# Neutrino-Nucleus Interactions and Oscillations

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# **Physics Beyond the Standard Model**

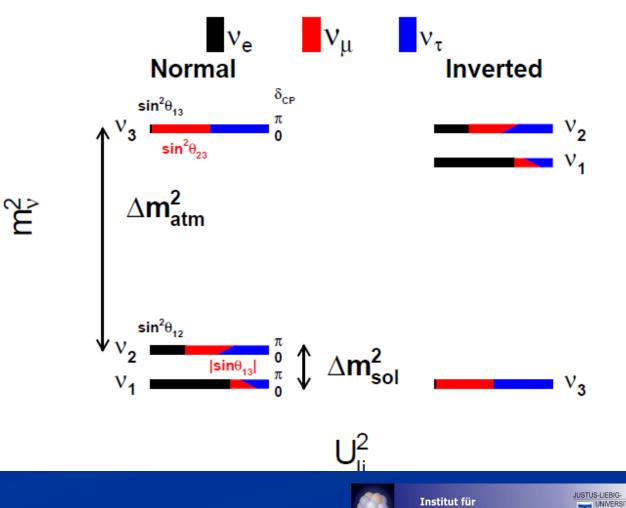
- Precision Physics at CERN LHC: 6.5 TeV protons, beam energy known within 0.1 % → No evidence for BSM physics!
- Beams with broad energy distributions (> 100%) found
   evidence for BSM physics: neutrino oscillations -> neutrinos are massive
   2005 Nobel Prize





# **Neutrino Hierarchy**

Know all mixing angles, not very precise Do not know CP violating phase Do not know mass hierarchy



Theoretische Physik

#### The impossible experiment:

- Beam lines are a few hundred kilometers long
- Beam energies are wide from a few 100 MeV 30 GeV, distribution not too well known
- Beam is wide: about I m at its source, km at the target
- Beam composition is not precisely known





#### Long-Baseline Experiment: T2K and NOvA



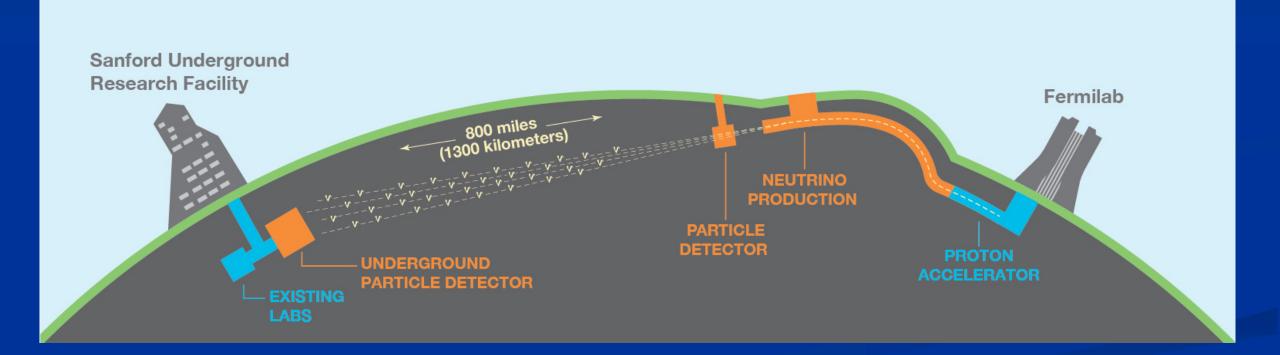




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#### Future (2027): DUNE, joint CERN-FNAL 1.5 B\$ project





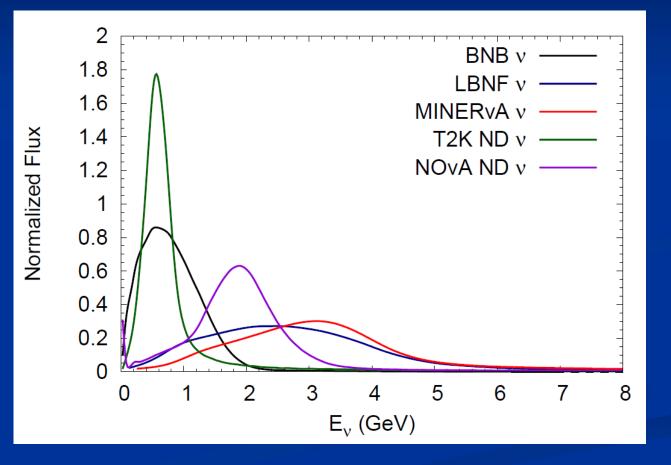


#### **Oscillations and Neutrino Energy**

#### **PROBLEM:**

Neutrinos are produced as secondary decay products of high-energy pA collisions

→ They have broad energy distributions
 Difference to any other high-energy and nuclear physics experiment!
 LHC: ΔE / E ~ 0.1 %







#### **Neutrino-Oscillations**

Simplified: 2 Flavors only

$$P(
u_{\mu} 
ightarrow 
u_{e}) = \sin^{2} 2\theta \sin^{2} \left( rac{\Delta m^{2} L}{4E_{
u}} 
ight)$$

 Energy must be reconstructed from hadronic final state, observed in less-than-perfect detectors
 → Compute backwards from final state to incoming neutrino Reaction mechanism must be known for reconstruction:
 Nuclear Physics is essential, because targets are nuclei: C, O, Ar



 $\Delta m^2 L$  $P(\nu_{\mu} \rightarrow \nu_{e}) = \sin^{2} 2\theta \sin^{2} \theta$  $4E_{\nu}$ b a 0.15 0.15  $v_{\mu}$  flux (AU)  $v_{\mu}$  flux (AU)  $\delta_{CP} = 0^{\circ}, \text{NH}$  $\delta_{CP} = 0^{\circ}$ , NH  $\delta_{CP} = 0^{\circ}, IH$  $\delta_{CP} = 0^{\circ}$ , IH  $P(\nu_{\mu} 
ightarrow \nu_{e})$  $\delta_{CP} = 90^{\circ}, \text{NH}$  $\delta_{CP} = 90^\circ, \text{NH}$ 0.10 ve) 0.10  $\delta_{CP} = 270^\circ$ , NH  $\delta_{CP} = 270^{\circ}, \text{NH}$  $\mathsf{P}(v_{\mu})$ 0.05 0.05 0.00 0.00 10 0.5 1.5 2.0 2.5 1.0 3.0 2 4 6 8 0.0 E<sub>v</sub> (GeV)  $E_v$  (GeV)

From: Diwan et al, Ann. Rev. Nucl. Part. Sci 66 (2016)

#### DUNE, 1300 km

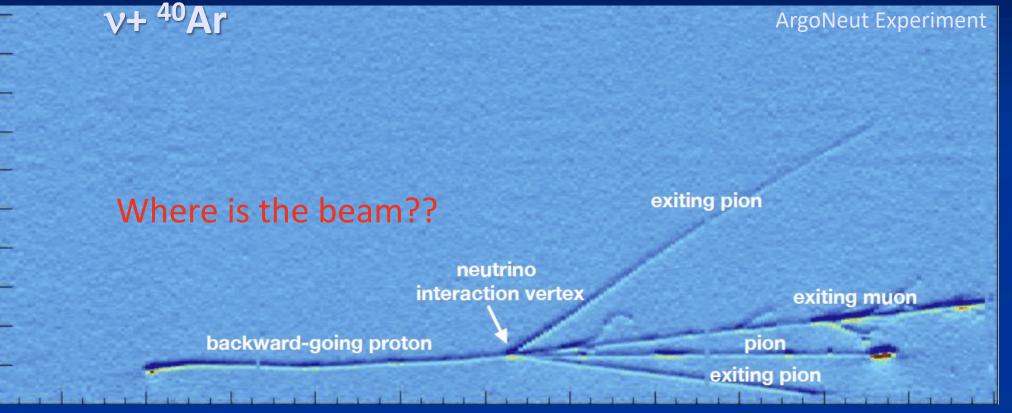
HyperK (T2K) 295 km

Energies have to be known within 100 MeV (DUNE) or 50 MeV (T2K) Ratios of event rates to about 20%





### **Neutrinos on Nuclei**



What is the ingoing state? Composition? Energy? Must reconstruct from final state!



# **Neutrino Cross Sections: Nucleus**

All targets in long-baseline experiments are nuclei: C, O, Ar, Fe Cross sections on the nucleus: QE + final state interactions (fsi) Resonance-Pion Production + fsi • Deep Inelastic Scattering  $\rightarrow$  Pions + fsi <u>Additional cross section on the nucleus:</u> Many-body effects, e.g., 2p-2h excitations Coherent neutrino scattering and coh. pion production 







GiBUU was constructed with the aim to encode the "best possible" theory: gibuu.hepforge.org
 "BEST POSSIBLE" requires
 All neutrino energies, -> relativistic from outset, includes resonances and DIS

- All targets
- Not just inclusive X-sections, but full events
- Reasonable bound nuclear ground states





#### **Quantum-kinetic Transport Theory** for FSI Off-shell transport term **Collision term On-shell drift term** $\mathcal{D}F(x,p) - \operatorname{tr}\left\{\Gamma f, \operatorname{Re}S^{\operatorname{ret}}(x,p)\right\}_{\operatorname{PB}} = C(x,p) \ .$ $\mathcal{D}F(x,p) = \{p_0 - H, F\}_{\rm PB} = \frac{\partial(p_0 - H)}{\partial x} \frac{\partial F}{\partial p} - \frac{\partial(p_0 - H)}{\partial p} \frac{\partial F}{\partial x}$ H contains mean-field potentials Describes time-evolution of F(x,p)— Spectral function $F(x,p) = 2\pi g f(x,p) \mathcal{P}(x,p)$ Phase space distribution

Kadanoff-Baym equations with BM offshell term







anti v

-0.65

-0.75

-0.85

-0.95

0.5

QE

2p2h

tot —

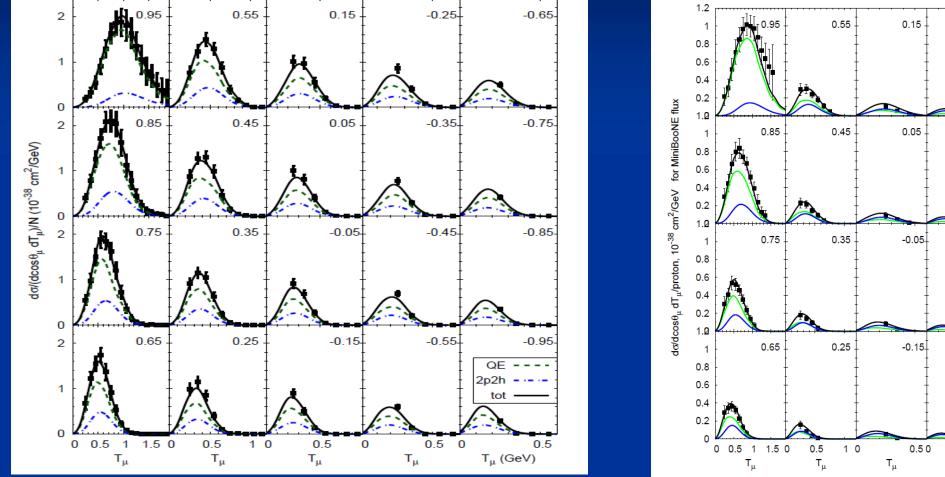
 $\mathsf{T}_{\mu}$ 

-0.25-

-0.35

-0.45

-0.55



ν

GiBUU 2016: no data adjustment

Bormio 01/2018

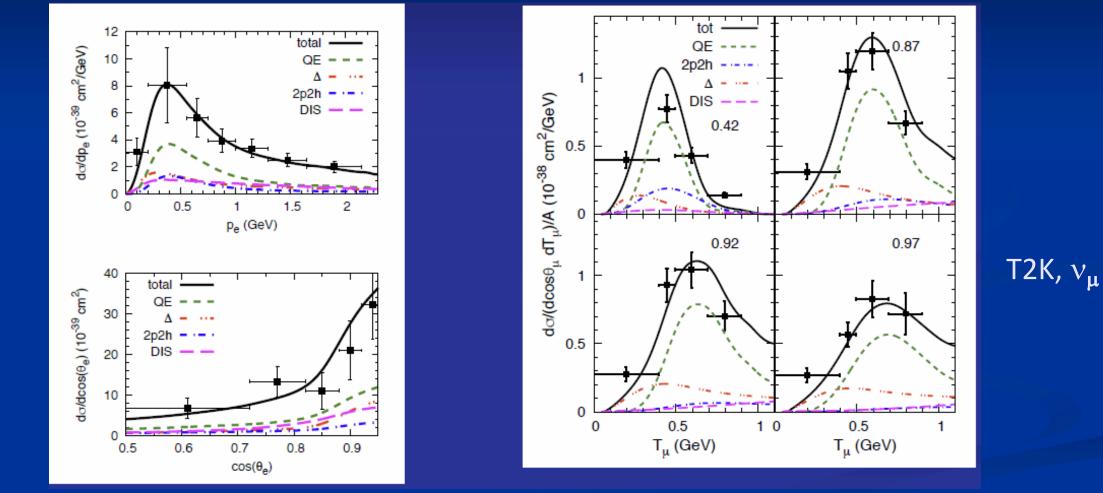


0.50

Tμ



# **Comparison with T2K incl. Data**



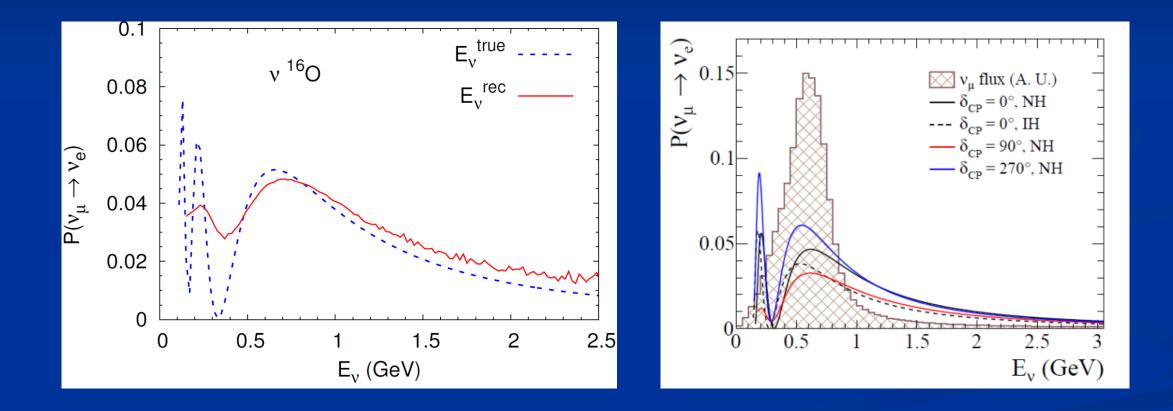
T2Κ, ν<sub>e</sub>

Agreement for different neutrino flavors





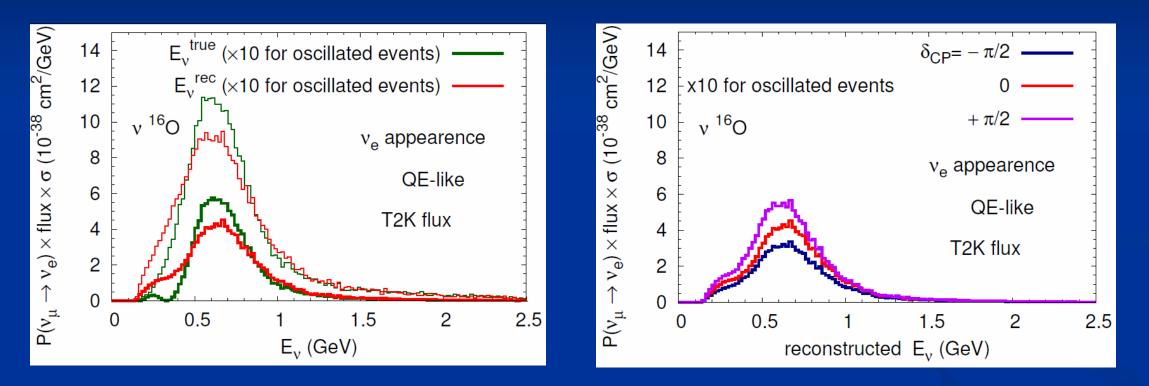
#### **Reconstruction in T2K**







# **Oscillation signal in T2K** $\delta_{CP}$ sensitivity of appearance exps

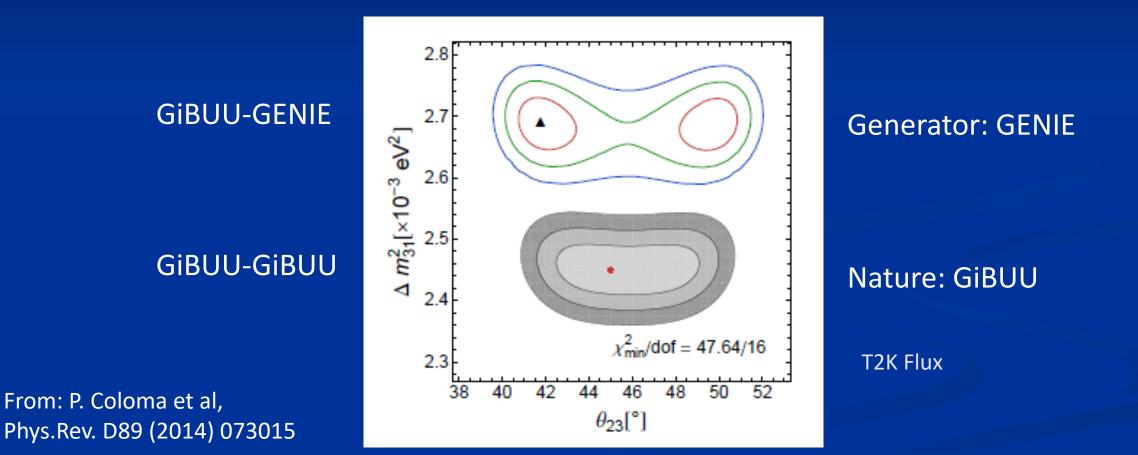


Uncertainties due to energy reconstruction as large as  $\delta_{\text{CP}}$  dependence





# Generator Dependence of Oscillation Parameters







# Summary

- Extraction of neutrino properties requires knowledge of neutrino energy to about 5% accuracy.
- In long-baseline experiments the incoming neutrino energy must be reconstructed from final state. Only partially known because detectors are less-than-perfect.
- Backwards calculation from this partially known final state requires command both of initial neutrino-nucleus reactions and of hadronic final state interactions
- Present models can do this to about 10- 20 %  $\rightarrow$  not good enough
- Precision neutrino long-baseline physics requires better state-of-the-art generators
- GiBUU is one such attempt
- BACK to PRECISION PHYSICS: not so much for experiment, but for theory







# **GiBUU:** References

#### Essential References:

- I. Buss et al, Phys. Rept. 512 (2012) I contains both the theory and the practical implementation of transport theory
- 2. Gallmeister et al., Phys.Rev. C94 (2016), 035502 contains the latest changes in GiBUU2016
- 3. Mosel, Ann. Rev. Nucl. Part. Sci. 66 (2016) 171 review, contains some discussion of generators
- 4. Mosel et al, Phys.Rev. C96 (2017) no.1, 015503 pion production comparison of MiniBooNE, T2K and MINERvA



