

# OLYMPUS and fits

Jan C. Bernauer

MITP Proton Radius Workshop, June 2014



# Overview

- TPE: Motivation
- Status of OLYMPUS
- Mainz fits
- TPE results from VEPP-3, (JLab)

# Cross section and form factors for elastic e-p scattering

The cross section:

$$\frac{\left(\frac{d\sigma}{d\Omega}\right)}{\left(\frac{d\sigma}{d\Omega}\right)_{Mott}} = \frac{1}{\varepsilon(1+\tau)} \left[ \varepsilon G_E^2(Q^2) + \tau G_M^2(Q^2) \right]$$

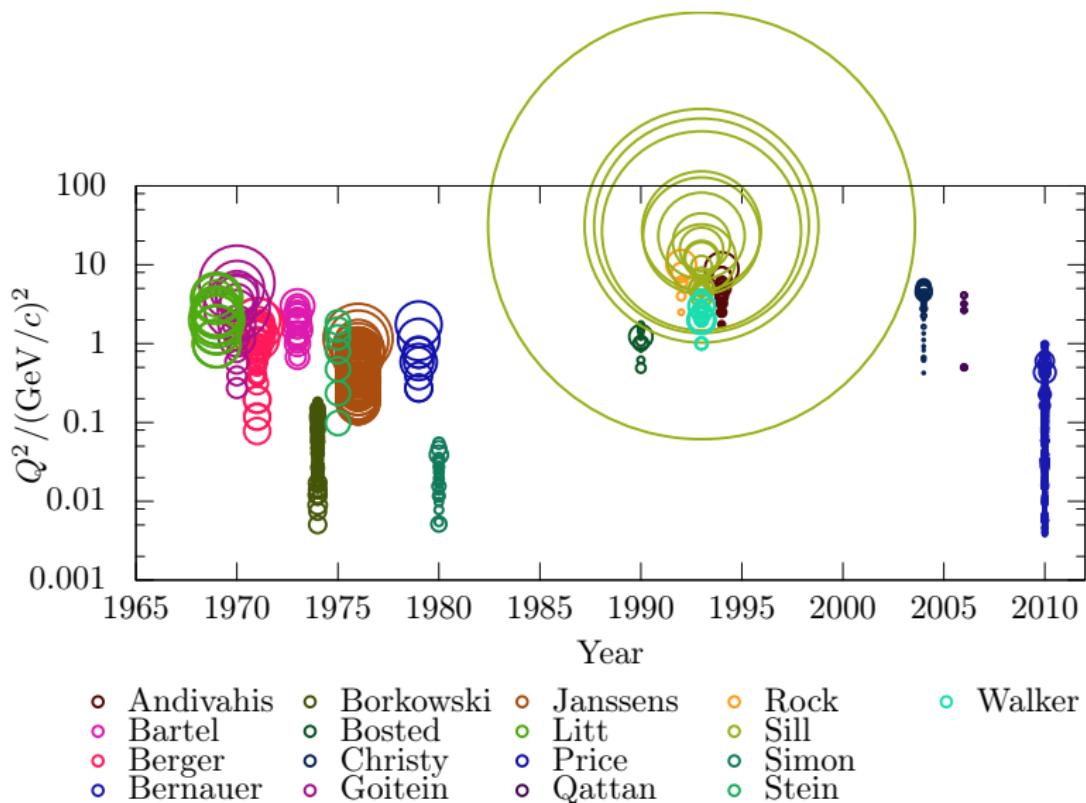
with:

$$\tau = \frac{Q^2}{4m_p^2}, \quad \varepsilon = \left( 1 + 2(1+\tau) \tan^2 \frac{\theta_e}{2} \right)^{-1}$$

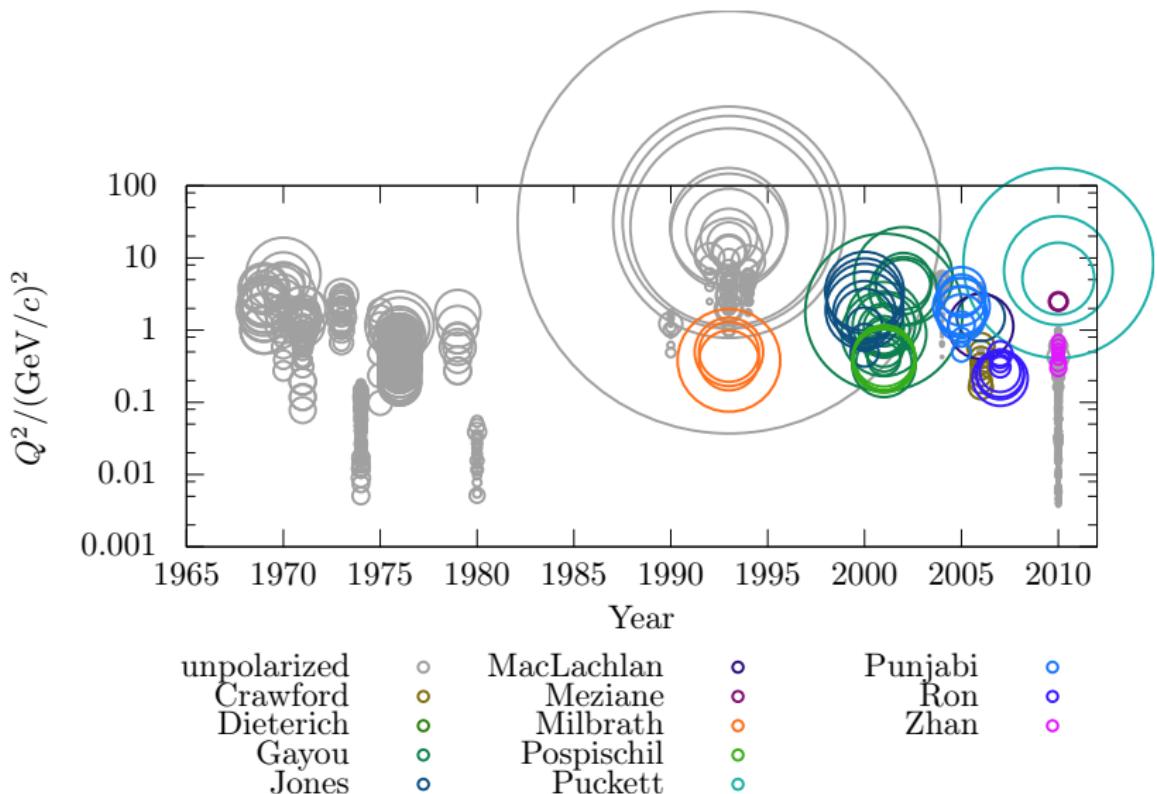
Fourier-transform of  $G_E$ ,  $G_M \rightarrow$  spatial distribution  
(Breit frame)

$$\langle r_E^2 \rangle = -6\hbar^2 \left. \frac{dG_E}{dQ^2} \right|_{Q^2=0} \quad \langle r_M^2 \rangle = -6\hbar^2 \left. \frac{d(G_M/\mu_p)}{dQ^2} \right|_{Q^2=0}$$

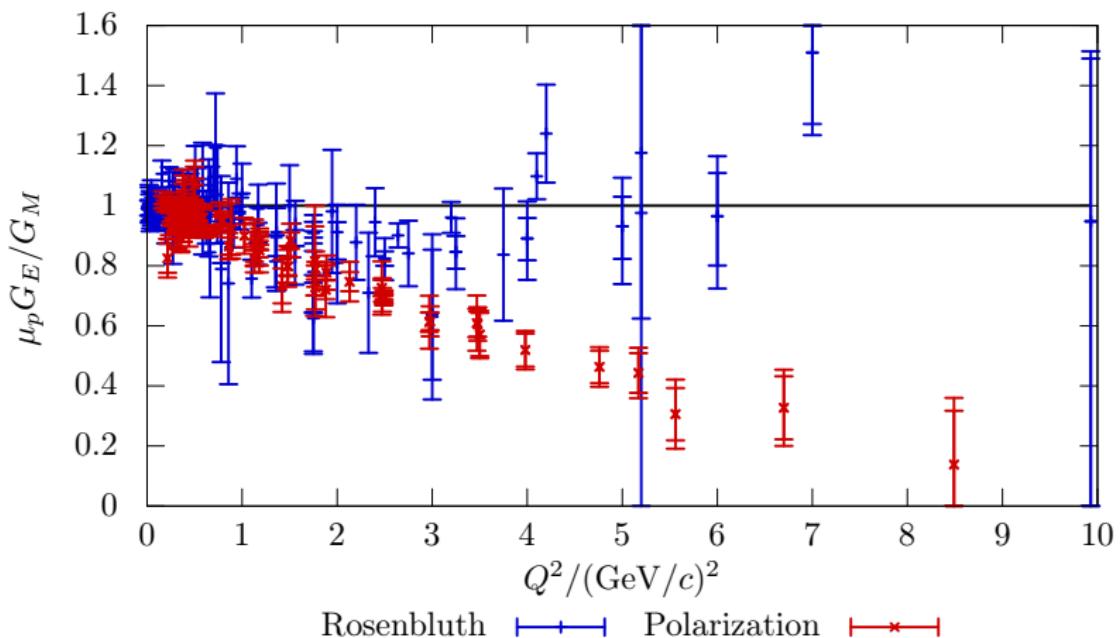
# Unpolarized: Rosenbluth



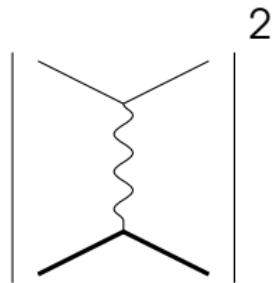
# Polarized: Ratio



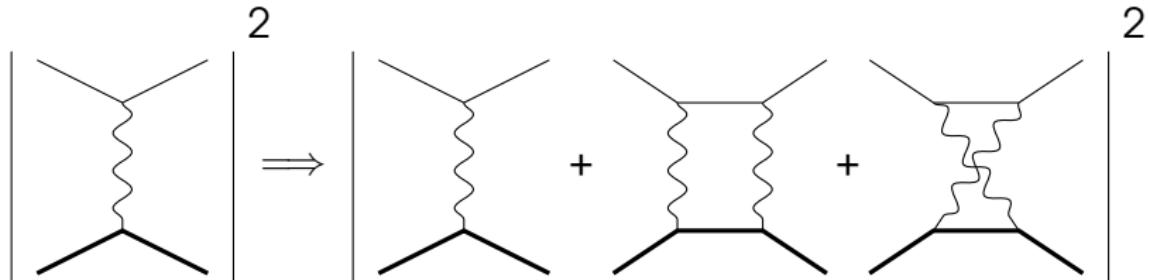
# Ratio: Difference!



# Most likely solution: Two Photon Exchange



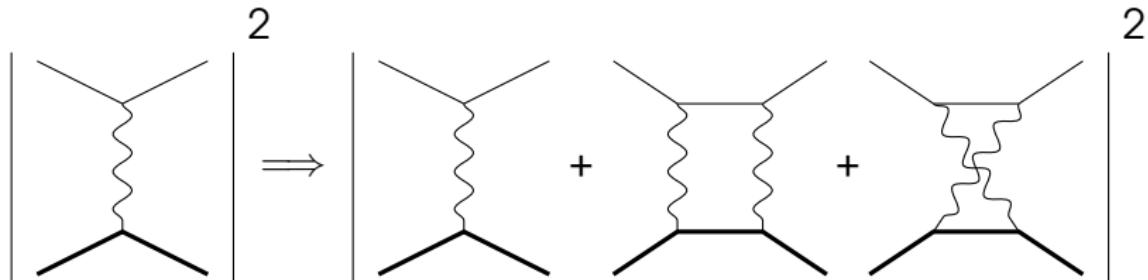
# Most likely solution: Two Photon Exchange



## Two-Photon-Exchange

- Not in standard radiative corrections
- Off-shell proton!
- How to handle high momenta in loop?

# Most likely solution: Two Photon Exchange



## Two-Photon-Exchange

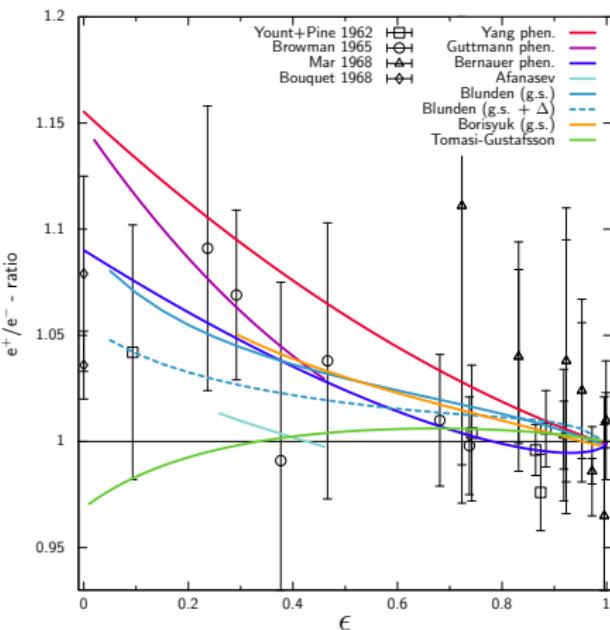
- Not in standard radiative corrections
- Off-shell proton!
- How to handle high momenta in loop?

## Measurement

- Rosenbluth/polarized reconciled?
- How to treat the hadron line?

# Measure Two Photon Exchange

- Interference term changes sign with lepton sign!
- Measured in the 1960s
- Not much data
- A lot of predictions!



# Three modern experiments

## CLAS

- $e^-$  to  $\gamma$  to  $e^{+/-}$ -beam

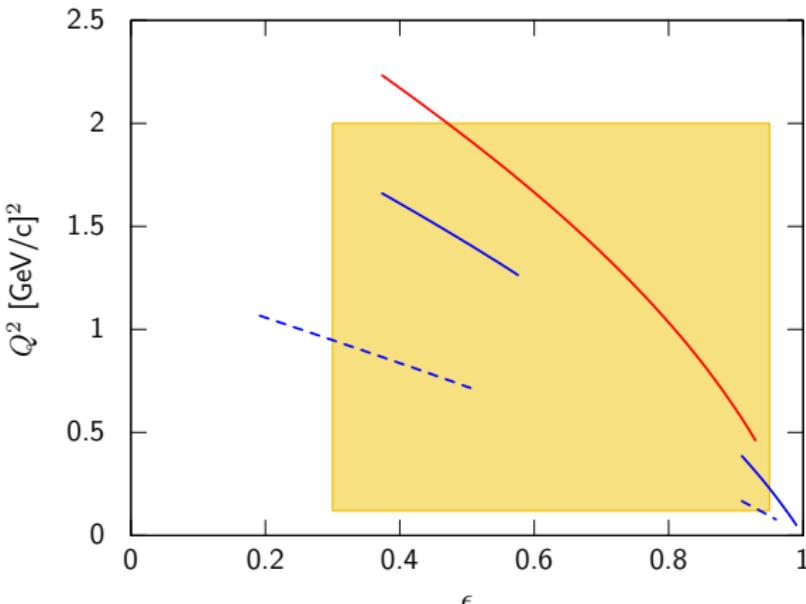
## VEPP-3

- 1.6/1 GeV beam
- no field
- preliminary results

## OLYMPUS

- DORIS @ DESY
- 2 GeV beam
- data taking finished 01/2013
- no results yet :(

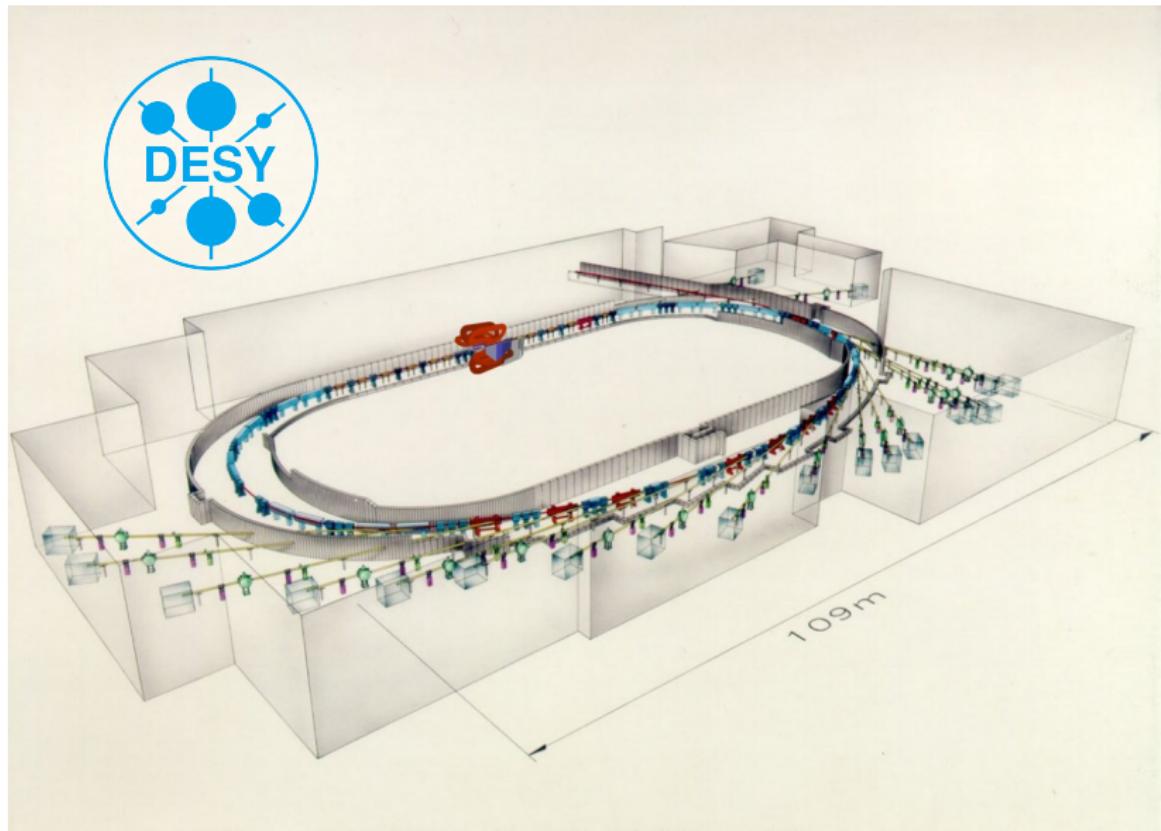
Kinematic Reach of Two-Photon Experiments



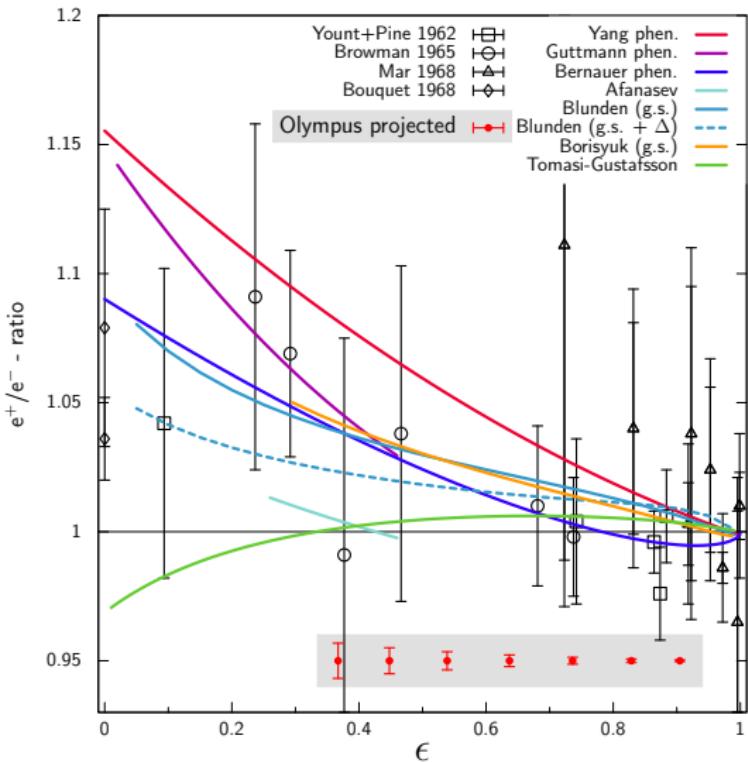
# The OLYMPUS collaboration

- Arizona State University, USA
- DESY, Hamburg, Germany
- Hampton University, USA
- INFN, Bari, Italy
- INFN, Ferrara, Italy
- INFN, Rome, Italy
- MIT Laboratory for Nuclear Science, Cambridge, USA
- Petersburg Nuclear Physics Institute, Gatchina, Russia
- University of Bonn, Bonn, Germany
- University of Glasgow, United Kingdom
- University of Mainz, Mainz, Germany
- University of New Hampshire, USA
- Yerevan Physics Institute, Armenia

# At DESY: DORIS

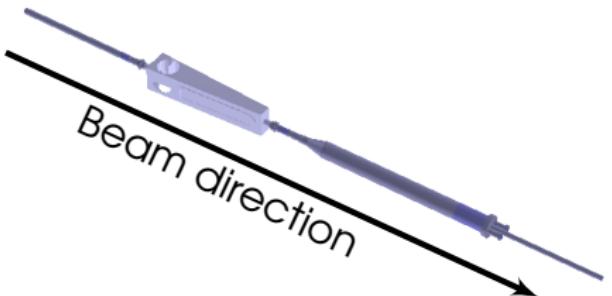


# Projected performance

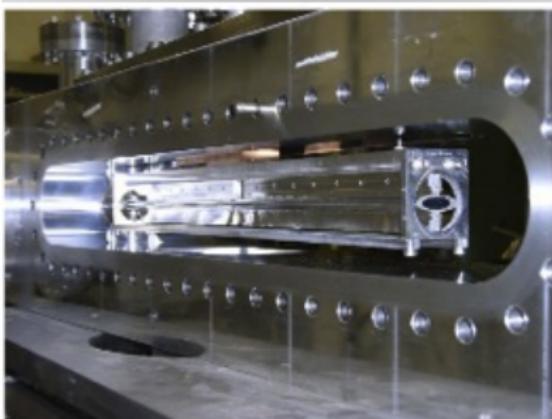
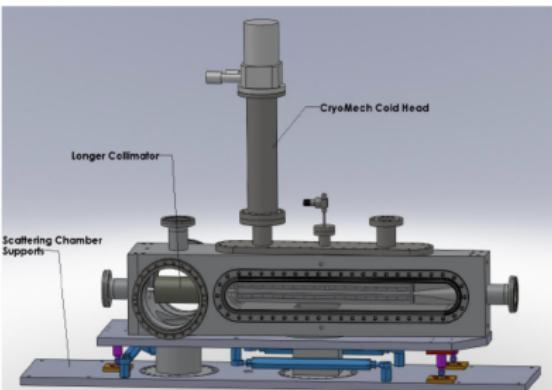


2 GeV beam,  $Q^2$ -range: 0.6 to 2.2  $\text{GeV}^2$

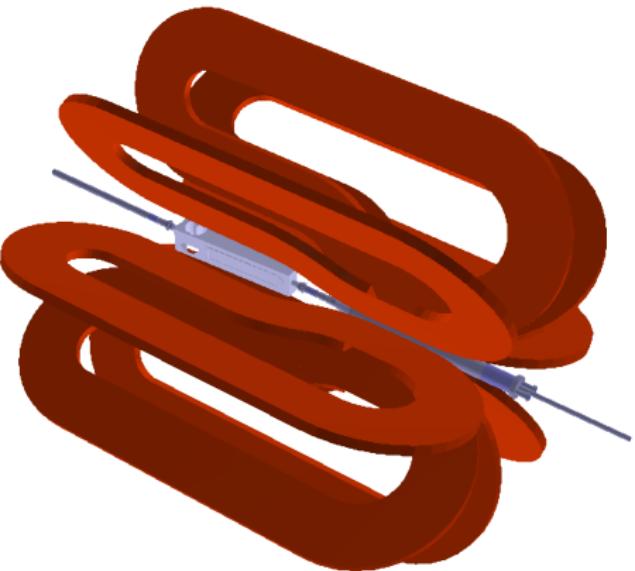
# Target / Vacuum



- Open cell design
- Cryogenic
- Target density:  $3 \cdot 10^{15} \text{ cm}^{-2}$
- Multi-stage pump system

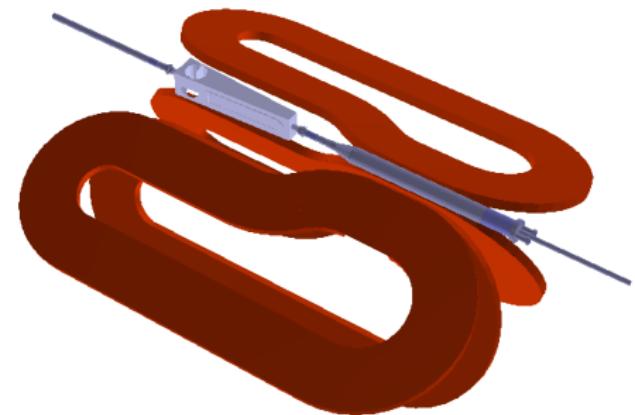


# Toroid



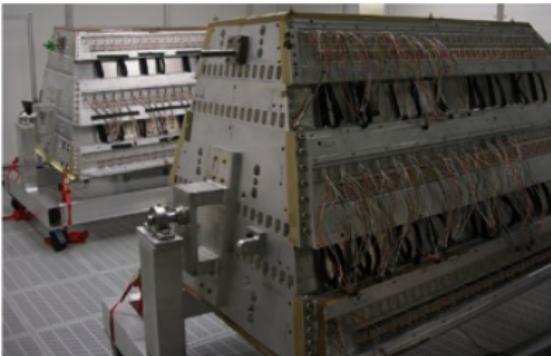
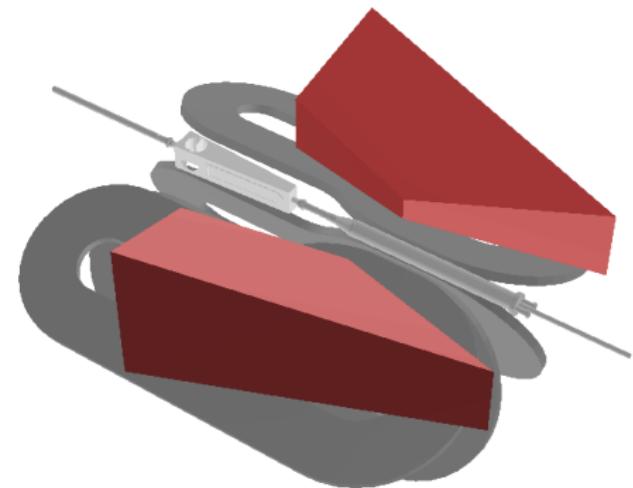
- From BLAST
- $\pm 5000 \text{ A} = 75\% \text{ of BLAST}$
- $\Rightarrow$  Peak field: 2.8 kG
- 8 coils

# Toroid



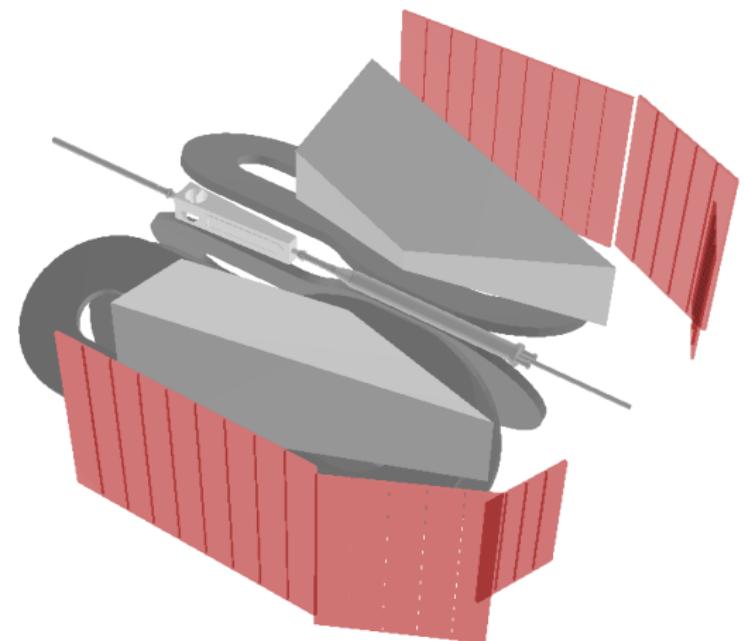
- From BLAST
- $\pm 5000 \text{ A} = 75\% \text{ of BLAST}$
- $\Rightarrow$  Peak field: 2.8 kG
- 8 coils
- 4 shown

# Wire chamber



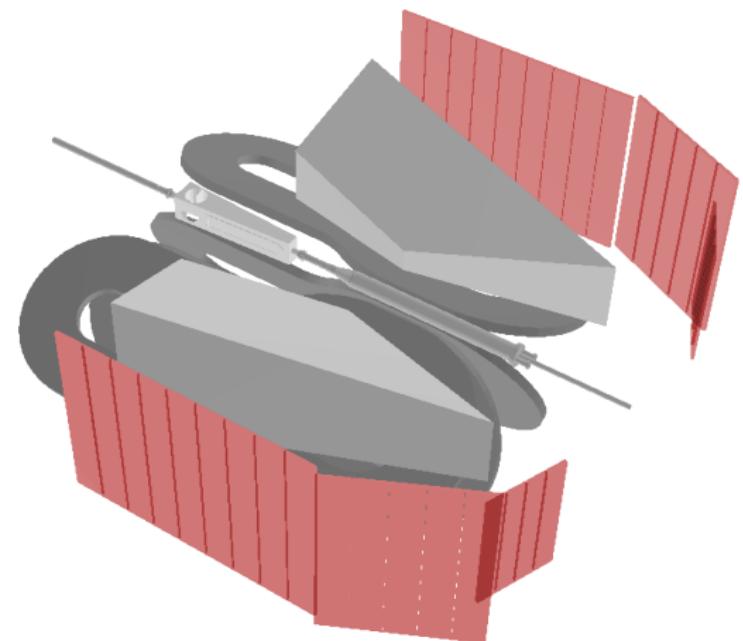
- From BLAST
- HDC design, 3 signal wires
- completely rewired
- $2 \cdot 3$  planes / chamber,  
3 chambers / side
- $10^\circ$  stereo angle

# Time Of Flight



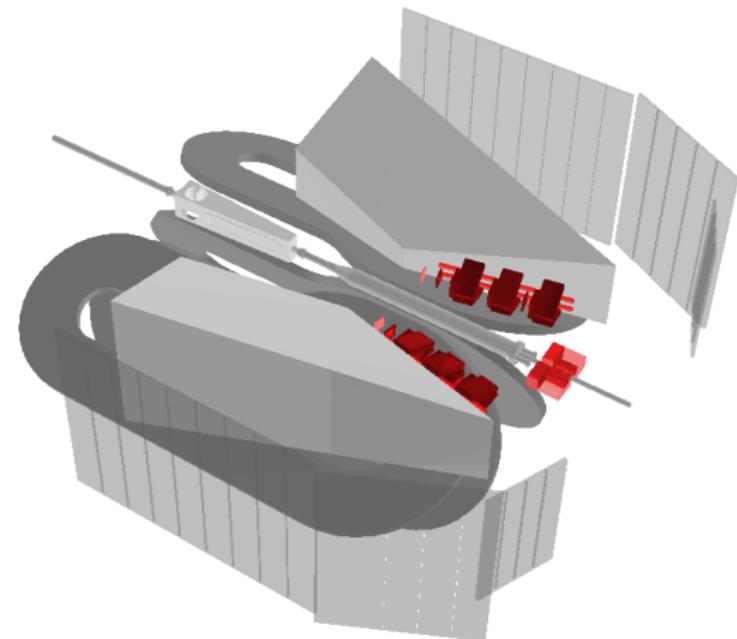
- From BLAST
- Rewrapped, tested
- Trigger
  - Top/bottom coinc.
  - kinematically constrained

# Time Of Flight



- From BLAST
- Rewrapped, tested
- Trigger
  - Top/bottom coinc.
  - kinematically constrained
  - + 2nd level WC

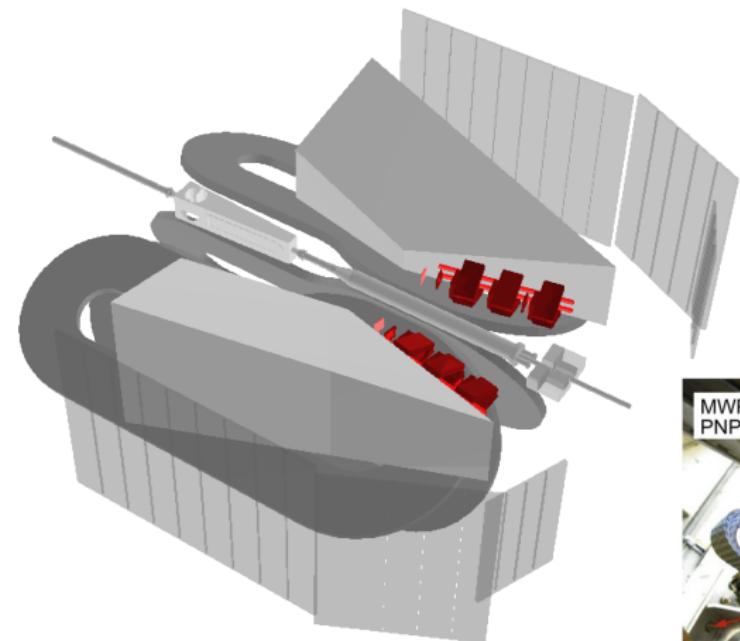
# Luminosity



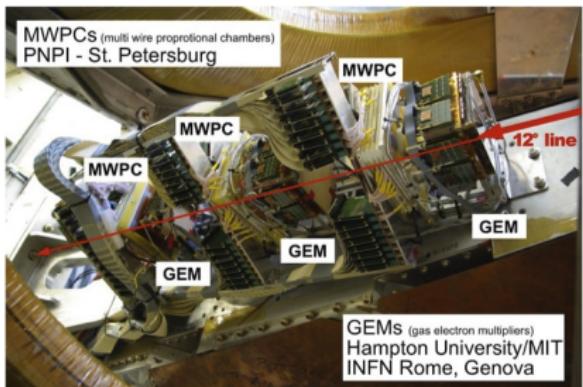
Tight control crucial!  
Redundant systems:

- $12^\circ$ -detector
- Symmetric  
Møller/Bhabha

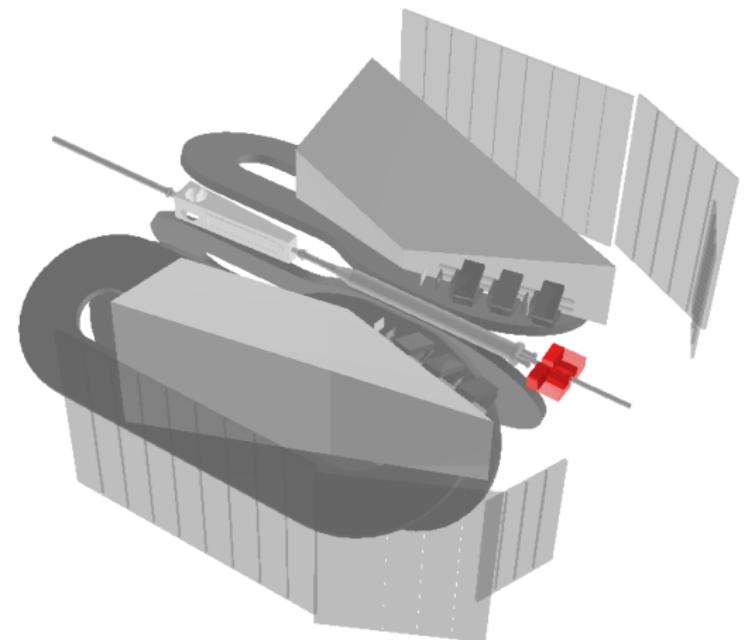
# 12°-detector



- 3 GEM (Hampton) + 3 MWPC (PNPI) each
- highly redundant
- SiPM trigger  
scintillators



# Symmetric Møller/Bhabha



- 2×9 crystals (Mainz)
- 1.2° symmetric angle
- high rate, no deadtime



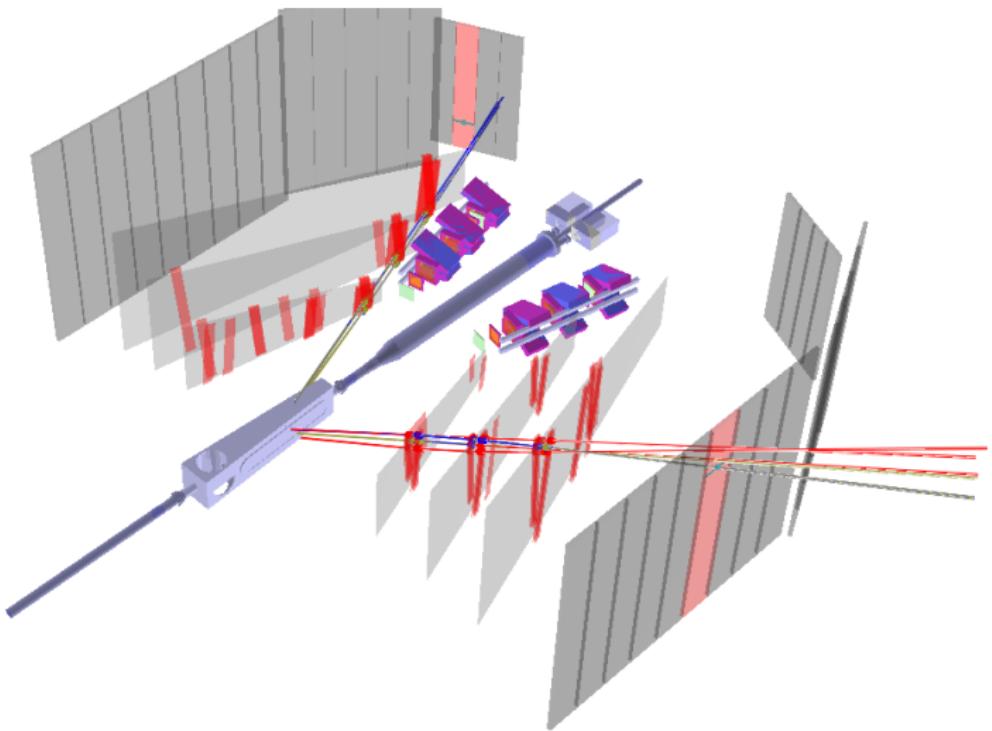
# Timeline

OLYMPUS full proposal	September 2008
Experiment funded by DOE	January 2010
BLAST moved to Germany	Spring 2010
Target test experiment	February 2011
Drift chambers installed	Spring 2011
Luminosity monitors installed	Summer 2011
Olympus roll-in	July 2011
First full Olympus test	August 2011
Sym. Møller/Bhabha installed	Fall 2011
First data run	January 2012
Second data run	October-December 2012
DORIS shut down	January 2013

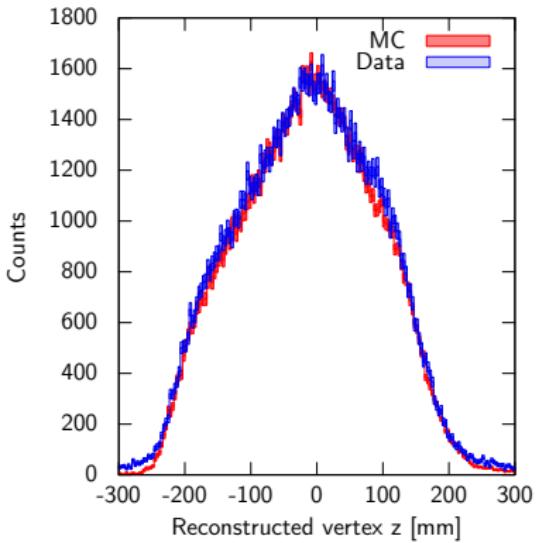
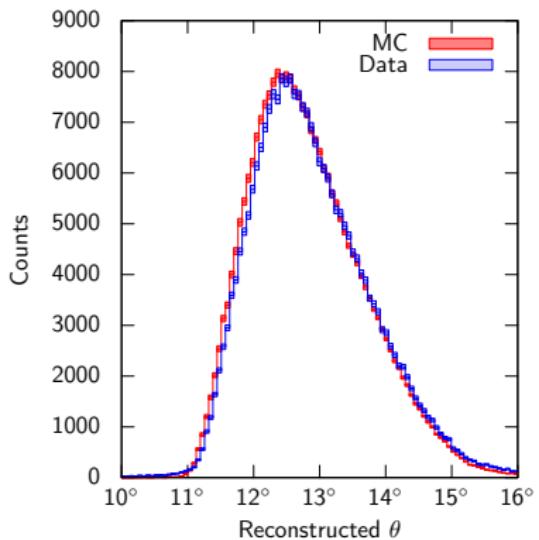
# Analysis

CAVEAT: The analysis is ongoing. All plots are preliminary.

# Wire chamber / Event-Display

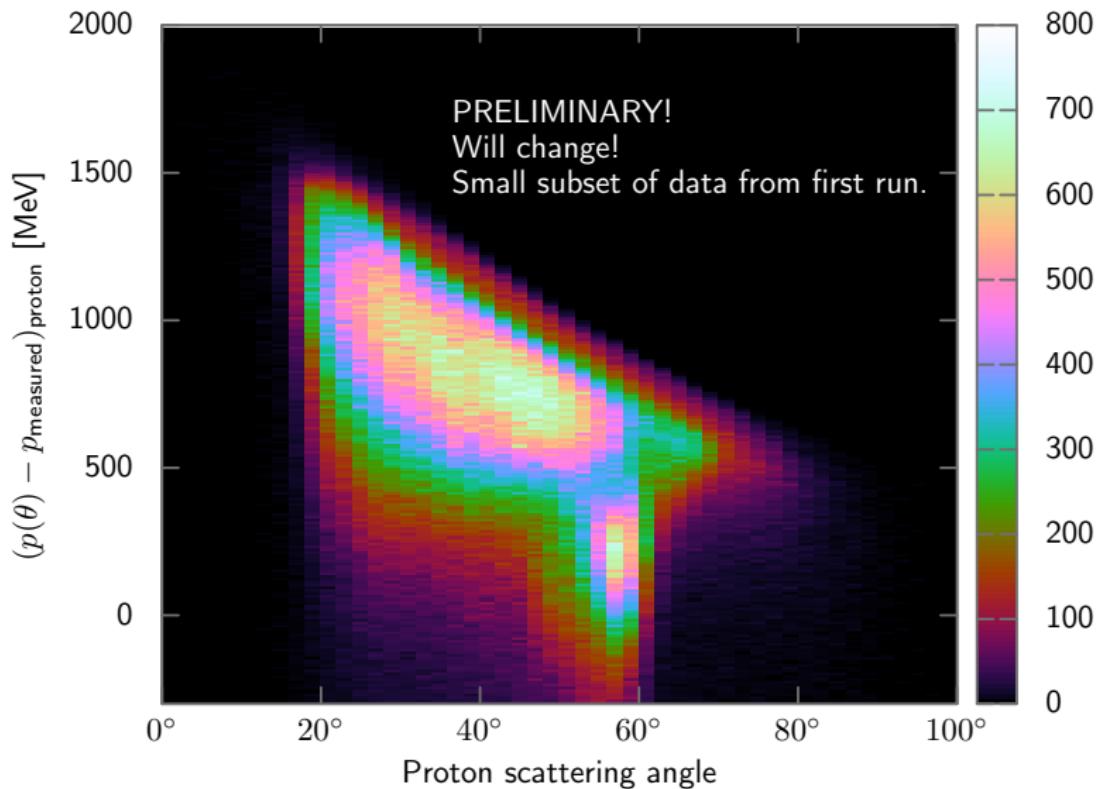


# Luminosity: Comparison Data/MC



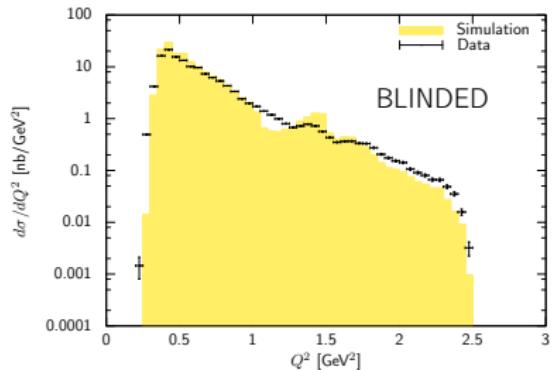
Analysis: Brian Henderson

# Reconstructed proton momentum

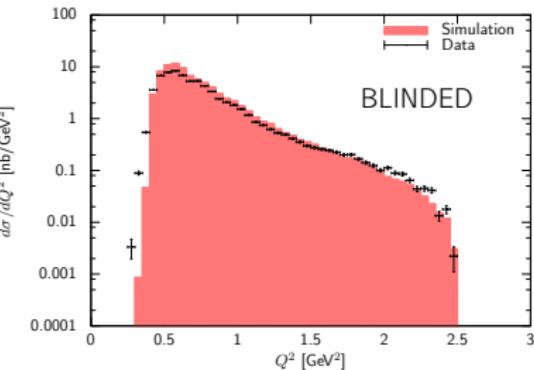


# Preliminary Analysis

Positrons



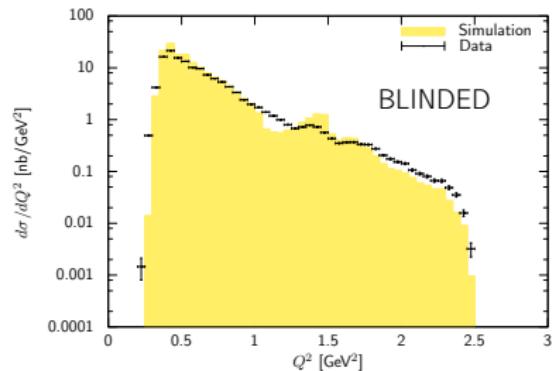
Electrons



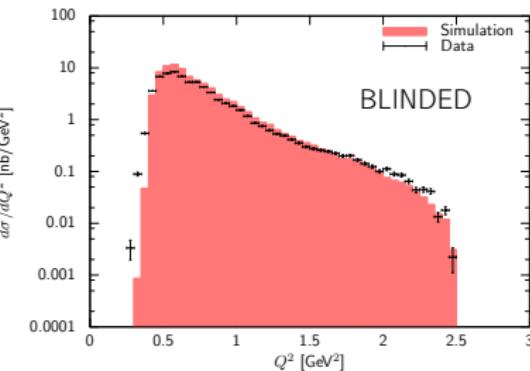
Analysis: Axel Schmidt

# Preliminary Analysis

Positrons



Electrons



Analysis: Axel Schmidt

Expect results end of 2014

# Form factor fits

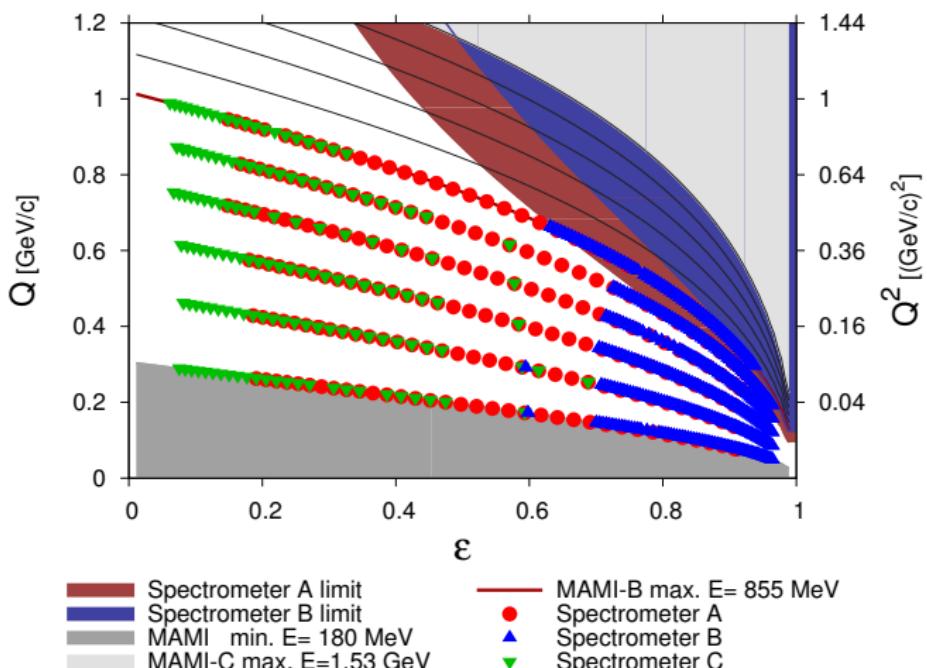
Form factor fits

# High-precision $p(e,e')p$ measurement at MAMI

Three spectrometer facility of the A1 collaboration:

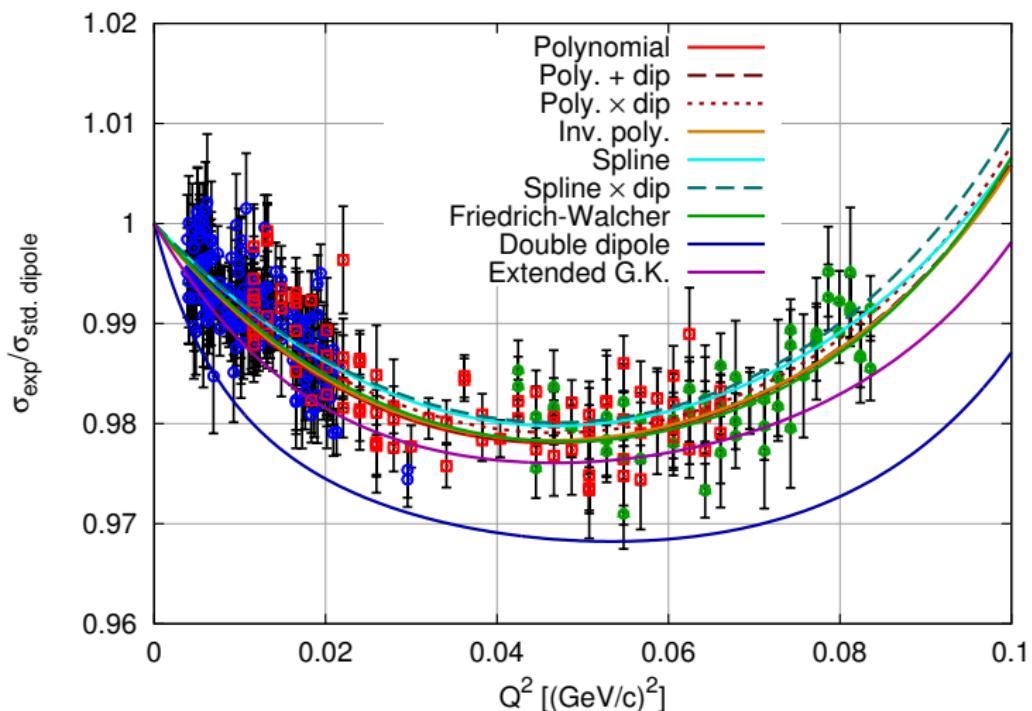


# Measured settings



1400 settings

# Cross sections: 180 MeV



## Inclusion of world data

- Extend data base with world data  
⇒ Cross check, extend  $Q^2$  reach

# Inclusion of world data

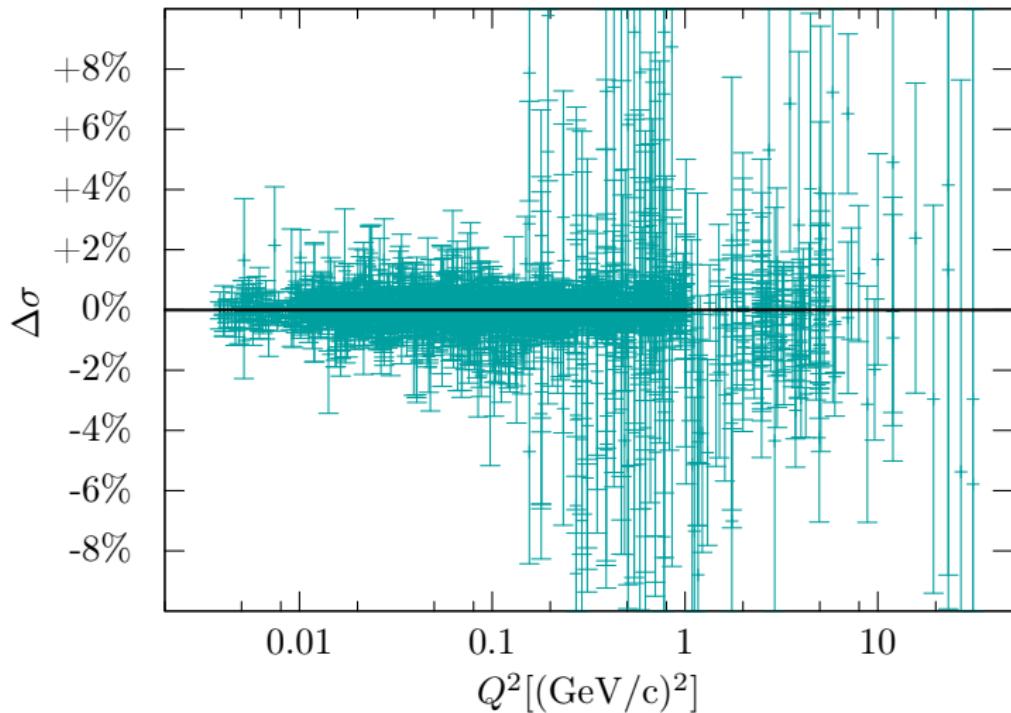
- Extend data base with world data  
⇒ Cross check, extend  $Q^2$  reach
- Take **cross sections** from Rosenbluth exp's
- Sidestep unknown error correlation
  - Update / standardize radiative corrections
  - One normalization parameter per source (Andivahis: 2)

- L. Andivahis *et al.*,  
Phys. Rev. D50, 5491 (1994).
- F. Borkowski *et al.*,  
Nucl. Phys. B93, 461 (1975).
- F. Borkowski *et al.*,  
Nucl.Phys. A222, 269 (1974).
- P. E. Bosted *et al.*,  
Phys. Rev. C 42, 38 (1990).
- M. E. Christy *et al.*,  
Phys. Rev. C70, 015206 (2004)
- M. Goitein *et al.*,  
Phys. Rev. D 1, 2449 (1970).
- T. Janssens *et al.*,  
Phys. Rev. 142, 922 (1966).
- J. Litt *et al.*,  
Phys. Lett. B31, 40 (1970).
- L. E. Price *et al.*,  
Phys. Rev. D4, 45 (1971).
- I. A. Qattan *et al.*,  
Phys. Rev. Lett. 94, 142301 (2005).
- S. Rock *et al.*,  
Phys. Rev. D 46, 24 (1992).
- A. F. Sill *et al.*,  
Phys. Rev. D 48, 29 (1993).
- G. G. Simon *et al.*,  
Nucl. Phys. A 333, 381 (1980).
- S. Stein *et al.*,  
Phys. Rev. D 12, 1884 (1975).
- R. C. Walker *et al.*,  
Phys. Rev. D 49, 5671 (1994).

# Inclusion of world data

- Extend data base with world data  
     $\Rightarrow$  Cross check, extend  $Q^2$  reach
  - Take **cross sections** from Rosenbluth exp's
  - Sidestep unknown error correlation
    - Update / standardize radiative corrections
    - One normalization parameter per source (Andivahis: 2)
  - Two models:
    - Splines with **variable** knot spacing  
     $\Rightarrow$  Adapt knot density to data density
    - Padé-Expansion  
     $\Rightarrow$  Low(er) flexibility, for comparison
- L. Andivahis *et al.*,  
Phys. Rev. D50, 5491 (1994).  
F. Borkowski *et al.*,  
Nucl. Phys. B93, 461 (1975).  
F. Borkowski *et al.*,  
Nucl.Phys. A222, 269 (1974).  
P. E. Bosted *et al.*,  
Phys. Rev. C 42, 38 (1990).  
M. E. Christy *et al.*,  
Phys. Rev. C70, 015206 (2004)  
M. Goitein *et al.*,  
Phys. Rev. D 1, 2449 (1970).  
T. Janssens *et al.*,  
Phys. Rev. 142, 922 (1966).  
J. Litt *et al.*,  
Phys. Lett. B31, 40 (1970).  
L. E. Price *et al.*,  
Phys. Rev. D4, 45 (1971).  
I. A. Qattan *et al.*,  
Phys. Rev. Lett. 94, 142301 (2005).  
S. Rock *et al.*,  
Phys. Rev. D 46, 24 (1992).  
A. F. Sill *et al.*,  
Phys. Rev. D 48, 29 (1993).  
G. G. Simon *et al.*,  
Nucl. Phys. A 333, 381 (1980).  
S. Stein *et al.*,  
Phys. Rev. D 12, 1884 (1975).  
R. C. Walker *et al.*,  
Phys. Rev. D 49, 5671 (1994).

It works!

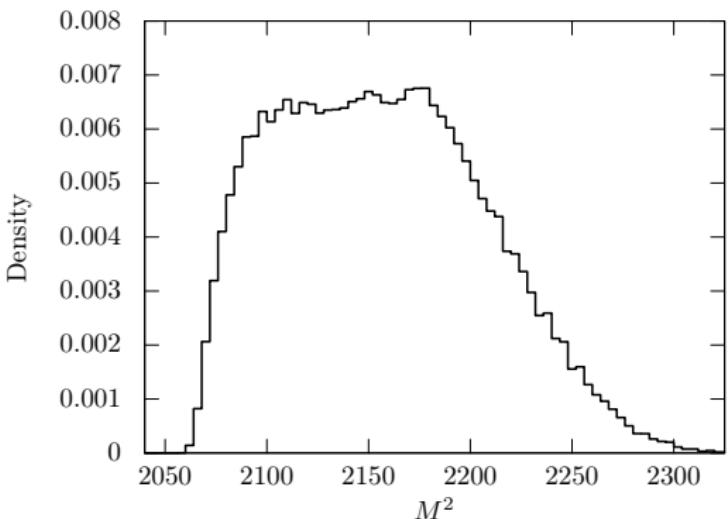


# Model dependence

- Spline model has variable knot spacing

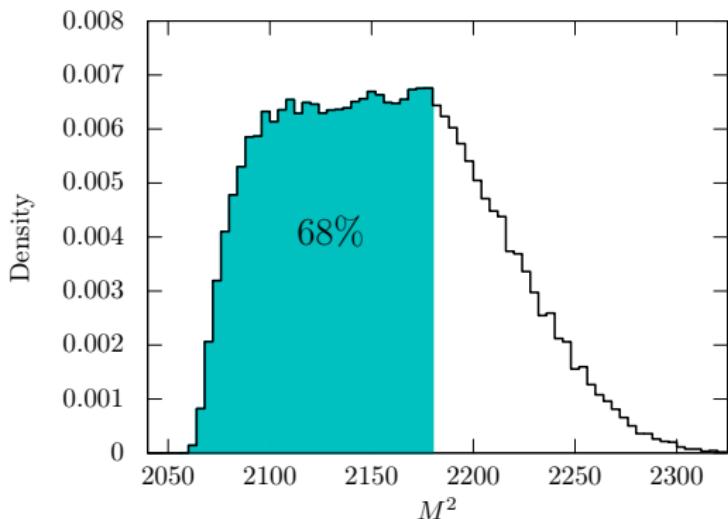
# Model dependence

- Spline model has variable knot spacing
- Vary knots, refit, record  $\chi^2$ .



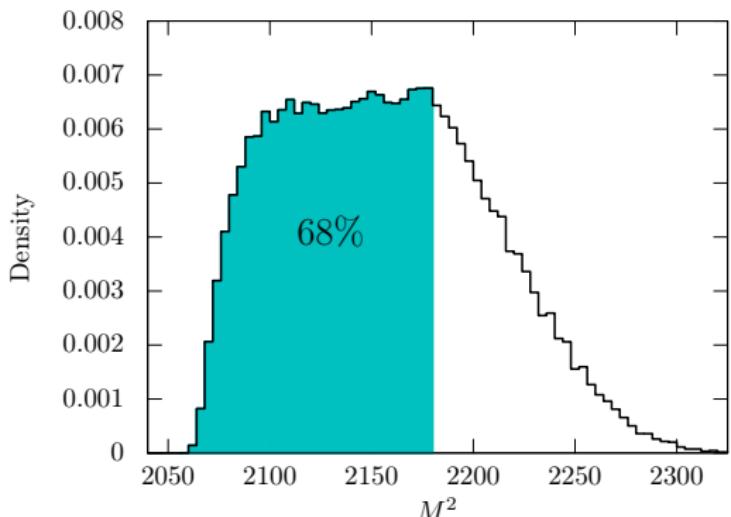
# Model dependence

- Spline model has variable knot spacing
- Vary knots, refit, record  $\chi^2$ .
- Select the 68% best tries.



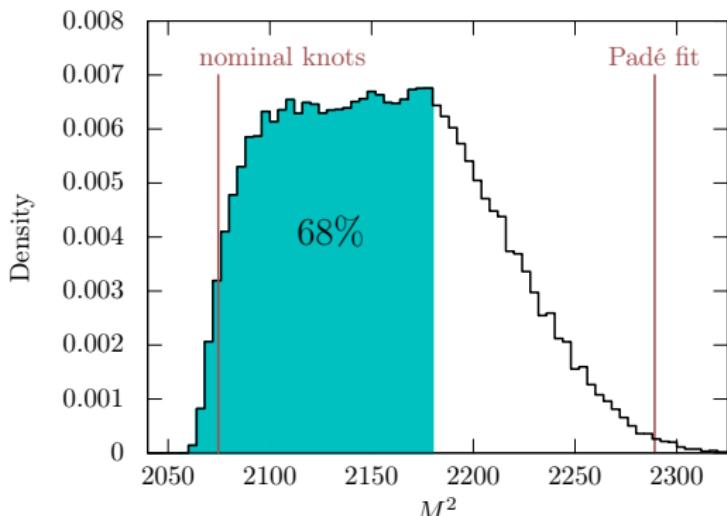
# Model dependence

- Spline model has variable knot spacing
- Vary knots, refit, record  $\chi^2$ .
- Select the 68% best tries.
- Construct envelope of models.



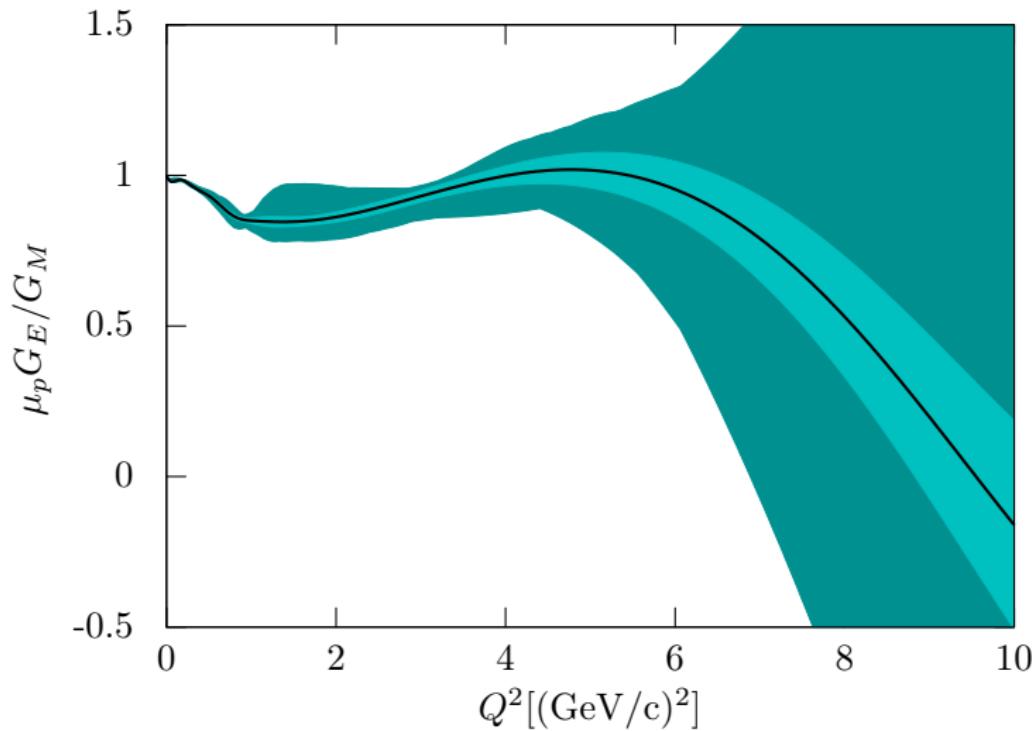
# Model dependence

- Spline model has variable knot spacing
- Vary knots, refit, record  $\chi^2$ .
- Select the 68% best tries.
- Construct envelope of models.

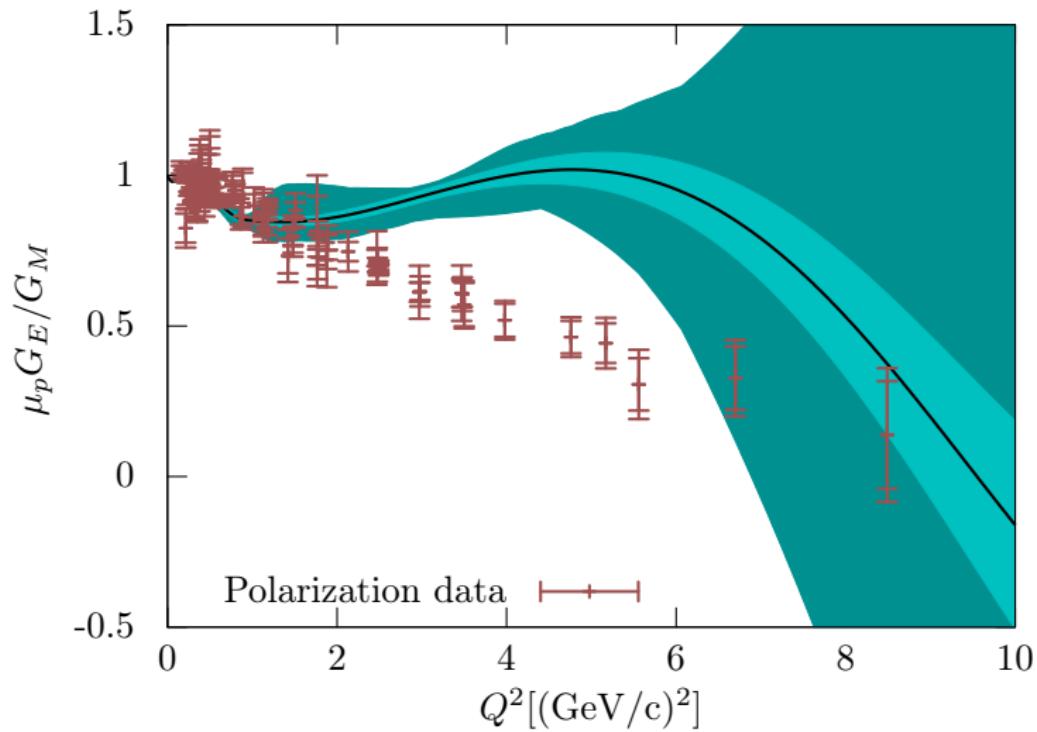


Band will cover at least 68% of all model variations!

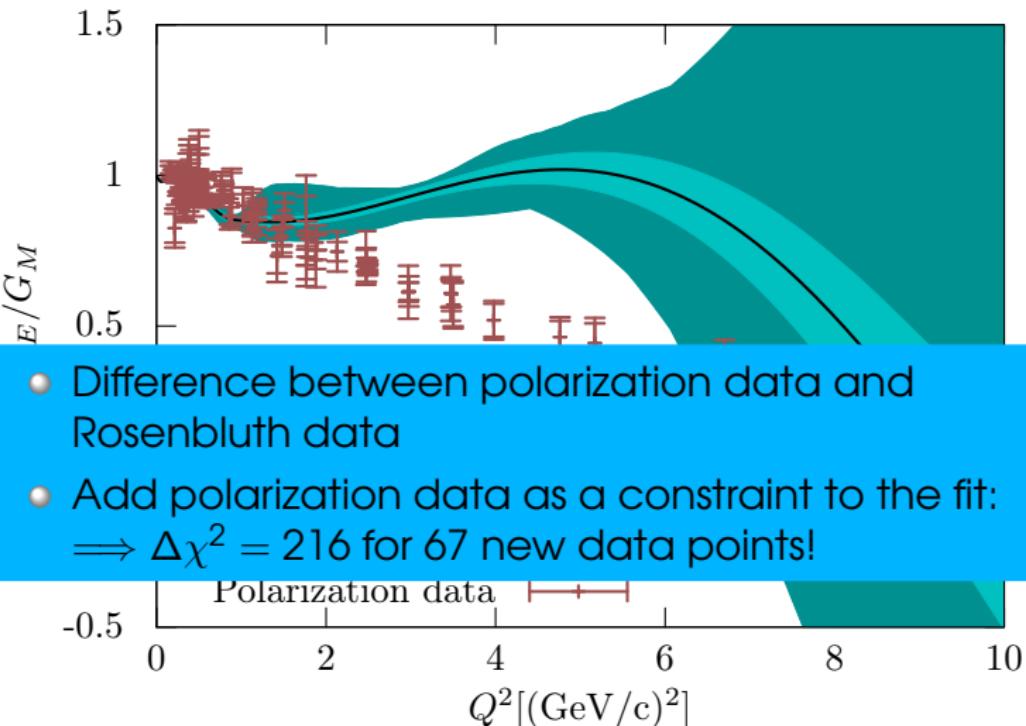
# Form factor ratio $G_E/G_M$



# Form factor ratio $G_E/G_M$



# Form factor ratio $G_E/G_M$



## Two Photon Exchange - A parametrization

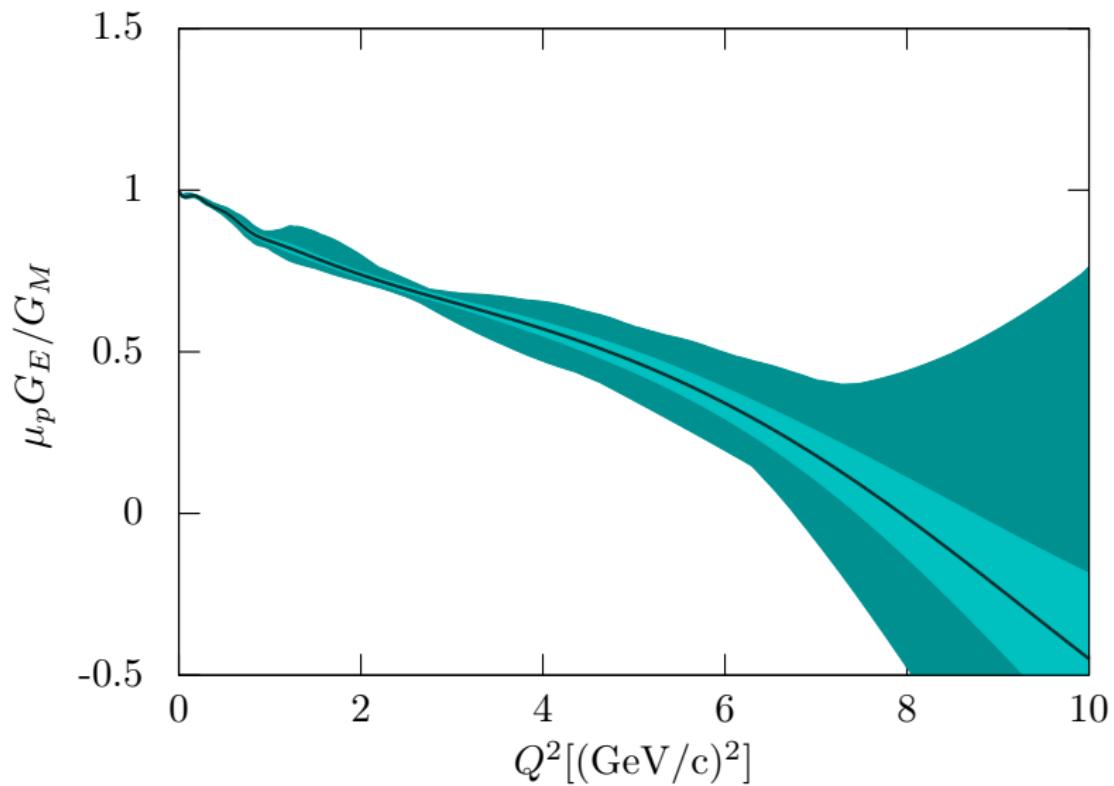
- Available data are sparse
- Mostly  $Q^2$  dependence
- Few data on  $\varepsilon$  dependence

# Two Photon Exchange - A parametrization

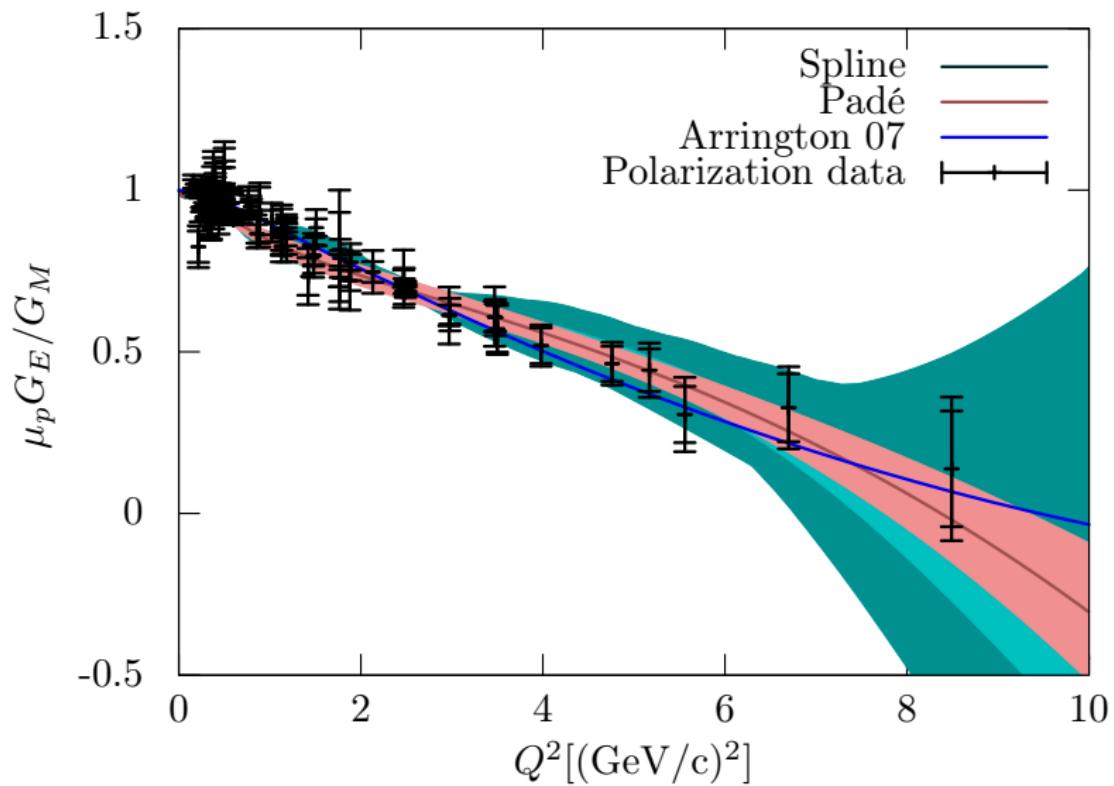
- Available data are sparse
- Mostly  $Q^2$  dependence
- Few data on  $\varepsilon$  dependence
- Only possible to fit simple model
- In addition to Feshbach Coulomb-correction!

$$\delta = a \cdot (1 - \varepsilon) \cdot \log(1 + b \cdot Q^2)$$

# $G_E/G_M$ fit incl. polarized data



# $G_E/G_M$ fit incl. polarized data



# How does this all affect the radius?

## Final result from flexible models

$$r_E = 0.879 \pm 0.005_{\text{stat.}} \pm 0.004_{\text{syst.}} \pm 0.002_{\text{model}} \pm 0.004_{\text{group}} \text{ fm},$$
$$r_M = 0.777 \pm 0.013_{\text{stat.}} \pm 0.009_{\text{syst.}} \pm 0.005_{\text{model}} \pm 0.002_{\text{group}} \text{ fm}.$$

# How does this all affect the radius?

## Final result from flexible models

$$r_E = 0.879 \pm 0.005_{\text{stat.}} \pm 0.004_{\text{syst.}} \pm 0.002_{\text{model}} \pm 0.004_{\text{group}} \text{ fm},$$
$$r_M = 0.777 \pm 0.013_{\text{stat.}} \pm 0.009_{\text{syst.}} \pm 0.005_{\text{model}} \pm 0.002_{\text{group}} \text{ fm}.$$

## More fits

	$r_E$ (fm)	$r_M$ (fm)
spline-type fits	0.875(5)(4)(2)	0.775(12)(9)(4)
polynomial-type fits	0.883(5)(5)(3)	0.778(15)(10)(6)

# How does this all affect the radius?

## Final result from flexible models

$$r_E = 0.879 \pm 0.005_{\text{stat.}} \pm 0.004_{\text{syst.}} \pm 0.002_{\text{model}} \pm 0.004_{\text{group}} \text{ fm},$$
$$r_M = 0.777 \pm 0.013_{\text{stat.}} \pm 0.009_{\text{syst.}} \pm 0.005_{\text{model}} \pm 0.002_{\text{group}} \text{ fm}.$$

## More fits

	$r_E$ (fm)	$r_M$ (fm)
spline-type fits	0.875(5)(4)(2)	0.775(12)(9)(4)
polynomial-type fits	0.883(5)(5)(3)	0.778(15)(10)(6)
var. spline +world cs.	0.878	0.772

# How does this all affect the radius?

## Final result from flexible models

$$r_E = 0.879 \pm 0.005_{\text{stat.}} \pm 0.004_{\text{syst.}} \pm 0.002_{\text{model}} \pm 0.004_{\text{group}} \text{ fm},$$
$$r_M = 0.777 \pm 0.013_{\text{stat.}} \pm 0.009_{\text{syst.}} \pm 0.005_{\text{model}} \pm 0.002_{\text{group}} \text{ fm}.$$

## More fits

	$r_E$ (fm)	$r_M$ (fm)
spline-type fits	0.875(5)(4)(2)	0.775(12)(9)(4)
polynomial-type fits	0.883(5)(5)(3)	0.778(15)(10)(6)
var. spline +world cs.	0.878	0.772
" " + polarization	0.878	0.769

# How does this all affect the radius?

## Final result from flexible models

$$r_E = 0.879 \pm 0.005_{\text{stat.}} \pm 0.004_{\text{syst.}} \pm 0.002_{\text{model}} \pm 0.004_{\text{group}} \text{ fm},$$
$$r_M = 0.777 \pm 0.013_{\text{stat.}} \pm 0.009_{\text{syst.}} \pm 0.005_{\text{model}} \pm 0.002_{\text{group}} \text{ fm}.$$

## More fits

	$r_E$ (fm)	$r_M$ (fm)
spline-type fits	0.875(5)(4)(2)	0.775(12)(9)(4)
polynomial-type fits	0.883(5)(5)(3)	0.778(15)(10)(6)
var. spline +world cs.	0.878	0.772
" " + polarization	0.878	0.769

### With other TPE calculations:

D. Borisuk et al.	0.876(5)(4)(2)(5)	0.803(13)(9)(5)(3)
Blunden et al.	0.875(5)(4)(2)(5)	0.799(13)(9)(5)(3)

# How does this all affect the radius?

## Final result from flexible models

$$r_E = 0.879 \pm 0.005_{\text{stat.}} \pm 0.004_{\text{syst.}} \pm 0.002_{\text{model}} \pm 0.004_{\text{group}} \text{ fm},$$
$$r_M = 0.777 \pm 0.013_{\text{stat.}} \pm 0.009_{\text{syst.}} \pm 0.005_{\text{model}} \pm 0.002_{\text{group}} \text{ fm}.$$

## More fits

	$r_E$ (fm)	$r_M$ (fm)
spline-type fits	0.875(5)(4)(2)	0.775(12)(9)(4)
polynomial-type fits	0.883(5)(5)(3)	0.778(15)(10)(6)
var. spline +world cs.	0.878	0.772
" " + polarization	0.878	0.769

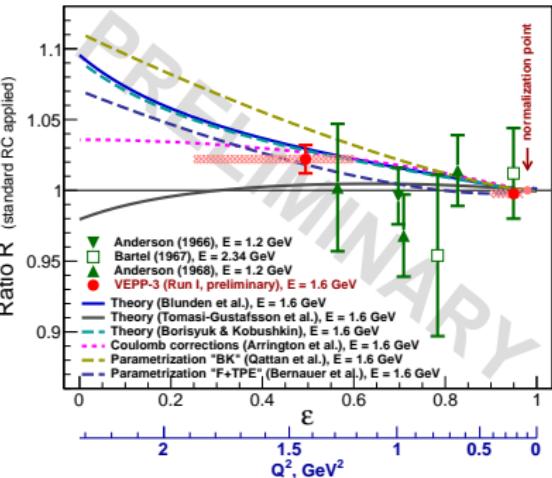
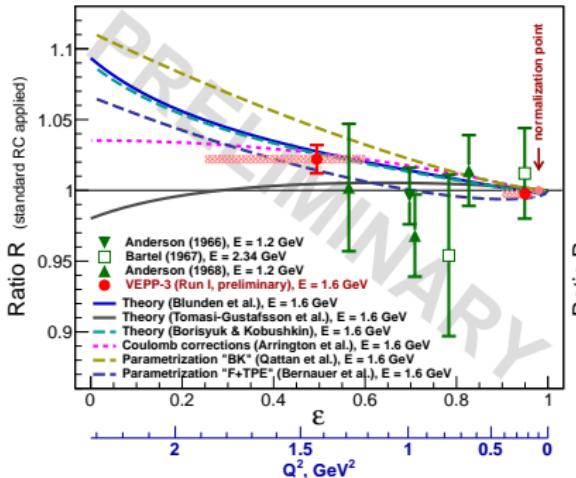
### With other TPE calculations:

D. Borisuk et al.	0.876(5)(4)(2)(5)	0.803(13)(9)(5)(3)
Blunden et al.	0.875(5)(4)(2)(5)	0.799(13)(9)(5)(3)
" "+ world (Bad $\chi^2$ )	0.875	0.787

## Comparison with TPE experiments

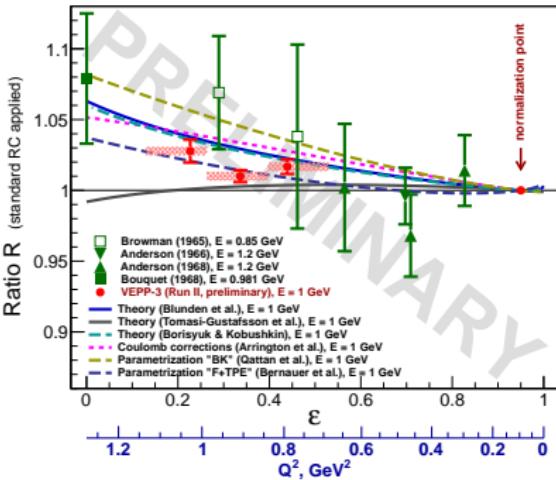
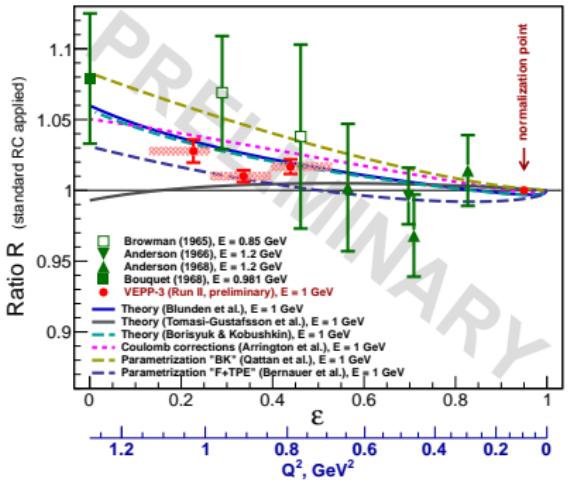
Both VEPP-3 and JLab have preliminary results - let's see.

# VEPP-3 result: Run I



Plots by Alexander Gramolin

# VEPP-3 result: Run II



Plots by Alexander Gramolin

# The take-away

- TPE prime candidate to resolve form factor discrepancy at high  $Q^2$
- Preliminary results from VEPP-3 and JLab + Mainz fit indicate effect is of right magnitude.
- OLYMPUS progresses nicely.
- TPE has so far small effect on extracted charge radius.
- Bigger on magnetic radius!

# Backup slides

# Model dependence: Charge Radius

Input		Analysis							
		Dipole	Dbl-D.	Poly.	P.+D.	P. $\times$ D.	Spline	S. $\times$ D.	F./W.
Std. dipole	811	0 $\pm$ 1	0 $\pm$ 1	0 $\pm$ 3	0 $\pm$ 3	0 $\pm$ 4	0 $\pm$ 5	0 $\pm$ 7	0 $\pm$ 1
Arrington 07	878	-18 $\pm$ 1	3 $\pm$ 3	-3 $\pm$ 3	-2 $\pm$ 3	-1 $\pm$ 4	-4 $\pm$ 5	-1 $\pm$ 6	-2 $\pm$ 3
Arr. 03 (P)	829	29 $\pm$ 1	10 $\pm$ 1	1 $\pm$ 3	1 $\pm$ 3	0 $\pm$ 4	-1 $\pm$ 5	0 $\pm$ 6	2 $\pm$ 6
Arr. 03 (R)	868	-9 $\pm$ 1	0 $\pm$ 2	0 $\pm$ 3	0 $\pm$ 3	0 $\pm$ 4	-3 $\pm$ 5	0 $\pm$ 6	-1 $\pm$ 3
FW	860	-4 $\pm$ 1	31 $\pm$ 14	-1 $\pm$ 3	-1 $\pm$ 3	1 $\pm$ 4	0 $\pm$ 5	0 $\pm$ 6	0 $\pm$ 3

Arrington 07: Phys. Rev. C 76 035205 (2007)

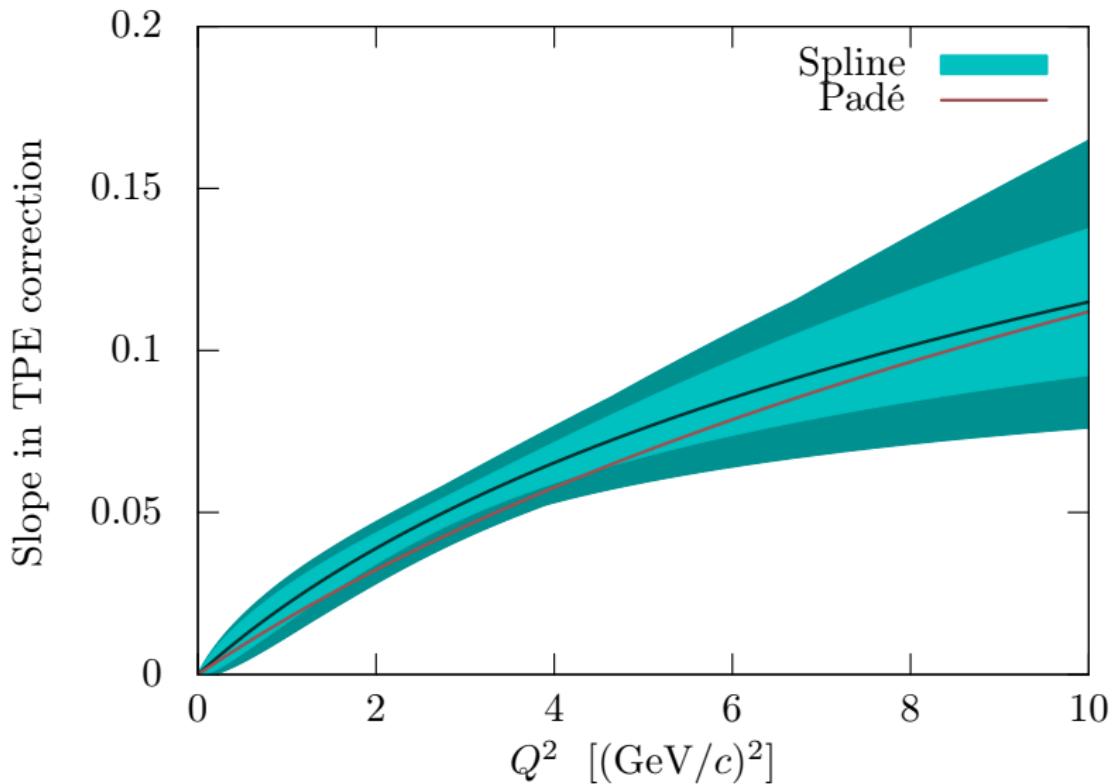
Arr. 03: Phys. Rev. C 68 034325 (2003)

FW: Eur. Phys. J. A 17 607 (2003)

# Model dependence: Magnetic Radius

Input		Analysis							
		Dipole	Dbl-D.	Poly.	P+D.	P×D.	Spline	S.×D.	F./W.
Std. dipole	811	0±1	0±1	-1±7	0±7	0±10	2±14	1±18	0±1
Arrington 07	858	-55±1	4±4	-5±6	-4±6	-1±9	2±13	0±17	-10±4
Arr. 03 (P)	837	-33±1	12±3	-1±7	0±7	0±9	2±13	0±19	-5±5
Arr. 03 (R)	863	-52±1	2±4	-4±6	-3±6	0±9	3±13	0±17	-8±4
FW	805	4±1	49±2	0±7	1±7	-1±10	1±13	-1±18	-1±4

# TPE dependence on $Q^2$



# TPE dependence on $\varepsilon$

