

Status and Prospects of the NO ν A Experiment

Daniel Cronin-Hennessy

August 29, 2014

Heavy Quarks and Leptons 2014

Schloß Waldhausen, Mainz Deutschland

Outline

Overview

NO ν A Physics

NO ν A Technology

Current Status

NO ν A: **NuMI Off-Axis ν_e Appearance Experiment**



Neutrinos at the Main Injector

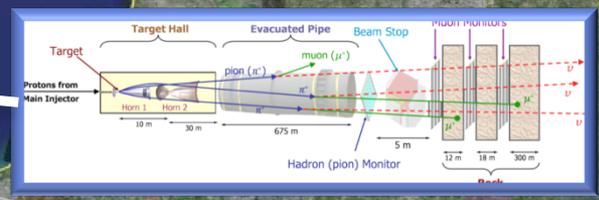
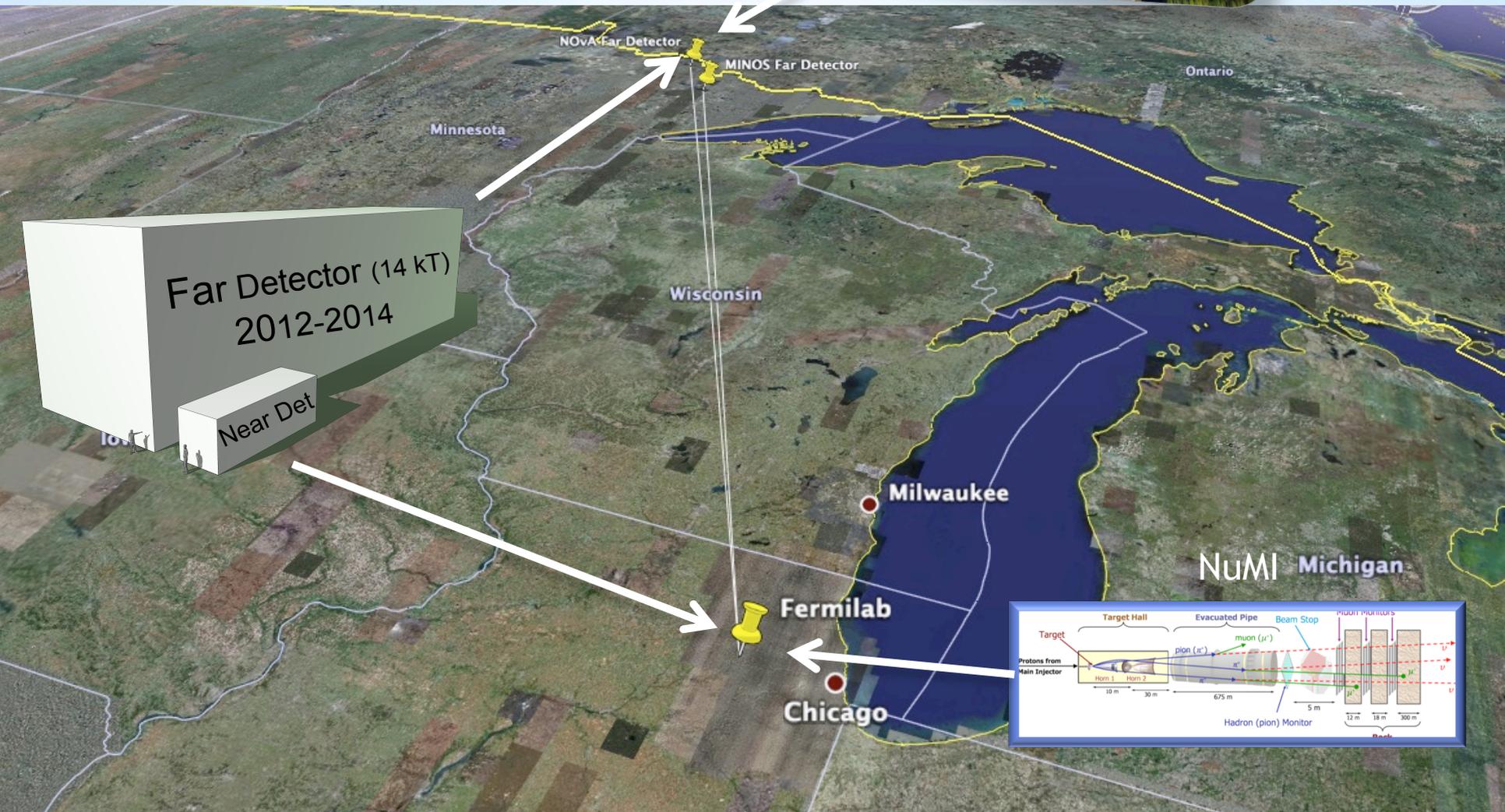
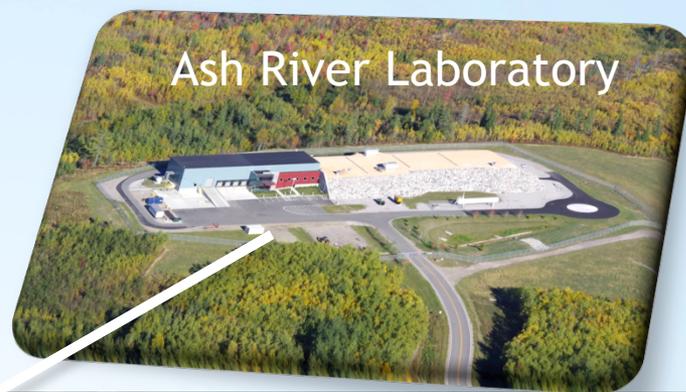
The NOvA Collaboration



38 Institutions from 7 Countries

205 members

Neutrino Beam from Fermilab to Minnesota



NO ν A Physics Impact

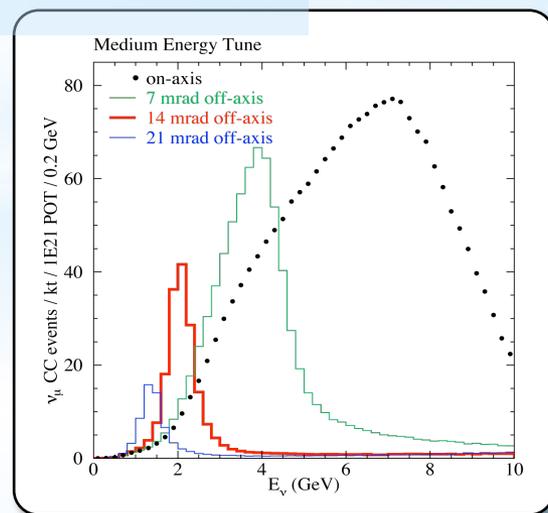
- * Mass Hierarchy from Matter Effect on ν_e Appearance
- * CP Violation from δ in the PNMS matrix
- * Improved Atmospheric Oscillation Parameters from ν_μ Disappearance
- * Tests of 3 Flavor mixing paradigm (ν_s ?)
- * θ_{23} Octant
- * Supernova ν 's

NO ν A Technique

Off-Axis position of detector results in a Narrow band beam peaked at ~ 2 GeV

At 810 km from source the L/E is optimal for electron neutrino appearance.

Beam upgrades that increase power from 320 kW to 700 kW (in progress)



NO ν A Oscillation Measurements

- * Survival Probability $P(\nu_\mu \rightarrow \nu_\mu)$
- * Appearance Probability $P(\nu_e \rightarrow \nu_e)$
- * Neutrinos and anti-neutrinos
- * Sensitive to Hierarchy, $\theta_{13}, \theta_{23}, \delta_{CP}$ and Δm_{31}^2
 - * Reactor experiments do not have all of these dependencies.

$$P(\bar{\nu}_\mu \rightarrow \bar{\nu}_e) \approx \sin^2 2\theta_{13} \sin^2 \theta_{23} \frac{\sin^2(A-1)\Delta}{(A-1)^2}$$

$$- 2\alpha \sin \theta_{13} \sin \delta_{CP} \sin 2\theta_{12} \sin 2\theta_{23} \frac{\sin A\Delta \sin(A-1)\Delta}{A(A-1)} \sin \Delta$$

$$+ 2\alpha \sin \theta_{13} \cos \delta_{CP} \sin 2\theta_{12} \sin 2\theta_{23} \frac{\sin A\Delta \sin(A-1)\Delta}{A(A-1)} \cos \Delta$$

Where: $\alpha = \frac{\Delta m_{21}^2}{\Delta m_{31}^2}$ $\Delta = \Delta m_{31}^2 \frac{L}{4E}$ $A = \begin{matrix} (-) \\ + \end{matrix} G_f N_e \frac{L}{\sqrt{2}\Delta}$

ν_e Appearance in Matter

$$P(\nu_\mu \rightarrow \nu_e) \simeq \left| \sqrt{P_{Atm}} e^{-i(\Delta_{32} + \delta)} + \sqrt{P_{Sol}} \right|^2$$

$$= P_{atm} + P_{sol} + 2\sqrt{P_{atm}P_{sol}}(\cos \Delta_{32} \cos \delta \mp \sin \Delta_{32} \sin \delta)$$

$$\sqrt{P_{atm}} = \sin \theta_{23} \sin 2\theta_{13} \frac{\sin(\Delta_{31} - aL)}{\Delta_{31} - aL} \Delta_{31}$$

$$\sqrt{P_{sol}} = \cos \theta_{23} \sin 2\theta_{12} \frac{\sin(aL)}{aL} \delta_{21}$$

$$a = G_F N_e / \sqrt{2} \simeq 1/3500 \text{ km}$$

This term has a dependence on the relative sign of Δ_{31}

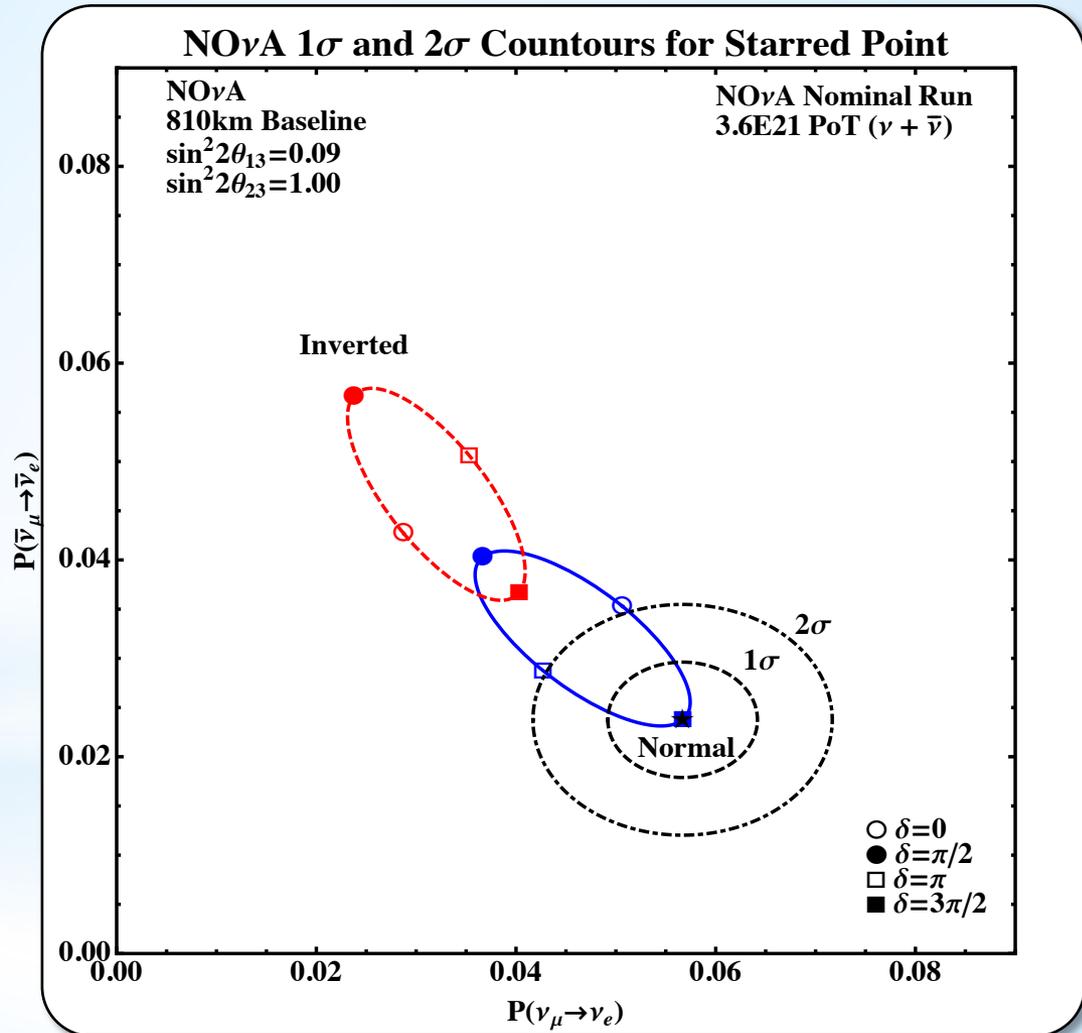
Effect increases with the baseline:

$$aL = 0.08 \text{ for } L = 295 \text{ km}$$

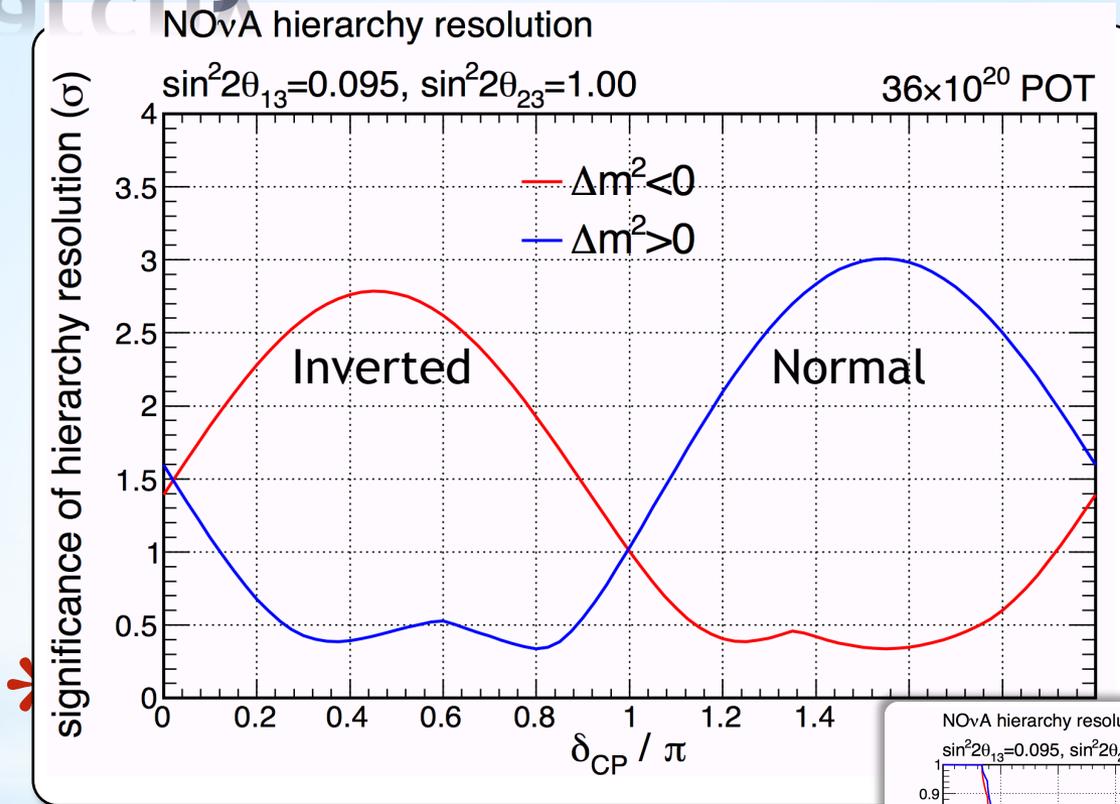
$$aL = 0.23 \text{ for } L = 810 \text{ km}$$

This gives NO ν A a method of determining the mass order for ν_1 / ν_3

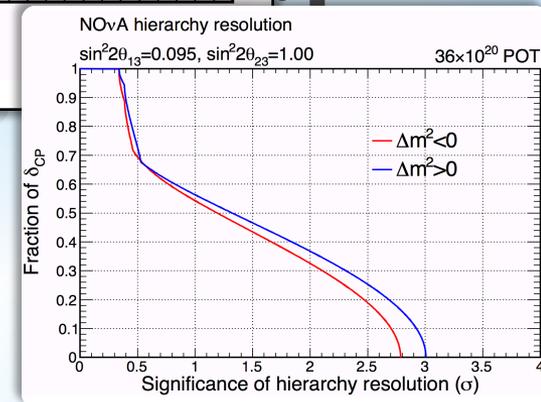
Measuring the Hierarchy



Measuring the Hierarchy



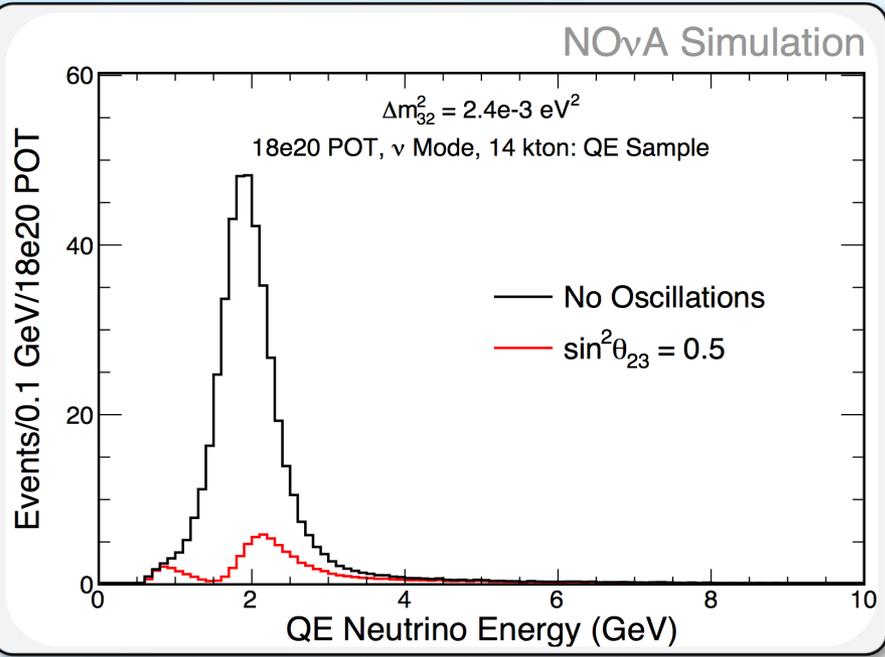
For a maximal θ_{23} , NOvA's sensitivity to the resolution of the hierarchy reaches 95% CL over a third of δ_{CP}



$$P(\nu_{\mu} \rightarrow \nu_{\mu})$$

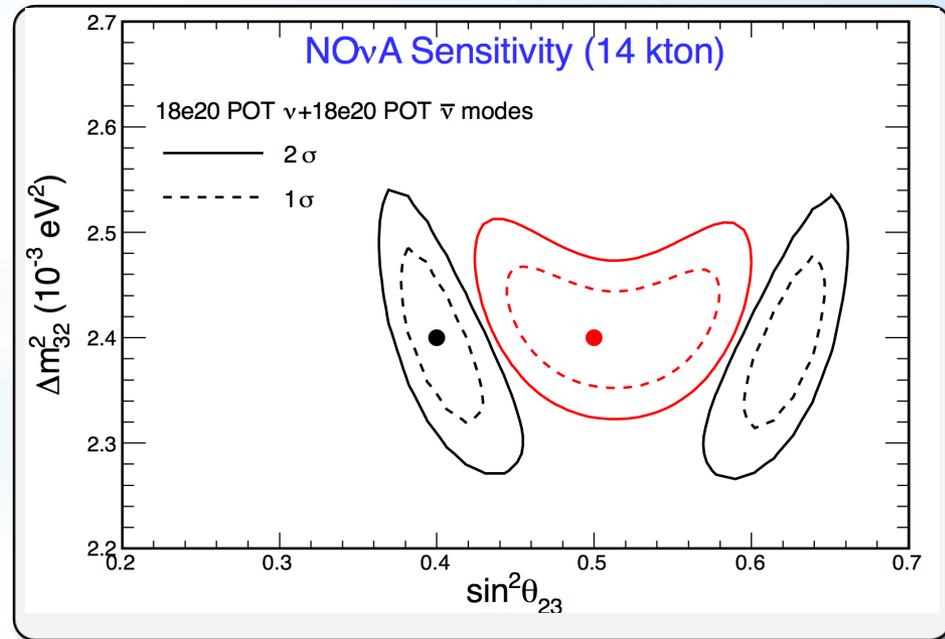
Sensitivities to non-maximal θ_{23} now include cosmic rejection efficiencies

For baseline exposure, any value θ_{23} the NOvA measurements will improve on current best fit values



Energy resolution on ν_{μ} CC Events from current detector performance & reconstruction techniques:

- 4.5% energy resolution for QE events
- 6% resolution for non-QE events

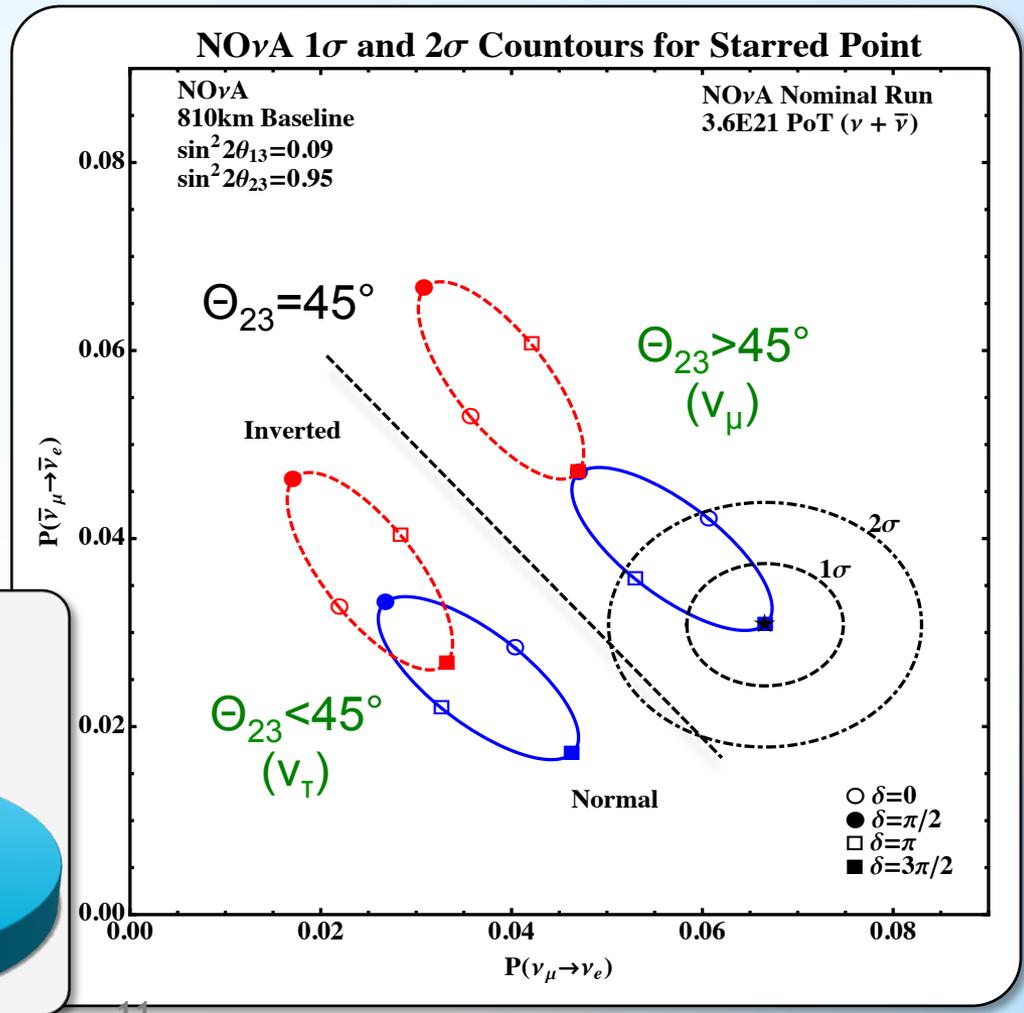
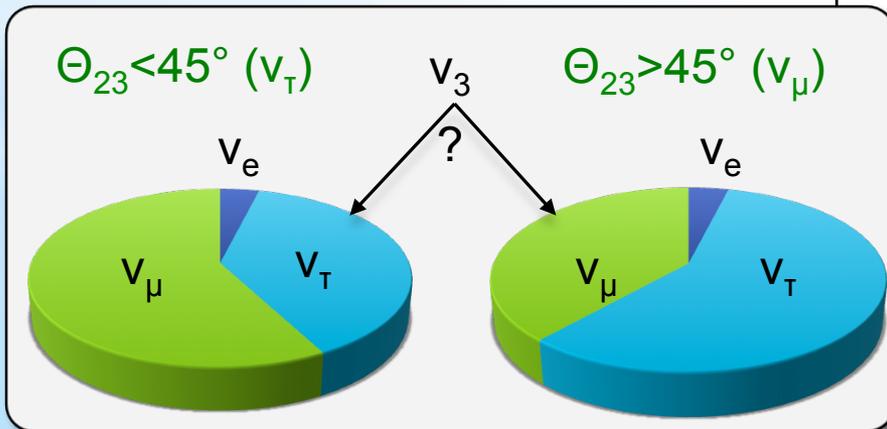


Non-Maximal θ_{23}

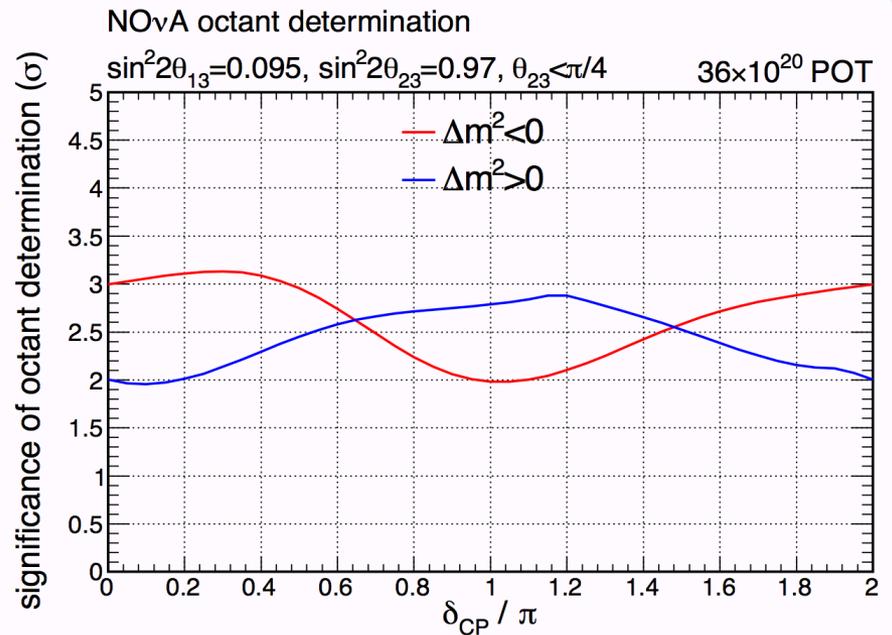
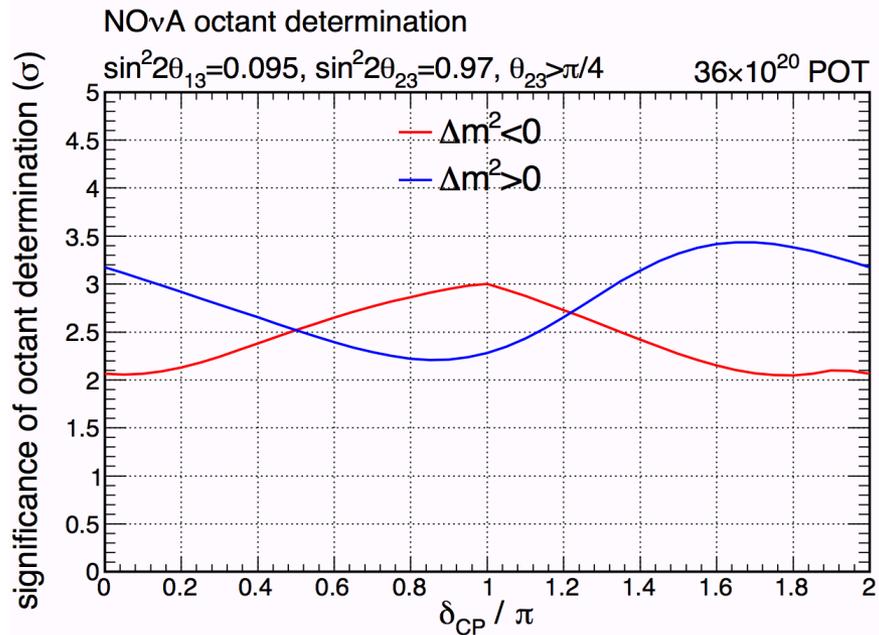
NOvA will make a precision measurement of $P(\nu_\mu \rightarrow \nu_\mu)$ which has the potential to establish $\theta_{23} \neq 45^\circ$ based on the $\sin^2 \theta_{23}$ dependence of the oscillation probability.

If this is the case then the $P(\nu_\mu \rightarrow \nu_e)$ ellipses shift based on how far θ_{23} differs from 45° .

This corresponds to the mass state ν_3 coupling more to ν_τ or ν_μ



Octant Determination



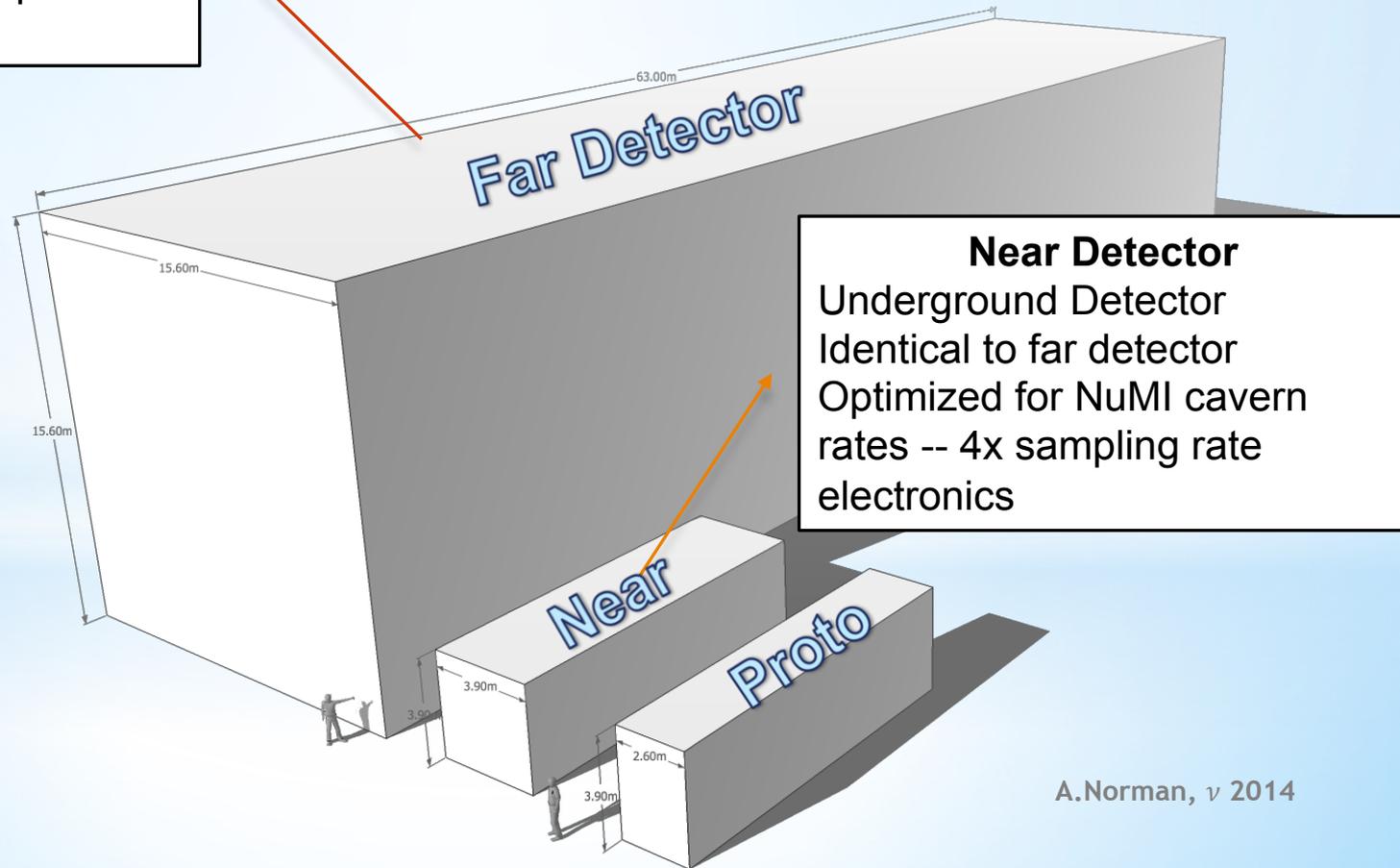
For Non-maximal θ_{23} octant determination by NO ν A

- For $\sin^2 2\theta_{23}=0.97$
- > 95% CL for all values of δ_{CP}

NOvA Detectors

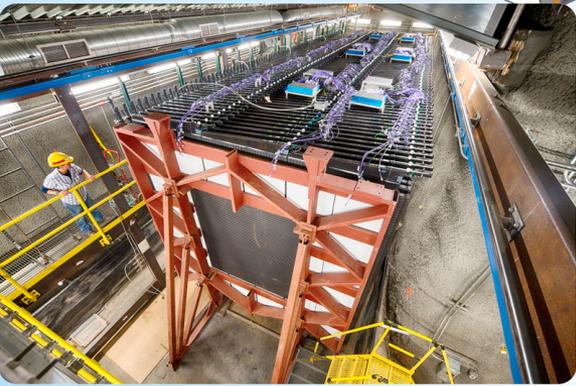
Far Detector

- *Surface Detector*
- 14 kt “Totally Active”, Low Z, Range Stack/Calorimeter
- Liquid Scintillator filled PVC
- 896 alternating X-Y planes

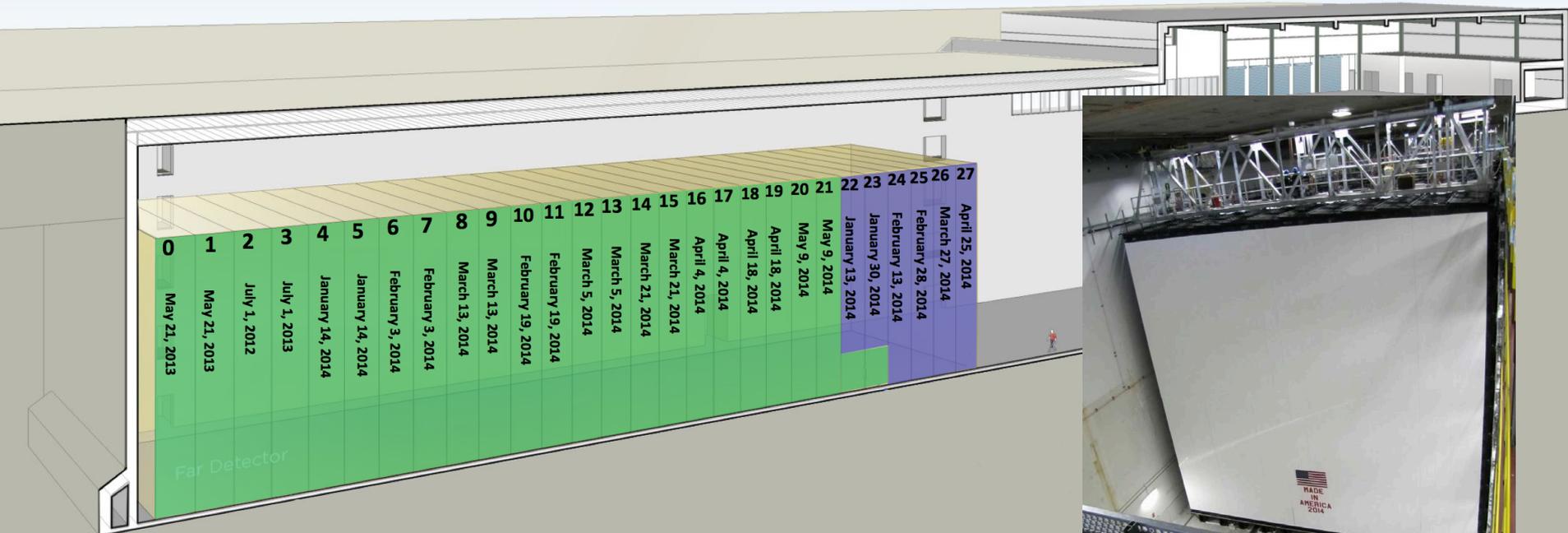


Near Detector

Underground Detector
Identical to far detector
Optimized for NuMI cavern
rates -- 4x sampling rate
electronics



* NOvA Detectors
are BIG



NOvA By the Numbers:

Over:

- ✓ 11 million liters of scintillator
- ✓ 10.4 million meters of Wavelength Shifting Fiber
- ✓ 305,000 meters of PVC Extrusion
- ✓ 11,000 Detector Modules, Front End Boards & APDs

Construction Completed:

Apr 25, 2014 (14 kt)

Electronics:

80% complete (11.25 kt)

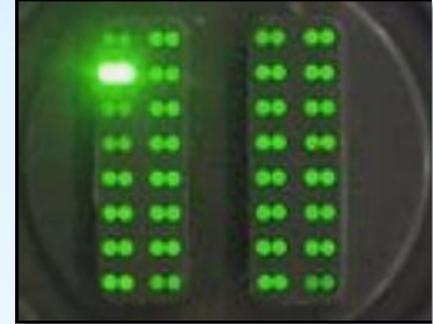
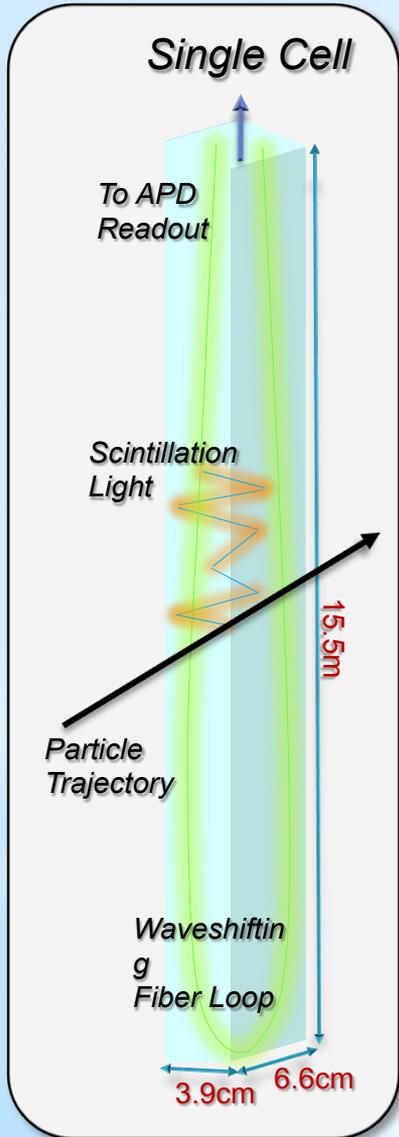
Estimated Completion:

July 2014

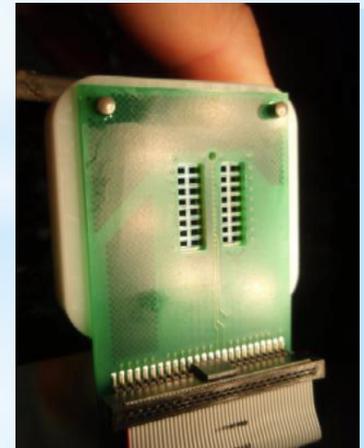
The Readout Cell

NOvA Detection Cell

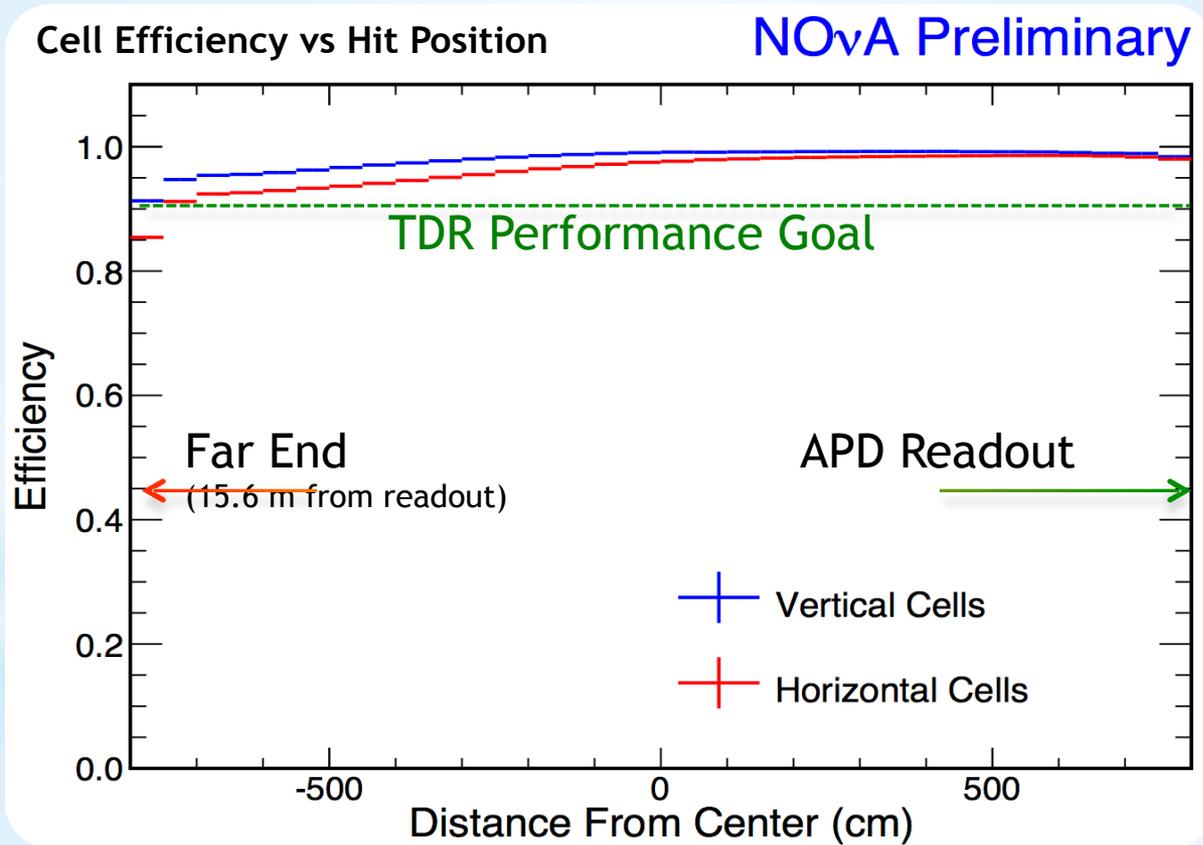
- * 15.5 m long, 3.9x6.6cm tube,
- * Made of reflective PVC structure,
- * Filled with liquid scintillator,
- * Wavelength shifting fiber for light transport,
- * Read out by an avalanche photodiode
 - * Passage of MIP through the cell results in $dE/dx \approx 12.9$ MeV across the cell.
- * Need the light output to be 30-38 p.e. from the FAR END end to give a 10:1 signal/noise



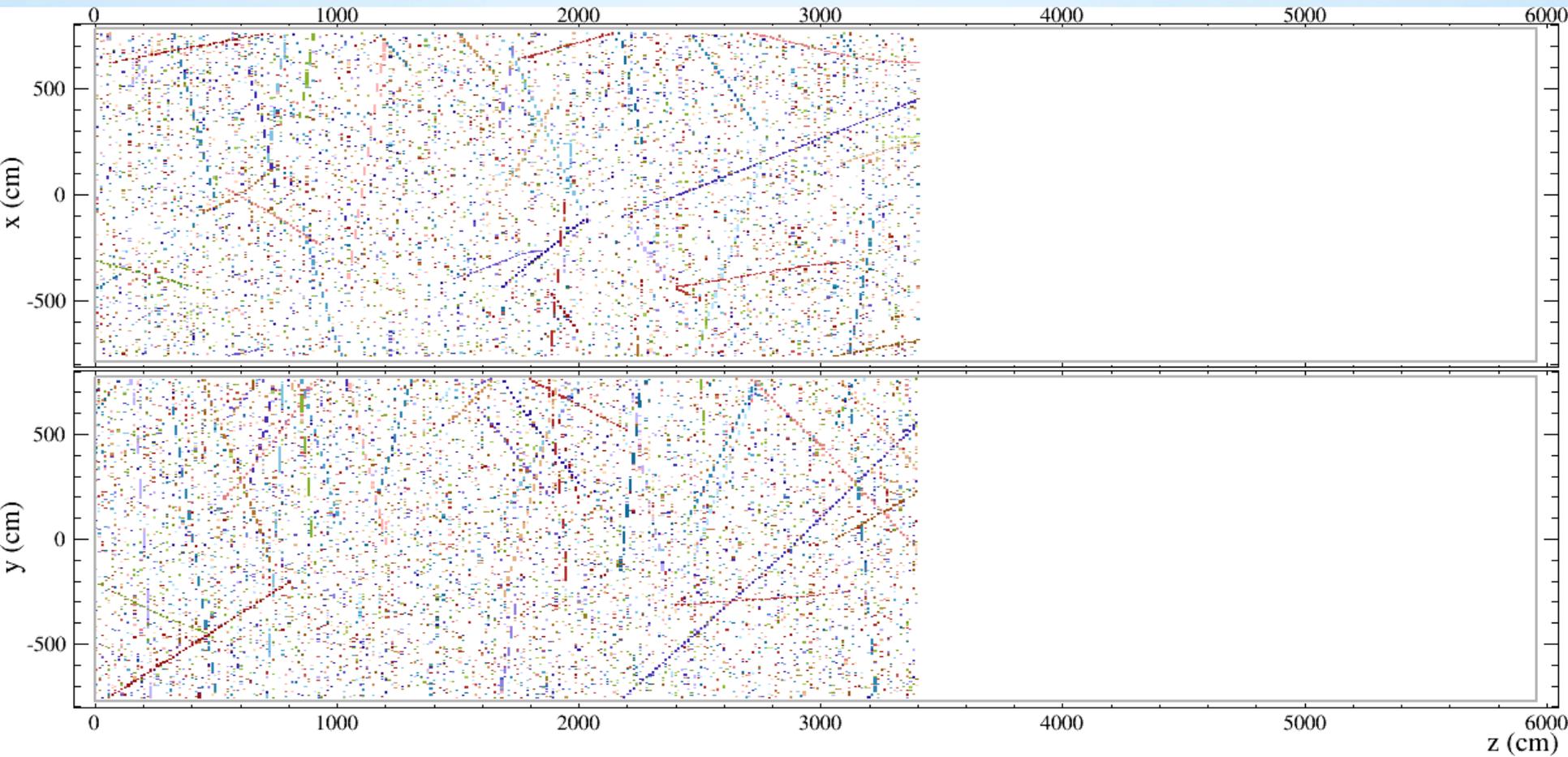
Waveshifting fiber readout and 32pixel APD used for cell digitization



NO_vA Cell Efficiency



Life on the surface: Cosmic ray rate $\sim 120\text{k Hz}$



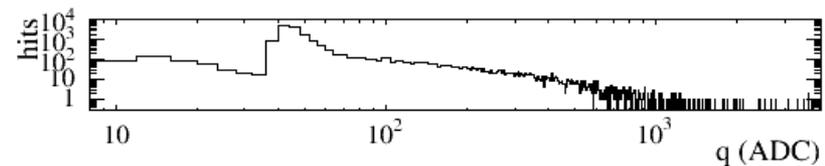
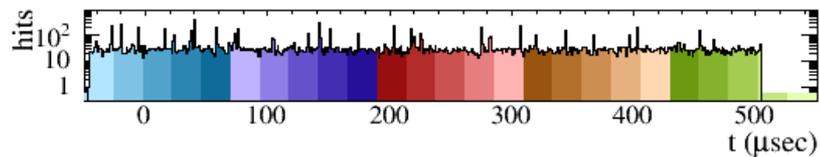
NOvA - FNAL E929

Run: 14828 / 38

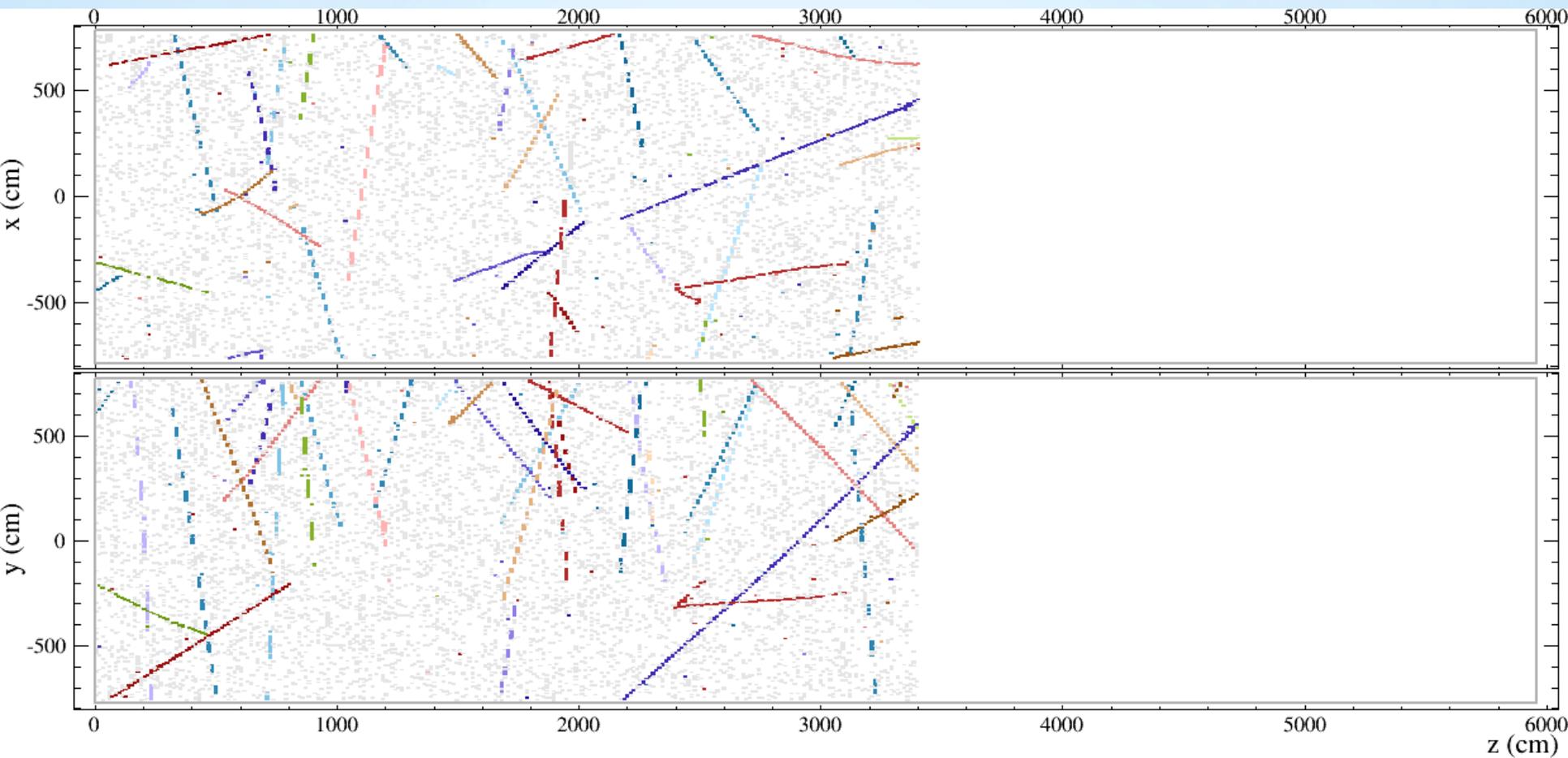
Event: 192569 / NuMI

UTC Tue Apr 22, 2014

21:41:51.422846016



Life on the surface: Reconstructed tracks



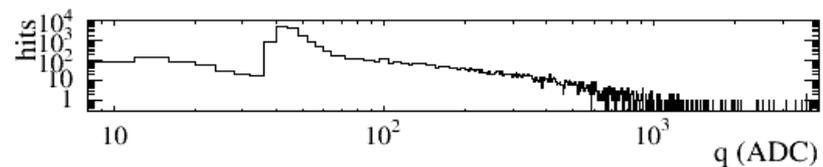
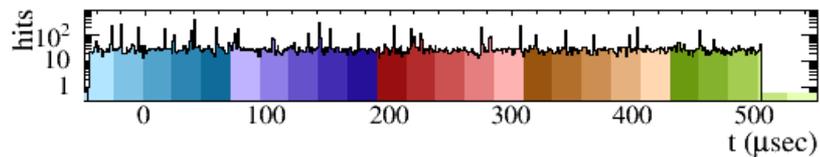
NOvA - FNAL E929

Run: 14828 / 38

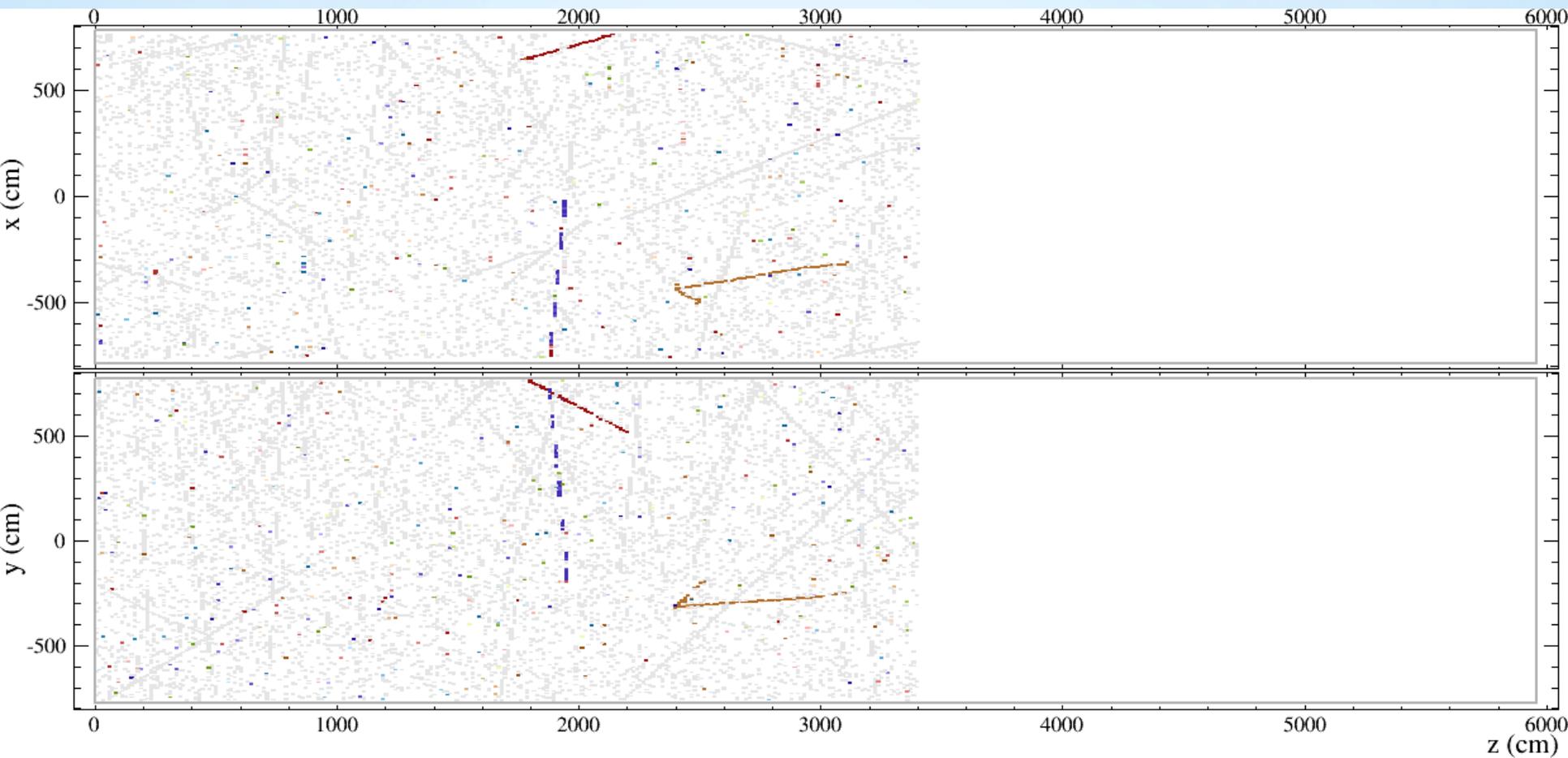
Event: 192569 / NuMI

UTC Tue Apr 22, 2014

21:41:51.422846016



Life on the surface: In time with accelerator signal



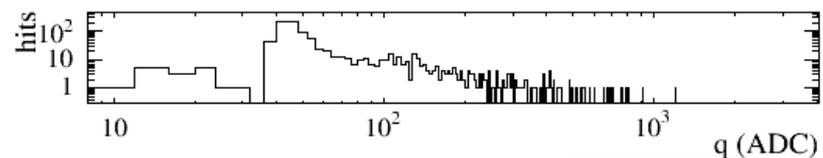
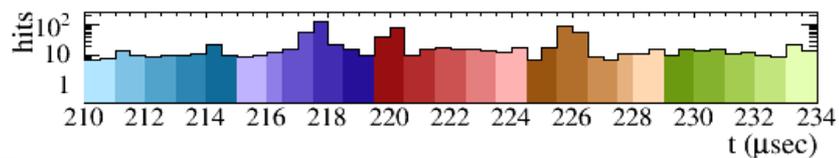
NOvA - FNAL E929

Run: 14828 / 38

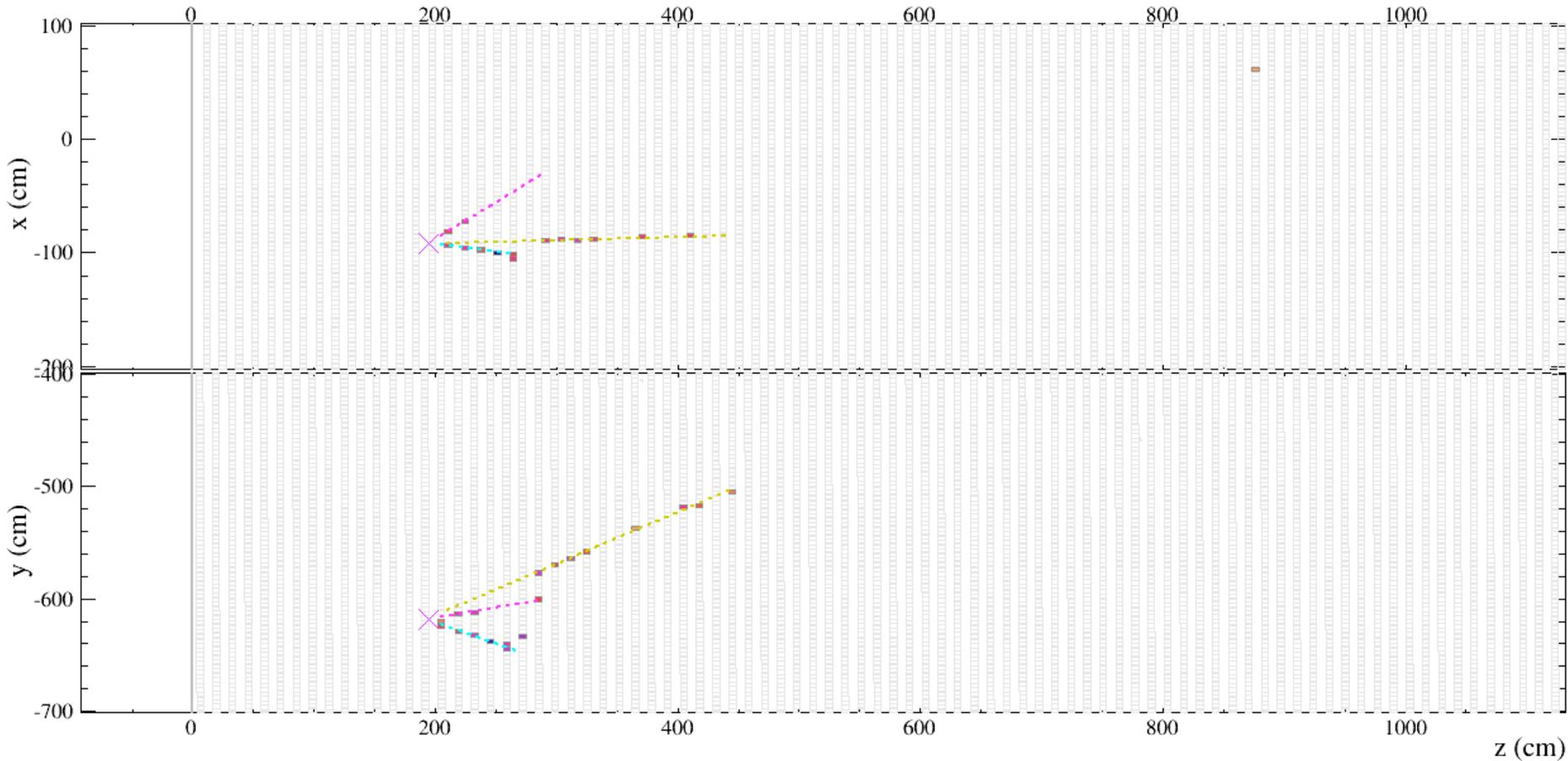
Event: 192569 / NuMI

UTC Tue Apr 22, 2014

21:41:51.422846016



First γ November 12, 2013



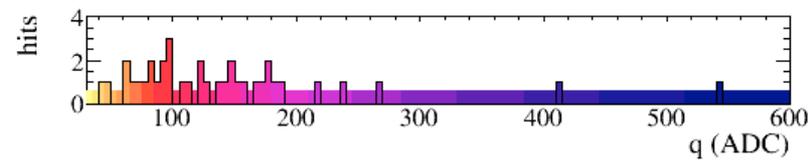
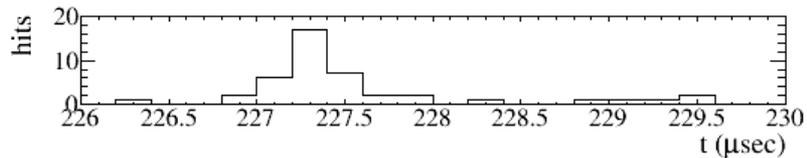
NOvA - FNAL E929

Run: 11654 / 9

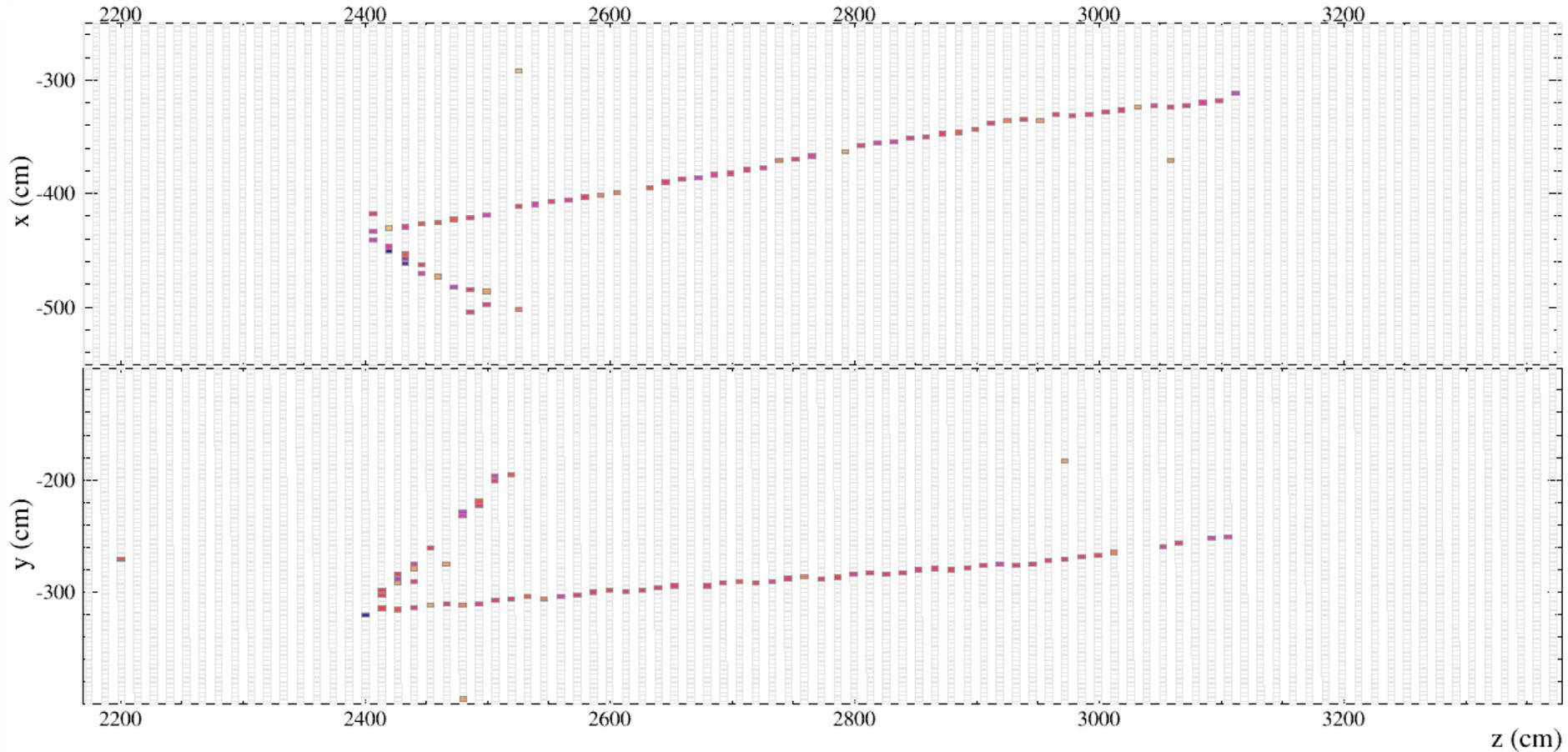
Event: 77385 / NuMI

UTC Tue Nov 12, 2013

13:25:44.976546176



First ν_{μ} -CC Candidate



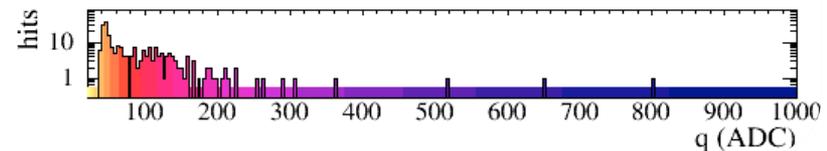
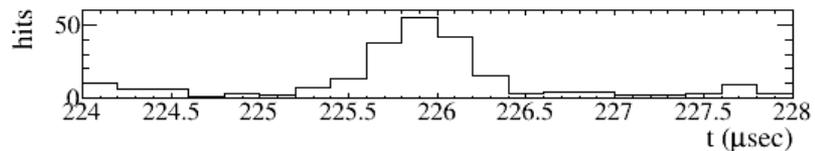
NOvA - FNAL E929

Run: 14828 / 38

Event: 192569 / NuMI

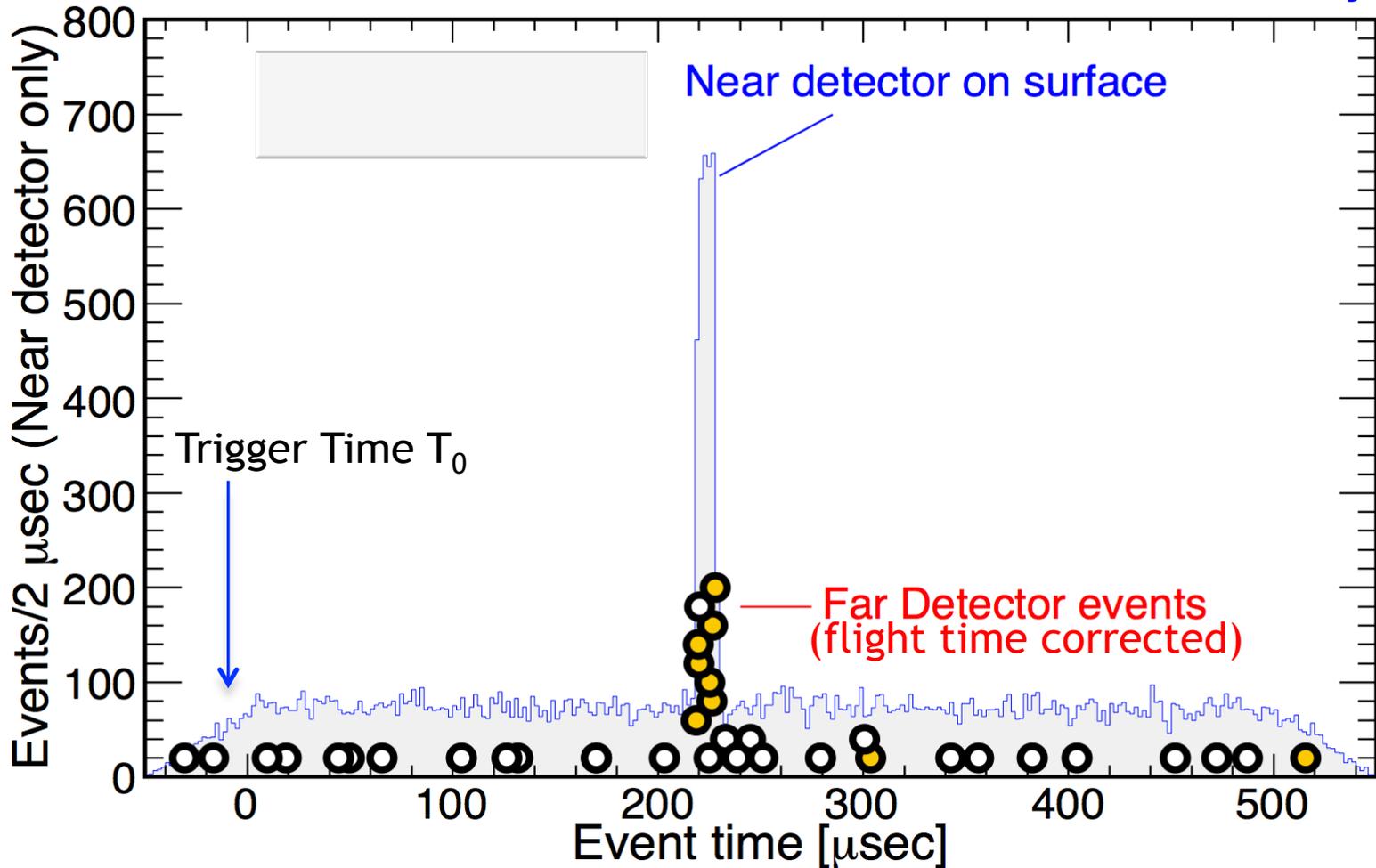
UTC Tue Apr 22, 2014

21:41:51.422846016

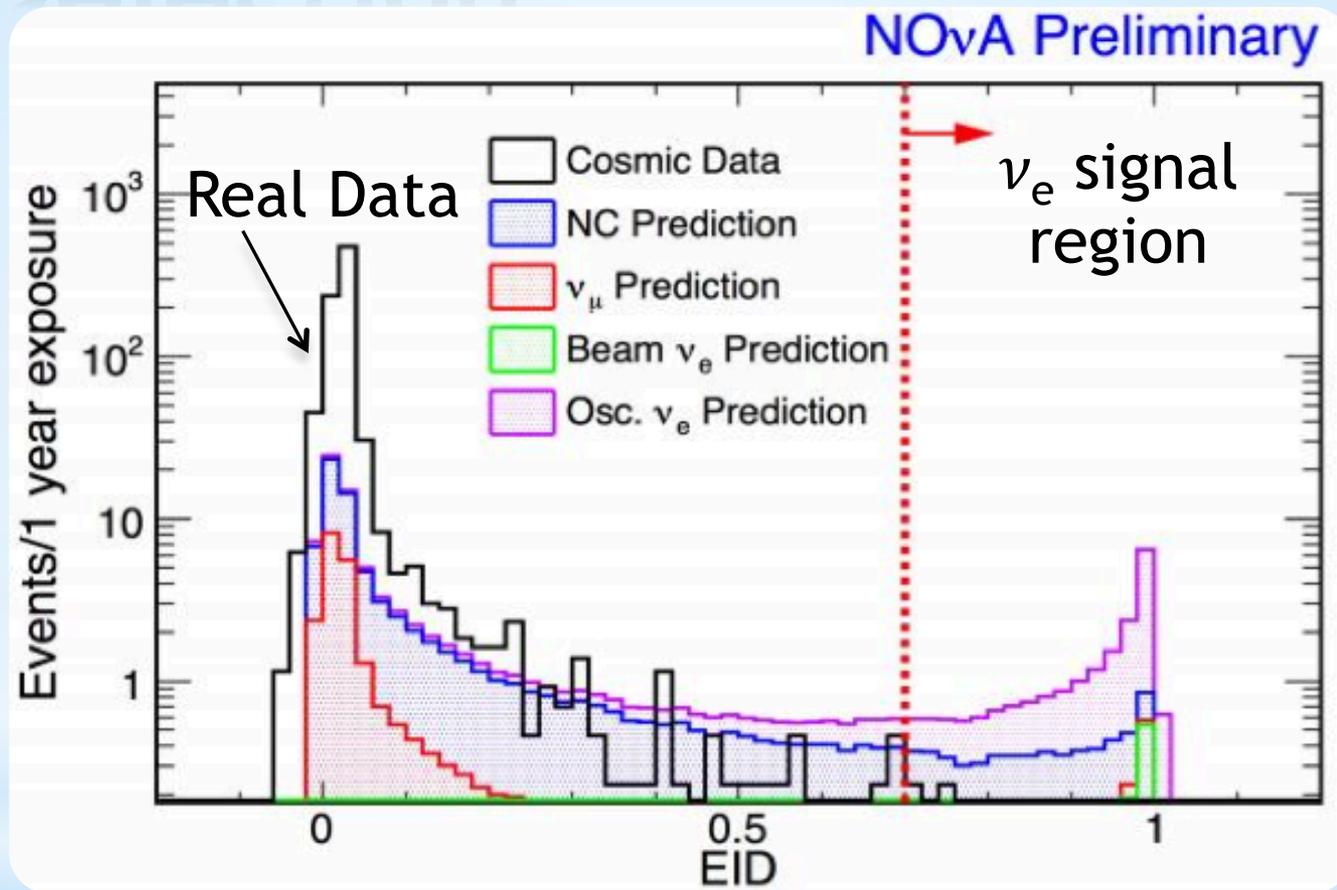


Far Detector γ Candidates

Neutrino Candidate Events NO ν A Preliminary



ν_e Selection



Complimentary Methods of PID

- EID - Neural net based on EM shower profile
- Library Event Matching (LEM) - Template method based on event topology

ν_e Selection

ν_e Signal and Background Estimates

Cut	Simulation				Data	
	ν_e Signal	NC	Beam ν_e	ν_μ CC	Cosmic Ray	All Background
All Events	36.7	380	28.1	557	19M	19M
Pre-selection	24.7	83.5	2.9	30.0	56k	56k
Vertex Gap	24.6	81.8	2.9	29.6	55k	55k
P_T/P	22.0	59.6	2.6	24.3	1248	1334
Maximum Y	21.2	57.4	2.5	23.0	834	917
Neutral Net	13.9	3.9	1.5	0.7	0.5	6.5
Library Template	14.0	3.5	1.5	1.1	0.9	7.0

- Exposure 6×10^{20} POT
- 1 yr at design mass/beam power
- 14 kt total detector mass
- Signal estimates are leading order
- Simple oscillation w/o matter effect
- Averaged over hierarchy and δ_{CP}

Summary

- * Near and Far Detectors are complete.
- * Detectors functioning as expected.
- * We are currently acquiring beam data.
- * Accelerator is ramping toward 700 kW.
- * By end of year we will have collected the data for our first oscillation results.