

Bayesian UQ in transfer reactions for astrophysics

Filomena Nunes
Michigan State University

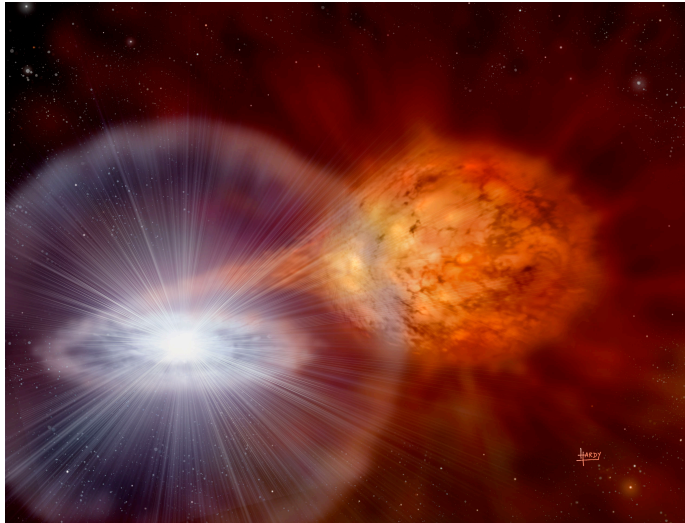
In collaboration with Manuel Catacora-Rios,
Garrett King and Amy Lovell



Outline

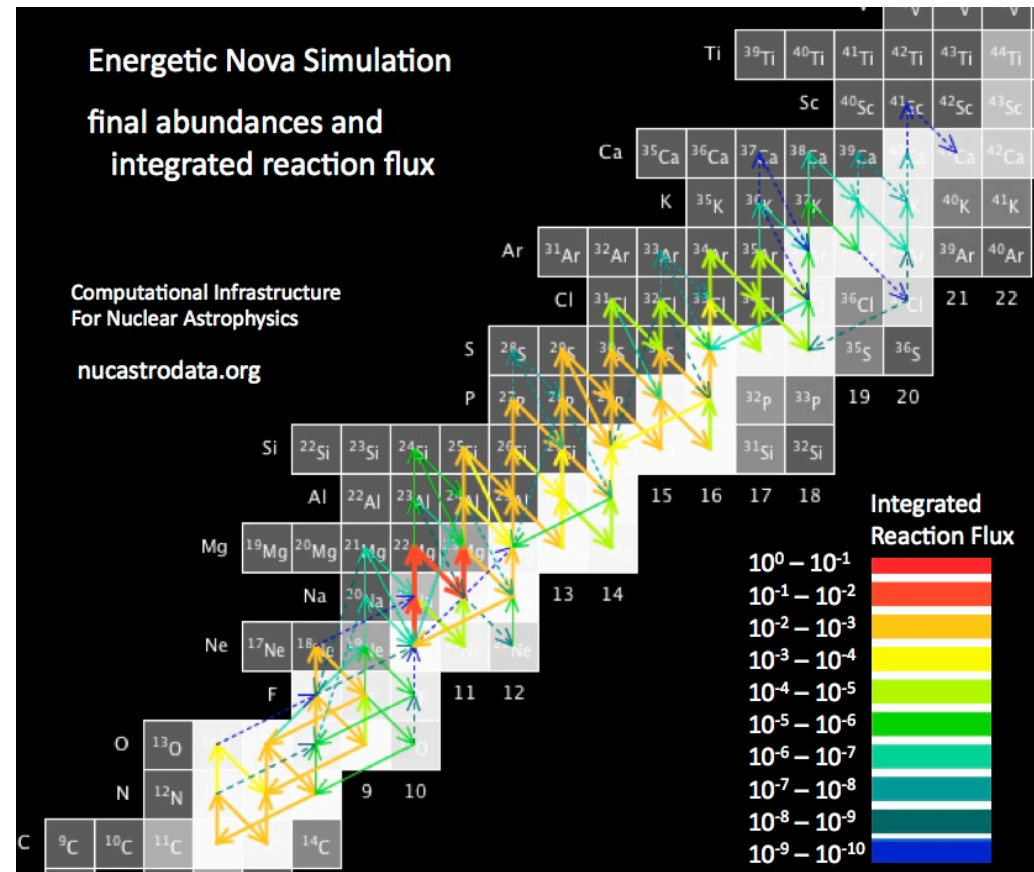
- ✧ Transfer reactions in the astrophysical context
- ✧ Source of uncertainties in reaction models
- ✧ Uncertainty quantification approaches
- ✧ Bayesian analysis to reduce uncertainties
- ✧ Challenges and Opportunities

From novae to nuclear reactions



apod.nasa.gov/

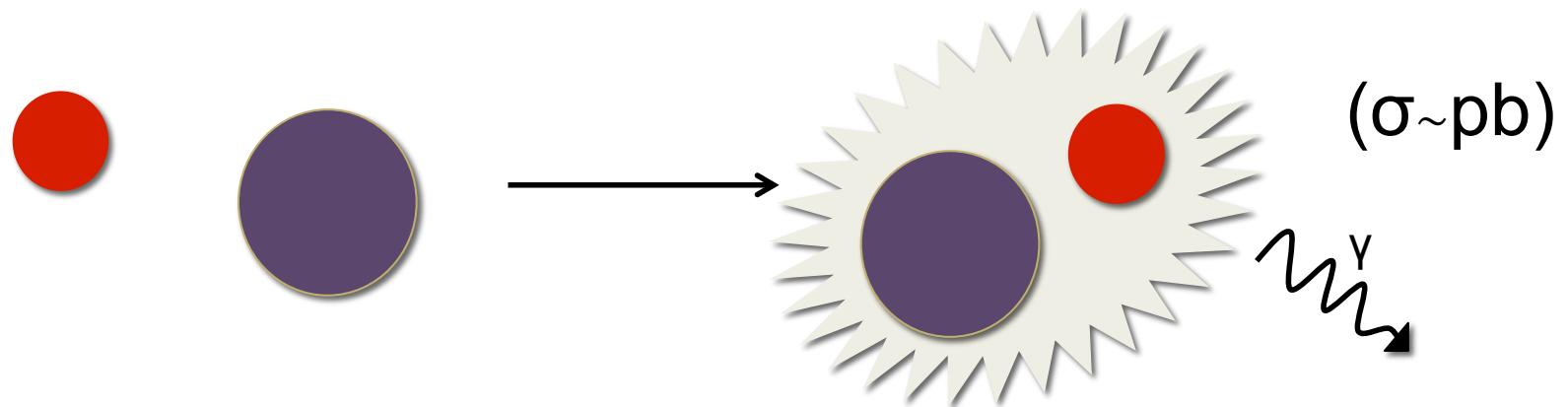
need:
proton capture rates on
proton-rich nuclei



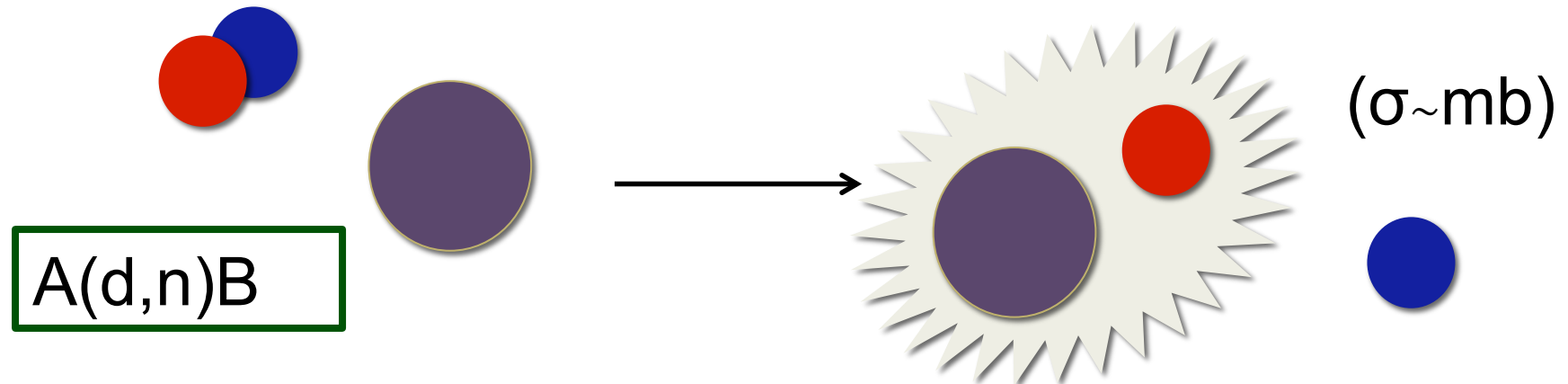
phy.ornl.gov/

how do we measure charged- particle capture in the Gamow window?

✧ (p,γ) cross sections are very low



✧ (d,n) cross sections are many orders of magnitude higher!



Applications to recent (p,g) measurements

$^{56}\text{Ni}(d, n\gamma)^{57}\text{Cu}$ to extract $^{56}\text{Ni}(p, \gamma)^{57}\text{Cu}$

NSCL: S800 for PID (^{57}Cu) + GRETINA for γ

$^{56}\text{Ni}(d, n)^{57}\text{Cu}$						
E_{ex}	J^π	ℓ	σ_{exp} (mb)	σ_{th} (mb)	$C^2S_{(d,n)}$	C^2S_{SM}
1.028	$5/2^-$	3	2.00(40)	2.62	0.76(28)	0.75
1.109	$1/2^-$	1	0.28(6)	0.45	0.62(22)	0.71
2.398	$5/2^-$	3	<0.2	2.61	$< 8 \times 10^{-2}$	1.8×10^{-3}
2.525	$7/2^-$	3	<0.2	14.5	–	3.9×10^{-2}
$^{56}\text{Ni}(d, p)^{57}\text{Ni}$						
E_{ex}	J^π	ℓ	σ_{exp} (mb)	σ_{th} (mb)	$C^2S_{(d,p)}$	C^2S_{SM}
0.768	$5/2^-$	3	2.10(60)	2.77	0.77(31)	0.74
1.122	$1/2^-$	1	0.50(15)	0.68	0.73(31)	0.69
2.443	$5/2^-$	3	<0.4	2.61	< 0.1	3×10^{-4}
2.579	$7/2^-$	3	1.24(36)	14.9	$8(3) \times 10^{-2}$	4.1×10^{-2}

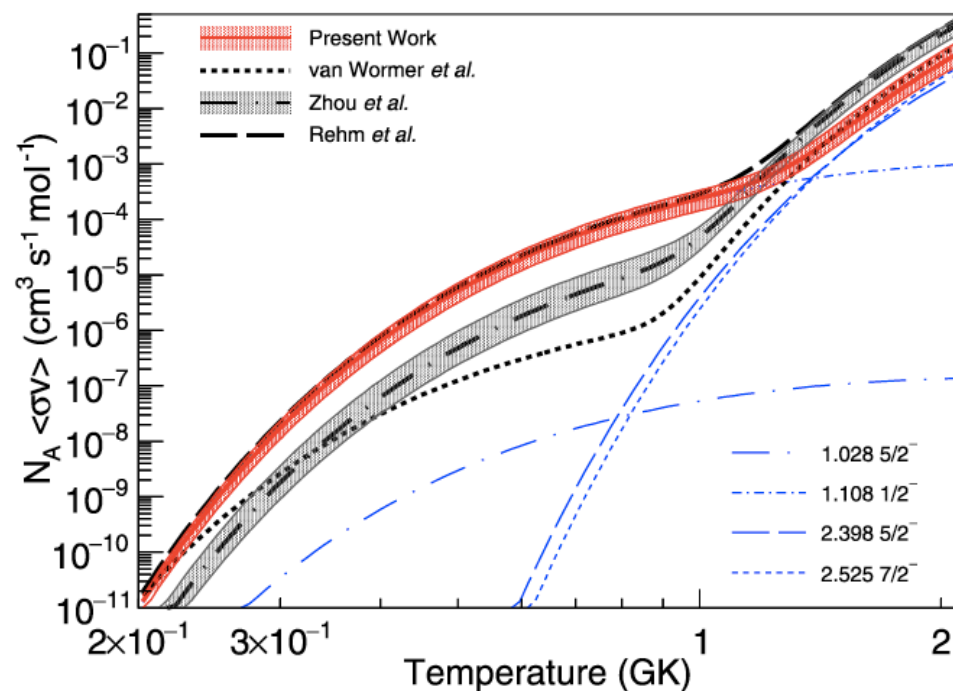
Angle
integrated
cross
section
only



Applications to recent (d,n) measurements

E_{ex} (keV)	E_{r} (keV)	J^{π}	Γ_p (eV)
1028(1)	338	$5/2^-$	5.7×10^{-12}
1108(2)	418	$1/2^-$	1.9×10^{-7}
2398(10)	1708	$5/2^-$	5.5×10^{-3}
2525(17)	1835	$7/2^-$	5.3×10^{-1}

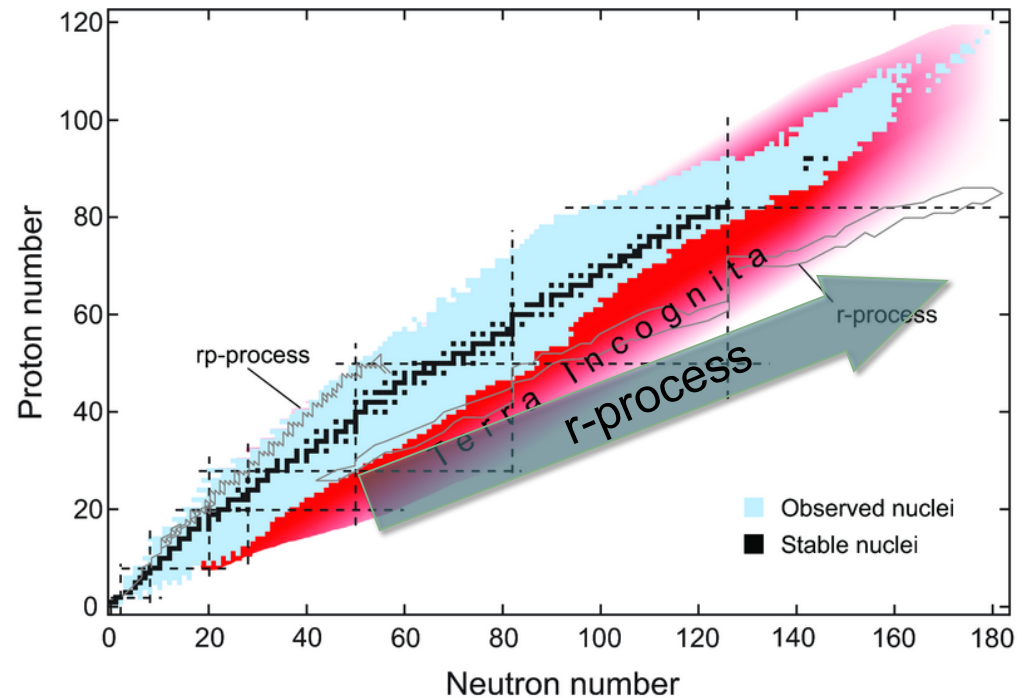
$^{56}\text{Ni}(p,\gamma)^{57}\text{Cu}$ rate



From neutron star mergers to reactions



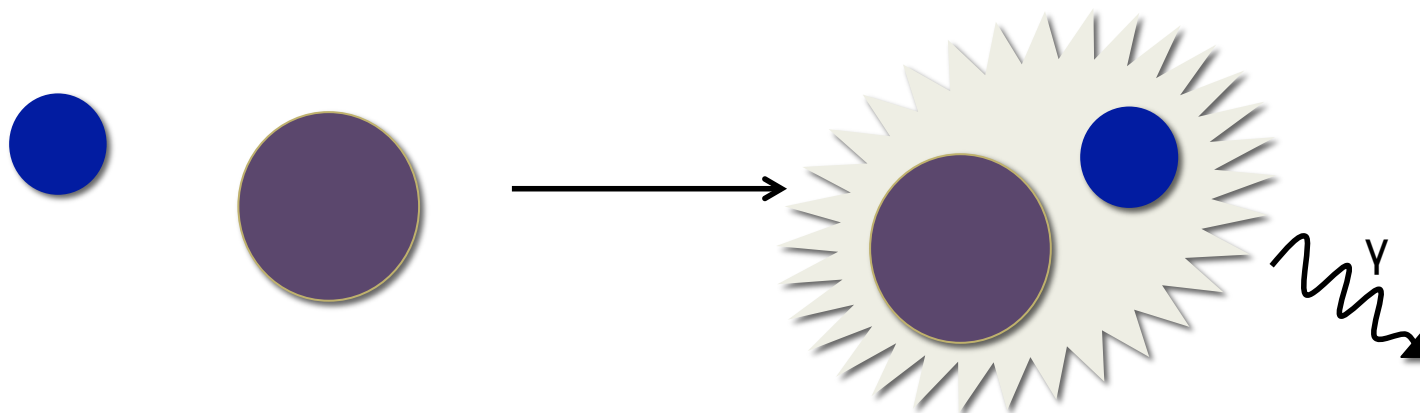
carnegiescience.edu/



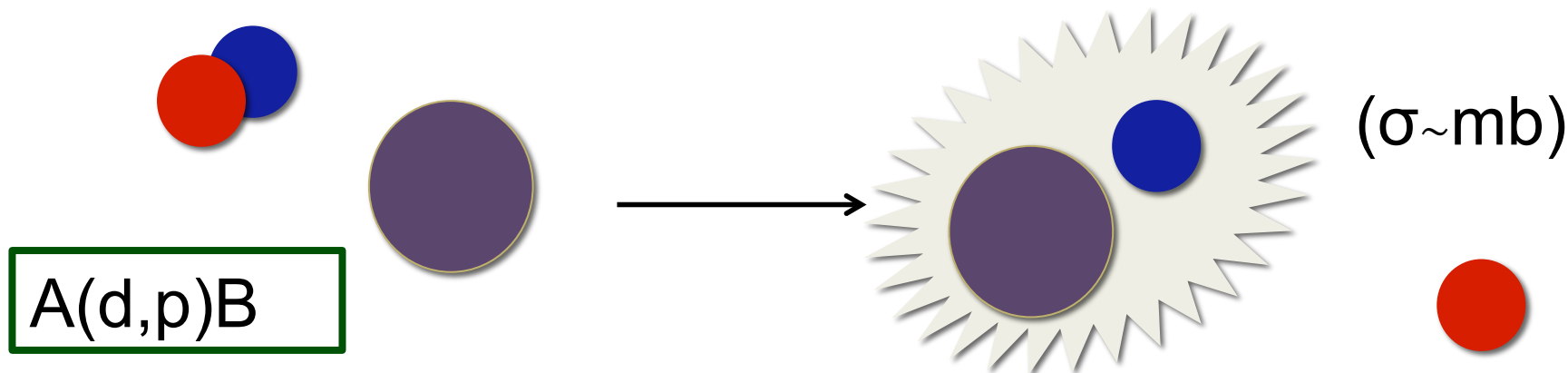
need:
neutron capture rates

how do we measure neutron capture on unstable nuclei?

✧ (n,g) cross sections on unstable nuclei: **Currently Impossible!**



✧ transfer offers an indirect measurement!



Applications to recent (n, γ) measurements

$^{95}\text{Mo}(d,p\gamma)^{96}\text{Mo}$ to extract $^{95}\text{Mo}(n,\gamma)^{96}\text{Mo}$

Compound nucleus (n, γ) is determined through:

$$\sigma_{\alpha\chi}(E_n) = \sum_{J,\pi} \sigma_{\alpha}^{\text{CN}}(E_{ex}, J, \pi) G_{\chi}^{\text{CN}}(E_{ex}, J, \pi)$$

xs for formation of CN
(depends on OP)

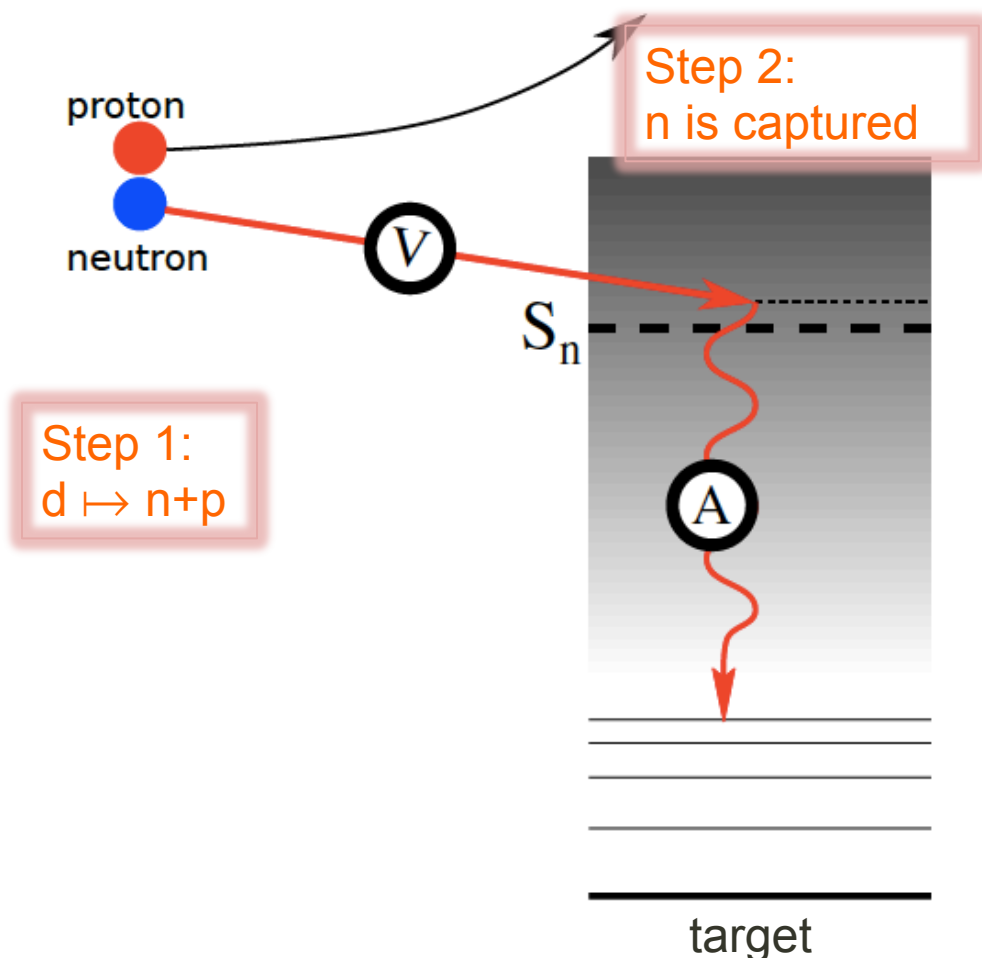
branching ratios
from surrogate
experiment

Compound nucleus (d,p γ) is determined through:

$$P_{\delta\chi}(E_{ex}, \theta_p) = \sum_{J,\pi} F_{\delta}^{\text{CN}}(E_{ex}, J, \pi, \theta_p) G_{\chi}^{\text{CN}}(E_{ex}, J, \pi)$$

Theory for deuteron induced transfer: populating compound states in continuum

✧ Two-step process



Source term generates
flux from breakup

$$S = (\chi_p | (U_{Ap} - U_{Ad} + U_{An}) | \chi_d \phi_d)$$

Neutron propagates in the field
of the target after breakup

$$\Phi_n = G_B^{\text{opt}} S$$



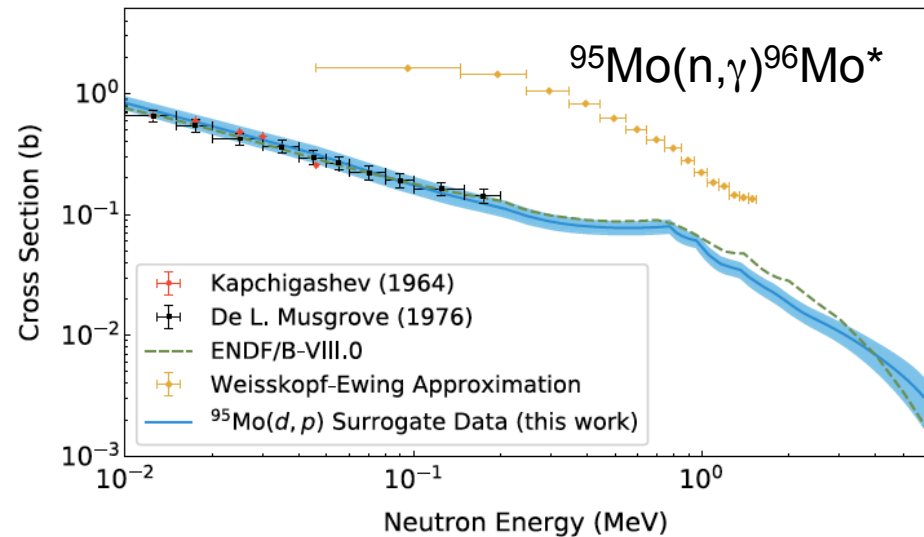
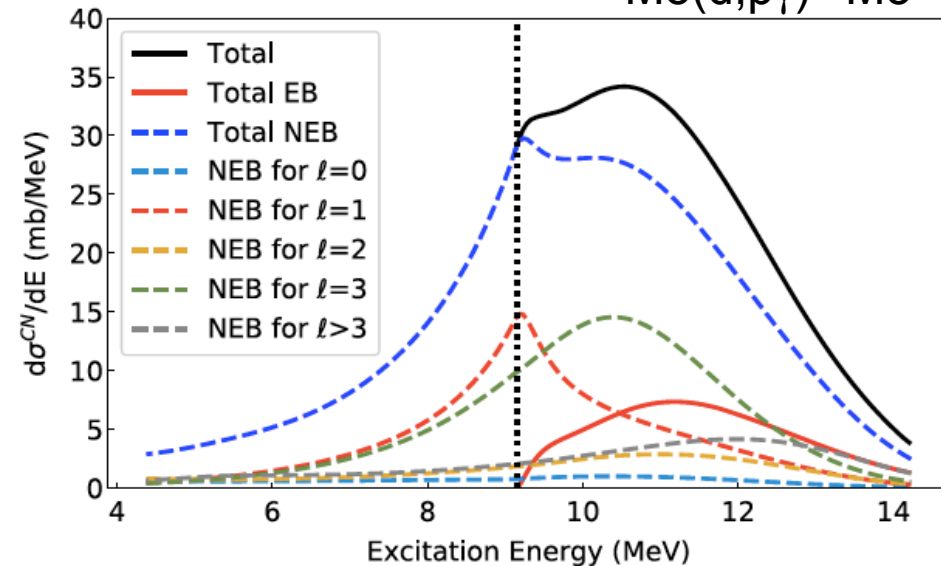
Applications to recent (n, γ) measurements

Compound nucleus (d,p γ) is determined through:

$$P_{\delta\chi}(E_{ex}, \theta_p) = \sum_{J,\pi} F_{\delta}^{\text{CN}}(E_{ex}, J, \pi, \theta_p) G_{\chi}^{\text{CN}}(E_{ex}, J, \pi)$$

L-distributions in (d,p) are different from those in (n,g)
reaction theory provides essential input

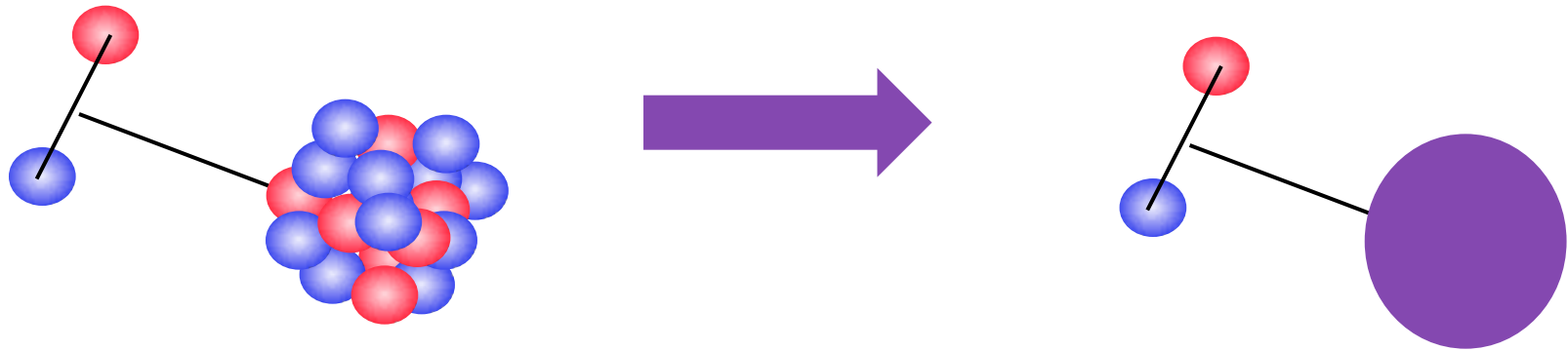
$^{95}\text{Mo}(d,p\gamma)^{96}\text{Mo}^*$



Outline

- ✧ Transfer reactions in the astrophysical context
- ✧ Source of uncertainties in reaction models
- ✧ **Uncertainty quantification approaches**
- ✧ Bayesian analysis to reduce uncertainties
- ✧ Challenges and Opportunities

Mapping the many-body onto the few-body

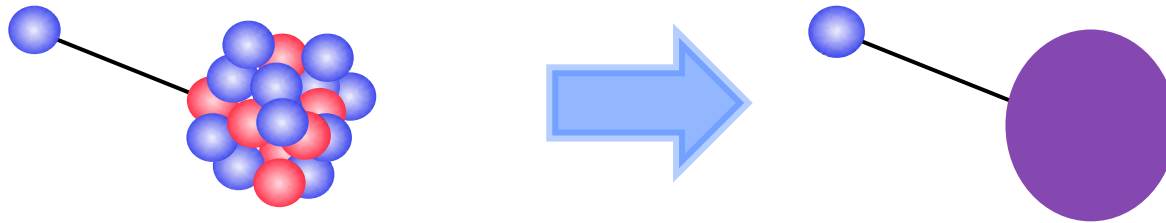


✓ solving the few-body scattering problem?

? the effective interactions (parametric uncertainties)

? target excitations/3-body force/other channels
(model uncertainties)

Effective interactions? Optical potentials



U_{opt} is intrinsically non-local, L-dependent, E-dependent

Phenomenological approach:

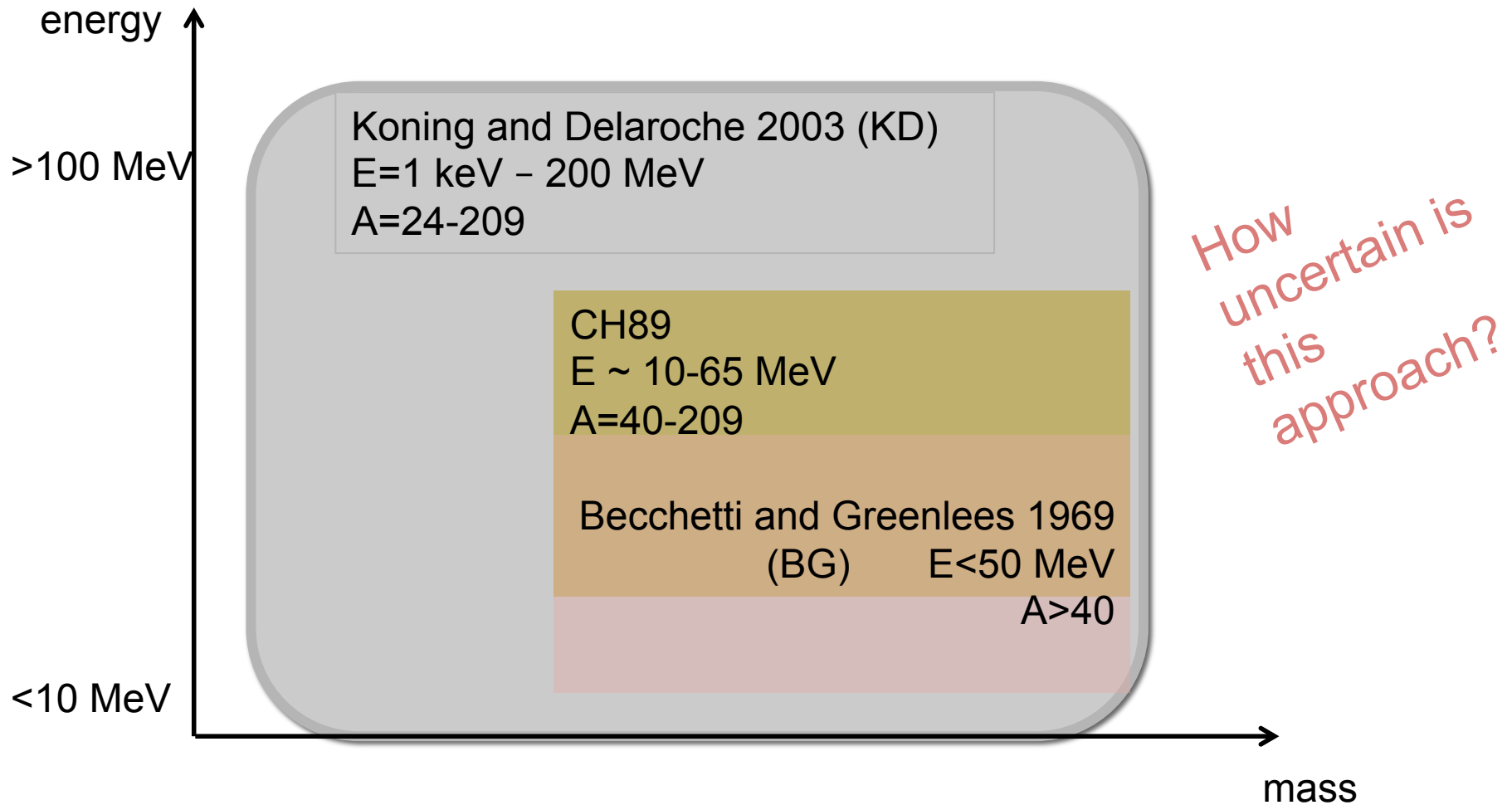
fit a set of data (typically elastic scattering)

extract model **parameters** of an assumed shape

typically local, L-independent, strongly E-dependent

$$U(R) = V f(R, r, a) + W f(R, r_w, a_w) + W_s f(R, r_s, a_s) + V_{so} + V_C$$

A map of global optical potentials



How to quantify uncertainties

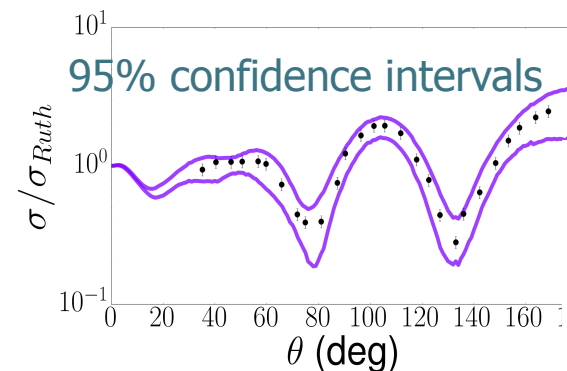


We develop a hypothesis (model) with a set of parameters (priors)

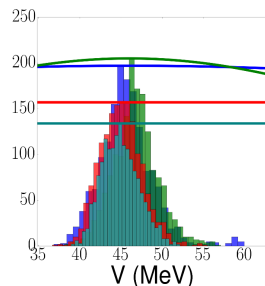
$$\text{optical model} \\ [T+U(R)-E]\psi=0$$

We confront it with reality (data) typically elastic scattering angular distributions (likelihood)

Use Bayes' Theorem + Markov Chain Monte Carlo to sample parameter space



Constraints on the model



Setting up the UQ part



Priors: Gaussians with mean at the BG global parameters and width 100% of the mean

Data: real data versus mock generated from KD global parameters with 10% error

Likelihood: assumption that data points are independent and errors are normally distributed

θ : parameters

σ : independent errors

x : angles

y : experimental cross section

f : model prediction for cross section

$$p(\mathbf{D}|\theta, f, \{\sigma_i^2\}) \propto \exp\left(-\frac{1}{2} \sum_{i=1}^n \frac{(y_i - f(x_i, \theta))^2}{\sigma_i^2}\right)$$

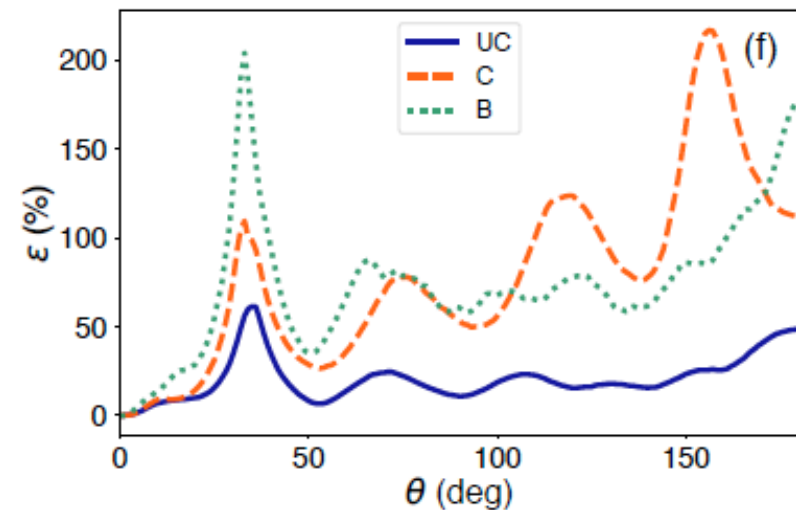
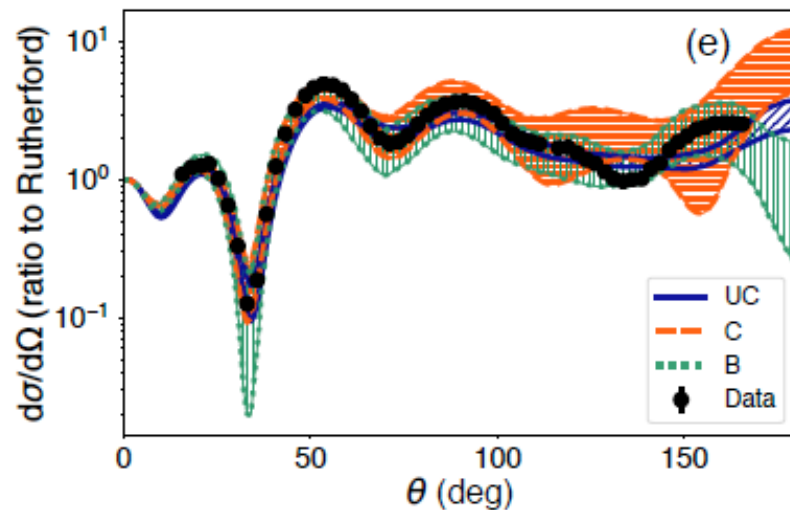
The formulation of the likelihood matters

UC – uncorrelated chi2 frequentist

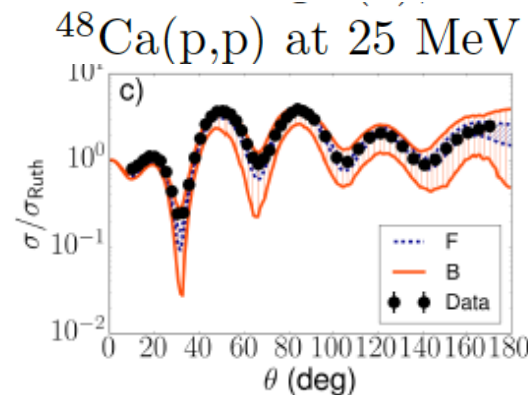
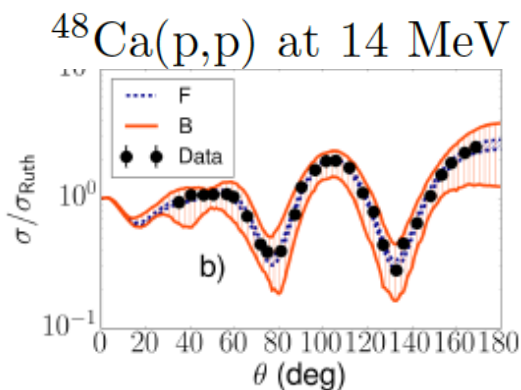
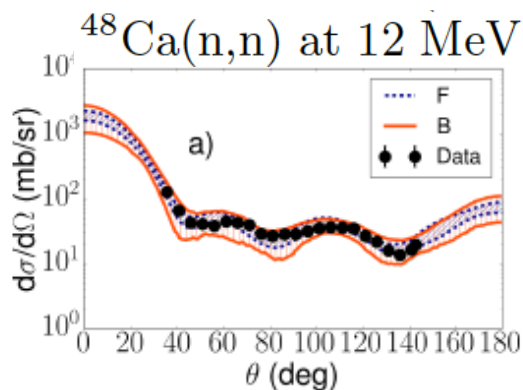
C – correlated chi2 frequentist

B - Bayesian

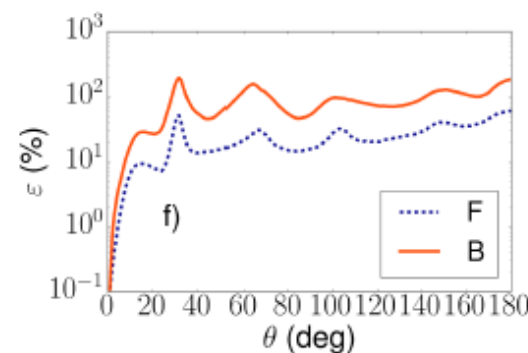
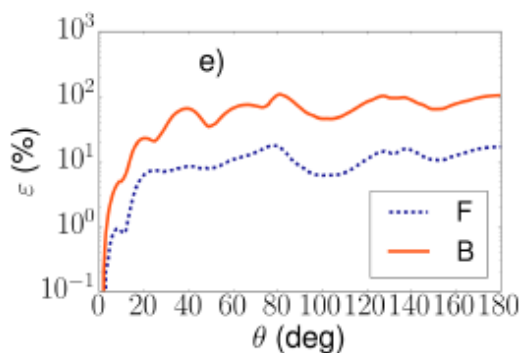
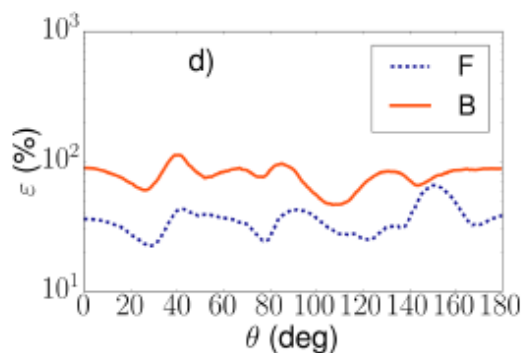
$^{40}\text{Ca}(p,p)^{40}\text{Ca}$ at 26.3 MeV



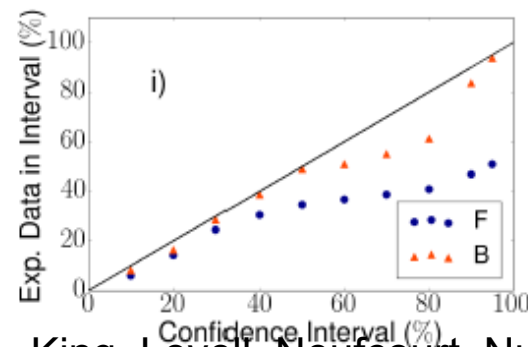
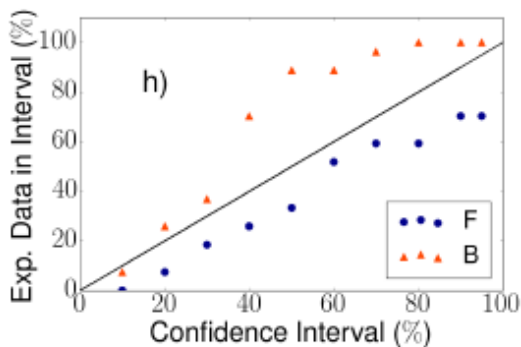
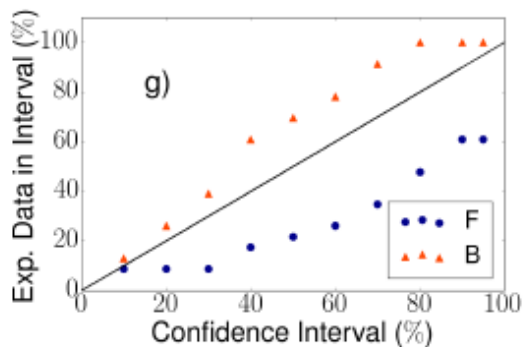
Optical model uncertainties: comparing frequentist and Bayesian



Cross section
angular
distributions



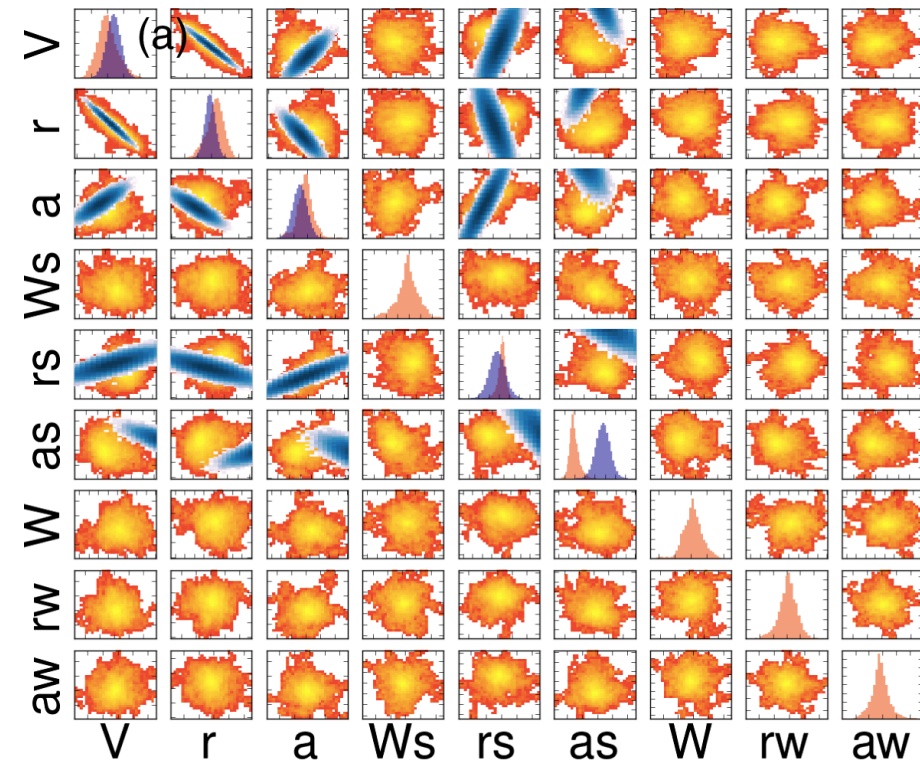
Percentage
uncertainty
width



Empirical
coverage

Optical model uncertainties: comparing frequentist and Bayesian

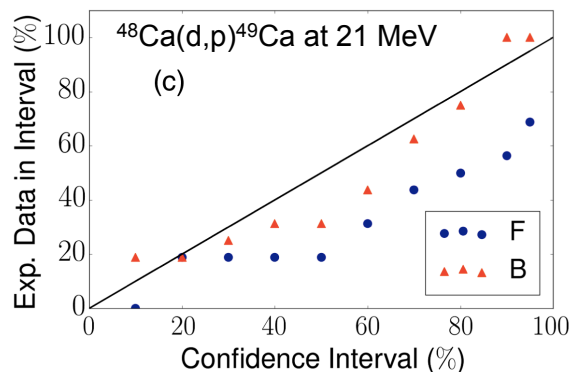
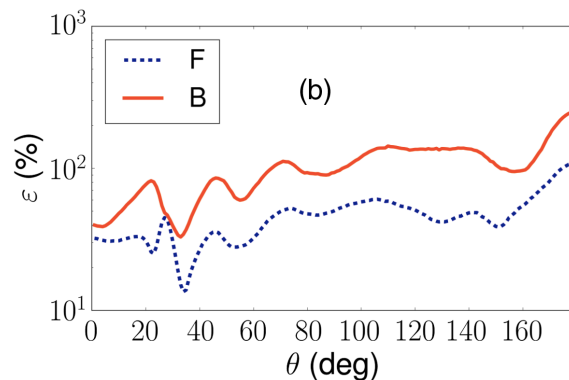
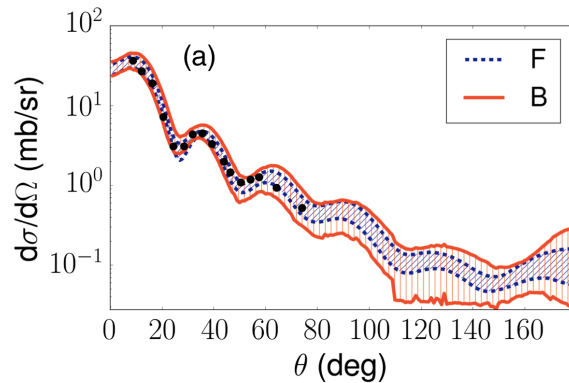
$^{48}\text{Ca}(n,n)$ at 12 MeV



parameter correlations in Bayesian
look very different to the frequentist
approach

blue (frequentist)
orange (Bayesian)

Propagating optical model uncertainties to (d,p) comparing frequentist and Bayesian



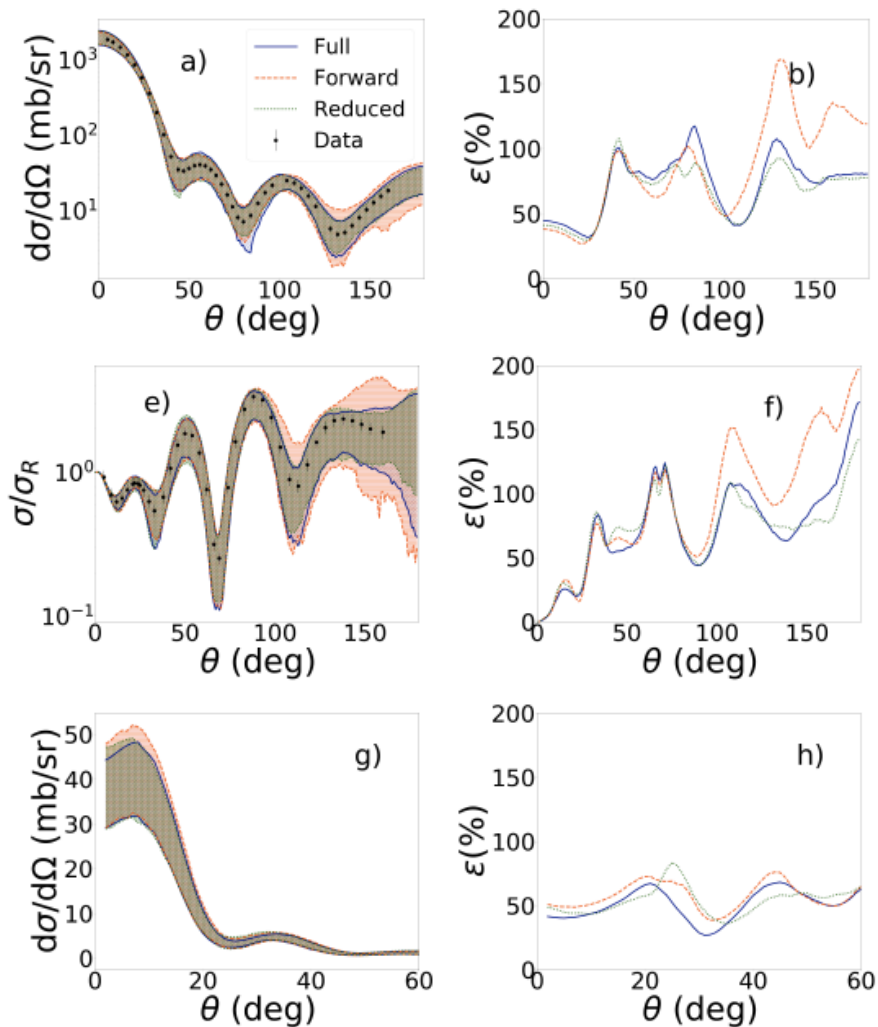
Uncertainties are larger than previously thought

Must explore other ways to reduce optical potential uncertainties

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- ✧ **Bayesian analysis to reduce uncertainties**
- ✧ **Challenges and Opportunities**

Angular information needed?

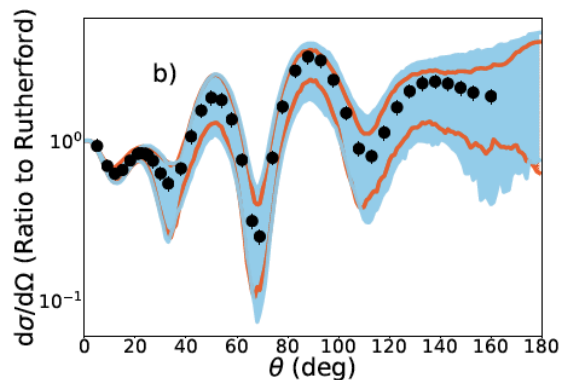
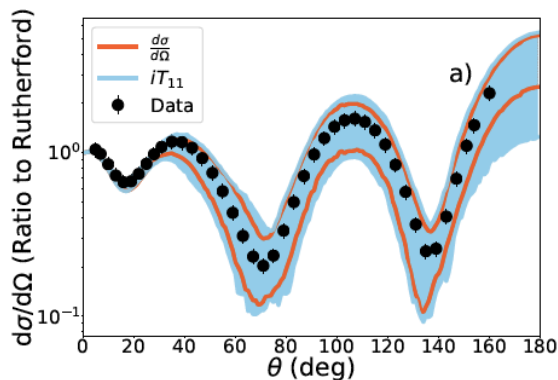


$^{48}\text{Ca}(n,n)^{48}\text{Ca}$ at 12 MeV

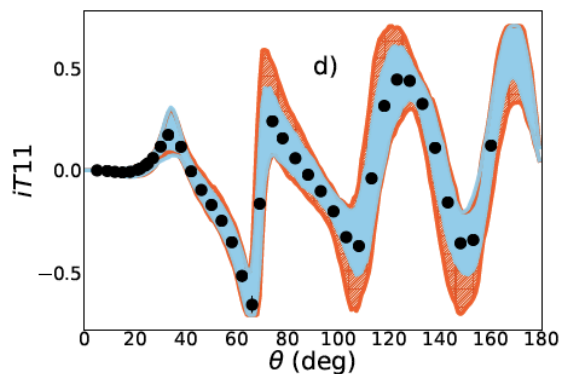
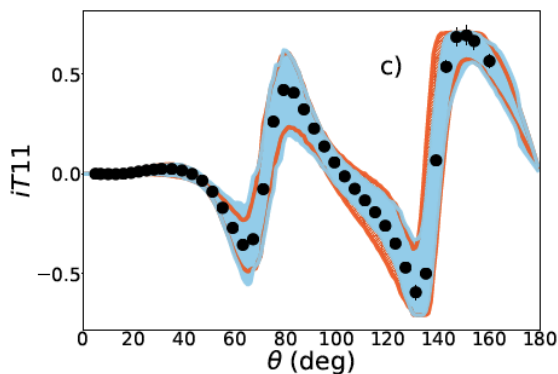
$^{48}\text{Ca}(p,p)^{48}\text{Ca}$ at 21 MeV

$^{48}\text{Ca}(d,p)^{49}\text{Ca}$ at 21 MeV

Which observable offers best constraint?



$^{48}\text{Ca}(p,p)^{48}\text{Ca}$ at 12 MeV



$^{48}\text{Ca}(p,p)^{48}\text{Ca}$ at 21 MeV

Observable	E (MeV)	V (MeV)	r (fm)	a (fm)	W_s (MeV)	r_s (fm)	a_s (fm)
$\frac{d\sigma}{d\Omega}$	12.0	59.48 (4.12)	1.173 (0.052)	0.699 (0.051)	9.476 (0.960)	1.294 (0.084)	0.571 (0.049)
iT_{11}	12.0	60.65 (5.22)	1.159 (0.057)	0.699 (0.067)	9.704 (0.954)	1.273 (0.079)	0.595 (0.080)
$\frac{d\sigma}{d\Omega}$	21.0	55.57 (4.11)	1.178 (0.052)	0.661 (0.057)	7.857 (0.767)	1.297 (0.083)	0.572 (0.051)
iT_{11}	21.0	57.16 (4.44)	1.165 (0.047)	0.691 (0.046)	8.011 (1.007)	1.260 (0.073)	0.579 (0.076)

What is the information content of the data?

Bayesian evidence: provides information contained in a data set.
Integral of the likelihood times the prior over full parameters space

$$p(d|\mathcal{M}) = \int_{\Omega_{\mathcal{M}}} p(d|\alpha, \mathcal{M})p(\alpha|\mathcal{M})d\alpha_{\mathcal{M}}$$

R	Strength of evidence
1 to 3.2	Not worth more than a bare mention
3.2 to 10	Substantial
10 to 100	Strong
> 100	Decisive

Kass and Raftery,
J. Amer. Stat. Assoc 9 (430) 791

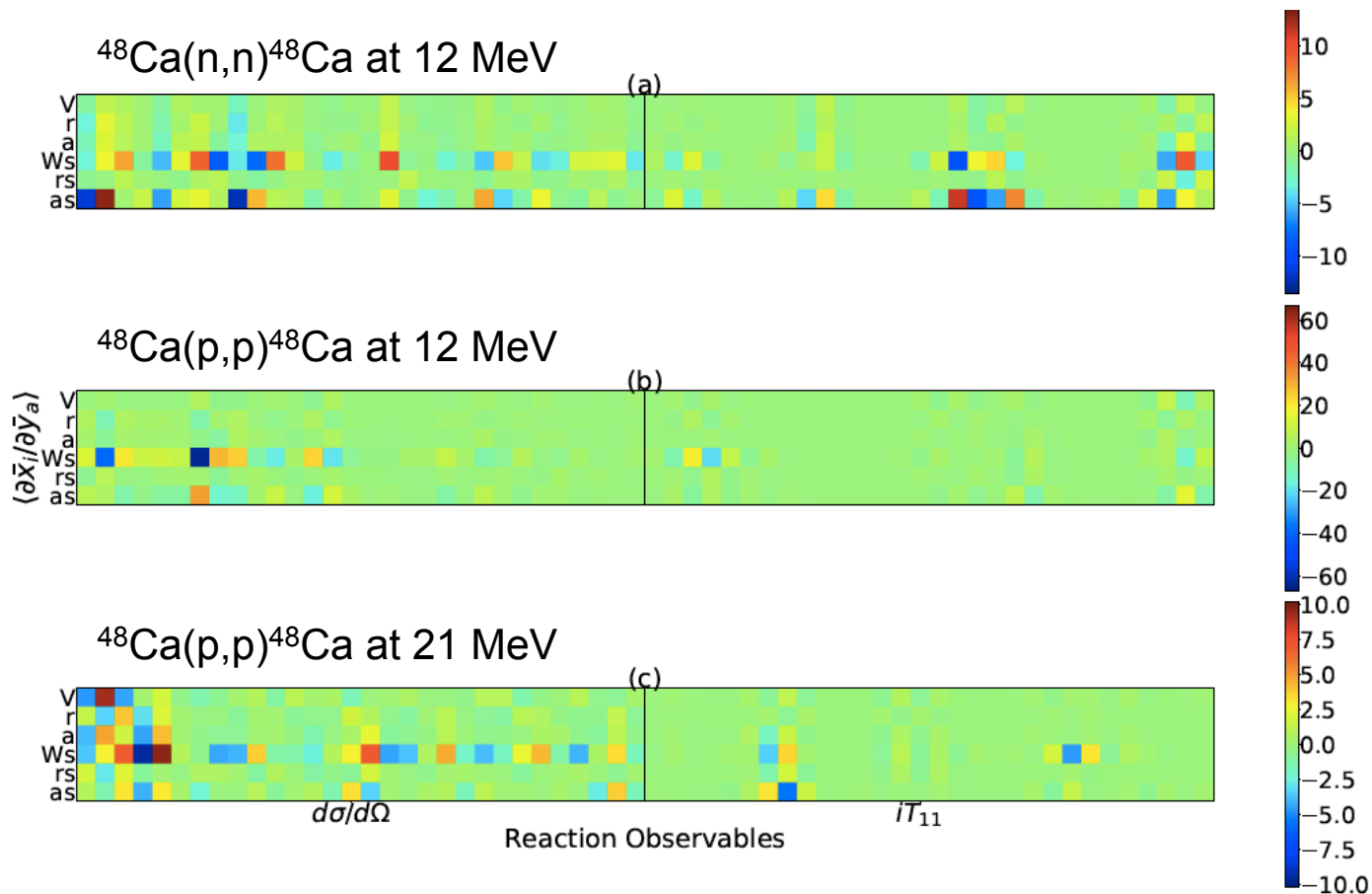
Reaction	$\bar{p}(d \mathcal{M})_{(d\sigma/d\Omega)}$	$\bar{p}(d \mathcal{M})_{(iT_{11})}$	\bar{R}
$^{48}\text{Ca}(n,n)$ at 12 MeV	0.198(0.017)	0.190(0.044)	1.04
$^{48}\text{Ca}(p,p)$ at 12 MeV	0.142(0.043)	0.112(0.035)	1.27
$^{48}\text{Ca}(p,p)$ at 21 MeV	0.171(0.036)	0.118(0.027)	1.44
$^{208}\text{Pb}(n,n)$ at 30 MeV	0.016(0.003)	0.039(0.010)	0.42
$^{208}\text{Pb}(p,p)$ at 30 MeV	0.233(0.044)	0.086(0.018)	2.72
$^{208}\text{Pb}(p,p)$ at 61 MeV	0.157(0.051)	0.196(0.049)	0.80

TABLE II: Bayesian evidence (multiplied by 10^{-3}) for the different reactions considered: using only cross section data (2nd column), using only polarization data (3rd column), and the ratio between the Bayesian evidence with cross section data over that with polarization data (the Bayes' factor).

What do we learn from sensitivities?

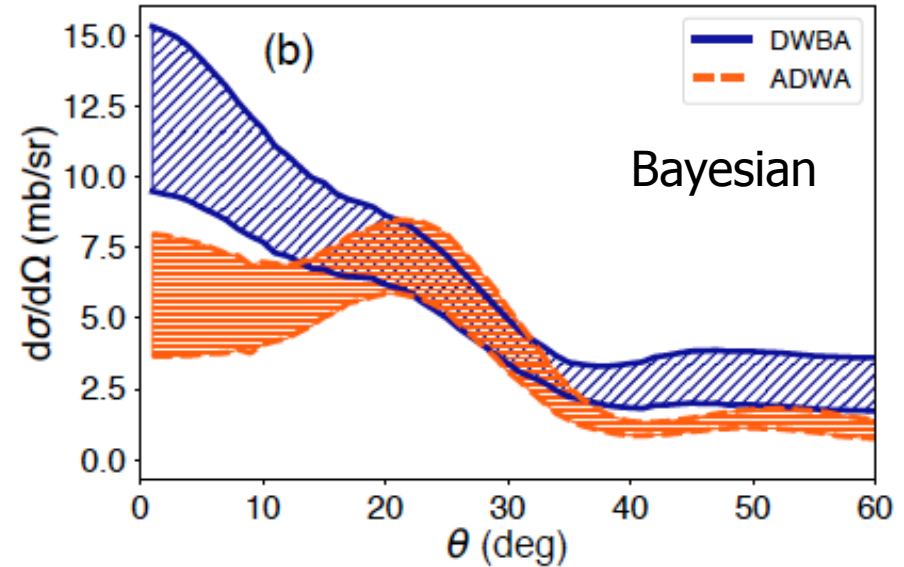
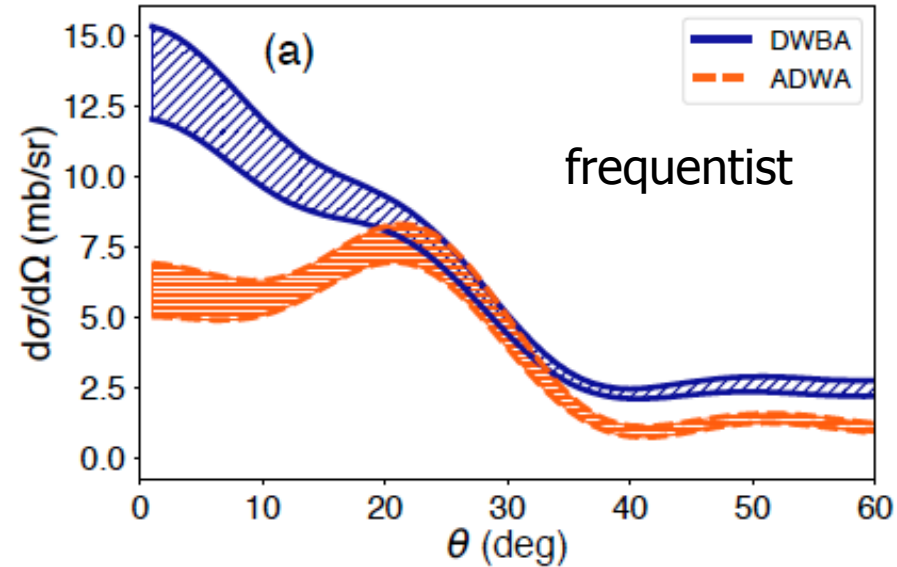
how much variation in parameter x_i is produced by a variation on observable y_a ?

$$\left\langle \frac{\partial \bar{x}_i}{\partial \bar{y}_a} \right\rangle = \tilde{\mathbf{C}}_{ib} \mathbf{C}_{ba}^{-1}$$



Comparing models

$^{40}\text{Ca}(d,p)^{41}\text{Ca}$ at 28.4 MeV



Can we discriminate between models?

Opportunities for the next 5yrs

- ❑ go beyond the simplest reaction models
- ❑ quantify model uncertainties
- ❑ perform model mixing

Computational challenge: need Emulators

Eigenvector continuation method for scattering
to reduce the dimensionality of the problem

Furnstahl et al. PLB 809 (2020) 135719

$$|\psi_{\text{trial}}\rangle = \sum_{i=1}^{N_b} c_i |\psi_E(\boldsymbol{\theta}_i)\rangle$$

QUILTR

Quantifying Uncertainties in Low energy Theory for Reactions

It's a suite of codes including Bayesian MCMC for optical model calibration and propagation of uncertainties to transfer reactions



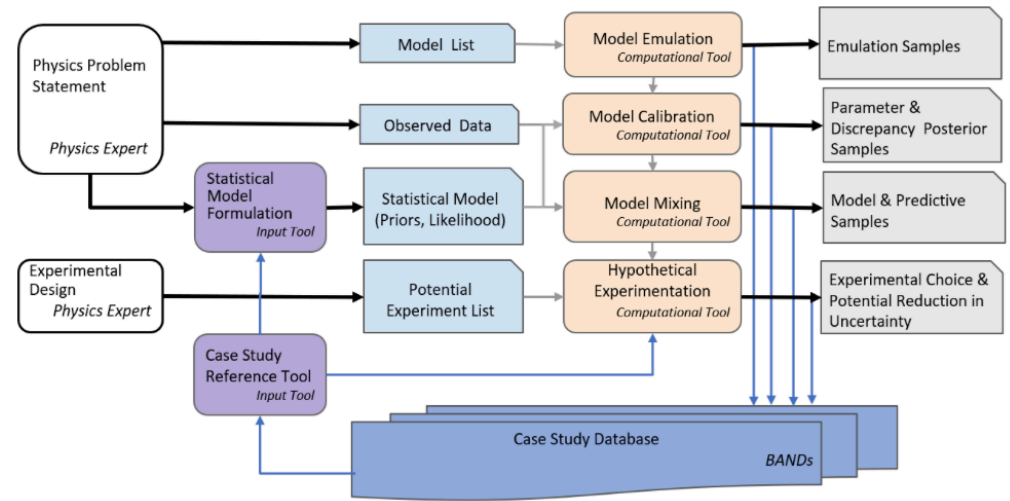
Wrappers built on:
FRESCO (by Ian Thompson)

(by Amy Lovell)

BAND

Bayesian Analysis of Nuclear Dynamics

<https://bandframework.github.io>



BAND CAMP

**Bayesian Analysis of Nuclear Dynamics (BAND)
FIRST ANNUAL BAND CAMP**

Monday, December 14, 2020
9 AM-5 PM, EST

In association with the [ISNET v8 Workshop](#), the BAND Collaboration is sponsoring a one-day series of on-line pedagogical lectures aimed at providing a foundation for nuclear physicists in modern Bayesian statistical methodologies. The program consists of three extended lectures, including discussion.

PROGRAM

Michael Grosskopf

Computer, Computational, and Statistical Sciences Division, Los Alamos National Laboratory
"Bayesian Basics"

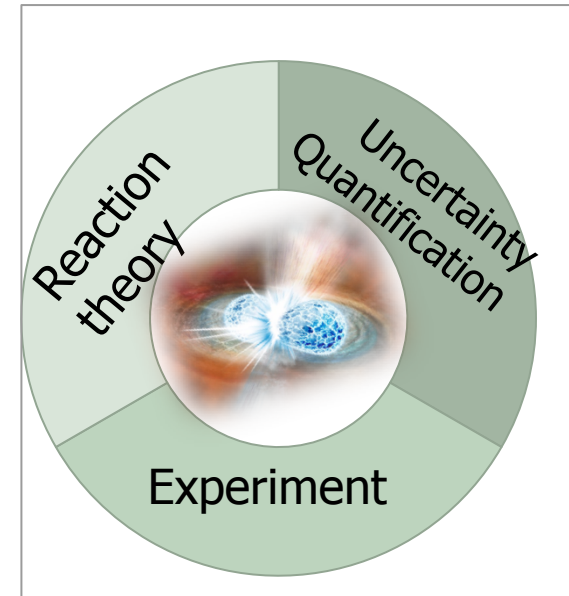
Simon Mak

Department of Statistical Science, Duke University
"Applications of Model Emulators for Parameter Estimation"

Matthew Pratola

Department of Statistics, The Ohio State University
"Model Mixing and Averaging"

In conclusion:



- Transfer reactions offer a versatile tool for extracting capture rates for astrophysics (many recent applications with impact on astrophysics)
- Reactions theories is needed to interpret the indirect measurements and obtain reliable capture rates
- Uncertainty quantification is an essential ingredient and Bayesian analysis offers many new avenues and a promising future of collaboration with experiment.

Thanks to my collaborators



Amy Lovell
(now staff at LANL)



Garrett King
(now GS WashU)



Manuel Catacora Rios
(now GS UChicago)



Michael
Quinonez

UQ reactions group

Transfer applications



Gregory Potel
(now at LLNL)



Terri Poxon-Pearson
(now at NNSA)

Many thanks to all for zooming in!