

Positivity Bounds in EFTs (of Gravity)

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①

Positivity

$$\langle \Omega | \hat{\phi}^2(f) | \Omega \rangle \geq 0$$

aka unitarity

$$\hat{\phi}(f) = \int d^4x \phi(x) f(x)$$

'Unitarity' + assumptions of 'analyticity' $\equiv$ causality
'polynomial boundedness' $\equiv$ locality

$\Rightarrow$ constraints on signs and magnitudes of coefficients in a LEEFT.

All 'cosmological' theories of inflation + dark energy are non-renormalizable.

e.g. Higgs inflation

$$\begin{aligned} & \propto R H^2 \\ \text{cutoff } \Lambda^2 & \sim \frac{M_{\text{Pl}}^2}{\epsilon} \end{aligned}$$

DBI / $\mathbb{R}$ -inflation
/ pseudo-Goldstone...

$$\mathcal{L} = -\Lambda^4 p(X)$$

$$X \sim \frac{(\partial\phi)^2}{\Lambda^4} \quad (2)$$

Wilsonian Effective action

Top DOWN

$$e^{\frac{i}{\hbar} S_W[\text{Light}]}$$

$$= \int D\text{Heavy} e^{\frac{i}{\hbar} S_{UV}[\text{Light}, \text{Heavy}]}$$

Bottom UP

$$S_W[\text{Light}] = \Lambda^4 F\left(\frac{\rho_{\text{Born}}}{\Lambda}, \frac{F_{\text{Fermion}}}{\Lambda^{3/2}}, \frac{\partial \mu}{\Lambda}\right)$$

↖
'cutoff' or 'strong coupling scale'.

Weak or Strong

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We introduce the notion of a weakly coupled EFT

Weakly coupled = New 'classical' tree level physics enters at scales below scale at which loops $\sim$ trees.

$$S_w = \frac{1}{g_*^2} \int d^D x \Lambda^D F_0 \left[\frac{\text{Boson}}{\Lambda^{\frac{D-2}{2}}} , \frac{\text{Fermion}}{\Lambda^{\frac{D-1}{2}}} , \frac{2}{\Lambda} \right]$$

$$+ g_*^0 \int d^D x \Lambda^D F_1 [\dots] \quad (\text{one-loop})$$

$$+ g_*^2 \int d^D x \Lambda^D F_2 [\dots] \quad (\text{two-loop}).$$

$$+ g_*^{2(n-1)} \dots \dots \dots (n\text{-loops})$$

Example ① - String theory

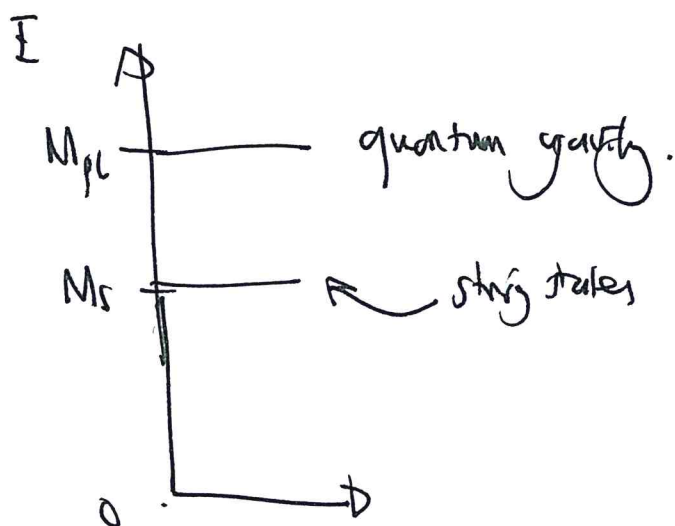
$$\Lambda = M_s$$

$$M_s^2 \sim \frac{1}{\alpha'}$$

$$g_* = g_s$$

$$S \sim \int d^{10}x \frac{M_s^8}{g_s^2} \left(R + \dots \alpha'^p R^{p+1} \dots \right)$$

$$M_{pl}^8 = \frac{M_s^8}{g_s^2}$$



Example ② - Kaluza-Klein theory

$$\int d^{4+N}x M_{4+N}^{2+N} R \sim \int \frac{M_{4+N}^{2+N}}{M_{KK}^N} R$$

$$M_{KK} \sim \frac{1}{L}$$

$\curvearrowright$ size of extra dimension

$$\frac{1}{g_*^2} = \left(\frac{M_{4+N}}{M_{KK}} \right)^N$$

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Strong coupling

$$g_* \sim 1$$

loops + trees are

of same order at cutoff

$$\text{eg } S = \int d^4x \frac{M_{\text{pl}}^2}{2} \left[R + \alpha \frac{R^2}{M_{\text{pl}}^2} + \beta \frac{R_{\mu\nu}^2}{M_{\text{pl}}^2} + \dots \right]$$

'traditional quantum gravity'.



Bottom up perspective - coefficients of local operators in LEEFT cannot be determined and are 'just matched' against observations.

CLAIM - this is not true entirely, positivity bounds impose conditions on 'signs' and 'magnitudes' of Wilson coefficients that imply not all LEEFTs admit UV completion.

Kallen-Lehman spectral rep

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Consider scalar operator $\hat{O}(x)$ in Minkowski.

Locality tells us

$$[\hat{O}(x), \hat{O}(y)] = 0, \quad (x-y)^2 > 0$$

Unitarity tells us that

$$\langle 4 | \hat{O}^\dagger \hat{O} | 4 \rangle \geq 0$$

$$\hat{O} | 4 \rangle = \int d^4x f(x) \hat{O}(x)$$

Together with Poincaré invariance these imply

$$\langle 0 | \hat{T} \hat{O}(x) \hat{O}(y) | 0 \rangle = \int \frac{d^4k}{(2\pi)^4} e^{ik \cdot (x-y)} G_0(k)$$

$$G_0(k) = \frac{Z_P}{k^2 + m^2 - i\epsilon} + \frac{S(k^2)}{N-1} + (-k^2)^N \int_{4m^2}^{\infty} d\mu \frac{\rho(\mu)}{\mu^N (k^2 + \mu - i\epsilon)}$$

$\rho(\mu) \geq 0$ is the KL spectral density.

$$S_{N-1}(k^2) = \sum_{r=0}^{N-1} \frac{C_r}{r!} (-k^2)^r$$

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ie. $\int \frac{d^d k}{(2\pi)^d} S_{N-1}(k^2) e^{ik \cdot (x-y)} = \sum_{r=0}^{N-1} \frac{C_r}{r!} \Box^r \delta^d(x-y)$

$N = \text{No. of subtractions}$

ie assumption is $p(\mu) \sim \mu^N$ (Polynomial boundedness)

Example - Conformal field theory

$$\langle \hat{\phi}(x) \hat{\phi}(y) \rangle = \frac{C_0}{(x^2 - y^2)^2}^\Delta \quad \Delta \text{ conformal dimension.}$$

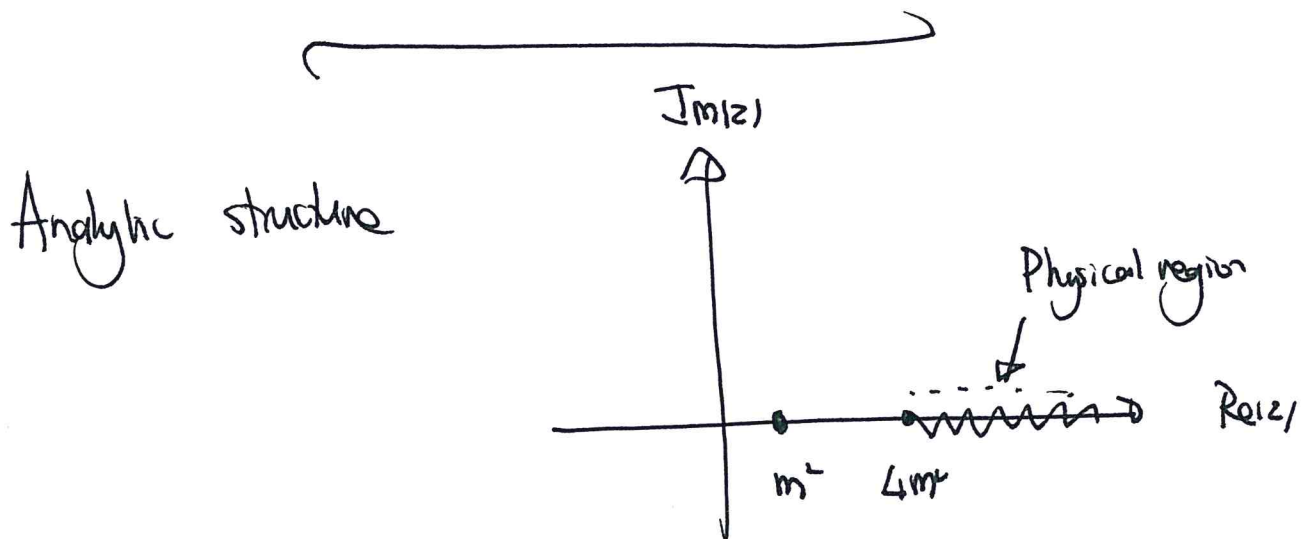
$$= \int \frac{d^d k}{(2\pi)^d} e^{ik \cdot (x-y)} \int_0^\infty d\mu \frac{\mu^{-\frac{d}{2} + \Delta}}{(k^2 + \mu - i\epsilon)}$$

$$\Rightarrow \int \frac{d^d k}{(2\pi)^d} e^{ik \cdot (x-y)} C (-k^2)^N \int_0^\infty d\mu \frac{\mu^{-\frac{d}{2} + \Delta - N}}{(k^2 + \mu - i\epsilon)}$$

So every time $\Delta = \frac{d}{2} + k$ for $k=0, 1, \dots$ we have to increase no of subtractions. (8)

$\therefore$ For any field theory with a UV fixed point

$$N_0 = \text{Integer } k \text{ just below } \left(\Delta - \frac{d}{2} \right).$$



$$G_0(z) = \frac{Z_p}{m^2 - z} + S_{N-1}(z) + z^N \int_{4m^2}^{\infty} d\mu \frac{\rho(\mu)}{\mu^N (\mu - z)}$$

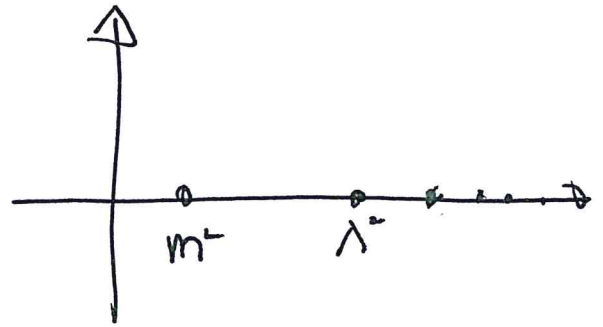
pole subtracted

$$\frac{1}{M!} \frac{d^M G_0}{dz^M} \Big|_{z=0} = \int_{4m^2}^{\infty} d\mu \frac{\rho(\mu)}{\mu^{M+1}} > 0$$

Assume weak coupling

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Branch cut $\Rightarrow$ sum of poles



$$\rho(\mu) = \sum_i g_i^2 \delta(\mu - m_i^2)$$

$$\Lambda^2 = \text{Min}(m_i^2)$$

What does this tell us about EFT 2.

$$S = \int d^4x \hat{\mathcal{O}}(x) \left[\square + a_1 \frac{\square^2}{\Lambda^2} + a_2 \frac{\square^3}{\Lambda^4} + \dots \right] \hat{\mathcal{O}}(x)$$

Tree level Feynman propagator is

$$G_0 = - \frac{1}{z + a_1 \frac{z^2}{\Lambda^2} + a_2 \frac{z^3}{\Lambda^4} + \dots}$$

$$G_0' = \frac{a_1}{\Lambda^2} + \frac{(a_2 - a_1^2)z}{\Lambda^4} + O(z^2) \quad (\text{assuming no subtraction})$$

Positivity bounds $a_1 > 0$ $a_2 > a_1^2$

More precisely

$$0 < a_2 - a_1^2 < a_1 \quad \text{since} \quad \int_{\Lambda^2}^0 \frac{\rho(\mu)}{\mu^2} < \frac{1}{\Lambda^2} \int_{\Lambda^2}^{\infty} \frac{\rho(\mu)}{\mu}$$

More generally ...

UV theory

$$G'_{0,UV}(z) = S_{N_0}(z) + z^{N_0} \int_{4m^2}^{\infty} d\mu \frac{\rho(\mu)}{\mu^{N_0}(\mu-z)}$$

LEFT

$$G_{0,LEFT}(z) \sim \frac{1}{z} \left(1 + \sum_m \frac{c_m}{\Lambda^{4m}} z (\ln(z))^m \dots \right)$$

$$G_{0,LEFT}(z) = S_{N_{EFT}}(z) + z^{N_{EFT}} \int_{4m^2}^{\infty} d\mu \frac{\rho(\mu)}{\mu^{N_{EFT}}(\mu-z)}$$

In general $N_{EFT} \gg N_0$

For $N_{EFT} > r \geq N_0$

$$\boxed{c_r = \int_{4m^2}^{\infty} d\mu \frac{\rho(\mu)}{\mu^r(\mu-z)} > \int_{4m^2}^{\infty} d\mu \frac{\rho(\mu)}{\mu^{r+1}} \quad (\epsilon \Lambda)}$$

What about gravity?

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eg extra dimensional models (Weakly coupled).
 $\Downarrow$
 4d gravity + ∞ tower of spin 2, 0 modes

$$S \approx \frac{M_p^2}{2} \int d^4x \sqrt{g_E} \left[R_E - \frac{1}{2} \sum_i (\nabla X_i)^2 - \frac{1}{2} m_{X_i}^2 (X_i^2 - X_i^2) \right. \\ \left. - \frac{1}{2} \sum_j (\nabla \phi_j)^2 - \frac{1}{2} m_{\phi_j}^2 \phi_j^2 \right. \\ \left. + \mathcal{L}_M[g_J, \chi_M] \right]$$

$$g_J \approx g_{E_J}^E + \sum_i \frac{\alpha_i}{\Lambda} X_{\mu\nu} + \sum_j \frac{\beta_j}{\Lambda} g_{\mu\nu}^E \chi_j$$

$$S = \int d^4x \sqrt{g_J} \frac{M_p^2}{2} \left[R_J + \frac{a}{\Lambda^2} \overbrace{(R_{\mu\nu}^J R^{\mu\nu}_J - \frac{1}{2} R_J^2)}^{\text{spin 2}} \right. \\ \left. + \frac{b}{\Lambda^2} R^2 \right] \quad \text{scalar part.}$$

1960's S-matrix assumptions

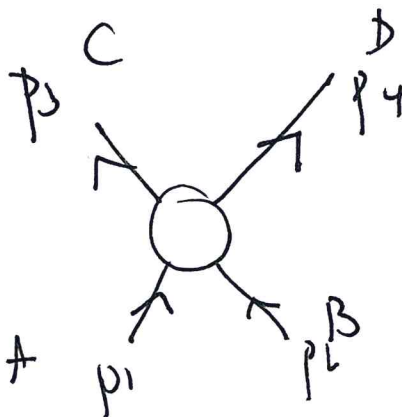
$$\hat{S}^\dagger \hat{S} = \hat{1}$$

1. Unitarity
2. Locality : Scattering amplitude polynomial (exponentially bounded) $|A(k)| < \alpha e^{\beta|k|}$
3. Causality : Analytic function of Mandelstam variables (modulo poles + cuts).

4. Poincaré Invariance

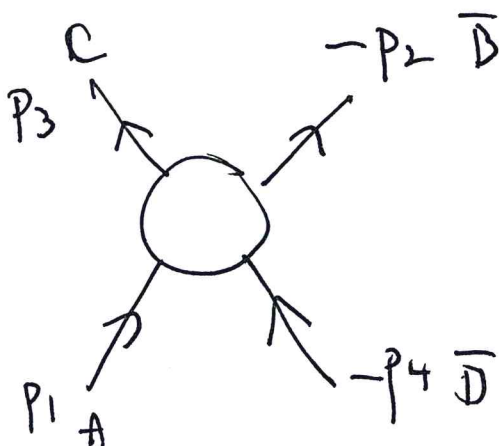
5. Crossing Symmetry : Follows from above

6. Mass-Gap :



s-channel

$$A + B \rightarrow C + D$$



u-channel

$$A + \bar{B} \rightarrow C + \bar{D}$$

$$s + t + u = 4m^2$$

$$s = (p_1 + p_2)^2$$

$$t = (p_1 - p_3)^2$$

$$u = (p_1 - p_4)^2$$

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Jin-Martin (generalization of Froissart).

$$-4m^2 < t < 4m^2 \quad |A(s, t)| < s^2$$

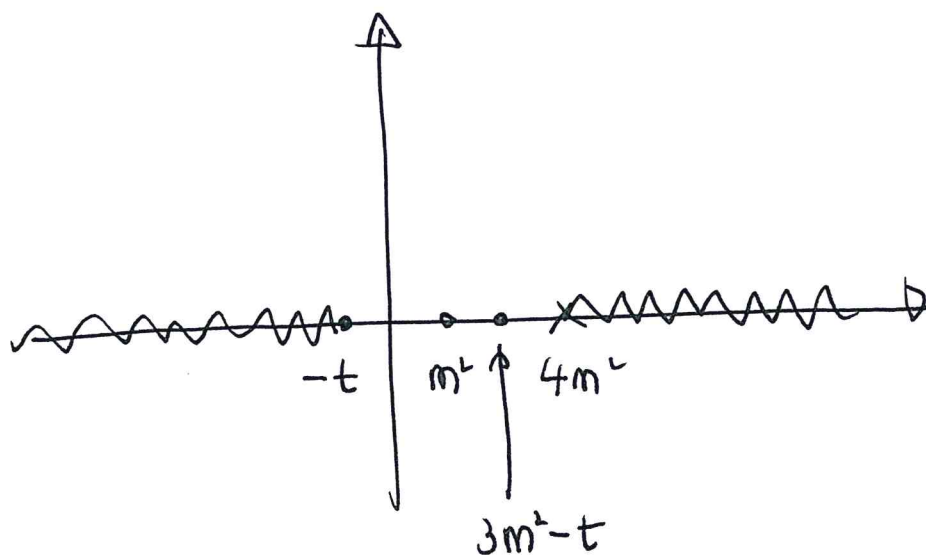
(usual Froissart $|A(s, 0)| < s \ln s$)

$$0 \leq t < 4m^2$$

$$\begin{aligned}
 A(s, t) = & a(t) + b(t) s \\
 & + \frac{s^2}{\pi} \int_{4m^2}^{\infty} d\mu \frac{\text{Im } A_{AB \rightarrow CD}(\mu, t)}{\mu^2 (\mu - s)} \quad \swarrow \text{s-channel} \\
 & + \frac{u^2}{\pi} \int_{4m^2}^{\infty} d\mu \frac{\text{Im } A_{A\bar{B} \rightarrow C\bar{D}}(\mu, t)}{\mu^2 (\mu - u)} \quad \swarrow \text{u-channel.}
 \end{aligned}$$

$$\left. \frac{\partial^n}{\partial t^n} \text{Im } A_{AB \rightarrow CD}(\mu, t) \right|_{t=0} > 0 \quad \forall n.$$

AND same for u-channel. $\left(\frac{\partial^n}{\partial t^n} (P_u(\cos \theta)) > 0 \right)$



Implications

$$\frac{\partial^{2+N}}{\partial s^{2+N}} A'_{AB \rightarrow \omega}(s, t) > 0 \quad \underline{0 < t < 4m^2}.$$

$$M = 2 + N$$

$$\frac{1}{M!} \frac{d^M}{ds^M} A'_{AB \rightarrow \omega}(s, t) \Big|_{s=2m^2} = \frac{1}{\pi} \int_{4m^2}^{\infty} \frac{\text{Im } A_S(\mu, t)}{(\mu - 2m^2 + t/L)^{M+1}} + \frac{1}{\pi} \int_{4m^2}^{\infty} \frac{\text{Im } A_U(\mu, t)}{(\mu - 2m^2 + t/L)^{M+1}} > 0.$$

$$t=0$$

$$S = -\frac{1}{2}(\phi)^2 + \frac{c}{\Lambda^4} (\phi)^2.$$

$$A \sim \frac{c}{\Lambda^4} (s^2 + t^2 + u^2) \Rightarrow c > 0$$

Non-forward light bounds

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define $f(s,t) = \frac{1}{2} \frac{\partial^2 A'(s,t)}{\partial s^2}$

We derive $f(s,t) > 0$

$$\frac{\partial f(s,t)}{\partial t} + \frac{3}{2\mu^2} f(s,t) > 0$$

$$\frac{1}{2} \frac{\partial^2 f(s,t)}{\partial t^2} + \frac{3}{2\mu^2} \left(\frac{\partial f(s,t)}{\partial t} + \frac{3}{2\mu^2} f(s,t) \right) > 0$$

$\vdots$
 $\checkmark$

where $\mu = 2m$ or Λ
(

Application to general spin

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Crossing complicated in helicity

Problems $\cos \theta = -\frac{2t}{s-4m^2}$
 Kinematic poles $s=4m^2$
 $\sqrt{s}u$ branch cuts
 For DF scatter $\sqrt{-s}u$ branch cuts

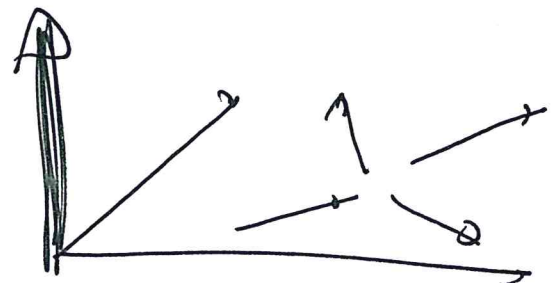
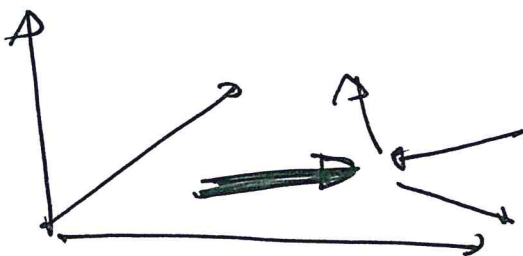
$$A_{\lambda_1 \lambda_2 \lambda_3 \lambda_4}^s(s, t) = \text{Phase} \times$$

$$\sum_{\lambda'_1 \lambda'_2 \lambda'_3 \lambda'_4} d_{\lambda'_1 \lambda_1}^{s_1}(\pi-x) d_{\lambda'_2 \lambda_2}^{s_2}(x) d_{\lambda'_3 \lambda_3}^{s_3}(x-\pi) d_{\lambda'_4 \lambda_4}^{s_4}(1-x)$$

Wigner J matrices

$$\times A_{\lambda'_1 \lambda'_4 \lambda'_3 \lambda'_2}^u(u, t)$$

$$\sin \chi = \frac{-2m\sqrt{t}}{\sqrt{(s-4m^2)(u-4m^2)}}$$



$$T_{T_1 T_2 T_3 T_4} = \sum_{\lambda_1 \lambda_2 \lambda_3 \lambda_4} U_{\lambda_1 T_1}^{s_1} U_{\lambda_2 T_2}^{s_2} U_{T_3 \lambda_3}^{s_3*} U_{T_4 \lambda_4}^{s_4*} A_{\lambda_1 \lambda_2 \lambda_3 \lambda_4}$$

$$T_{T_1 T_2 T_3 T_4}^s(s, t, u) = e^{-i \sum_i T_i \chi} T_{-T_1 -T_4 -T_2 -T_3}^u(u, t, s)$$

Regularized transversely amplitudes

$$\mathcal{T}_{T\pi LT_4}^+(s, \theta) = (\sqrt{s_u})^\varepsilon (s(s-4m^2))^{s_1+s_2} \\ \times (\mathcal{T}_{T\pi LT_4}(s, \theta) + \mathcal{T}_{T\pi LT_4}(s, -\theta))$$

def $N_s = 2 + \varepsilon + 2(s_1+s_2)$

$$f_{T\pi L}(s, t) = \frac{1}{N_s!} \frac{d^{N_s}}{ds^{N_s}} \mathcal{T}_{T\pi LT\pi L}^+(s, t)$$

$$s = u + 2m^2 - t/2$$

$$f_{T\pi L}(s, t) = \frac{1}{\pi} \int_{4m^2}^{\infty} d\mu \frac{\text{Abs}_s \mathcal{T}_{T\pi LT\pi L}^+(\mu, t)}{(\mu - 2m^2 + t/2 - u)^{N_s+1}} \\ + \frac{1}{\pi} \int_{4m^2}^{\infty} d\mu \frac{\text{Abs}_u \mathcal{T}_{T\pi LT\pi L}^+(\mu, t)}{(\mu - 2m^2 + t/2 + u)^{N_s+1}}$$

Application 1 - Weak Gravity conjecture.

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Bellazzini et al.

$$S = \int d^4x \sqrt{-g} \left[\frac{M_{pl}^2}{2} R - \frac{1}{4} F_{\mu\nu}^2 \right]$$

$$+ \frac{\alpha_1}{4M_{pl}^4} (F_{\mu\nu}^2)^2 + \frac{\alpha_2}{4M_{pl}^4} (\tilde{F}^{\mu\nu} F_{\mu\nu})^2$$

$$+ \frac{\alpha_3}{2M_{pl}^2} F_{\mu\nu} F_{\alpha\beta} W^{\mu\nu\alpha\beta}$$

Negl

$$\mathcal{A}_{4D} = \underbrace{\frac{-S^2}{M_{pl}^2 t}}_{\text{if removed}} - \frac{S}{M_{pl}^2} + \frac{2S^2(2\alpha_1 - \alpha_3)}{M_{pl}^4}$$

$$\Rightarrow (2\alpha_1 - \alpha_3) > 0$$

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$$\left(\frac{\sqrt{2}|Q|}{M/M_{Pl}} \right)_{\text{est}} = 1 + \frac{4}{5} \frac{(4\pi)^2 M_{Pl}^2 (2\alpha_4 - \alpha_3)}{M^2} > 1$$

Implies Mild form of weak gravity conjecture

(If no $Q > M$ since $M > \sum_i m_i$ $Q = \sum q_i$
 kinematically forbidden).

