

Holographic Vacua of $\mathcal{N} = 1^*$

Friðrik Freyr Gautason

Holography, Generalized Geometry and Duality
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based on [1805.03623 + 190x.xxxxx] with Nikolay Bobev, Benjamin Niehoff and
Jesse Van Muiden



KU LEUVEN

CONFINEMENT IN ADS/CFT

Not many explicit duals to confining $\mathcal{N} = 1$ theories.

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It would be great to have an explicit type IIB background dual to a deformation of $\mathcal{N} = 4$ that exhibits confinement in the IR and is regular (or mildly singular due to explicit branes).

$\mathcal{N} = 1^*$ Gauge theory

$\mathcal{N} = 4$ IN $\mathcal{N} = 1$ LANGUAGE

The $\mathcal{N} = 4$ vectormultiplet consists of

$$A_\mu, \quad X_I, \quad \psi_a,$$

where μ is a Lorentz index, $I \in \mathbf{6}$ and $a \in \mathbf{4}$ of $\mathfrak{su}(4) \simeq \mathfrak{so}(6)$. All transform in the adjoint of the gauge group $SU(N)$

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vector: $V = (A_\mu, \psi_4),$

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Only $\mathfrak{u}(1)_r \times \mathfrak{su}(3)_f \subset \mathfrak{su}(4)$ manifest. $i \in \mathbf{3}$ of $\mathfrak{su}(3)$. Kähler and super-potentials for $\mathcal{N} = 4$

$$K = \frac{1}{g_{\text{YM}}^2} \text{Tr } \Phi_i^\dagger \Phi_i, \quad W = \frac{1}{g_{\text{YM}}^2} \text{Tr } [\Phi_1, \Phi_2] \Phi_3.$$

$$\mathcal{N} = 1^*$$

Add mass terms

$$\delta W = \frac{1}{g_{\text{YM}}^2} \text{Tr} (m_1 \Phi_1^2 + m_2 \Phi_2^2 + m_3 \Phi_3^2) .$$

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Three special cases

- ❖ $m_1 = m_2 \neq 0, m_3 = 0 \Rightarrow \mathcal{N} = 2^*$,
- ❖ $m_1 = m_2 = 0, m_3 \neq 0 \Rightarrow$ Leigh-Strassler fixed point,
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Focus on the last case which displays rich vacuum structure.

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Classical vacuum equations

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Most of the time the IR theory will have unbroken $\mathrm{U}(1)$, i.e. free Electro-magnetism. "Coulomb vacuum".

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Quantum vacua can be labelled by N -dimensional sublattice in $\mathbf{Z}_N^e \times \mathbf{Z}_N^m$. These are generated by two elements

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Each element on the sublattice denotes a dyon that condenses in the vacuum. The dual lattice denotes dyons that are confined. Terminology: all vacua specified by x, y are *massive*

- ▶ $D = 1$: Higgs vacuum (classically massive).
- ▶ $D = N, b = 0$: Confining vacuum.
- ▶ all other : oblique confining vacuum.

OEIS[A000203]=1, 3, 4, 7, 6, 12, 8, 15, 13, 18, 12, 28, 14, 24, 24, 31,
18, 39, 20, 42, 32, 36, 24, 60, 31, 42, 40, 56, 30, 72, 32, 63, ...

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$$G_3 \sim m_1 d\bar{z}_1 \wedge dz_2 \wedge dz_3 + m_2 dz_1 \wedge d\bar{z}_2 \wedge dz_3 + m_3 dz_1 \wedge dz_2 \wedge d\bar{z}_3 ,$$

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m_i are proportional to the masses in the field theory. Probe D3-branes in this background “polarize” to spherical (p, q) -fivebranes as described by Myers

$$\partial V_{\text{D3}} = 0 \quad \Rightarrow \quad [\phi_i, \phi_j] + m \epsilon_{ij}^k \phi_k = 0 .$$

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PS argued that massive vacua are dual to a single stack of polarized fivebranes. Hard to find full non-linear asymptotically AdS_5 solutions of the IIB EOM. PS used linearized approximations to approach a solution.

5D SUPERGRAVITY

We use $\mathcal{N} = 8$ gauged supergravity in 5D. It contains all supergravity modes dual to the operators of the EM tensor multiplet of $\mathcal{N} = 4$ SYM.

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The 5D SUGRA has $\mathfrak{su}(4)$ *gauge* symmetry and among others 42 scalar fields that transform in

$$\mathbf{20}'_0 \oplus \mathbf{10}_{-2} \oplus \overline{\mathbf{10}}_2 \oplus \mathbf{1}_4 \oplus \mathbf{1}_{-4} ,$$

under $\mathfrak{su}(4) \times \mathfrak{u}(1)_Y$. The $\mathfrak{u}(1)_Y$ is a remnant of $\mathfrak{sl}(2, \mathbf{R})$ of type IIB.

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The invariant subsector is a $\mathcal{N} = 2$ supergravity with two hypermultiplets. The eight scalars parametrize the scalar manifold

$$\frac{G_{2(2)}}{\mathrm{SU}(2) \times \mathrm{SU}(2)} .$$

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scalar	$\mathfrak{su}(4)$ rep	dual operator
φ	$\mathbf{1}$	$\mathrm{Tr} F ^2$ and $\mathrm{Tr} (F \wedge F)$
ϕ_4	$\mathbf{10} \oplus \overline{\mathbf{10}}$	$\mathrm{Tr} (\psi_4 \psi_4)$
ϕ	$\mathbf{10} \oplus \overline{\mathbf{10}}$	$\mathrm{Tr} (\psi_i \psi_i)$
α	$\mathbf{20}'$	$\mathrm{Tr} (\phi_i \phi_i)$

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Two ways to truncate the theory by imposing discrete symmetries. Both lead to 4-scalar truncations. One of those can be further truncated using residual $u(1)$ symmetries.

$$\begin{aligned} \text{self-dual} & : \phi_4, \phi \rightarrow \text{Re } \phi_4, \text{Re } \phi, \\ \text{parity-invariant} & : \text{Re } \varphi, \text{Re } \phi_4, \text{Re } \phi, \text{Re } \alpha. \end{aligned}$$

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First of these was used by GPPZ to find a supersymmetric domain wall solution.

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The second was used to construct a Euclidean domain wall solution on S^4 .

Bobev, Elvang, Kol, Olson and Pufu (2016)

5D SUPERGRAVITY

Bobev, Elvang, Kol, Olson and Pufu (2016)

I will focus on the second truncation with metric-scalar sector

$$\mathcal{L} = \frac{1}{4\pi G_N} \sqrt{|g|} \left(\frac{1}{4} R + \frac{1}{2} \mathcal{K}_{i\bar{j}} \partial_\mu z^i \partial^\mu \bar{z}^{\bar{j}} - \mathcal{P} \right),$$

with

$$\mathcal{K}_{i\bar{j}} = \partial_i \partial_{\bar{j}} \mathcal{K}, \quad \mathcal{K} = -\log \left[(1 - z_1 \bar{z}_1) (1 - z_2 \bar{z}_2)^3 \right],$$

$$\mathcal{W} = \frac{3g}{4} (1 + z_1 z_2) (1 - z_2^2), \quad \mathcal{P} = \frac{1}{2} e^{\mathcal{K}} \left[\mathcal{K}^{i\bar{j}} D_i \mathcal{W} D_{\bar{j}} \bar{\mathcal{W}} - \frac{8}{3} \mathcal{W} \bar{\mathcal{W}} \right].$$

and

$$z_1 = \tanh \frac{1}{2} (3\alpha + \varphi - 3i\phi + i\phi_4),$$
$$z_2 = \tanh \frac{1}{2} (\alpha - \varphi - i\phi - i\phi_4).$$

BPS EQUATIONS ON $\mathbf{R}^{1,3}$ AND S^4

$$ds_5^2 = dr^2 + \mathcal{R}^2 e^{2A} d\Omega_4^2,$$

The BPS equations are

$$(A')^2 = \mathcal{R}^{-2} e^{-2A} + \frac{4}{9} e^{\mathcal{K}} \mathcal{W} \widetilde{\mathcal{W}},$$

$$(A' + \mathcal{R}^{-1} e^{-A})(z^i)' = -\frac{2}{3} e^{\mathcal{K}} \mathcal{W} \mathcal{K}^{i\bar{j}} D_{\bar{j}} \widetilde{\mathcal{W}},$$

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Flat space domain wall obtained by

$$\mathcal{R}^2 d\Omega_4^2 \rightarrow ds_{\mathbf{R}^{1,3}}^4, \quad \tilde{\cdot} \rightarrow \bar{\cdot}, \quad \mathcal{R} \rightarrow \infty.$$

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Gubser : $\mathcal{P} < \infty$, Maldacena-Nuñez : $-G_{tt}^{\text{Einst}} < \infty$.

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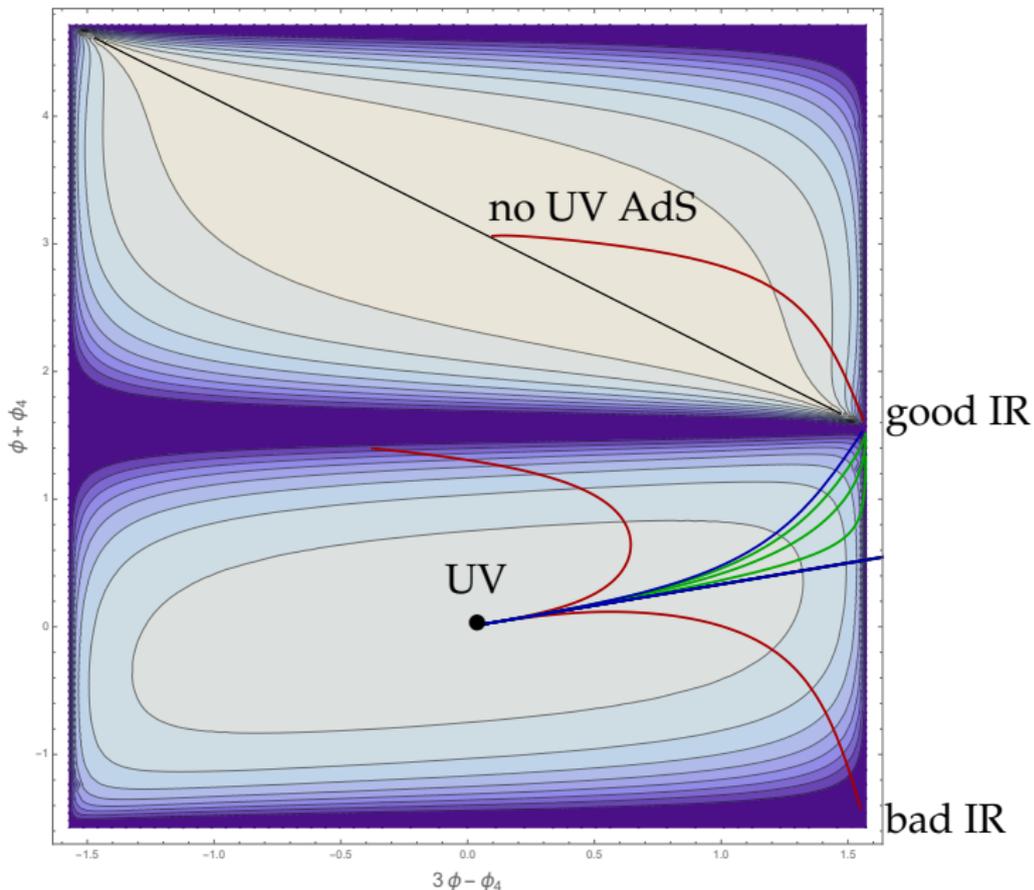
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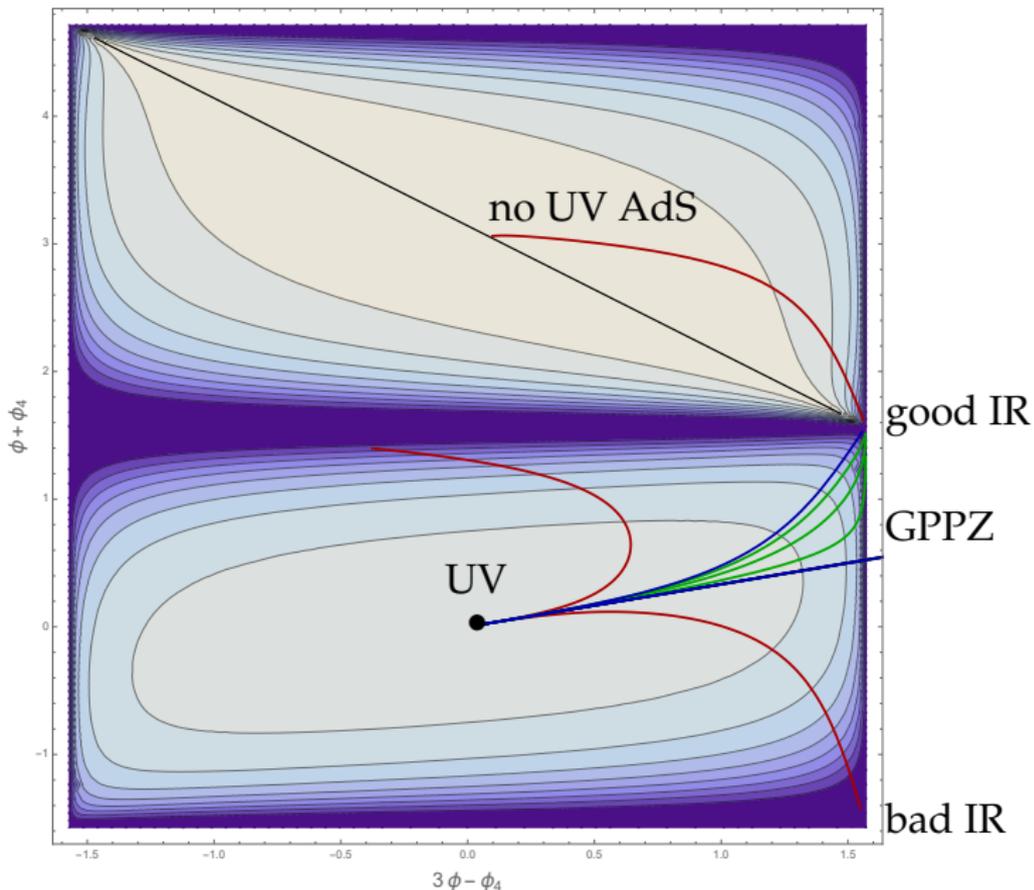
Maldacena and Nuñez (2000)

These agree in all cases I have checked.

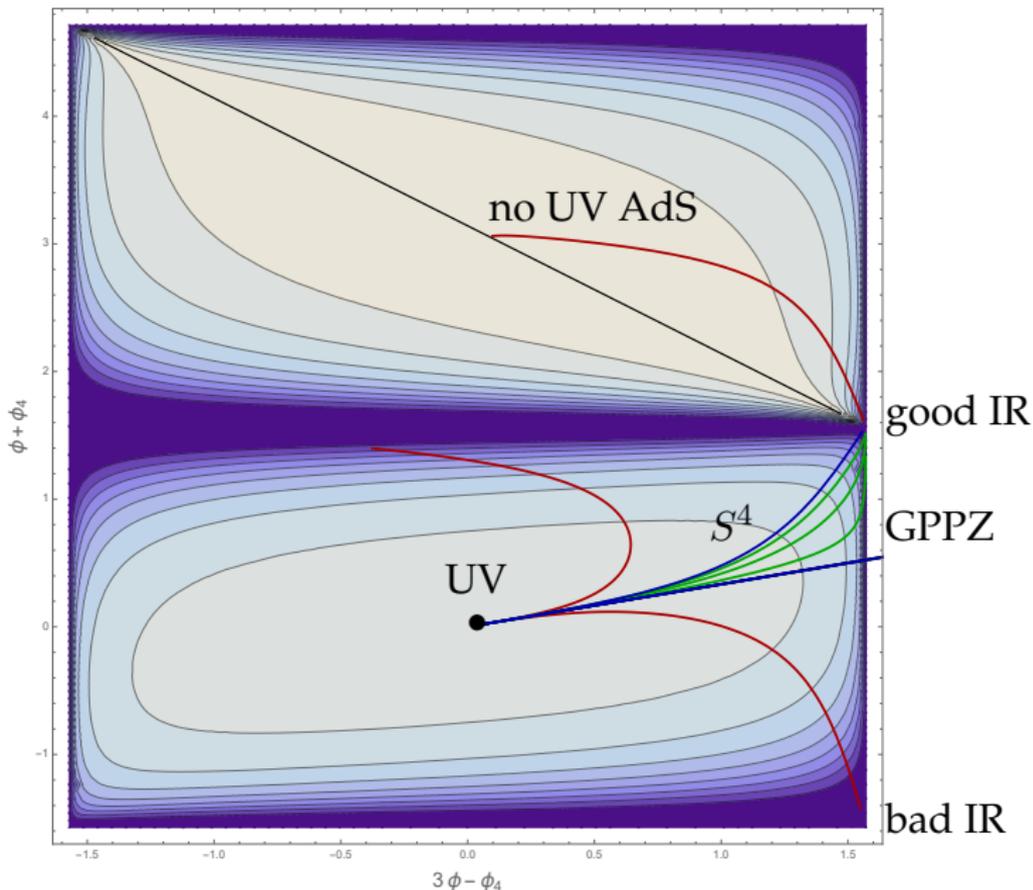
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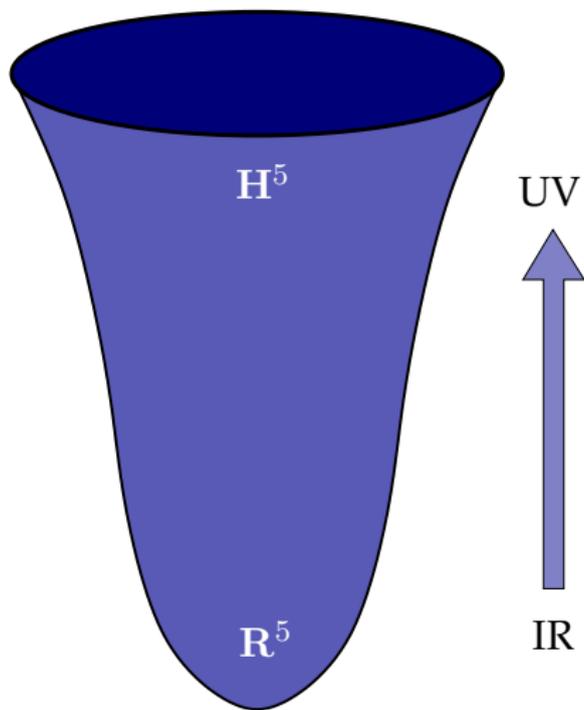
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UV-IR MATCHING



10D SOLUTION

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Partial uplift done by Pilch and Warner. Simultaneous uplift done by Petrini, Samtleben, Schmidt and Skenderis. We use different coordinates which are more suited for the near-singularity analysis.

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Bobev, FFG, Niehoff and van Muiden (2018)

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χ, α : angles on S^5 ,

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$\sigma_{1,2,3}$: left-invariant one-forms.

$$ds_{10}^2 = \frac{4(K_1 K_2 - K_3^2)^{1/4}}{g^2 \sqrt{g_s}} \left(\frac{dt^2 + (1-t^2)(1-\lambda^2 t^6)^{1/3} ds_4^2}{t^2(1-t)(1-\lambda^2 t^6)^{1/2}} + \frac{(1-\lambda^2 t^6)^{1/2}}{K_1 K_2 - K_3^2} d\Omega_5^2 \right)$$

$$\begin{aligned} d\Omega_5^2 = & K_4 d\chi^2 - 4\lambda t^4 (1-t^2)^2 (\cos 2\alpha d\chi - \sin 2\alpha \cos 2\chi \sigma_3)^2 \\ & + 4\lambda t^6 d(\cos 2\alpha \cos 2\chi)^2 + \frac{(1-\lambda^2 t^8)^2 (1-t^2)}{(1-\lambda^2 t^6)} (d\alpha + \sin 2\chi \sigma_3)^2 \\ & + \cos^2 2\chi (1+\lambda t^4)^2 (4t^2 d\alpha^2 + (1-t^2)^2 \sigma_3^2) \\ & + (1-t^2) (\sin^2 \chi K_1 \sigma_1^2 + \sin 2\chi K_3 \sigma_1 \sigma_2 + \cos^2 \chi K_2 \sigma_2^2) . \end{aligned}$$

$$e^\Phi = \frac{g_s(1+\lambda t^4)}{\sqrt{K_1 K_2 - K_3^2}} \left((1+t^2)(1-\lambda t^4) + 2t^2(1-\lambda t^2) \cos 2\chi \cos 2\alpha \right) ,$$

10D SOLUTION

$$C_0 = -\frac{2t^2(1+\lambda t^2)(1-\lambda t^4)\cos 2\chi \sin 2\alpha}{g_s(1+\lambda t^4)((1+t^2)(1-\lambda t^4)+2t^2(1-\lambda t^2)\cos 2\chi \cos 2\alpha)},$$

$$B_2 + ig_s C_2 = \frac{4}{g^2} \frac{te^{-i\alpha}}{K_1 K_2 - K_3^2} \\ \times \left[(a_1 d\chi + a_2 \sigma_3 - i(1-\lambda^2 t^8)(K_1 + K_2) \sin 2\chi d\alpha) \wedge \Sigma \right. \\ \left. - (a_3 d\chi + a_4 \sigma_3 - i(1-\lambda^2 t^8)(K_1 - K_2 - 2iK_3) \sin 2\chi d\alpha) \wedge \bar{\Sigma} \right],$$

$$F_5 = -\frac{1}{g^4 g_s} (1 + \star_{10}) d \left[\frac{(1-t^2)(1-\lambda^2 t^8)}{t^4(1-\lambda^2 t^6)^{1/3}} dx_0 \wedge dx_1 \wedge dx_2 \wedge dx_3 \right].$$

$$K_1 = (1+t^2)(1-\lambda^2 t^8) + 2t^2((1-\lambda^2 t^6) + \lambda t^2(1-t^2)\cos(4\alpha)) \cos 2\chi,$$

$$K_2 = (1+t^2)(1-\lambda^2 t^8) - 2t^2((1-\lambda^2 t^6) + \lambda t^2(1-t^2)\cos(4\alpha)) \cos 2\chi,$$

$$K_3 = 2\lambda t^4(1-t^2)\cos 2\chi \sin 4\alpha,$$

$$K_4 = (1+t^2)^2(1+\lambda t^4)^2 - 4t^4(1+\lambda t^2)^2 \cos^2 2\chi,$$

$$\Sigma = i \sin \chi \sigma_1 + \cos \chi \sigma_2,$$

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Solution parametrized by one integration constant $-1 \leq \lambda \leq 1$ that determines the gaugino vev:

$$\langle \psi_4 \psi_4 \rangle \sim N^2 m^3 \lambda .$$

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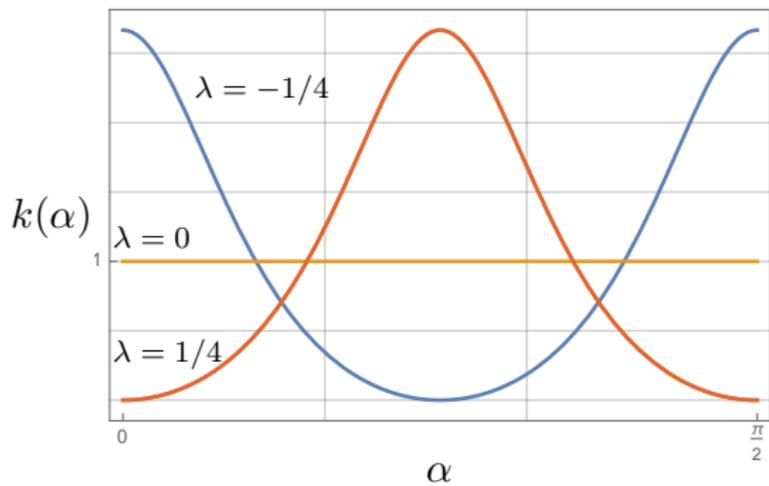
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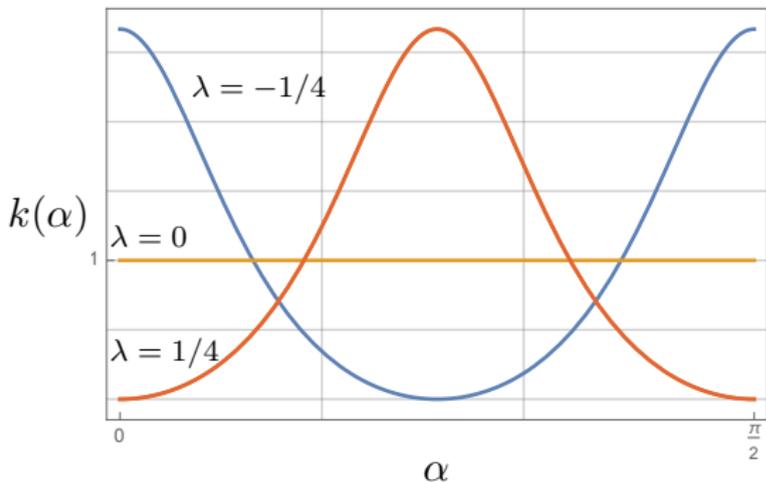
Tension of the (p, q) -fivebranes not constant

$$k(\alpha) = \frac{1 - \lambda^2}{1 + 2\lambda \cos 4\alpha + \lambda^2} , \quad \int k(\alpha) d\alpha = 2\pi .$$

(p, q) -FIVEBRANES



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As $|\lambda| \rightarrow 1$ we get δ -function peaks at $(1 + \lambda + 4n)\pi/8$,
 $n = 0, 1, 2, 3$.

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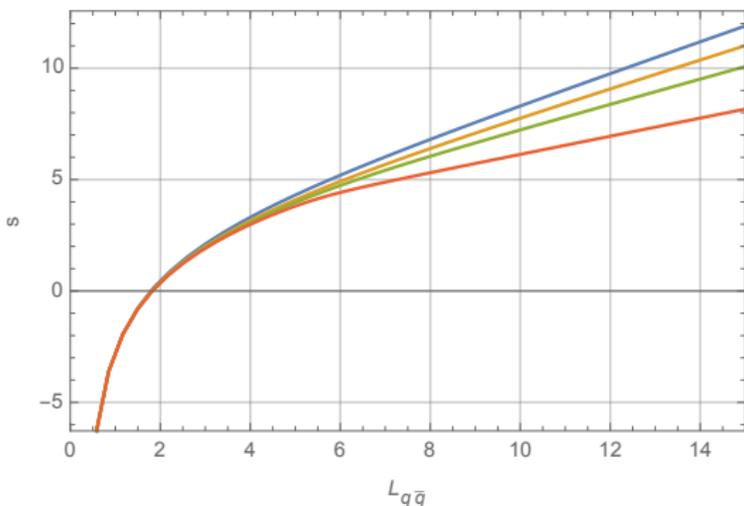
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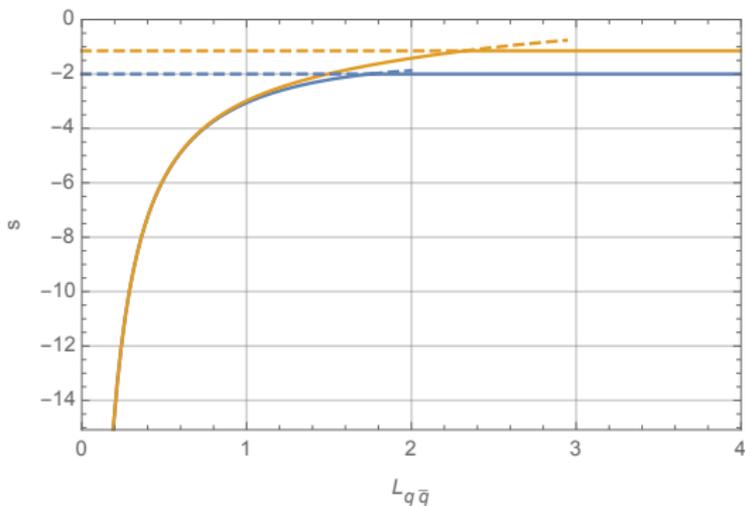
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This suggests that the limit $|\lambda| \rightarrow 1$ could be dual to a massive vacuum.

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It would be nice to generalize this story to include $20'$ vevs. Seems that we must include all eight scalars. Finally we would like to compare the Gaugino vev with computations in the field theory.

Dorey (1999)
Dorey and Kumar (2000)

Thank you