

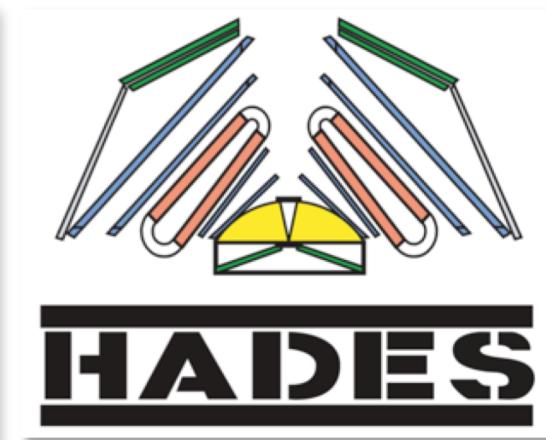
# Recent Results from HADES

---

Christoph Blume  
for the HADES Collaboration



58<sup>th</sup> International Winter Meeting  
on Nuclear Physics  
Bormio, Italy, Jan. 20. – 24., 2020



### Open questions

Origin of hadron masses  
Role of condensates  
QCD-Confinement  
Equation-of-state of dense matter

Super-dense matter in the laboratory

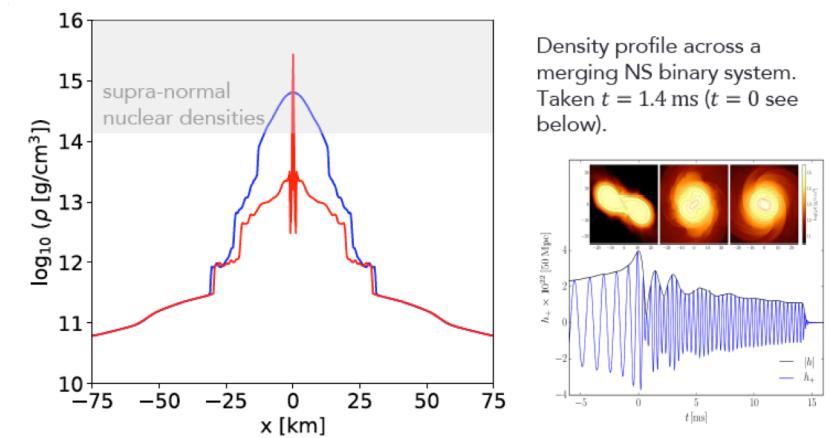
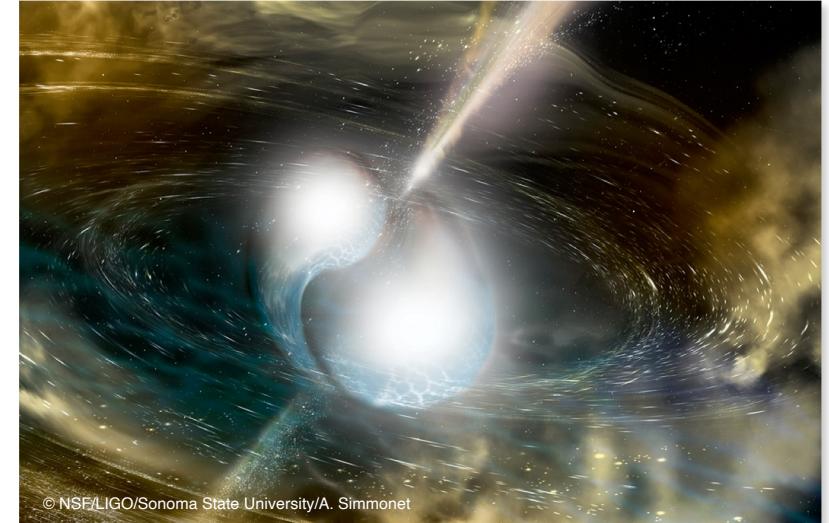
### Neutron Star Merger

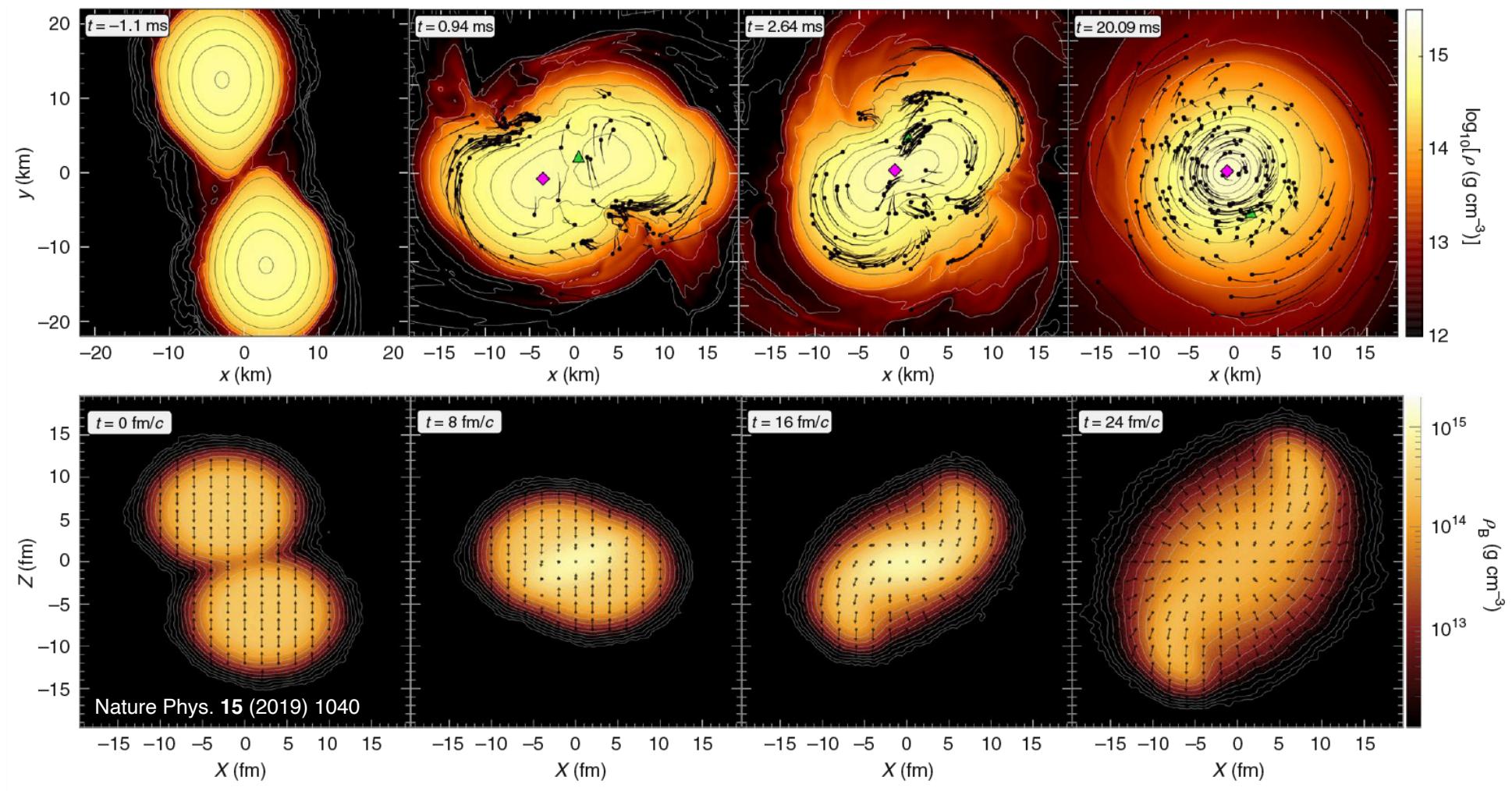
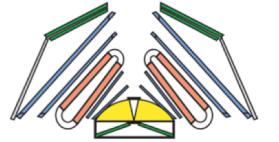
Observation via gravitational waves

**GW170817:** B.P. Abott et al. (LIGO + VIRGO)  
**PRL 119** (2017) 1611001

Sensitivity to equation-of-state

Super-dense matter in the universe

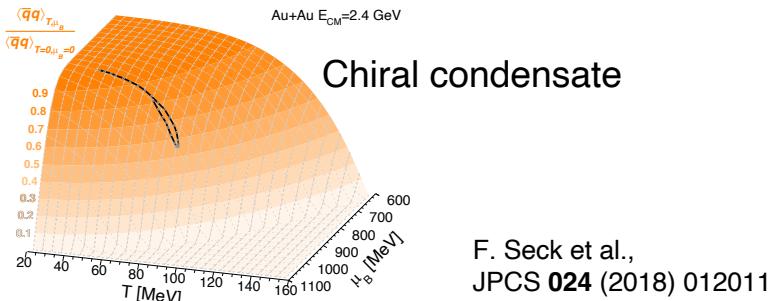




### Heavy-ion collisions

QCD phase diagram in the region of high  $\mu_B$

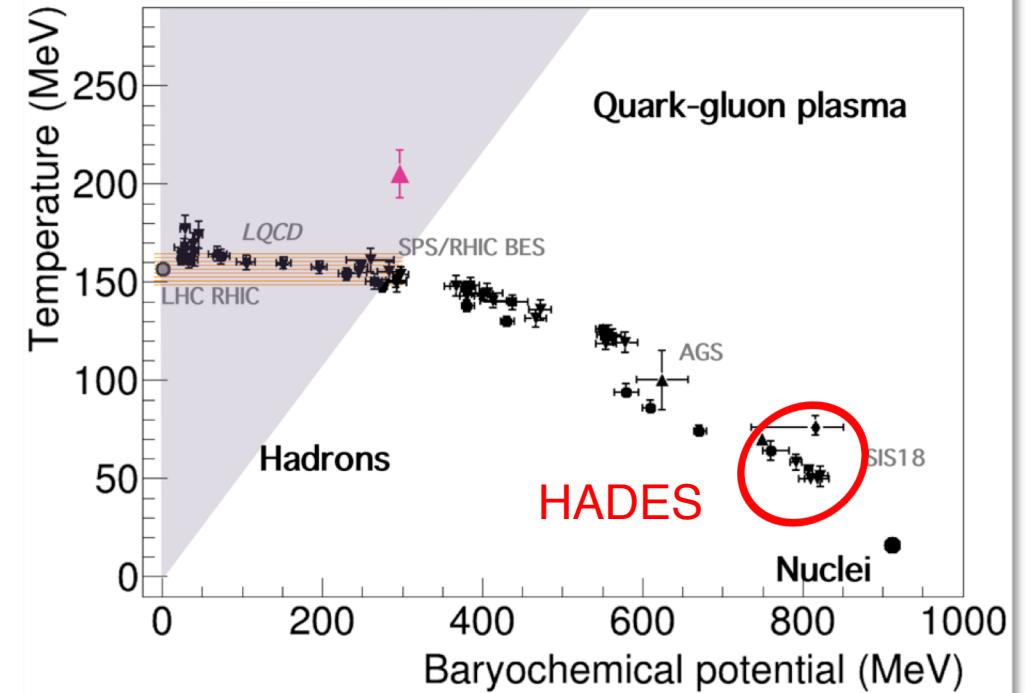
### Rare and penetrating probes



### Bulk properties of dense fireball

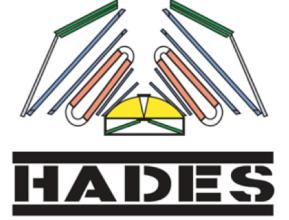
### Pion + proton beams

Properties of baryon resonances  
(vacuum, cold QCD matter)



# HADES

## The Experiment



### Fast detectors

Trigger rates up to 50 kHz

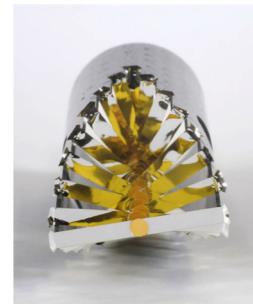
### Large acceptance

Full azimuthal angle

Polar angle coverage:  $18^\circ - 85^\circ$

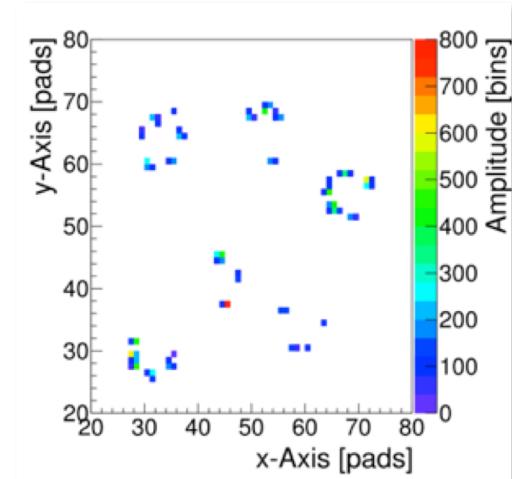
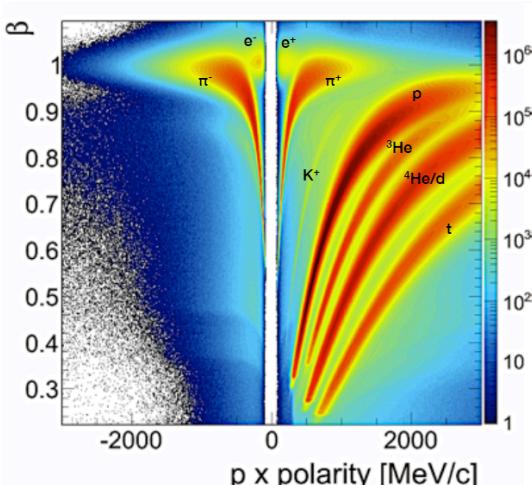
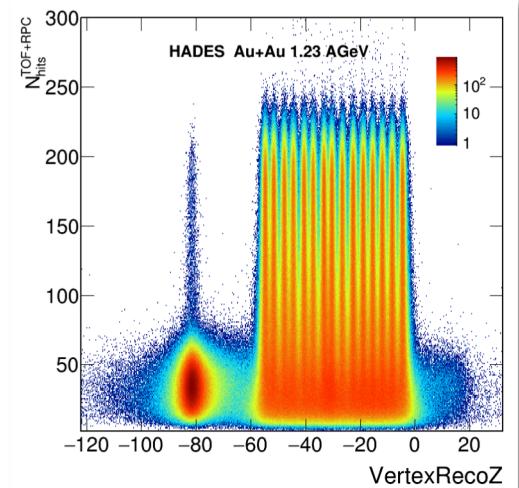
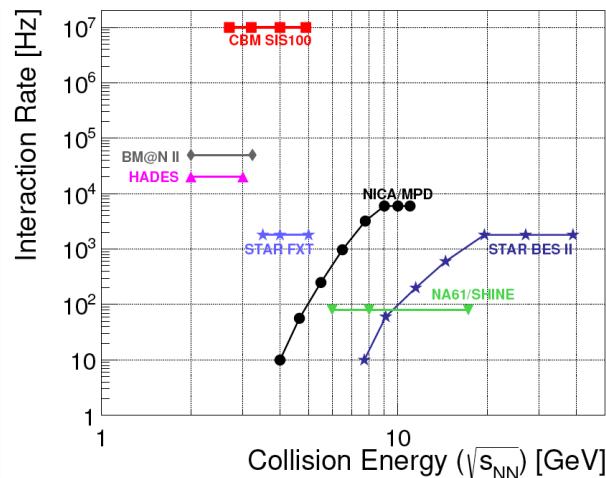
### Segmented Au target

2% Interaction probability  
Very low material budget



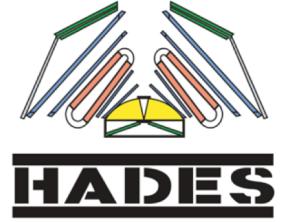
### Particle identification

Time-of-flight (TOF + RPC)  
 $dE/dx$  (MDC + TOF)  
Cherenkov-effect (RICH)



# HADES

## The Experiment



Au+Au at  $\sqrt{s_{NN}} = 2.42$  GeV

Beam:  $1.5 \times 10^6$  ions / s

$7 \times 10^9$  events recorded

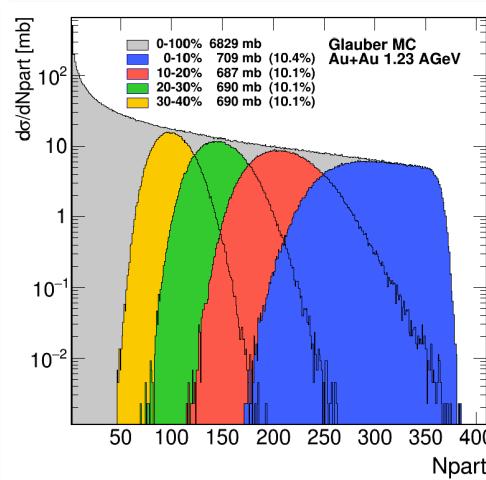
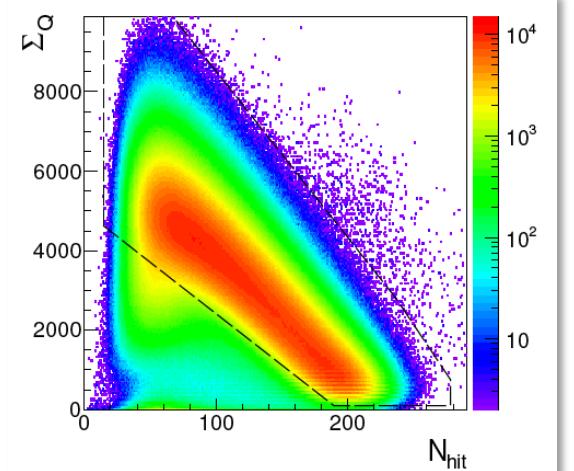
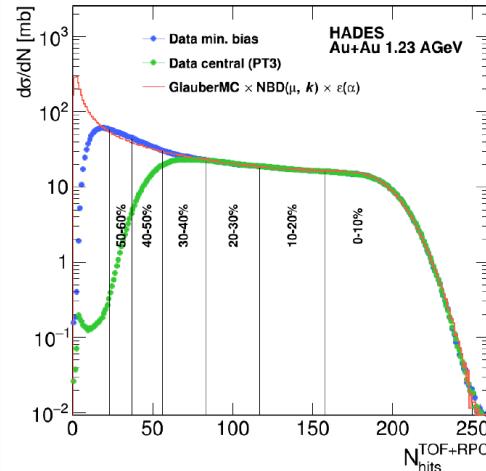
LVL1 trigger: 40% most central

### Centrality selection

$N_{\text{hits}}$  in TOF+RPC  
(standard analysis)

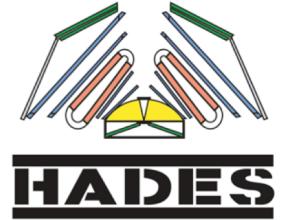
Sum of charges in Forward Wall  $\Sigma_Q$   
(E-by-e fluctuation analysis)

Volume determination  
 $N_{\text{part}}$  from Glauber MC  
Eur. Phys. J. **A54** (2018) 85



# HADES

## The Experiment



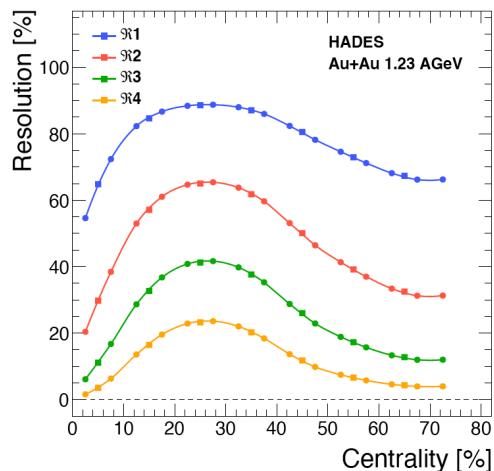
### Event plane Reconstruction

1<sup>st</sup>-Order event plane

Projectile spectators in Forward Wall

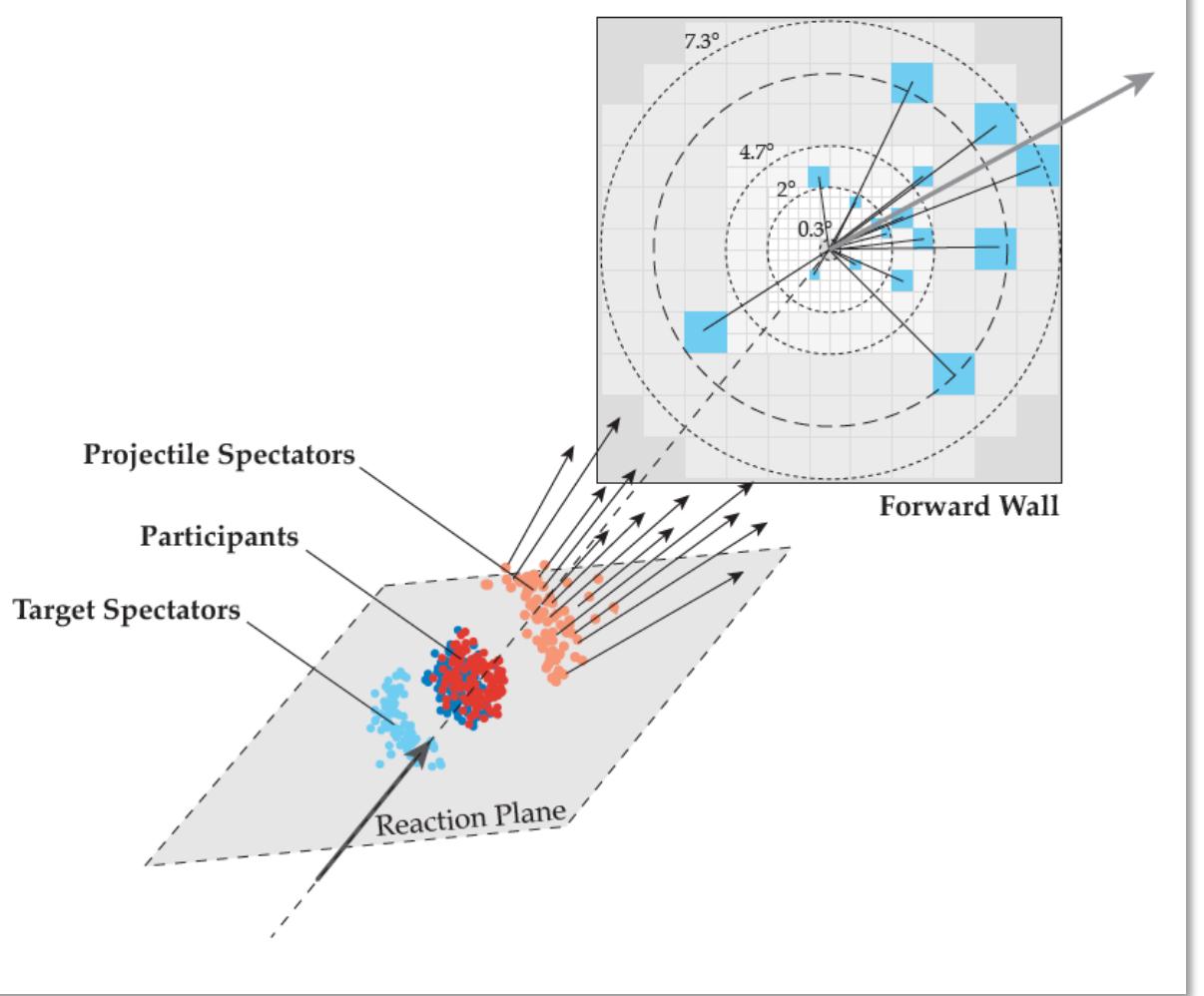
EP-resolution via sub-event method

J.-Y. Ollitrault, arXiv:nucl-ex/9711003



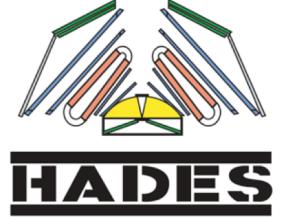
$$v_n = v_n^{obs} / \Re_n$$

$$\Re_n = \langle \cos[n(\Psi_n - \Psi_{RP})] \rangle$$



# HADES

## The Experiment



Au+Au at 1.23A GeV

( $\sqrt{s_{NN}} = 2.42$  GeV)

Particle production

Virtual photons

Collective effects

Femtoscopy

Global  $\Lambda$  polarization

Proton number fluctuations

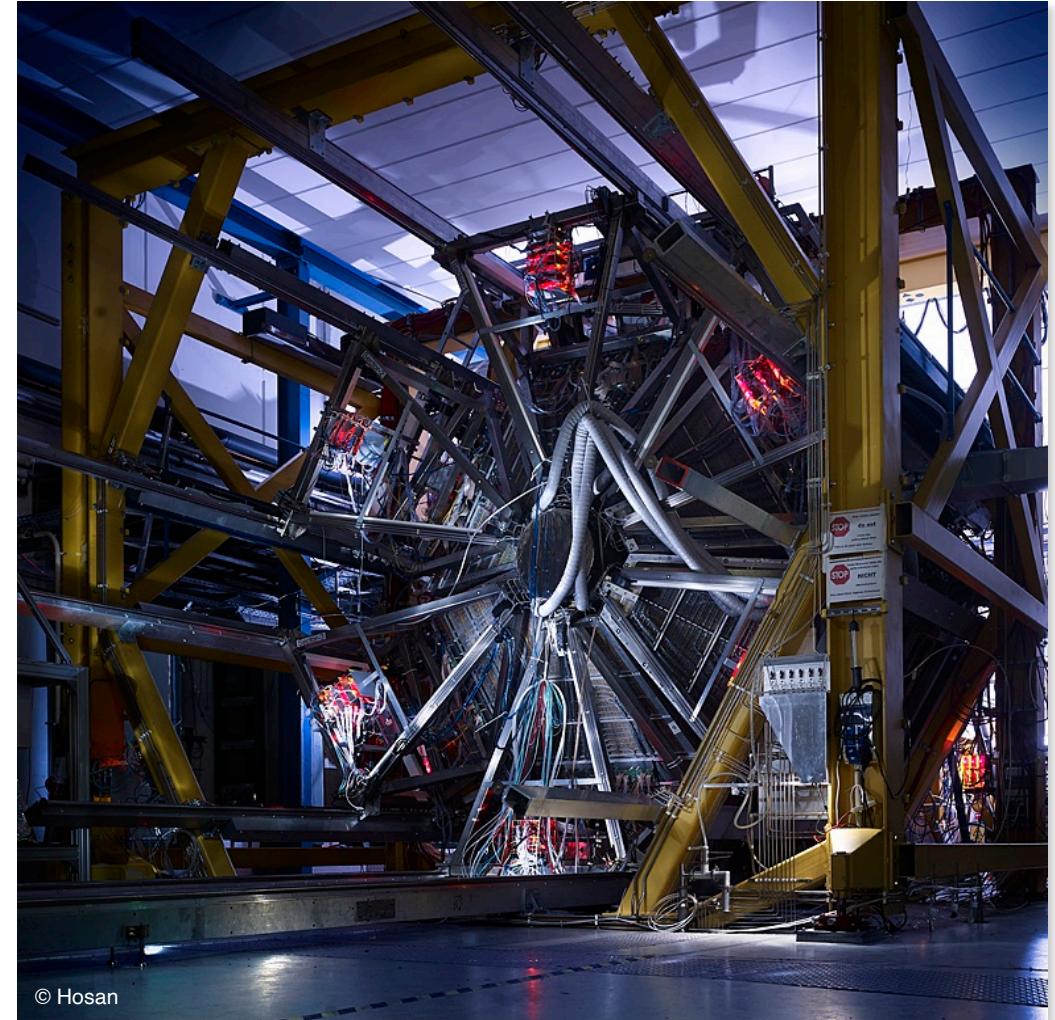
$\pi^- + N$  at 1.7 GeV

Strangeness in cold nuclear matter

Ag+Ag at 1.58A GeV

( $\sqrt{s_{NN}} = 2.6$  GeV)

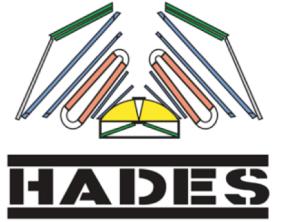
FAIR Phase-0: first look



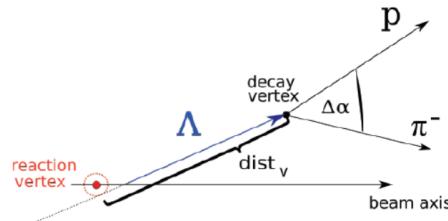
© Hosan

# Particle Production

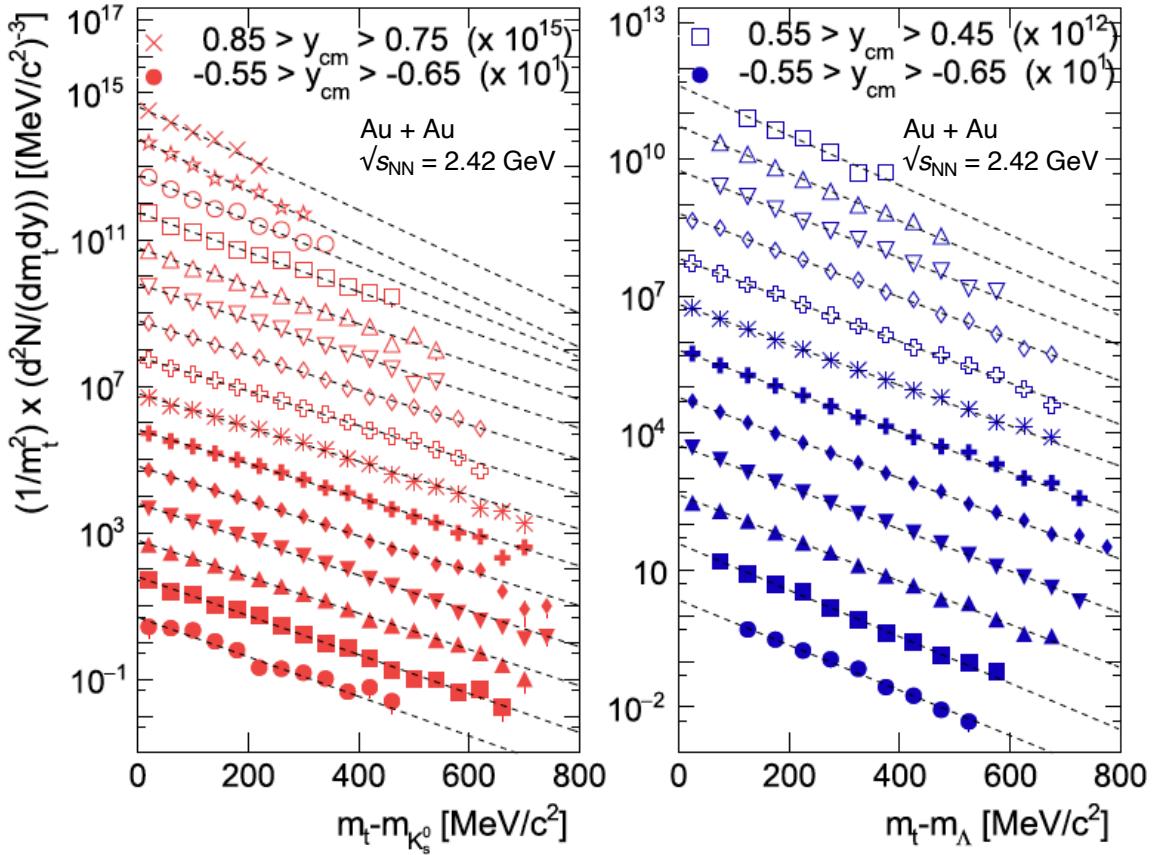
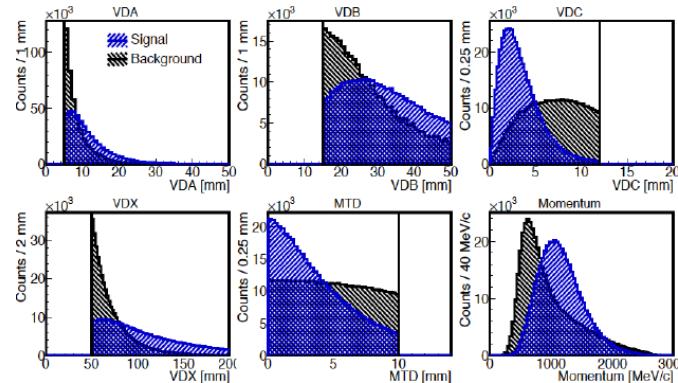
## Strange Particles: $K_s^0$ and $\Lambda$



Weak decay topology  
Large phase space coverage



New developments  
Further improvements with ANN



# Particle Production

## Strange Particles: $K^0_s$ and $\Lambda$

### Sub-threshold production

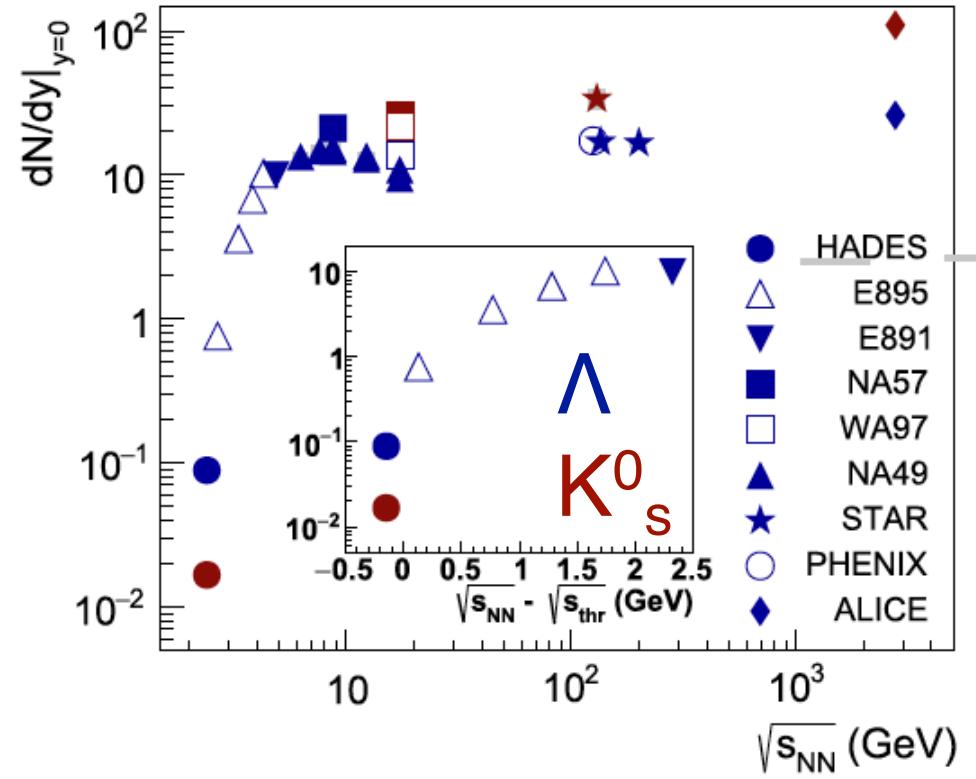
No strangeness production in elementary NN-collisions at  $\sqrt{s_{NN}} = 2.4$  GeV  
 (Lowest thres.: NN  $\rightarrow$  N $\Lambda$ K $^+$   $\sqrt{s_{NN}} = 2.55$  GeV)

### Sensitive to system properties

Energy transfer via strangeness exchange  
 $(\pi Y \rightarrow N K^-)$   
 Coupling of K $^-$  to baryons  
 Potentials: K $^+$ N repulsive, K-N attractive

### Steep excitation function

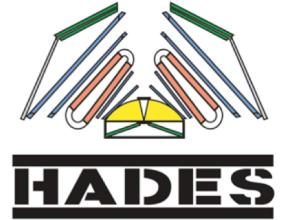
Phys. Lett. **B793** (2019) 457



Data compilation: CB and C. Markert,  
 Prog. Part. Nucl. Phys. **66** (2011) 834

# Particle Production

## Strange Particles: System Size Dependence



### Sub-threshold production

No strangeness production in elementary NN-collisions at  $\sqrt{s_{NN}} = 2.4$  GeV  
(Lowest thres.:  $NN \rightarrow N\Lambda K^+$   $\sqrt{s_{NN}} = 2.55$  GeV)

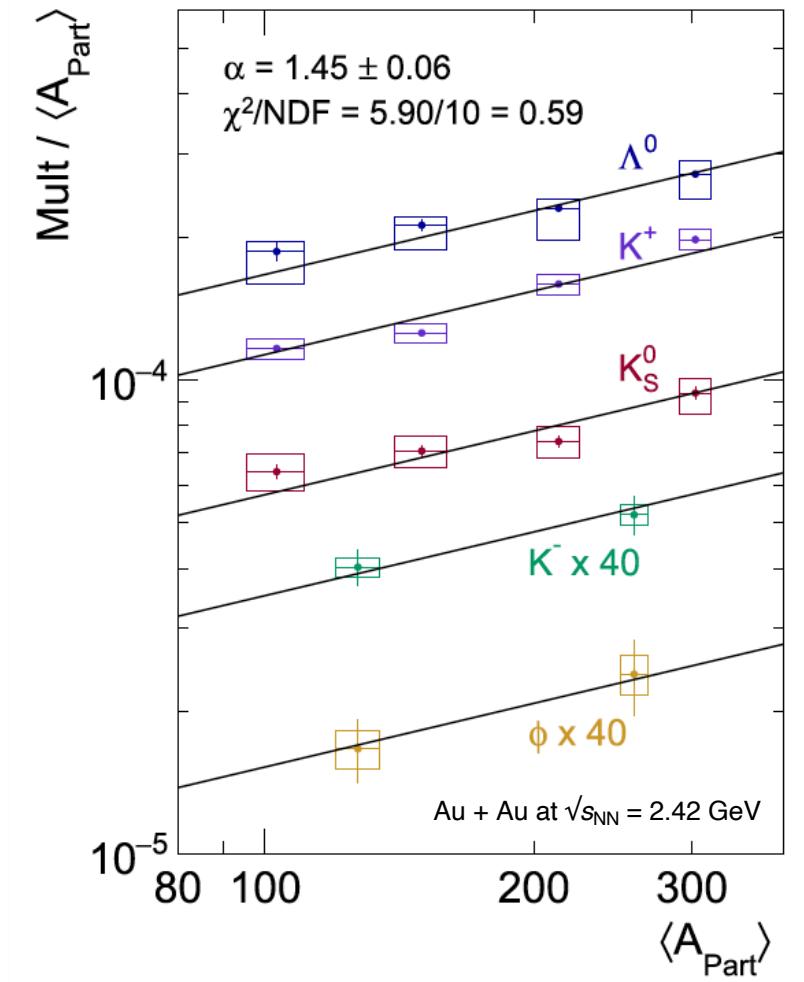
### System size dependence

Parameterization:  $\langle Mult \rangle \sim \langle A_{part} \rangle^\alpha$   
Multiplicities rise stronger than linear ( $\alpha > 1$ )

### Universal $A_{part}$ dependence

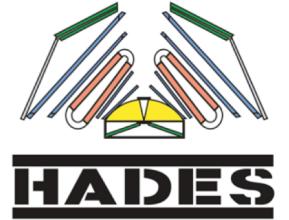
Expected hierarchy due to different thresholds not seen

Phys. Lett. **B793** (2019) 457



# Particle Production

## Strangeness in Cold Matter: Pion-Nucleus

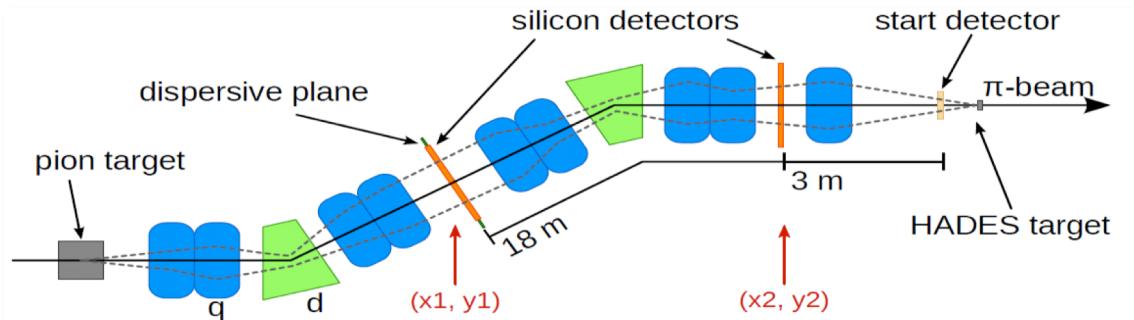


### Pion beam facility

Secondary  $\pi^-$  beam at  $1.7 \text{ GeV}/c$

Intensity:  $\sim 3 \cdot 10^5 \pi^-/\text{s}$

CERBEROS tracking system:  $0.3\%(\sigma)$  res.  
Eur. Phys. J. **A53** (2017) 188

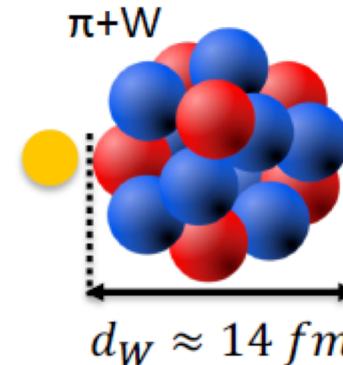
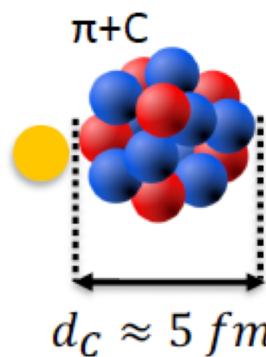


### Target nuclei: C and W

Comparison of  $K^+$ ,  $K^-$  and  $\phi$  production  
Phys. Rev. Lett. **123** (2019) 022002

$\pi$ -induced production close to surface  
 $\rightarrow$  Long path lengths inside nuclear matter

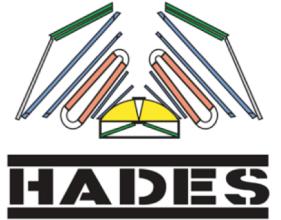
Study of absorption in cold nuclear matter



Mean free path:  $\lambda_\pi \approx 1.5 \text{ fm}$

# Particle Production

## Strangeness in Cold Matter: Pion-Nucleus



### Suppression of $K^-$ relative to $K^+$

Average value (in HADES acceptance):

$$(K^-/K^+)_{\text{W}}/(K^-/K^+)_{\text{C}} = 0.319 \pm 0.009(\text{stat})^{+0.014}_{-0.012}(\text{syst})$$

### Suppression of $\phi$

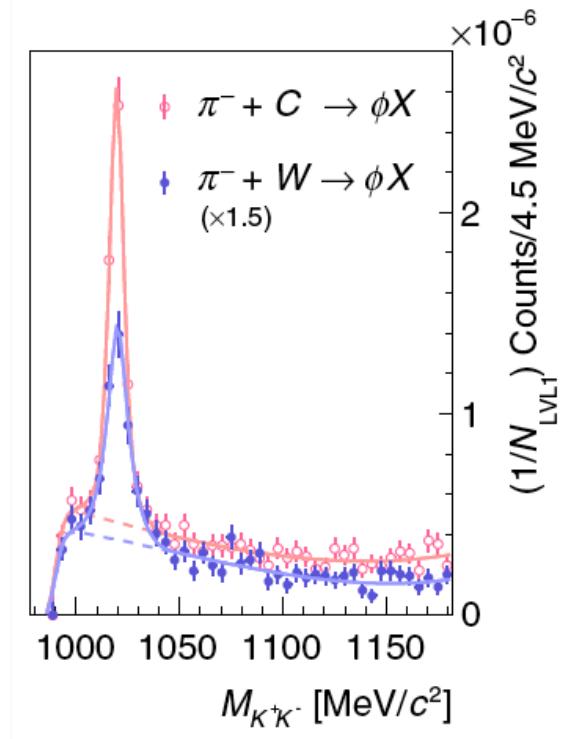
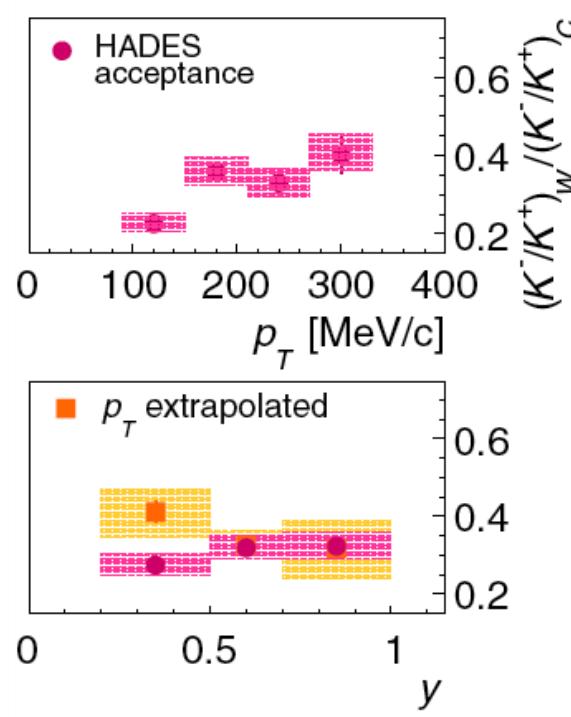
$\phi/K^-$  ratios comparable in  $\pi+C$  and  $\pi+W$ :

$$(\phi/K^-)_{\text{C}} = 0.55 \pm 0.04(\text{stat})^{+0.06}_{-0.07}(\text{sys})$$

$$(\phi/K^-)_{\text{W}} = 0.63 \pm 0.06(\text{stat}) \pm 0.11(\text{sys})$$

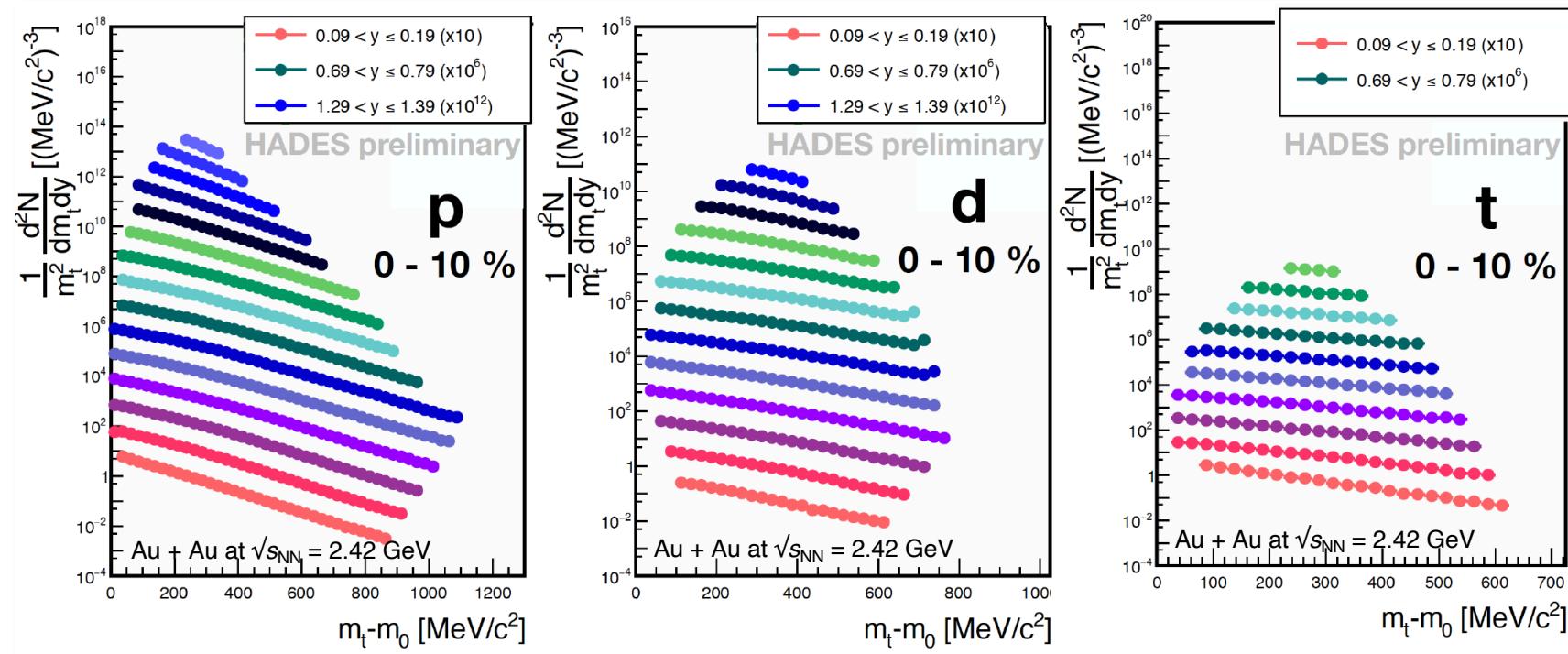
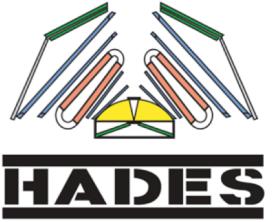
→ Similar suppression for  $\phi$  than for  $K^-$   
In line with measured transparency ratio:

$$\begin{aligned} T &= [(12/184)(\Delta\sigma_{\text{W}}^\phi/\Delta\sigma_{\text{C}}^\phi)] \\ &= 0.18 \pm 0.02 \pm 0.01^{+0.04}_{-0.03} \end{aligned}$$



# Particle Production

## Protons, Deuterons and Tritons

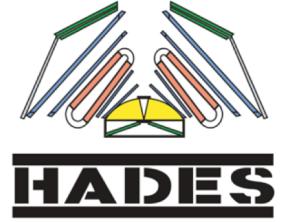


New high statistics data on protons and light nuclei  
Multi-differential analysis

M. Szala,  
SQM19

# Particle Production

## Transverse Mass Spectra → Kinetic Freeze-Out



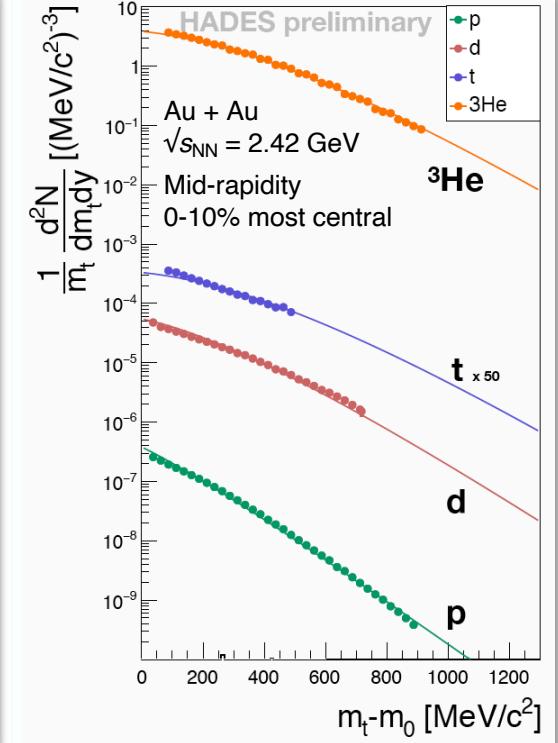
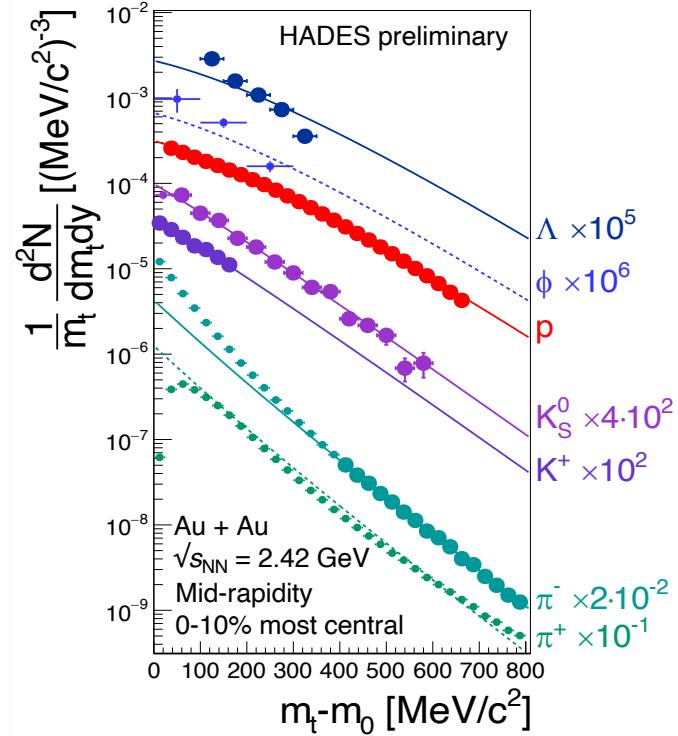
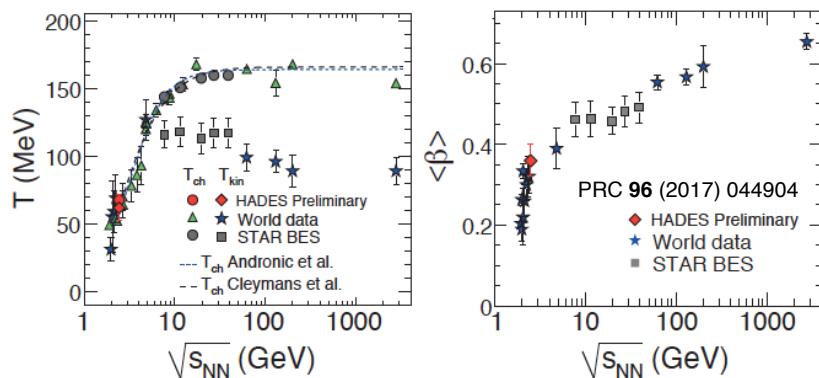
### Fit with Blast-Wave model

Phys. Rev. **C48** (1993) 2462

$\pi, p, K, \Lambda, \phi$ :  $T_{\text{kin}} = 62 \pm 10 \text{ MeV}$   
 $\langle \beta_T \rangle = 0.36 \pm 0.04$

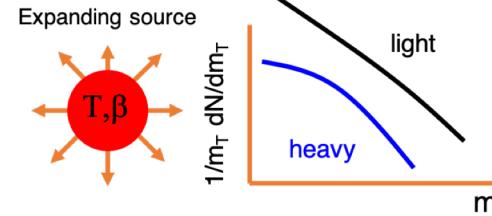
Light nuclei::  $T_{\text{kin}} = 71 \pm 8 \text{ MeV}$   
 $\langle \beta_T \rangle = 0.30 \pm 0.02$

### World systematics



### Blast-Wave model

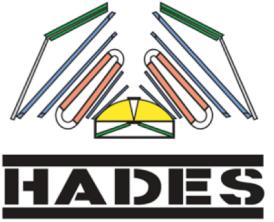
⇒ Kin. freeze-out params.:  
Temperature:  $T_{\text{kin}}$   
Average transverse  
expansion velocity:  $\langle \beta_T \rangle$



M. Szala,  
ECT\* Workshop

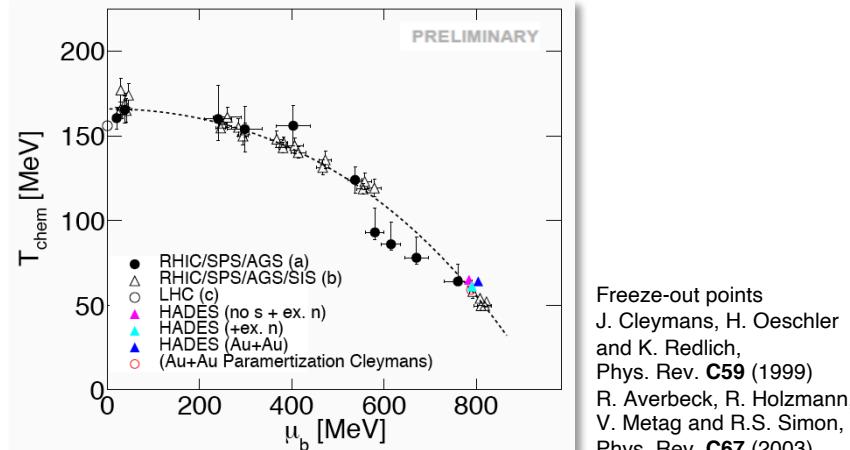
# Particle Production

## Particle Yields → Chemical Freeze-Out



### Fit with THERMAL-FIST

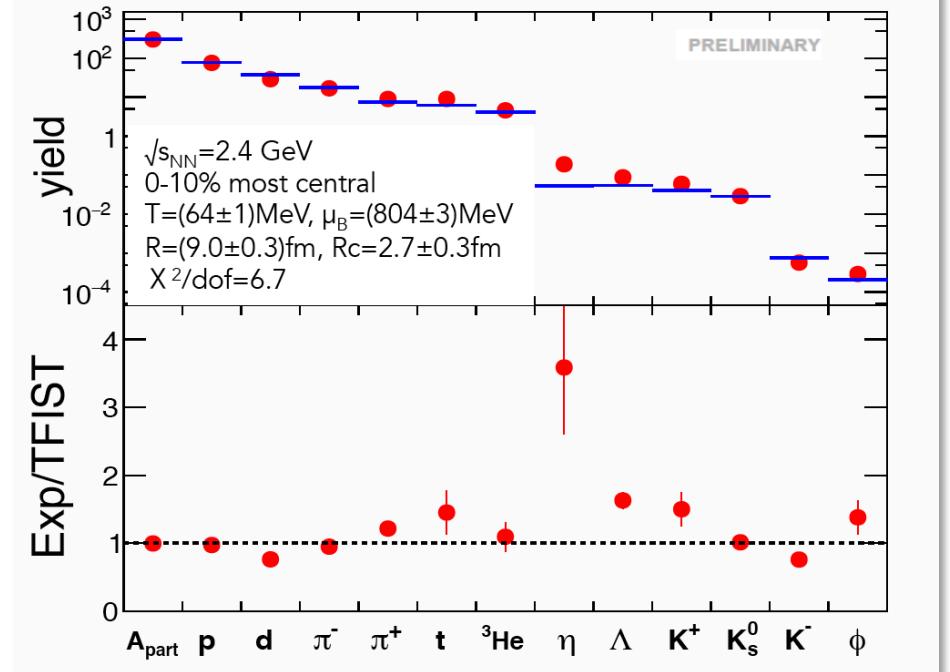
Freeze-Out parameters ( $T_{\text{chem}}$ ,  $\mu_B$ ) follow universal chemical freeze-out curve



Relatively high  $\chi^2$  for full hadron spectrum

Inclusion of excited nuclei states helps  
Add. exclusion of strangeness  $\Rightarrow$  small  $\chi^2$   
(see backup slides)

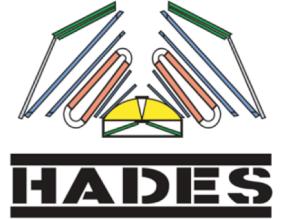
Thermal Fist: V. Vovchenko H. Stoecker, Comput. Phys. Commun. 244 (2019) 295.



M. Lorenz,  
EMMI-Workshop on Anti-matter,  
hyper-matter and exotica

# Virtual Photons

## Di-Electron Spectrum



1<sup>st</sup> heavy system at low  $\sqrt{s_{NN}}$

Two analysis strategies

Efficiency corrected

Nature Phys. **15** (2019) 1040

Large excess radiation

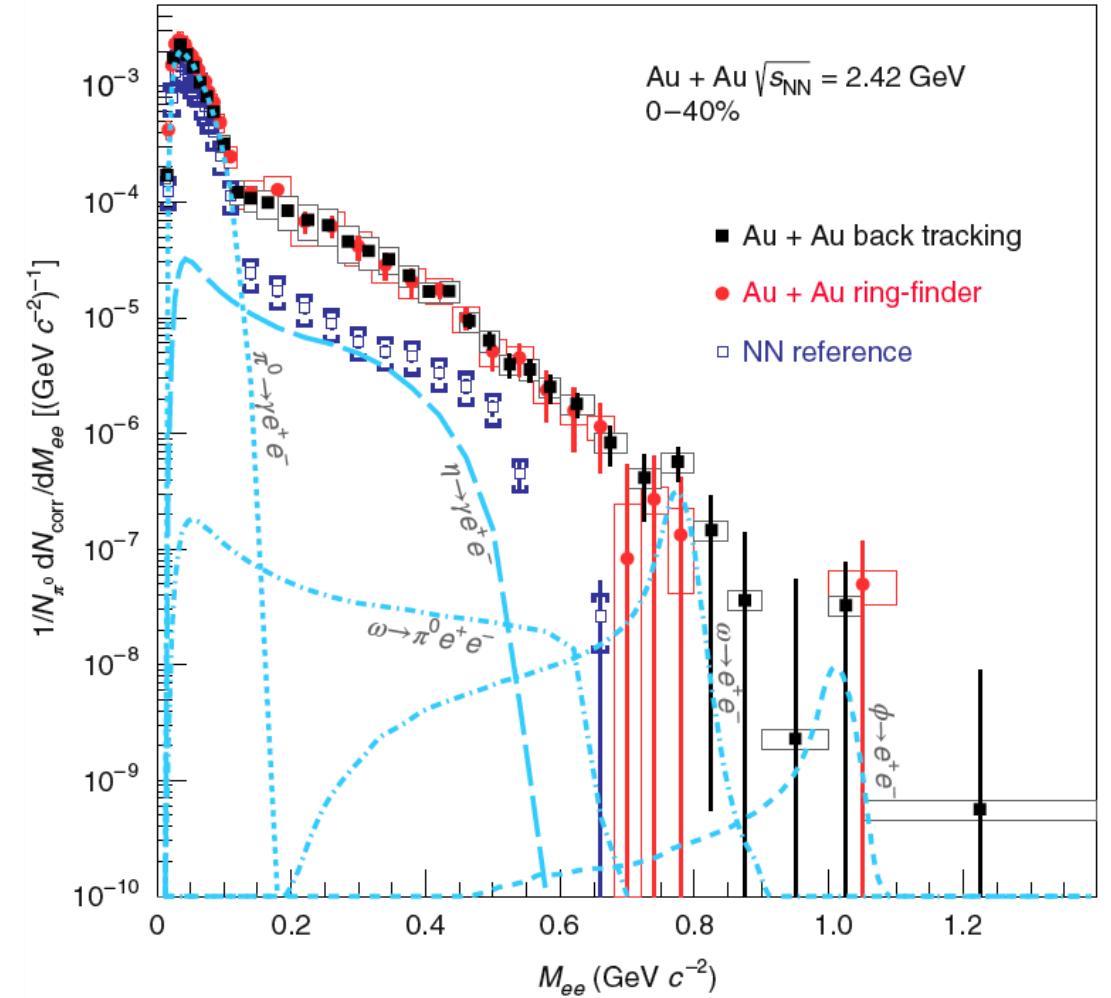
$0.15 < M_{ee} < 0.7 \text{ GeV}/c^2$

Relative to meson decays and NN reference

$$\frac{dN_{\text{ref}}^{\text{NN}}}{dM_{ee}} = \left( 0.54 \frac{dN^{pp}}{dM_{ee}} + 0.46 \frac{dN^{np}}{dM_{ee}} \right) \langle A_{\text{part}} \rangle$$

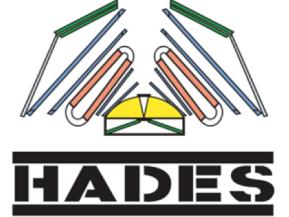
Subtraction of reference

Extraction of spectrum of excess radiation



# Virtual Photons

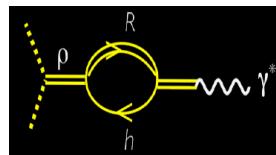
## Di-Electron Spectrum



### Spectrum of excess radiation

Not described with vacuum  $\rho$

Broadening of  $\rho$  due to  $\rho$ -baryon scattering



### Thermal radiation of medium

Thermal radiation folded with fireball evolution from coarse-grained transport

In-medium propagator: R. Rapp and H. van Hees, PLB **753** (2016) 586

$\rho - a_1$  chiral mixing: R. Rapp and P. Hohler, PLB **731** (2014) 103

CG FRA: S. Endres et al., PRC **92** (2015) 014911

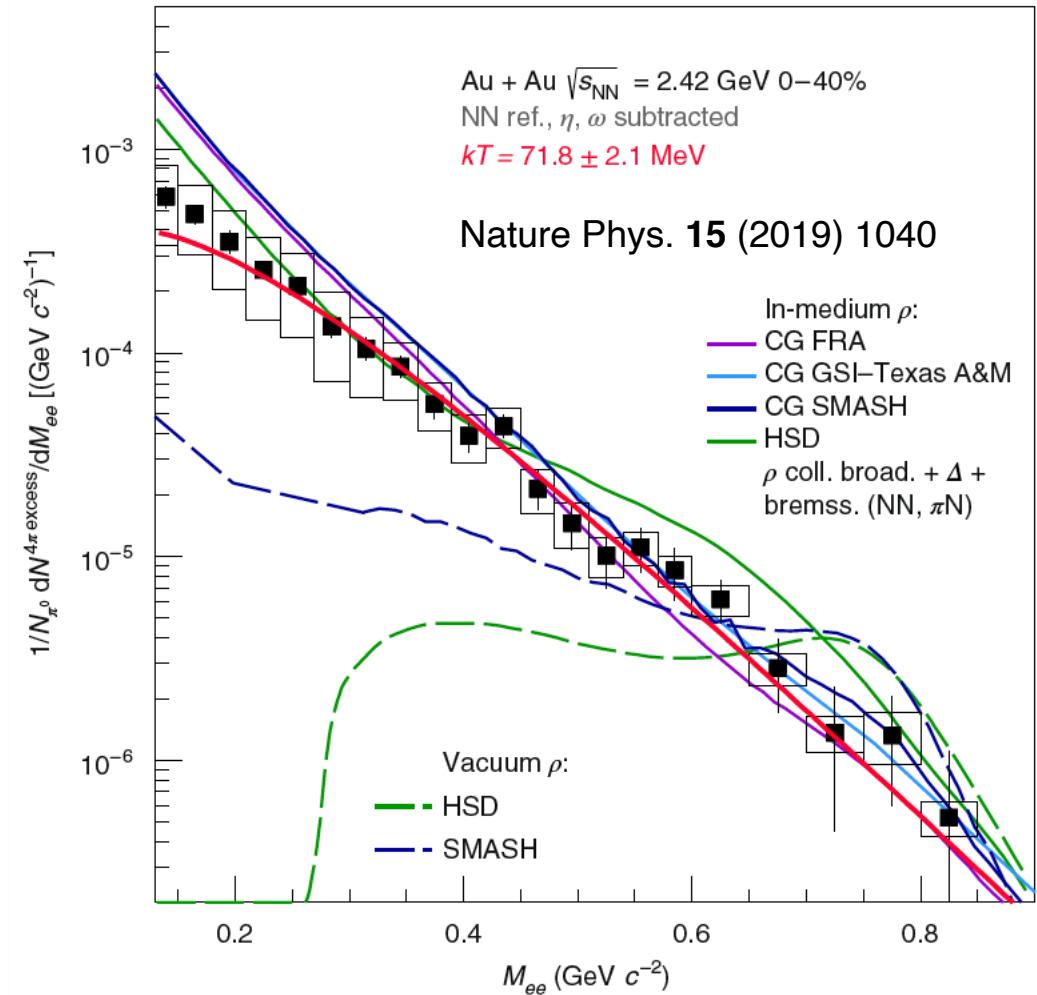
CG GSI-TAMU: T. Galatyuk, F. Seck et al., EPJA **52** (2016) 131

CG SMASH: J. Staudenmaier et al., PRC **98** (2018) 054908.

HSD transport model: E. Bratkovskaya et al., PRC **87** (2013) 064907.

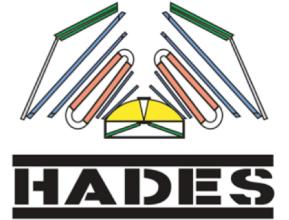
Exponential spectrum  $\rightarrow$  black body rad.

Fit result:  $T_{\text{rad}} = 71.8 \pm 2.1$  MeV



# Virtual Photons

## Di-Electron Spectrum



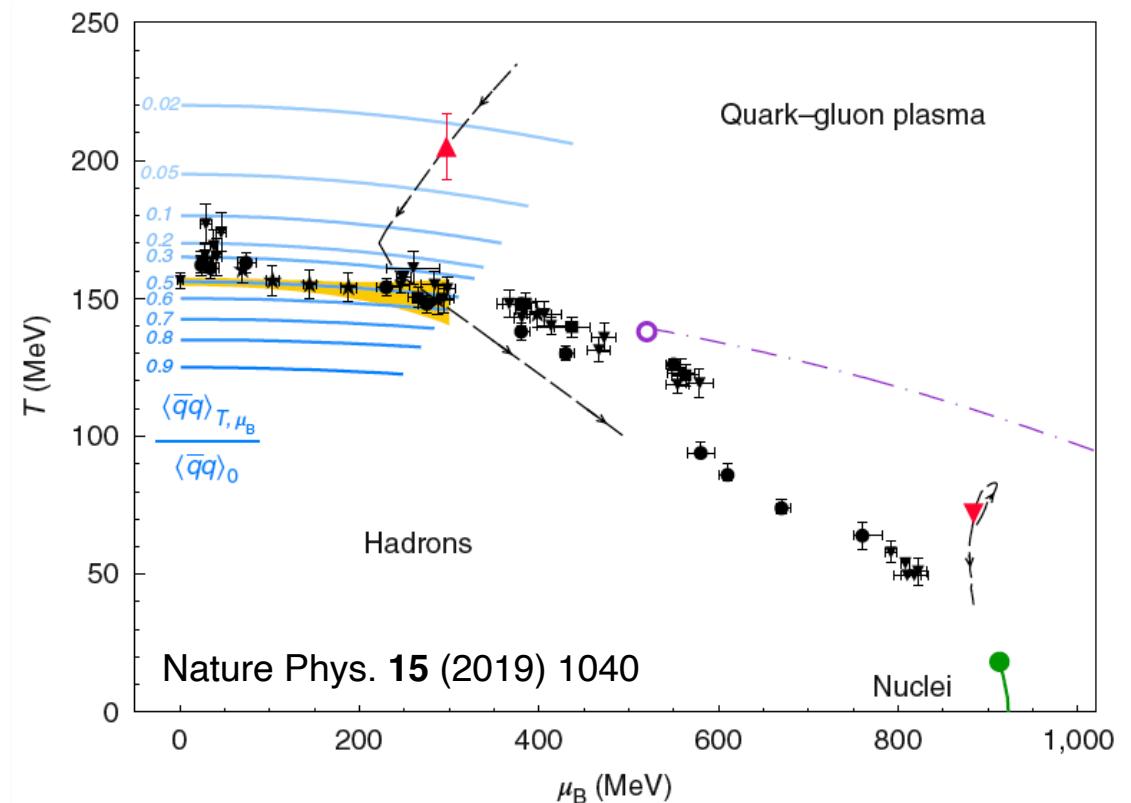
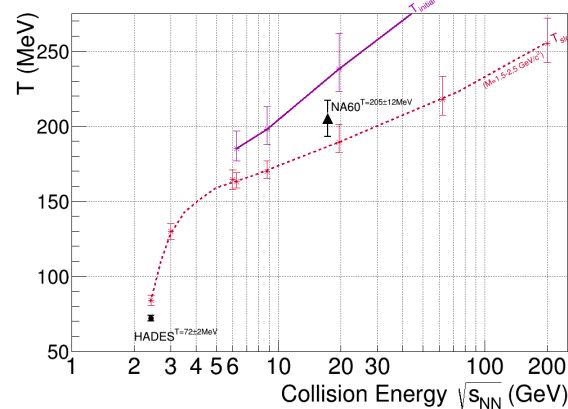
### Thermal radiation of medium

Exponential spectrum  $\rightarrow$  black body rad.  
Fit result:  $T_{\text{rad}} = 71.8 \pm 2.1 \text{ MeV}$

### QCD phase diagram

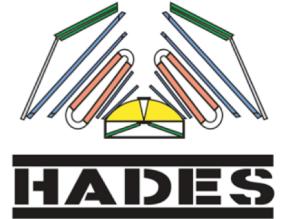
3<sup>rd</sup> data point from HADES  
 $T_{\text{rad}} > T_{\text{chem}} \approx T_{\text{kin}}$

### SPS measurement by NA60



# Collective Effects

## Principle



### Emission relative to event plane

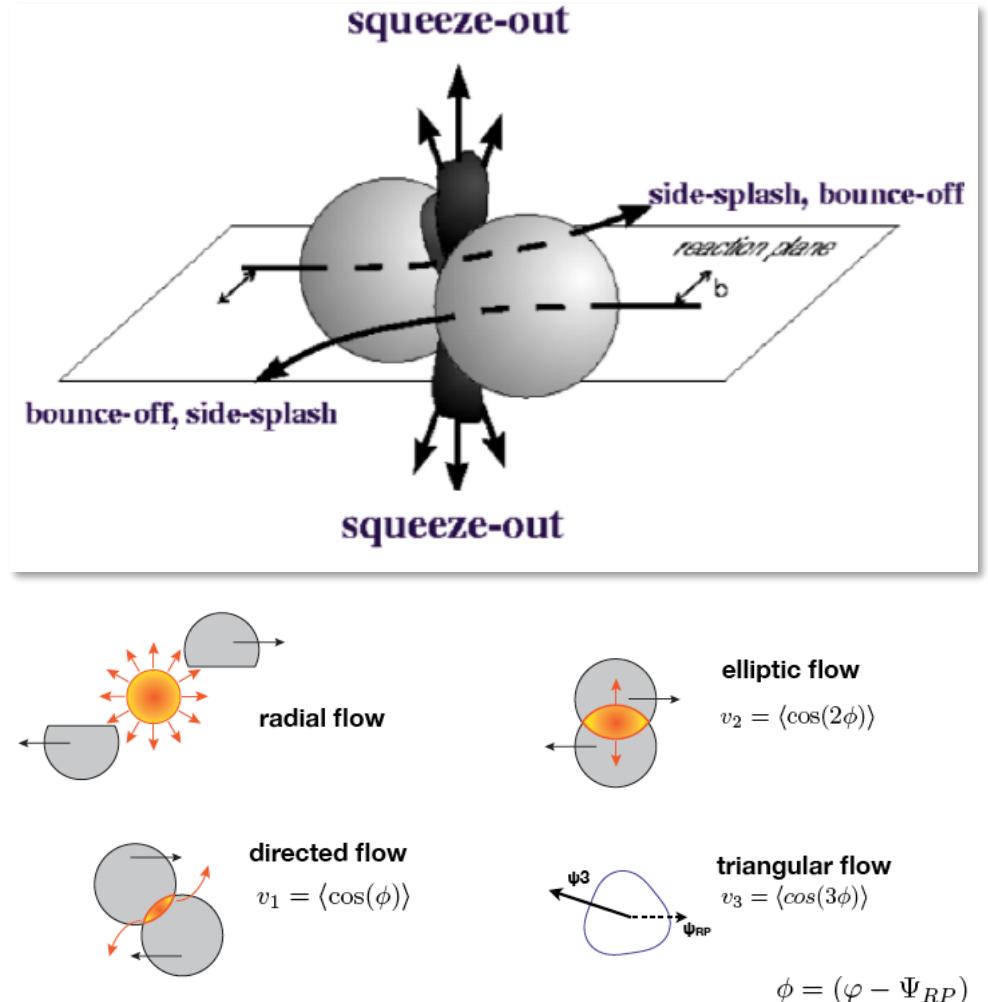
Interactions in medium  $\Rightarrow$  different pressure gradients in different directions

Access to medium properties, e.g. viscosity  
Equation-of-state

### Fourier-Decomposition

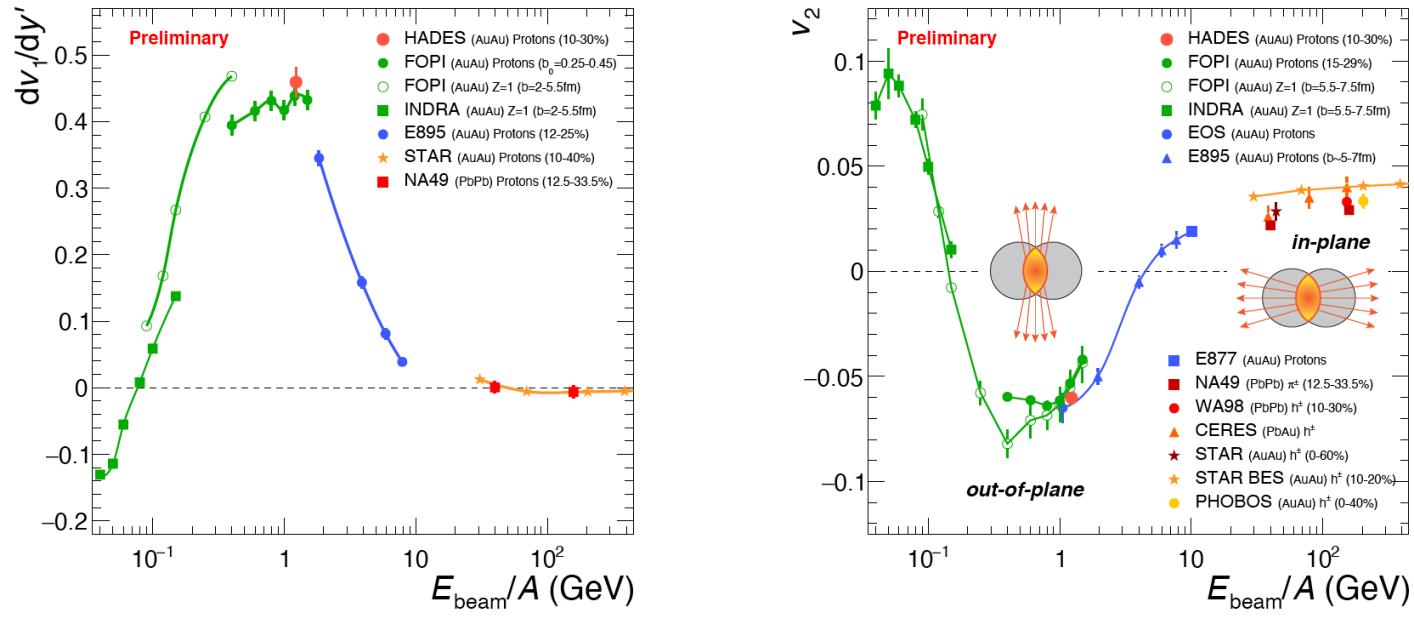
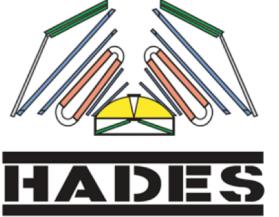
Extraction of moments  $v_n$

$$E \frac{d^3 N}{d^3 p} = \frac{1}{2\pi} \frac{d^2 N}{p_t dp_t dy} \left( 1 + 2 \sum_{n=1}^{\infty} v_n \cos[n(\varphi - \Psi_{RP})] \right)$$



# Collective Effects

## Energy Dependence



### Comparison to world data

Good agreement of integrated  $dv_1/dy$  (directed flow) and  $v_2$  (elliptic flow)

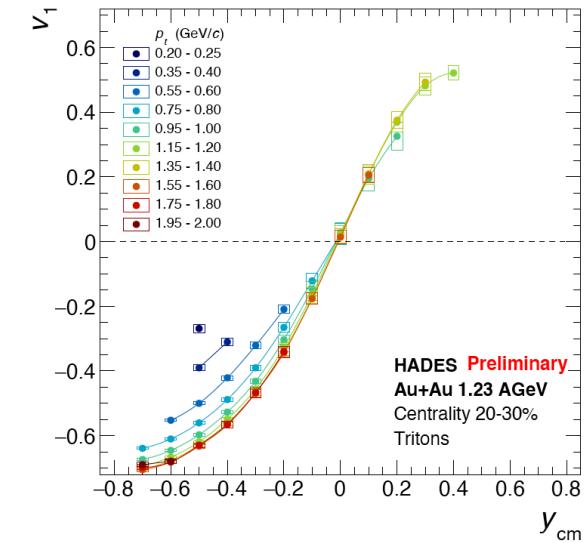
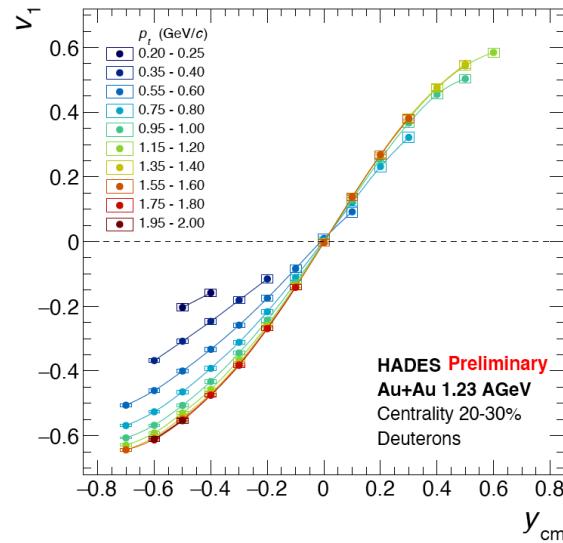
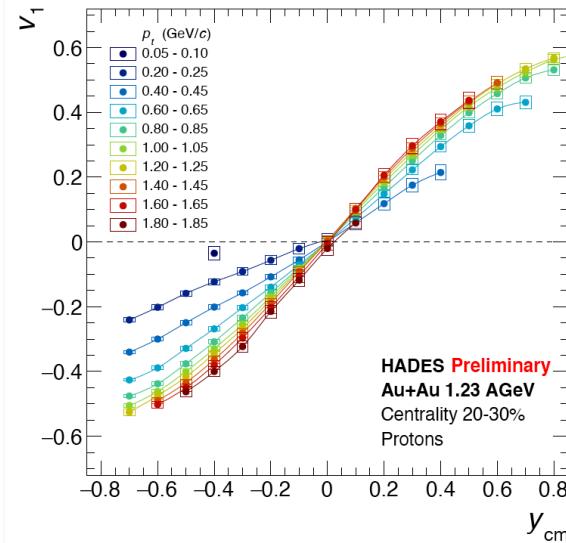
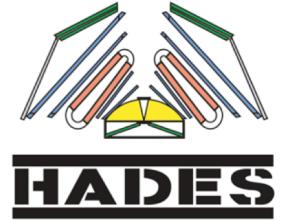
### Out-of-plane $v_2$

Long spectator passing time  $\tau_{\text{passing}} \approx \tau_{\text{expansion}}$   $\Rightarrow$  “squeeze-out”

B. Kardan, QM18

# Collective Effects

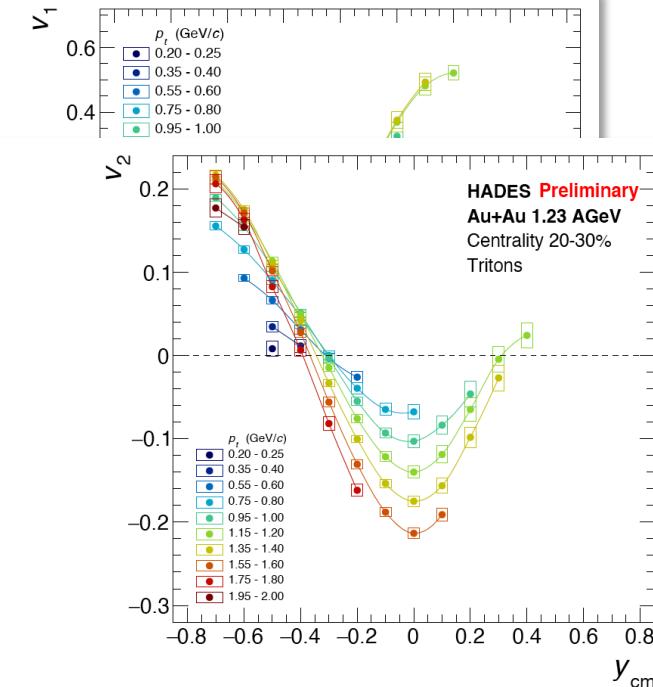
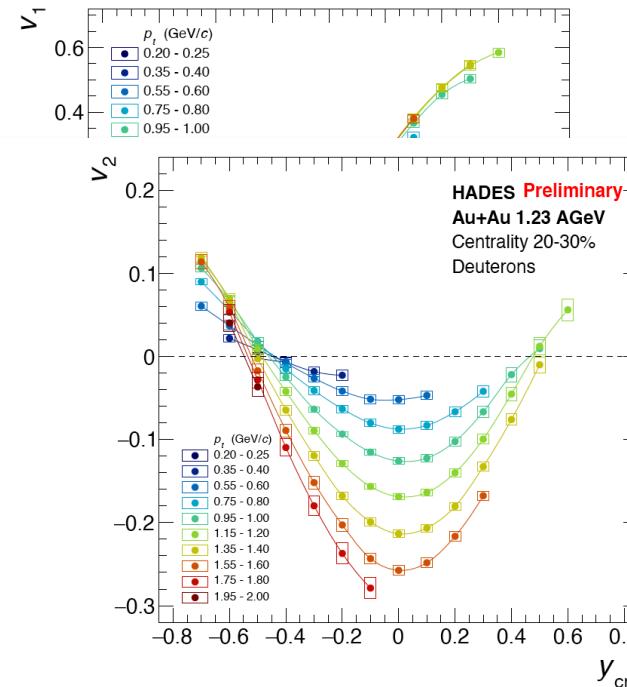
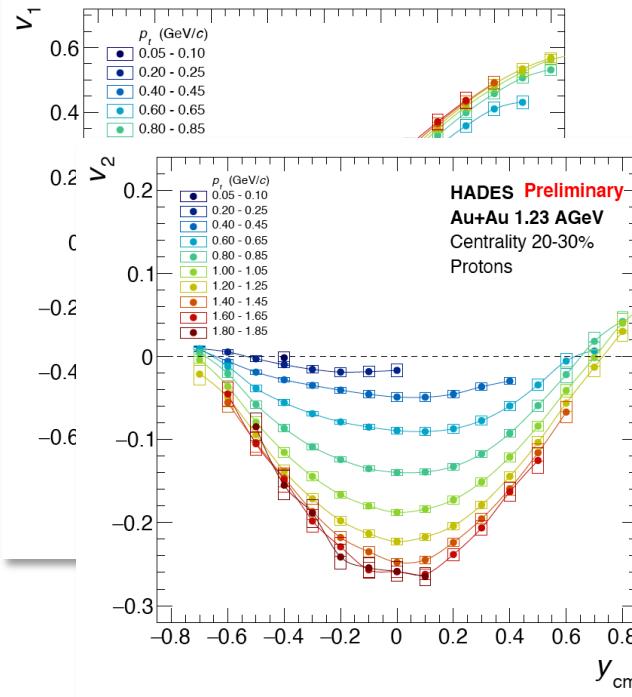
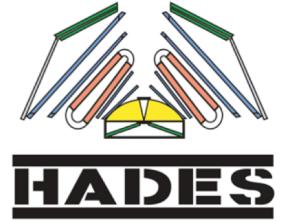
Results on  $v_1$ ,  $v_2$ ,  $v_3$  and  $v_4$  for Protons, Deuterons and Tritons



B. Kardan,  
QM18

# Collective Effects

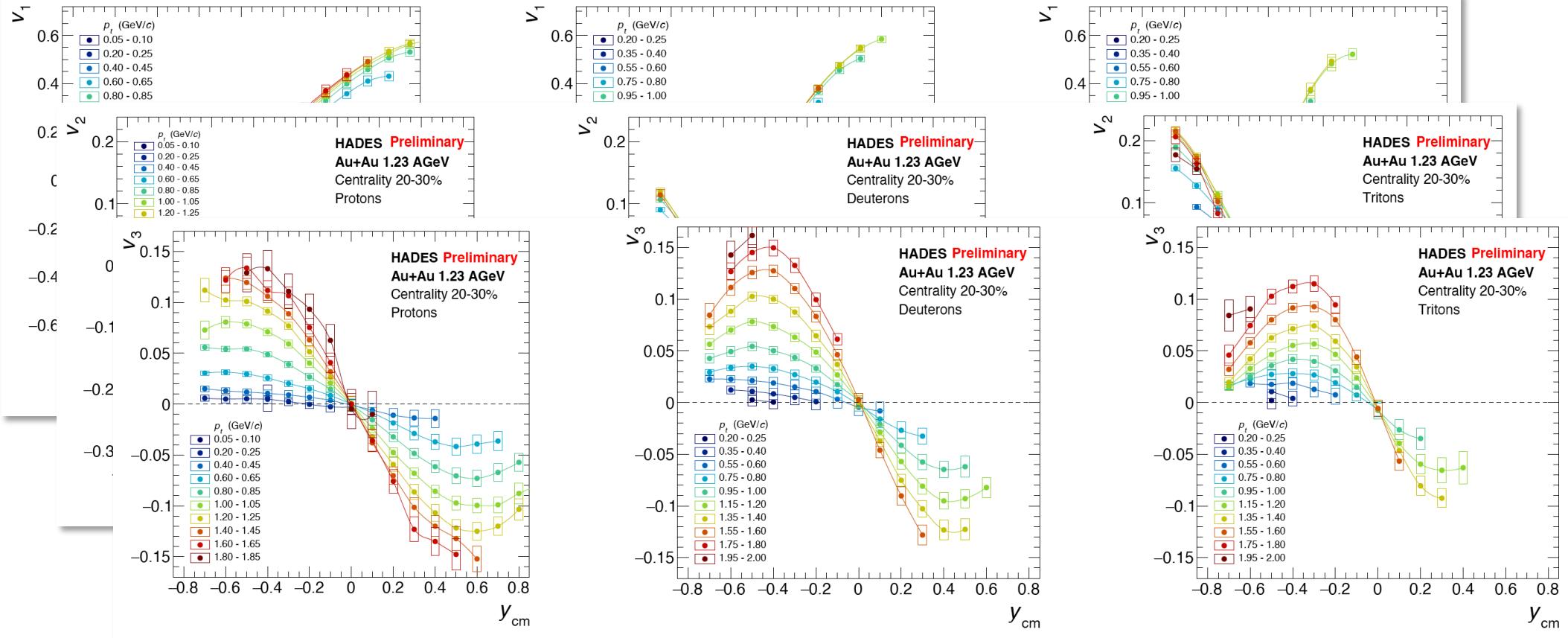
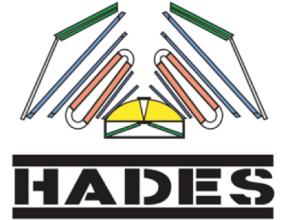
Results on  $v_1$ ,  $v_2$ ,  $v_3$  and  $v_4$  for Protons, Deuterons and Tritons



B. Kardan,  
QM18

# Collective Effects

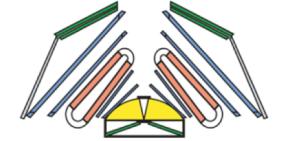
Results on  $v_1$ ,  $v_2$ ,  $v_3$  and  $v_4$  for Protons, Deuterons and Tritons



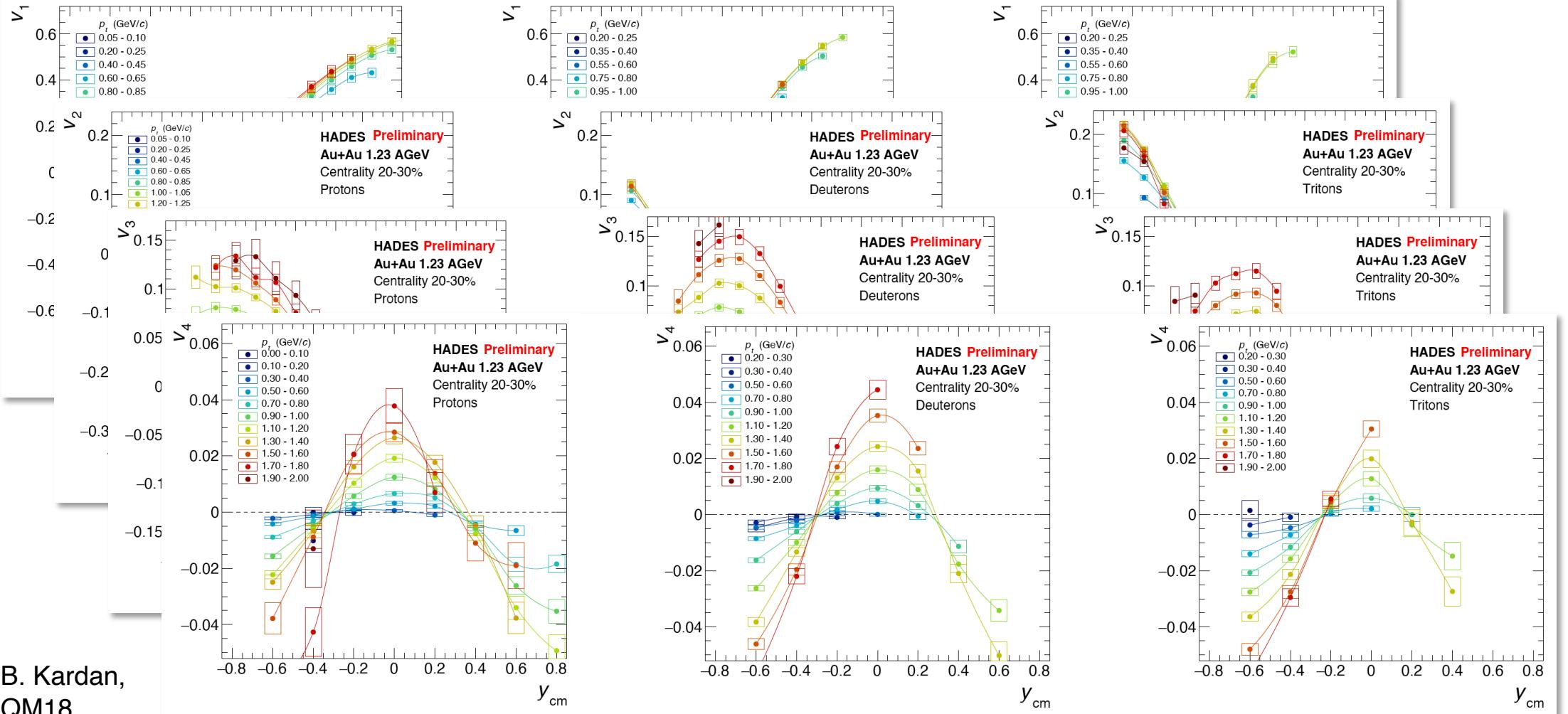
B. Kardan,  
QM18

# Collective Effects

## Results on $v_1$ , $v_2$ , $v_3$ and $v_4$ for Protons, Deuterons and Tritons



**HADES**



B. Kardan,  
QM18

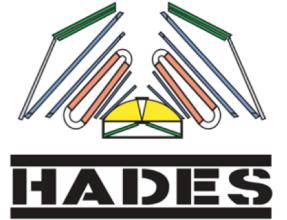
Christoph Blume

Bormio, Italy, January 2020

24

# Collective Effects

## Relation between $v_2$ and $v_4$



### Scaling properties

Prediction for ideal fluid:

$$\frac{v_4(p_t)}{v_2^2(p_t)} = \frac{1}{2}$$

P.F. Kolb, PRC **67** (2003) 031902

N. Borghini and J.-Y. Ollitrault, PLB **642** (2006) 227

C. Gombeaud and J.-Y. Ollitrault, PRC **81** (2010) 014901

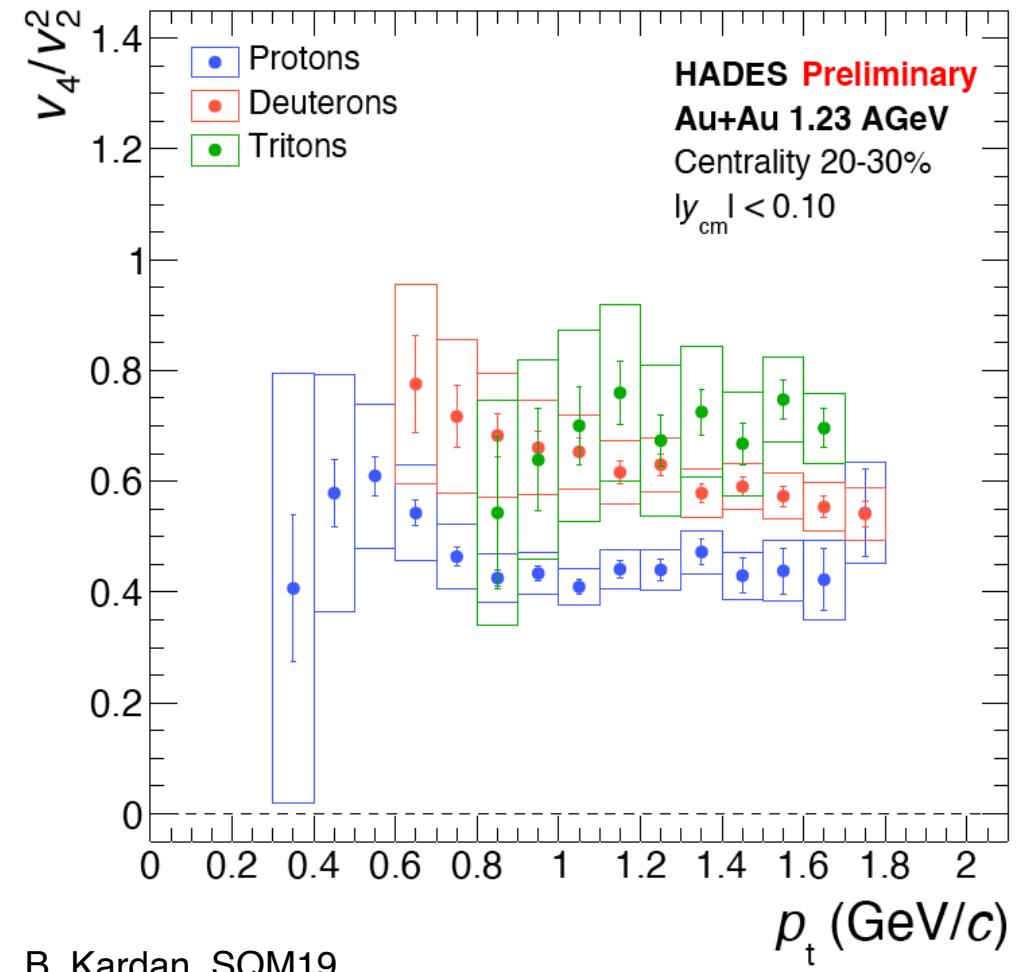
Slightly higher values ( $\sim 0.6$ ) expected in more realistic scenario

### Observed ratios for p, d and t

Independent of  $p_t$

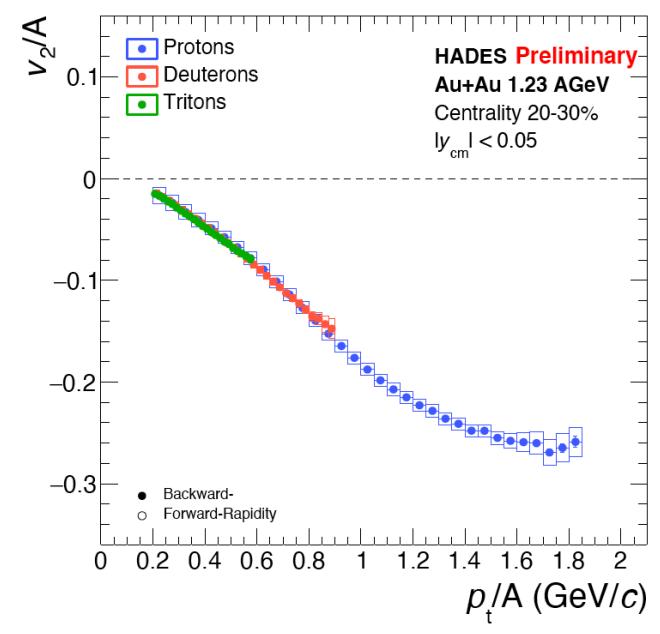
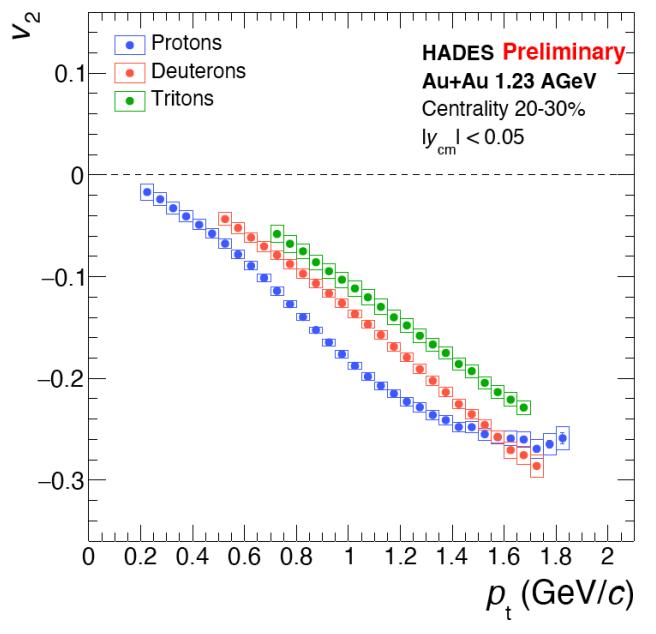
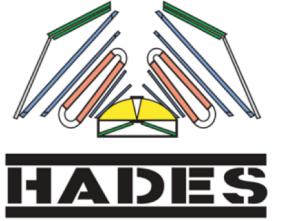
Close to predicted value of  $\sim 0.6$

Hydro-like matter at HADES energies?



# Collective Effects

## Scaling Properties of $v_2$ at Mid-Rapidity



### Nucleon Coalescence

Scaling of  $v_2$  and  $p_t$  with nuclear mass number  $A$

Works as expected in naive coalescence picture (only at mid-rapidity!)

B. Kardan, QM18  
Nucl. Phys. **A982** (2019) 431

# Femtoscopy

## Principle

### Bose-Einstein-Correlations

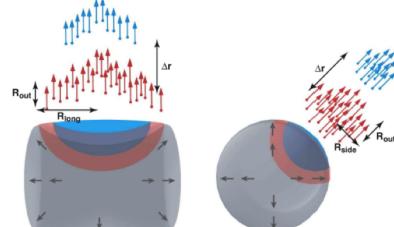
Pairs of identical bosons (charged pions)

Widths of correlation inversely proportional to size  $R$  of pion source

$$\Delta q = \frac{\hbar c}{R} \approx \frac{200 \text{ GeV fm}}{R} \frac{c}{c}$$

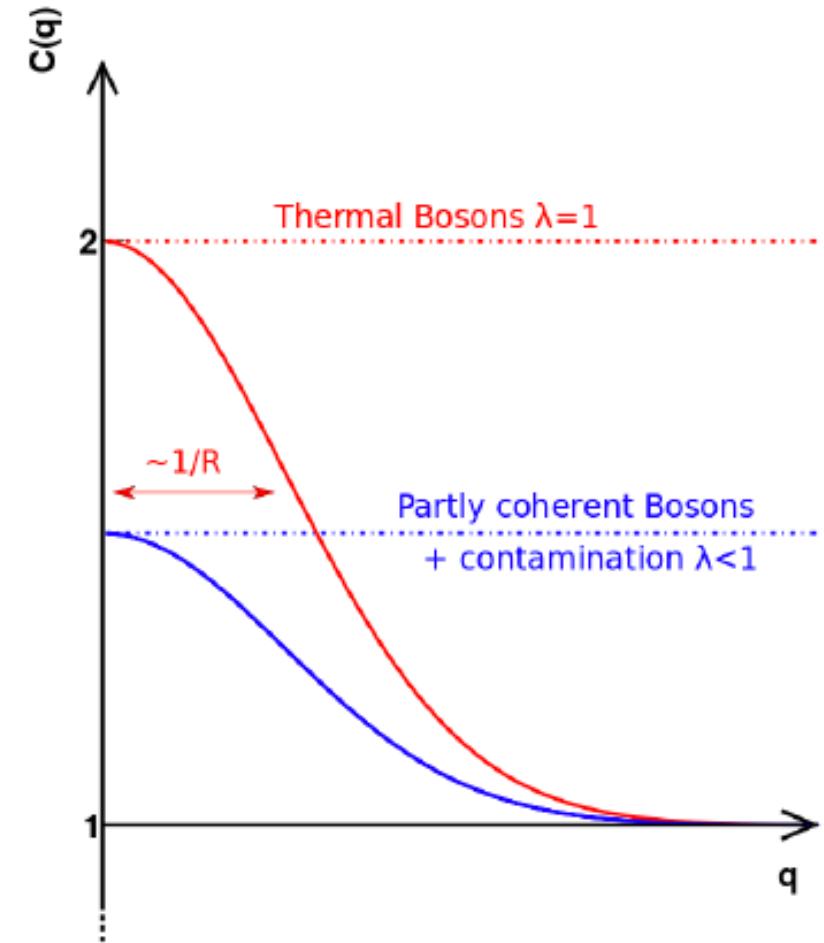
### Multi-dimensional analysis

Access to fireball evolution  
(radial flow)



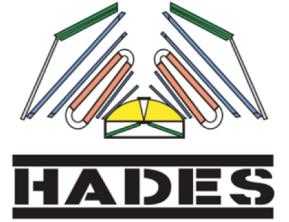
M.A. Lisa and S. Pratt,  
arXiv:0811.1352

Analysis relative to event plane  
Access to event shapes

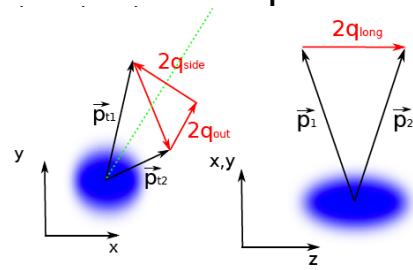


# Femtoscopy

## Charged-Pion Correlations

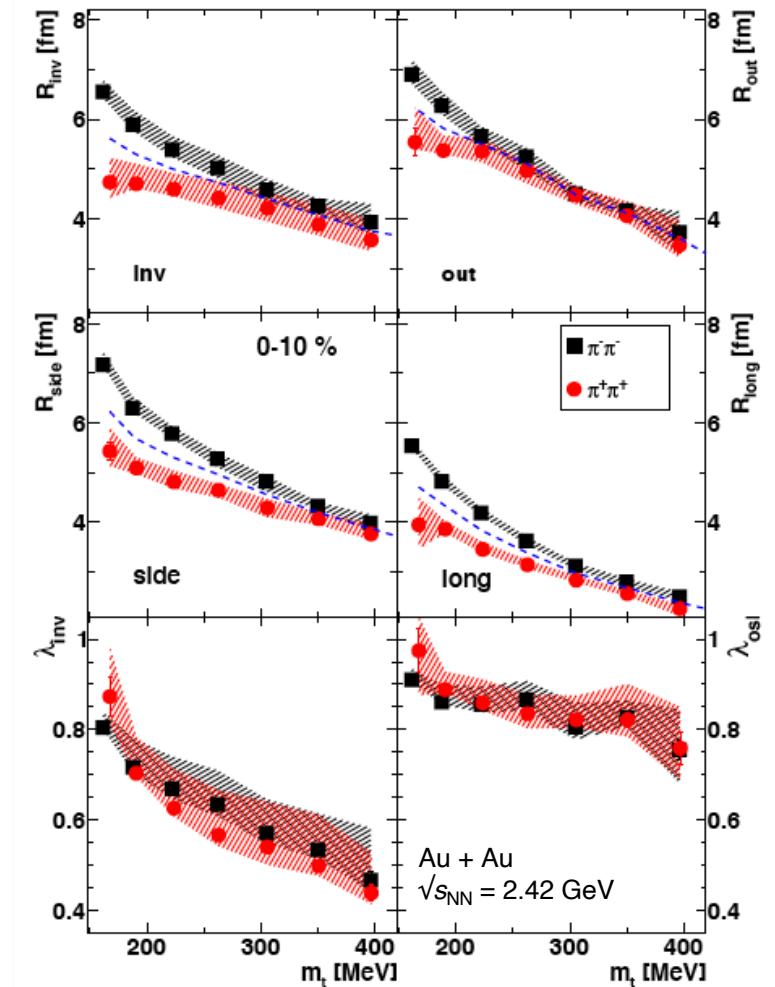


Space-time-extend of fireball  
3D correlation function  
Bertsch-Pratt parameterization



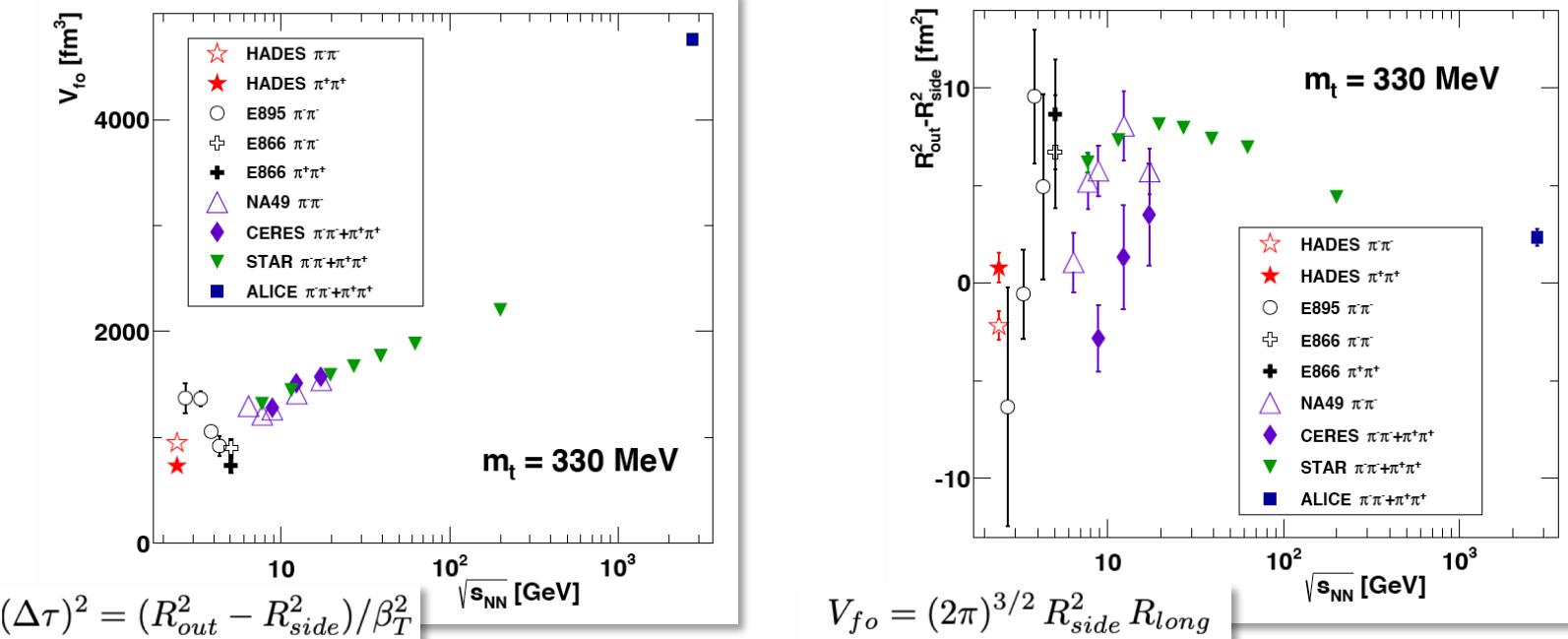
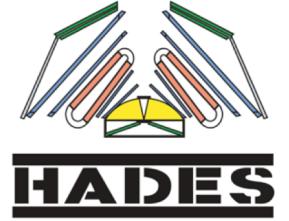
Radius parameters  
Strong  $m_t$ -dependence → radial flow  
1<sup>st</sup> observation of charge sign difference

R. Greifenhagen, QM19  
Phys. Lett. **B795** (2019) 446  
arXiv:1910.07885



# Femtoscopy

## Radius Parameters



### Energy dependences

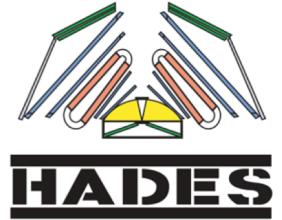
Freeze-out volume: HADES result follows trend from higher energies (SPS, RHIC)  
Room for structures at low energies as indicated by E895 data?

Difference between  $R_{out}$  and  $R_{side}$  close to zero (HADES), maximal for intermediate energies (top-SPS, RHIC)

R. Greifenhagen, QM19  
Phys. Lett. **B795** (2019) 446  
arXiv:1910.07885

# Femtoscopy

## Azimuthal Dependence



### Fits relative to event plane

Rotation of osl-system relative to EP-system

Formulas: PLB 496 (2000) 1, PRC 57 (1998) 266

Corrected for EP-resolution

⇒ Access to event shape parameters

### Eccentricity $\epsilon$ in xy-plane

Compare to initial participant eccentricity

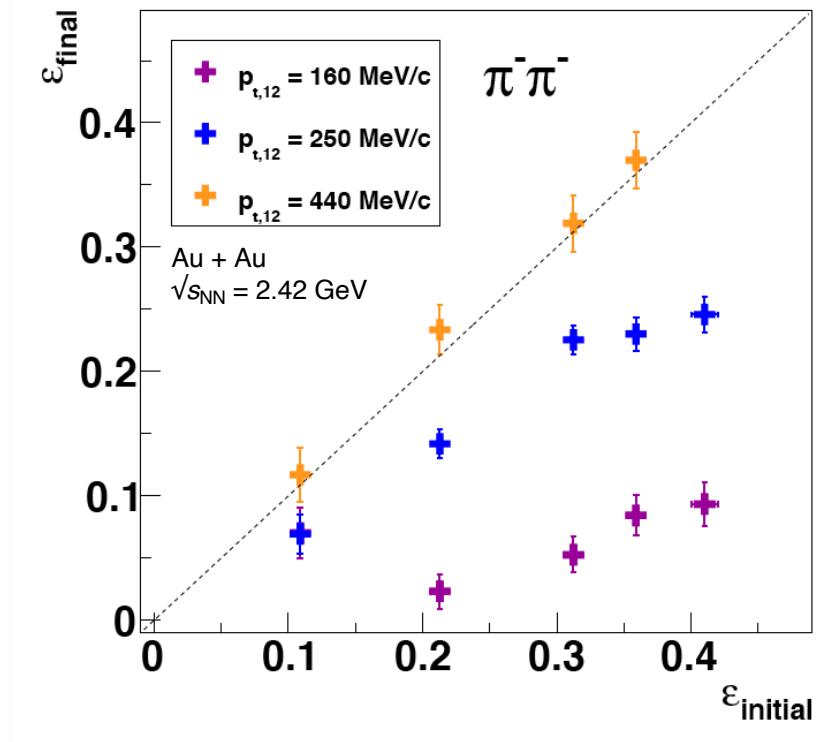
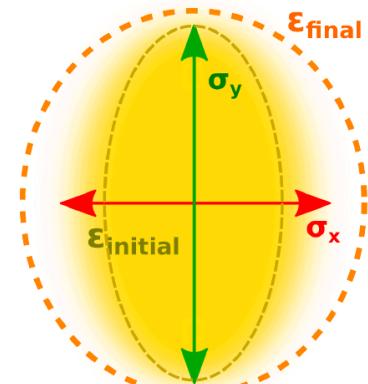
$\epsilon_{\text{initial}}$  from Glauber MC

Early stage (high  $p_{t,12}$ ):

$\epsilon_{\text{final}} \approx \epsilon_{\text{initial}}$

Late stage (low  $p_{t,12}$ ):

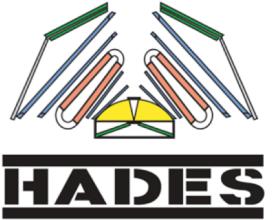
$\epsilon_{\text{final}} \rightarrow 0$



$$\epsilon_{\text{final}} = \frac{\sigma_y^2 - \sigma_x^2}{\sigma_y^2 + \sigma_x^2}$$

R. Greifenhagen, QM19  
arXiv:1910.07885

# Global $\Lambda$ Polarization Principle



## Global polarization

Large angular momenta  $|L| \sim 10^5 \hbar$   
Extreme vorticities possible ( $\omega \approx 10^{21} \text{ s}^{-1}$ )

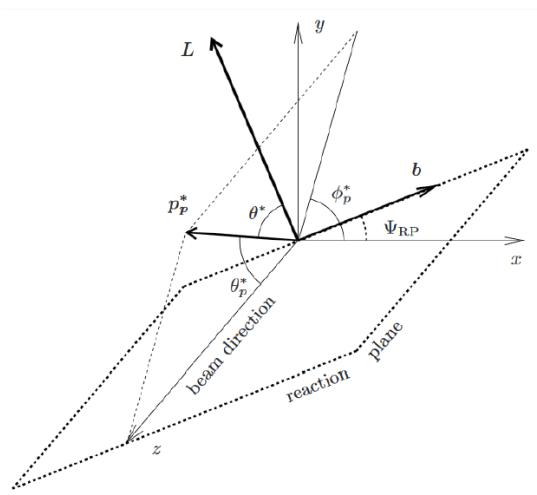
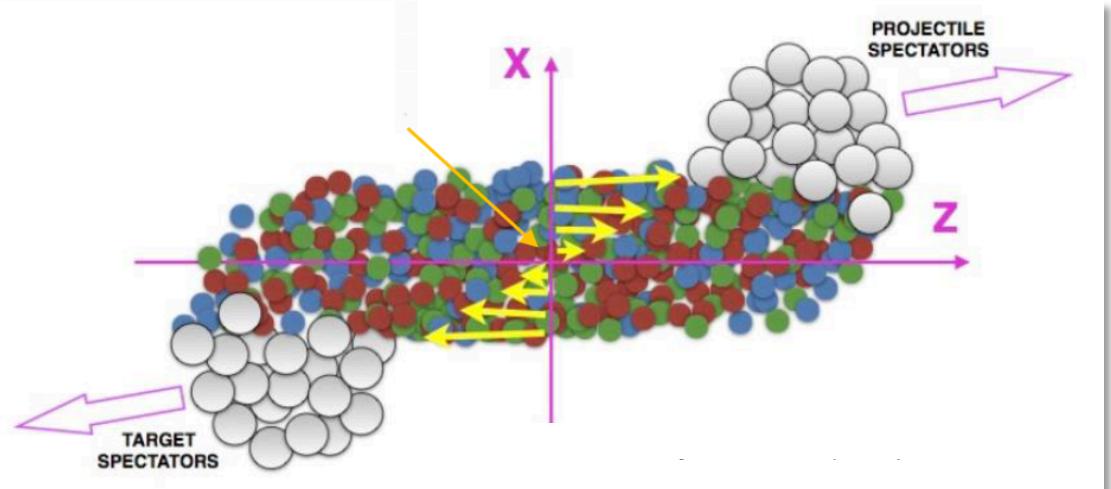
Observable via polarization of spins  
relative to event plane  
(spin-orbit coupling, e.m.-coupling)

## Observable

Weak decay:  $\Lambda \rightarrow p + \pi^-$   
Proton preferentially in spin direction  
 $\Rightarrow$  Polarization  $P_\Lambda$ :

$$P_\Lambda = \frac{8}{\pi \alpha_\Lambda} \frac{\langle \sin(\Psi_{EP} - \phi_p^*) \rangle}{R_{EP}}$$

$\Lambda$  decay parameter:  $a_\Lambda = 0.643 \pm 0.013$   
 $\Psi_{EP}$  = event plane angle,  $R_{EP}$  = EP-resolution  
 $\phi_p^*$  = proton azimuth angle relative to EP



Z. Liang and X.N. Wang,  
PRL 94 (2005) 102301

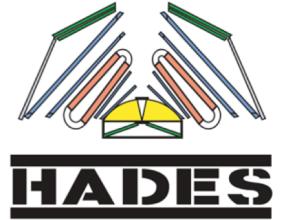
F. Becattini et al.,  
PRC 95 (2017) 054902

STAR Collaboration,  
PRC 76 (2007) 024915

F. Kornas,  
SQM19

# Global $\Lambda$ Polarization

## Measurement at $\sqrt{s_{NN}} = 2.42$ GeV



### Analysis procedure

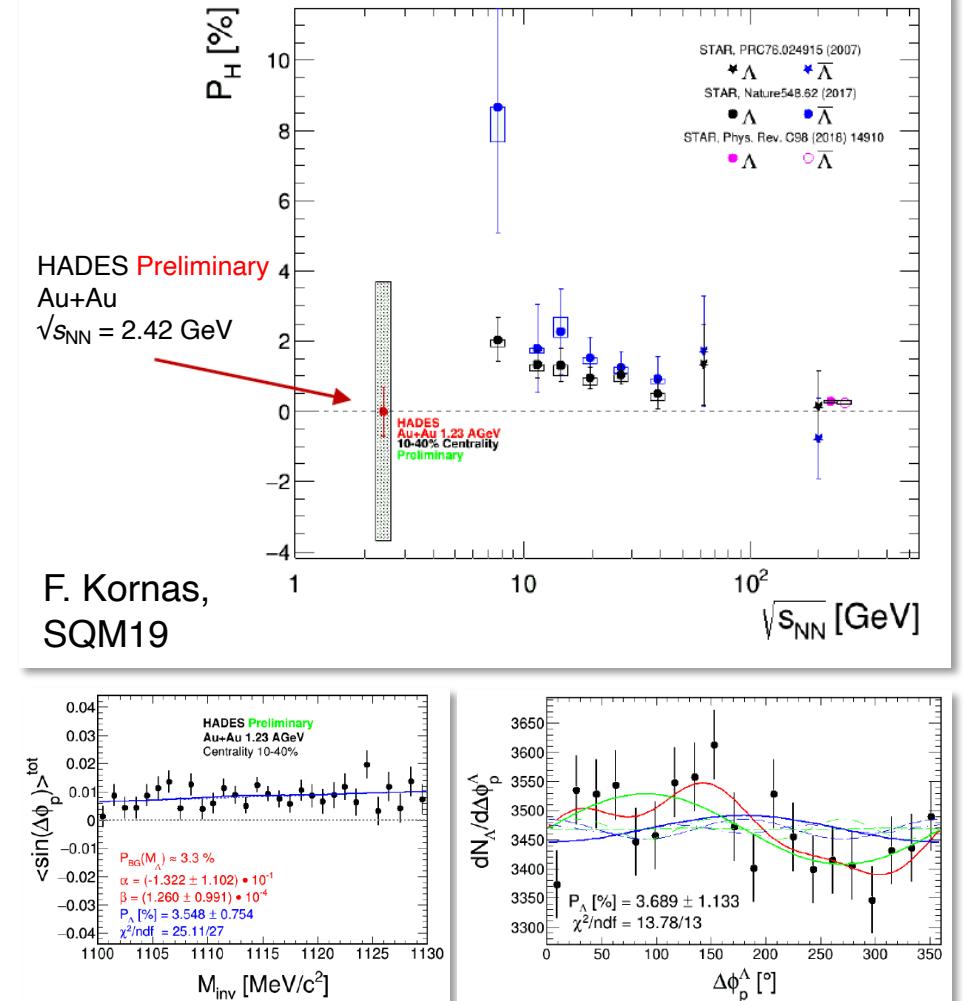
EP estimation from spectators in FW  
Optimized  $\Lambda$  reconstruction with ANN  
 $N_\Lambda = 1.9 \cdot 10^5$  (10 – 40%)

Two methods:  
Invariant mass fit and EP-method

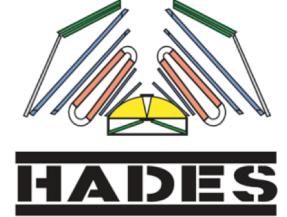
### Preliminary Results

$P_\Lambda^{\text{EP}} = (3.762 \pm 0.699 \text{ (stat.)}) \%$   
 $P_\Lambda^{\text{MINV}} = (3.548 \pm 0.754 \text{ (stat.)}) \%$   
Background:  
 $P_\Lambda^{\text{BG}} = (3.689 \pm 1.133 \text{ (stat.)}) \%$

On-going investigations:  
Systematic effects  
Influence of finite detector acceptance



# Proton Number Fluctuations



## Motivation

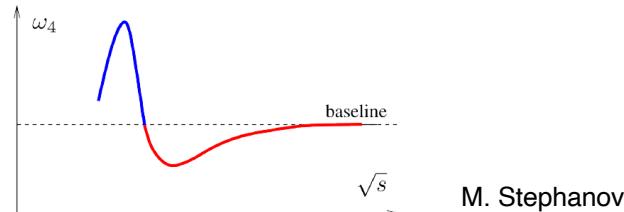
### Search for critical point

- Susceptibilities diverge
- Enhanced fluctuations

Fluctuations of conserved quantities  
(strangeness, baryon number, charge)

Higher moments should be more sensitive

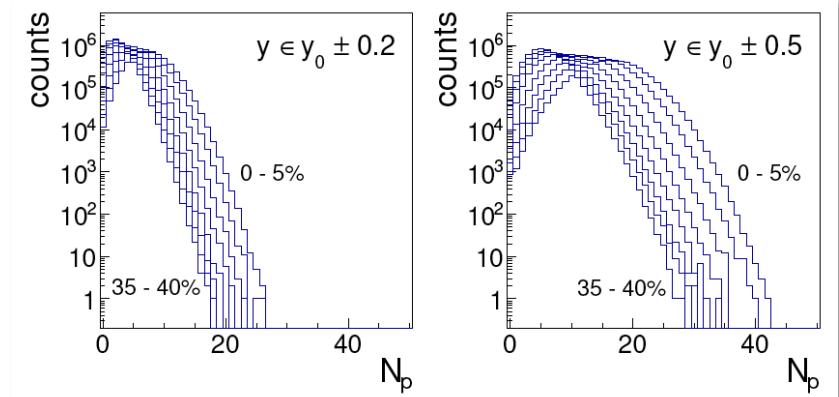
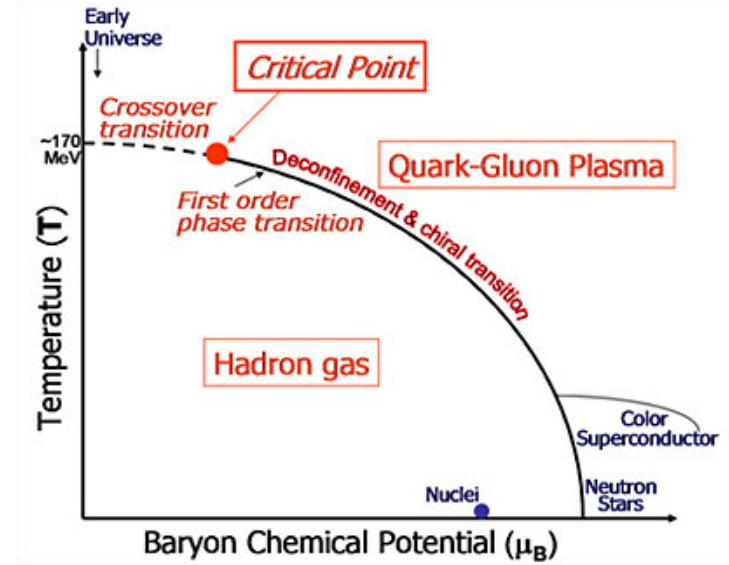
Vary freeze-out conditions via energy scan



M. Stephanov

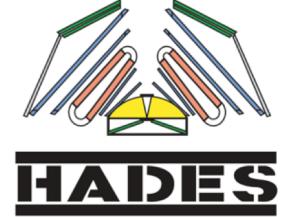
### Observable

Moments of net-proton distributions  
(proxy for baryon number)



# Proton Number Fluctuations

## Corrections



### Observable (HADES)

Moments of proton multiplicity distributions  
Centrality selection with Forward Wall (5%)

### Efficiency correction

Track density dependent efficiency

Evaluated via three different methods

- Binomial E-by-E: A. Bzdak and V. Koch, PRC **86** (2012) 044904  
S. He and X. Luo, Chin. Phys. **C42** (2018) 104001
- Unfolding: P. Garg et al., J. Phys. G **40** (2013) 055103
- Moment expansion: T. Nonaka et al., NIM **A906** (2018) 10

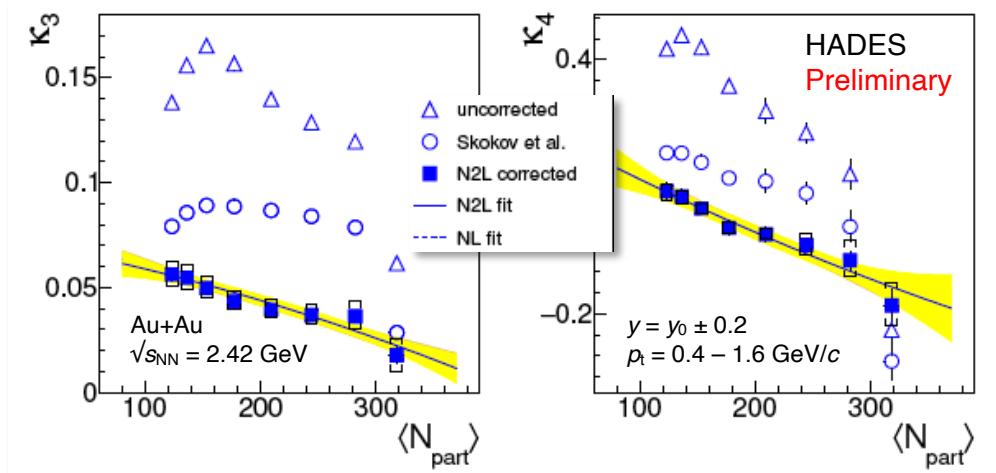
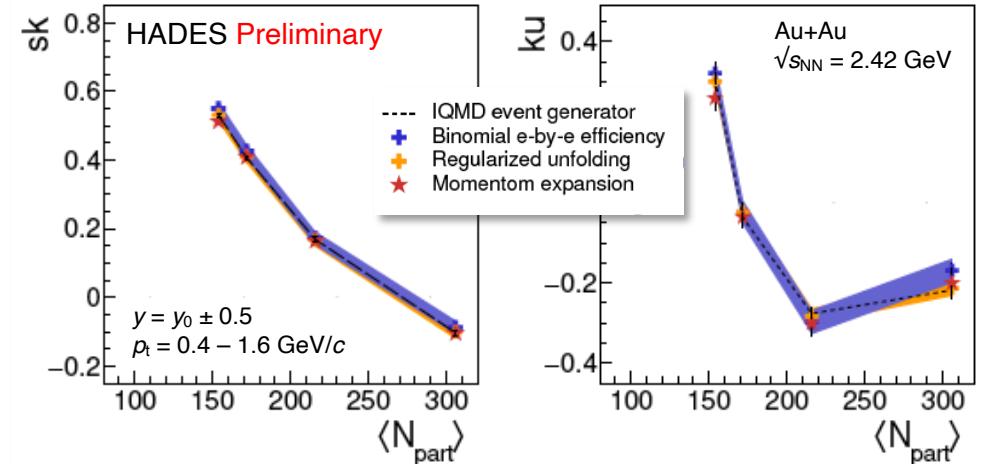
### Volume corrections

Extension of assumption of constant  
reduced cumulants  $\kappa_n(V) = K_n/V = \text{const.}$

- V. Skokov et al., PRC **88** (2013) 034911  
P. Braun-Munzinger et al., NPA **960** (2017) 114

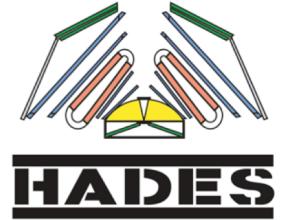
Here: 2<sup>nd</sup>-order approach:

$$\kappa_n(V) = \kappa_n + \kappa'_n(V - \langle V \rangle) + \kappa''_n(V - \langle V \rangle)^2$$



# Proton Number Fluctuations

## Energy Dependence of Scaled Cumulants



### Extension of STAR-BES results

Skewness ( $Sk \cdot \sigma$ ): smooth trend

Kurtosis ( $Ku \cdot \sigma^2$ ): change of sign (0-5%)

### Contribution from spectators

Fluctuation sources: fireball  $\leftrightarrow$  spectators

Relative admixture energy dependent!

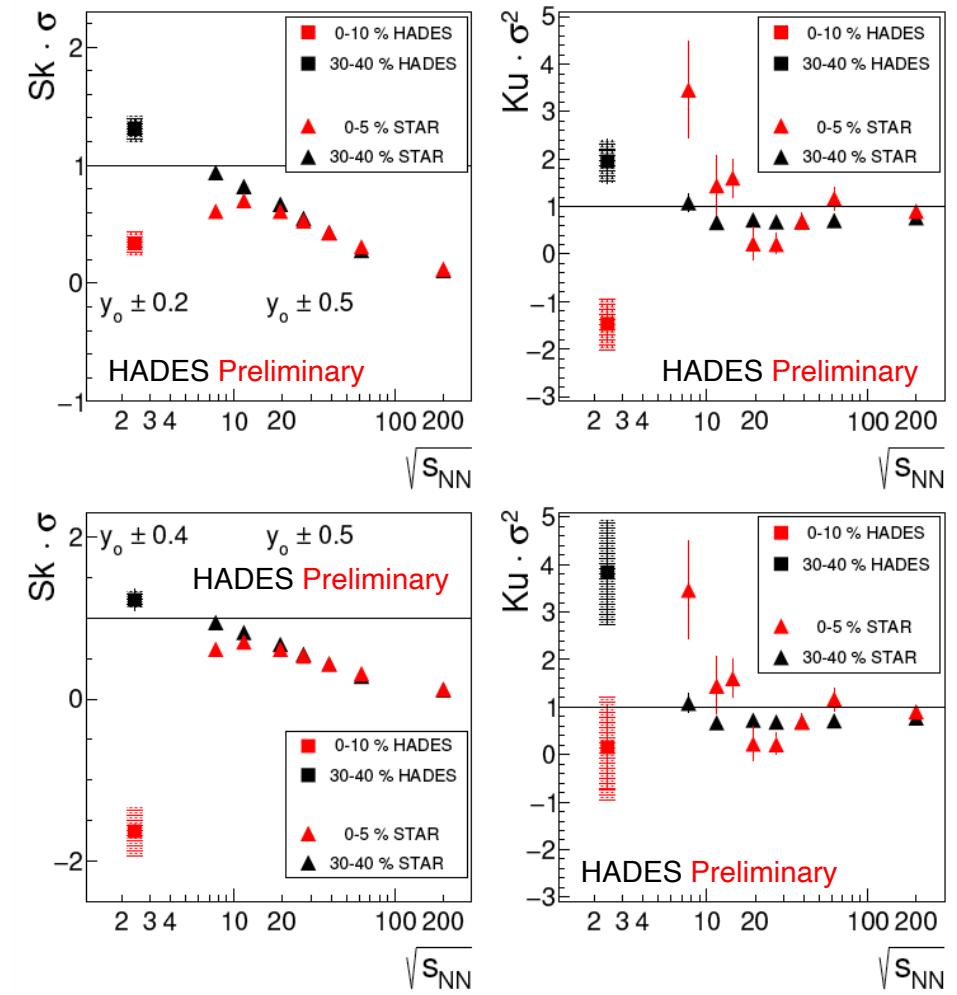
$\Rightarrow$  Two rapidity intervals shown:

$y_0 \pm 0.2$  and  $y_0 \pm 0.4$  (STAR:  $y_0 \pm 0.5$ )

### Outlook

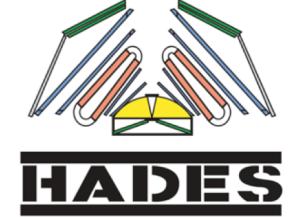
Include bound protons (d, t, He)

Ag+Ag data



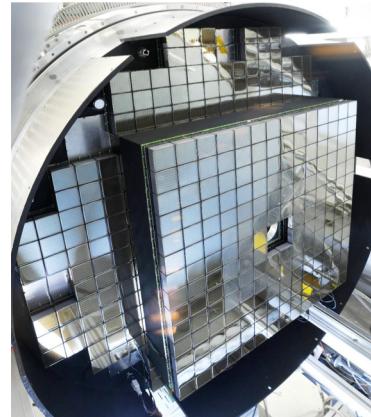
# FAIR Phase-0

## Ag+Ag at $\sqrt{s_{NN}} = 2.6$ GeV

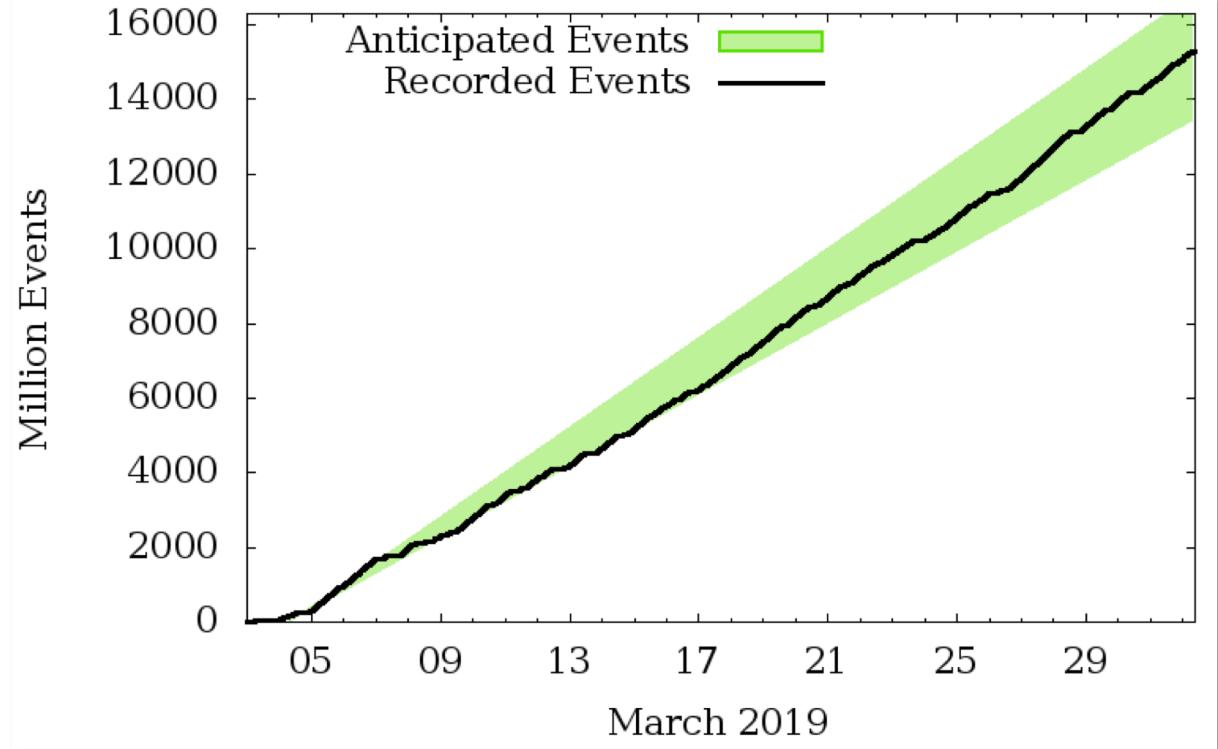


New data set from March 2019  
~ 15 billion events collected

Upgraded setup  
New RICH photon detector  
(same MAPTs as for CBM)

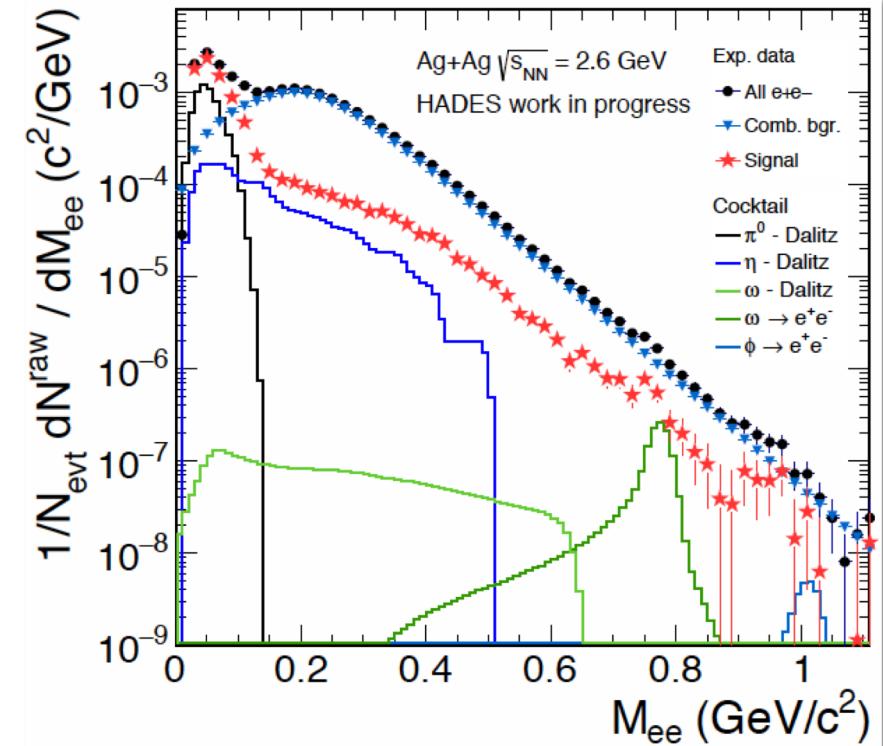
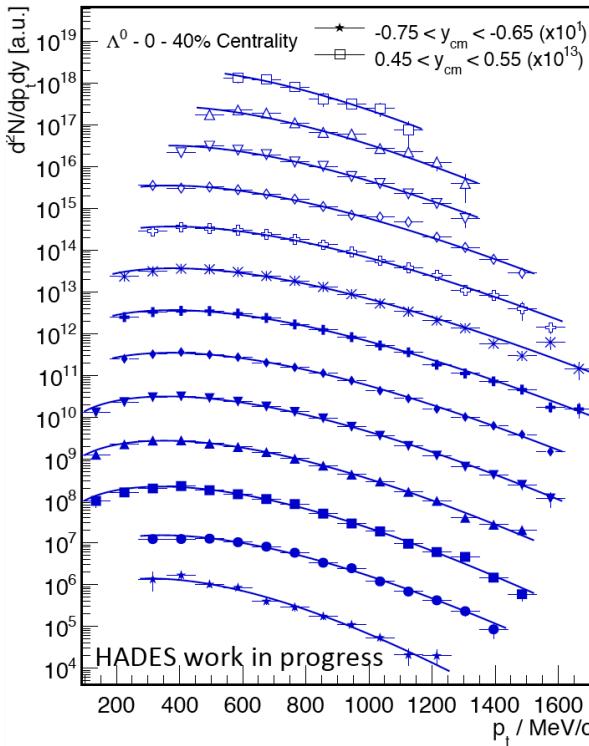
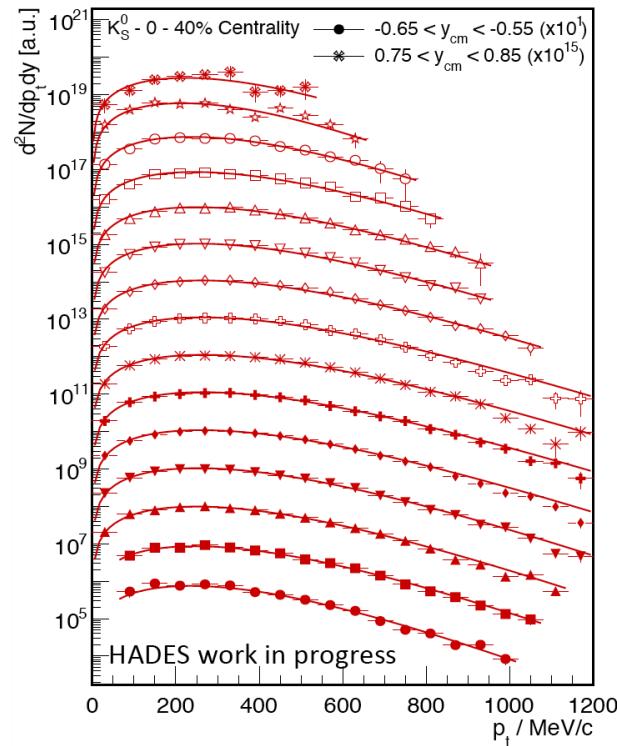


Electromagnetic calorimeter  
⇒ Poster by P. Chudoba  
⇒ Talk by A. Rost



# FAIR Phase-0

## Ag+Ag at $\sqrt{s_{NN}} = 2.6$ GeV



First glimpse into new data set  
 $K_0^s$  and  $\Lambda$  spectra at the NN-threshold  
1<sup>st</sup> Dilepton spectrum

# Conclusions

## “Thermalized” system at high baryon densities

- Particle yields at freeze-out
- Flow properties
- Thermal di-electron excess

## Low energy point in excitation functions

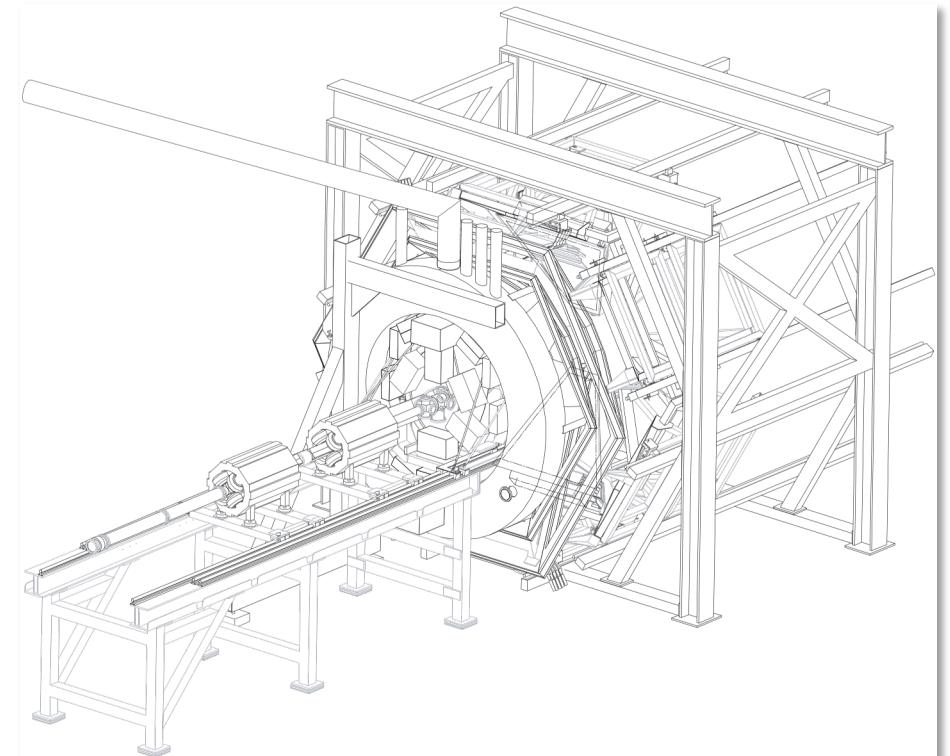
- Femtoscopy radii
- Global  $\Lambda$  polarization
- Proton fluctuations

## In-medium effects in cold nuclear matter

- Evidence for kaon and  $\phi$  absorption

## Strong scientific program for FAIR Phase-0

- Ag+Ag at  $\sqrt{s_{NN}} = 2.6 \text{ GeV}$
- $\pi + A$  Program at SIS18



WYDZIAŁ FIZYKI ASTRONOMII I INFORMATYKI STOSOWANEJ

Many Thanks!

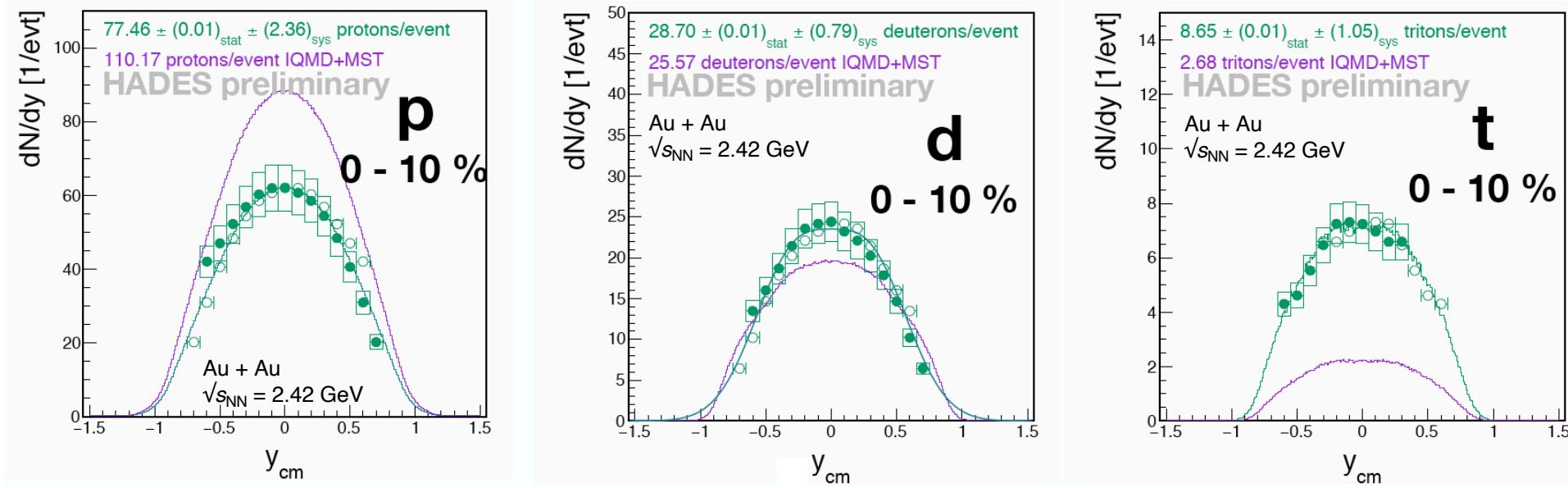
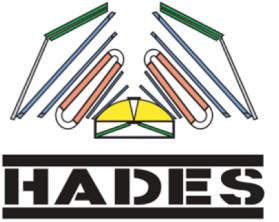


The HADES Collaboration

# BACKUP

# Particle Production

## Protons, Deuterons and Tritons



### Comparison to simple coalescence approach

Transport model + clustering afterburner (IQMD + MST\*,  $r = 5$  fm in pos. space,  $t < 140$  fm/c)  
Unable to reproduce triton data

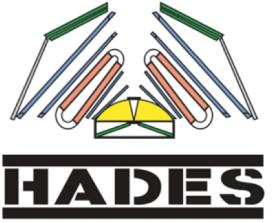
M. Szala,  
SQM19

Comparisons to more involved calculations on-going

\*Thanks to  
Y. Leifels

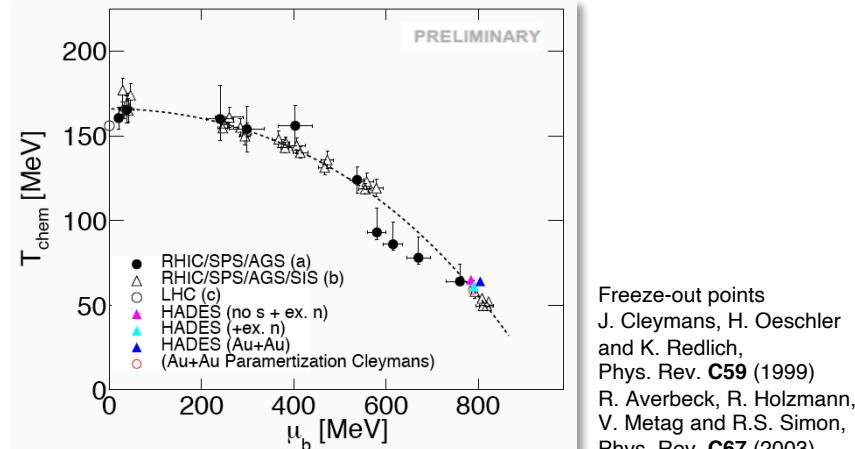
# Particle Production

## Particle Yields → Chemical Freeze-Out



### Fit with THERMAL-FIST

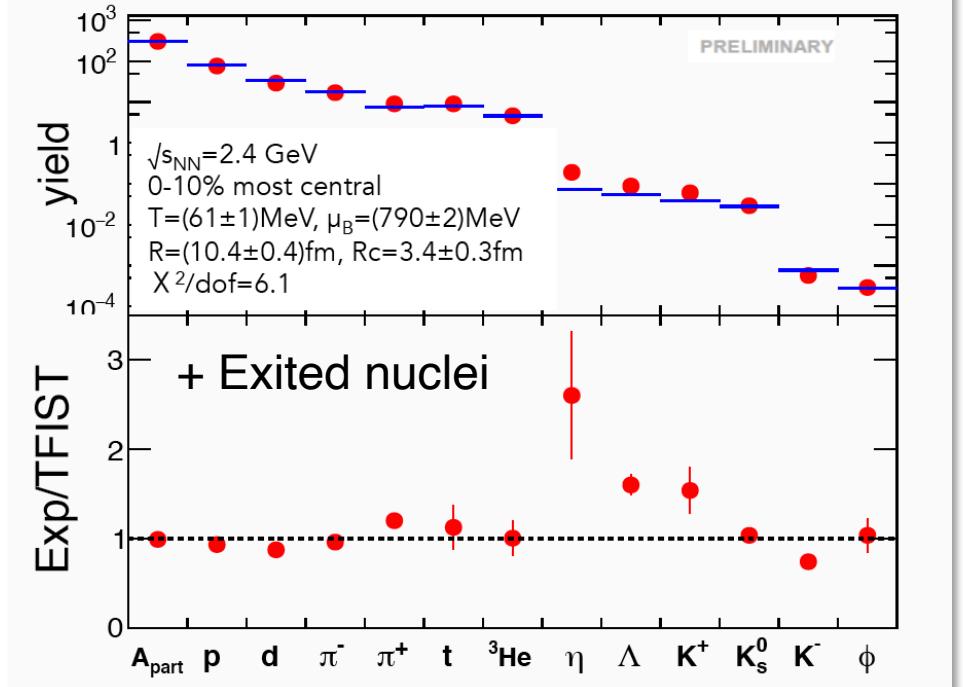
Freeze-Out parameters ( $T_{\text{chem}}$ ,  $\mu_B$ ) follow universal freeze-out curve



Relatively high  $\chi^2$  for full hadron spectrum

Inclusion of excited nuclei states helps  
Add. exclusion of strangeness  $\Rightarrow$  small  $\chi^2$   
(see backup slides)

Thermal Fist: V. Vovchenko H. Stoecker, Comput. Phys. Commun. 244 (2019) 295.



M. Lorenz,  
EMMI-Workshop on Anti-matter,  
hyper-matter and exotica

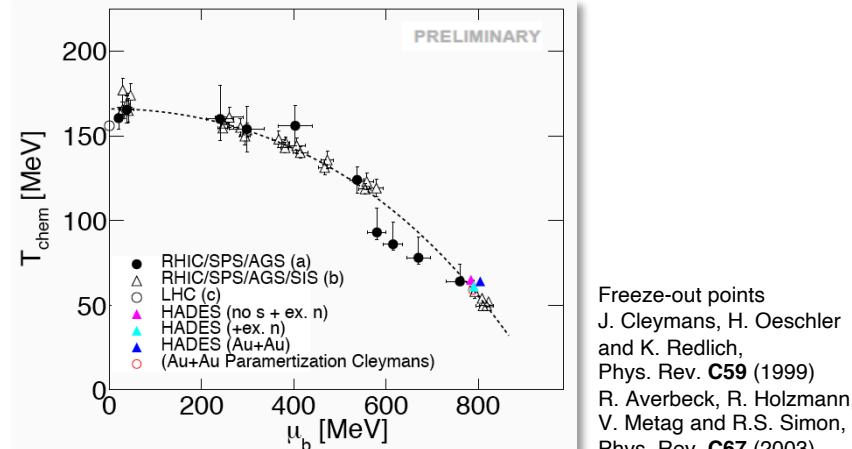
# Particle Production

## Particle Yields → Chemical Freeze-Out



### Fit with THERMAL-FIST

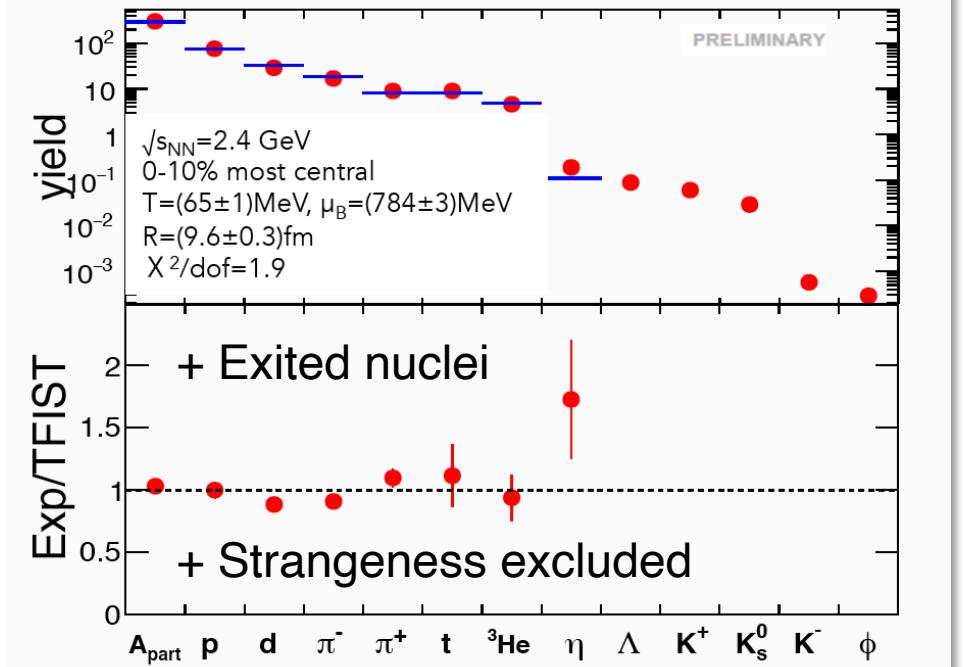
Freeze-Out parameters ( $T_{\text{chem}}$ ,  $\mu_B$ ) follow universal freeze-out curve



Relatively high  $\chi^2$  for full hadron spectrum

Inclusion of excited nuclei states helps  
Add. exclusion of strangeness  $\Rightarrow$  small  $\chi^2$   
(see backup slides)

Thermal Fist: V. Vovchenko H. Stoecker, Comput. Phys. Commun. 244 (2019) 295.



M. Lorenz,  
EMMI-Workshop on Anti-matter,  
hyper-matter and exotica

# Femtoscopy

## Azimuthal Dependence

### Fits relative to event plane

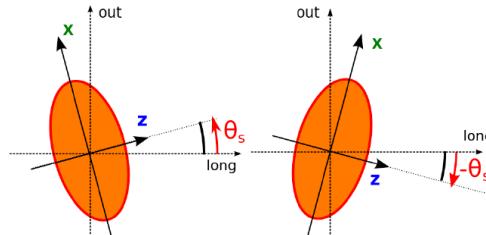
Rotation of osl-system relative to EP-system

Formulas: PLB 496 (2000) 1, PRC 57 (1998) 266

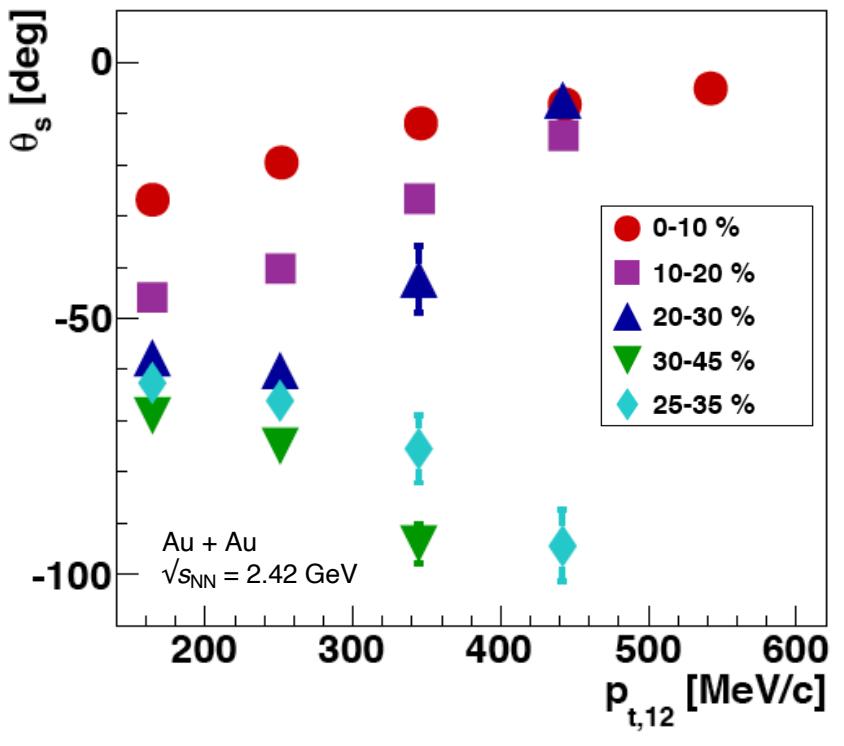
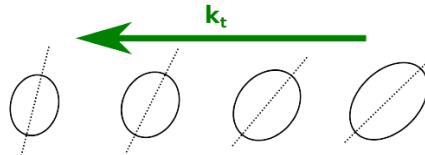
Corrected for EP-resolution

⇒ Access to event shape parameters

### Tilt angle $\theta_s$ in xz-plane



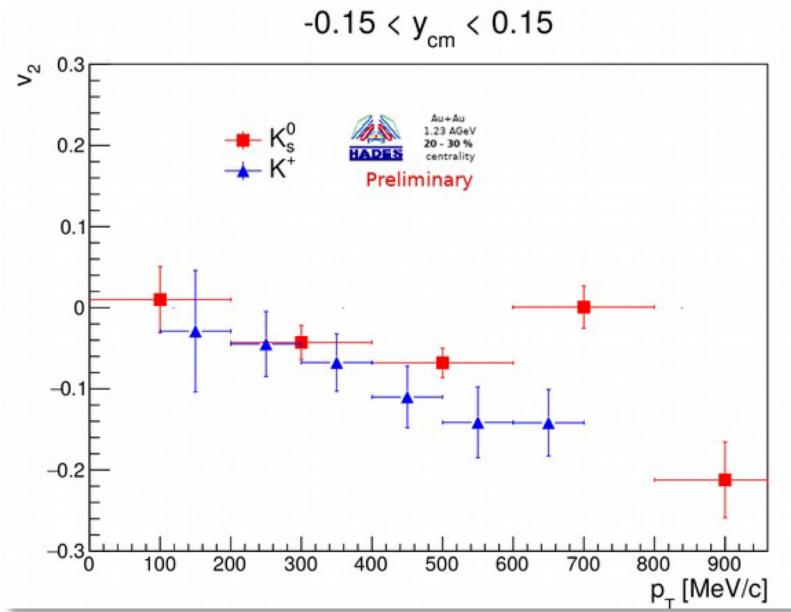
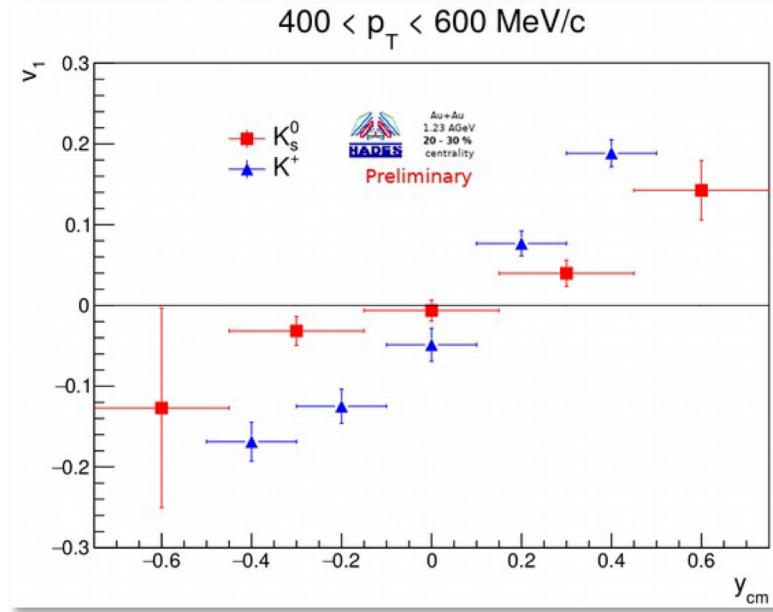
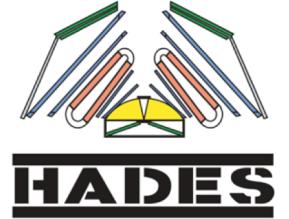
$|\theta_s|$  tends to decrease with increasing  $p_{t,12}$



R. Greifenhagen, QM19  
arXiv:1910.07885

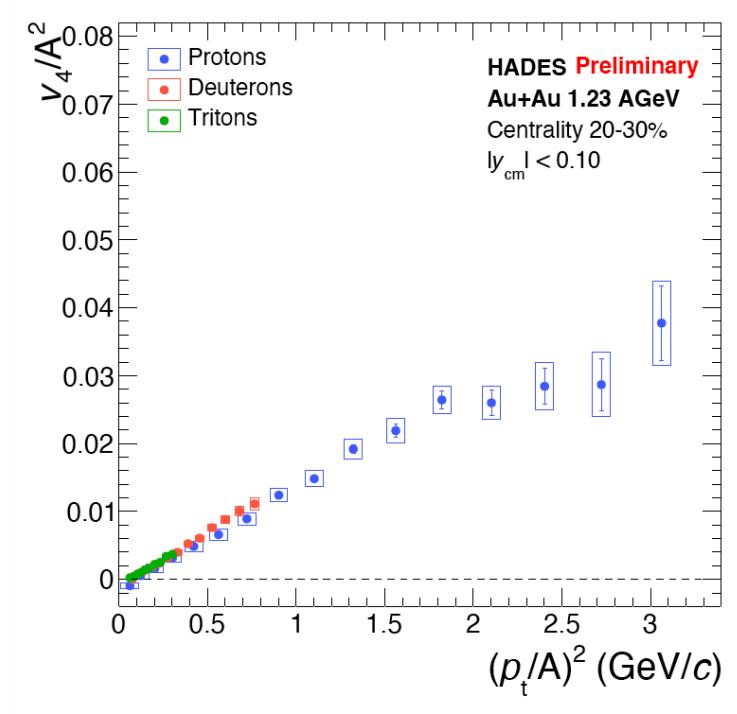
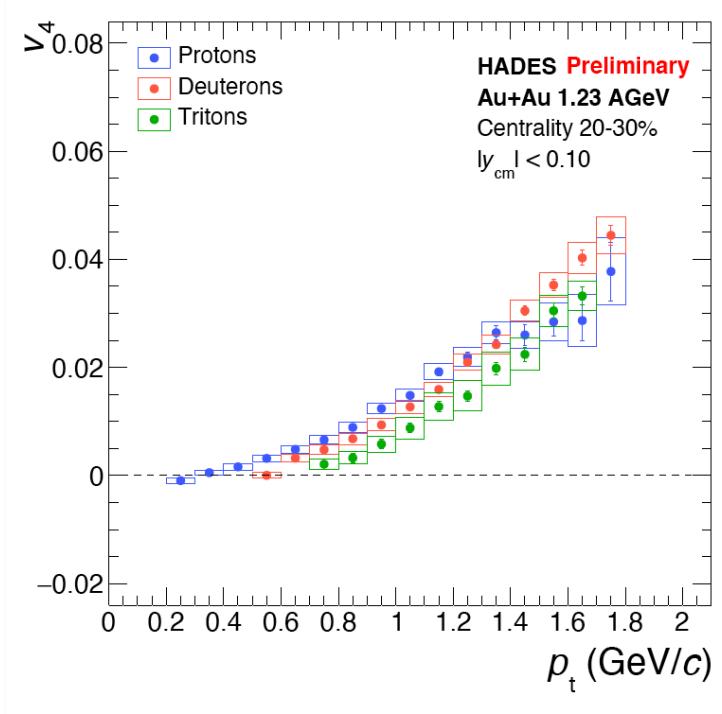
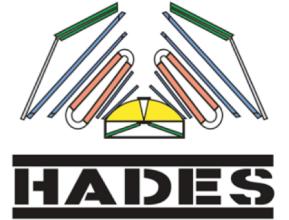
# Collective Effects

## Directed and Elliptic Flow of Kaons



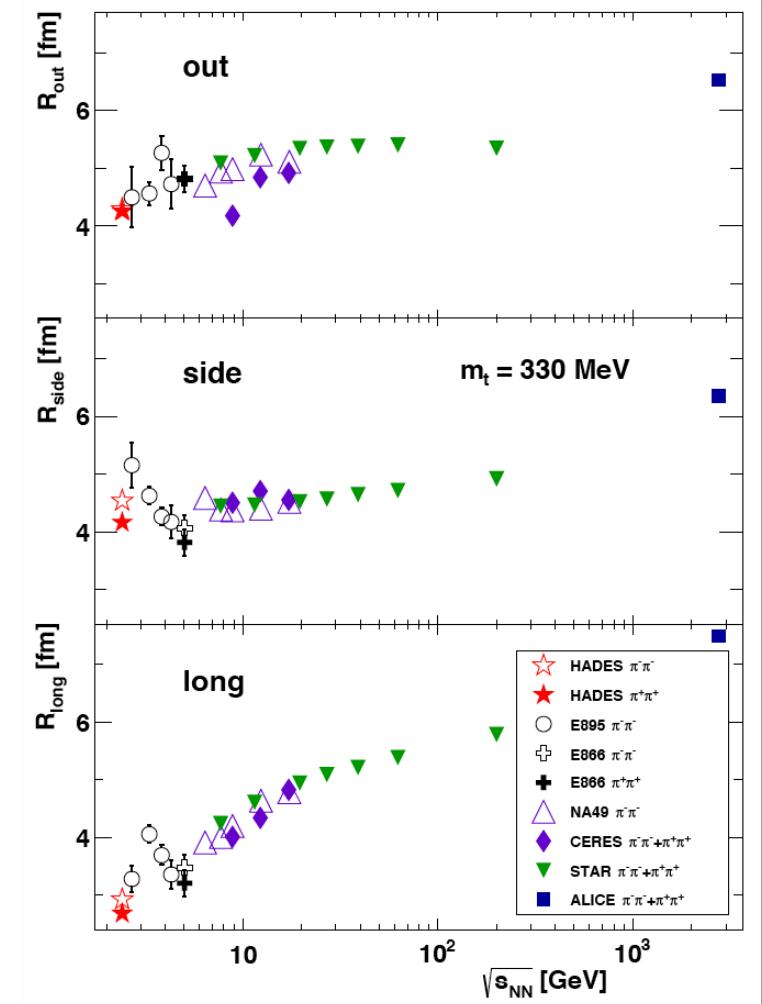
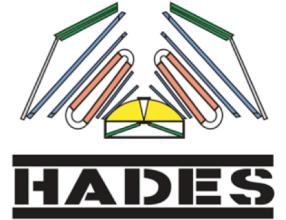
# Collective Effects

## Scaling Properties of $v_4$ at Mid-Rapidity



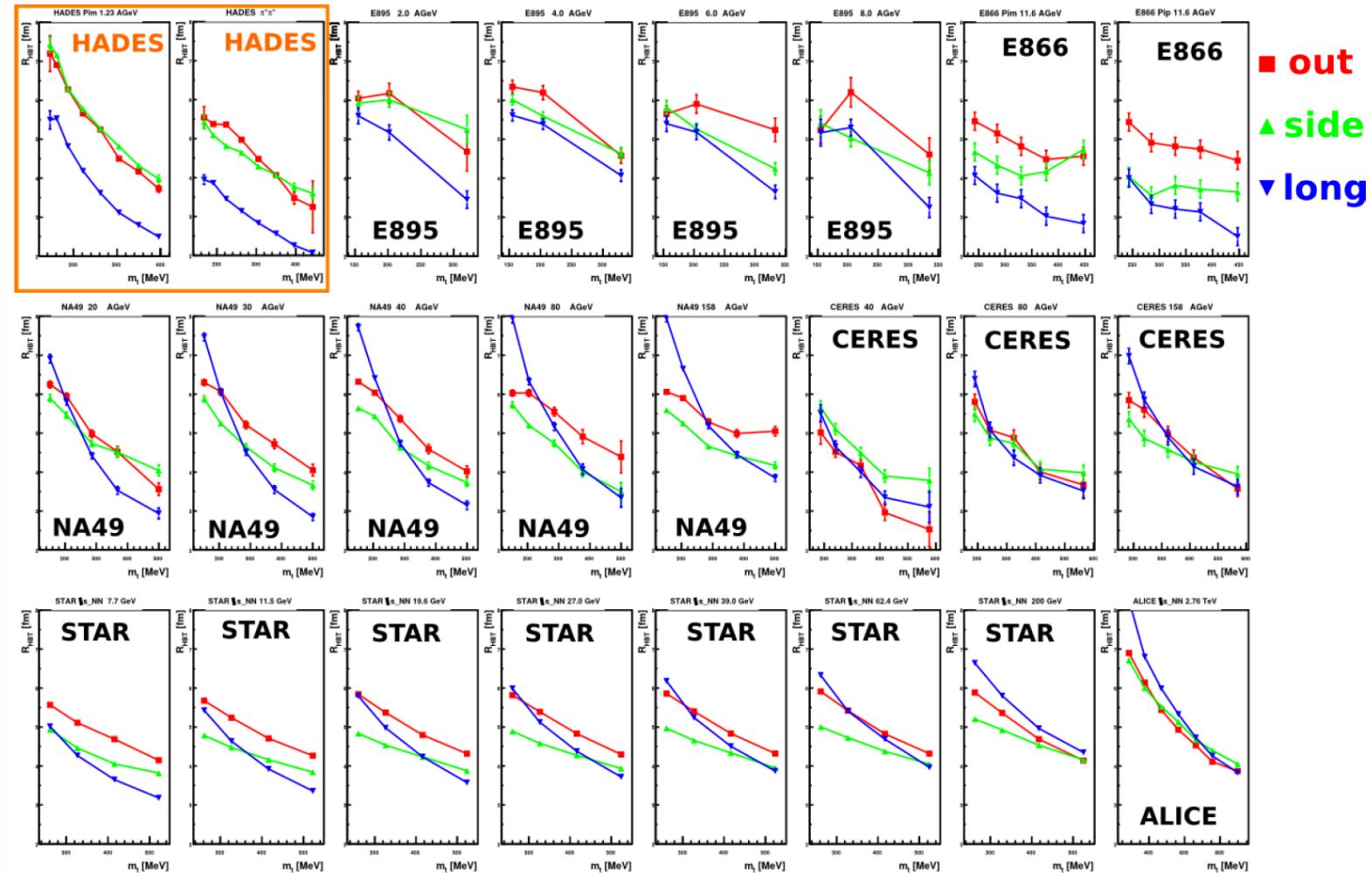
# Femtoscopy

## Energy Dependence of Radius Parameters



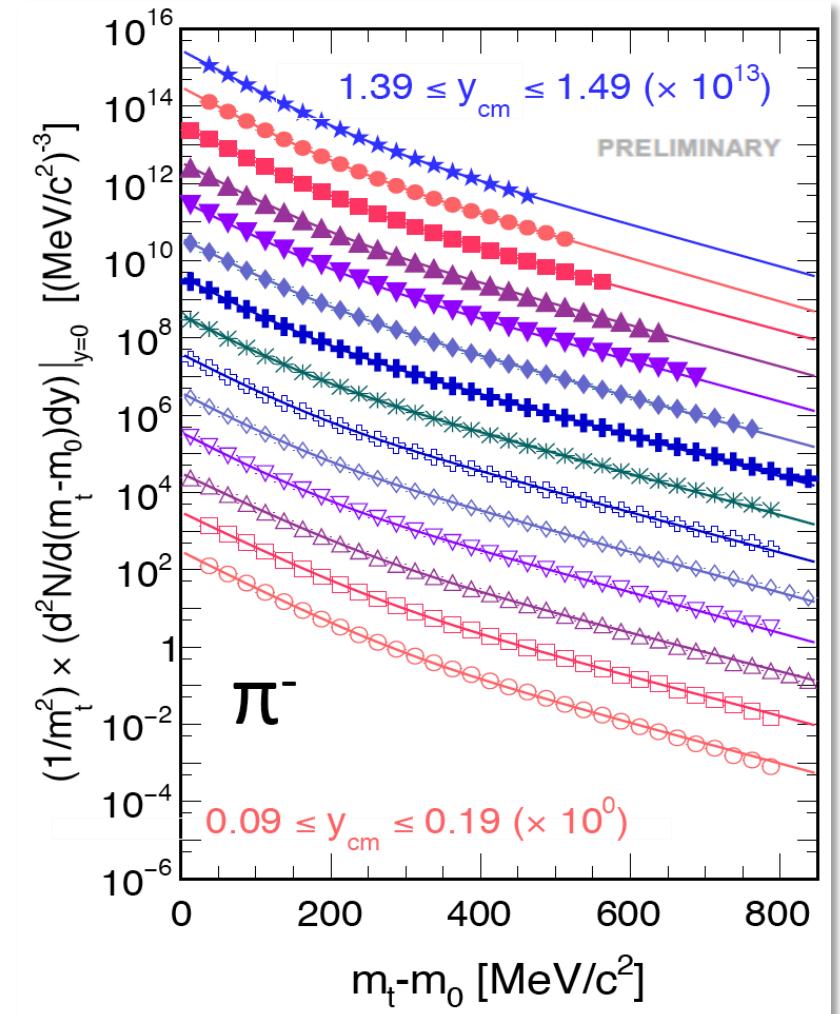
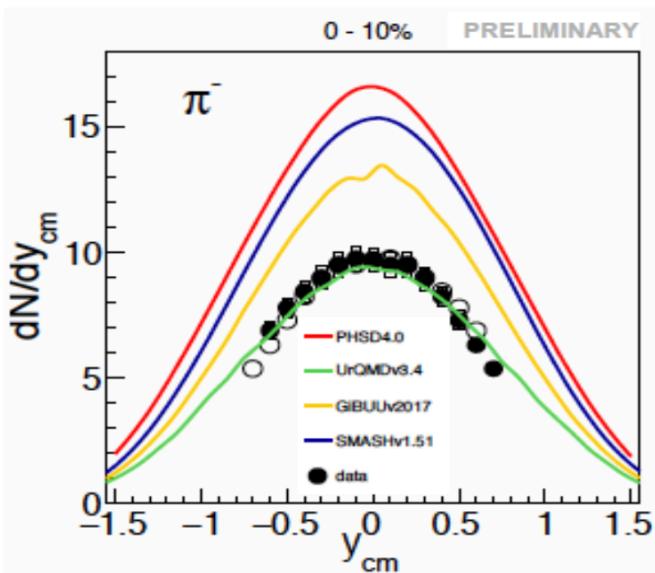
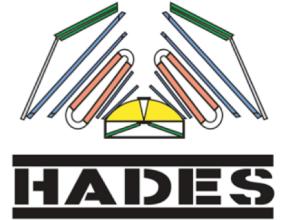
# Femtoscopy

## Radius Parameters



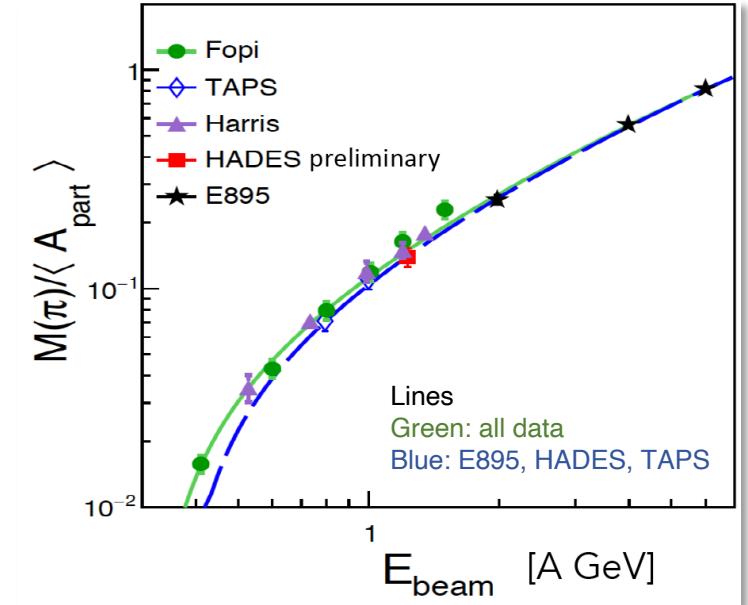
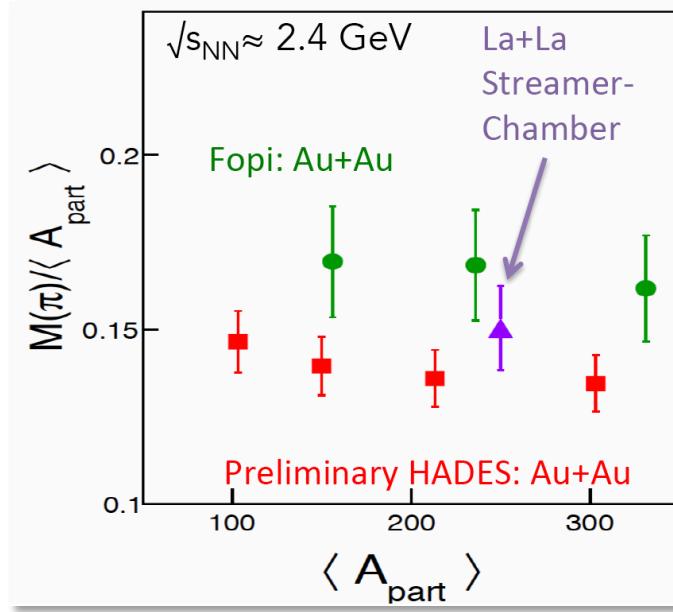
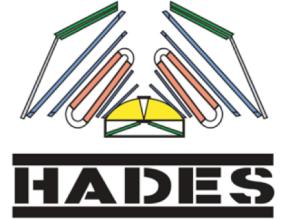
# Particle Production

## Pions



# Particle Production

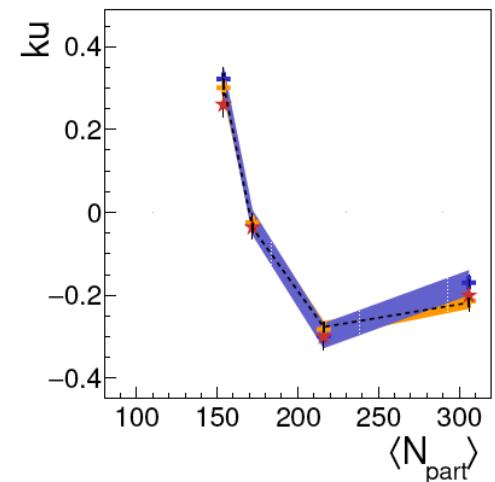
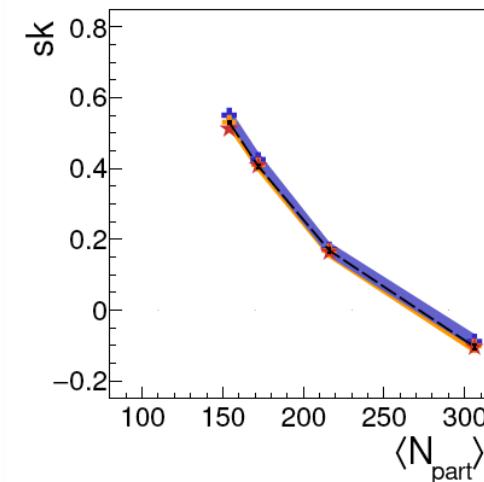
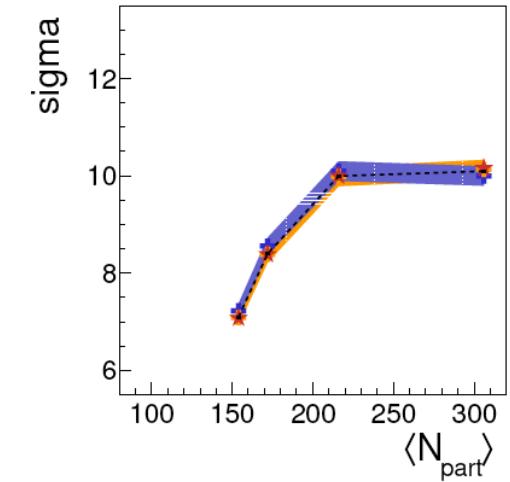
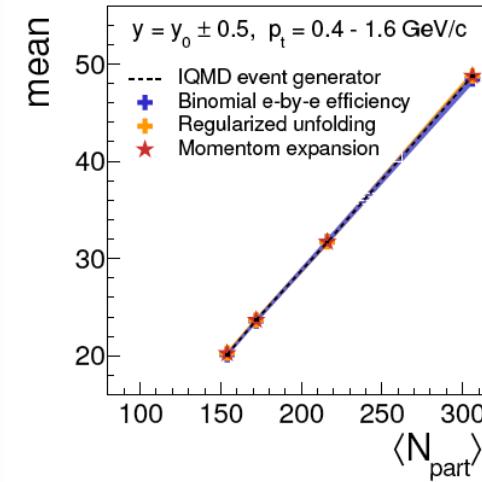
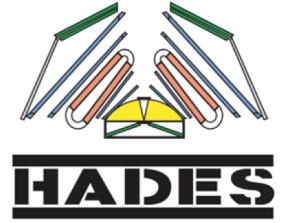
## Pions

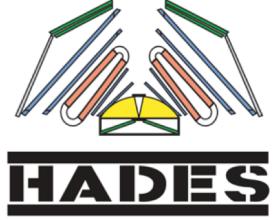


J.W. Harris et al., Phys. Rev. Lett. **58** (1987) 463  
R. Averbeck et al. [TAPS], Phys. Rev. **C67** (2003) 024903  
J.L. Klay et al. [E895], Phys. Rev. **C68** (2003) 054905  
W. Reisdorf et al. [FOPI], Nucl. Phys. **A781** (2007) 459

# Proton Number Fluctuations

## Efficiency Correction





# Proton Number Fluctuations

## NNLO Volume Corrections

---

$$\begin{aligned}
\tilde{\kappa}_1 &= \kappa_1 + v_2 \kappa'_1, \\
\tilde{\kappa}_2 &= \kappa_2 + \kappa_1^2 v_2 + \kappa'_2 v_2 + 2\kappa_1 \kappa'_1 V_2 + 2\kappa_1 \kappa'_1 v_3 + 2\kappa_1'^2 v_2 V_2 + \kappa_1'^2 V_1 V_2 + 2\kappa_1'^2 V_3 + \kappa_1'^2 v_4, \\
\tilde{\kappa}_3 &= \kappa_3 + \kappa_1^3 v_3 + 3\kappa_1 \kappa_2 v_2 + 3(\kappa_1 \kappa'_2 + \kappa'_1 \kappa_2) v_3 + 6\kappa'_1 (\kappa_1^2 + \kappa'_2) v_2 V_2 + 3\kappa'_1 (\kappa_1^2 + 2\kappa'_2) V_3 \\
&\quad + 3\kappa'_1 (\kappa_1^2 + \kappa'_2) v_4 + 12\kappa_1 \kappa_1'^2 V_2^2 + 3\kappa_1 \kappa_1'^2 V_1 V_3 + 24\kappa_1 \kappa_1'^2 v_2 V_3 + 6\kappa_1 \kappa_1'^2 V_4 + 3\kappa_1 \kappa_1'^2 v_5 \\
&\quad + 3(\kappa_1 \kappa'_2 + \kappa'_1 \kappa_2) V_2 + 8\kappa_1'^3 v_2 V_2^2 + 6\kappa_1'^3 V_1 V_2^2 + 10\kappa_1'^3 v_3 V_3 + \kappa_1'^3 V_1^2 V_3 + 24V_2 V_3 \kappa_1'^3 \\
&\quad + 3\kappa_1'^3 V_1 V_4 + 12\kappa_1'^3 v_2 V_4 + 3\kappa_1'^3 V_5 + \kappa_1'^3 v_6 + 3\kappa'_1 \kappa'_2 V_1 V_2 + \kappa'_3 v_2, \\
\tilde{\kappa}_4 &= \kappa_4 + \kappa_1^4 v_4 + 6\kappa_1^2 \kappa_2 v_3 + (4\kappa_1 \kappa_3 + 3\kappa_2^2) v_2 + 24(\kappa_1^3 \kappa'_1 + 4\kappa_1 \kappa'_1 \kappa'_2 + 2\kappa_1'^2 \kappa_2) v_2 V_3 \\
&\quad + 4(\kappa_1^3 \kappa'_1 + 6\kappa_1 \kappa'_1 \kappa'_2 + 3\kappa_1'^2 \kappa_2) V_4 + 2(2\kappa_1^3 \kappa'_1 + 6\kappa_1 \kappa'_1 \kappa'_2 + 3\kappa_1'^2 \kappa_2) v_5 \\
&\quad + 48(\kappa_1^2 \kappa_1'^2 + \kappa_1'^2 \kappa'_2) v_2 V_2^2 + 12(4\kappa_1^2 \kappa_1'^2 + 5\kappa_1'^2 \kappa'_2) v_3 V_3 + 72(\kappa_1^2 \kappa_1'^2 + 2\kappa_1'^2 \kappa'_2) V_2 V_3 \\
&\quad + 6(\kappa_1^2 \kappa_1'^2 + 3\kappa_1'^2 \kappa'_2) V_1 V_4 + 72(\kappa_1^2 \kappa_1'^2 + \kappa_1'^2 \kappa'_2) v_2 V_4 + 6(2\kappa_1^2 \kappa_1'^2 + 3\kappa_1'^2 \kappa'_2) V_5 \\
&\quad + 6(\kappa_1^2 \kappa_1'^2 + \kappa_1'^2 \kappa'_2) v_6 + 2(6\kappa_1^2 \kappa'_2 + 12\kappa_1 \kappa'_1 \kappa_2 + 4\kappa'_1 \kappa'_3 + 3\kappa_2'^2) v_2 V_2 \\
&\quad + 2(3\kappa_1^2 \kappa'_2 + 6\kappa_1 \kappa'_1 \kappa_2 + 4\kappa'_1 \kappa'_3 + 3\kappa_2'^2) V_3 + 2(3\kappa_1^2 \kappa_2 + 2\kappa_1 \kappa'_3 + 2\kappa'_1 \kappa_3 + 3\kappa_2 \kappa'_2) v_3 \\
&\quad + (6\kappa_1^2 \kappa'_2 + 12\kappa_1 \kappa'_1 \kappa_2 + 4\kappa'_1 \kappa'_3 + 3\kappa_2'^2) v_4 + 96\kappa_1 \kappa_1'^3 V_2^3 + 96\kappa_1 \kappa_1'^3 V_3^2 + 288\kappa_1 \kappa_1'^3 v_3 V_2^2 \\
&\quad + 72\kappa_1 \kappa_1'^3 V_1 V_2 V_3 + 4\kappa_1 \kappa_1'^3 V_1^2 V_4 + 144\kappa_1 \kappa_1'^3 V_2 V_4 + 128\kappa_1 \kappa_1'^3 v_3 V_4 + 12\kappa_1 \kappa_1'^3 V_1 V_5 \\
&\quad + 72\kappa_1 \kappa_1'^3 v_2 V_5 + 12\kappa_1 \kappa_1'^3 V_6 + 4\kappa_1 \kappa_1'^3 v_7 + 24(2\kappa_1 \kappa'_1 \kappa'_2 + \kappa_1'^2 \kappa_2) V_2^2 + 6(2\kappa_1 \kappa'_1 \kappa'_2 + \kappa_1'^2 \kappa_2) V_1 V_3 \\
&\quad + 2(2\kappa_1 \kappa'_3 + 2\kappa'_1 \kappa_3 + 3\kappa_2 \kappa'_2) V_2 + 48\kappa_1'^4 v_2 V_2^3 + 48\kappa_1'^4 V_1 V_2^3 + 48\kappa_1'^4 V_1 V_3^2 + 240\kappa_1'^4 v_2 V_3^2 \\
&\quad + 32\kappa_1'^4 v_4 V_4 + 288\kappa_1'^4 V_2^2 V_3 + 24\kappa_1'^4 V_1^2 V_2 V_3 + \kappa_1'^4 V_1^3 V_4 + 144\kappa_1'^4 v_4 V_2^2 + 72\kappa_1'^4 V_1 V_2 V_4 \\
&\quad + 128\kappa_1'^4 V_3 V_4 + 4\kappa_1'^4 V_1^2 V_5 + 72\kappa_1'^4 V_2 V_5 + 56\kappa_1'^4 v_3 V_5 + 6\kappa_1'^4 V_1 V_6 + 24V_2 V_6 \kappa_1'^4 v_2 V_6 + 4\kappa_1'^4 V_7 \\
&\quad + \kappa_1'^4 v_8 + 36\kappa_1'^2 \kappa'_2 V_1 V_2^2 + 6\kappa_1'^2 \kappa'_2 V_1^2 V_3 + 4\kappa'_1 \kappa'_3 V_1 V_2 + 3\kappa_2'^2 V_1 V_2 + \kappa'_4 v_2.
\end{aligned}$$

# Proton Number Fluctuations

## Reduced Proton Cumulants (Fully Corrected)

