

# Constraining cosmological parameters with **BAO** and **RSD** from **BOSS** & **eBOSS** galaxy clustering datasets

**Héctor Gil-Marín** (*Institute Lagrange de Paris Fellow, LPNHE Sorbonne University*)  
*Tensions in the  $\Lambda$ CDM paradigm*  
*Mainz, 16th May 2018*



**SDSS** III



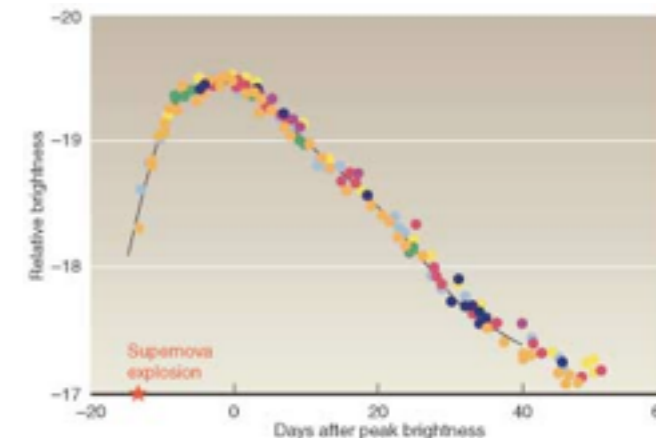
# Cosmological Standard Model

Probes to test Dark Energy on late-time Universe

- Standard Candle Supernovae

- Weak lensing

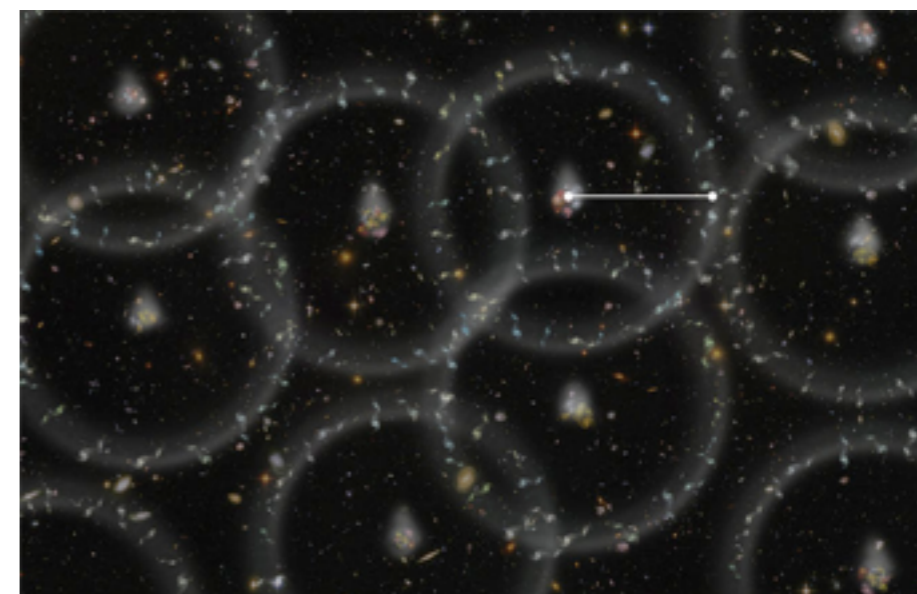
- Cluster counting



- **Standard Ruler BAO & RSD**

↓ How?

Observe dark matter tracers with high precision redshift covering a large area/volume of the sky



# Main Goal of redshift surveys: BAO & RSD

- **Standard Ruler BAO & RSD**

**Growth of structure:** Ratio monopole to quadrupole  $\sim f\sigma_8$

**BAO peak position:** in monopole  $\sim (D_A^2/H)^{1/3} / r_s \rightarrow D_V / r_s$

**BAO relative peak position:** monopole, 'quadrupole'  $\sim D_A H$

## Cosmological Parameters

$f\sigma_8(z)$   $H(z)r_s$   $D_A(z)/r_s$  ( $D_V/r_s$ )

## non-Cosmological Parameters

**Galaxy bias physics:  $b_1\sigma_8$   $b_2\sigma_8$   $\sigma_{\text{FoG}}$  .....**

# BOSS in a nutshell

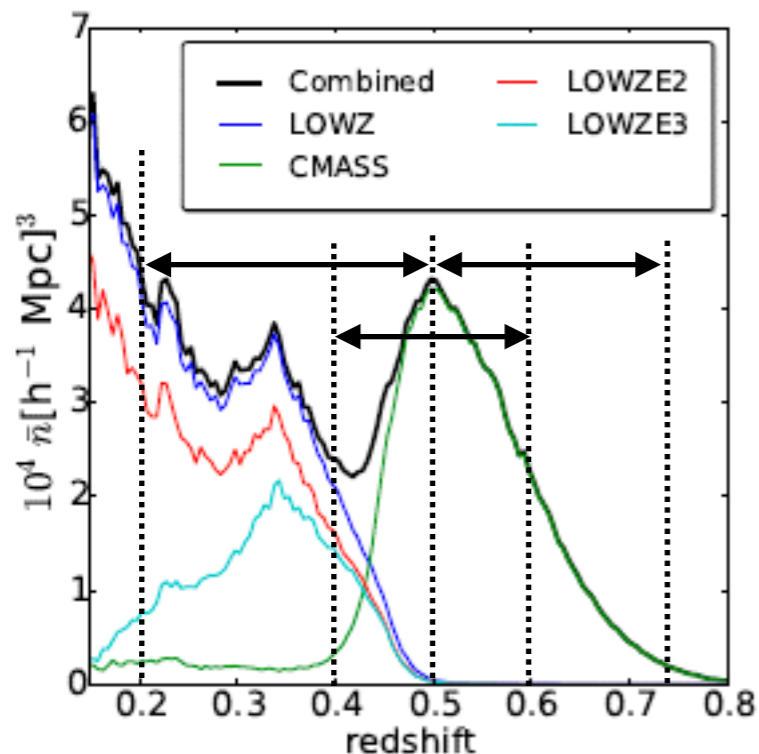
- Part of SDSS-III collaboration
- Apache Point Telescope 2.5m
- 2009 - 2014 observing period
- LOWZ-LRGs ( $0.15 < z < 0.43$ ),
- CMASS-LRGs ( $0.43 < z < 0.70$ ),
- Ly- $\alpha$  ( $z \sim 2.5$ )

3 overlapping z-bins

$0.2 < z < 0.5$

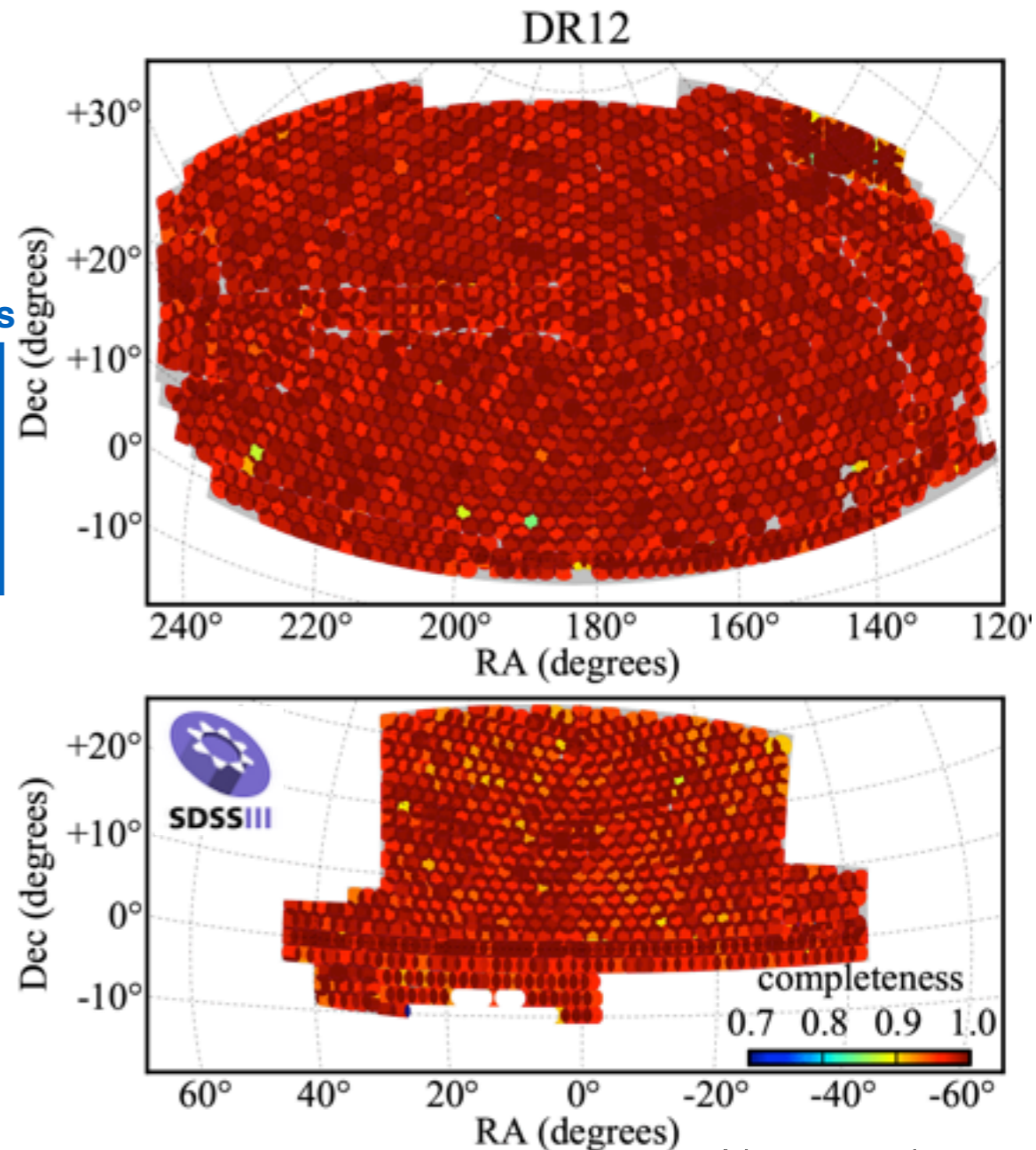
$0.4 < z < 0.6$

$0.5 < z < 0.75$



- $0.2 < z < 0.75$
- $\sim 10^6$  LRG targets
- high density of tracers  
 $4 \times 10^{-4} h/\text{Mpc}$
- High density variation

DR12 footprint for the LRG sample



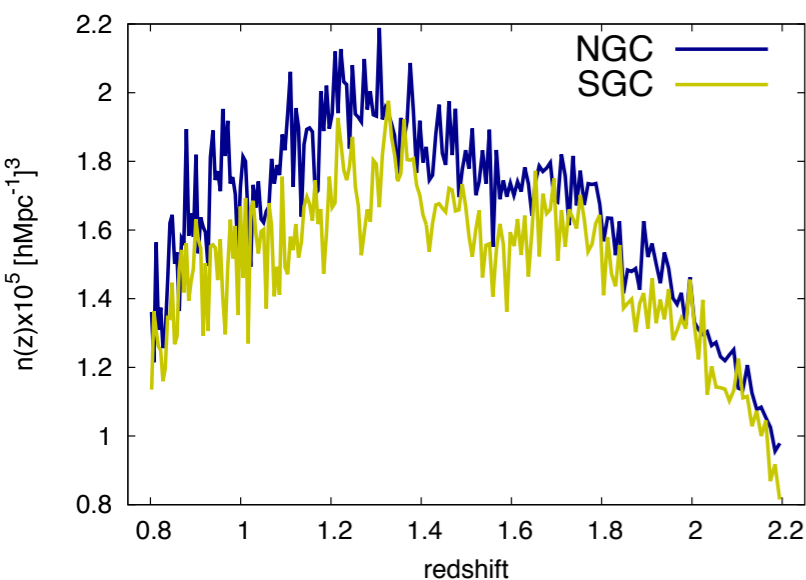
*Alam et al. 2016*

**Area = 9376 deg<sup>2</sup>**

# eBOSS in a nutshell

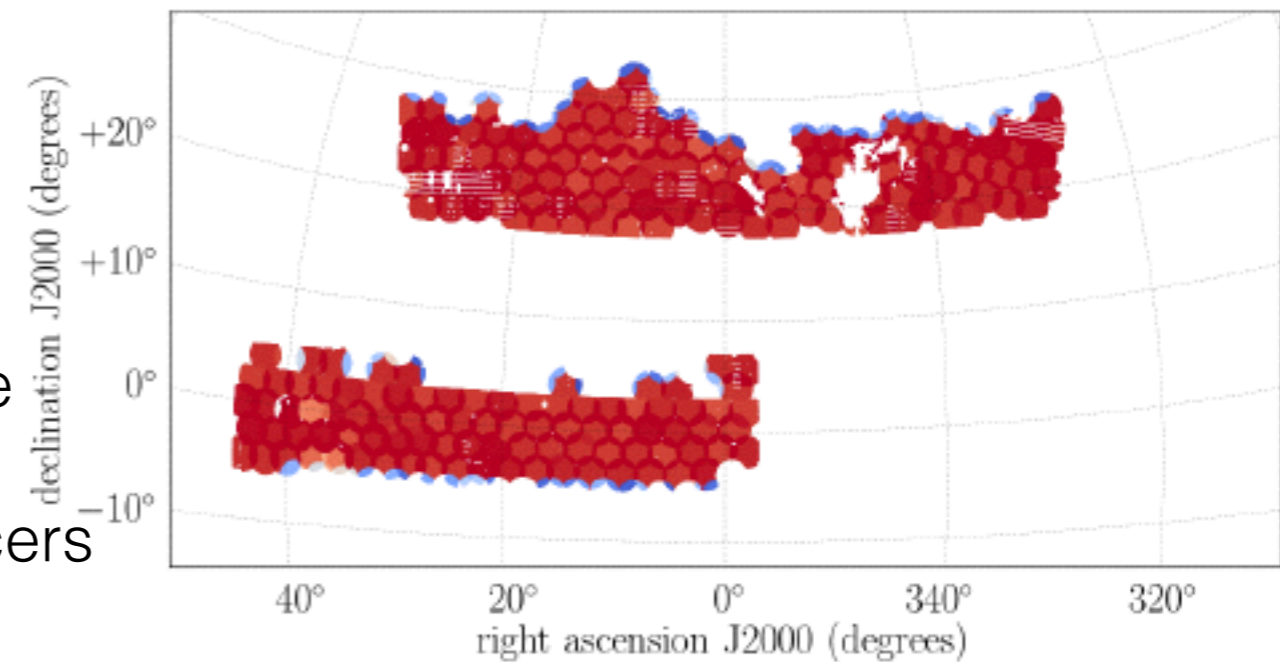
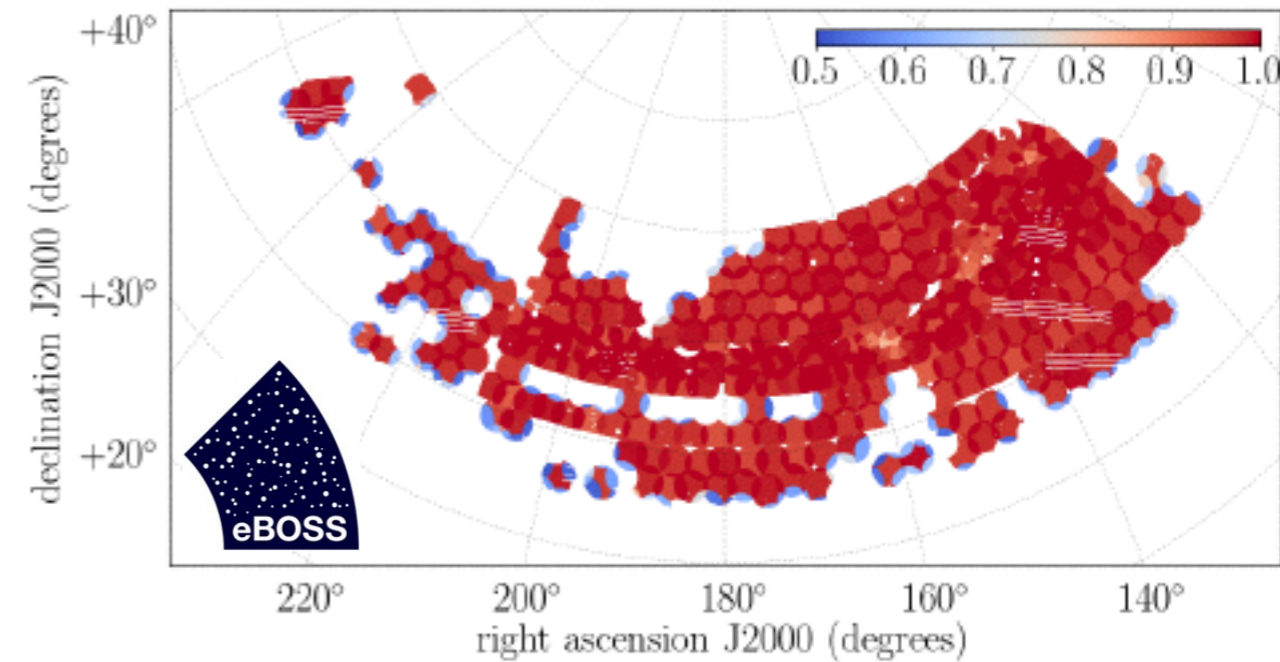
- Part of SDSS-IV collaboration
- Apache Point Telescope 2.5m
- 2014 - 2019 observing period
- LRGs & ELGs ( $0.6 < z < 1.1$ ),
- Quasars ( $0.8 < z < 2.2$ )
- Ly- $\alpha$  ( $z \sim 2.5$ )

## Quasar Sample (DR14)



- $0.8 < z < 2.2$
- Wide redshift range
- 148,659 quasars
- Low density of tracers  
 $2 \times 10^{-5} \text{ h/Mpc}$
- Low density variation

## DR14 footprint for the quasar sample



*Ata et al. 2017*

**Area = 2112.9 deg<sup>2</sup>**

# Clustering of Tracers: What do we measure?

The redshift survey catalogues deliver: **angles** and **redshifts** for each galaxy

$$r(z) = \int_0^z \frac{cdz'}{H(z', \Omega_m)}$$

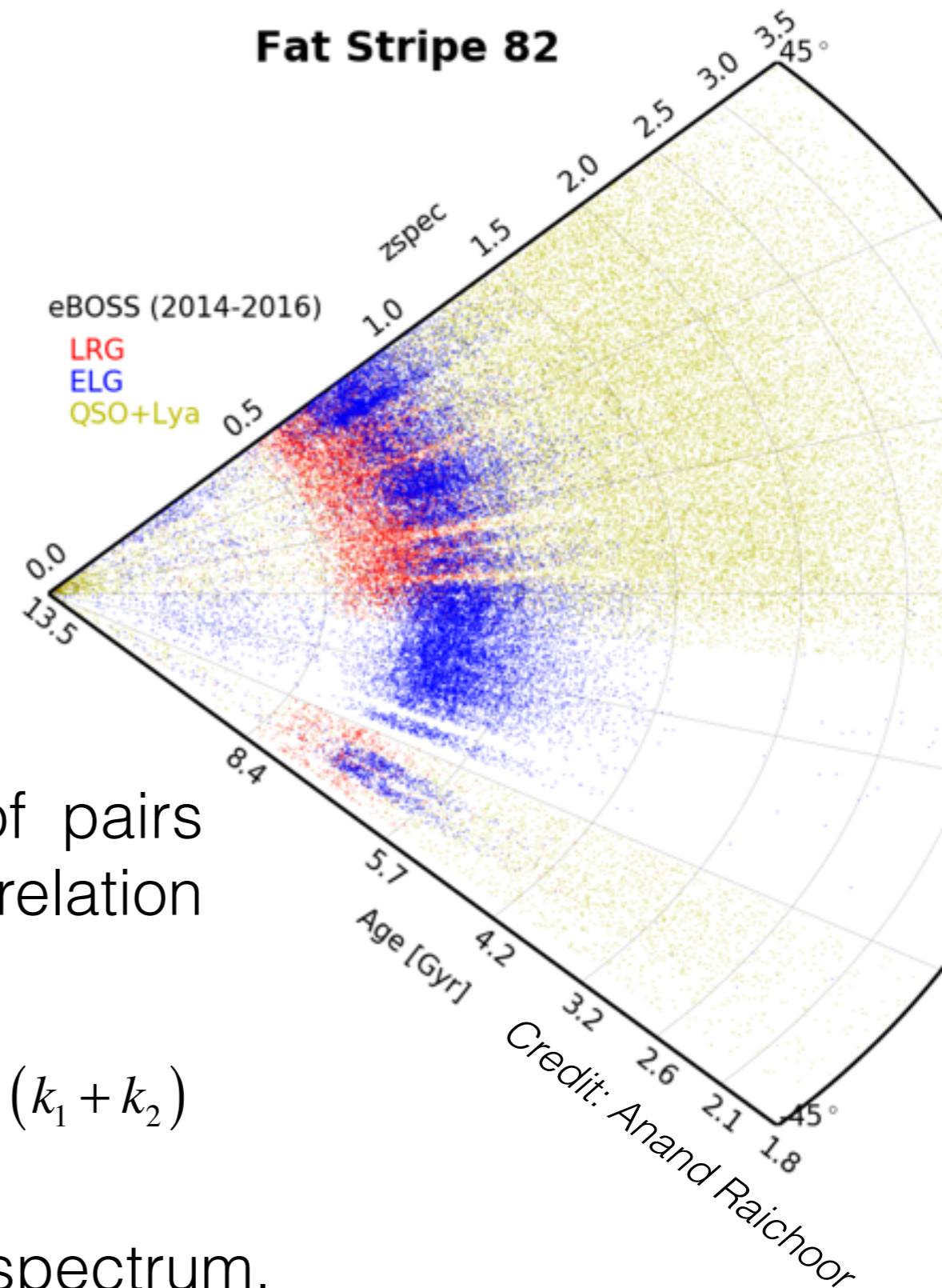
$$H(z, \Omega_m) = H_0 \sqrt{\Omega_m (1+z)^3 + 1 - \Omega_m}$$

3D galaxy maps

**Clustering strength:** Quantify number of pairs over a uniform random distribution: correlation function,  $\xi(R)$ , or Power Spectrum,  $P(k)$

$$\langle \delta(r_1) \delta(r_2) \rangle = \xi(r_1 - r_2) \quad \langle \delta(k_1) \delta(k_2) \rangle = P(k_1) \delta^D(k_1 + k_2)$$

... and higher order functions, such as bispectrum.



# Baryonic Acoustic Oscillation peak

## Before recombination

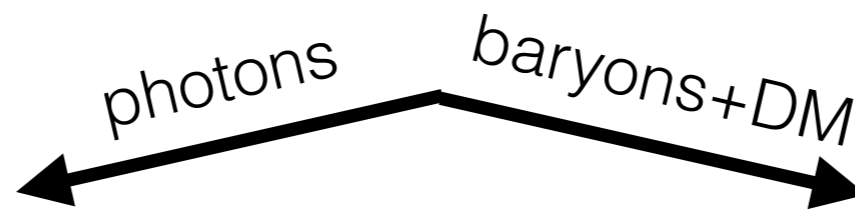
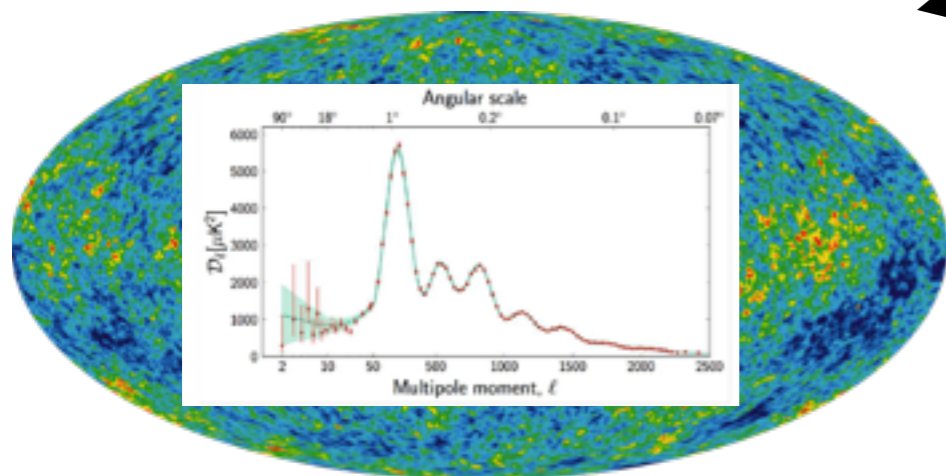
- Photons and baryons are coupled
- Radiation pressure makes them oscillate
- Pressure waves within the fluid travelling at certain speed

## Decoupling at $z \sim 1000$

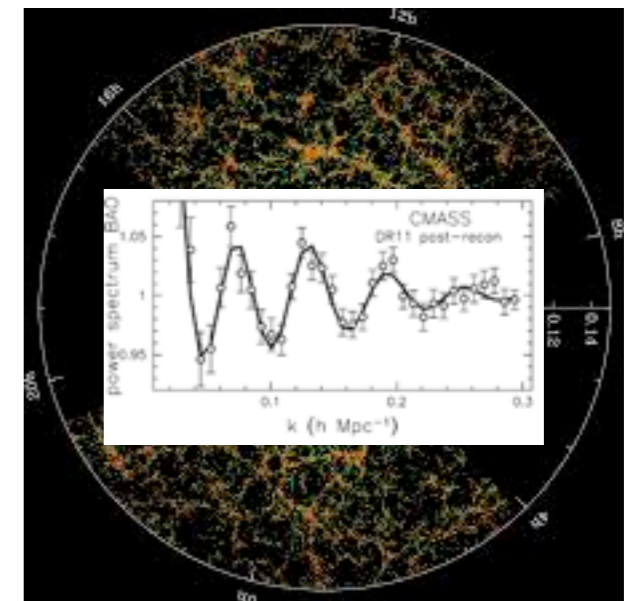
- Universe become optically thin
- Decoupling between baryons and photons
- Sound speed within the fluid decreases
- Travelling wave freezes

Baryons & photons are left a distinct imprint, a spherical peak, at a specific scale:  
**The sound horizon scale at recombination,  $\sim 150$  Mpc.**

## CMB

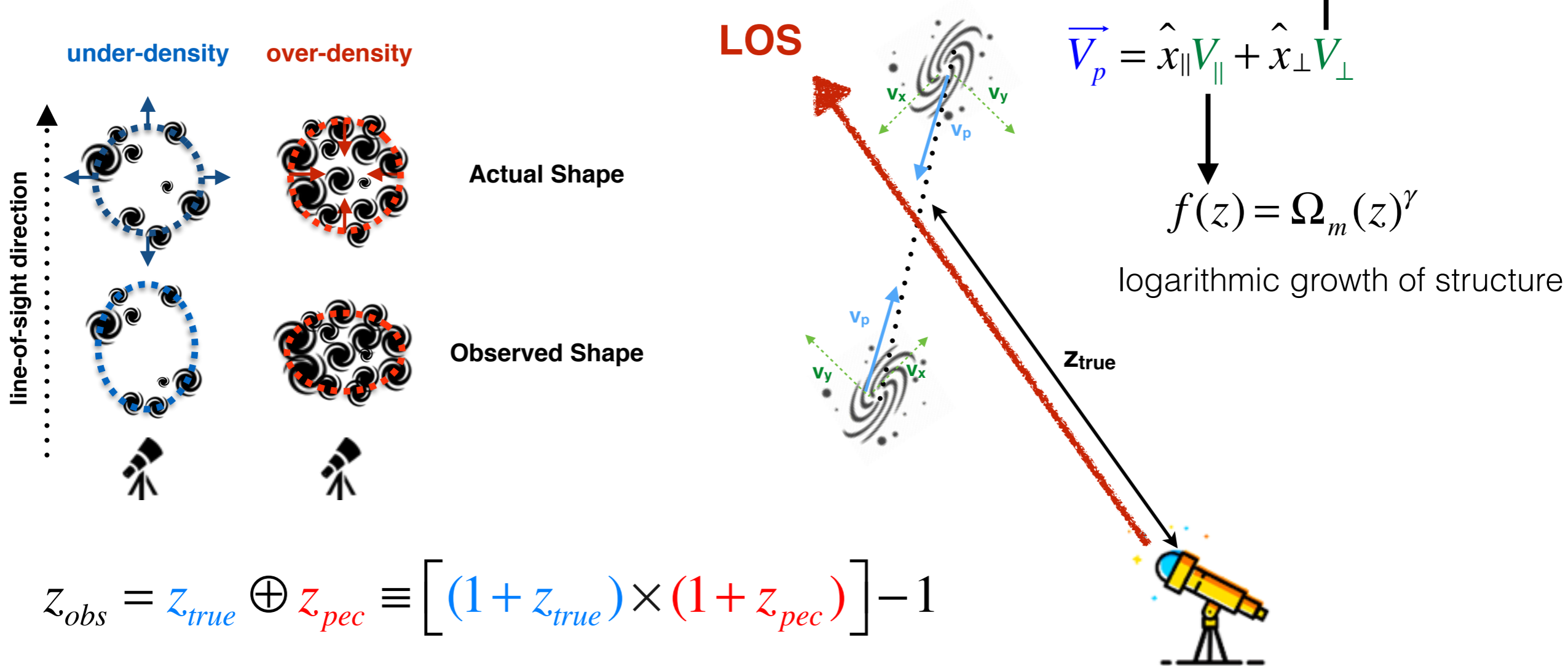


## Galaxies



# Redshift Space Distortions

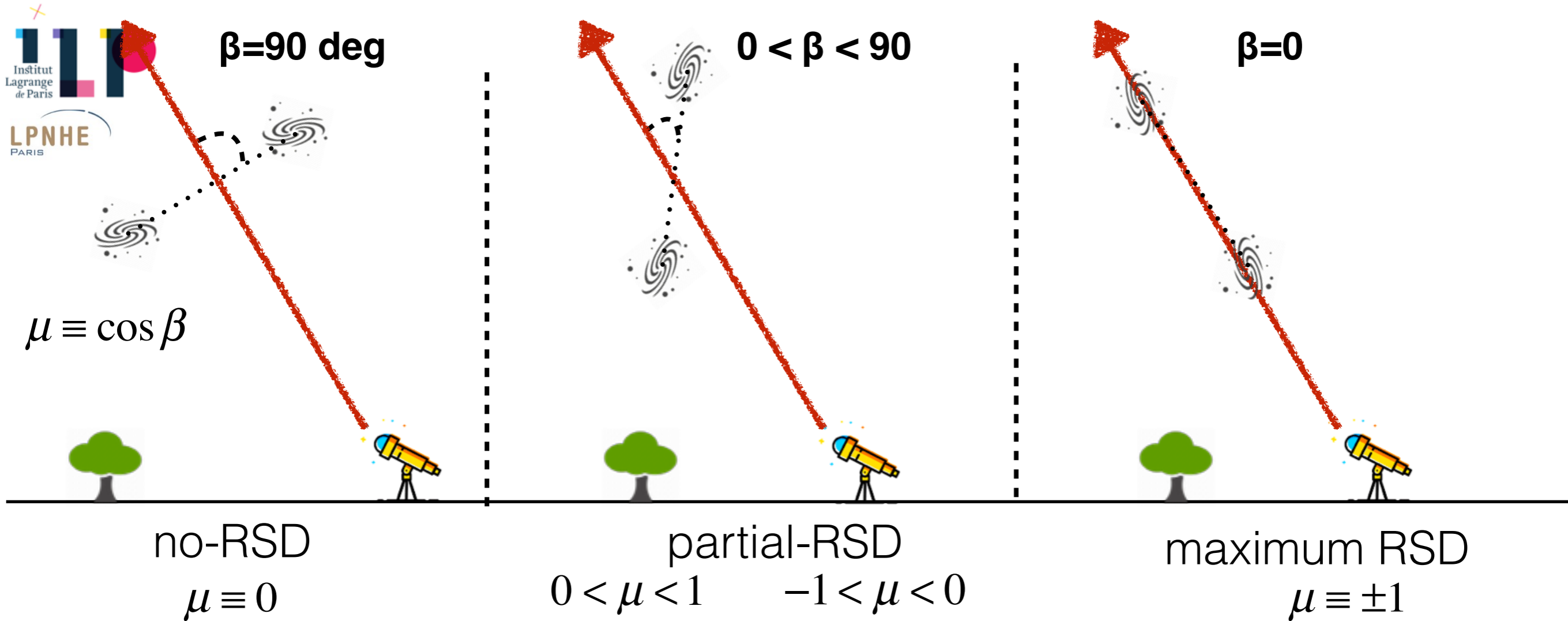
- Universe assumed **isotropic** and **homogeneous**
- **RSD**: Enhancement / reduction of the clustering along the line-of-sight (LOS) direction due to peculiar velocities (Kaiser 1987)



## 1. Hubble flow

## 2. Coherent with growth of structure





$$P_g^{(s)}(k, \mu; z) = \underbrace{[b(z) + f(z)\mu^2]^2}_{\text{boost}} P_m(k; z) \quad (\text{Kaiser 87})$$

- Break isotropy:  $\mu$ -dependence: monopole, quadrupole, hexadecapole
- Modify the clustering: boost depends on  $f(z)$
- **Does not** change shape of  $P_m(k)$

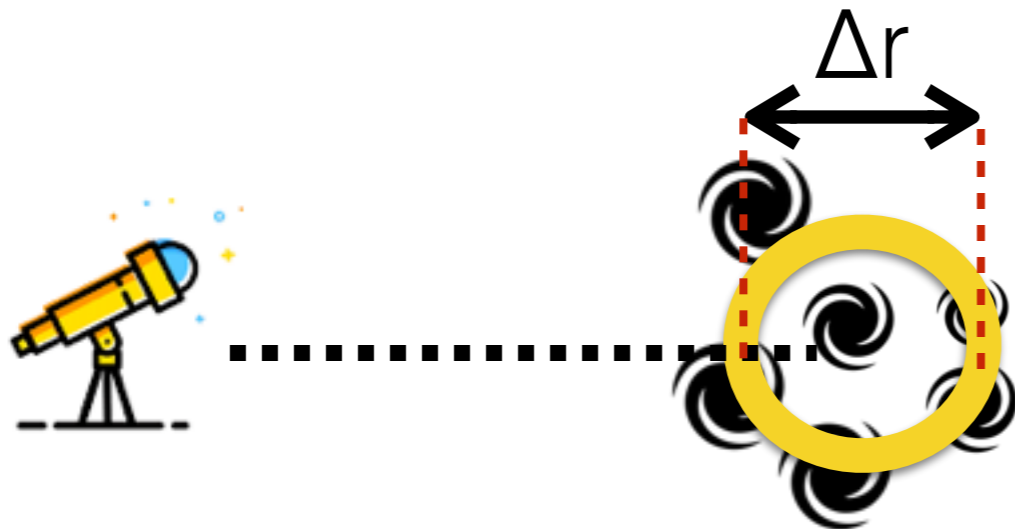
# Alcock-Paczynski effect

- Universe assumed **isotropic** and **homogeneous**
- **AP effect:** Anisotropy induced by transforming redshifts into comoving distances assuming a wrong cosmology

Both transverse and longitudinal modes are modified by  $\Omega_m$

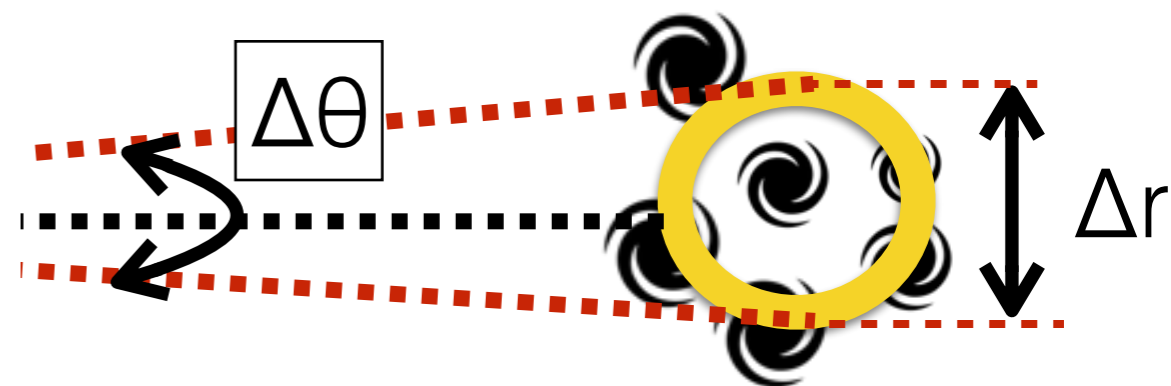
## Radial distance

$$\Delta r_{\parallel}(z_1, z_2; \Omega_m) = \int_{z_1}^{z_2} \frac{cdz'}{H_0 \sqrt{\Omega_m (1+z')^3 + 1 - \Omega_m}} \approx \frac{c\Delta z}{H(\bar{z}, \Omega_m)}$$



## Angular diameter distance

$$\Delta r_{\perp}(\theta_1, \theta_2; z, \Omega_m) = \Delta\theta \int_0^z \frac{cdz'}{H(z', \Omega_m)}$$



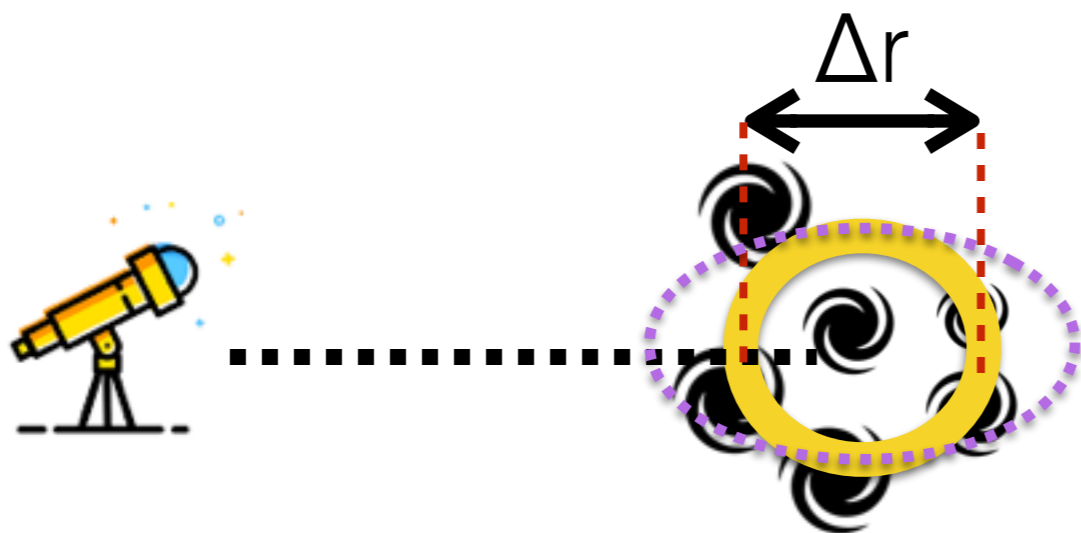
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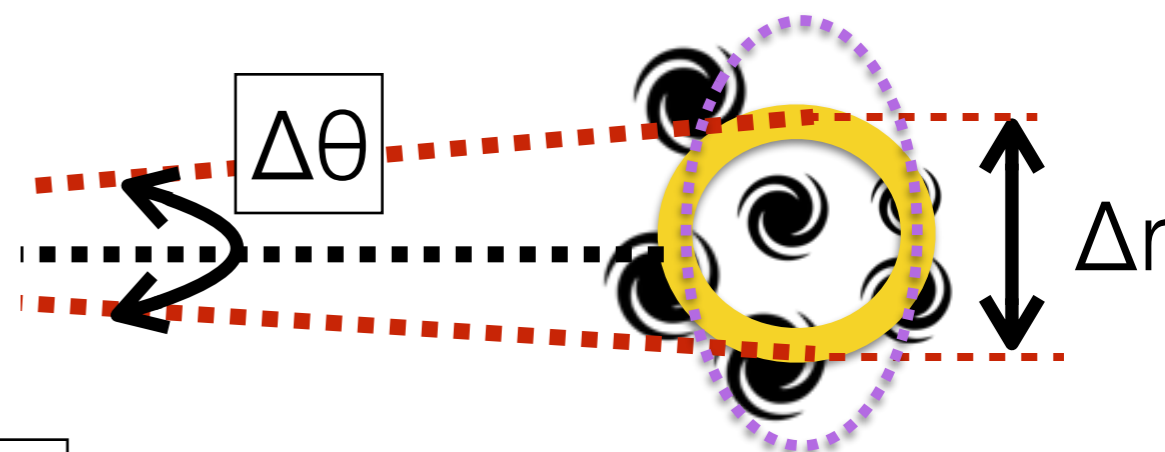
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## Angular diameter distance

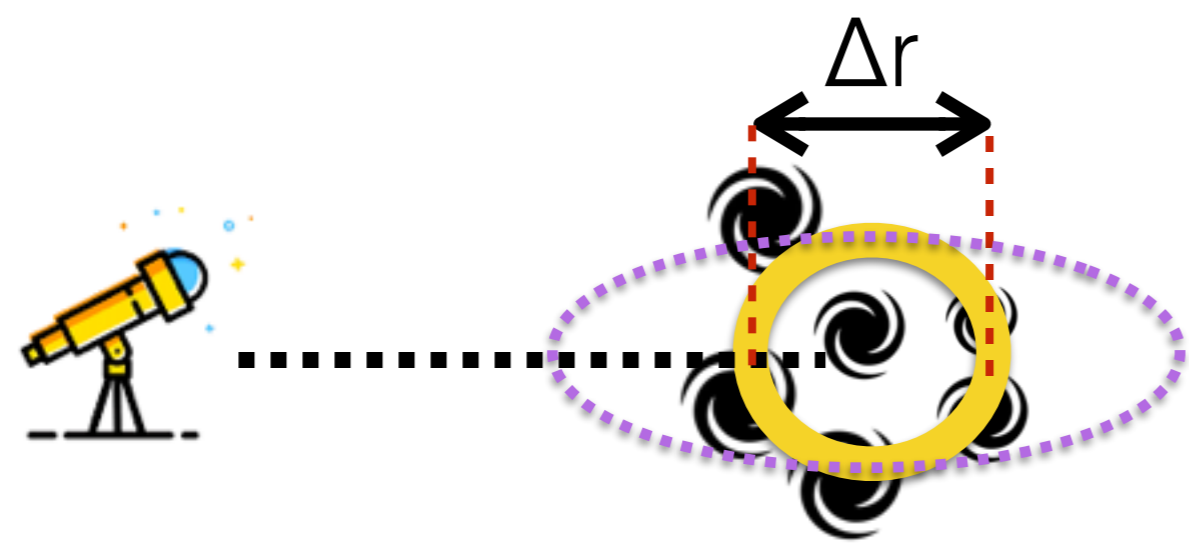
$$\Delta r_{\perp}(\theta_1, \theta_2; z, \Omega_m) = \Delta\theta \int_0^z \frac{cdz'}{H(z', \Omega_m)}$$



cluster is modified along both directions

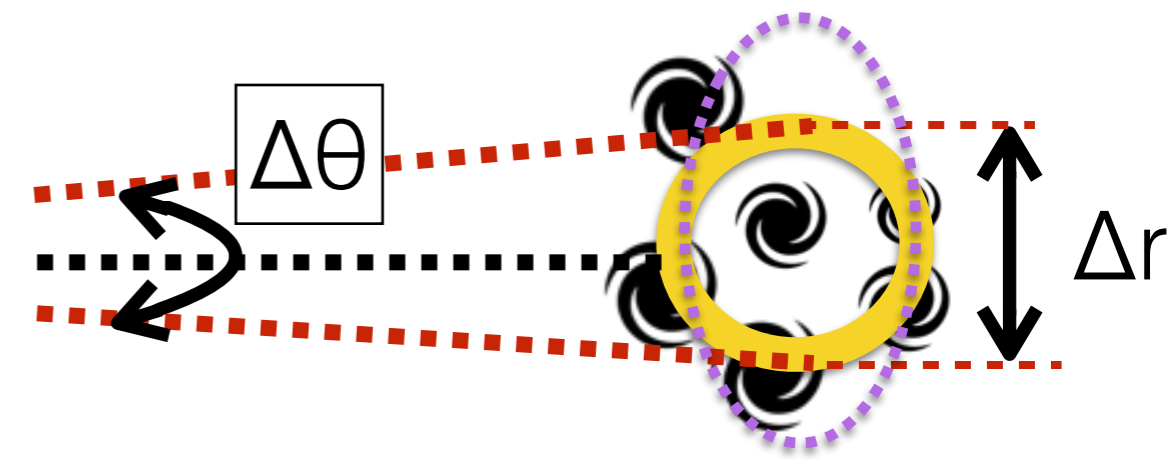
### Radial distance

$$\Delta r_{\parallel}(z_1, z_2; \Omega_m) = \int_{z_1}^{z_2} \frac{cdz'}{H_0 \sqrt{\Omega_m (1+z')^3 + 1 - \Omega_m}} \approx \frac{c\Delta z}{H(\bar{z}, \Omega_m)}$$



### Angular diameter distance

$$\Delta r_{\perp}(\theta_1, \theta_2; z, \Omega_m) = \Delta\theta \int_0^z \frac{cdz'}{H(z', \Omega_m)}$$

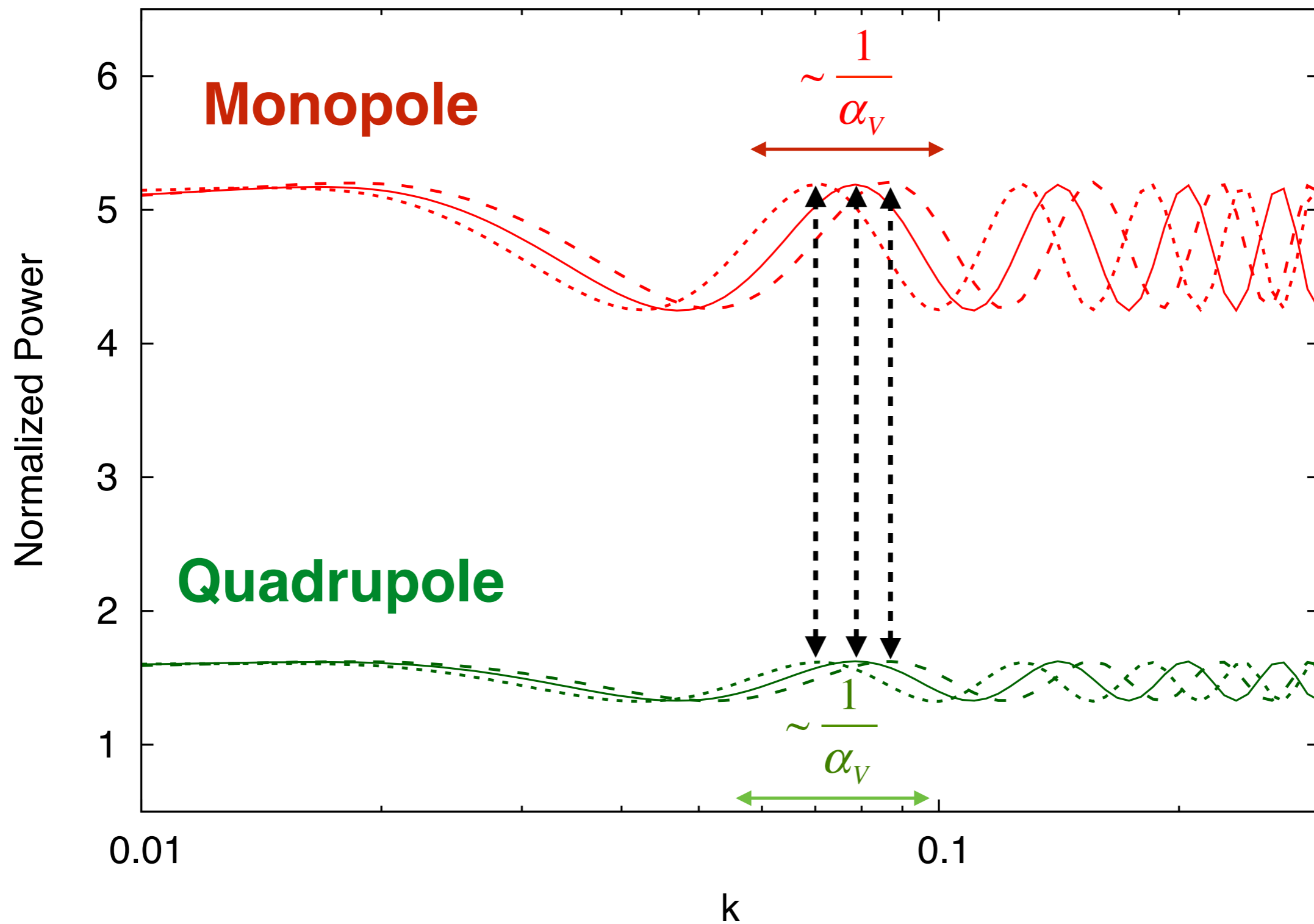


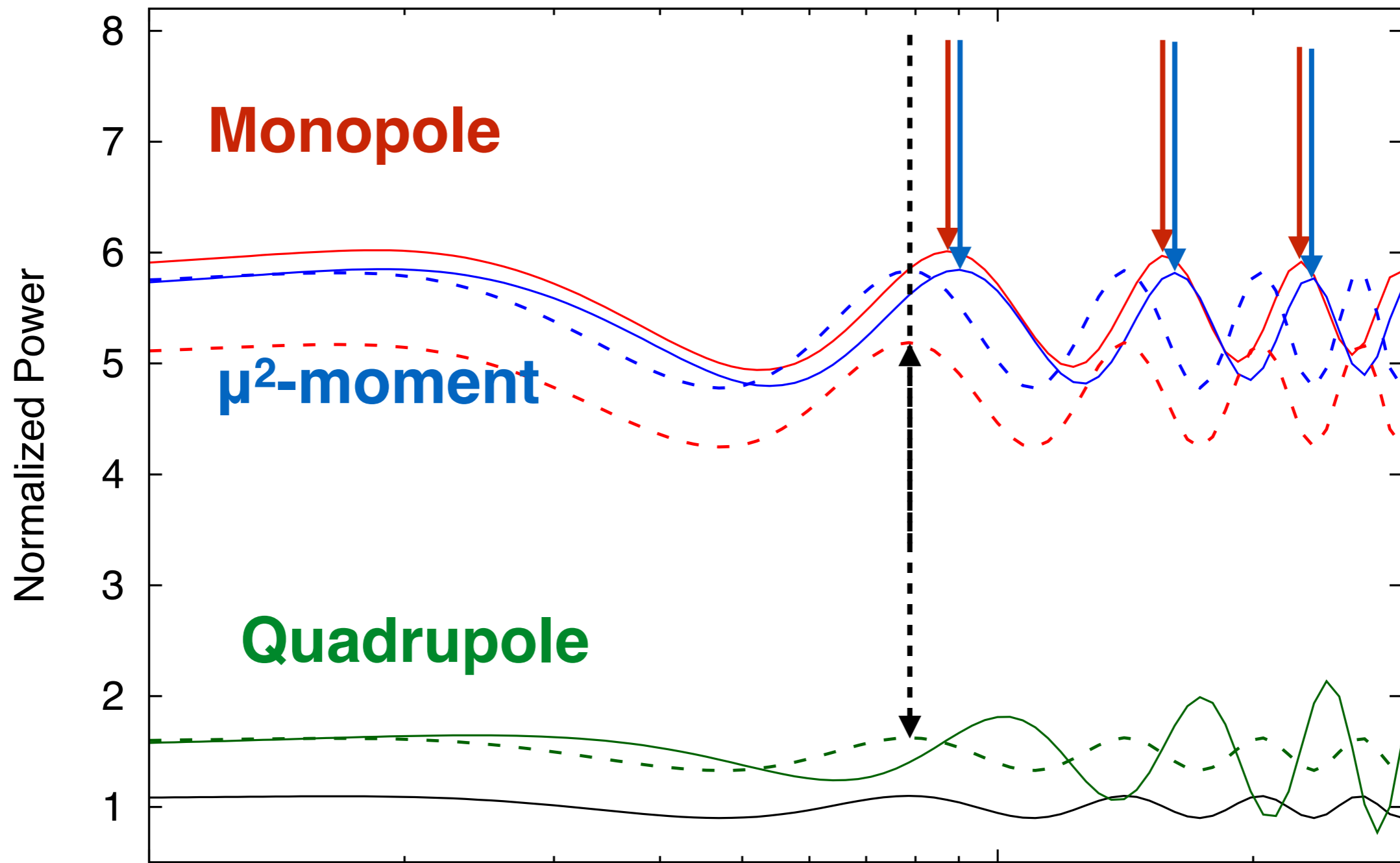
$$\begin{aligned}
 k_{\parallel} &= \alpha_{\parallel} k'_{\parallel} \\
 k_{\perp} &= \alpha_{\perp} k'_{\perp}
 \end{aligned}
 \longrightarrow
 \begin{array}{l}
 \alpha_{\parallel} = \frac{r_s^{fid} H^{fid}}{r_s H} \\
 \alpha_{\perp} = \frac{D_A / r_s}{D_A^{fid} / r_s^{fid}}
 \end{array}
 \xrightarrow{\text{new variables}}
 \begin{array}{l}
 \alpha_V = \sqrt[3]{\alpha_{\perp}^2 \alpha_{\parallel}} \\
 \alpha_{\varepsilon} = \alpha_{\perp} / \alpha_{\parallel}
 \end{array}
 \begin{array}{l}
 \text{Isotropic dilation} \\
 \text{AP-effect}
 \end{array}$$

$\alpha_V$  modify the  $k$  vector in the monopole  
 $\alpha_{\varepsilon}$  generates an anisotropy (distort symmetric 3D-features along and across the LOS)

# Physical Interpretation

$$\alpha_V \sim (D_A^2/H)^{1/3} / r_s$$

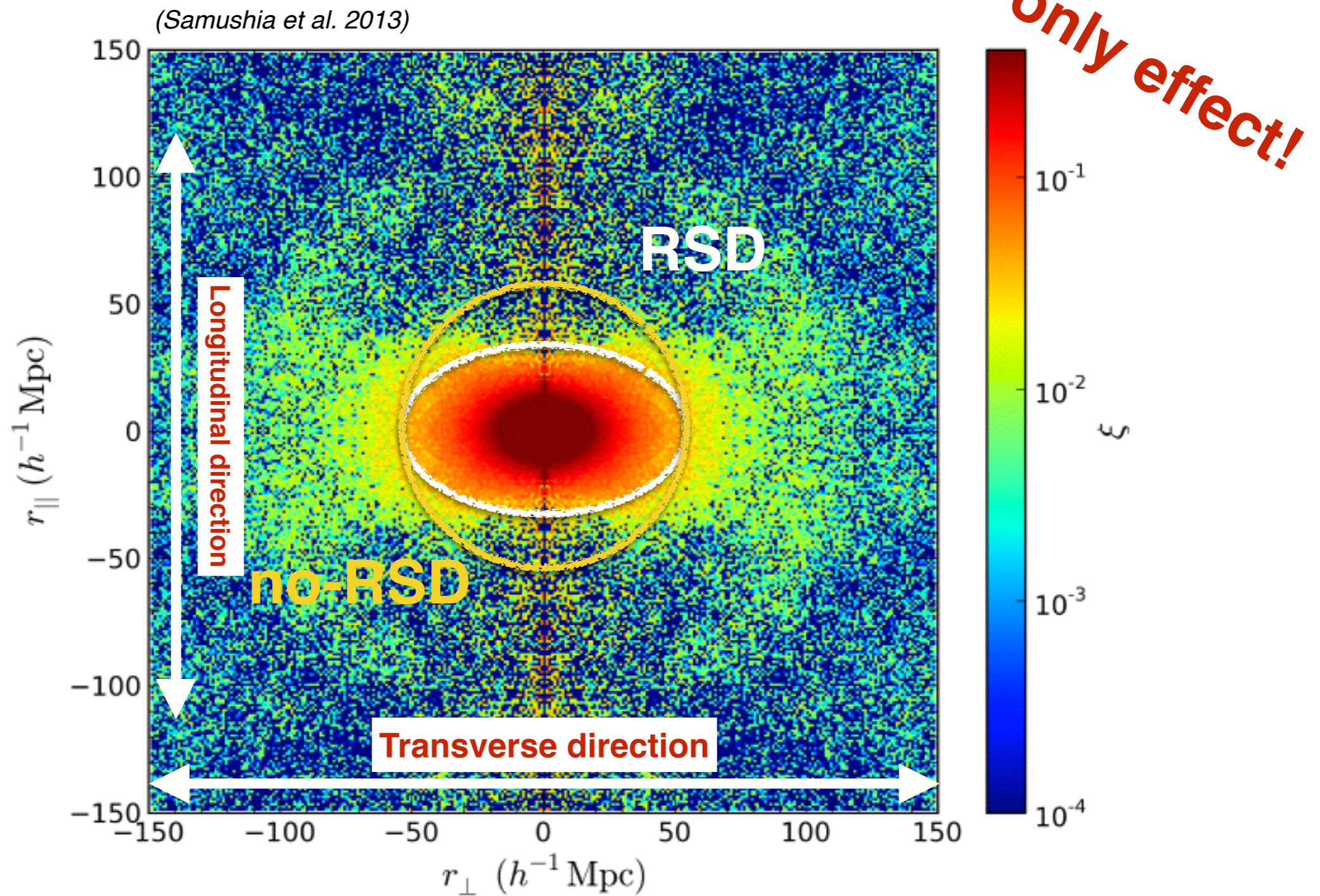




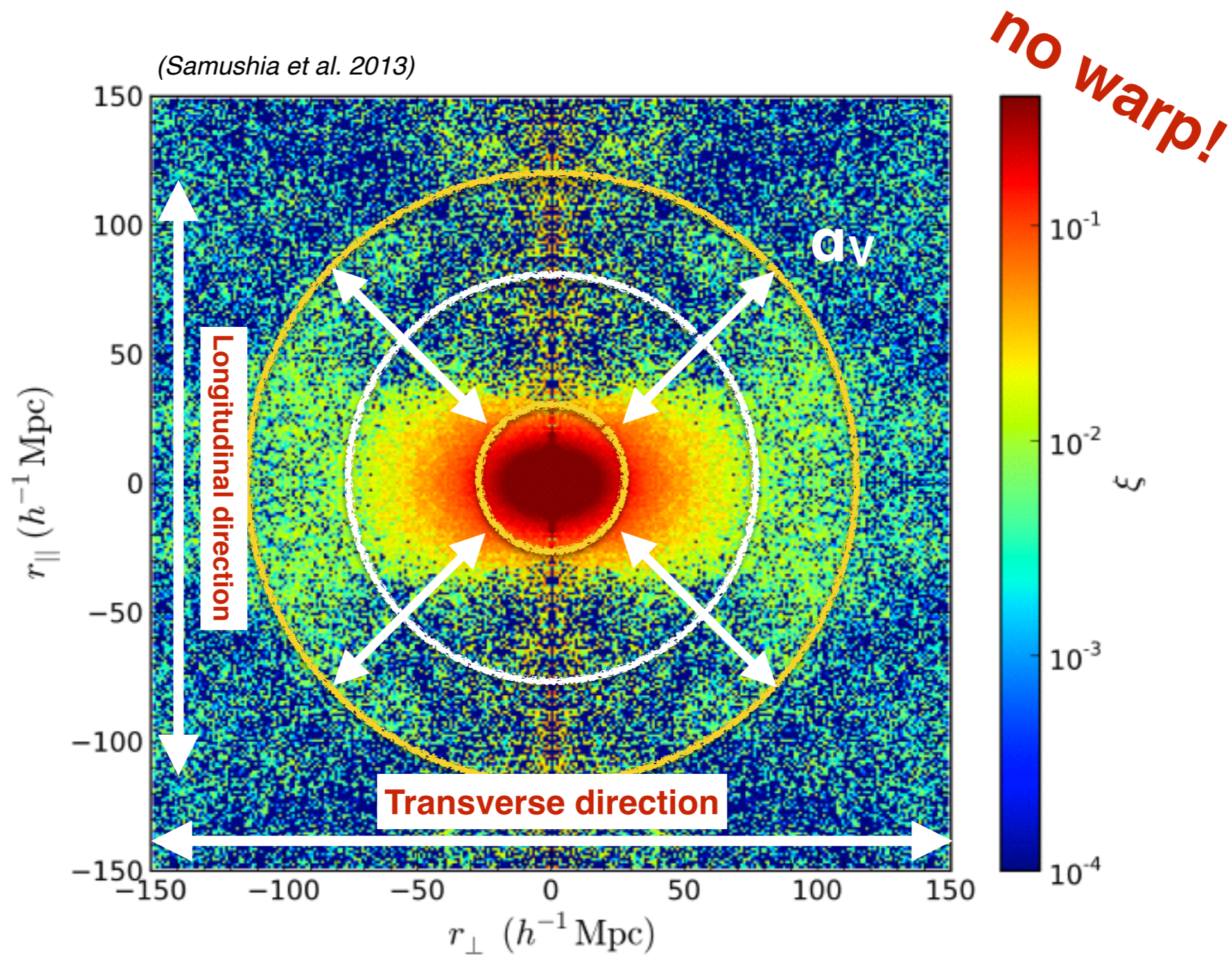
$$P^{(\mu^2)} \equiv P^{(0)} + \frac{2}{5} P^{(2)}$$

k

# Physical Interpretation

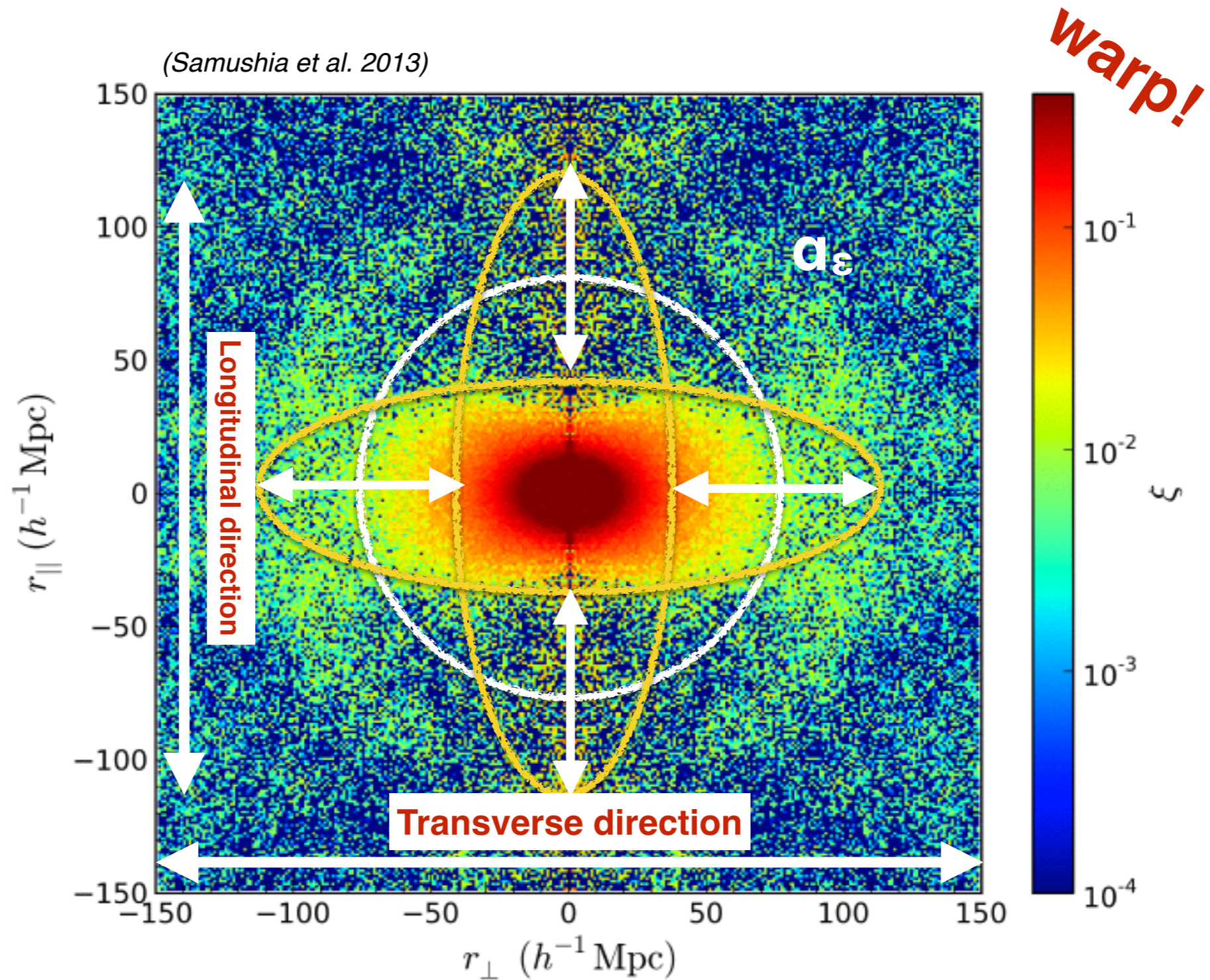


# Physical Interpretation





# Physical Interpretation

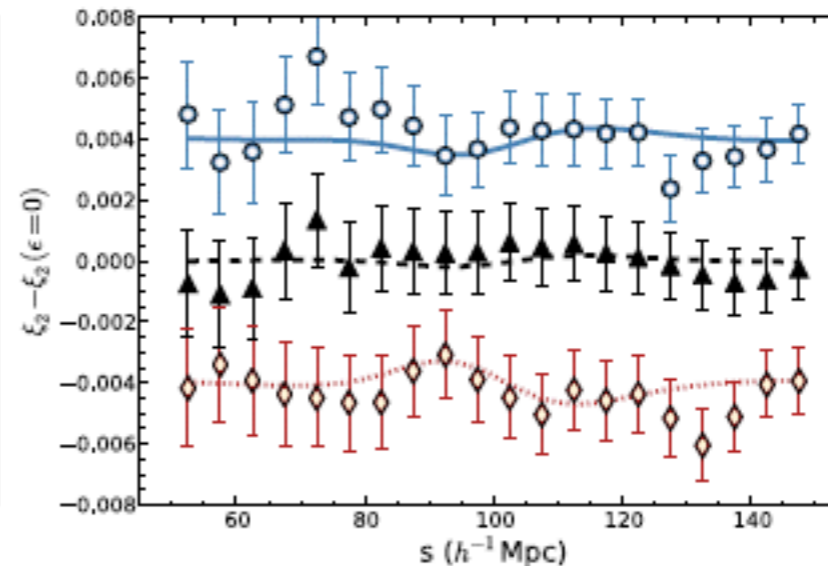
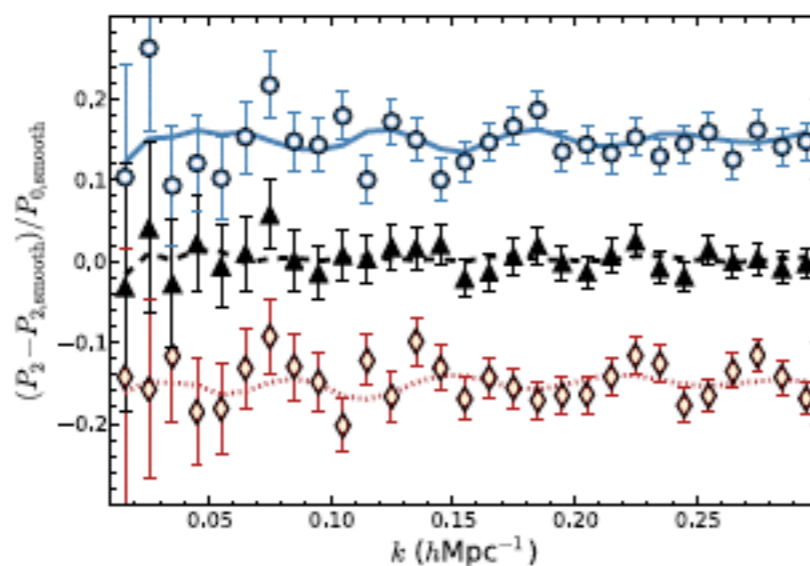
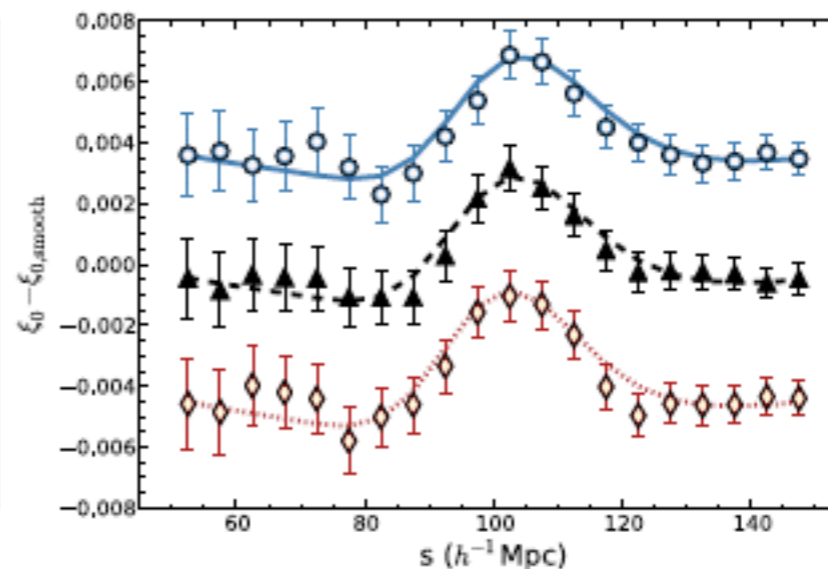
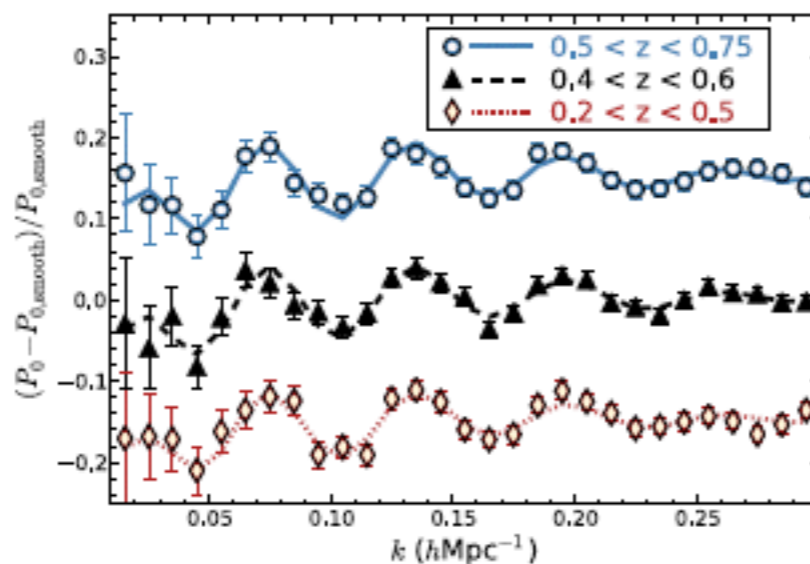


RSD & AP are degenerated!

# LRGs BAO/RSD from BOSS

Power Spectrum

Correlation Function



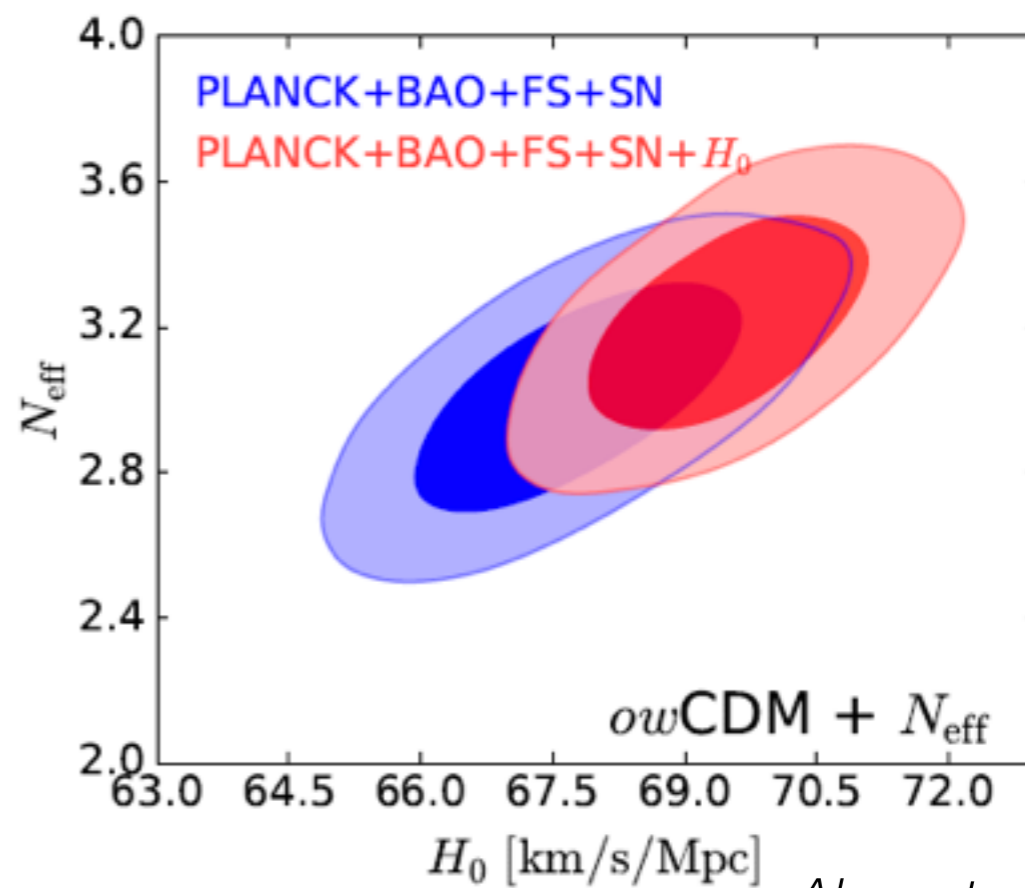
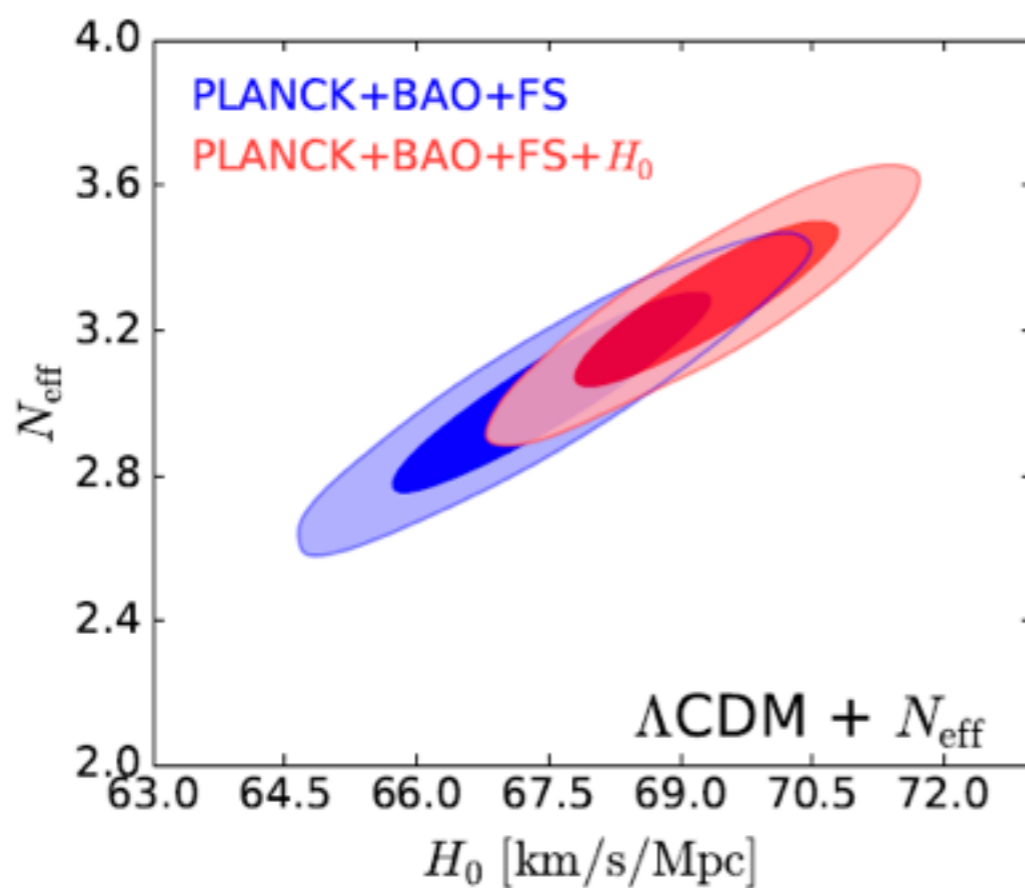
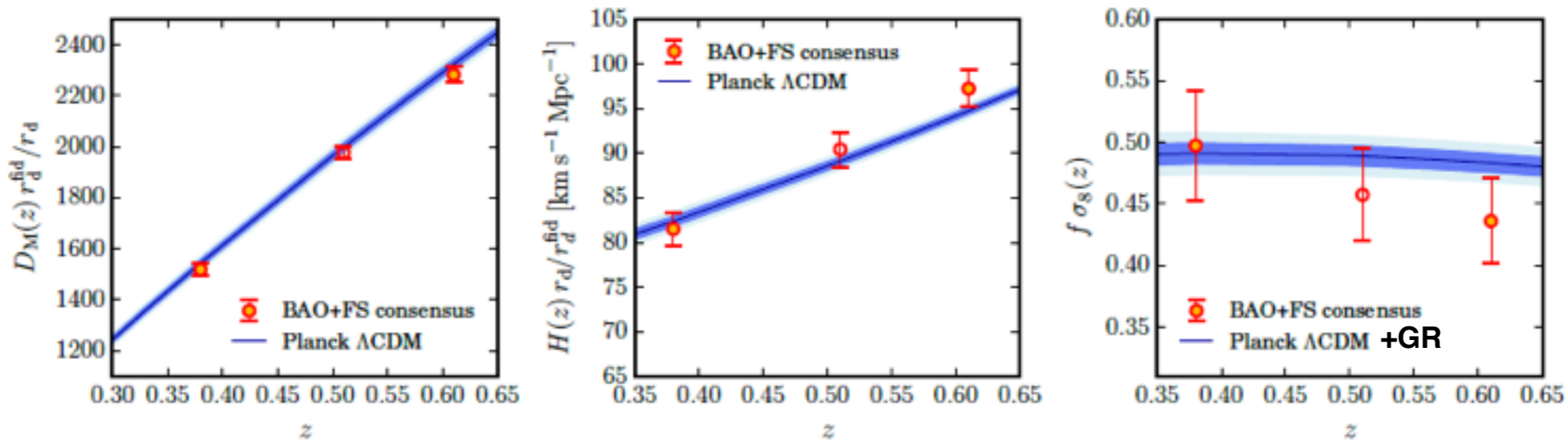
3 overlapping z-bins

**0.2 < z < 0.5**  
**0.4 < z < 0.6**  
**0.5 < z < 0.75**

*Alam et al. 2016*

Type of analyses:

- RSD (full shape) vs. BAO,
- pre-recon vs. post-recon,
- configuration space vs. Fourier space

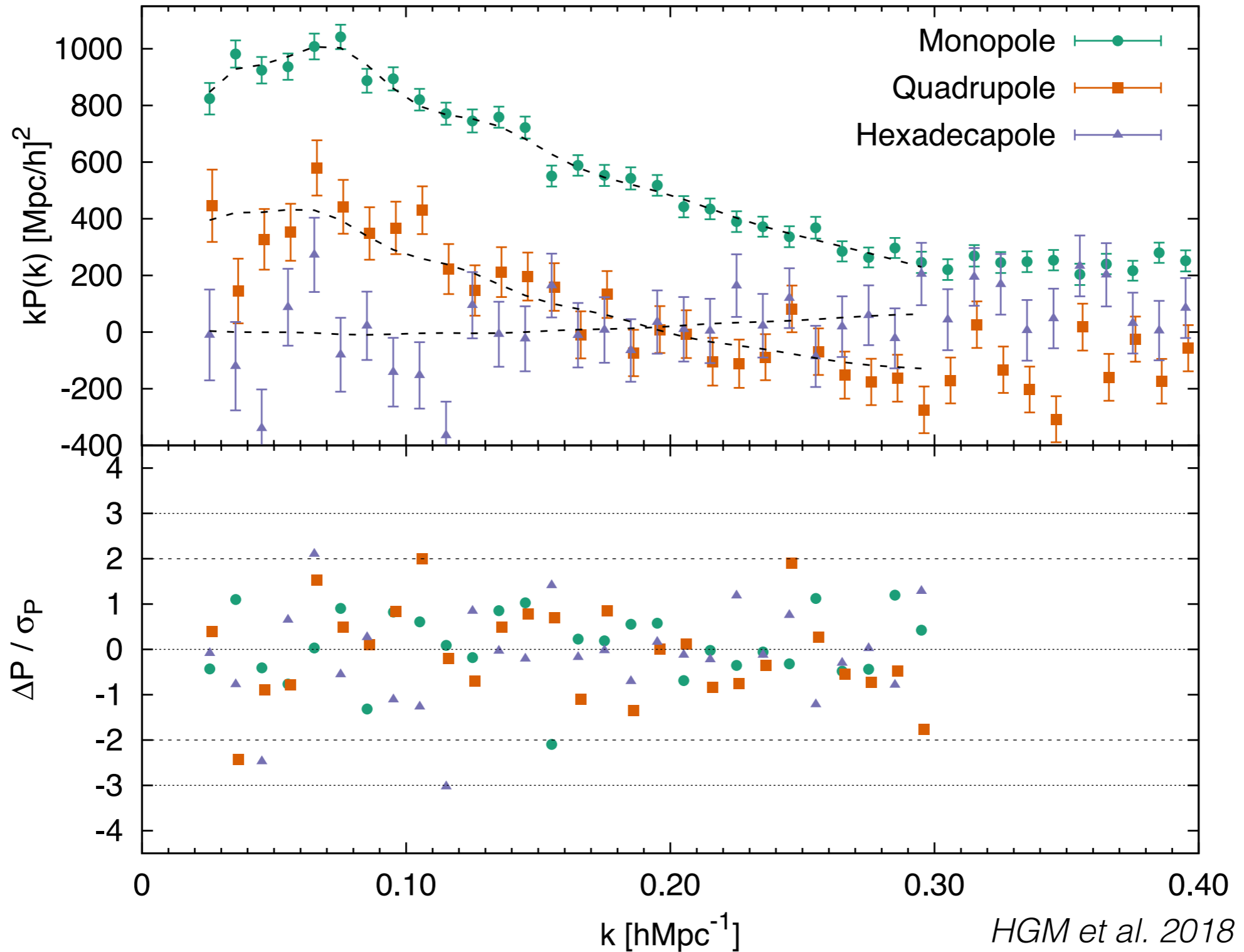


$H_0 = 73.0 \pm 1.8 \text{ km s}^{-1} \text{ Mpc}^{-1}$  measurement of Riess et al.

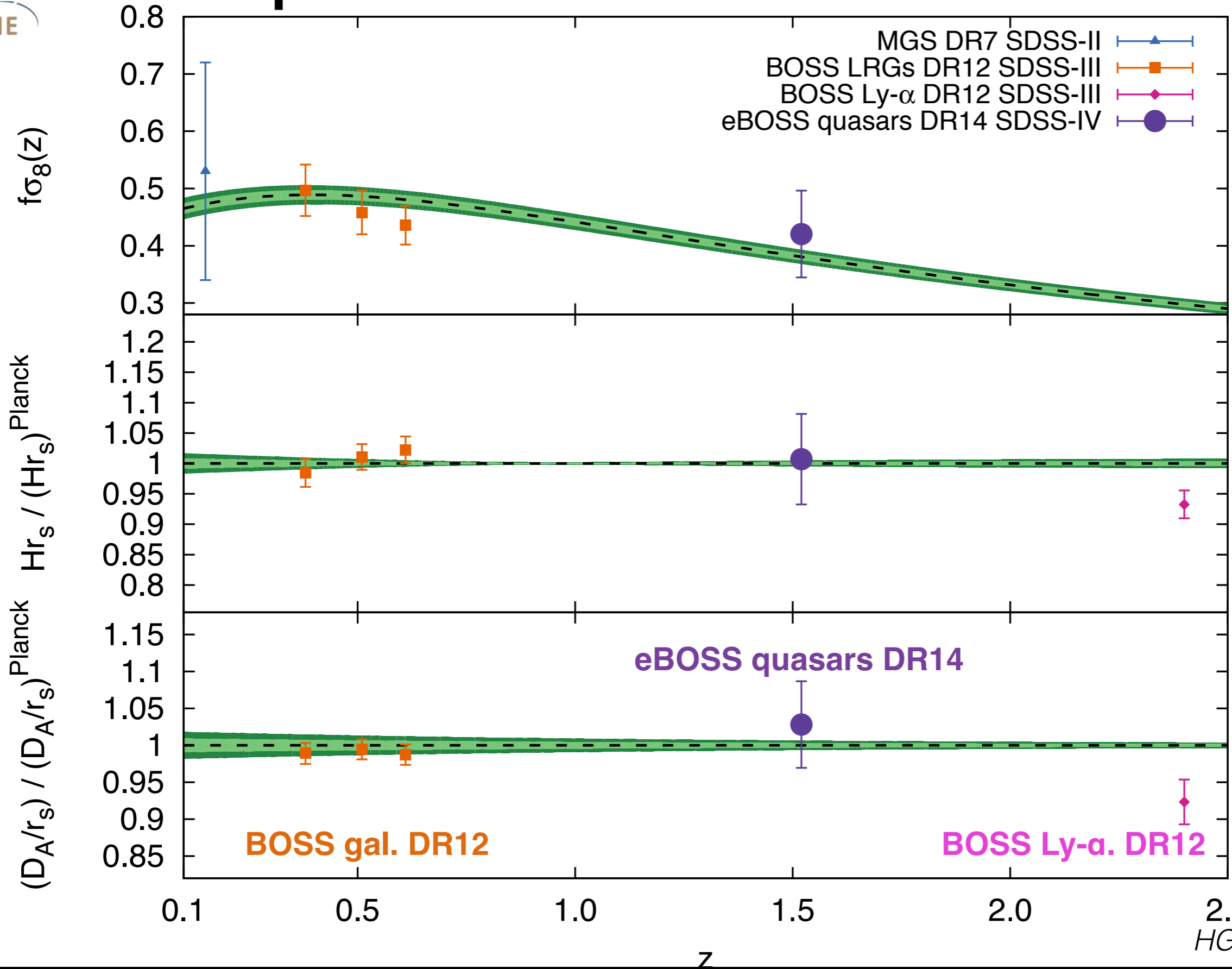
Alam et al. 2016

# quasar BAO/RSD from eBOSS

DR14Q  $0.8 < z < 2.2$

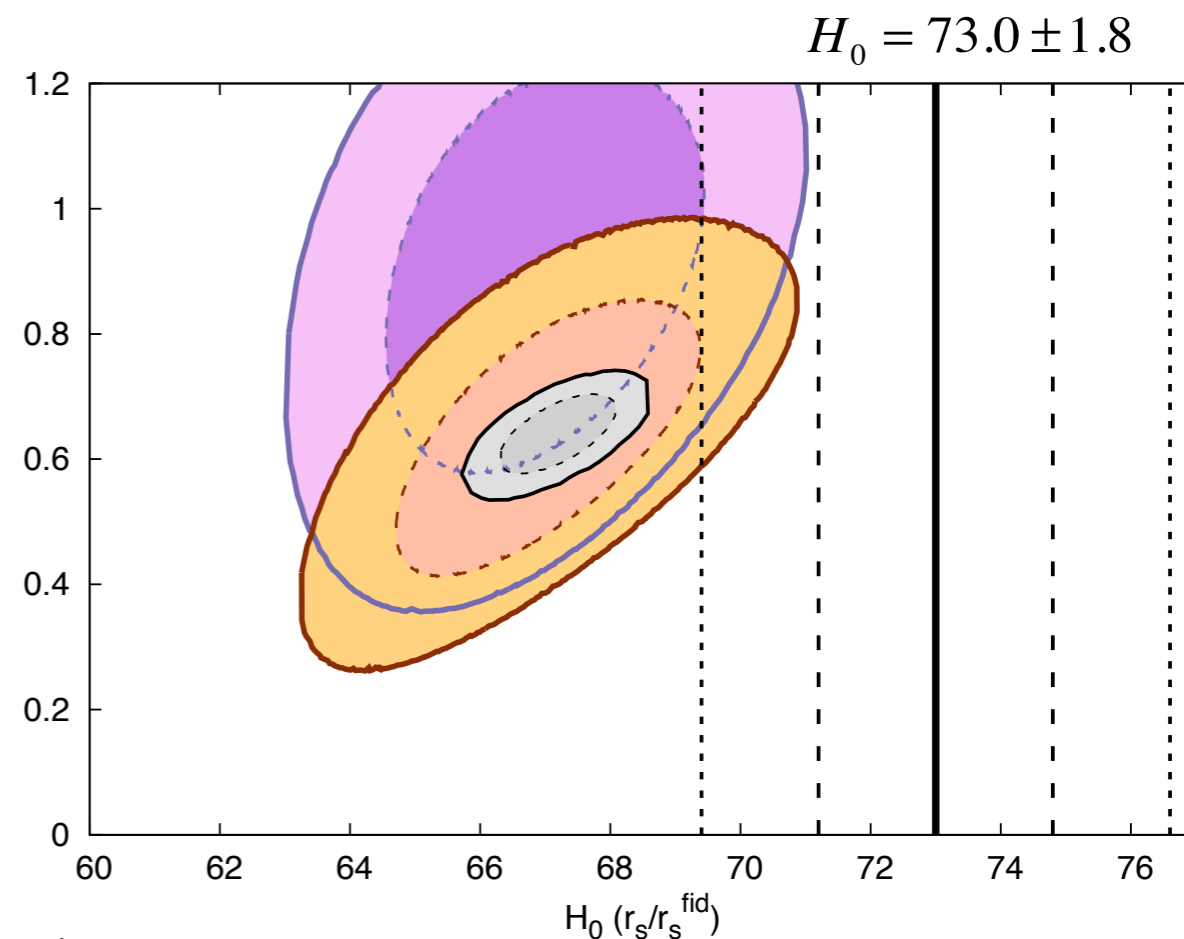
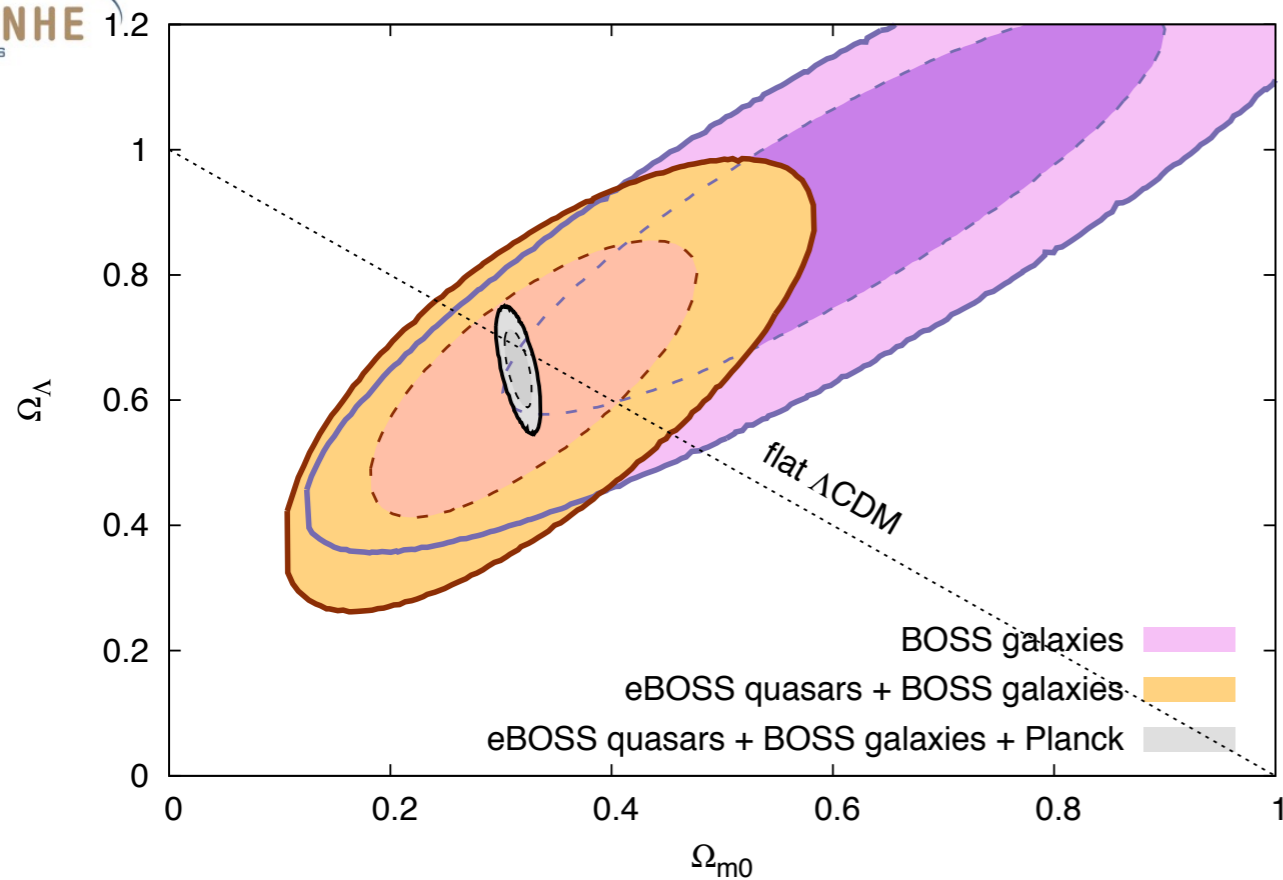


# quasar BAO/RSD from eBOSS



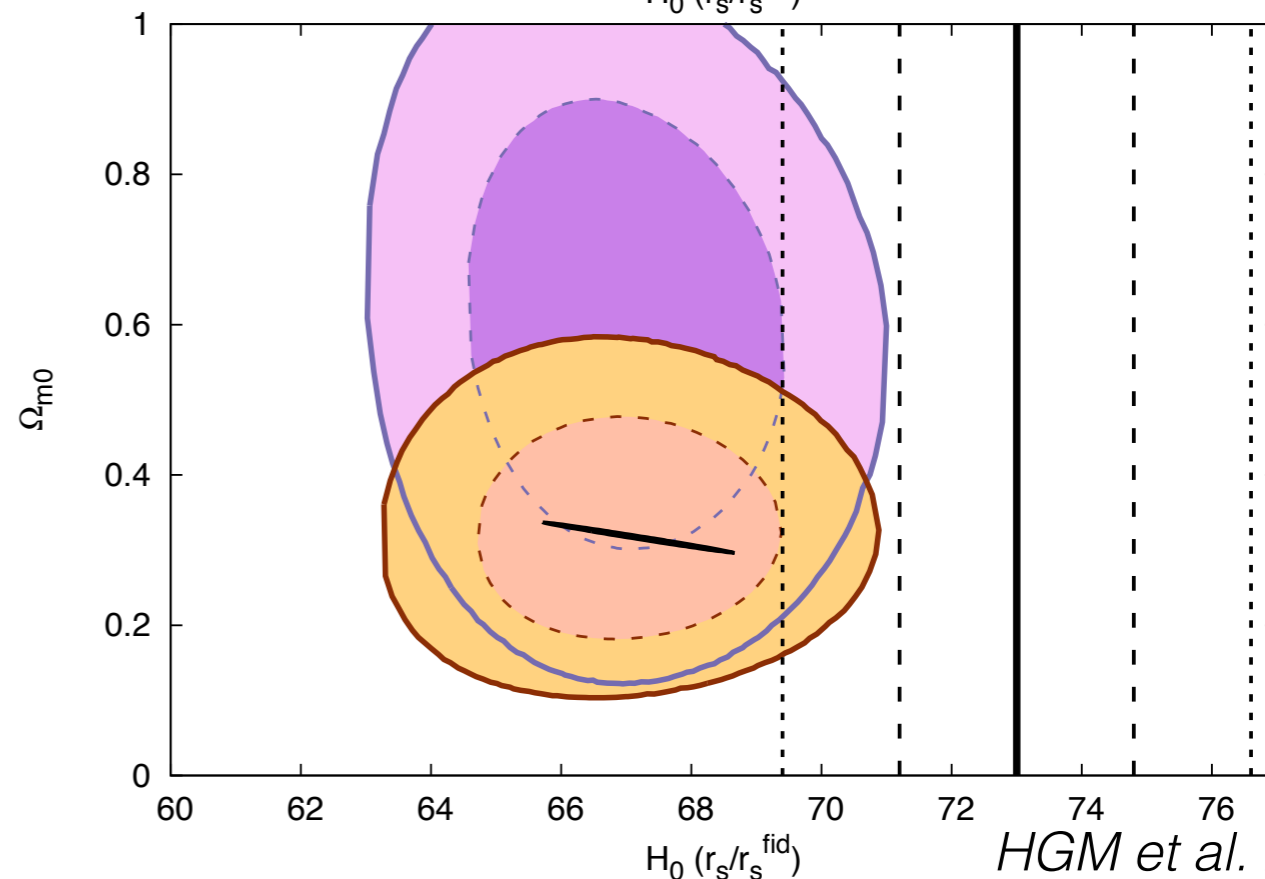
HGM et al. 2018

# Tensions with $H_0$



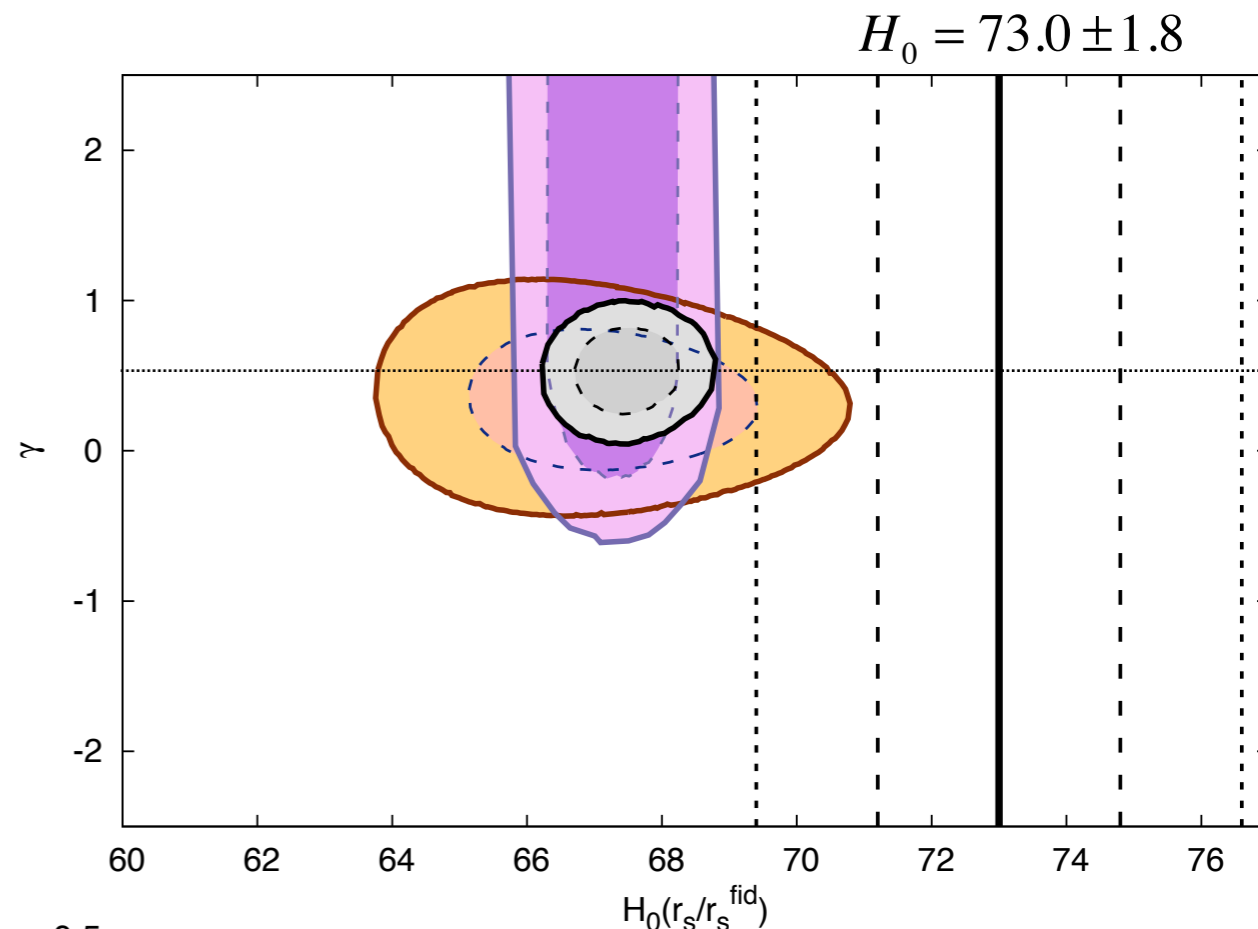
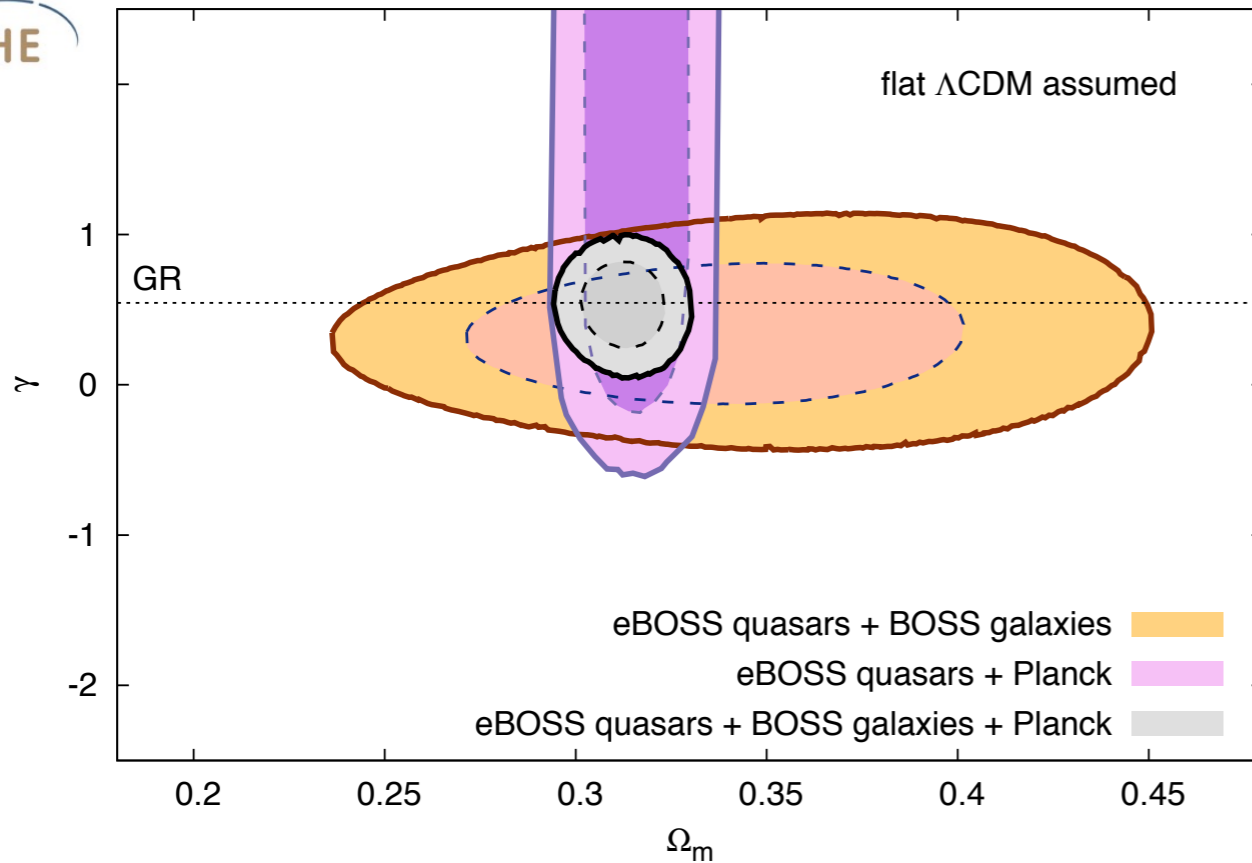
**GR &  $\Lambda$ CDM assumed**  
**relax flatness**

$D_A(z)$ ,  $H(z)$ ,  $f\sigma_8(z)$  from BOSS galaxies / eBOSS quasars



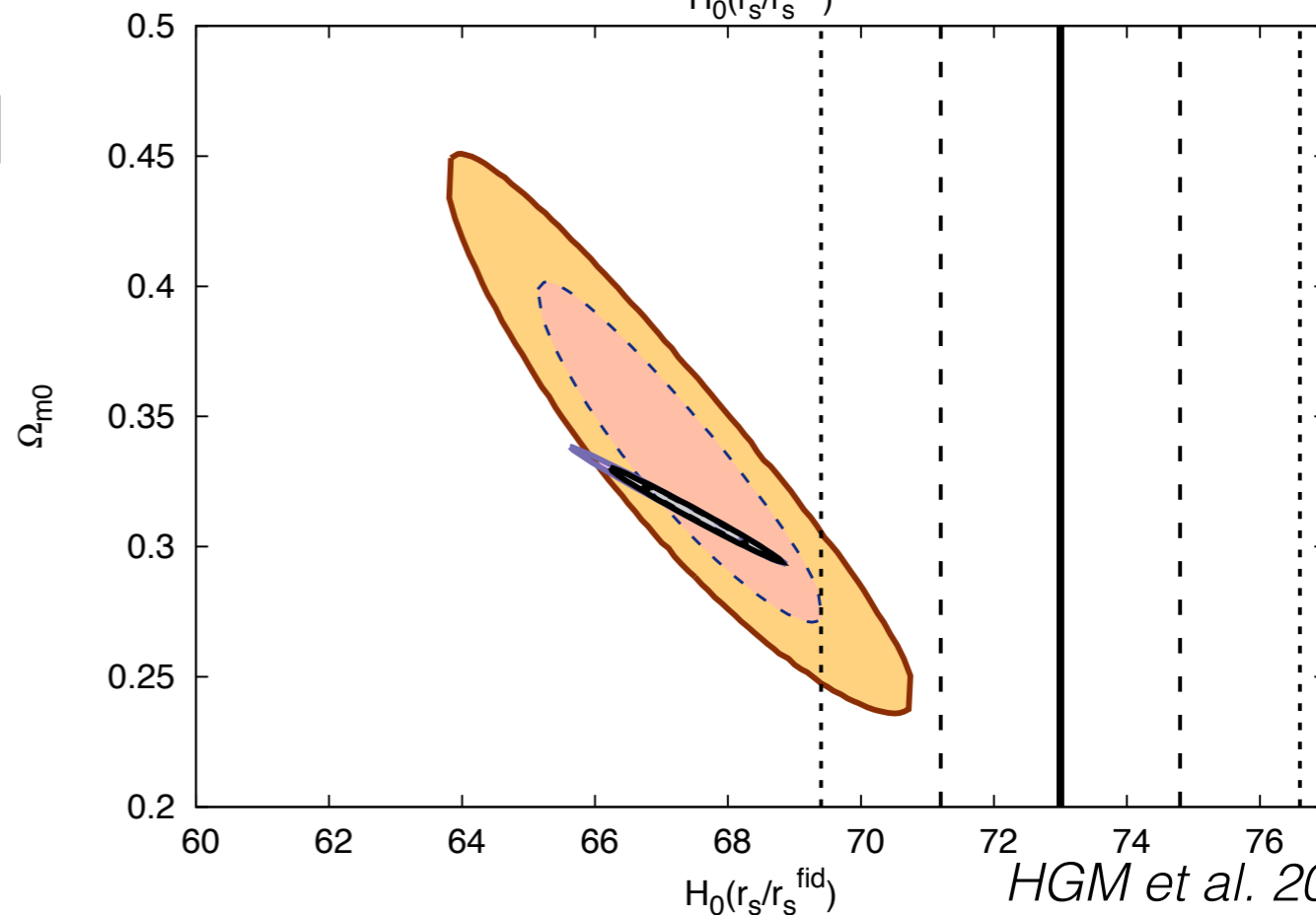
*HGM et al. 2018*

# Tensions with $H_0$



**flatness &  $\Lambda$ CDM assumed**  
**relax GR**

$D_A(z)$ ,  $H(z)$ ,  $f\sigma_8(z)$  from BOSS galaxies / eBOSS quasars



# Summary

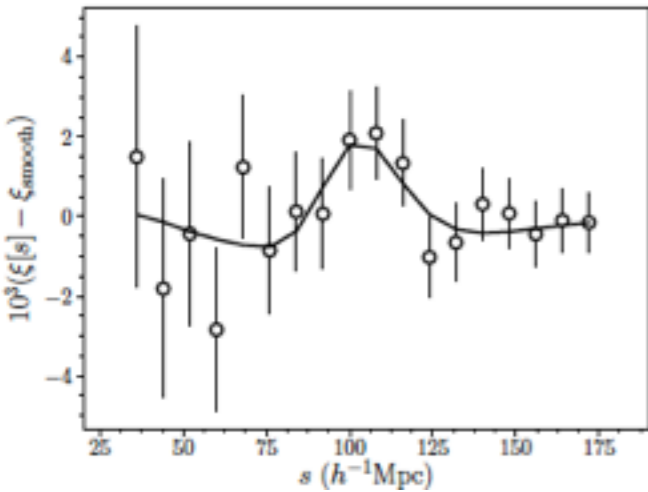
- BOSS measured  $H(z)r_s$  in the range  $0.2 < z < 0.75$  using LRGs (high density)
- **First results from eBOSS** quasars measuring  $H(z)r_s$  in the range  $0.8 < z < 2.2$
- If  $r_s$  taken from Planck or BBN,  $H(z)r_s$  from eBOSS+BOSS galaxies in agreement with Planck+ $\Lambda$ CDM ('tension' with cosmic ladder at  $\sim 3\sigma$ )
- This disagreement doesn't reduce when flatness, GR,  $N_{\text{eff}}=3$  conditions are relaxed
- More eBOSS data coming in the next 1.5yr (ELG, LRG, quasars + Ly $\alpha$ ): errors of quasars to  $\sim 1/2$ , measurements at  $z \sim 0.75$  from LRGs + ELGs
- Ly- $\alpha$  BAO at  $z=2.4$  at  $2.3\sigma$  from Planck (see next talk by Andreu)
- DESI will deliver more precise results within 5yr from now.



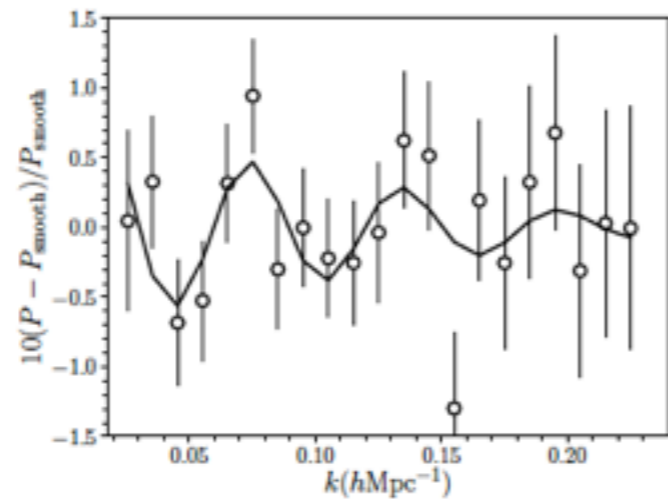
Back up slides

# DR14Q BAO results

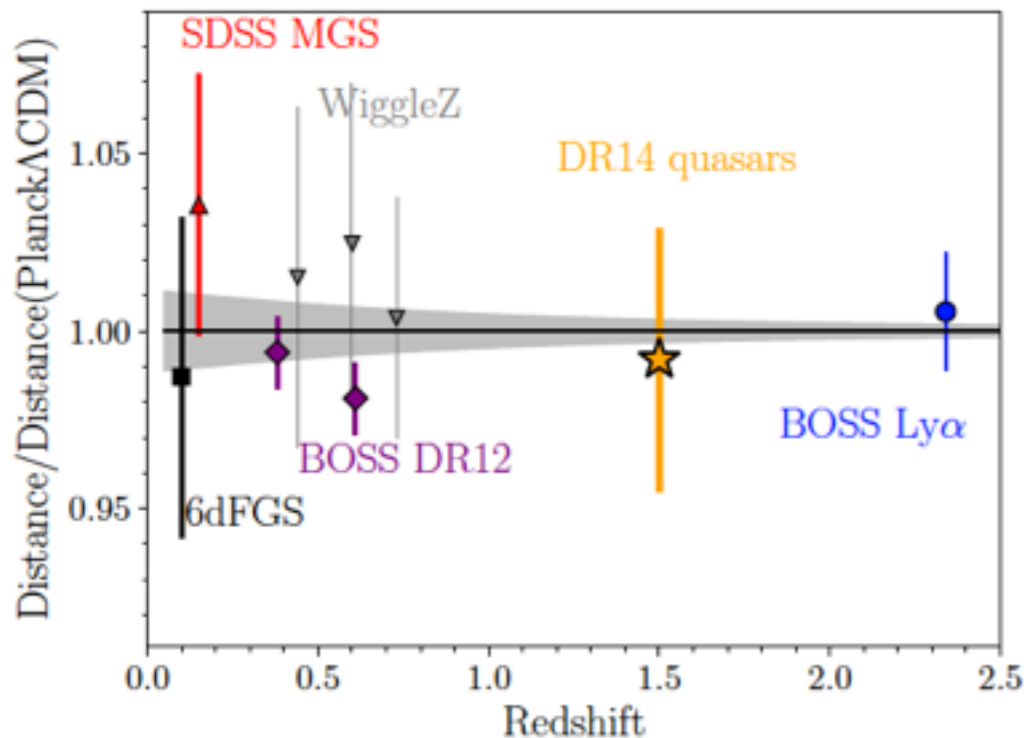
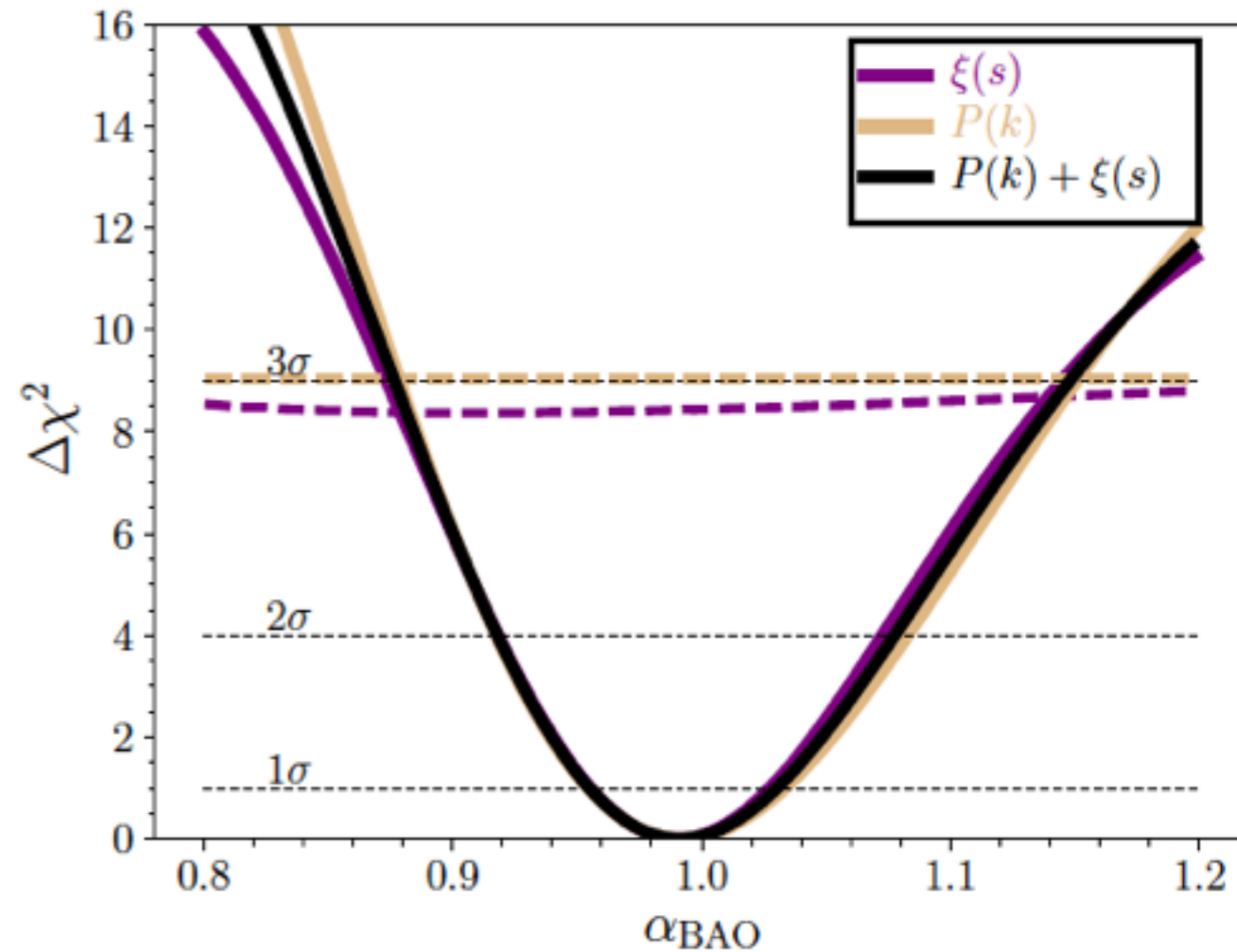
## Correlation Function



## Power Spectrum



## Significance of BAO peak



- Correlation factor  $\rho=0.97$
- $3\sigma$  detection
- In good agreement with Planck+GR
- $D_V(z=1.52)=3843 \pm 147$  Mpc (3.8%)
- $\chi^2=6.2/13$  for  $\xi(R)$  and  $27.7/33$  for  $P(k)$

Ata et al. 2017

# Impact of potential systematics

## BAO Systematics

- Very robust
- $\sim 0.1\%$  non-linear shift at  $z=1.5$
- relative velocity between DM & bar (?)
- Reconstruction assumptions (bias and  $f$ )

## RSD systematics

- model dependent
- Intrinsic alignments
- failures & collisions

# Impact of systematics

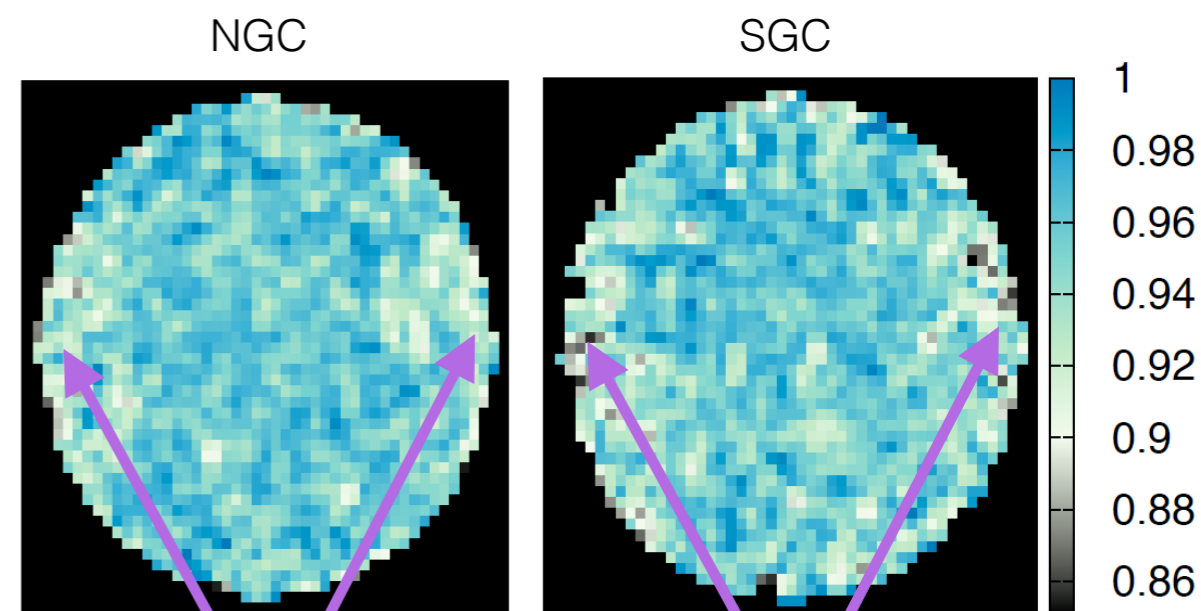
- **Redshift Failures:** *i)* Weight the nearest neighbour (NN), use in BOSS analysis. *ii)* Weight all observed galaxies by their position in the plate,

$$w_{spec}(x_{foc}, y_{foc}) \sim \frac{1}{P_{sucess}(x_{foc}, y_{foc})}$$

- **Collision Pairs:** Traditional nearest neighbour weighting (NN)

Imprint such effects on the mocks and check how these correction schemes perform

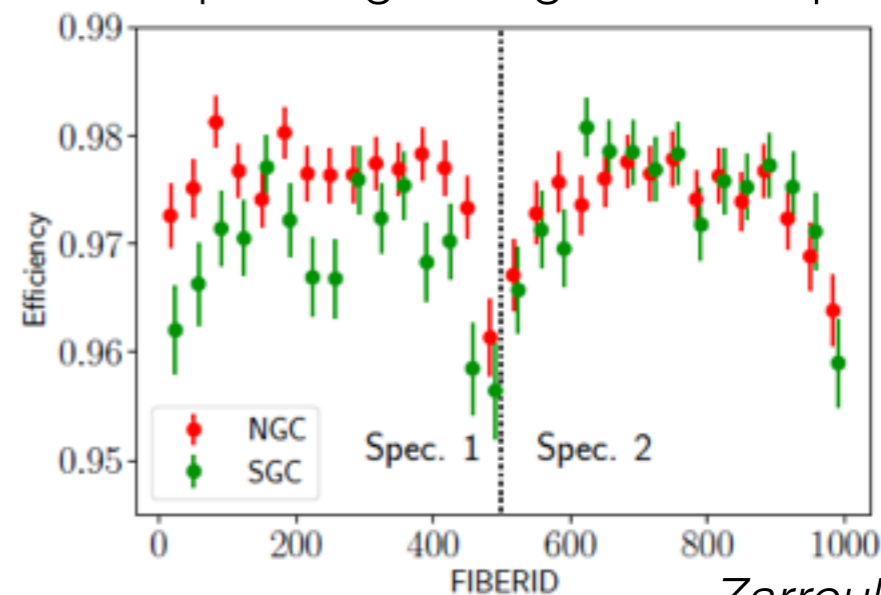
## Redshift efficiency pattern en eBOSS DR14Q



*lower at the edges*

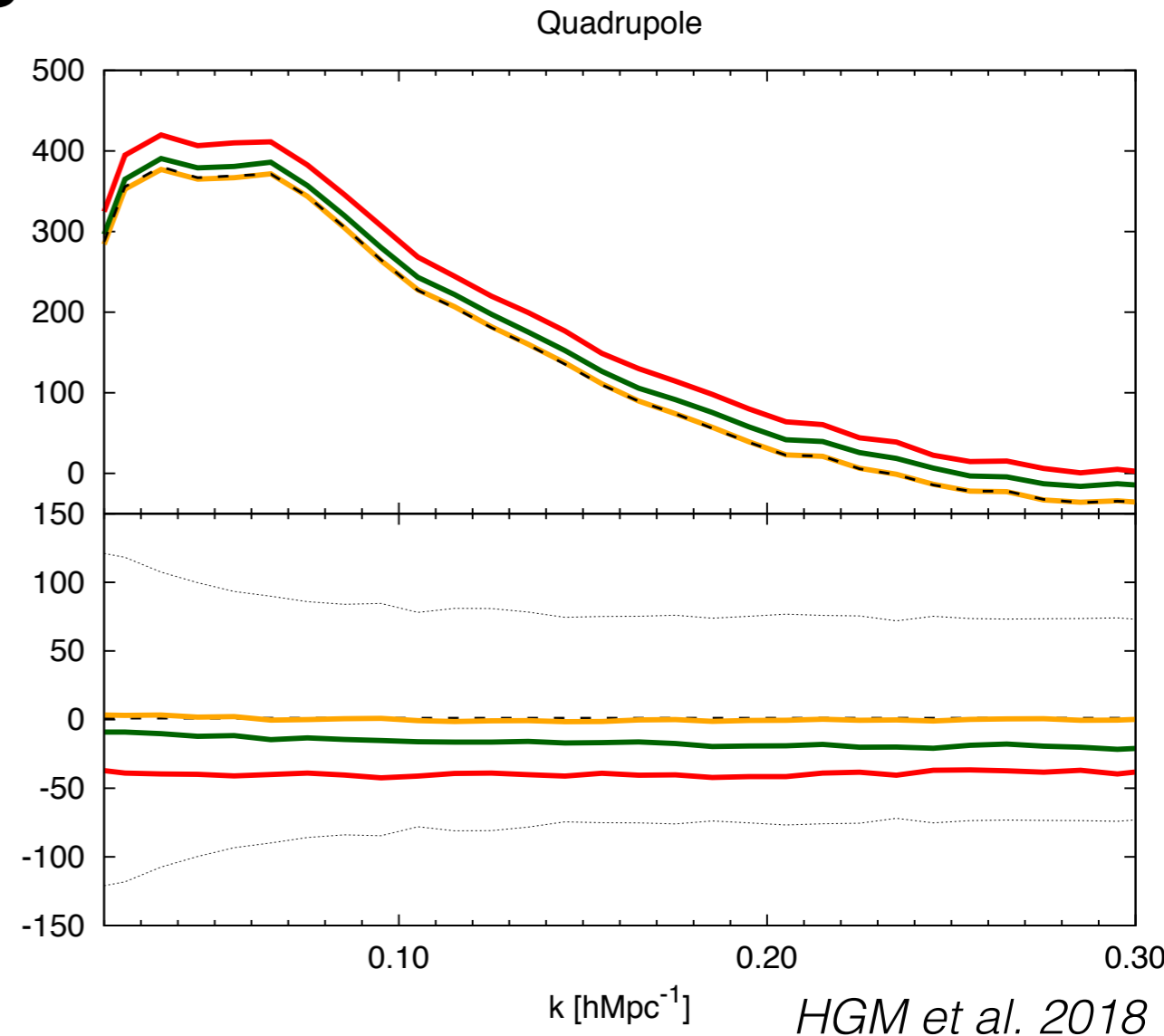
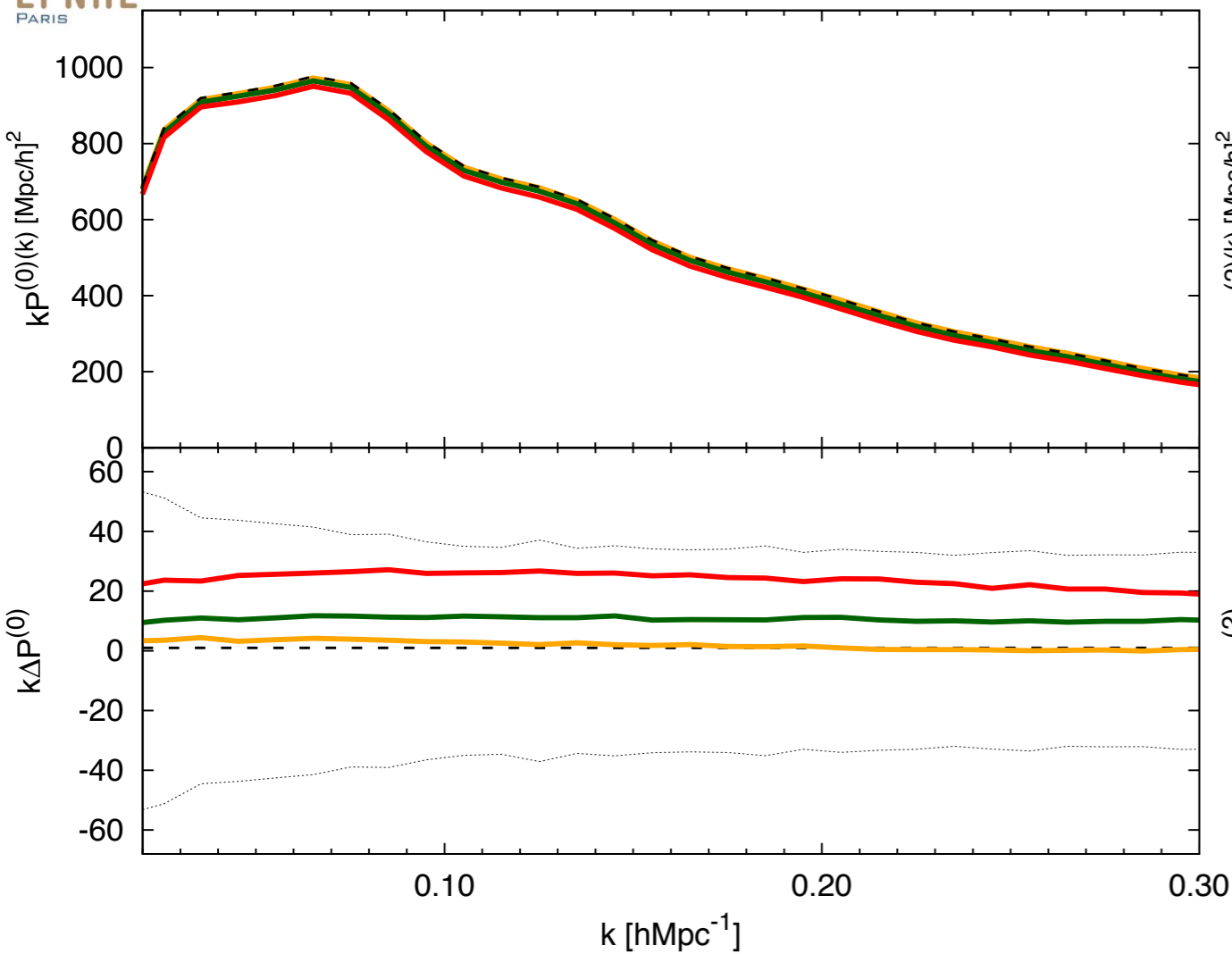
*HGM et al. 2018*

Fibres corresponding to edges of the spectrograph



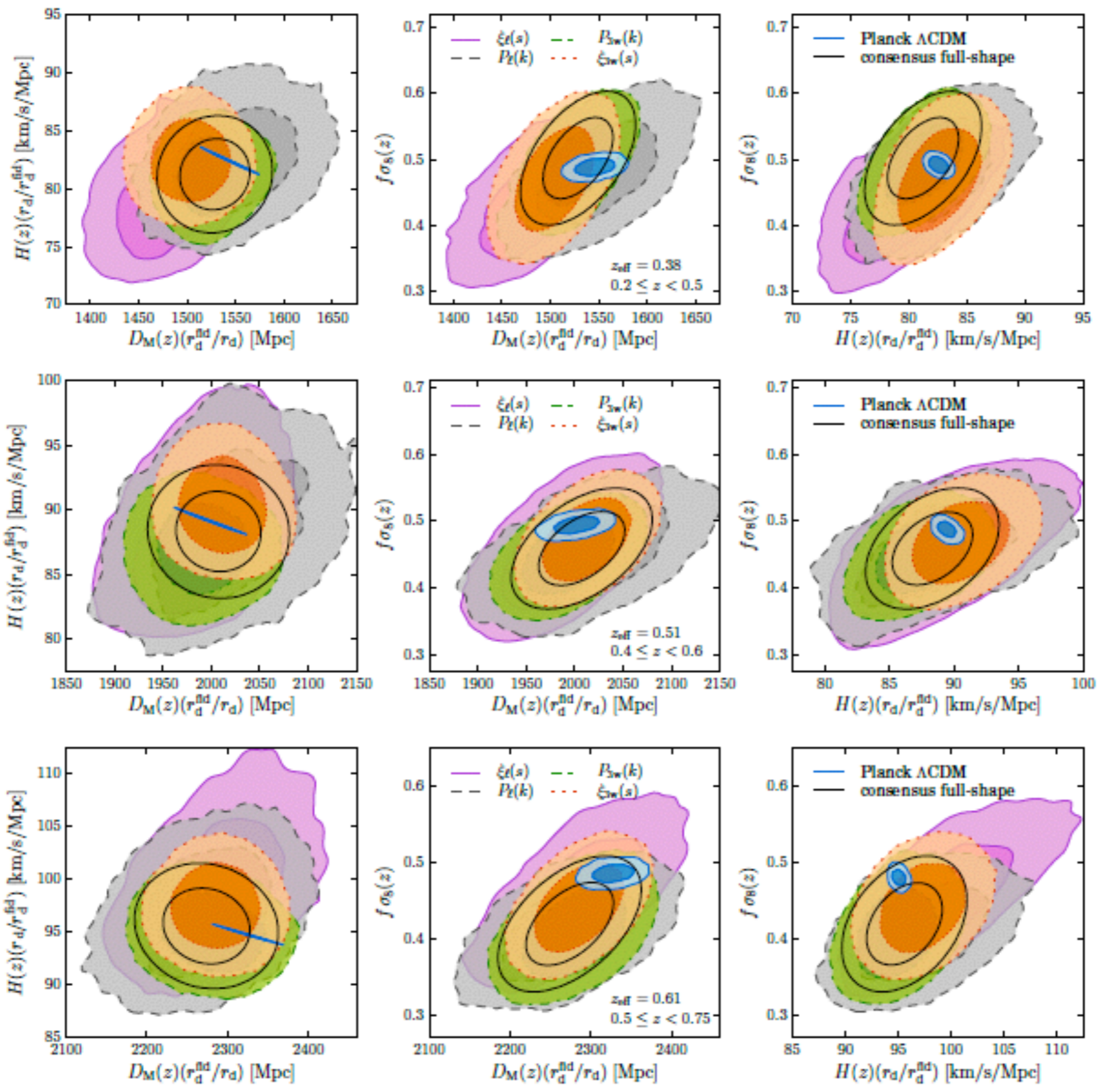
*Zarrouk et al. 2018*

# Impact of systematics



*HGM et al. 2018*

- True signal (systematic effect not applied)
- Corrected: redshift failures (focal weight) + close pairs (NN)
- Corrected: redshift failures (NN) + close pairs (NN)
- Corrected: redshift failures (focal weight) [close pairs not applied]



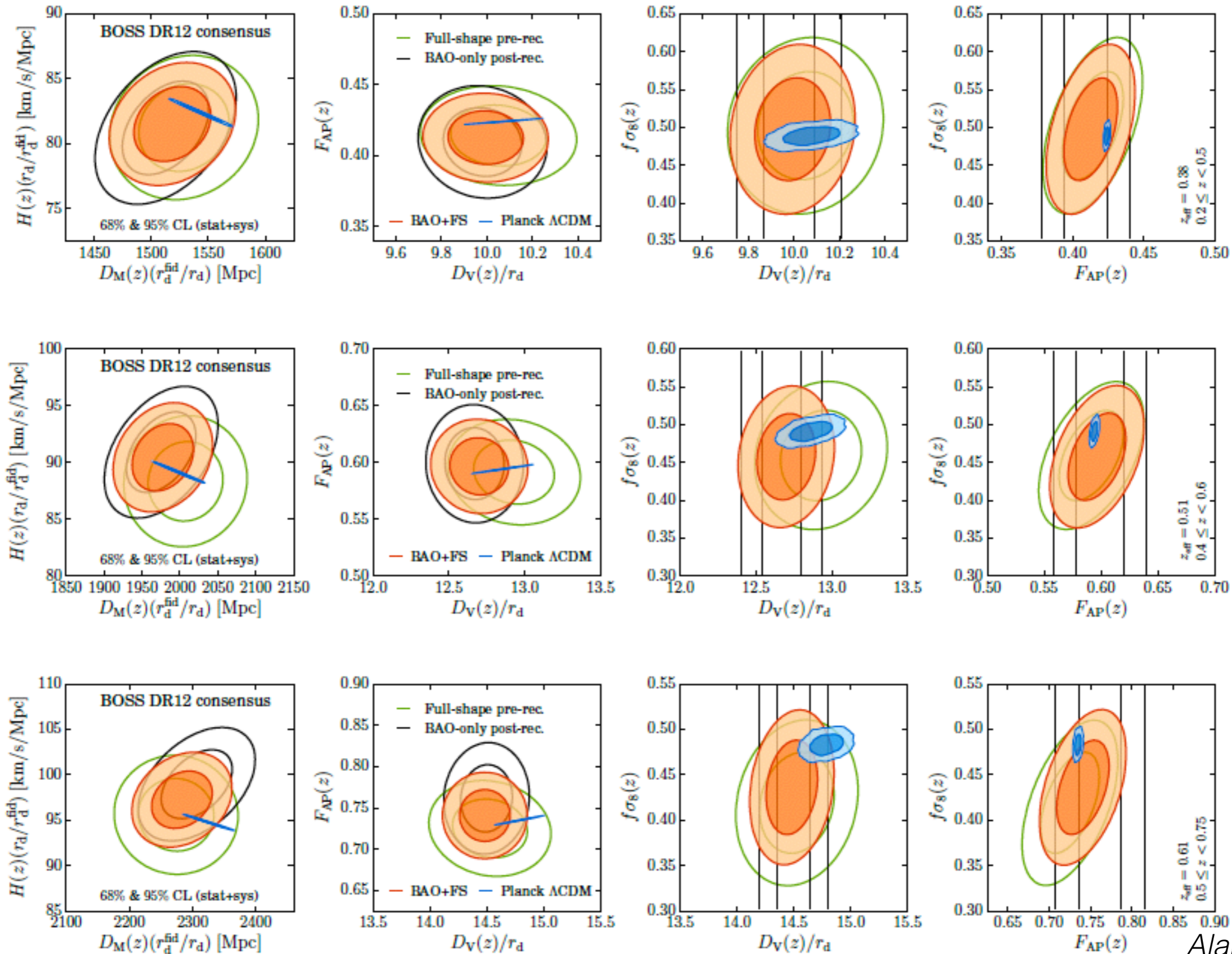
Alam et al. 2016

BAO post-recon

RSD Full Shape pre-recon

Consensus

Planck

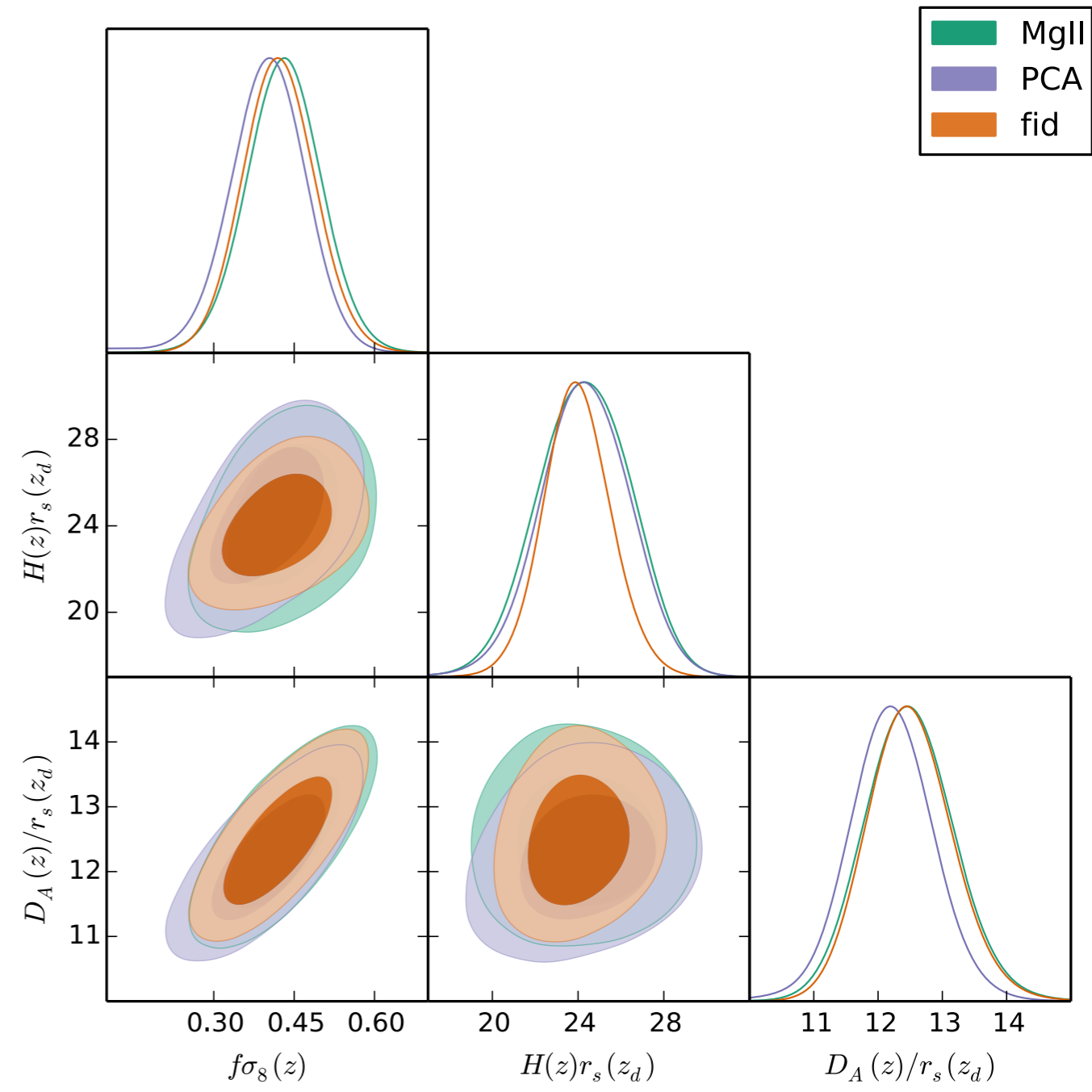
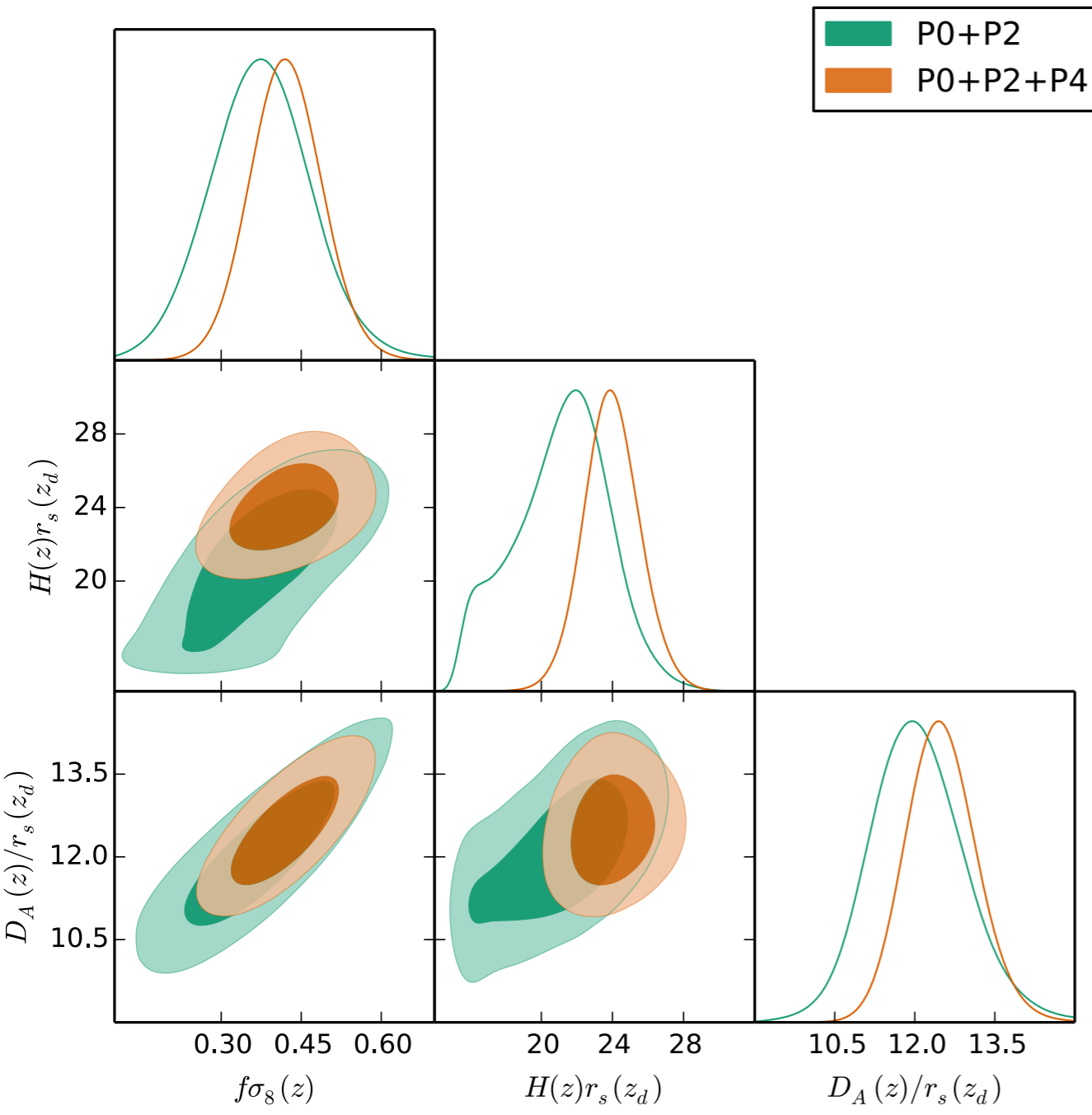


0.2 < z < 0.5

0.4 < z < 0.6

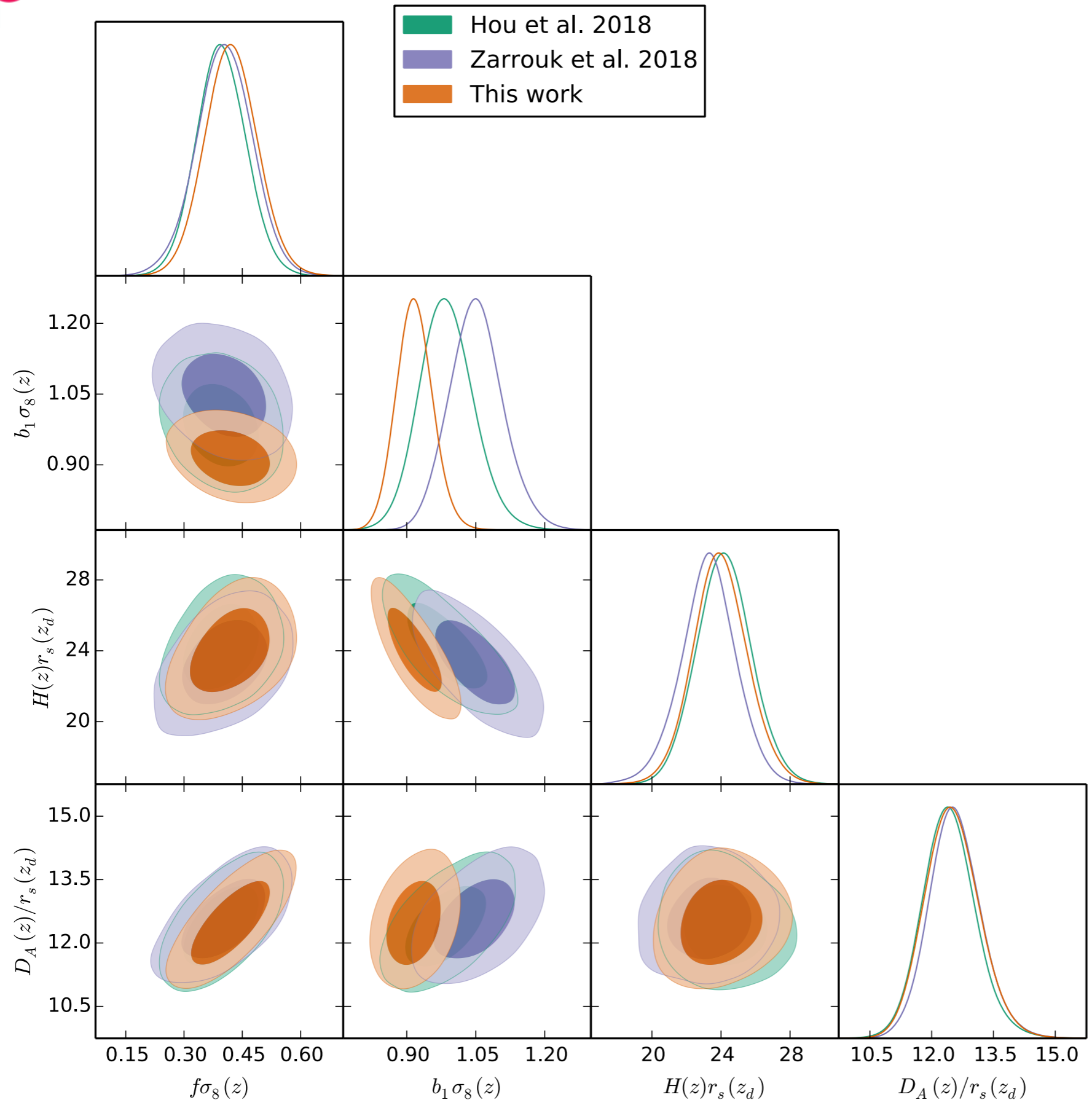
0.5 < z < 0.75

Alam et al. 2016



*HGM et al. 2018*





*HGM et al. 2018*