

TENSIONLESS STRINGS AND CARROLLIAN THINGS

BMS ON THE WORLDSHEET

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- ▶ Of interest to higher spin holographic dualities. [Klebanov-Polyakov '03, Gaberdiel-Gopakumar '10]
- ▶ Our programme: Initial goals.
 - Aim(1):** Understand string theory in this “ultra-stringy” regime.
 - Aim(2):** Make connection between tensionless strings and higher spins concrete.

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- ▶ Our programme: Initial goals.
 - Aim(1):** Understand string theory in this “ultra-stringy” regime.
 - Aim(2):** Make connection between tensionless strings and higher spins concrete.
- ▶ **Lacking:** A worldsheet organising principle (like 2d CFT for string theory).

SUMMARY OF RESULTS

References:

- ▶ A. Bagchi. JHEP 1305 (2013) 141 [arXiv:1303.0291].
- ▶ A. Bagchi, S. Chakraborty, P. Parekh. JHEP 1601 (2016) 158 [arXiv:1507.04361].
- ▶ A. Bagchi, S. Chakraborty, P. Parekh. JHEP 1610 (2016) 113 [arXiv:1606.09628].
- ▶ AB, Banerjee, Chakraborty, Parekh. JHEP 1802 (2018) 065 [arXiv:1710.03482].

Results:

- ▶ 3d *Bondi-Metzner-Sachs* algebra arises on the worldsheet as opposed to usual two copies of the Virasoro algebra.
- ▶ **Classical tensionless strings**: properties obtained intrinsically and as limit of relativistic strings and superstrings.
- ▶ **Quantum tensionless strings**: limiting & intrinsic description differ. New physics!
- ▶ Interesting connections to strings near the Hagedorn transition.

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CLASSICAL CLOSED STRINGS

Isberg, Lindstrom, Sundborg, Theodoridis 1993

Start with Nambu-Goto action:

$$S = -T \int d^2\xi \sqrt{-\det \gamma_{\alpha\beta}}. \quad (1)$$

To take the tensionless limit, first switch to Hamiltonian framework.

- ▶ **Generalised momenta:** $P_m = T \sqrt{-\gamma} \gamma^{0\alpha} \partial_\alpha X_m$.
- ▶ **Constraints:** $P^2 + T^2 \gamma \gamma^{00} = 0$, $P_m \partial_\sigma X^m = 0$.
- ▶ **Hamiltonian:** $\mathcal{H}_T = \mathcal{H}_C + \rho^i (\text{constraints})_i = \lambda (P^2 + T^2 \gamma \gamma^{00}) + \rho P_m \partial_\sigma X^m$.

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Action after integrating out momenta:

$$S = \frac{1}{2} \int d^2\xi \frac{1}{2\lambda} \left[\dot{X}^2 - 2\rho\dot{X}^m\partial_\sigma X_m + \rho^2\partial_\sigma X^m\partial_\sigma X_m - 4\lambda^2 T^2\gamma\gamma^{00} \right] \quad (2)$$

Identifying

$$g^{\alpha\beta} = \begin{pmatrix} -1 & \\ \rho & -\rho^2 + 4\lambda^2 T^2 \end{pmatrix},$$

action takes the familiar Weyl-invariant form

$$S = -\frac{T}{2} \int d^2\xi \sqrt{-g} g^{\alpha\beta} \partial_\alpha X^m \partial_\beta X^n \eta_{mn}. \quad (3)$$

CLASSICAL TENSIONLESS CLOSED STRINGS

Isberg, Lindstrom, Sundborg, Theodoridis 1993

- ▶ Tensionless limit can now be taken systematically.
- ▶ $T \rightarrow 0 \Rightarrow$

$$g^{\alpha\beta} = \begin{pmatrix} -1 & \rho \\ \rho & -\rho^2 \end{pmatrix}.$$

- ▶ Metric is degenerate. $\det g = 0$.
- ▶ Replace degenerate metric density $T\sqrt{-g}g^{\alpha\beta}$ by a rank-1 matrix $V^\alpha V^\beta$ where V^α is a vector density

$$V^\alpha \equiv \frac{1}{\sqrt{2\lambda}}(1, \rho) \quad (4)$$

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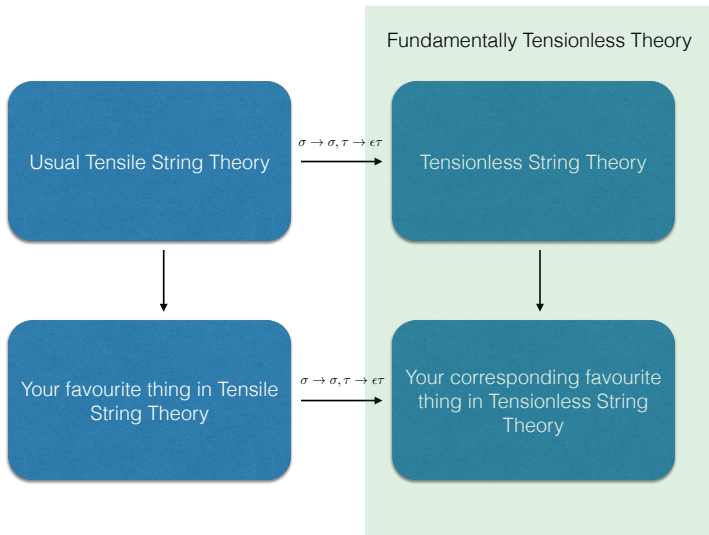
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- ▶ Action in $T \rightarrow 0$ limit

$$S = \int d^2\xi V^\alpha V^\beta \partial_\alpha X^m \partial_\beta X^n \eta_{mn}. \quad (5)$$

- ▶ Starting point of tensionless strings.
- ▶ Need not refer to any parent theory. Treat this as action of fundamental objects.

CLOSURE?



SYMMETRIES OF TENSIONLESS CLOSED STRINGS

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- ▶ Tensionless action is invariant under world-sheet diffeomorphisms.
- ▶ **Fixing gauge:** "Conformal" gauge: $V^\alpha = (v, 0)$ (v : constant).
- ▶ **Tensile:** Residual symmetry after fixing conformal gauge = $\text{Vir} \otimes \text{Vir}$.
Central to understanding string theory.
- ▶ **Tensionless:** Similar residual symmetry left over after gauge fixing.

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- ▶ **Tensionless:** Similar residual symmetry left over after gauge fixing.
- ▶ For world-sheet diffeomorphism: $\xi^\alpha \rightarrow \xi^\alpha + \varepsilon^\alpha$, change in vector density:

$$\delta_\varepsilon V^\alpha = -V \cdot \partial \varepsilon^\alpha + \varepsilon \cdot \partial V^\alpha + \frac{1}{2}(\partial \cdot \varepsilon)V^\alpha$$

- ▶ Tensionless residual symmetries: for $V^\alpha = (v, 0)$,

$$\varepsilon^\alpha = \{f'(\sigma)\tau + g(\sigma), f(\sigma)\} \quad (6)$$

SYMMETRIES OF TENSIONLESS CLOSED STRINGS

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- ▶ Define: $L(f) = f'(\sigma)\tau\partial_\tau + f(\sigma)\partial_\sigma$, $M(g) = g(\sigma)\partial_\tau$.
- ▶ Expand: $f = \sum a_n e^{in\sigma}$, $g = \sum b_n e^{in\sigma}$
- ▶ Therefore we have:

$$L(f) = \sum_n a_n e^{in\sigma} (\partial_\sigma + in\tau\partial_\tau) = \sum_n a_n L_n, \quad (7)$$

$$M(g) = \sum_n b_n e^{in\sigma} \partial_\tau = \sum_n b_n M_n. \quad (8)$$

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- ▶ Symmetry algebra in terms of Fourier modes:

$$\begin{aligned} [L_m, L_n] &= (m-n)L_{m+n} + \frac{c_L}{12}(m^3 - m)\delta_{m+n,0}, & [M_m, M_n] &= 0. \\ [L_m, M_n] &= (m-n)M_{m+n} + \frac{c_M}{12}(m^3 - m)\delta_{m+n,0}. \end{aligned} \quad (9)$$

Isberg et al find $c_L = c_M = 0$.

- ▶ [3d Bondi-Metzner-Sachs algebra](#) or [2d Galilean Conformal Algebra](#).
- ▶ Various other applications: Holography of 3d flat space, Galilean field theories, non-relativistic limit of AdS/CFT.

TENSIONLESS STRINGS: SYMMETRIES AS A LIMIT

A Bagchi 2013

- ▶ **Tensile string:** Residual symmetry in conformal gauge $g_{\alpha\beta} = e^{\phi}\eta_{\alpha\beta}$:

$$\begin{aligned} [\mathcal{L}_m, \mathcal{L}_n] &= (m-n)\mathcal{L}_{m+n} + \frac{c}{12}m(m^2-1)\delta_{m+n,0} \\ [\mathcal{L}_m, \bar{\mathcal{L}}_n] &= 0, \quad [\bar{\mathcal{L}}_m, \bar{\mathcal{L}}_n] = (m-n)\bar{\mathcal{L}}_{m+n} + \frac{\bar{c}}{12}m(m^2-1)\delta_{m+n,0} \end{aligned} \quad (10)$$

- ▶ World-sheet is a cylinder. Symmetry best expressed as 2d conformal generators on the cylinder.

$$\mathcal{L}_n = ie^{in\omega} \partial_{\omega}, \quad \bar{\mathcal{L}}_n = ie^{in\bar{\omega}} \partial_{\bar{\omega}} \quad (11)$$

where $\omega, \bar{\omega} = \tau \pm \sigma$. Vector fields generate centre-less Virasoros.

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- ▶ **Tensionless limit** \Rightarrow length of string becomes infinite ($\sigma \rightarrow \infty$).
- ▶ Ends of closed string identified \Rightarrow limit best viewed as ($\sigma \rightarrow \sigma, \tau \rightarrow \epsilon\tau, \epsilon \rightarrow 0$).

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- ▶ Ends of closed string identified \Rightarrow limit best viewed as ($\sigma \rightarrow \sigma, \tau \rightarrow \epsilon\tau, \epsilon \rightarrow 0$).
- ▶ Define

$$L_n = \mathcal{L}_n - \bar{\mathcal{L}}_{-n}, \quad M_n = \epsilon(\mathcal{L}_n + \bar{\mathcal{L}}_{-n}). \quad (12)$$

- ▶ New vector fields (L_n, M_n) well-defined in limit and given by:

$$L_n = ie^{in\sigma}(\partial_\sigma + in\tau\partial_\tau), \quad M_n = ie^{in\sigma}\partial_\tau. \quad (13)$$

- ▶ These are *exactly the generators defined previously*. Close to form BMS_3 .

TENSIONLESS STRINGS AND CARROLLIAN STRUCTURES

- ▶ Tensionless limit on the worldsheet: $\sigma \rightarrow \sigma, \tau \rightarrow \epsilon\tau, \epsilon \rightarrow 0$
- ▶ Worldsheet velocities $v = \frac{\sigma}{\tau} \rightarrow \infty$. Effectively, $\frac{v}{c} \rightarrow \infty$
- ▶ Hence worldsheet speed of light $\rightarrow 0$. Carrollian limit.

- ▶ Degenerate worldsheet metric.
- ▶ Riemannian tensile worldsheet \rightarrow Carrollian tensionless worldsheet.

- ▶ Action for tensionless string \Rightarrow a massless spin-0 particle coupled to a Carrollian background. Compare, e.g. with [\[Bergshoeff, Gomis, Rollier, Rosseel, Veldhuis 2017\]](#)
- ▶ BMS symmetries are conformal Carroll symmetries.
- ▶ Very clear why the BMS_3 algebra appears here.

- ▶ Similar discussions: [\[Duval, Gibbons, Horvathy 2014\]](#)

TENSIONLESS EM-TENSOR

A Bagchi 2013

- ▶ Spectrum of tensile string theory (in conformal gauge in flat space)
 - ▶ **Quantise** worldsheet theory as a theory free scalar fields.
 - ▶ **Constraint**: vanishing of EOM of metric (which is fixed to be flat).
 - ▶ **Op form**: Physical states vanish under action of modes of E-M tensor.
- ▶ EM tensor for 2d CFT on cylinder:

$$T_{cyl} = z^2 T_{plane} - \frac{c}{24} = \sum_n \mathcal{L}_n e^{in\omega} - \frac{c}{24}; \quad \bar{T}_{cyl} = \sum_n \bar{\mathcal{L}}_n e^{in\bar{\omega}} - \frac{\bar{c}}{24} \quad (14)$$

- ▶ **The Ultra-relativistic EM tensor**

$$T_{(1)} = \lim_{\epsilon \rightarrow 0} \left(T_{cyl} - \bar{T}_{cyl} \right) = \sum_n (L_n - in\tau M_n) e^{in\sigma} - \frac{c_L}{24} \quad (15)$$

$$T_{(2)} = \lim_{\epsilon \rightarrow 0} \epsilon \left(T_{cyl} + \bar{T}_{cyl} \right) = \sum_n M_n e^{in\sigma} - \frac{c_M}{24} \quad (16)$$

- ▶ **Classical constraint** on the tensionless string: $T_{(1)} = 0$, $T_{(2)} = 0$.
- ▶ Quantum version: **physical spectrum of tensionless strings** restricted by

$$\langle \text{phys} | T_{(1)} | \text{phys}' \rangle = 0, \quad \langle \text{phys} | T_{(2)} | \text{phys}' \rangle = 0. \quad (17)$$

INTRINSIC ANALYSIS: EOM AND SOLUTIONS

AB, Chakraborty, Parekh 2015

- ▶ **Equation of motion** in $V^a = (v, 0)$ gauge: $\ddot{X}^\mu = 0$.
- ▶ **Solution:**

$$X^\mu(\sigma, \tau) = x^\mu + \sqrt{2c'} A_0^\mu \sigma + \sqrt{2c'} B_0^\mu \tau + i\sqrt{2c'} \sum_{n \neq 0} \frac{1}{n} (A_n^\mu - in\tau B_n^\mu) e^{in\sigma} \quad (18)$$

- ▶ **Closed string b.c.:** $X^\mu(\sigma, \tau) = X^\mu(\sigma + 2\pi, \tau) \Rightarrow A_0^\mu = 0$.
- ▶ **Constraints:**

$$\dot{X}^2 = 2c' \sum_{m,n} B_{-m} \cdot B_{m+n} e^{in\sigma} = 0, \quad \dot{X} \cdot X' = 2c' \sum_{m,n} (A_{-m} - in\tau B_{-m}) \cdot B_{m+n} e^{in\sigma} = 0$$

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- ▶ **Define:**

$$L_n = \sum_m A_{-m} \cdot B_{m+n}, \quad M_n = \sum_m B_{-m} \cdot B_{m+n}$$

- ▶ **Classical constraints in terms of modes:**

$$\sum_n (L_n - in\tau M_n) e^{in\sigma} = 0 = T_{(1)}, \quad \sum_n M_n e^{in\sigma} = 0 = T_{(2)}. \quad (19)$$

- ▶ **Familiar form obtained earlier from purely algebraic considerations.**

INTRINSIC ANALYSIS: EOM AND SOLUTIONS

AB, Chakraborty, Parekh 2015

- ▶ The algebra of the modes are:

$$\{A_m^\mu, A_n^\nu\} = 0, \quad \{B_m^\mu, B_n^\nu\} = 0, \quad \{A_m^\mu, B_n^\nu\} = -im\delta_{m+n,0} \eta^{\mu\nu}. \quad (20)$$

Note: this is *not* the algebra of harmonic oscillator modes. (More later.)

- ▶ The worldsheet symmetry algebra of tensionless strings, now constructed from the quadratics of the modes:

$$\{L_m, L_n\} = -i(m-n)L_{m+n}, \quad \{L_m, M_n\} = -i(m-n)M_{m+n}, \quad \{M_m, M_n\} = 0. \quad (21)$$

- ▶ Quantization: $\{, \}_{PB} \rightarrow -\frac{i}{\hbar} [,]$ leads to the 2d Galilean Conformal Algebra.

LIMITING ANALYSIS: MODES

AB, Chakraborty, Parekh 2015

- ▶ Tensile string mode expansion:

$$X^\mu(\sigma, \tau) = x^\mu + 2\sqrt{2\alpha'}\alpha_0^\mu\tau + i\sqrt{2\alpha'}\sum_{n\neq 0}\frac{1}{n}[\tilde{\alpha}_n^\mu e^{-in(\tau+\sigma)} + \alpha_n^\mu e^{-in(\tau-\sigma)}].$$

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- The limiting procedure: $\tau \rightarrow \epsilon\tau$, $\sigma \rightarrow \sigma$, $\alpha' = c'/\epsilon$ with $\epsilon \rightarrow 0$

$$\begin{aligned} X^\mu(\sigma, \tau) &= x^\mu + 2\sqrt{\frac{2c'}{\epsilon}}\alpha_0^\mu\epsilon\tau + i\sqrt{\frac{2c'}{\epsilon}}\sum_{n\neq 0}\frac{1}{n}[\tilde{\alpha}_n^\mu e^{-in\sigma}(1 - in\epsilon\tau) + \alpha_n^\mu e^{in\sigma}(1 - in\epsilon\tau)], \\ &= x^\mu + 2\sqrt{2c'}(\sqrt{\epsilon})\alpha_0^\mu\tau + i\sqrt{2c'}\sum_{n\neq 0}\frac{1}{n}\left[\frac{\alpha_n^\mu - \tilde{\alpha}_{-n}^\mu}{\sqrt{\epsilon}} - in\tau\sqrt{\epsilon}(\alpha_n^\mu + \tilde{\alpha}_{-n}^\mu)\right]e^{in\sigma}. \end{aligned}$$

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- ▶ Thus we get a relation between the tensionless and tensile modes:

$$A_n^\mu = \frac{1}{\sqrt{\epsilon}}(\alpha_n^\mu - \tilde{\alpha}_{-n}^\mu), \quad B_n^\mu = \sqrt{\epsilon}(\alpha_n^\mu + \tilde{\alpha}_{-n}^\mu). \quad (22)$$

- ▶ The equivalent of the Virasoro constraints are now related as:

$$L_n = \mathcal{L}_n - \bar{\mathcal{L}}_{-n}, \quad M_n = \epsilon[\mathcal{L}_n + \bar{\mathcal{L}}_{-n}] \quad (23)$$

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QUANTUM TENSIONLESS STRINGS FROM THE LIMIT

AB, Chakraborty, Parekh 2015

- ▶ Think of Tensionless Hilbert space as one obtained from the tensile theory. Metric degenerates. So constraints on the system change.
- ▶ Impose Tensionless Constraints on the Hilbert space of the Tensile theory.
- ▶ Physical Hilbert space of tensile theory: Built on the tensile vacuum.

$$|0\rangle_\alpha : \quad \alpha_n^\mu |0\rangle_\alpha = 0 = \tilde{\alpha}_n^\mu |0\rangle_\alpha \quad \forall n > 0. \quad (24)$$

- ▶ The states at a general level: $|\Phi\rangle = \prod_{i,j} \alpha_{-m_i}^{\mu_i} \tilde{\alpha}_{-m_j}^{\mu_j} |0\rangle_\alpha$

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- ▶ The states at a general level: $|\Phi\rangle = \prod_{i,j} \alpha_{-m_i}^{\mu_i} \tilde{\alpha}_{-m_j}^{\mu_j} |0\rangle_\alpha$
- ▶ The mass² operator acting on this state:

$$m^2 = \frac{2}{c'} \left(\sum_{m \neq 0} B_{-m} \cdot B_m \right) \prod_{i,j} \alpha_{-m_i}^{\mu_i} \tilde{\alpha}_{-m_j}^{\mu_j} |0\rangle_\alpha$$

$$\Rightarrow m^2 = \frac{2\epsilon}{c'} \left(\sum_{m \neq 0} (\alpha_{-m} + \tilde{\alpha}_m) \cdot (\alpha_m + \tilde{\alpha}_{-m}) \right) \prod_{i,j} \alpha_{-m_i}^{\mu_i} \tilde{\alpha}_{-m_j}^{\mu_j} |0\rangle_\alpha \rightarrow 0$$

- ▶ **Massless higher spin states** (rather trivially).

FUNDAMENTALLY QUANTUM TENSIONLESS STRINGS

AB, Chakraborty, Parekh 2015

- ▶ Wish to return to harmonic oscillator basis for the tensionless string. Define:

$$C_n^\mu = \frac{1}{2}(A_n^\mu + B_n^\mu), \quad \tilde{C}_n^\mu = \frac{1}{2}(-A_{-n}^\mu + B_{-n}^\mu) \quad (25)$$

- ▶ The algebra: $[C_m^\mu, C_n^\nu] = m\delta_{m+n}\eta^{\mu\nu}$, $[\tilde{C}_m^\mu, \tilde{C}_n^\nu] = m\delta_{m+n}\eta^{\mu\nu}$.
- ▶ The tensile and tensionless raising and lowering operators are related by

$$C_n^\mu(\epsilon) = \beta_+ \alpha_n^\mu + \beta_- \tilde{\alpha}_{-n}^\mu, \quad \text{where: } \beta_\pm = \frac{1}{2} \left(\sqrt{\epsilon} \pm \frac{1}{\sqrt{\epsilon}} \right)$$

$$\tilde{C}_n^\mu(\epsilon) = \beta_- \alpha_{-n}^\mu + \beta_+ \tilde{\alpha}_n^\mu. \quad (26)$$

- ▶ **The fundamental tensionless vacuum** $|0\rangle_c$:

$$C_n^\mu |0\rangle_c = 0 = \tilde{C}_n^\mu |0\rangle_c \quad \forall n > 0. \quad (27)$$

- ▶ **Different from tensile vacuum**: mixing of tensile raising & lowering op in C, \tilde{C} .
- ▶ “Fundamental” states built by acting raising operators on $|0\rangle_c$:

$$|\Psi\rangle = \prod_{i,j} C_{-m_i}^{\mu_i} \tilde{C}_{-m_j}^{\nu_j} |0\rangle_c \quad (28)$$

THE “OTHER” VACUUM

AB, Chakraborty, Parekh 2015

The relation between operators is a Bogoliubov transformation

$$C_n^\mu = e^{iG} \alpha_n e^{-iG} = \cosh \theta \alpha_n^\mu - \sinh \theta \tilde{\alpha}_{-n}^\mu, \quad G = i \sum_{n=1}^{\infty} \theta [\alpha_{-n} \cdot \tilde{\alpha}_{-n} - \alpha_n \cdot \tilde{\alpha}_n]$$

$$\tilde{C}_n^\mu = e^{iG} \tilde{\alpha}_n e^{-iG} = -\sinh \theta \alpha_{-n}^\mu + \cosh \theta \tilde{\alpha}_n^\mu, \quad \tanh \theta = \frac{1 - \epsilon}{1 + \epsilon} \quad (29)$$

- Relation between the two vacua:

$$|0\rangle_c = \exp[iG] |0\rangle_\alpha = \left(\frac{1}{\cosh \theta} \right)^{1+1+\dots} \prod_{n=1}^{\infty} \exp[\tanh \theta \alpha_{-n} \tilde{\alpha}_{-n}] |0\rangle_\alpha \quad (30)$$

- Using the regularisation: $1 + 1 + 1 + \dots \infty = \zeta(0) = -\frac{1}{2}$

$$|0\rangle_c = \sqrt{\cosh \theta} \prod_{n=1}^{\infty} \exp[\tanh \theta \alpha_{-n} \tilde{\alpha}_{-n}] |0\rangle_\alpha \quad (31)$$

- This new vacuum is a squeezed state w.r.t $|0\rangle_\alpha$.

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TWO DIFFERENT TENSIONLESS SUPERSTRINGS

AB, (Banerjee), Chakraborty, Parekh 2016-17

- ▶ Inspired by the RNS superstring, the tensionless action:

$$S = \int d^2\sigma \left[V^a V^b \partial_a X \cdot \partial_b X + i\bar{\psi} \cdot \rho^a \partial_a \psi \right]. \quad (32)$$

- ▶ Fundamental change in the fermionic sector:

$$\{\rho^\alpha, \rho^\beta\} = 2\eta^{\alpha\beta} \rightarrow \boxed{\{\rho^\alpha, \rho^\beta\} = 2V^\alpha V^\beta} \quad (33)$$

- ▶ In the “conformal” gauge $V^\alpha = (1, 0)$,

$$\{\rho^0, \rho^1\} = \mathbb{O}, \quad (\rho^0)^2 = \mathbb{I}, \quad (\rho^1)^2 = \mathbb{O}. \quad (34)$$

- ▶ Modified Clifford algebra: invariant under similarity trans $\rho^a \rightarrow S^{-1} \rho^a S$.
- ▶ Two different classes of solutions, disconnected by similarity transformation.
- ▶ Homogeneous and Inhomogeneous tensionless superstrings.
- ▶ Different contraction of 2 copies of Super-Virasoros.
- ▶ Also called the “democratic” and the “despotic” limits [Lodato, Merbis 2016].
- ▶ Algebras appear also as asymptotic symmetries of 3d flat Supergravity.

HOMOGENEOUS TENSIONLESS SUPERSTRINGS

AB, Chakraborty, Parekh 2016

- ▶ Arises from the “trivial” solution:

$$\rho^0 = \mathbb{I}, \quad \rho^1 = \mathbb{O}. \quad (35)$$

- ▶ Residual symmetry [Lindstrom et al 1991, Gamboa et al 1989]:

$$\begin{aligned} [L_n, L_m] &= (n - m)L_{n+m}, & [L_n, M_m] &= (n - m)M_{n+m} \\ [L_n, Q_r^\alpha] &= \left(\frac{n}{2} - r\right)Q_{n+r}^\alpha, & \{Q_r^\alpha, Q_s^\beta\} &= \delta^{\alpha\beta}M_{r+s} \end{aligned} \quad (36)$$

- ▶ This is a “homogeneous” contraction of two copies of Super Virasoro.

$$\begin{aligned} L_n &= \mathcal{L}_n - \bar{\mathcal{L}}_{-n}, & M_n &= \epsilon (\mathcal{L}_n + \bar{\mathcal{L}}_{-n}), \\ Q_r^+ &= \sqrt{\epsilon} Q_r, & Q_r^- &= \sqrt{\epsilon} \bar{Q}_{-r}. \end{aligned} \quad (37)$$

- ▶ Homogeneous as both supercharges are scaled in the same way.
- ▶ Also arises as asymptotic symmetries of usual supergravity in 3d flat space. [Barnich et al. 2014; Lodato, Merbis 2016].

INHOMOGENEOUS TENSIONLESS SUPERSTRINGS

AB, Banerjee, Chakraborty, Parekh 2017

- ▶ Arises from the solution:

$$\rho^0 = \begin{pmatrix} 1 & 0 \\ 0 & -1 \end{pmatrix}, \quad \rho^1 = \begin{pmatrix} 0 & 0 \\ 1 & 0 \end{pmatrix}. \quad (38)$$

- ▶ Residual symmetry:

$$\begin{aligned} [L_n, L_m] &= (n - m)L_{n+m}, & [L_n, M_m] &= (n - m)M_{n+m}, & (39) \\ \{G_r, G_s\} &= 2L_{r+s}, & \{G_r, H_s\} &= 2M_{r+s}, \\ [L_n, G_r] &= \left(\frac{n}{2} - r\right)G_{n+r}, & [L_n, H_r] &= \left(\frac{n}{2} - r\right)H_{n+r}, & [M_n, G_r] &= \left(\frac{n}{2} - r\right)H_{n+r}. \end{aligned}$$

- ▶ This is another contraction of the Super Virasoro:

$$\begin{aligned} L_n &= \mathcal{L}_n - \bar{\mathcal{L}}_{-n}, & M_n &= \epsilon(\mathcal{L}_n + \bar{\mathcal{L}}_{-n}), \\ G_r &= Q_r - i\bar{Q}_{-r}, & H_r &= \epsilon(Q_r + i\bar{Q}_{-r}). \end{aligned} \quad (40)$$

- ▶ Inhomogeneous as supercharges are scaled in the different ways.
- ▶ Asymptotic symmetries of “twisted” supergravity in 3d flat space. [Lodato, Merbis 2016].

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HAGEDORN PHYSICS

AB, Chakraborty, Parekh 2015

- ▶ String theory: exponential growth of states. # of states at n th level $\sim \exp(4\pi n^{1/2})$.
- ▶ Canonical partition function diverges above Hagedorn temp: $\mathcal{T}_H = \frac{1}{4\pi\sqrt{\alpha'}}$.
New phase? New degrees of freedom?
- ▶ **Effective string tension near Hagedorn temperature** [Pisarski, Alvarez 1982; Olesen 1985]:

$$T_{eff} = T_0 \sqrt{1 - \frac{\mathcal{T}^2}{\mathcal{T}_H^2}} \Rightarrow T_{eff} \rightarrow 0 \text{ as } \mathcal{T} \rightarrow \mathcal{T}_H \quad (41)$$

- ▶ Our analysis is useful here: we provide a worldsheet description which was previously considered impossible.
- ▶ In the approach to \mathcal{T}_H , it becomes thermodynamically favourable to form a long string as opposed to heating up a gas of strings.
- ▶ **Conjecture:** Vacuum $|0\rangle_C$ is the worldsheet analogue of this emergent long string.
- ▶ New degrees of freedom as the excitations of $|0\rangle_C$.
- ▶ Worldsheet description doesn't break down: Carrollian structures appear!

OTHER DIRECTIONS AND APPLICATIONS

- ▶ **Thermal nature of the tensionless vacuum:** result of **left-right entanglement of the world-sheet 2d CFT** and its limit?
- ▶ Bogoliubov transformations: **Rindler physics on the worldsheet?** Observer in the tensionless vacuum always sees a thermal state.
- ▶ **Strings near spacetime singularities:** another domain for applying tensionless techniques?

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SUMMARY

Classical Tensionless Strings

- ▶ Limiting procedure reproduced earlier work by Isberg et al.
- ▶ Mode expansions done in fundamental and limit methods matched well.
- ▶ Supersymmetric generalisations: two different algebras.

Quantum Tensionless Strings

- ▶ Limiting case showed link to massless higher spins.
- ▶ Fundamental tensionless strings naturally defined a different vacuum.
- ▶ This was a coherent or a squeezed state w.r.t. the tensile vacuum.
- ▶ Connections to Hagedorn physics.
- ▶ Tensionless vacuum = emergent long string near \mathcal{T}_H .

QUESTIONS AND FUTURE DIRECTIONS

What happens to spacetime?

QUESTIONS AND FUTURE DIRECTIONS

What happens to spacetime?

- ▶ We have taken a weird limit on the worldsheet.
- ▶ Doing something drastic to the string would also affect the background where it propagates.
- ▶ Should be a deformation of Minkowski spacetime.
- ▶ Implement on worldsheet by tensionless analogue of vanishing of beta functions for maintaining conformal invariance.

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Others

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Others

- ▶ Open strings?
- ▶ Exploring Hagedorn physics with Carrollian structures.
- ▶ Connections to higher spin holography (do this in AdS!)
- ▶ String amplitudes in the tensionless limit.
- ▶ Connections to the Ambitwistor string. [\[Casali-Tourkine 2016-17\]](#)
- ▶ Many more possible avenues..

Thank you!