

Physics Case of Measuring the Weak Charge of Carbon-12 at MESA



Misha Gorshteyn

C-12 Workshop - IF UNAM, Mexico City - April 1st, 2019



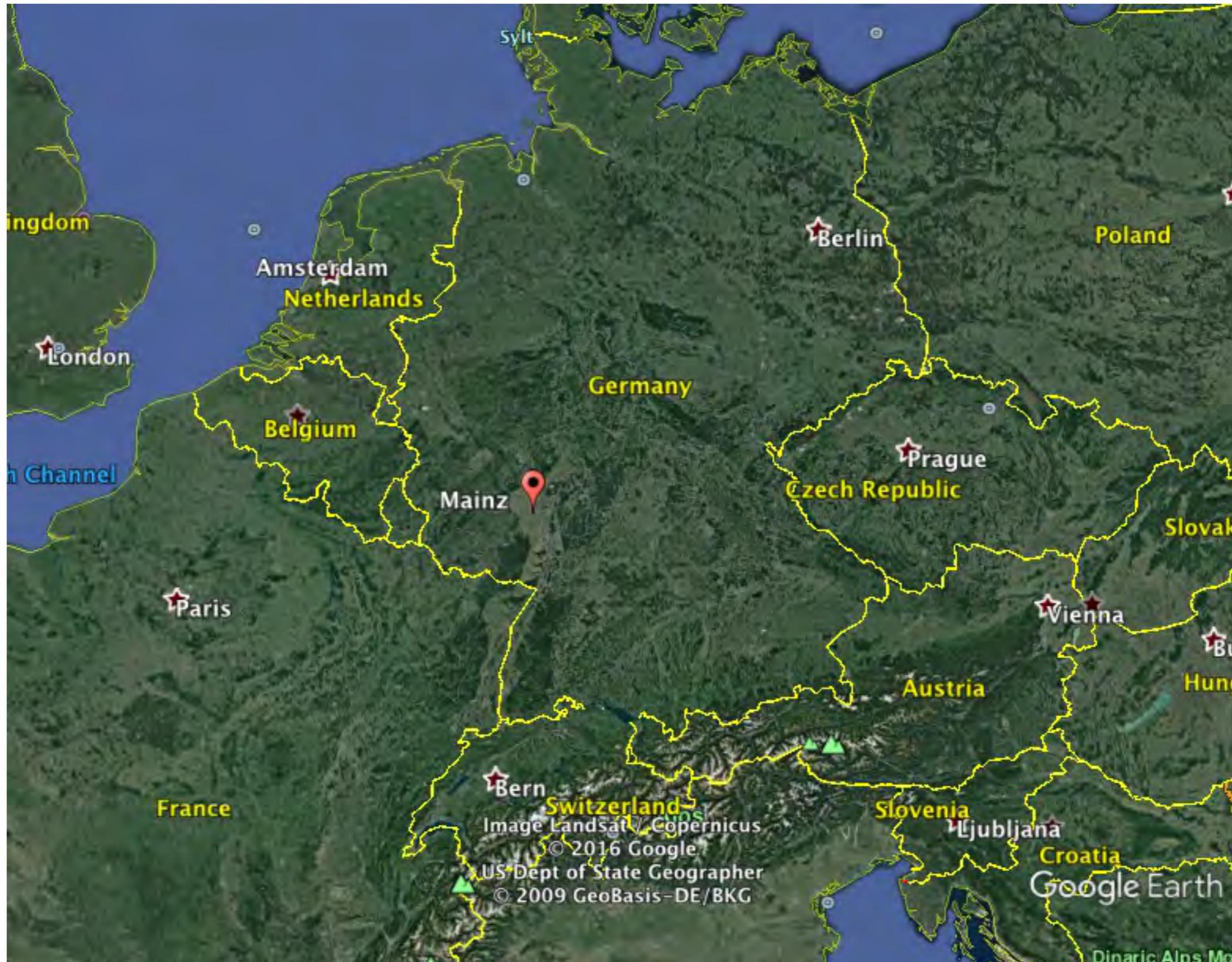
Why are we here?

DFG (Mainz) - CONACyT (UNAM) Cooperation Project

“Constraints on Physics Beyond the Standard Model from Lower Energy Electron Scattering”

- Granted in Fall 2017 - 3 years duration
- The goal: interpretation of a future experiment at Mainz
- In parallel: experimentalists study the feasibility of the measurement

Europe - Germany - Mainz



University of Mainz:

- 1477: founded by Archbishop of Mainz, Prince-Elector Adolf II von Nassau;
- 1792: came to a halt due to French revolutionary wars and political turmoil;
- 1823: last academic program (medicine) ceases;
- 1946: re-founded during French occupation after WWII;

MAinz MIcrotron MAMI:

- 1975: first proposal of a microtron;
- 1979: first electron beam MAMI A1 - 14 MeV;
- 1983: upgrade MAMI A2 - 183 MeV;

- 1990: second upgrade MAMI B - 855 MeV;
- 1992: polarized electrons source;

- 2006: upgrade MAMI C - 1.5 GeV, 100 μ A

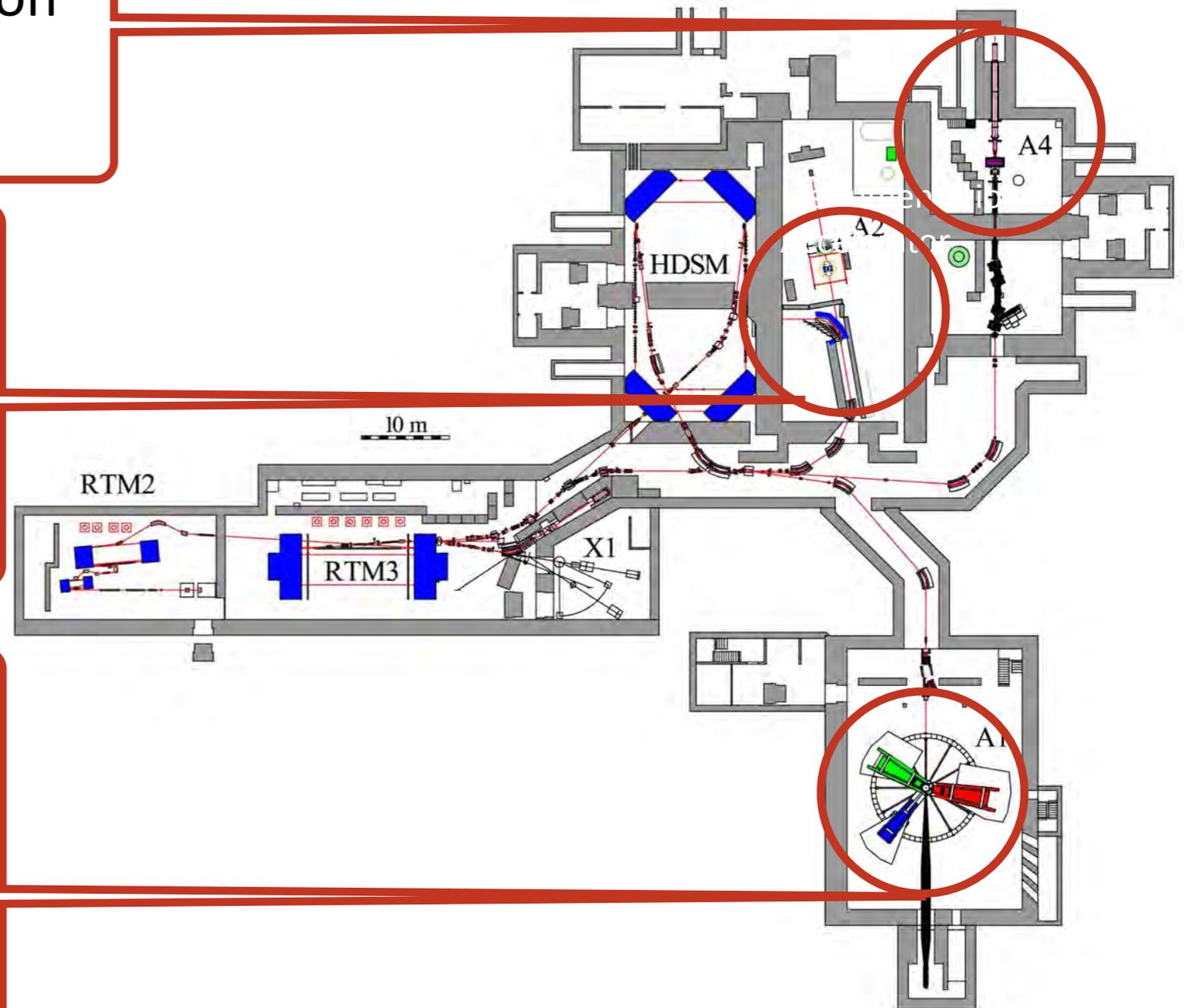


Research highlights @ MAMI

A4 Collaboration: polarized electrons
PV asymmetries $\sim 10^{-5/6}$
Strange form factors of the nucleon
Two-photon exchange
Compton polarimetry

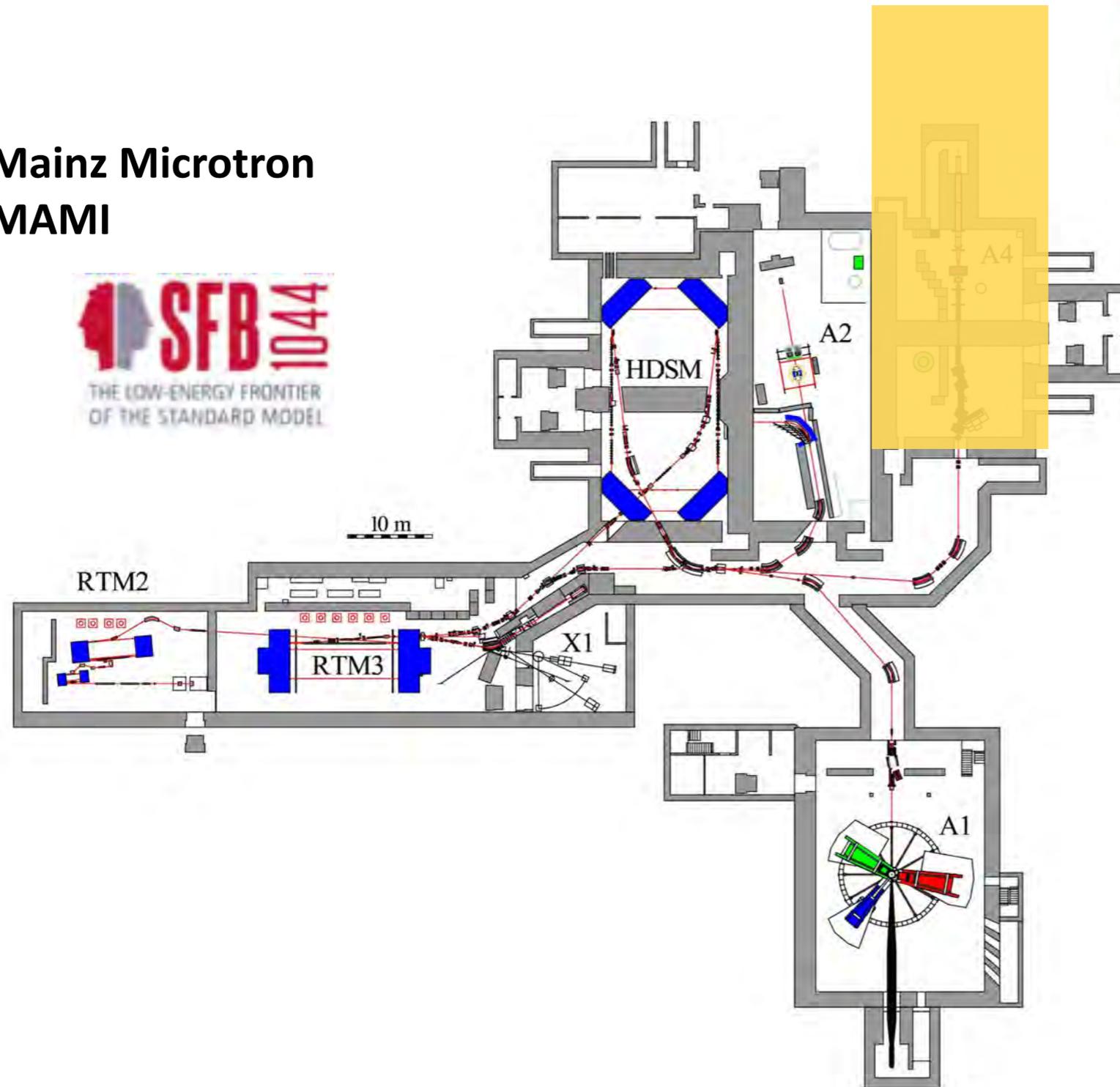
A2 Collaboration: tagged real γ 's
4 π -detectors (TAPS, Crystal Ball)
Nucleon polarizabilities;
GDH sum rule verification;
Pion photoproduction

A1 Collaboration: e-scattering
3 spectrometer arrangement
QE e-scattering; e-m nucleon FFs;
VCS & generalized nucleon pol.
Dark photon search



Low-energy physics in Mainz: old and new

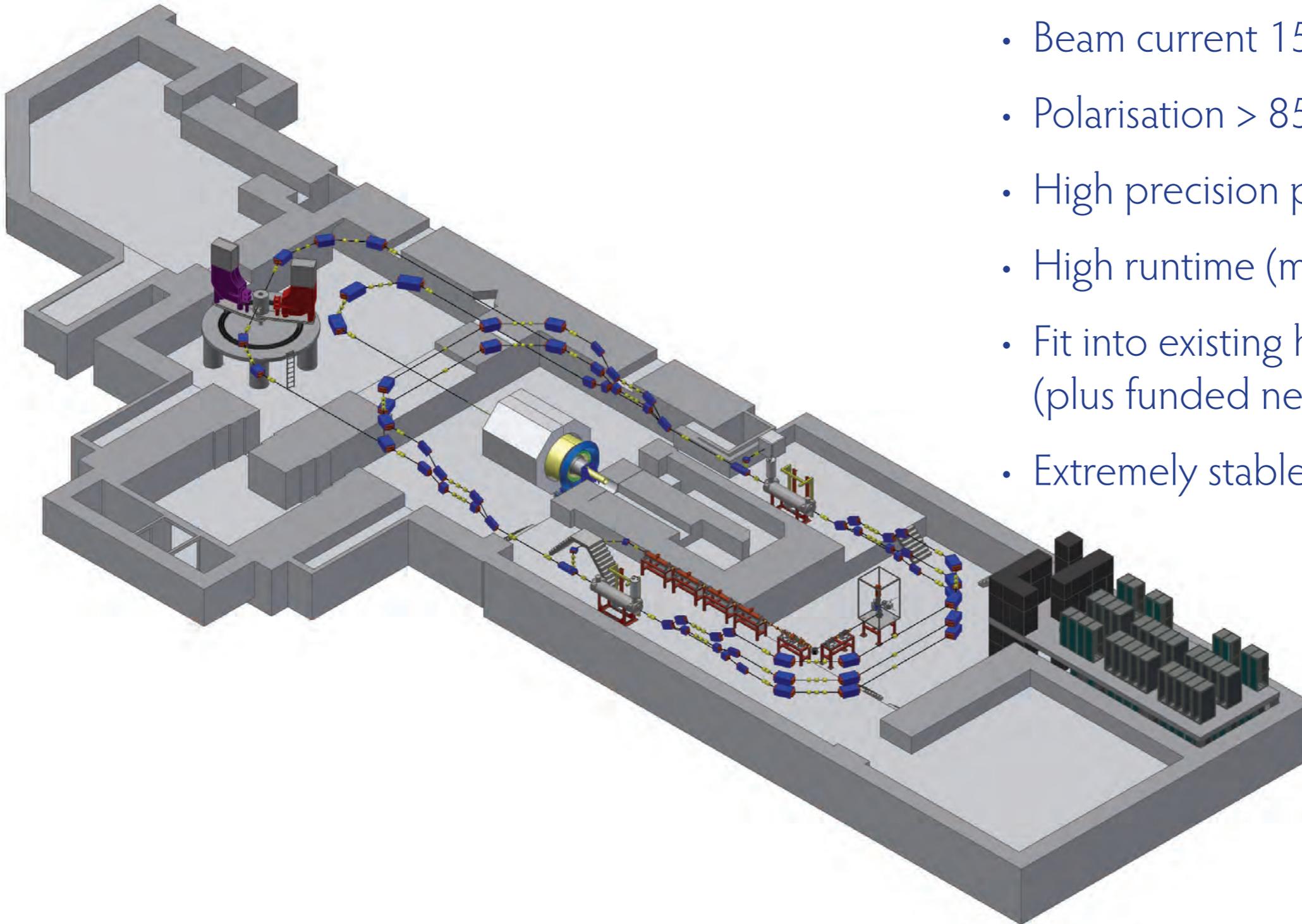
Mainz Microtron
MAMI



MESA
Accelerator

MESA

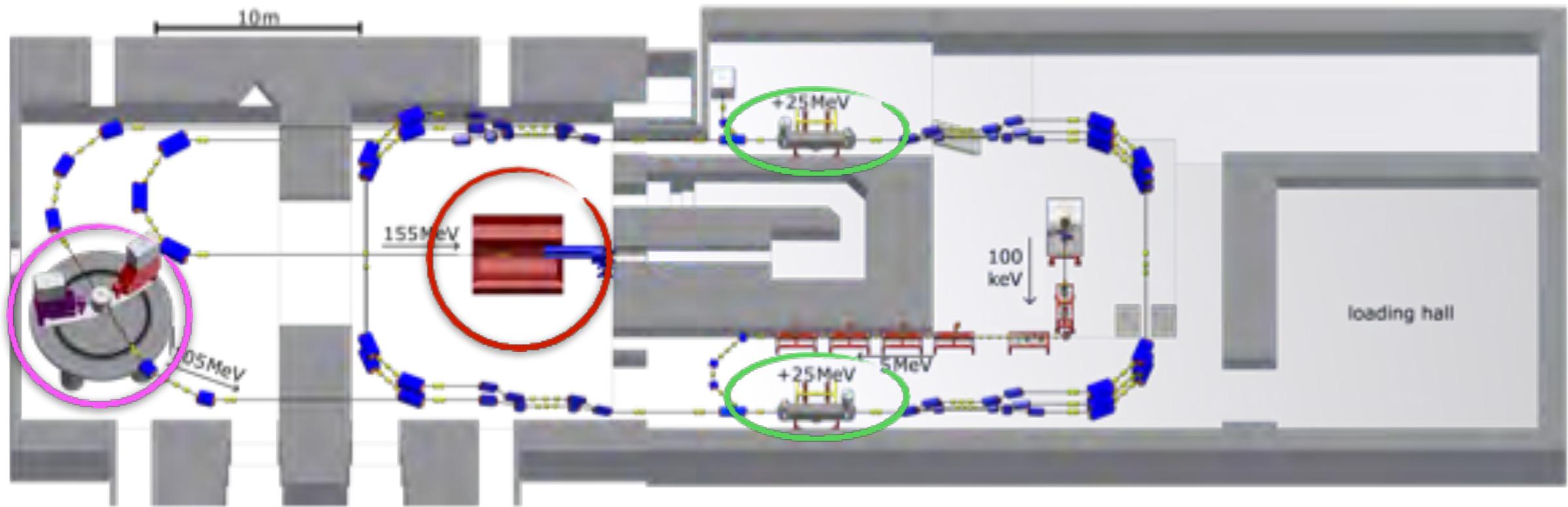
Mainz Energy-recovery Superconducting Accelerator



- Beam current 150 μA
- Polarisation > 85%
- High precision polarimetry
- High runtime (more than 4000 h/year)
- Fit into existing halls at MAMI (plus funded new hall)
- Extremely stable

MESA

Mainz Energy-recovery Superconducting Accelerator



MAGIX

P2

Superconducting cavities

“Extracted Beam” vs. “Energy Recovery” modes:
155 MeV, 150 μ A vs. 105 MeV, up to 10 mA

Collaborative Research Center CRC 1044

The low-energy frontier of the Standard Model

From quarks and gluons to hadrons and nuclei

Sub-projects:

P: Implications of hadron physics to **Precision** tests of Standard Model

P1: $(g-2)_\mu$ and $\alpha_{em}(M_Z^2)$

P2: proton's weak charge and weak mixing angle at low energy

S: Hadron **Structure** via spacelike and timelike processes

S1: Baryon form factors

S2: Polarizabilities

S3: Light hadron spectroscopy

M: **Meson** structure, decays and spectroscopy

M1: $\gamma\gamma$ physics

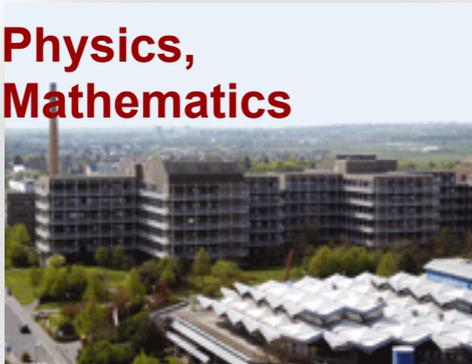
M2: Meson dynamics

N: Interactions in **Nuclear** systems, baryon interactions from QCD

Cluster of Excellence PRISMA+

Participating institutes:

Physics,
Mathematics



Nuclear Physics



Helmholtz Institute
Mainz



Nuclear Chemistry



★ Precision Physics, Fundamental Interactions and Structure of Matter

Research Areas:

- A: Exploring intensity frontier with MESA
- B: Precision physics at low-energy frontier
- C: Exploring the weakly-interacting universe
- D: Physics at high-energy accelerators
- E: Theory and phenomenology of fundamental interactions

Structural Initiatives:

- Electron Accelerator MESA
- TRIGA Reactor and User Facility
- Mainz Institute for Theoretical Physics
- Detector Laboratory

Precise measurement of the
Weak Mixing Angle
with polarized e-p scattering
@ MESA

Weak mixing angle:
central role in the SM

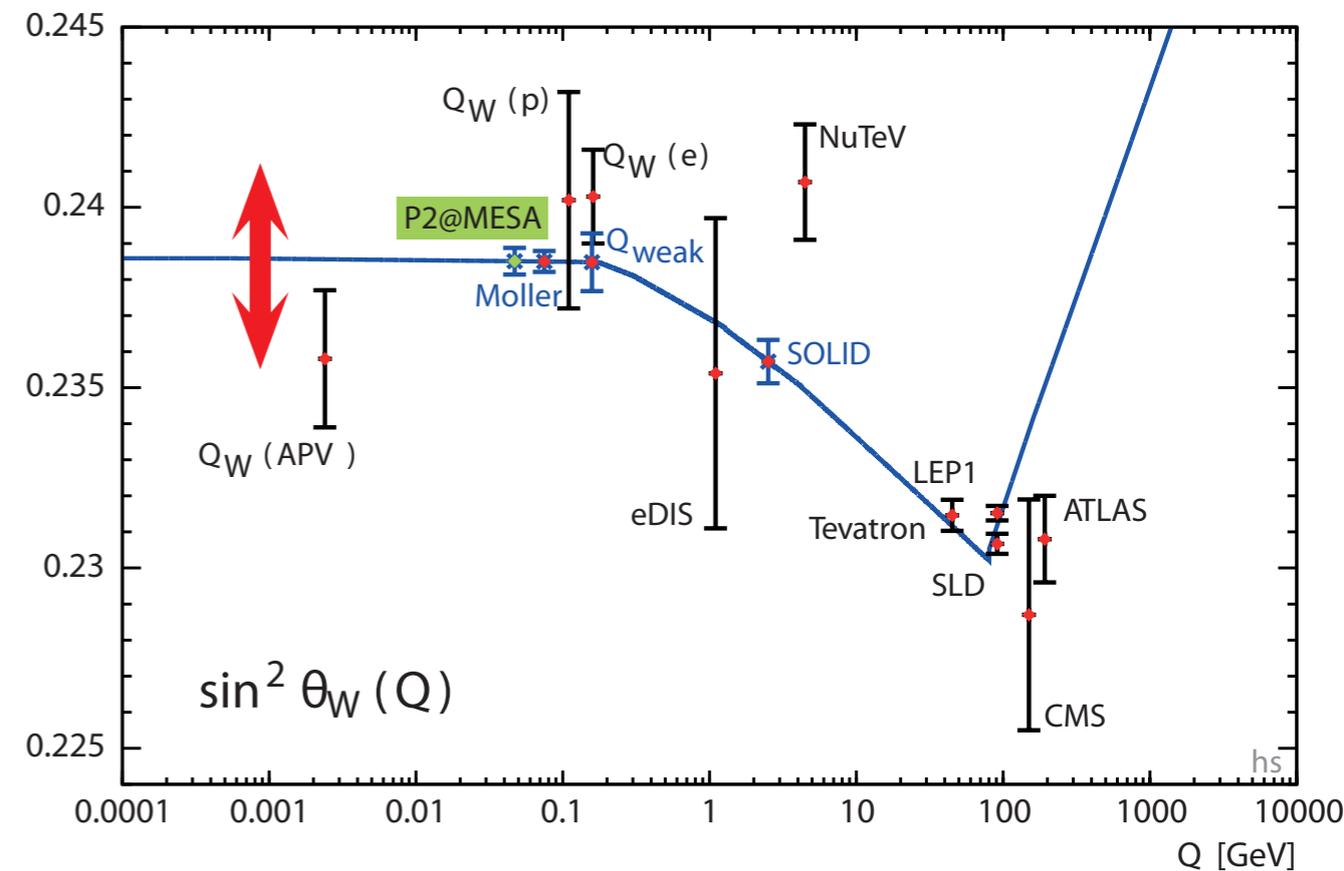
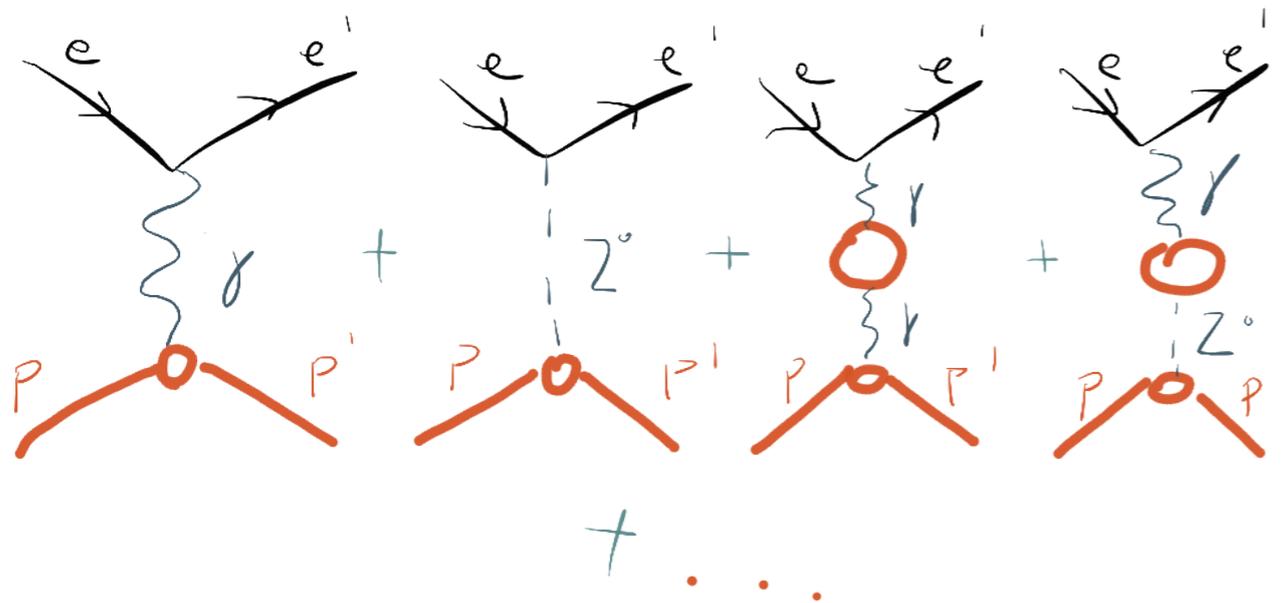
$$\begin{pmatrix} \gamma \\ Z^0 \end{pmatrix} = \begin{pmatrix} \cos \theta_W & \sin \theta_W \\ -\sin \theta_W & \cos \theta_W \end{pmatrix} \begin{pmatrix} B^0 \\ W^0 \end{pmatrix}$$

Tree-level:

$$\sin^2 \theta_W = 1 - \frac{M_W^2}{M_Z^2} = \frac{g'^2}{g^2 + g'^2}$$

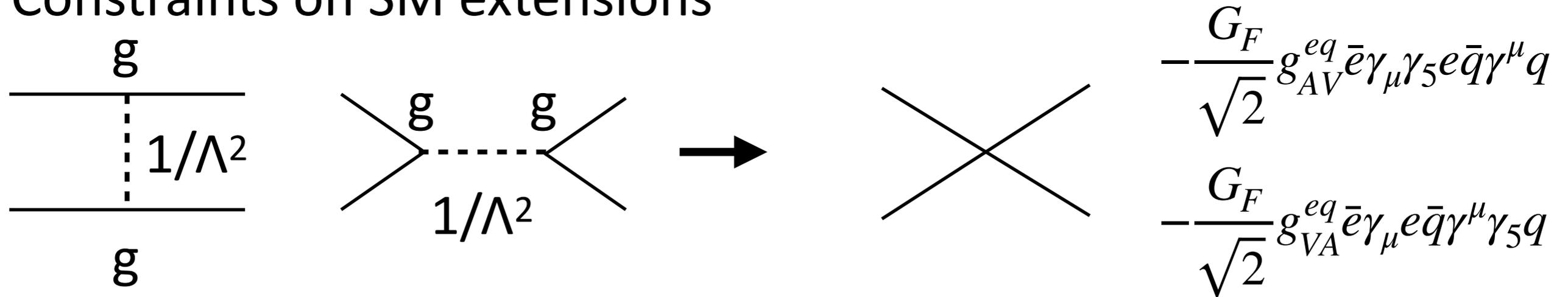
Radiative Corrections: scale dependence

The running is specific for each theory - test of SM



Test of the Standard Model running of $\sin^2\theta_W$

Constraints on SM extensions



In SM at tree level: $(g_{AV}^{eu})^{SM, tree} = \frac{1}{2} - \frac{4}{3} \sin^2 \theta_W$ $(g_{AV}^{ed})^{SM, tree} = -\frac{1}{2} + \frac{2}{3} \sin^2 \theta_W$

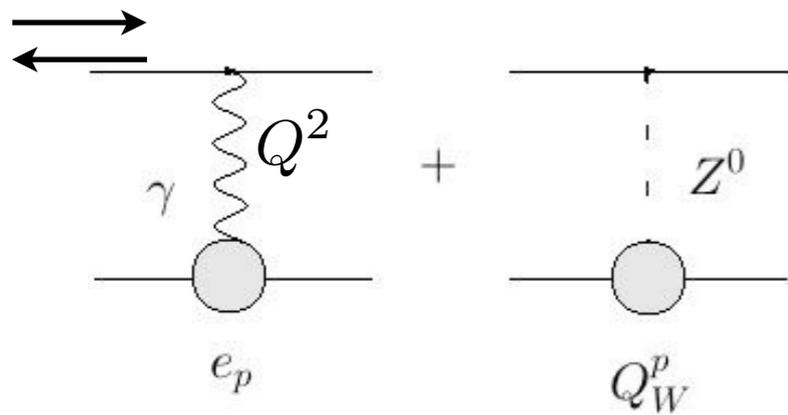
$Q_W^p = 2(2g_{AV}^{eu} + g_{AV}^{ed}) = 1 - 4 \sin^2 \theta_W$ $Q_W^n = 2(g_{AV}^{eu} + 2g_{AV}^{ed}) = -1$

Importantly: vector current is conserved - is not renormalized by QCD

Heavy New Physics: $\delta g_{AV}^{eq} \sim \frac{g_{NP}^2}{g_{SM}^2} \frac{M_Z^2}{\Lambda^2} \sim g_{NP}^2 \frac{v^2}{\Lambda^2}$

Measure the deviations from SM values - discover/constrain BSM

Weak charge of the proton with PV electron scattering



$$A^{\text{PV}} = \frac{\sigma_R - \sigma_L}{\sigma_R + \sigma_L} = -\frac{G_F Q^2}{4\sqrt{2}\pi\alpha} [Q_W^p + Q^2 B(Q^2)]$$

Q_W^p in SM accidentally suppressed $Q_W^{p, \text{tree}} = 1 - 4 \sin^2 \theta_W \approx 0.07$

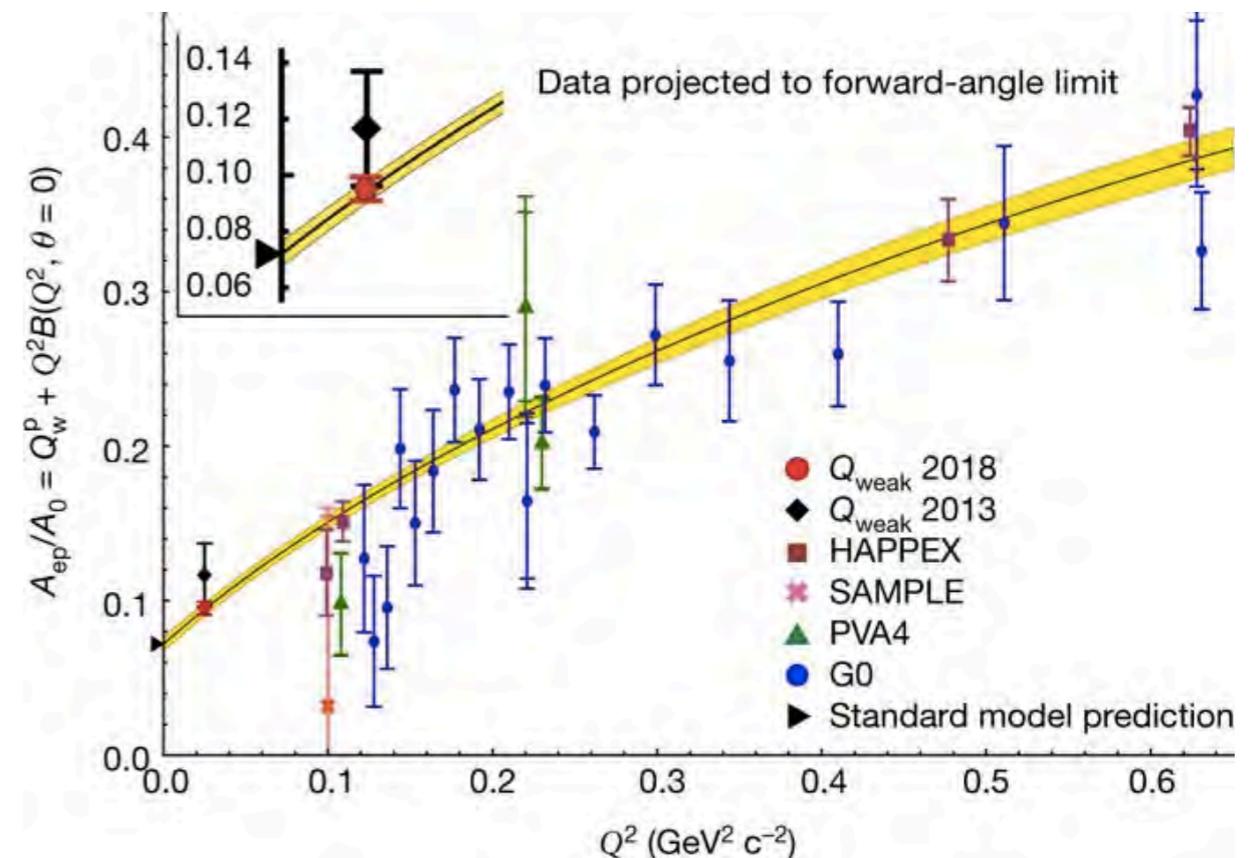
Sensitive test of running of θ_W at low energy $\frac{\delta Q_W^p}{Q_W^p} = \frac{4 \sin^2 \theta_W}{1 - 4 \sin^2 \theta_W} \frac{\delta \sin^2 \theta_W}{\sin^2 \theta_W} \approx 12 \frac{\delta \sin^2 \theta_W}{\sin^2 \theta_W}$

$B(Q^2)$ - from non-forward PVES data

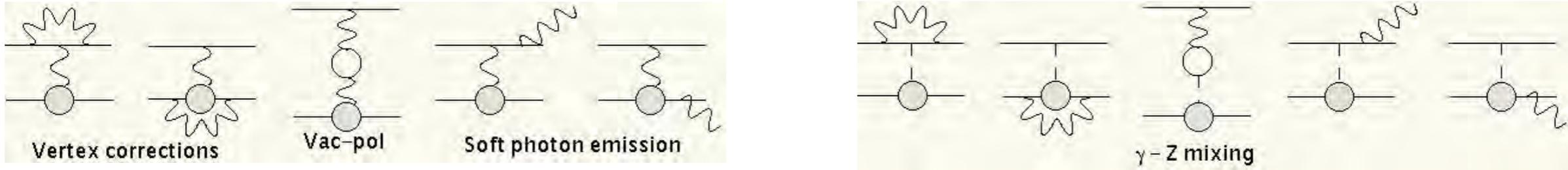
Young et al. '07;

Androic et al. [Qweak Coll.], '13

Plot from Q-Weak publication
(Nature 2018)



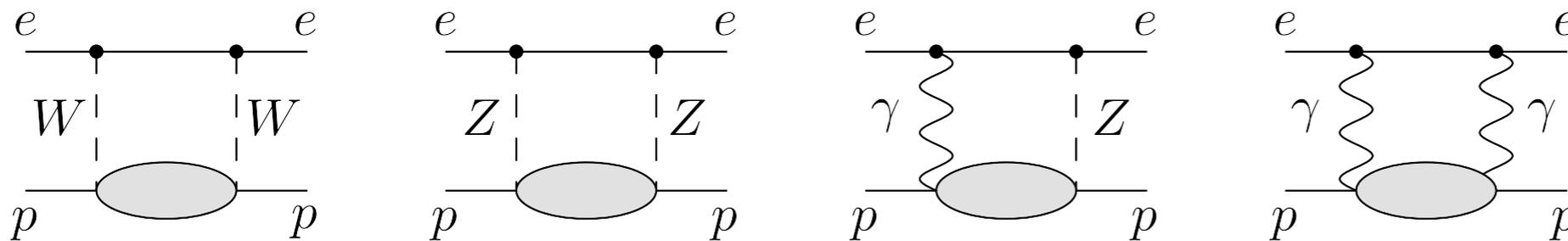
Need to worry about radiative corrections



Hadronic structure effects are under control

$$Q_W^p = (1 + \Delta_\rho + \Delta_e)(1 - 4 \sin^2 \hat{\theta}_W + \Delta'_e) + \square_{WW} + \square_{ZZ} + \square_{\gamma Z}$$

Marciano, Sirlin, '85; Ramsey-Musolf, '99



Q-Weak@JLab

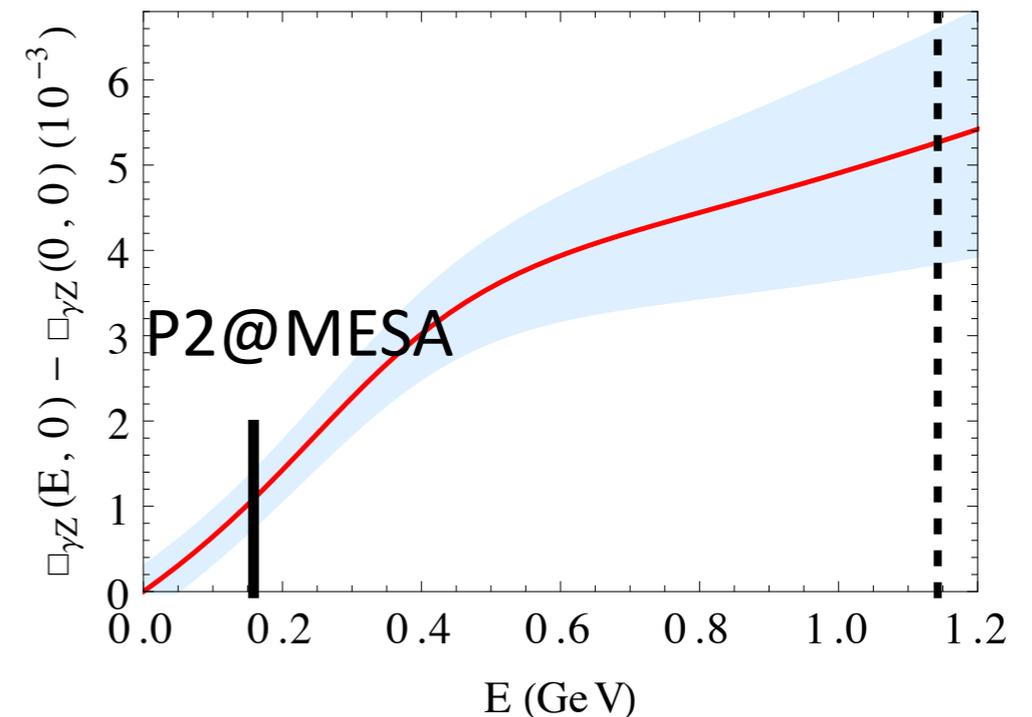
EW boxes: special attention

γZ -box: steep energy dependence, correction & uncertainty rise with energy

MG, Horowitz '09;

MG, Horowitz, Ramsey-Musolf '11;

....



WMA with Qweak@JLab and P2@MESA

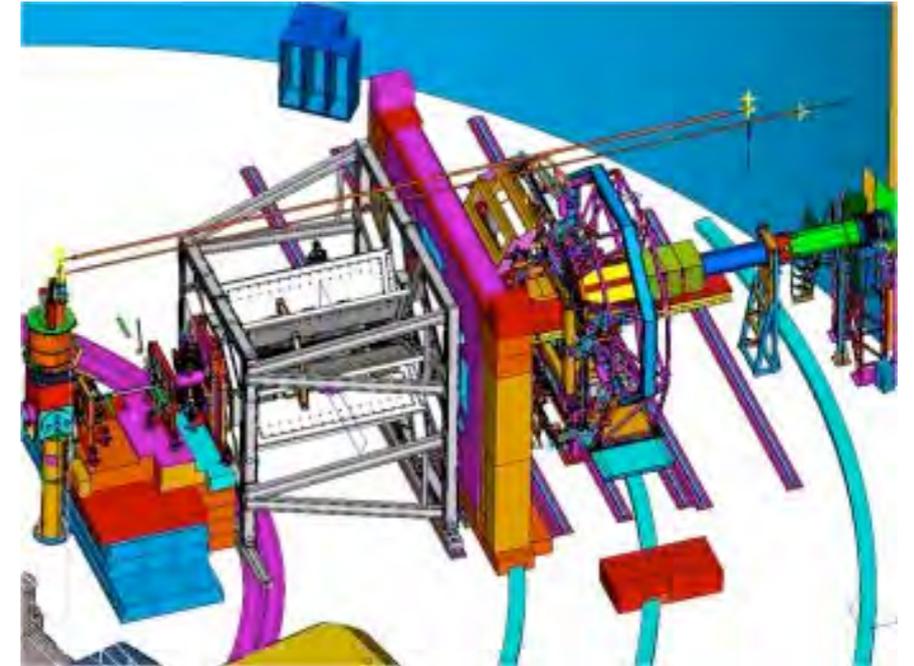
Qweak - published in Nature 2018

1.16 GeV beam; $Q^2=0.03 \text{ GeV}^2$

PV asymmetry to 5% precision;

Combined Theo+Exp. uncertainty - 6%

$\Delta\sin^2\theta_W/\sin^2\theta_W = 0.5\%$

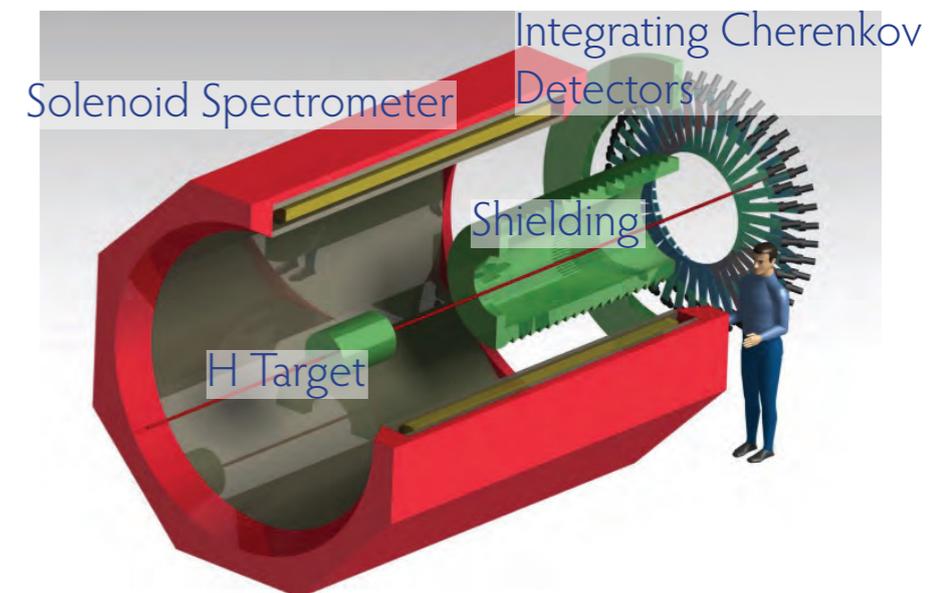


MESA/P2:

155 MeV beam; $Q^2=0.0045 \text{ GeV}^2$

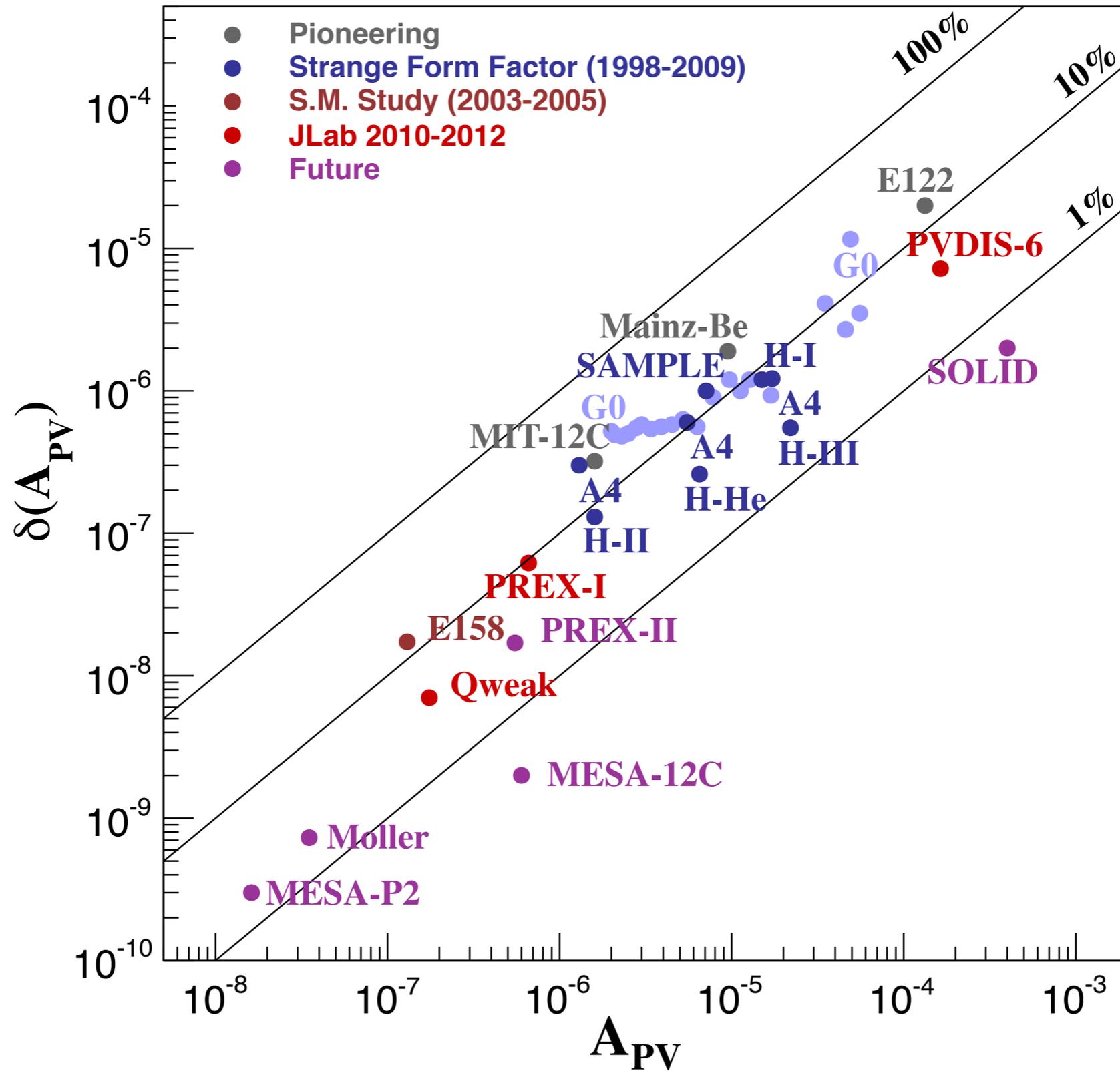
PV asymmetry to 1.5% precision (Exp+Th);

$\Delta\sin^2\theta_W/\sin^2\theta_W = 0.13\%$

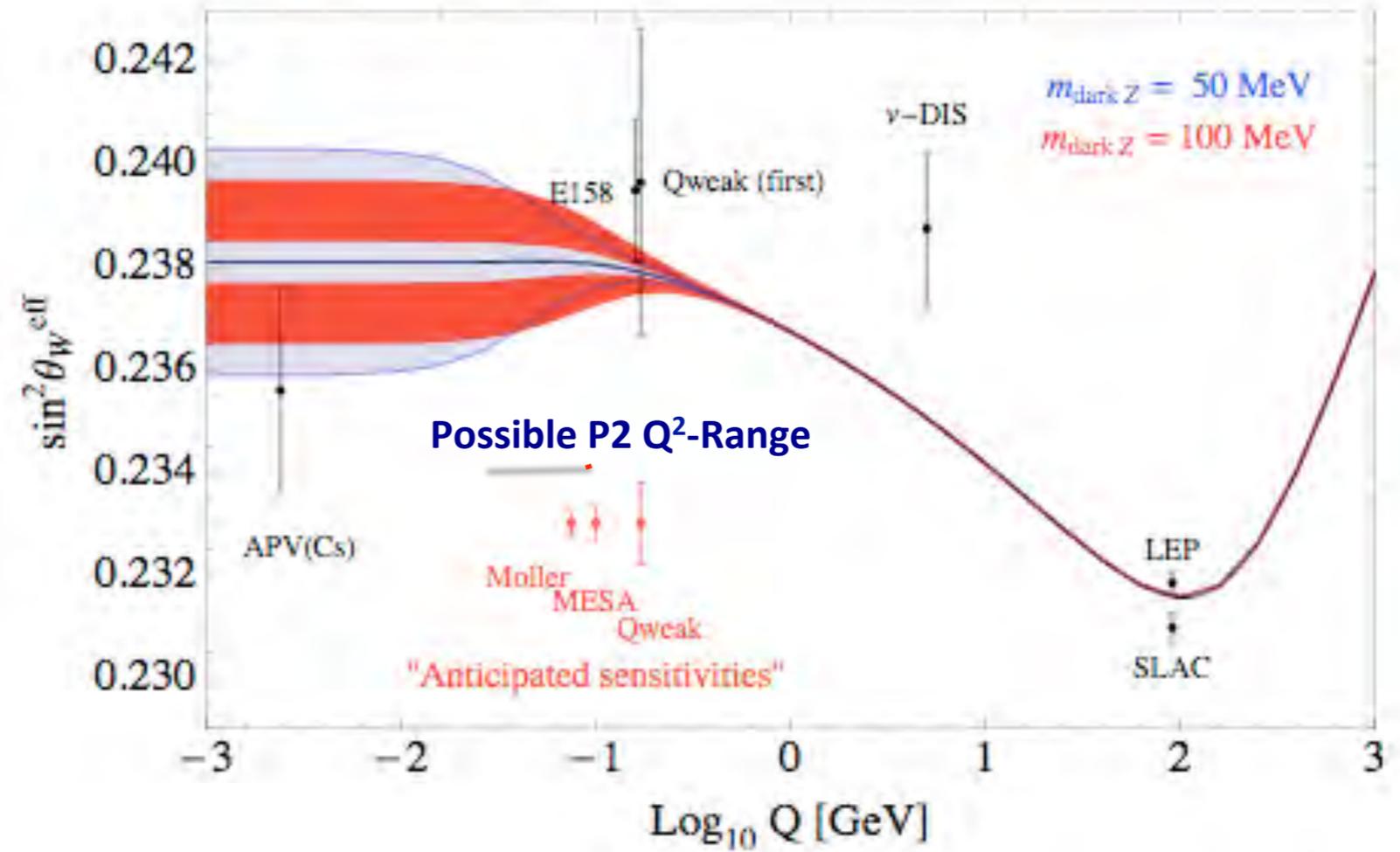


Physics reach up to 49 TeV (comparable to LHC)

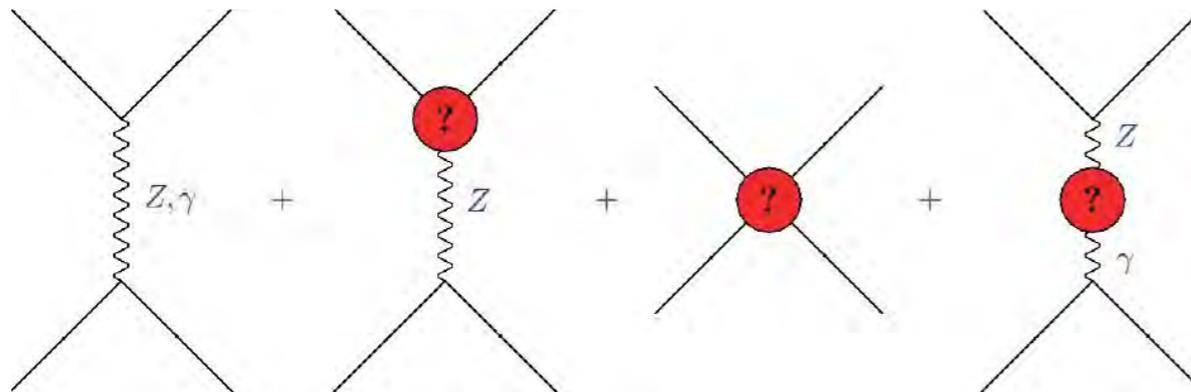
Evolution of PVES experiments



Running $\sin^2 \theta_W$ and Dark Parity Violation



Bill Marciano



Dark Z

Mixing
w. dark Z or γ

New contact
interactions

New fermions

Proposal to enhance the physics reach of P2 experiment

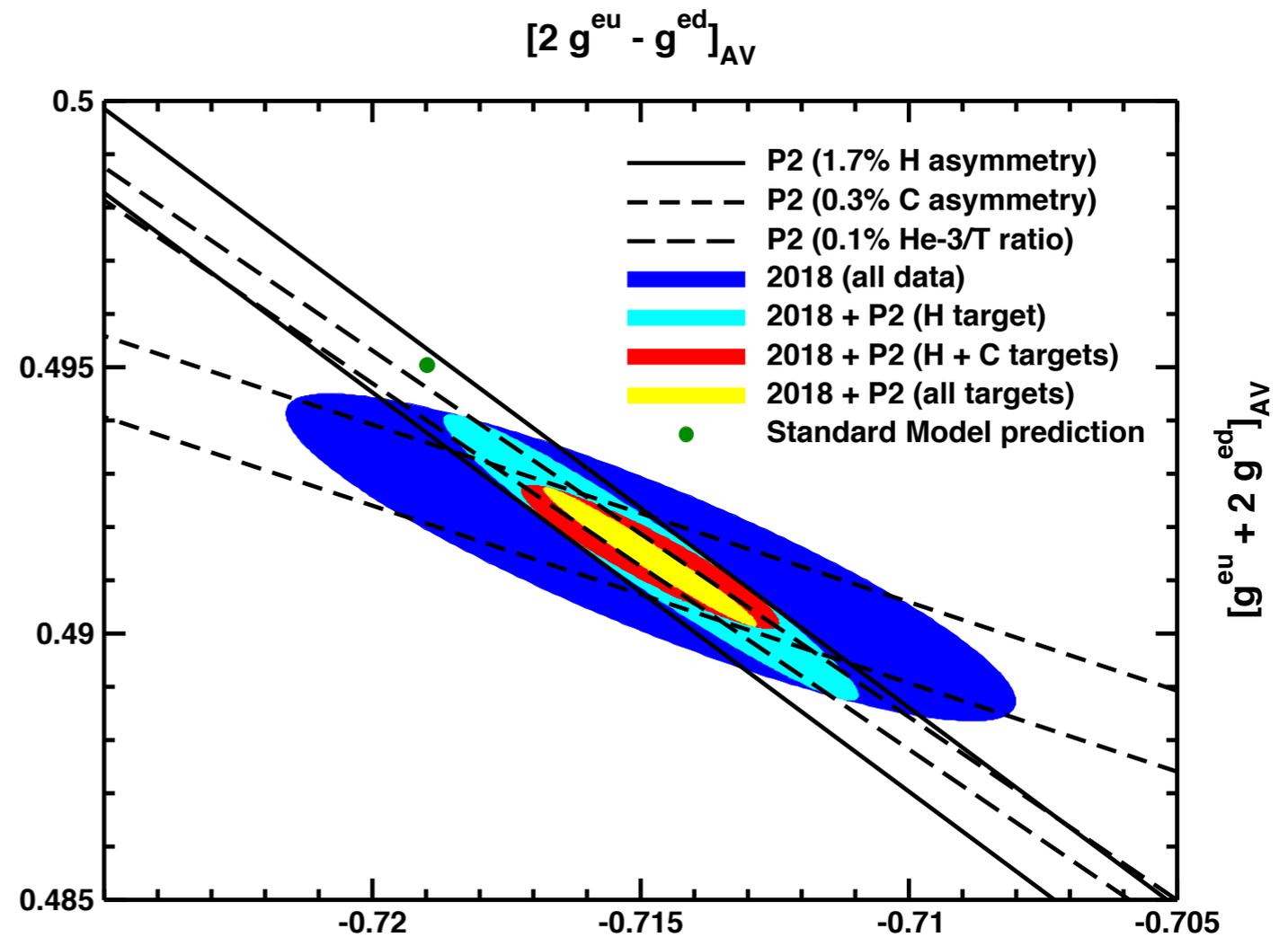
Additional measurement of PV asymmetry on Carbon-12

Why carbon-12? $Q_W^{C-12} = 6(Q_W^p + Q_W^n) = -24 \sin^2 \theta_W$

- A different isospin combination of g_{AV}^{eq}
- Spin-0 target - simpler interpretation (e.g. smaller correction B)
- Larger cross section, larger asymmetry - “easy” experimentally
- Tree-level term large — corrections relatively small
- But: need a sub-% measurement - is everything under control???

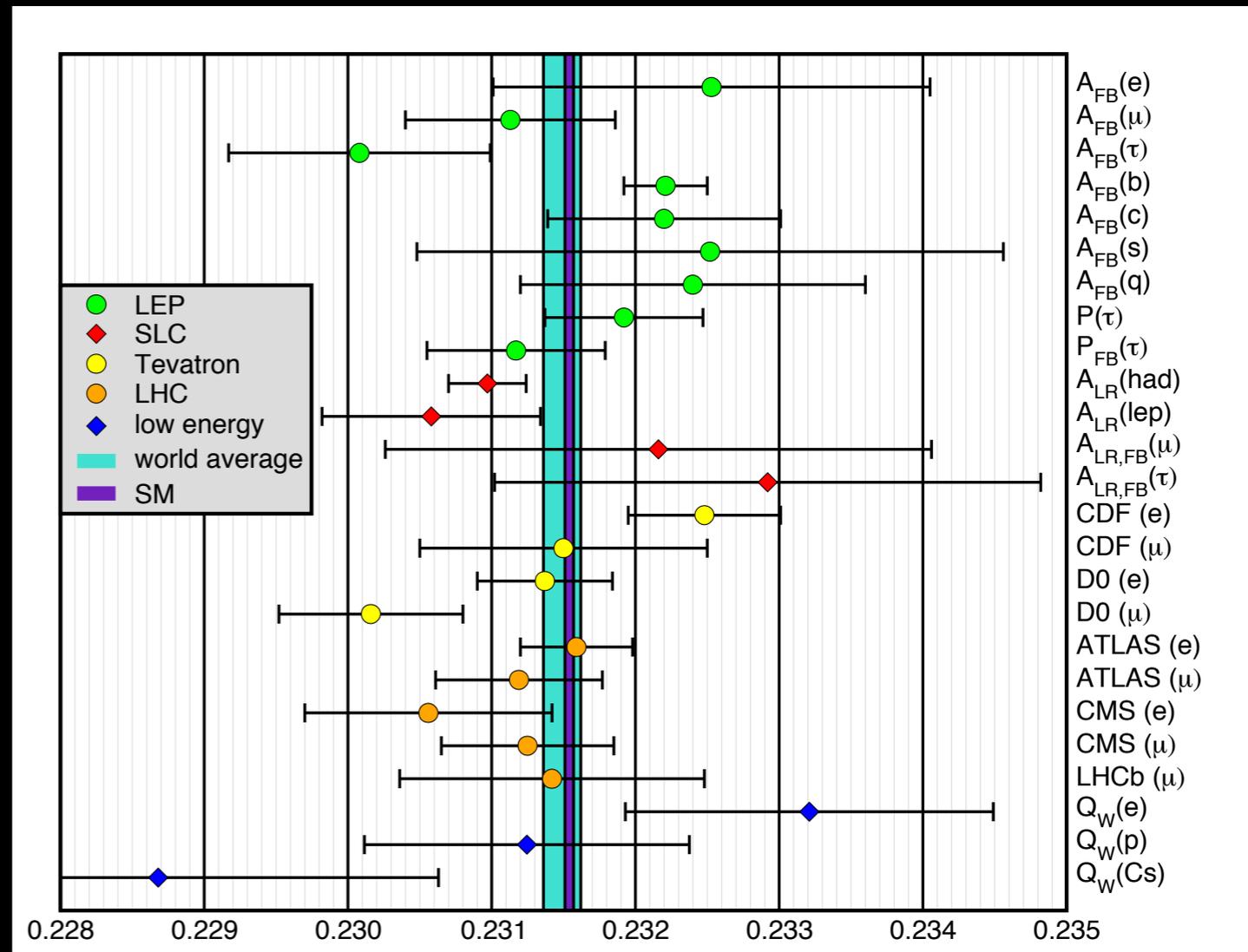
Impact of Qweak and MESA on effective e-q operators:

MESA - C12:
 a 0.3% measurement of A^{PV}
 = 0.3% meas. of $\sin^2\theta_W$
 Access the isoscalar combination
 of g_{AV} 's



SOME CONTEXT

- weak mixing angle
 $\sin^2\theta_W = g'^2/(g^2 + g'^2)$
- many ways to measure in SM (flavor, Q^2 , ...)
- differences may reveal BSM physics (heavy or light Z' , dim-6 operators)



Slide stolen from Jens's talk

Goal of this project and this workshop:

Ensure theory uncertainties for PVES on ^{12}C are below 0.3% level;
Hope that all experimental issues are solvable (and solved);

Experiment:

- Tracking (PV asymmetry $\sim Q^2$ - should be known to below 0.3%)
- Polarimetry (should be known to below 0.3%)
- Electronics
- ...
- **Frank Maas talk on Thursday, April 4 + Krishna Kumar in the crowd**

Theory side:

Overview talks:

Jens (Opening talk **Tuesday, March 19**)

Hubert (e-p scattering from low to high energy, **Wed. April 3**)

QED corrections:

Bremsstrahlung (correct Q^2 reconstruction)

- **Razvan's talk Monday April 1 (today)**

Coulomb distortion and interplay with neutron skin program

- **Sasha's talk Thursday April 4**

Electroweak corrections:

SM running (get the connection $\sin^2\theta_W(0) \longleftrightarrow \sin^2\theta_W(M_Z^2)$ right)

- **Rodolfo's talk Friday April 5** (+ *new ideas on BSM ν interactions*)

Precise heavy quark masses from QCD sum rules for running $\sin^2\theta_W$

- **Pere's talk on Tuesday, March 26**

Precise γZ -box calculation (connection to β decays and ν scattering)

- **Chien Yeah's talk Tuesday April 2**

Theory continued:

Nuclear structure and nuclear astrophysics:

Neutron stars \longleftrightarrow neutron skins from PVES and PV in atoms

- **Jorge P.'s seminar Wednesday March 20**

BSM constraints between low-energy and high energy frontier:

Interplay with colliders and flavor physics

- **Jorge M. C.'s colloquium Thursday March 21**

Connection to other low-energy tests:

Coherent neutrino scattering

- **Eric V.-J.'s talk Thursday April 4**

Neutrinoless double-beta decay

- **KK's colloquium Friday April 5**

Extracurricular activities:

Scalar charges of the nucleon @ JLab

- **Aurore's talk Tuesday April 2**

T-violation with ultra cold neutrons

- **Libertad's talk Thursday April 4**

D-meson physics

- **Genaro's talk Thursday April 4**

All the talks are/will be available on the indico page!

<https://indico.mitp.uni-mainz.de/event/193/>

Further highlights:

- Amazing history, sightseeing, food, hospitality
- Key to successful work: excellent espresso all day long
- We are having great time - many thanks to local organizers!