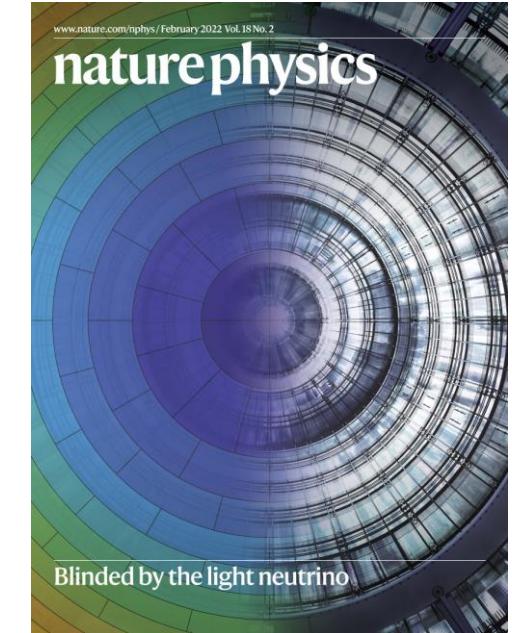


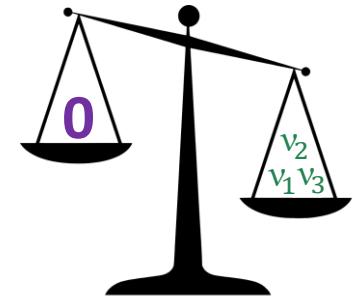
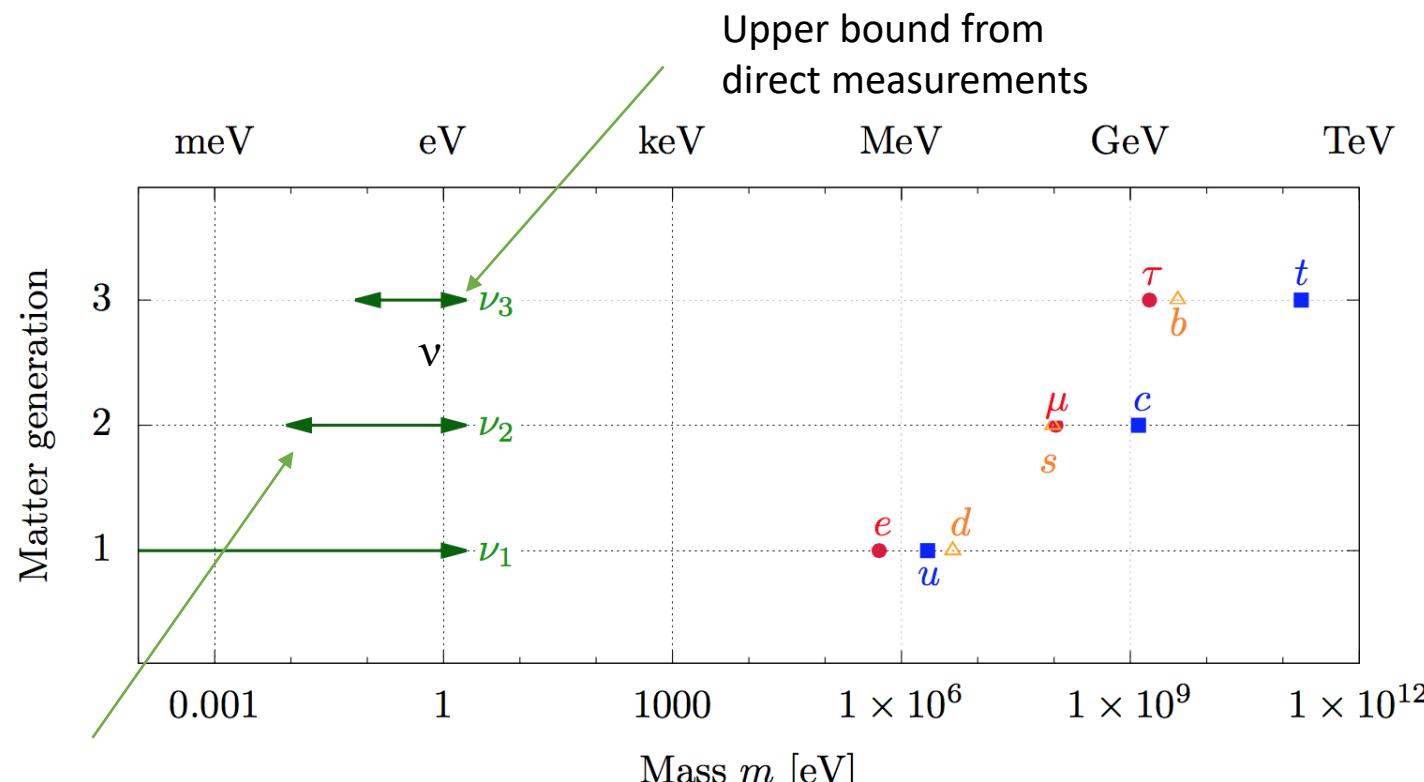
# Recent results from the KATRIN experiment



Christoph Köhler  
Technical University of Munich  
59<sup>th</sup> International Winter Meeting on Nuclear Physics



# Neutrino mass



- Neutrino oscillations → Non-zero mass
- Smallness and origin of mass?

# Determination methods

## Cosmology

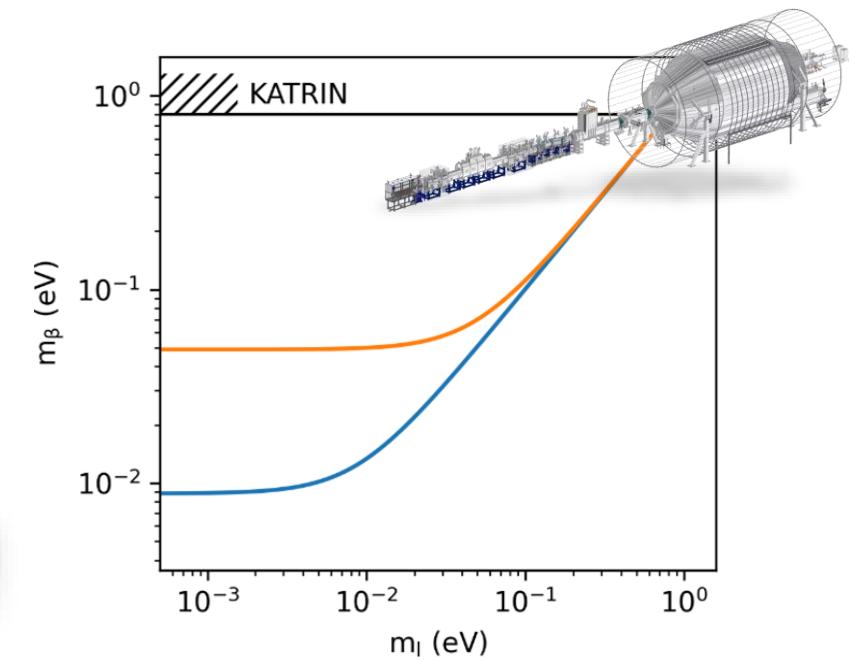
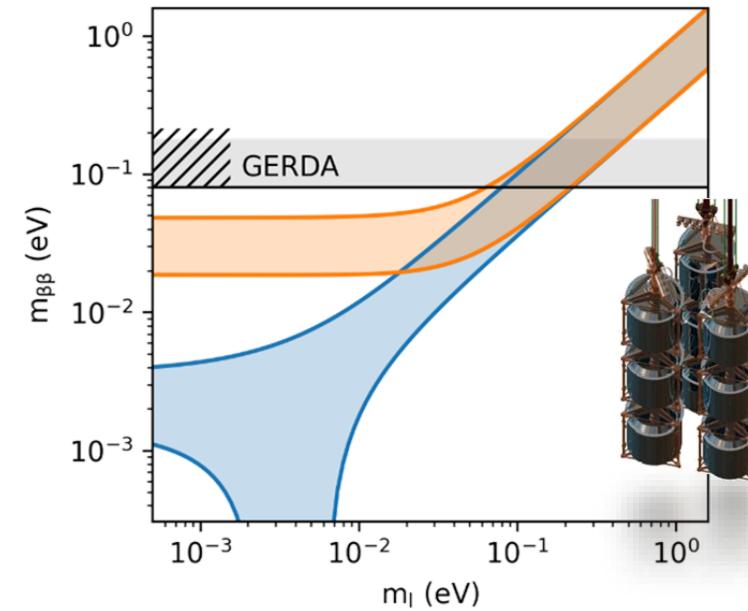
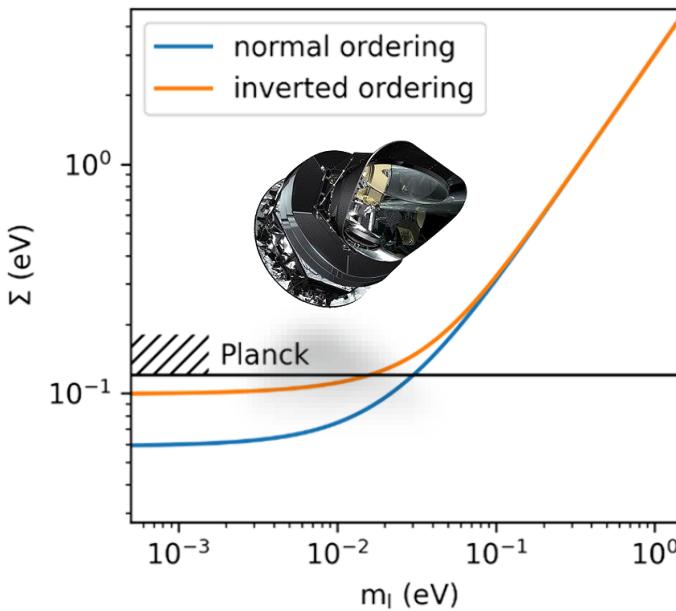
$$\Sigma = \sum_i m_i$$

## Neutrinoless $\beta\beta$ -decay

$$m_{\beta\beta} = \left| \sum_i U_{ei}^2 m_i \right|$$

## $\beta$ -decay kinematics

$$m_\beta = \sqrt{\sum_i |U_{ei}^2| m_i^2}$$



# Determination methods

## Cosmology

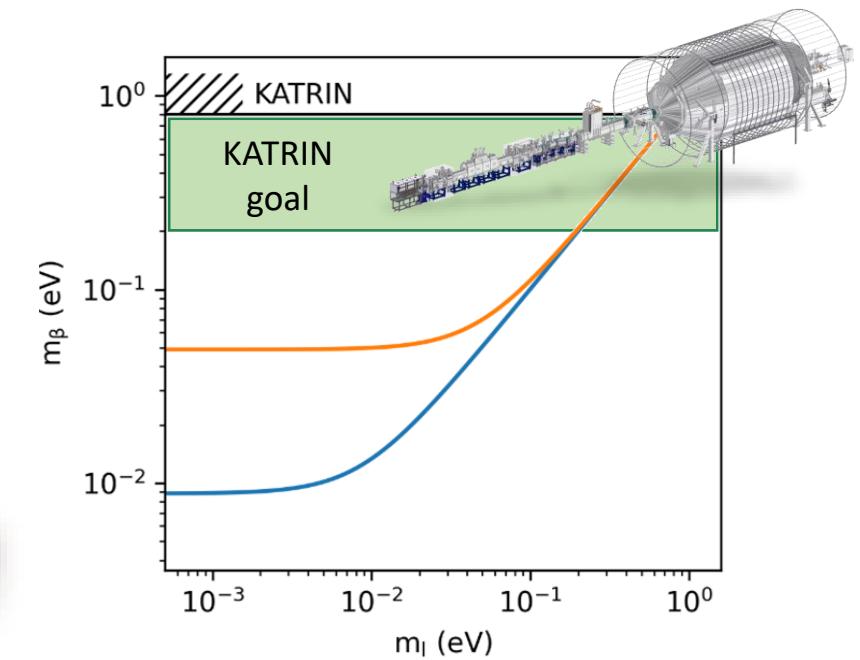
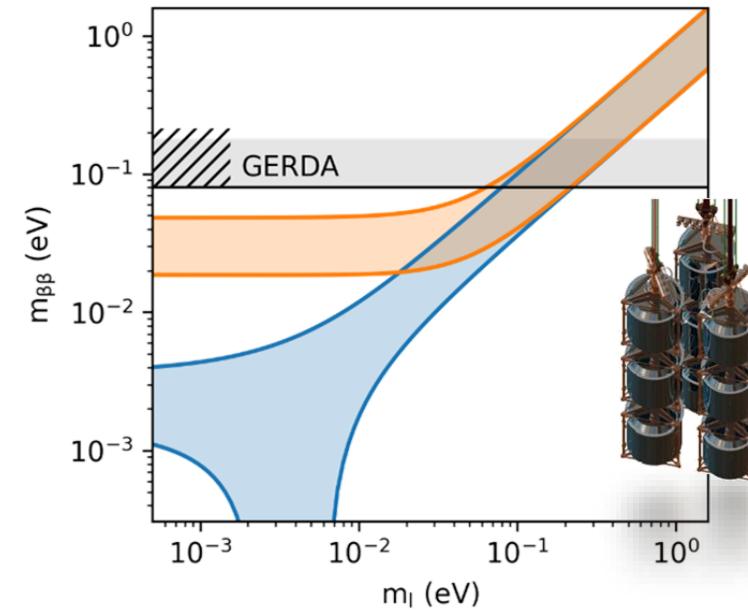
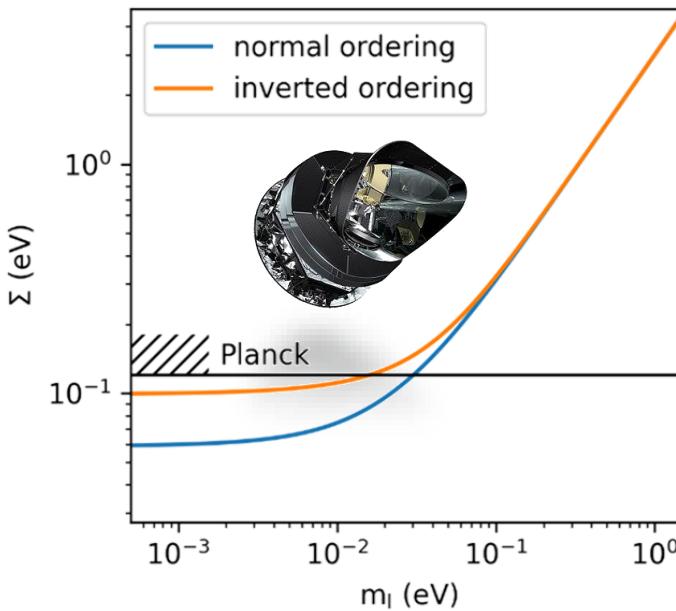
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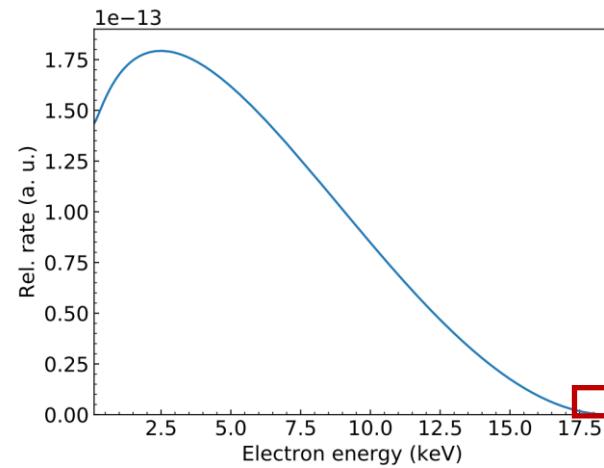
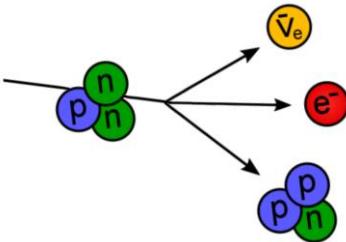
## $\beta$ -decay kinematics

$$m_\beta = \sqrt{\sum_i |U_{ei}^2| m_i^2}$$

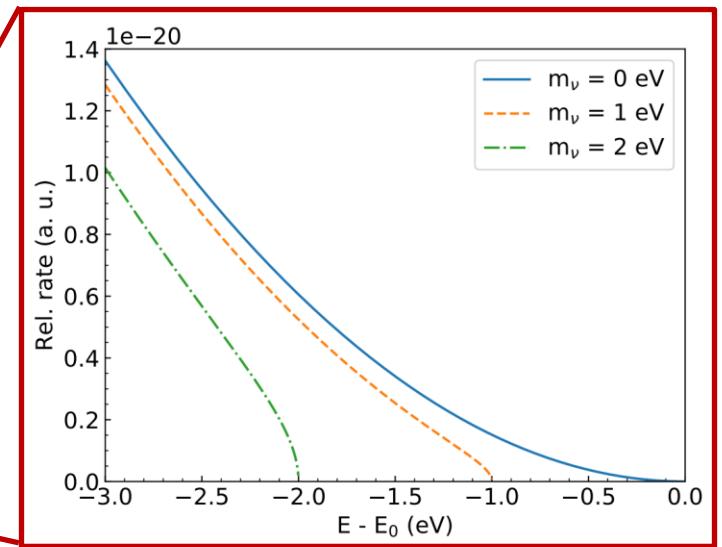


# Neutrino mass signal in $\beta$ -decay

- Tritium  $\beta$ -decay:
  - $E_0 = 18.56 \text{ keV}$
  - $T_{1/2} = 12.3 \text{ years}$



$$m^2(\nu_e) = \sum_i |U_{ei}|^2 \cdot m_i^2$$

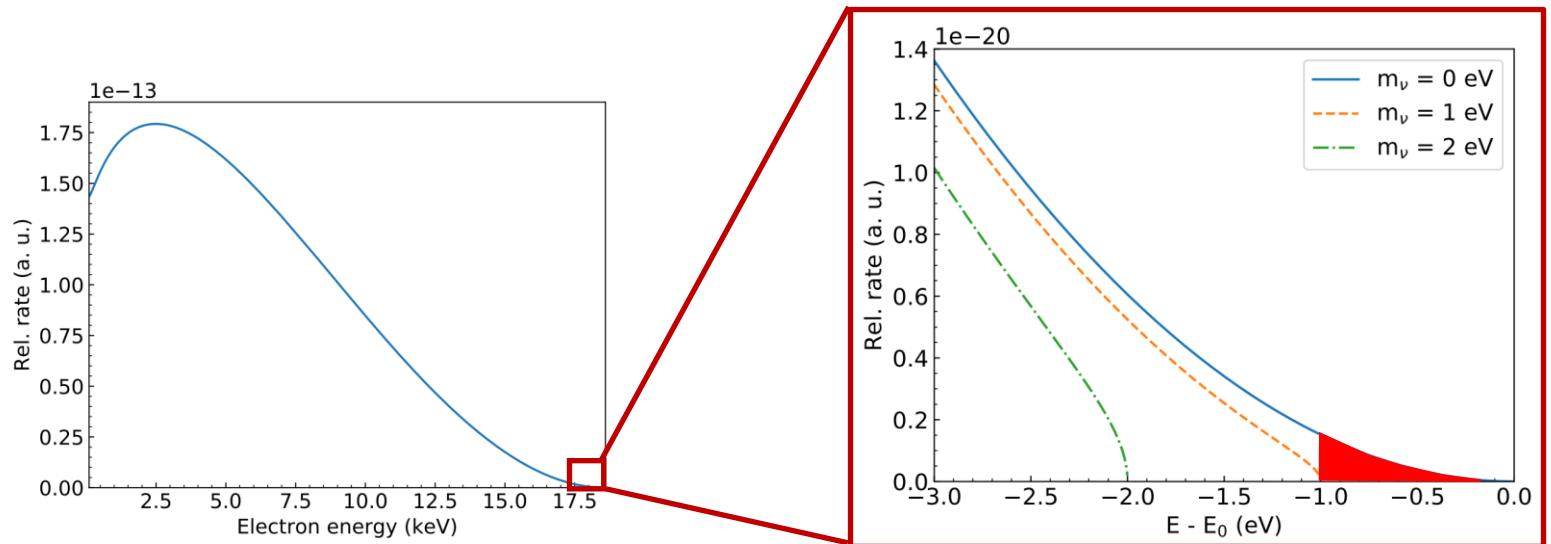


- Distortion of electron spectrum near the endpoint
- Independent of cosmology and neutrino nature

# Experimental challenge

- Only  $10^{-13}$  of all decays in last 1 eV,  
 $10^{-8}$  in the last 40 eV
- Strong tritium source:  $10^{11}$  decays/s
- Low background level < 0.1 cps
- Excellent energy resolution  $\sim 1$  eV
- Precise understanding of the spectrum shape

$$m^2(\nu_e) = \sum_i |U_{ei}|^2 \cdot m_i^2$$



# KArlsruhe TRItium Neutrino Experiment



- International collaboration (150 members)
- Design sensitivity: 0.2 eV (90 % C.L.)  
(1000 days of measurement time)



# Experimental overview

## Windowless gaseous tritium source

- 30 µg molecular tritium in closed loop
- $10^{11} T_2$  decay/s

## Transport section

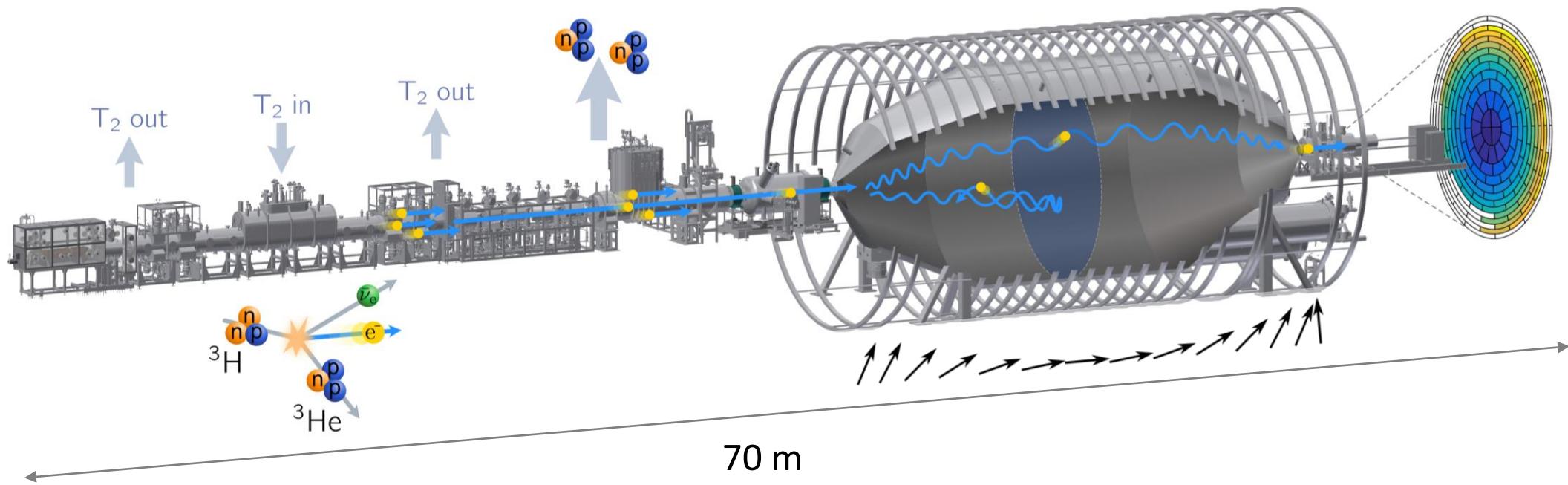
- Tritium gas/ion removal  $> 10^{14}$

## Spectrometer

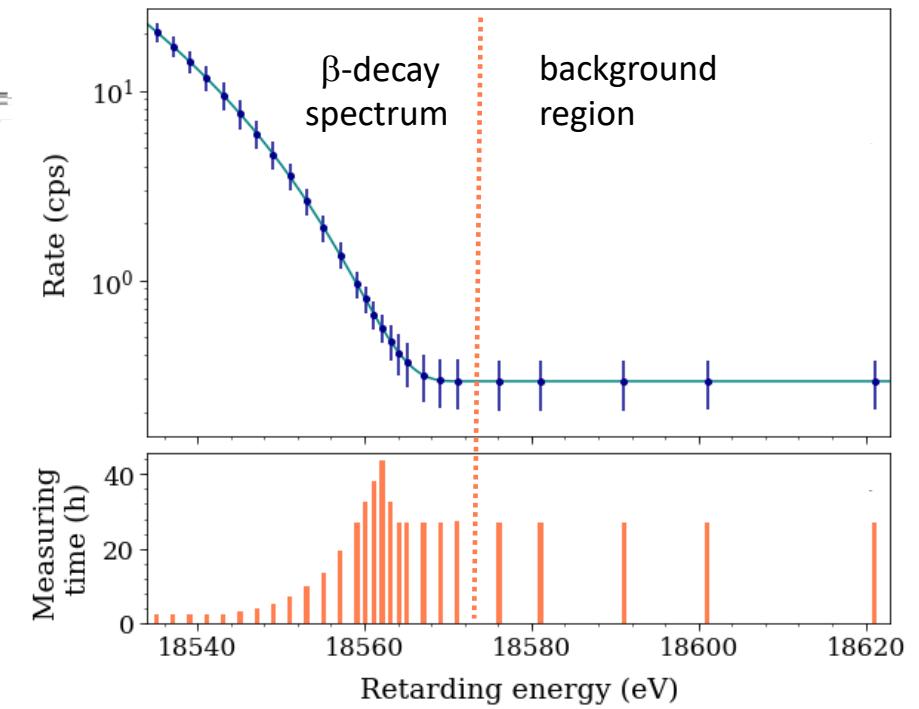
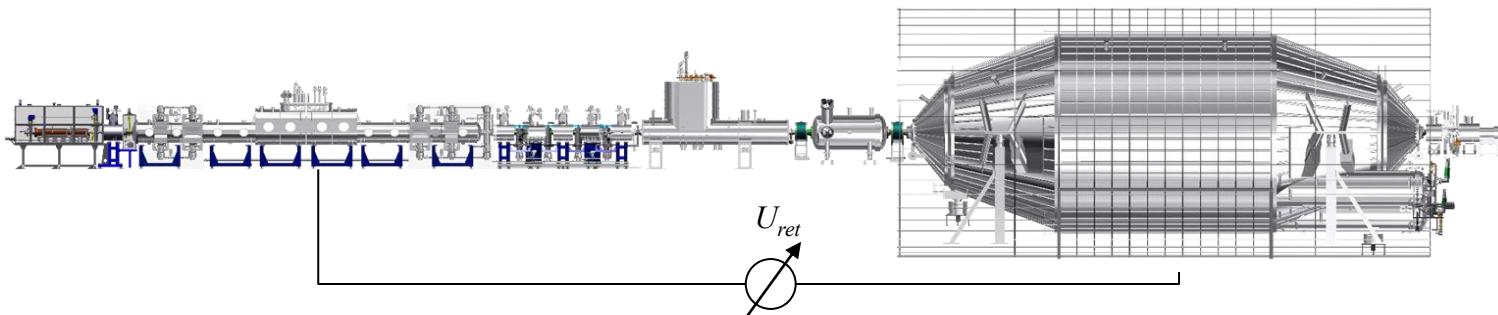
- MAC-E filter principle
- High resolution:  $O(1)$  eV

## Detector

- Focal plane detector, **148 pixel** PIN-diode
- Counts electrons



# Measurement strategy

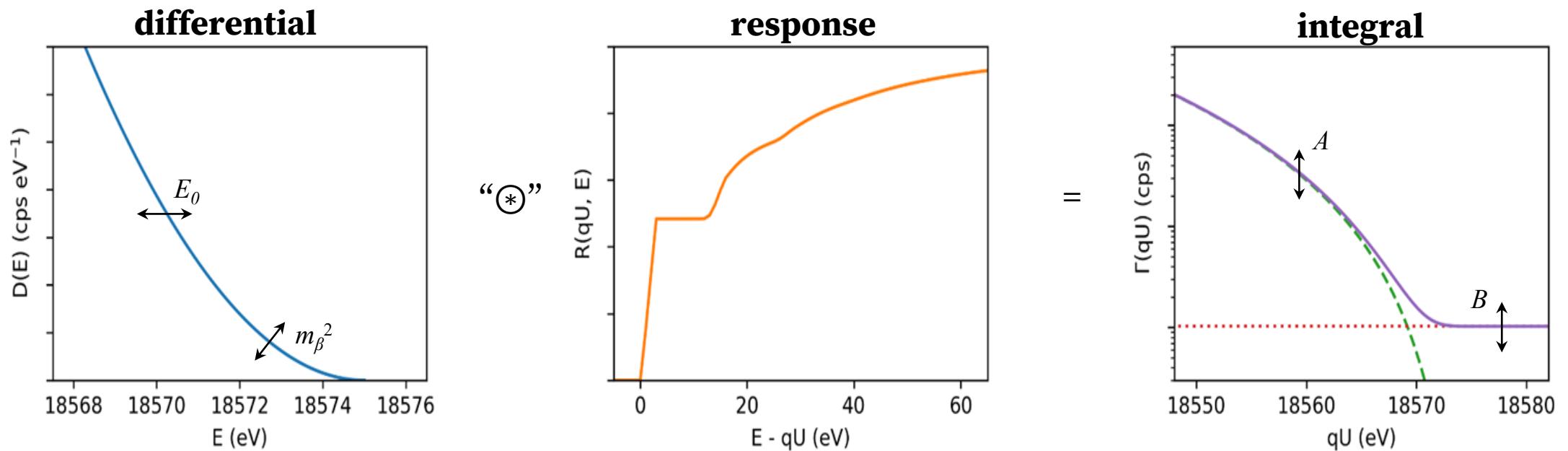


- **~30 HV set points** with varying duration
- Scan interval:  $E_0 - 40 \text{ eV}$ ,  $E_0 + 130 \text{ eV}$
- **~2 h scan time**
- Several **campaigns** per year O(100) scans

# Analysis strategy

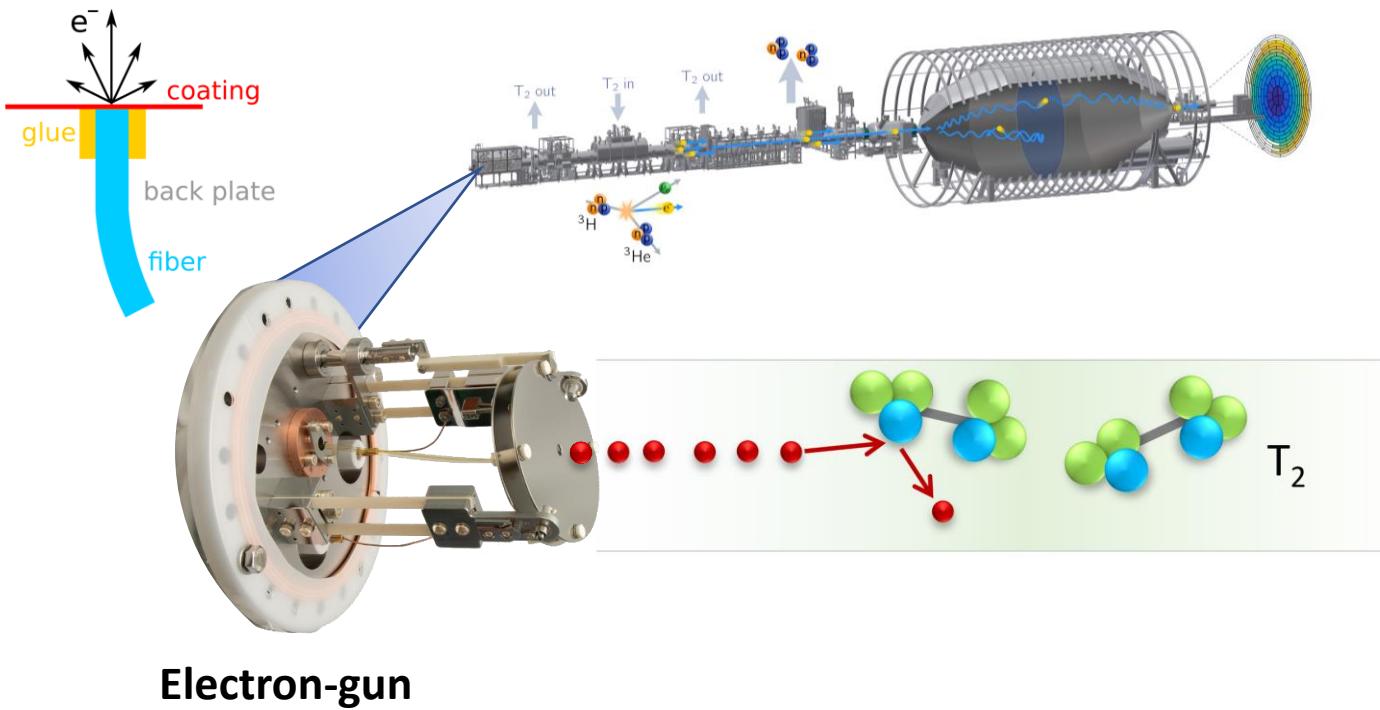
- Maximum likelihood fit of model

$$\Gamma(qU) \propto A \int_{qU}^{E_0} D(E; m_\beta^2, E_0) R(qU, E) dE + B$$



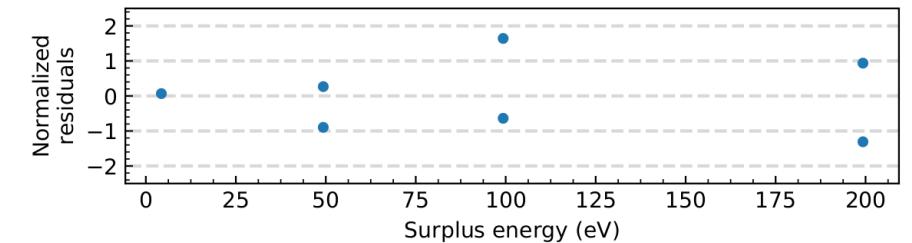
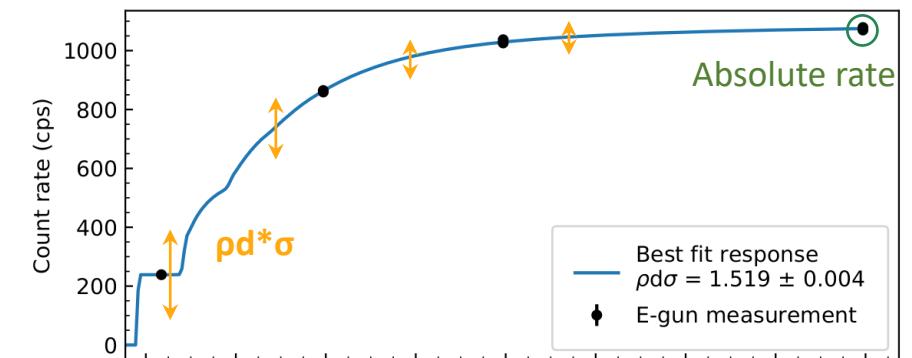
- Free parameters: Amplitude **A**, squared neutrino mass  $m_\beta^2$ , endpoint  $E_0$  and background **B**
- Theoretical** (Fermi theory, molecular excitations) and **experimental inputs** (calibration measurements)

# Column density determination

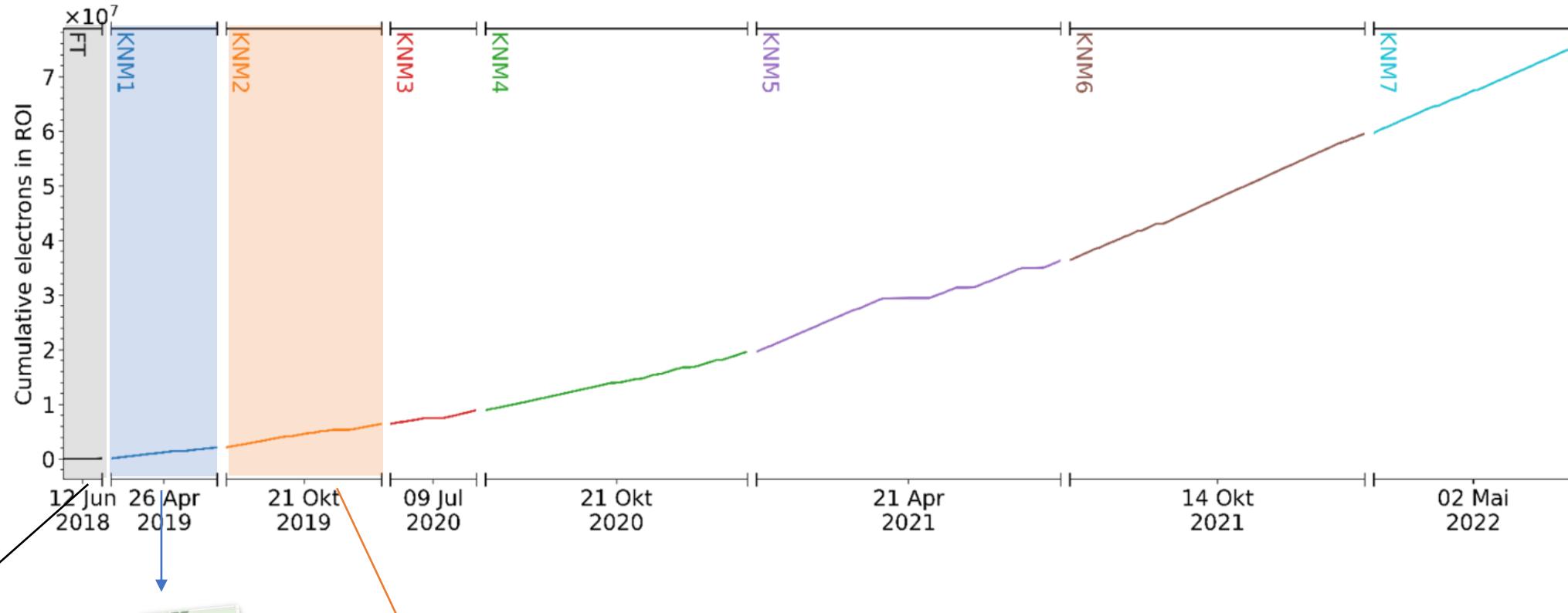


- Electron rate measured at different retarding potentials:  $\sim 30$  min
- Fit with model response function:
  - Absolute **electron rate**
  - **$\rho d^* \sigma$**  (column density \* cross section)

- Source scattering depending on:
    - Electron path
    - **Column density**
    - Cross section
- Input for **response function**



# Data taking overview

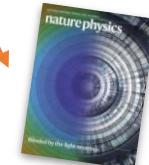


**Commissioning**  
[EPJ C 80, 264 (2020)]



**< 1.1 eV (90% CL)**

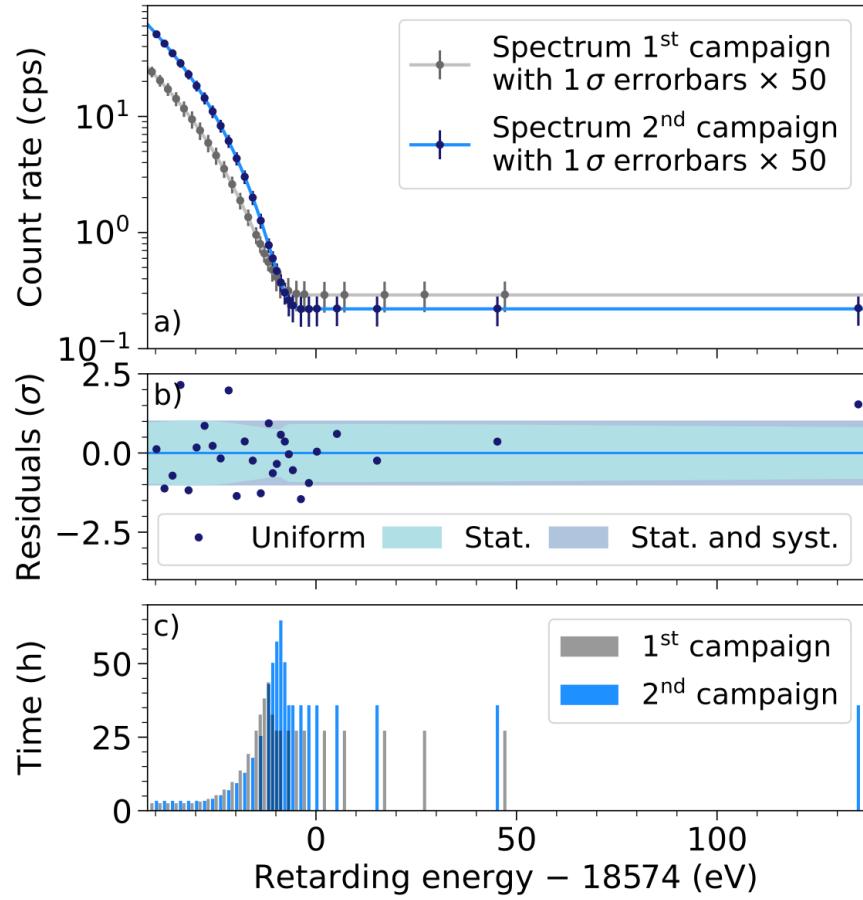
[Aker et al., PRL 123 (2019) 22, 221802]



**< 0.8 eV (90% CL)**

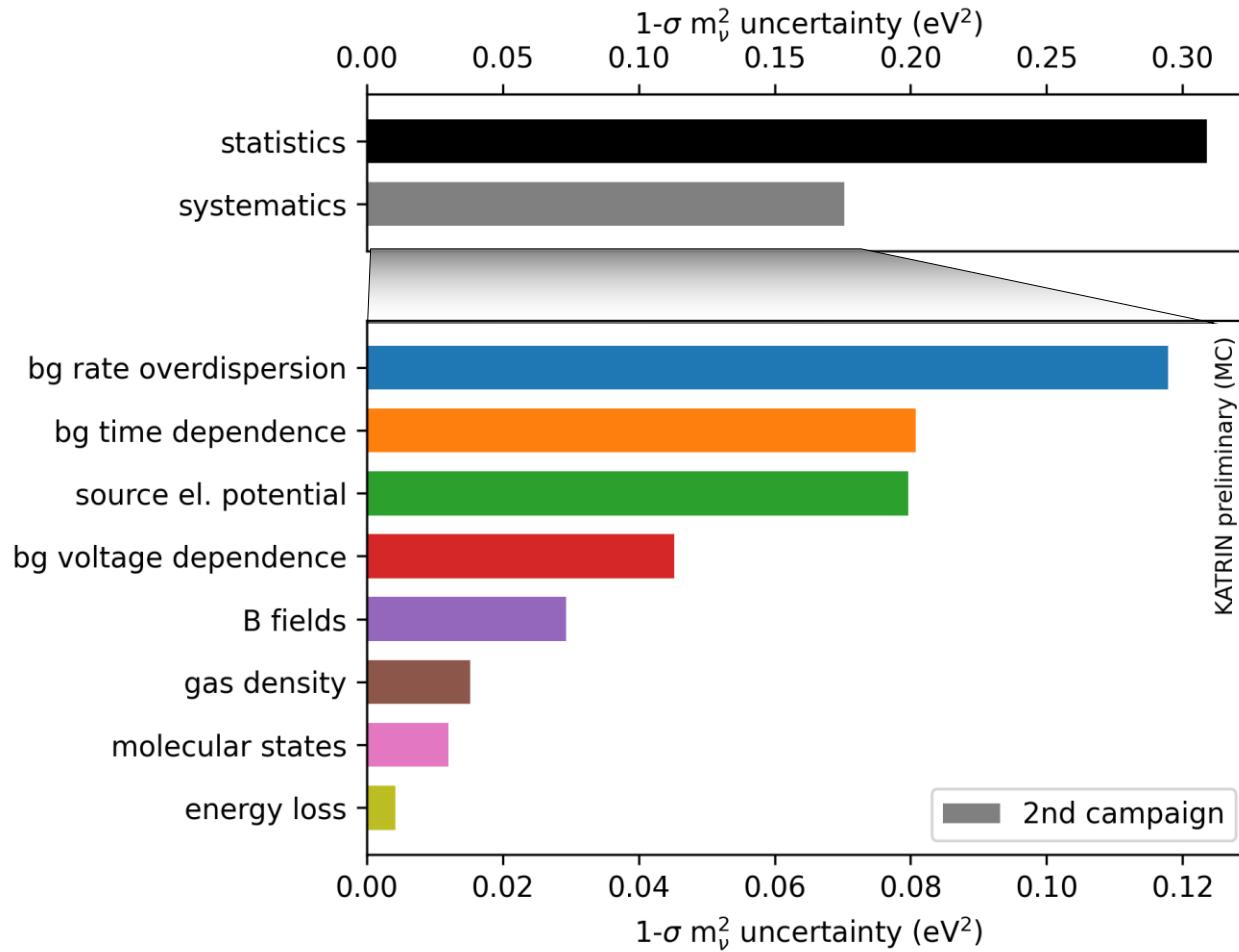
[Aker et al., Nature Phys. 18 (2022) 2, 160-166]

# Latest neutrino mass result

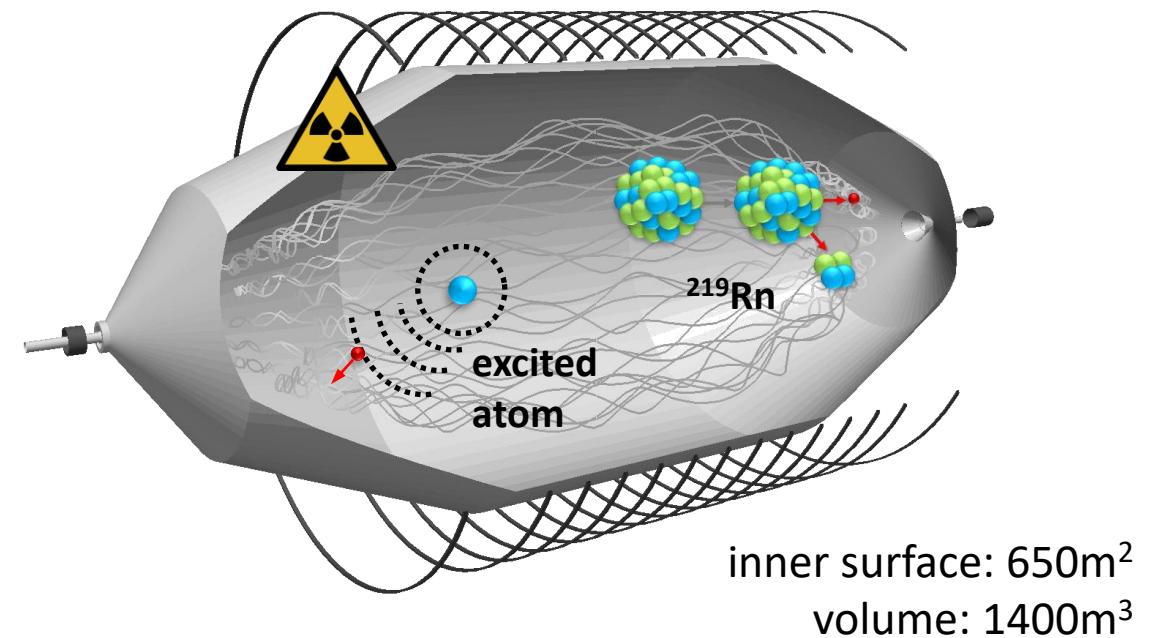


- KNM1:
  - Best fit:  $m_\nu^2 = (-1.0^{+0.9}_{-1.1}) \text{ eV}^2$
  - Limit:  $m_\nu < 1.1 \text{ eV (90\% CL)}$
- KNM2:
  - Best fit:  $m_\nu^2 = (0.26^{+0.34}_{-0.34}) \text{ eV}^2$
  - Limit:  $m_\nu < 0.9 \text{ eV (90\% CL)}$
- Combined result:  $m_\nu < 0.8 \text{ eV (90\% CL)}$

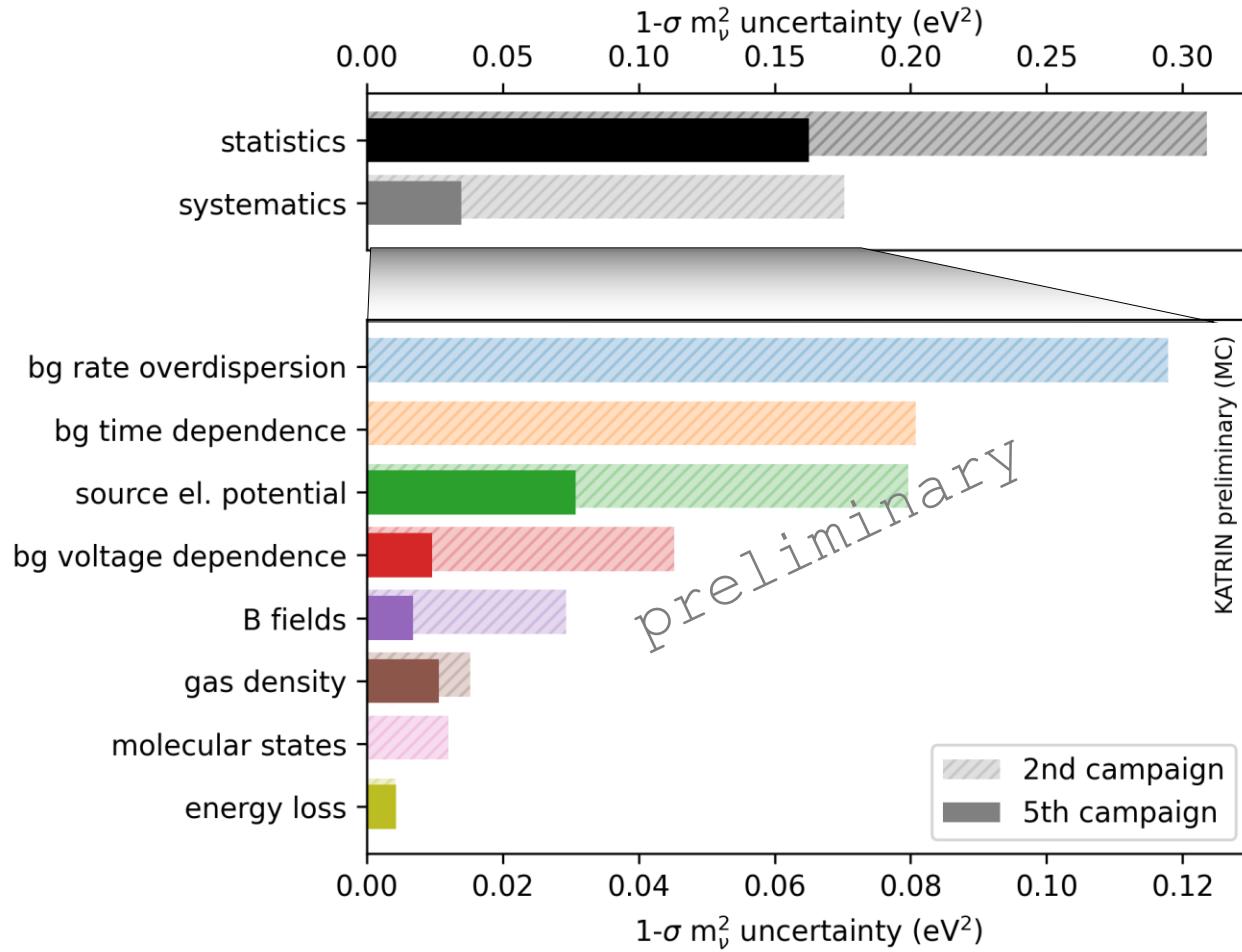
# Uncertainty breakdown: 2<sup>nd</sup> campaign



- **Statistical uncertainty dominates**, systematics non-negligible
- **Background-related uncertainties** dominate systematics budget

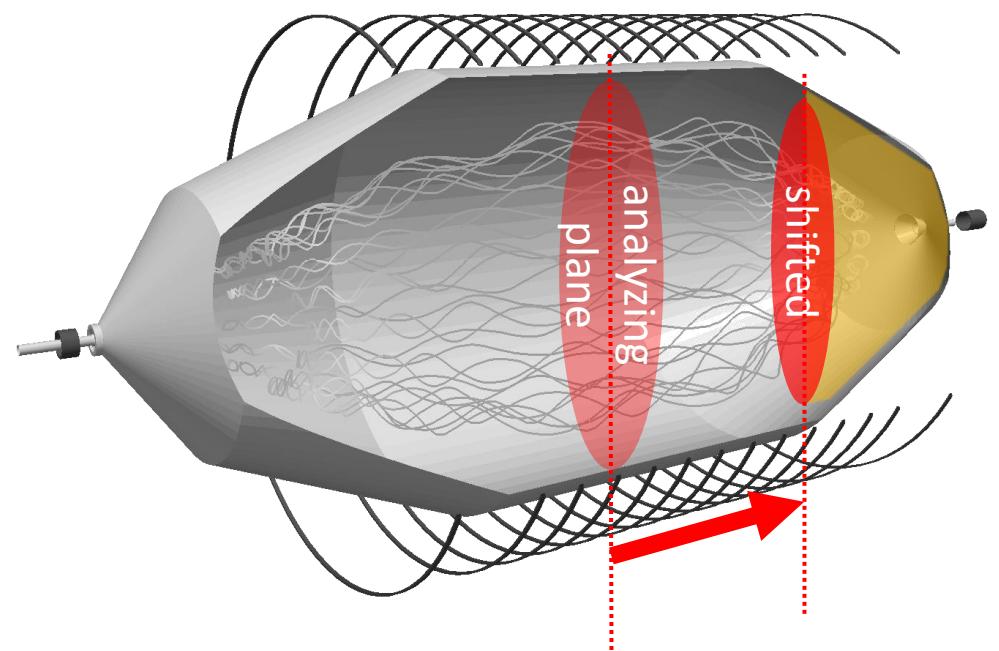


# Preview: 5<sup>th</sup> campaign

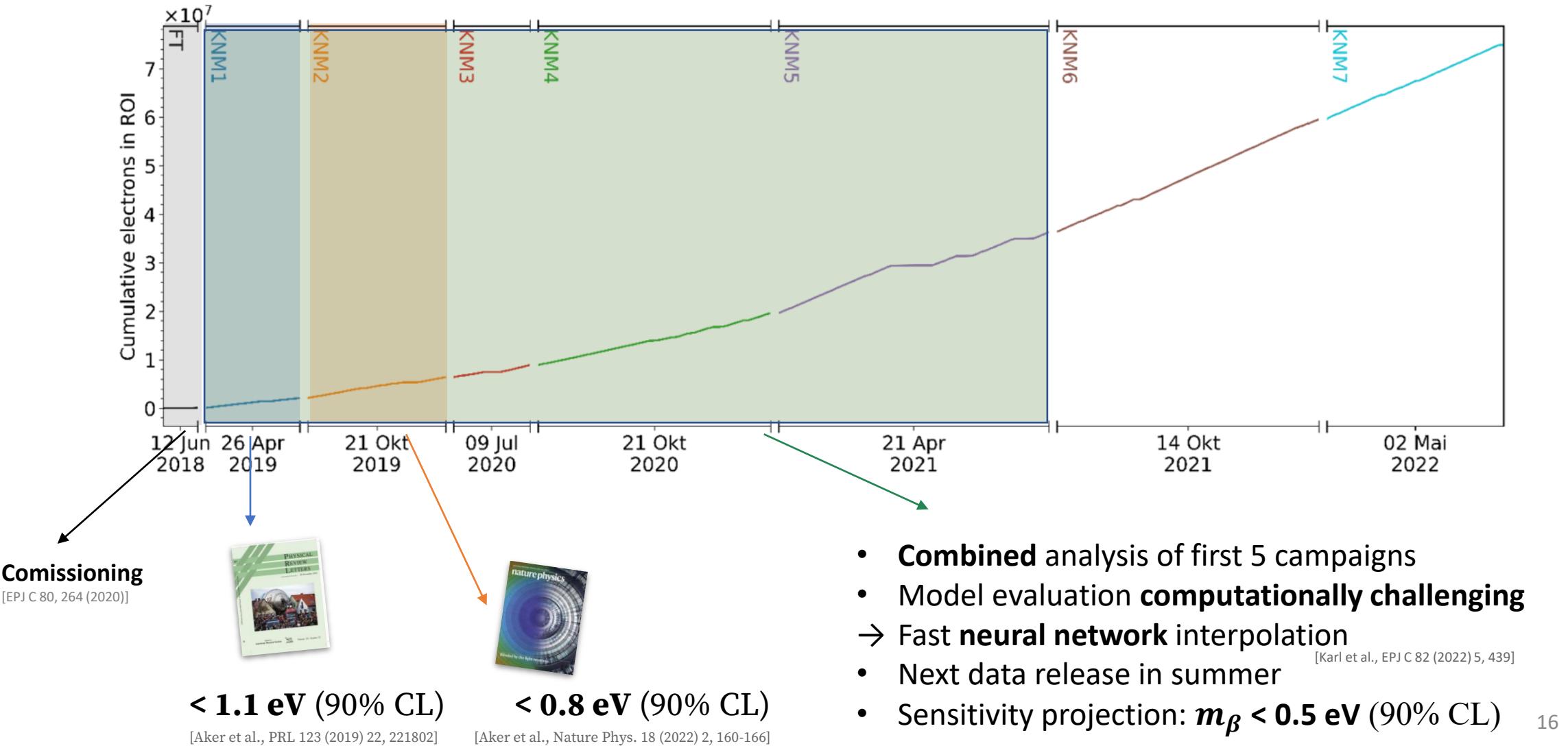


- Still statistics dominated, **systematics largely improved**
- **Major improvement:** Background reduction ( $\div 2$ ) via new EM field layout

Lokhov et al arXiv:2201.11743 (2022)

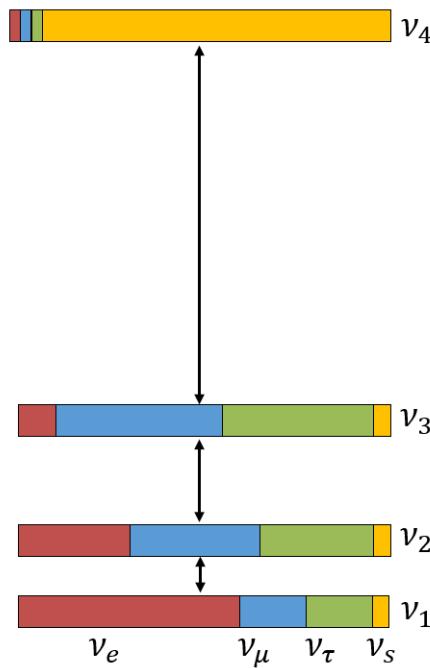


# Outlook

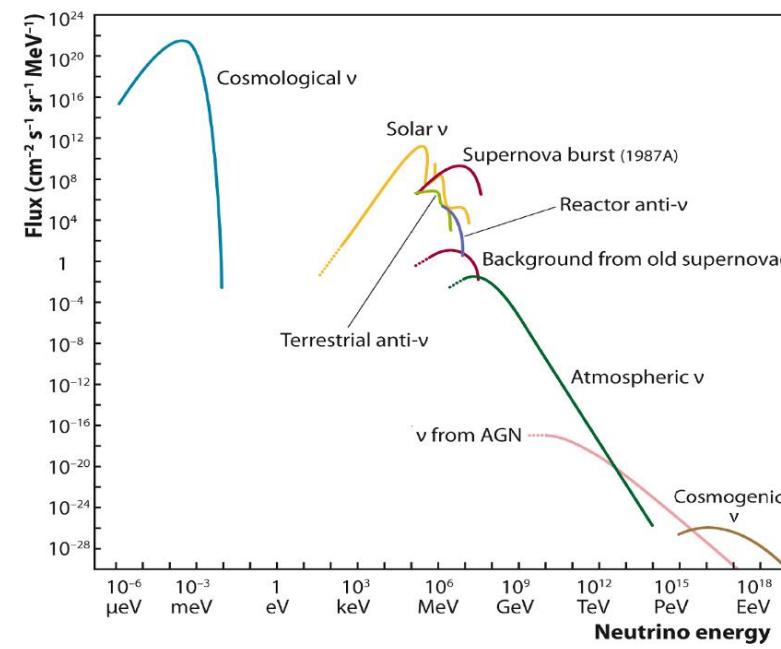


# Physics beyond neutrino mass

Search for light sterile neutrinos



Search for Big Bang neutrinos

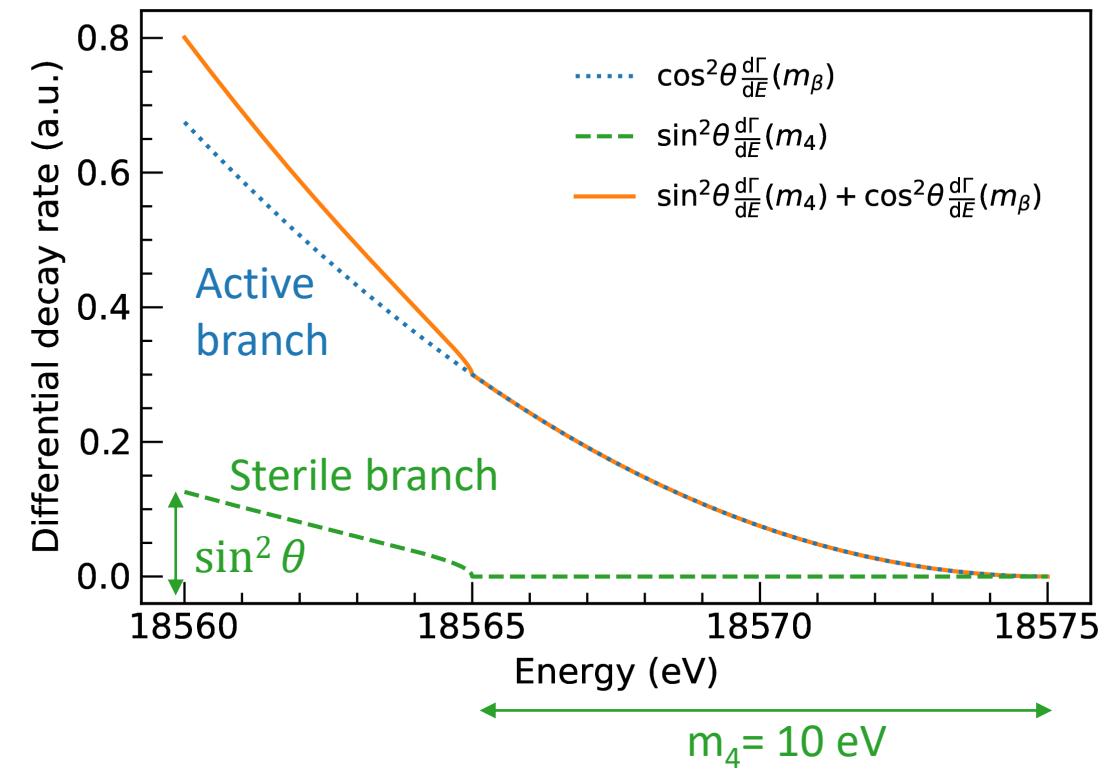
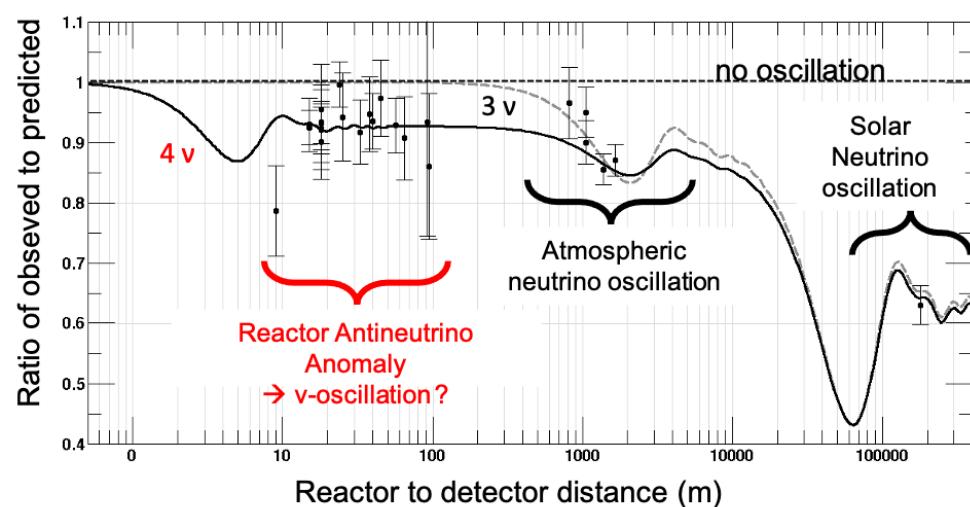


and more ...  
(not in this talk)

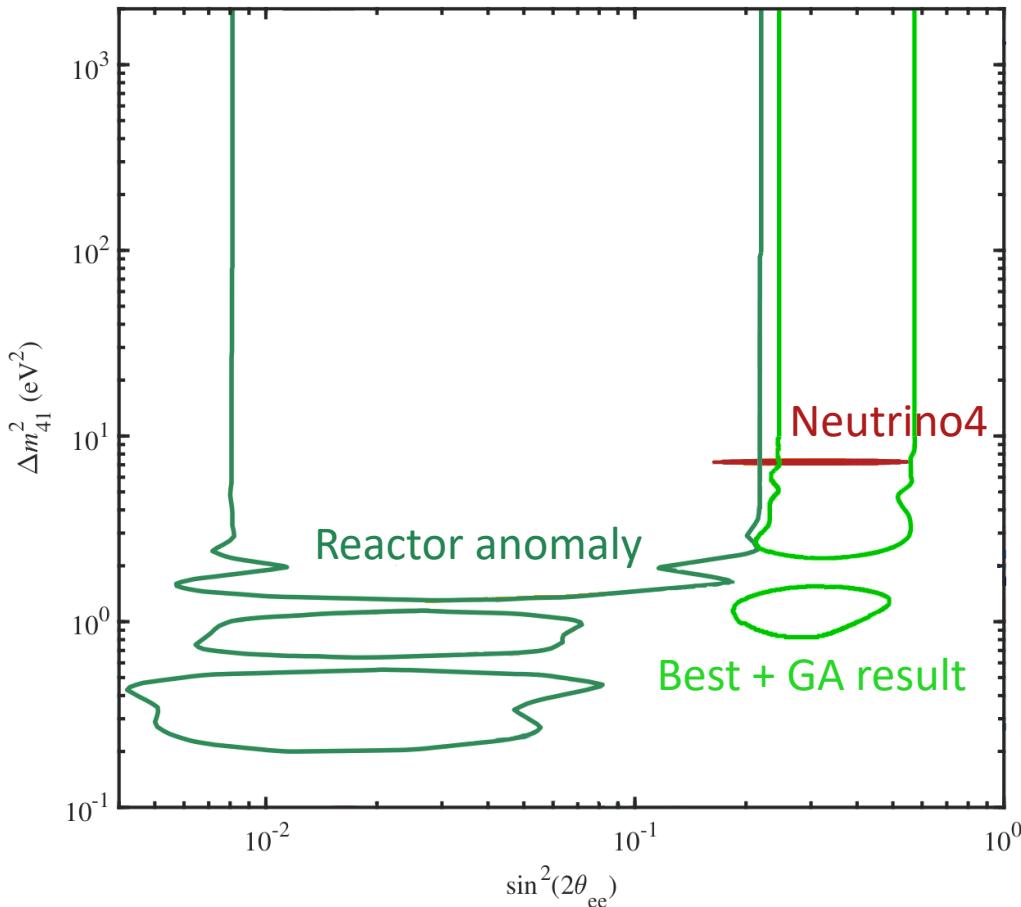
# Light sterile neutrinos

- Motivated by anomalies in  $\nu$ -oscillation experiments → eV-scale

G. Mention et al Phys. Rev. D 83, 073006 (2011)



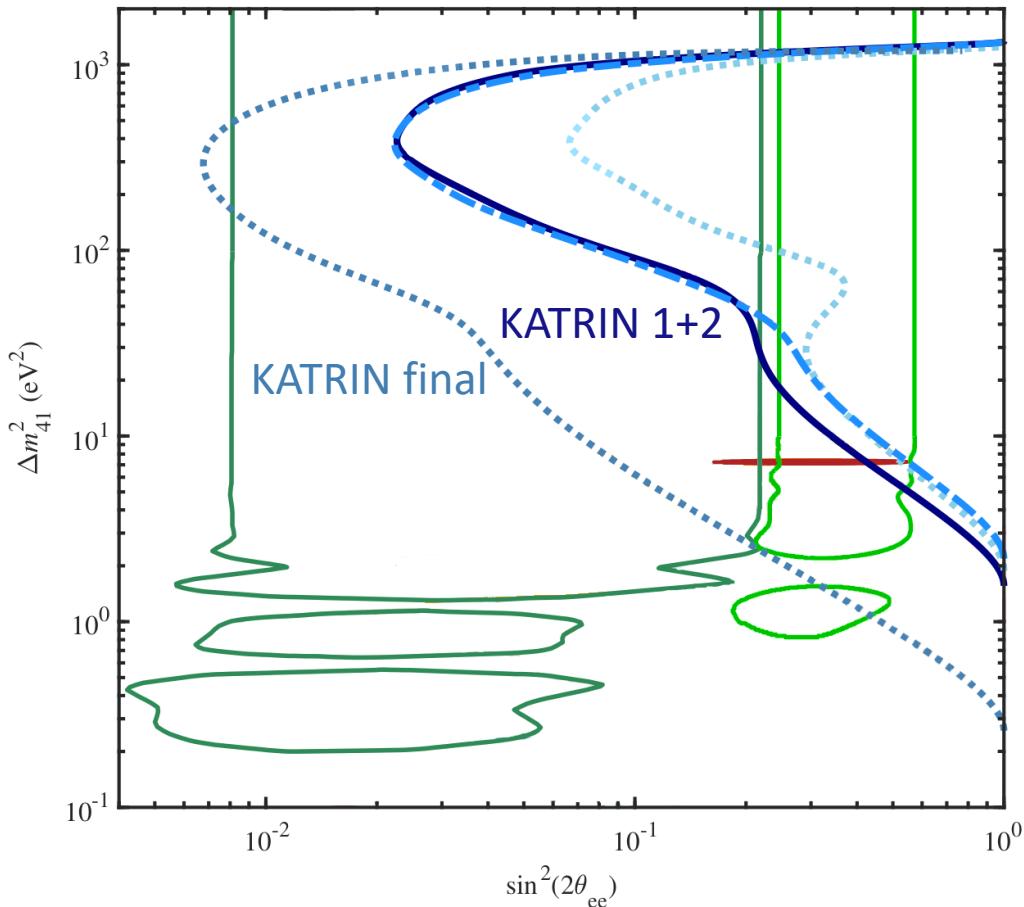
# Sterile neutrino parameter space



- Neutrino4 result → disputed
- Reactor anomaly → resolved?
- Gallium anomaly → strengthened with BEST result

G. Mention et al Phys. Rev. D **83**, 073006 (2011)  
A. P. Serebrov et al., Pisma Zh. Eksp. Teor. Fiz. 109, 209 (2019)  
V. V. Barinov et al. (BEST), arXiv:2109.11482 (2021)

# Sterile neutrino parameter space

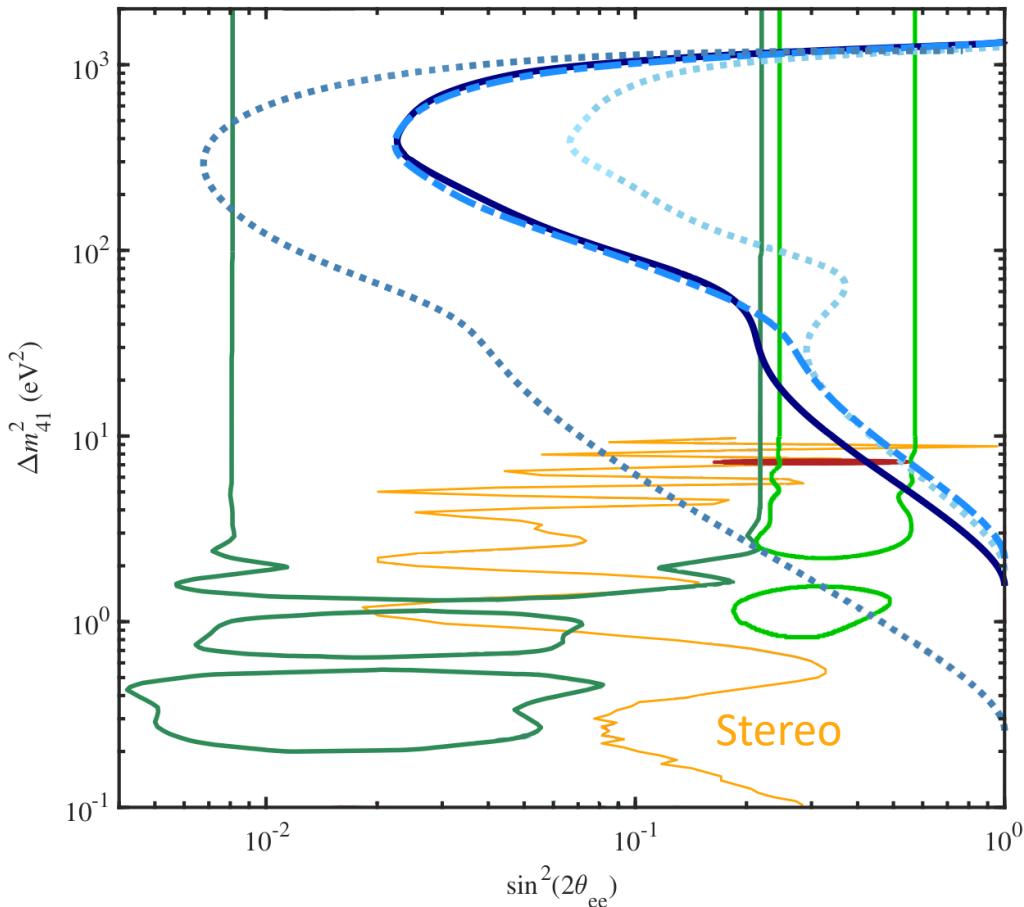


- **Partly exclusion** of Neutrino4 and GA with recent KATRIN result
- Potential to probe Neutrino4 and GA to **great extend**

KATRIN Collab., PRL. 126, 091803 (2021)

KATRIN Collab. arXiv:2201.11593 (2022)

# Sterile neutrino parameter space

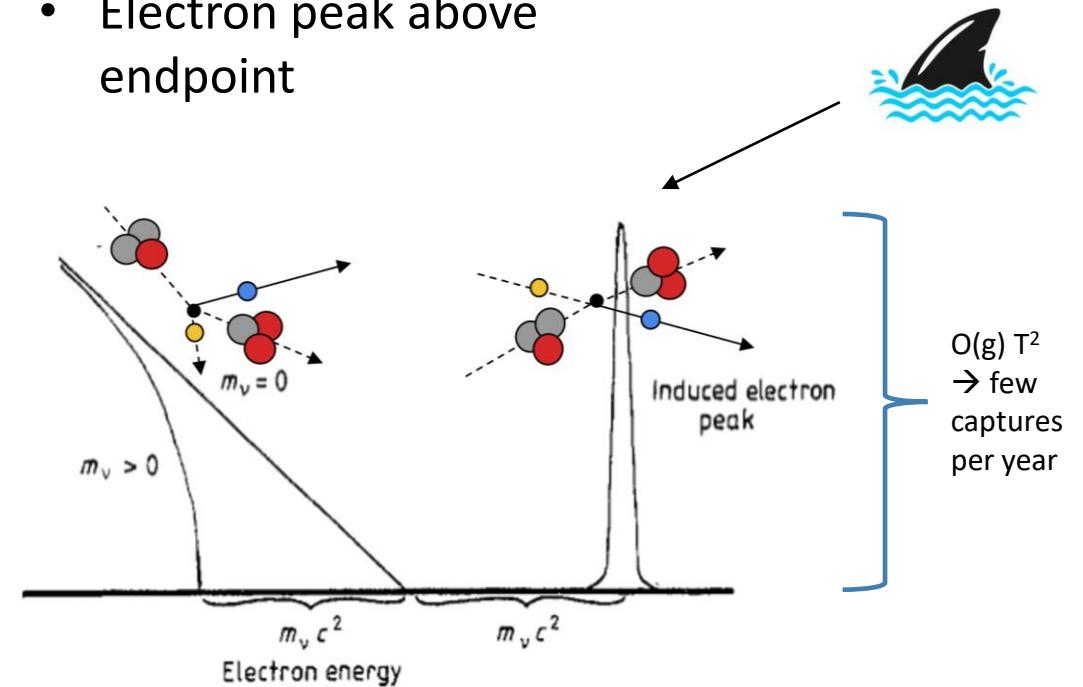
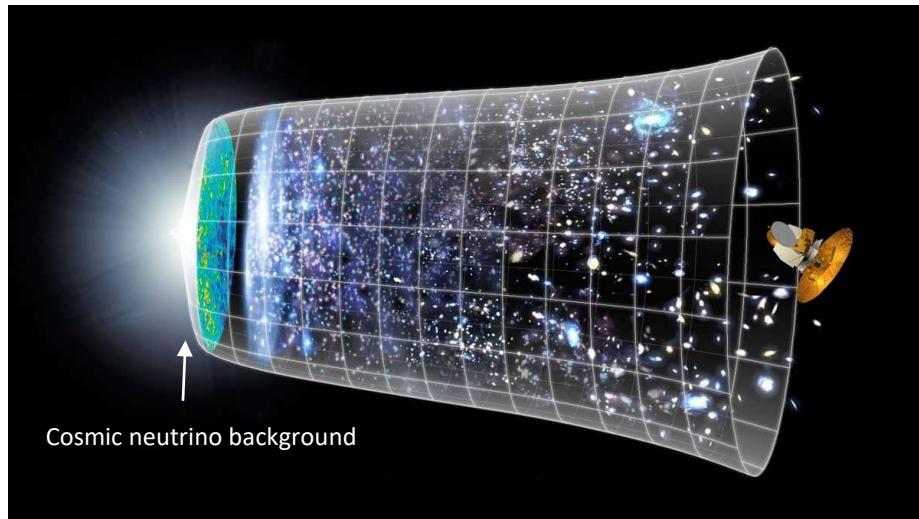


- Complementary probe to oscillation-based experiments

DANSS, arXiv:1911.10140 (2019)  
STEREO, Phys. Rev. D 102, 052002 (2020)  
PROSPECT, Phys. Rev. D 103, 032001 (2021)

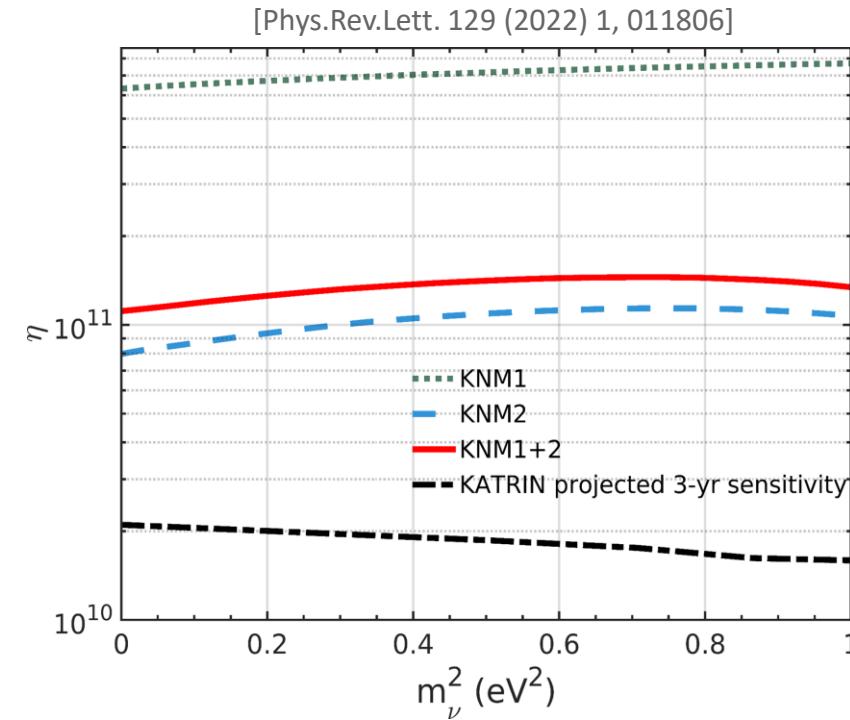
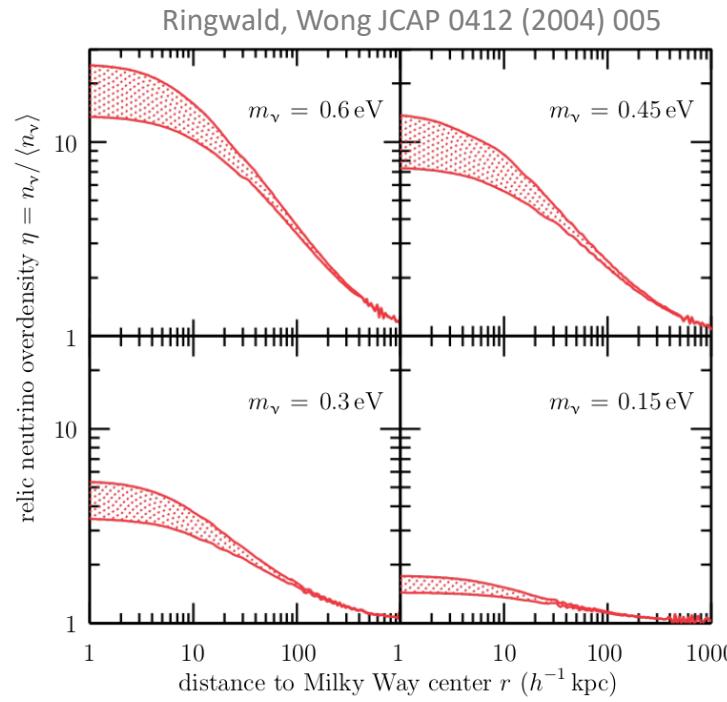
# Relic neutrino signature

- Cosmic neutrino background (**CvB**)
    - Neutrino decoupling 1 s after Big Bang
    - High abundance:  $400 \text{ v/cm}^3$  today
    - Challenge: Tiny cross-section
- 
- **Capture on tritium, no energy threshold**
  - Electron peak above endpoint



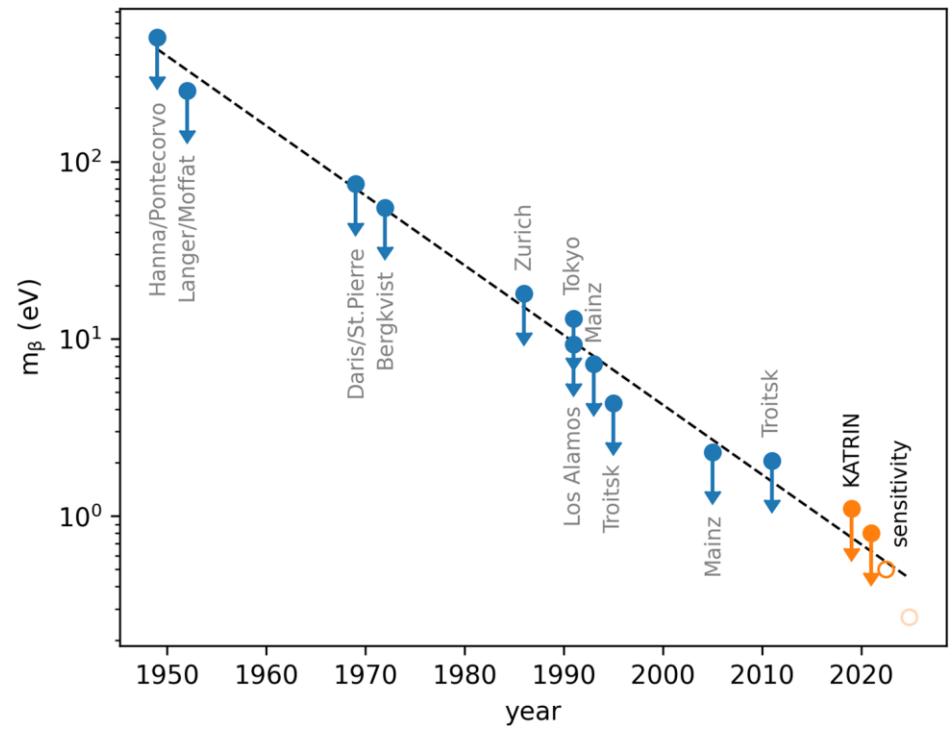
# Relic neutrino overdensity

- KATRIN data:
  - constraint on **local overdensity**  $\eta < 1.1 \cdot 10^{11}$  (95% CL)
  - **x100 improvement** over previous laboratory bounds



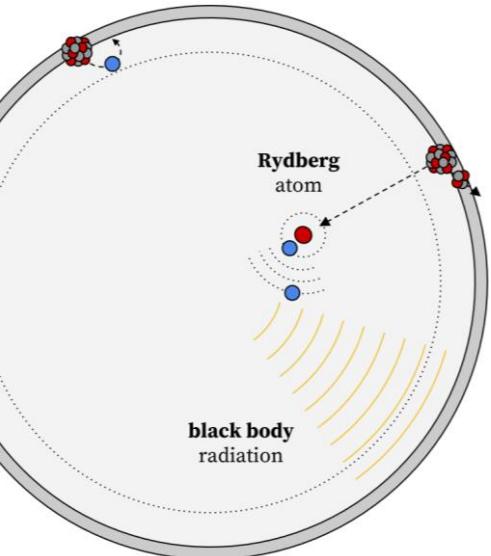
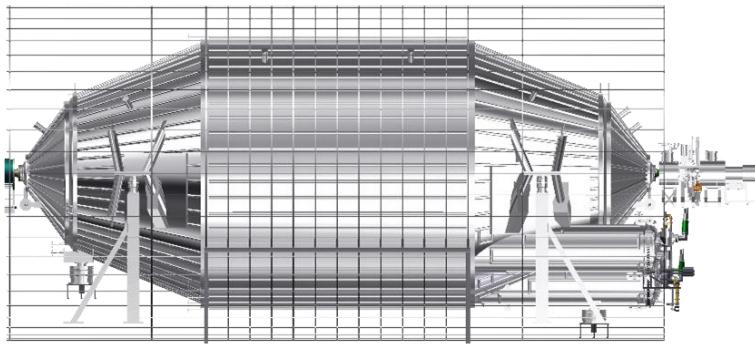
# Summary

- First **sub-eV neutrino mass limit** from a direct experiment:  $m_\beta < 0.8$  eV
- Projected sensitivity for next data release:  $m_\beta < 0.5$  eV
- Beyond the neutrino mass:
  - Complementary search for eV-scale neutrinos
  - New limit on relic neutrino overdensity

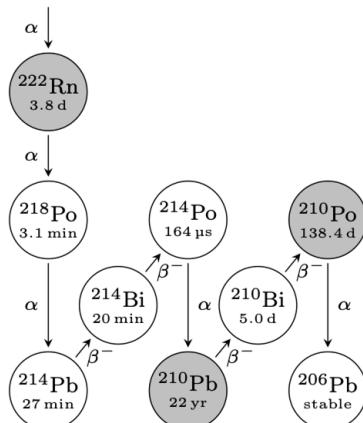


*Backup*

# Spectrometer background



- $^{219}Rn$  decays in spectrometer → trapped electrons, non-Poisson rate



- $^{222}Rn$  on spectrometer walls ( $650 \text{ m}^2$ )
- Sputtering of Rydberg atoms in  $^{210}Po$  decay
- Ionization by black body radiation
- Shifted-analyzing plane (SAP)

