

Non-standard neutrino interactions with light mediators

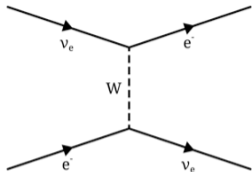
Pedro A N Machado

in collaboration with

K Babu, A Friedland, and I Mocioiu

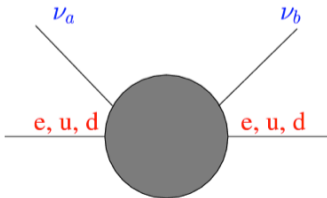


Neutrino matter effects and NSIs



$$\mathcal{H} = \frac{1}{2E} U \begin{pmatrix} 0 & 0 & 0 \\ 0 & \Delta m_{21}^2 & 0 \\ 0 & 0 & \Delta m_{31}^2 \end{pmatrix} U^\dagger + \frac{1}{2E} \begin{pmatrix} V & 0 & 0 \\ 0 & 0 & 0 \\ 0 & 0 & 0 \end{pmatrix}$$

New physics could induce new matter effects



Neutrino NSIs

Typically effective operators can be used to parametrize such interactions

Dimension 6 operators are expected to be dominant, for instance

$$\frac{1}{\Lambda^2} (\bar{L}_\alpha \gamma^\rho L_\beta) (\bar{L}_\gamma \gamma_\rho L_\delta)$$

SU(2) invariance ensures that **the same FCNCs in the neutrino sector are present in the charged lepton sector**

Neutrino NSIs

Now compare neutrino oscillation precision with rare decays

Neutrino NSIs

Now compare neutrino oscillation precision with rare decays

$$\text{BR}(\mu \text{ to } 3e) < 10^{-12}$$

$$\text{BR}(\tau \text{ to } 3\mu) < 2 \cdot 10^{-8}$$

$$\text{BR}(\tau \text{ to } 3e) < 3 \cdot 10^{-8}$$

Oscillation errors: 1~10%

Neutrino NSIs

Now compare neutrino oscillation precision with rare decays

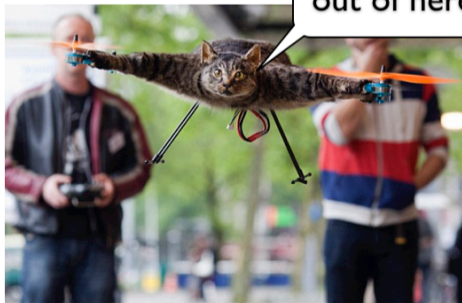
$$\text{BR}(\mu \text{ to } 3e) < 10^{-12}$$

$$\text{BR}(\tau \text{ to } 3\mu) < 2 \cdot 10^{-8}$$

$$\text{BR}(\tau \text{ to } 3e) < 3 \cdot 10^{-8}$$

Oscillation errors: 1~10%

That's it, I'm
out of here...



Neutrino NSIs

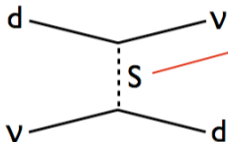
Is it possible to have large neutrino NSIs while respecting $SU(2)$ invariance and still being consistent with all flavor data?

Neutrino NSIs

One could try specific models, for instance:

Neutrino NSIs

One could try specific models, for instance:

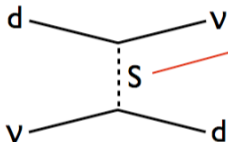


This would be a leptoquark scalar triplet with coupling $Q.S.L$

Different triplet components would induce FCNC for neutrinos and charged leptons

Neutrino NSIs

One could try specific models, for instance:



This would be a leptoquark scalar triplet with coupling $Q.S.L$

Different triplet components would induce FCNC for neutrinos and charged leptons

Different masses

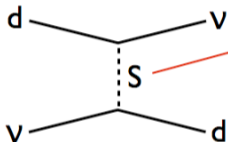
Different FCNC effects

T parameter...

Bergmann et al 1999

Neutrino NSIs

One could try specific models, for instance:



This would be a leptoquark scalar triplet with coupling $Q.S.L$

Different triplet components would induce FCNC for neutrinos and charged leptons

Different masses

Different FCNC effects

T parameter...

Bergmann et al 1999

This argument used to work on the last century, but now with more data (h, B factories) it is not valid anymore

Friedland et al 2012

Neutrino NSIs

Is it possible to have large neutrino NSIs while respecting SU(2) invariance and still being consistent with all flavor data?

D=6 operators induce too large FCNC for charged sector

Fine tuning would be needed to suppress D=6
while maintaining D=8 operators dominant

Gavela et al 2009

Even D=8 operators would be severely constrained

Antusch et al 2008

Neutrino NSIs and light mediators

There is another possibility: a weakly coupled light messenger

EFT would not apply

Generically, an experiment probes a mediator which mass is close to the experiment energy scale

$$A \propto \frac{1}{q^2 - m^2}$$

Neutrino matter effect is a zero momentum phenomenon

Neutrino NSIs and light mediators

We need **flavor changing** and **well behaved** interactions
We gauge $B - L$ number of the third family (call it X)

Neutrino NSIs and light mediators

We need **flavor changing** and **well behaved** interactions
We gauge $B - L$ number of the third family (call it X)

	q_1	q_2	q_3	\rightarrow X charge $1/3$	
q_1	•	•	•	\rightarrow Needs a doublet with X charge $1/3$	ϕ_2
q_2	•	•	•		
q_3	•	•	•	\rightarrow Needs a doublet with no X charge	ϕ_1

Due to details on the scalar potential and neutrino masses we also need a scalar singlet s with X charge $1/3$

Neutrino NSIs and light mediators

Due to the presence of a doublet that also has X charge, there is a mass mixing in the gauge boson sector

$$M_{\text{gauge}} = \frac{1}{4} \begin{pmatrix} g'^2 v^2 & -gg'v^2 & -g'g_X v_2^2/3 \\ -gg'v^2 & g^2 v^2 & gg_X v_2^2/3 \\ -g'g_X v_2^2/3 & gg_X v_2^2/3 & g_X^2 (v_2^2 + v_s^2)/9 \end{pmatrix}$$

Neutrino NSIs and light mediators

Due to the presence of a doublet that also has X charge, there is a mass mixing in the gauge boson sector

Usual B-W³ mass matrix

$$M_{\text{gauge}} = \frac{1}{4} \begin{pmatrix} g'^2 v^2 & -gg'v^2 & -g'g_X v_2^2/3 \\ -gg'v^2 & g^2 v^2 & gg_X v_2^2/3 \\ -g'g_X v_2^2/3 & gg_X v_2^2/3 & g_X^2 (v_2^2 + v_s^2)/9 \end{pmatrix}$$

Neutrino NSIs and light mediators

Due to the presence of a doublet that also has X charge, there is a mass mixing in the gauge boson sector

linear combination: massless state

$$M_{\text{gauge}} = \frac{1}{4} \begin{pmatrix} g'^2 v^2 & -gg'v^2 & -g'g_X v_2^2/3 \\ -gg'v^2 & g^2 v^2 & gg_X v_2^2/3 \\ -g'g_X v_2^2/3 & gg_X v_2^2/3 & g_X^2 (v_2^2 + v_s^2)/9 \end{pmatrix}$$

Neutrino NSIs and light mediators

Due to the presence of a doublet that also has X charge, there is a mass mixing in the gauge boson sector

Mixing between X and Z

$$M_{\text{gauge}} = \frac{1}{4} \begin{pmatrix} g'^2 v^2 & -gg'v^2 & -g'g_X v_2^2/3 \\ -gg'v^2 & g^2 v^2 & gg_X v_2^2/3 \\ -g'g_X v_2^2/3 & gg_X v_2^2/3 & g_X^2 (v_2^2 + v_s^2)/9 \end{pmatrix}$$

Neutrino NSIs and light mediators

Due to the presence of a doublet that also has X charge, there is a mass mixing in the gauge boson sector

$$M_{\text{gauge}} = \frac{1}{4} \begin{pmatrix} g'^2 v^2 & -gg'v^2 & -g'g_X v_2^2/3 \\ -gg'v^2 & g^2 v^2 & gg_X v_2^2/3 \\ -g'g_X v_2^2/3 & gg_X v_2^2/3 & g_X^2 (v_2^2 + v_s^2)/9 \end{pmatrix}$$

Neutrino NSIs and light mediators

Due to the presence of a doublet that also has X charge, there is a mass mixing in the gauge boson sector

$$M_{\text{gauge}} = \frac{1}{4} \begin{pmatrix} g'^2 v^2 & -gg'v^2 & -g'g_X v_2^2/3 \\ -gg'v^2 & g^2 v^2 & gg_X v_2^2/3 \\ -g'g_X v_2^2/3 & gg_X v_2^2/3 & g_X^2 (v_2^2 + v_s^2)/9 \end{pmatrix}$$

$$Z_\mu \approx -s_w B_\mu + c_w W_\mu^3 + s_X X_\mu^0 \quad M_X^2 = \frac{1}{36} g_X^2 \left(\frac{v_1^2 v_2^2}{v^2} + v_s^2 \right)$$

$$X_\mu \approx -s_X (-s_w B_\mu + c_w W_\mu^3) + X_\mu^0$$

$$s_X \equiv \frac{1}{3} \frac{g_X}{\sqrt{g^2 + g'^2}} \frac{v_2^2}{v^2}$$

Neutrino NSIs and light mediators

Due to the presence of a doublet that also has X charge, there is a mass mixing in the gauge boson sector

$$M_{\text{gauge}} = \frac{1}{4} \begin{pmatrix} g'^2 v^2 & -gg'v^2 & -g'g_X v_2^2/3 \\ -gg'v^2 & g^2 v^2 & gg_X v_2^2/3 \\ -g'g_X v_2^2/3 & gg_X v_2^2/3 & g_X^2 (v_2^2 + v_s^2)/9 \end{pmatrix}$$

$$Z_\mu \approx -s_w B_\mu + c_w W_\mu^3 + s_X X_\mu^0 \quad M_X^2 = \frac{1}{36} g_X^2 \left(\frac{v_1^2 v_2^2}{v^2} + v_s^2 \right)$$

$$X_\mu \approx -s_X (-s_w B_\mu + c_w W_\mu^3) + X_\mu^0$$

$$s_X \equiv \frac{1}{3} \frac{g_X}{\sqrt{g^2 + g'^2}} \frac{v_2^2}{v^2}$$

We could also have kinetic mixing

Neutrino NSIs and light mediators

$$\mathcal{L}_{yuk}^q = \bar{\mathbf{Q}}_L \begin{pmatrix} y_{11}^u \tilde{\phi}_1 & y_{12}^u \tilde{\phi}_1 & y_{13}^u \tilde{\phi}_2 \\ y_{21}^u \tilde{\phi}_1 & y_{22}^u \tilde{\phi}_1 & y_{23}^u \tilde{\phi}_2 \\ 0 & 0 & y_{33}^u \tilde{\phi}_1 \end{pmatrix} \mathbf{u}_R + \bar{\mathbf{Q}}_L \begin{pmatrix} y_{11}^d \phi_1 & y_{12}^d \phi_1 & 0 \\ y_{21}^d \phi_1 & y_{22}^d \phi_1 & 0 \\ y_{31}^d \phi_2 & y_{32}^d \phi_2 & y_{33}^d \phi_1 \end{pmatrix} \mathbf{d}_R + \text{h.c.}$$

Neutrino NSIs and light mediators

$$\mathcal{L}_{yuk}^q = \bar{\mathbf{Q}}_L \begin{pmatrix} y_{11}^u \tilde{\phi}_1 & y_{12}^u \tilde{\phi}_1 & y_{13}^u \tilde{\phi}_2 \\ y_{21}^u \tilde{\phi}_1 & y_{22}^u \tilde{\phi}_1 & y_{23}^u \tilde{\phi}_2 \\ 0 & 0 & y_{33}^u \tilde{\phi}_1 \end{pmatrix} \mathbf{u}_R + \bar{\mathbf{Q}}_L \begin{pmatrix} y_{11}^d \phi_1 & y_{12}^d \phi_1 & 0 \\ y_{21}^d \phi_1 & y_{22}^d \phi_1 & 0 \\ y_{31}^d \phi_2 & y_{32}^d \phi_2 & y_{33}^d \phi_1 \end{pmatrix} \mathbf{d}_R + \text{h.c.}$$

We choose to generate the CKM mixing on the up sector

$$R_{12}^{uL} \cdot M_u \cdot R_{12}^{uR\dagger} \approx \begin{pmatrix} m_u & 0 & V_{ub}^0 m_t \\ 0 & m_c & V_{cb}^0 m_t \\ 0 & 0 & m_t \end{pmatrix} \quad \text{and} \quad R_{12}^{dL} \cdot M_d \cdot R_{12}^{dR\dagger} \approx \begin{pmatrix} m_d^0 & 0 & 0 \\ 0 & m_s^0 & 0 \\ 0 & 0 & m_b^0 \end{pmatrix}$$

Neutrino NSIs and light mediators

$$\mathcal{L}_{yuk}^q = \bar{\mathbf{Q}}_L \begin{pmatrix} y_{11}^u \tilde{\phi}_1 & y_{12}^u \tilde{\phi}_1 & y_{13}^u \tilde{\phi}_2 \\ y_{21}^u \tilde{\phi}_1 & y_{22}^u \tilde{\phi}_1 & y_{23}^u \tilde{\phi}_2 \\ 0 & 0 & y_{33}^u \tilde{\phi}_1 \end{pmatrix} \mathbf{u}_R + \bar{\mathbf{Q}}_L \begin{pmatrix} y_{11}^d \phi_1 & y_{12}^d \phi_1 & 0 \\ y_{21}^d \phi_1 & y_{22}^d \phi_1 & 0 \\ y_{31}^d \phi_2 & y_{32}^d \phi_2 & y_{33}^d \phi_1 \end{pmatrix} \mathbf{d}_R + \text{h.c.}$$

We choose to generate the CKM mixing on the up sector

$$R_{12}^{uL} \cdot M_u \cdot R_{12}^{uR\dagger} \approx \begin{pmatrix} m_u & 0 & V_{ub}^0 m_t \\ 0 & m_c & V_{cb}^0 m_t \\ 0 & 0 & m_t \end{pmatrix} \quad \text{and} \quad R_{12}^{dL} \cdot M_d \cdot R_{12}^{dR\dagger} \approx \begin{pmatrix} m_d^0 & 0 & 0 \\ 0 & m_s^0 & 0 \\ 0 & 0 & m_b^0 \end{pmatrix}$$

$$\mathcal{L}_{X\text{-FCNC}} = \frac{g_X}{3} \bar{\mathbf{Q}}_L \begin{pmatrix} 0 & 0 & 0 \\ 0 & 0 & 0 \\ 0 & 0 & 1 \end{pmatrix} \gamma^\mu X_\mu \mathbf{Q}_L$$

$$\mathcal{L}_{X\text{-FCNC}} \approx \frac{g_X}{3} \bar{\mathbf{u}}_L \begin{pmatrix} V_{ub}^2 & V_{ub} V_{cb} & V_{ub} \\ V_{ub} V_{cb} & V_{cb}^2 & V_{cb} \\ V_{ub} & V_{cb} & 1 \end{pmatrix} \gamma^\mu X_\mu \mathbf{u}_L + \frac{g_X}{3} \bar{\mathbf{d}}_L \begin{pmatrix} 0 & 0 & 0 \\ 0 & 0 & 0 \\ 0 & 0 & 1 \end{pmatrix} \gamma^\mu X_\mu \mathbf{d}_L$$

Neutrino NSIs and light mediators

$$\mathcal{L}_{yuk}^q = \bar{\mathbf{Q}}_L \begin{pmatrix} y_{11}^u \tilde{\phi}_1 & y_{12}^u \tilde{\phi}_1 & y_{13}^u \tilde{\phi}_2 \\ y_{21}^u \tilde{\phi}_1 & y_{22}^u \tilde{\phi}_1 & y_{23}^u \tilde{\phi}_2 \\ 0 & 0 & y_{33}^u \tilde{\phi}_1 \end{pmatrix} \mathbf{u}_R + \bar{\mathbf{Q}}_L \begin{pmatrix} y_{11}^d \phi_1 & y_{12}^d \phi_1 & 0 \\ y_{21}^d \phi_1 & y_{22}^d \phi_1 & 0 \\ y_{31}^d \phi_2 & y_{32}^d \phi_2 & y_{33}^d \phi_1 \end{pmatrix} \mathbf{d}_R + \text{h.c.}$$

We choose to generate the CKM mixing on the up sector

$$R_{12}^{uL} \cdot M_u \cdot R_{12}^{uR\dagger} \approx \begin{pmatrix} m_u & 0 & V_{ub}^0 m_t \\ 0 & m_c & V_{cb}^0 m_t \\ 0 & 0 & m_t \end{pmatrix} \quad \text{and} \quad R_{12}^{dL} \cdot M_d \cdot R_{12}^{dR\dagger} \approx \begin{pmatrix} m_d^0 & 0 & 0 \\ 0 & m_s^0 & 0 \\ 0 & 0 & m_b^0 \end{pmatrix}$$

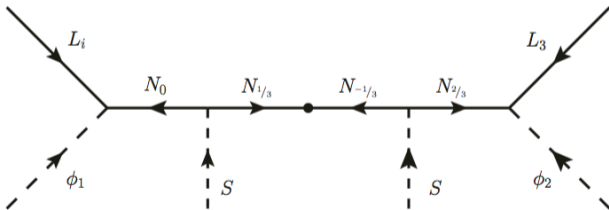
$$\mathcal{L}_{X\text{-FCNC}} = \frac{g_X}{3} \bar{\mathbf{Q}}_L \begin{pmatrix} 0 & 0 & 0 \\ 0 & 0 & 0 \\ 0 & 0 & 1 \end{pmatrix} \gamma^\mu X_\mu \mathbf{Q}_L$$

$$\mathcal{L}_{X\text{-FCNC}} \approx \frac{g_X}{3} \bar{\mathbf{u}}_L \begin{pmatrix} V_{ub}^2 & V_{ub} V_{cb} & V_{ub} \\ V_{ub} V_{cb} & V_{cb}^2 & V_{cb} \\ V_{ub} & V_{cb} & 1 \end{pmatrix} \gamma^\mu X_\mu \mathbf{u}_L + \frac{g_X}{3} \bar{\mathbf{d}}_L \begin{pmatrix} 0 & 0 & 0 \\ 0 & 0 & 0 \\ 0 & 0 & 1 \end{pmatrix} \gamma^\mu X_\mu \mathbf{d}_L$$

Neutrino NSIs and light mediators

PMNS mixing cannot be generated in the charged fermion sector:
tree level FCNC free!

Neutrino masses and mixing: Froggatt-Nielsen mechanism



Presence of TeV-ish fermion singlets

Phenomenology

Electron and muon anomalous magnetic moment

Fixed target experiments

Upsilon and kaon decays

B meson and kaon oscillations

LSND

NUTEV

Borexino and GEMMA

Neutrino oscillations

Phenomenology

Electron and muon anomalous magnetic moment

Fixed target experiments

Upsilon and kaon decays

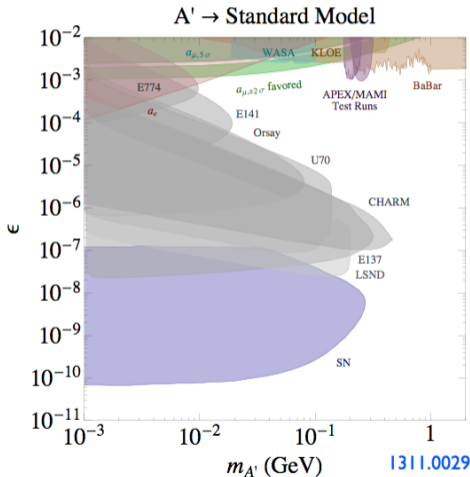
B meson and kaon oscillations

LSND

NUTEV

Borexino and GEMMA

Neutrino oscillations



Phenomenology

Electron and muon anomalous magnetic moment

Fixed target experiments

Upsilon and kaon decays

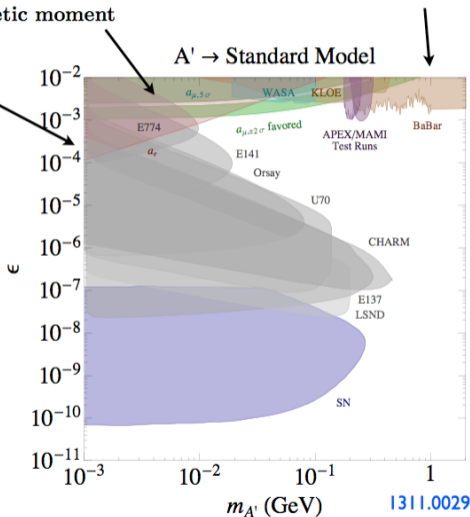
B meson and kaon oscillations

LSND

NUTEV

Borexino and GEMMA

Neutrino oscillations



Phenomenology

Electron and muon anomalous magnetic moment

Fixed target experiments

Upsilon and kaon decays

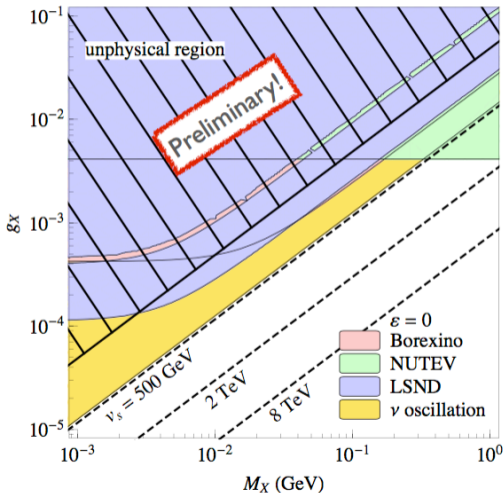
B meson and kaon oscillations

LSND

NUTEV

Borexino and GEMMA

Neutrino oscillations



Phenomenology

Electron and muon anomalous magnetic moment

Fixed target experiments

Upsilon and kaon decays $g_X \lesssim 0.018$

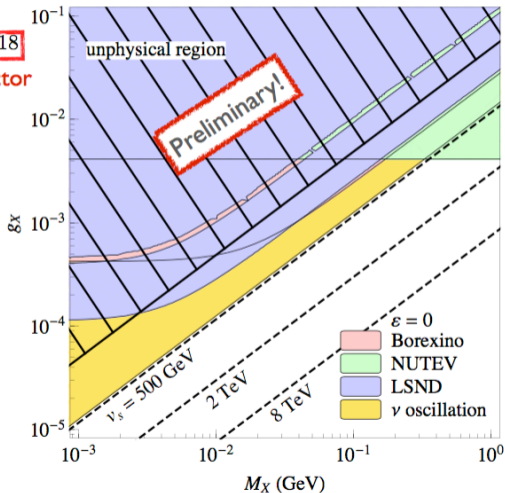
B meson and kaon oscillations **up sector**

LSND

NUTEV

Borexino and GEMMA

Neutrino oscillations



Phenomenology

Electron and muon anomalous magnetic moment

Fixed target experiments

Upsilon and kaon decays $g_X \lesssim 0.018$

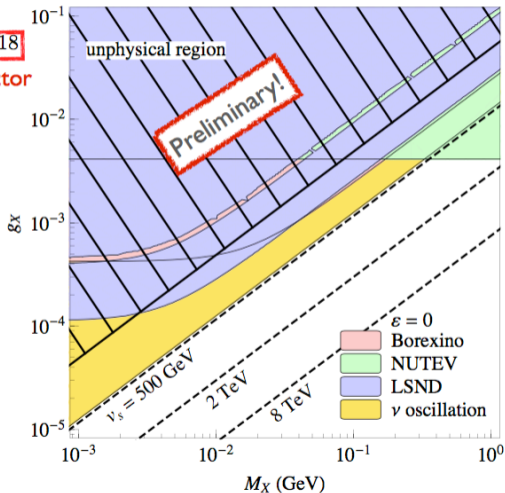
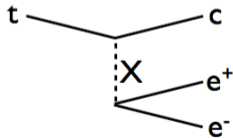
B meson and kaon oscillations **up sector**

LSND

NUTEV

Borexino and GEMMA

Neutrino oscillations



Conclusions

We elaborated a *renormalizable, UV complete* model which can provide large NSI in the neutrino sector and pass all flavor bounds

This simple realization may indicate a whole class of flavorful light mediator models which may have an interesting phenomenology

We found that the strongest bound comes from neutrino oscillation experiments