,
 Neutrons carry most nuclear spin

$$\mathbf{S}_n = \sum_{i=1}^N \boldsymbol{\sigma}_i / 2, \quad \mathbf{S}_p = \sum_{i=1}^Z \boldsymbol{\sigma}_i / 2$$

$$\frac{S_A(0)}{2J+1} = \frac{(J+1)}{\pi J} |a_p \langle \mathbf{S}_p \rangle + a_n \langle \mathbf{S}_n \rangle|^2$$

$$a_{n/p} = (a_0 \mp a_1)/2,$$

$$S_n(0) \propto |\langle \mathbf{S}_n \rangle|^2 \quad S_p(0) \propto |\langle \mathbf{S}_p \rangle|^2.$$

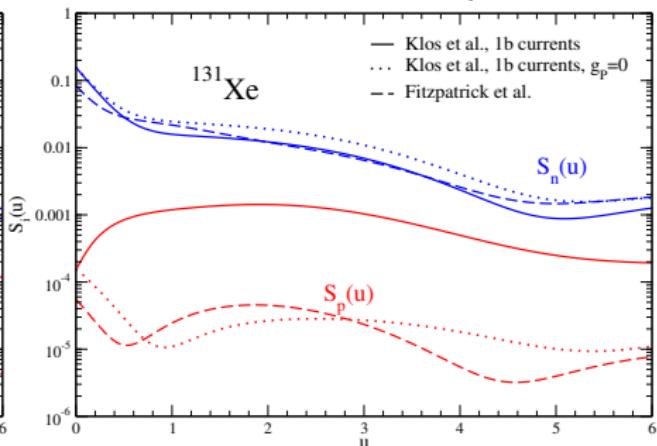
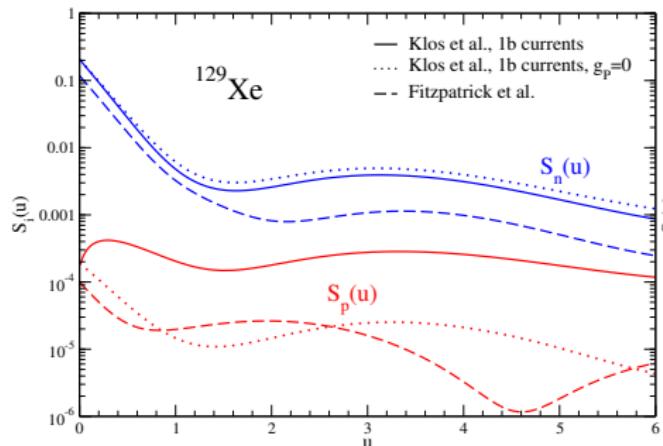
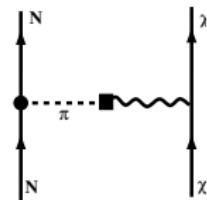
Couplings more sensitive to
protons ($a_0 = a_1$) or neutrons ($a_0 = -a_1$)

JM, Gazit, Schwenk, PRD86 103511(2012)

Klos, JM, Gazit, Schwenk, PRD88 083516(2013)

Comparison to Fitzpatrick et al.

SD structure factor compared to Fitzpatrick et al. JCAP02 004 (2013) at the level of 1b currents, with no pseudoscalar term, in JCAP02 004 (2013) coupling independent (O_6) to axial term (O_4)



Differences between shell model calculations larger than in SI case:
SD scattering sensitive to nuclear structure details: $\langle \mathbf{S}_n \rangle, \langle \mathbf{S}_p \rangle$

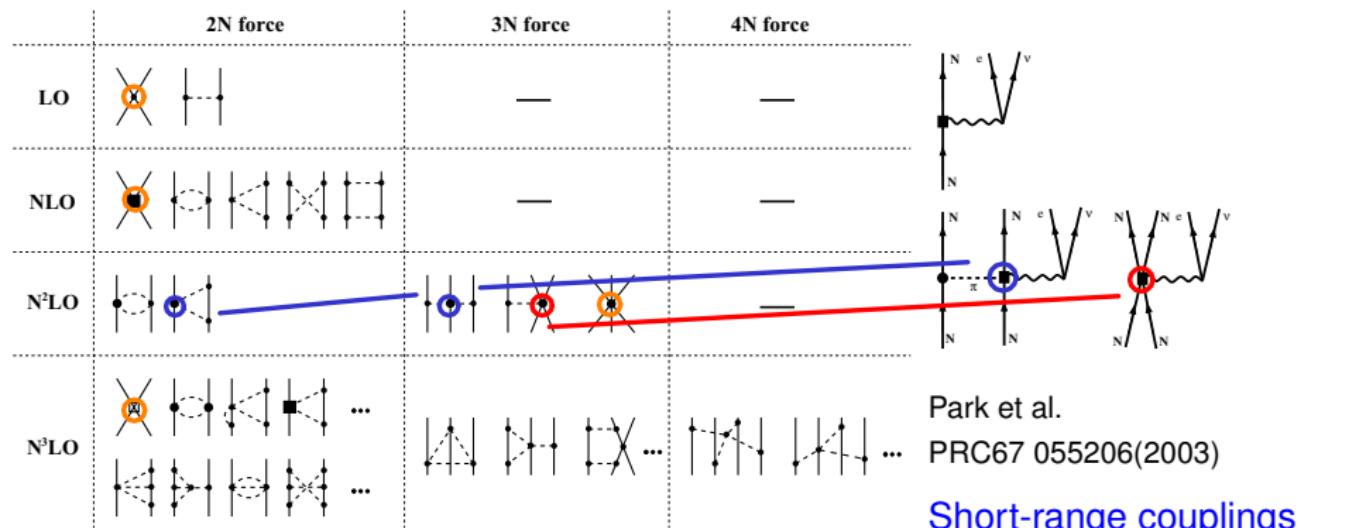
Reasonable agreement between different calculations: nuclear structure not limiting in extracting of relevant physics from WIMP scattering data

Chiral effective field theory

Chiral EFT: low energy approach to QCD, nuclear structure energies

Approximate chiral symmetry: pion exchanges, contact interactions

Systematic expansion: nuclear forces and electroweak currents



Weinberg, van Kolck, Kaplan, Savage, Weise, Epelbaum, Meißner...

2b currents and light nuclei

2b currents (meson-exchange currents) tested in light nuclei:

^3H β decay

Gazit, Quaglioni, Navrátil
PRL103 102502(2009)

$A \leq 9$ magnetic moments

Pastore et al. PRC87 035503(2013) \Rightarrow

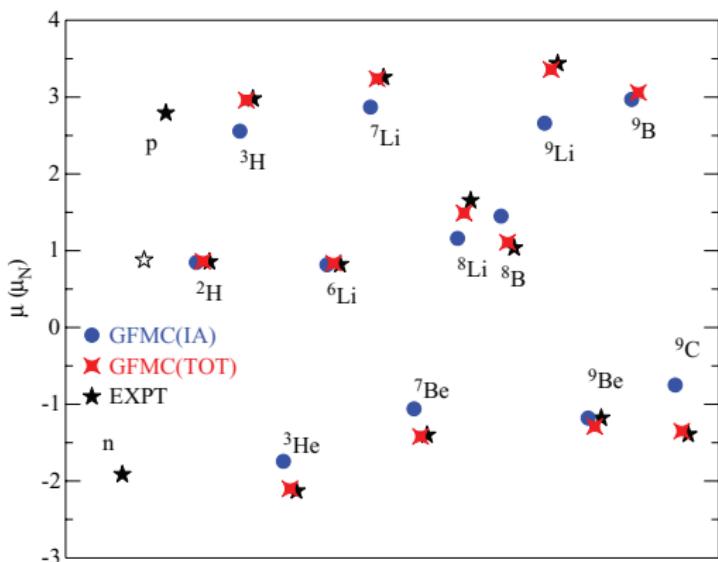
^8Be EM transitions

Pastore et al. PRC90 024321(2014)

^3H μ capture

Gazit PLB666 472(2008)

Marcucci et al. PRC83 014002(2011)



2b currents studied in electromagnetic and weak sectors

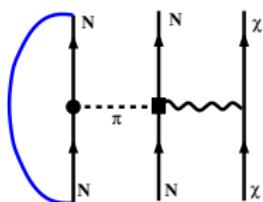
Chiral EFT 2b currents

Leading correction: 2b currents

Approximate in medium-mass nuclei: normal-ordered 1b part
with respect to spin/isospin symmetric Fermi gas

$$\begin{aligned}\mathbf{J}_{12}^3 = & -\frac{g_A}{2F_\pi^2} (\tau_1 \times \tau_2)^3 \left(\frac{\sigma_2 \cdot \mathbf{k}_2}{m_\pi^2 + k_2^2} \left[c_4(\sigma_1 \times \mathbf{k}_2) + \frac{\hat{c}_6}{12m} (\sigma_1 \times \mathbf{q}) - (1 \leftrightarrow 2) \right] \right) \\ & - \frac{g_A}{F_\pi^2} c_3 \left[\tau_1^3 \frac{(\sigma_1 \cdot \mathbf{k}_1) \mathbf{k}_1}{m_\pi^2 + k_1^2} + \tau_2^3 \frac{(\sigma_2 \cdot \mathbf{k}_2) \mathbf{k}_2}{m_\pi^2 + k_2^2} \right]\end{aligned}$$

Normal-ordered long-range two-body currents:



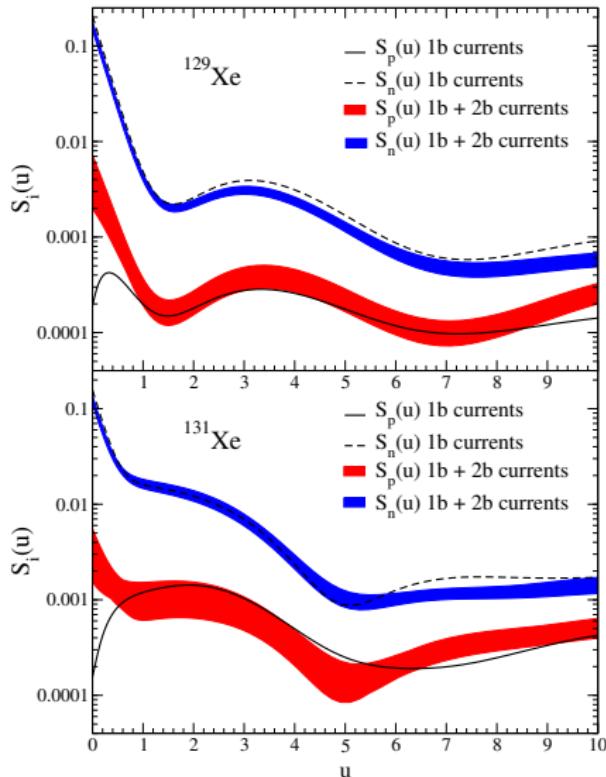
$$\mathbf{J}_{i,2b}^{\text{eff}} = -g_A \frac{\tau_i^3}{2} \frac{\rho}{3F_\pi^2} \left[c_4 \hat{l}_{12}^\sigma(\rho, p) - c_3 l_1^\sigma(\rho, p) - \frac{\hat{c}_6}{4m} l_{c6}(\rho, p) \right] \sigma_i = -g_A \frac{\tau_i^3}{2} \delta a_1 \sigma_i$$

$$\mathbf{J}_{i,2b}^{\text{eff}, P} = -g_A \frac{\tau_i^3}{2} \frac{\rho}{F_\pi^2} \left[\frac{-2c_3 p^2}{m_\pi^2 + p^2} + \frac{c_3 + c_4}{3} l^P(\rho, p) - \frac{\hat{c}_6}{12m} l_{c6}(\rho, p) \right] (\mathbf{p} \cdot \sigma_i) \mathbf{p} = -g_A \frac{\tau_i^3}{2} \frac{\delta a_1^P(p^2)}{p^2} (\mathbf{p} \cdot \sigma_i) \mathbf{p}$$

2b currents renormalize isovector couplings:

axial (reduction $\sim 20\%$, $p = 0$), pseudoscalar (enhancement $\sim 40\%$, $p = m_\pi$)

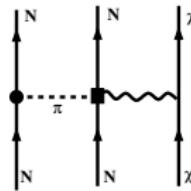
SD Structure Factors with 1b+2b currents



In $^{129,131}_{54}\text{Xe}$ $\langle S_n \rangle \gg \langle S_p \rangle$,
Neutrons carry most nuclear spin

Couplings more sensitive to
protons ($a_0 = a_1$) or neutrons ($a_0 = -a_1$)

2b currents naturally involve both
neutrons and protons:



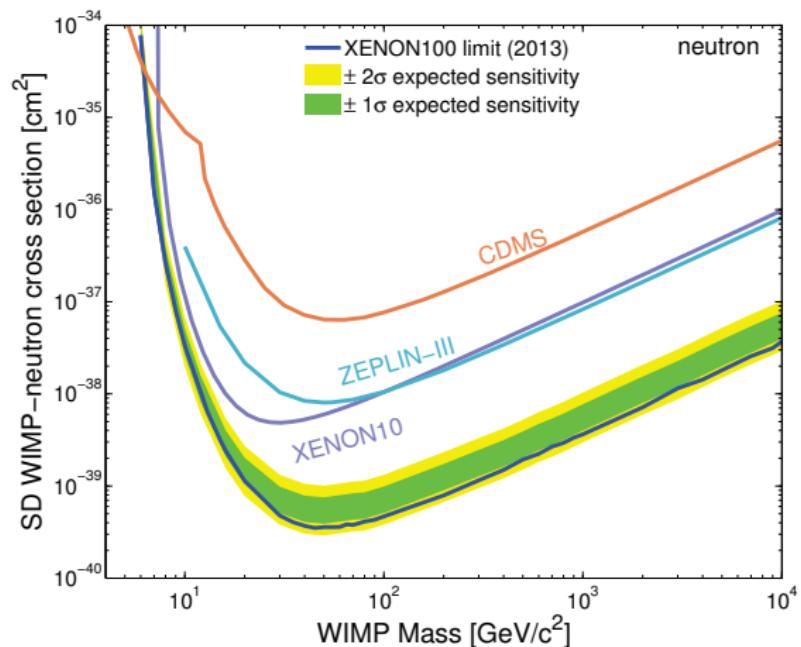
Neutrons always contribute with 2b
currents, dramatic increase in $S_p(u)$

See talk by P. Klos on Tuesday 24!

JM, Gazit, Schwenk, PRD86 103511(2012)

Klos, JM, Gazit, Schwenk, PRD88 083516(2013)

Application to experiment: XENON100



Our calculations used by XENON100 Collaboration to set limits on WIMP-nucleus cross-sections

XENON100 obtained world best limits for spin-dependent scattering with “neutron” couplings

Soon will be improved by LUX when they complete their spin-dependent analysis

Aprile et al. PRL111 021301 (2013)

Summary

WIMP scattering off Xenon for direct dark matter detection experiments

- State-of-the-art large-scale nuclear structure calculations
 - with tested valence spaces and nuclear interactions,
 - very good agreement with experimental spectra for all Xe isotopes
 - outlook: ab initio calculations based on chiral EFT forces
- Spin-Independent scattering: coherent enhancement
 - good agreement to phenomenological Helm form factor
 - not very sensitive to nuclear structure details
 - possible impact of SI 2b currents to be explored
- Spin-Dependent scattering:
 - more sensitive to nuclear structure: spin expectation values
 - 2b currents predicted by chiral EFT
 - Reduce the isovector structure factor at low p , $\sim 20\%$ at $p = 0$
 - Smaller reduction/enhancement at $p \sim m_\pi$,
depending on dominant multipoles (longitudinal enhanced)
 - Large increase in “proton” $S_A(p)$ (subleading species)