

Electroweak cosmology

Thomas Konstandin

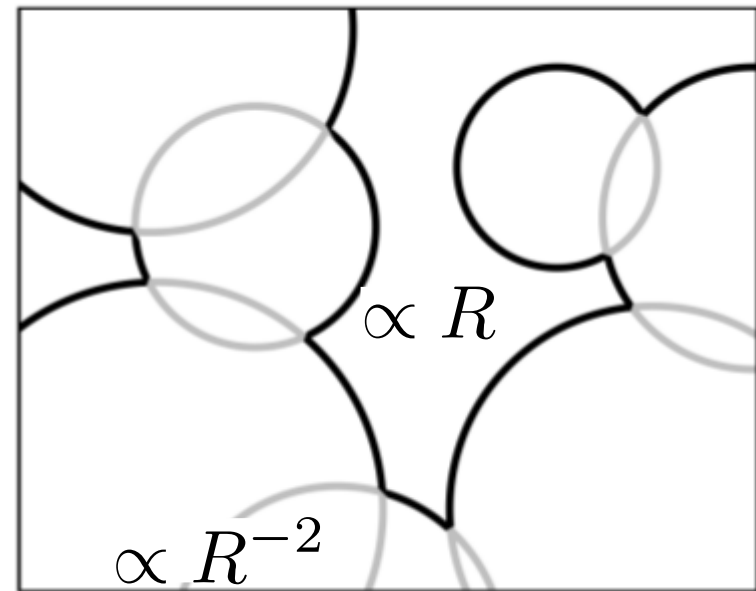
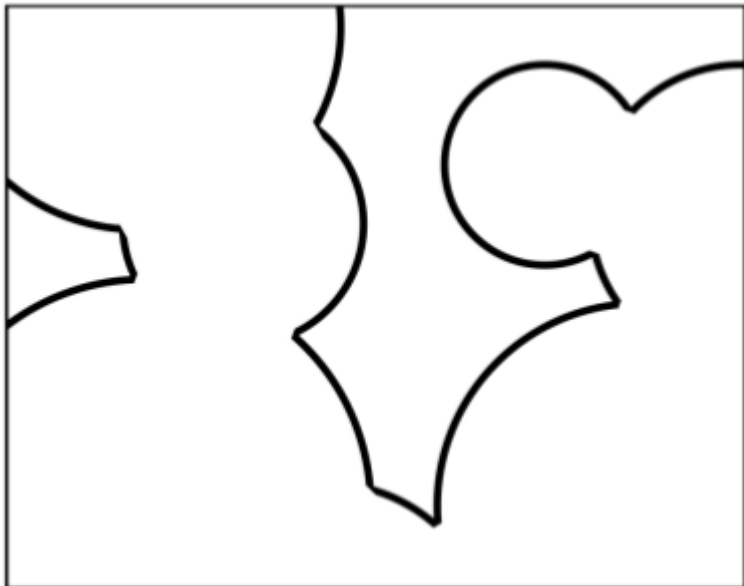


LISA WG, MITP Mainz, Oct 17, 2017

Motivation

Gravitational waves from bubble dynamics: Beyond the Envelope

Ryusuke Jinno^{a,b} and Masahiro Takimoto^{b,c}



Motivation

3.2 Analytic expressions

Now we present the analytic expressions for the GW spectrum. After a short calculation (see Appendix A–C), we obtain the single-bubble spectrum

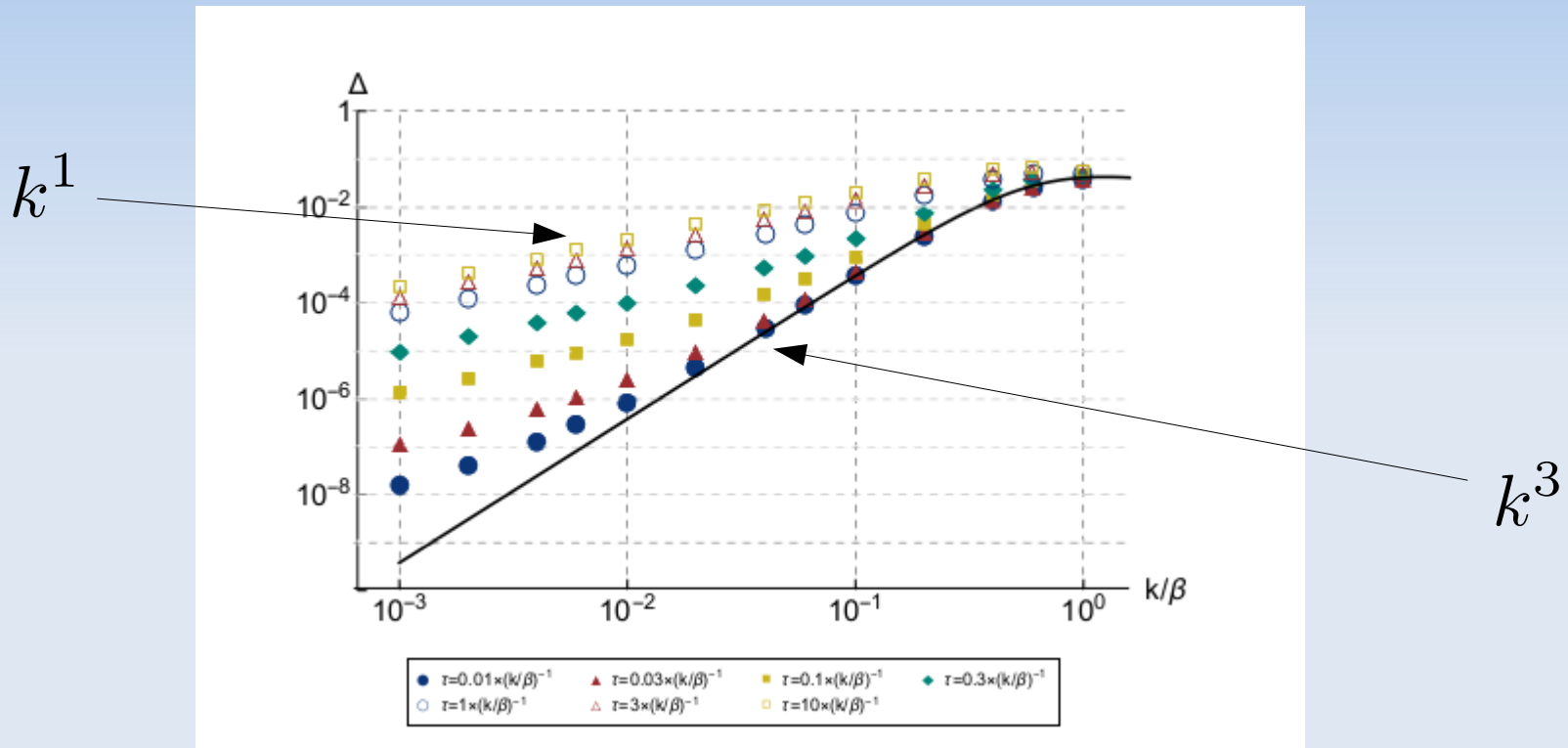
$$\Delta^{(s)} = \int_{-\infty}^{\infty} dt_x \int_{-\infty}^{\infty} dt_y \int_{v|t_x, y|}^{\infty} dr \int_{-\infty}^{t_{\max}} dt_n \int_{t_n}^{t_x} dt_{xi} \int_{t_n}^{t_y} dt_{yi} \left[\begin{aligned} & e^{-I(x_i, y_i)} \Gamma(t_n) \frac{r}{r_{xn}^{(s)} r_{yn}^{(s)}} \\ & \times \left[j_0(kr) \mathcal{K}_0(n_{xn}, n_{yn}) + \frac{j_1(kr)}{kr} \mathcal{K}_1(n_{xn}, n_{yn}) + \frac{j_2(kr)}{(kr)^2} \mathcal{K}_2(n_{xn}, n_{yn}) \right] \\ & \times \partial_{t_{xi}} [r_B(t_{xi}, t_n)^3 D(t_x, t_{xi})] \partial_{t_{yi}} [r_B(t_{yi}, t_n)^3 D(t_y, t_{yi})] \cos(kt_{x,y}) \end{aligned} \right], \quad (3.2)$$

and the double-bubble spectrum

$$\Delta^{(d)} = \int_{-\infty}^{\infty} dt_x \int_{-\infty}^{\infty} dt_y \int_0^{\infty} dr \int_{-\infty}^{t_x} dt_{xn} \int_{-\infty}^{t_y} dt_{yn} \int_{t_{xn}}^{t_x} dt_{xi} \int_{t_{yn}}^{t_y} dt_{yi} \int_{-1}^1 dc_{xn} \int_{-1}^1 dc_{yn} \int_0^{2\pi} d\phi_{xn, yn} \left[\begin{aligned} & \Theta_{\text{sp}}(x_i, y_n) \Theta_{\text{sp}}(x_n, y_i) e^{-I(x_i, y_i)} \Gamma(t_{xn}) \Gamma(t_{yn}) \\ & \times r^2 \left[j_0(kr) \mathcal{K}_0(n_{xn}, n_{yn}) + \frac{j_1(kr)}{kr} \mathcal{K}_1(n_{xn}, n_{yn}) + \frac{j_2(kr)}{(kr)^2} \mathcal{K}_2(n_{xn}, n_{yn}) \right] \\ & \times \partial_{t_{xi}} [r_B(t_{xi}, t_{xn})^3 D(t_x, t_{xi})] \partial_{t_{yi}} [r_B(t_{yi}, t_{yn})^3 D(t_y, t_{yi})] \cos(kt_{x,y}) \end{aligned} \right]. \quad (3.3)$$

10 integrations?

Motivation



Motivation

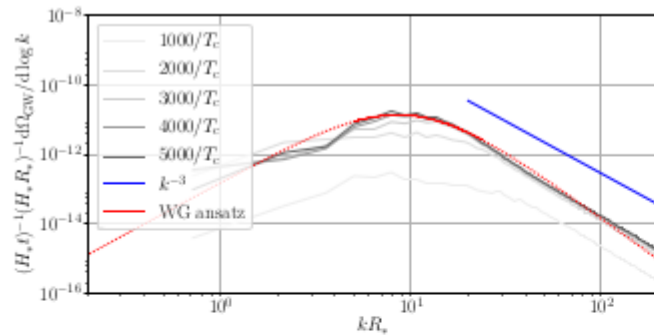
Shape of the acoustic gravitational wave power spectrum from a first order phase transition

Mark Hindmarsh,^{1,2,*} Stephan J. Huber,^{1,†} Kari Rummukainen,^{2,‡} and David J. Weir^{2,§}

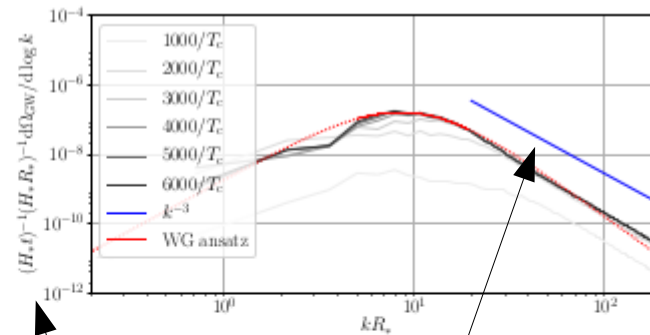
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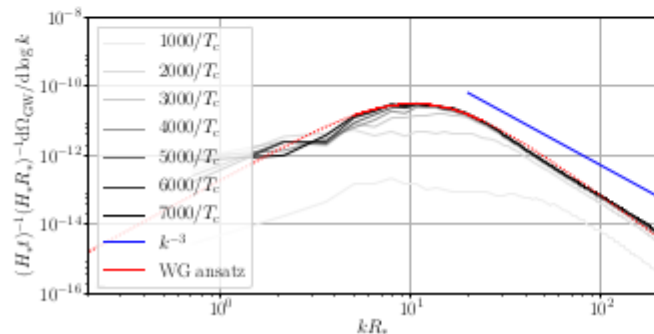
(Dated: April 20, 2017)



(a) Weak, $v_w = 0.92$



(b) Intermediate, $v_w = 0.92$



enhancement β/H

k^{-3}

Envelope redux

Gravitational Wave Production by Collisions: More Bubbles

Stephan J. Huber⁽¹⁾ and Thomas Konstandin^{(2)*}

$$\frac{dE_{GW}}{d\omega d\Omega} = 2G\omega^2 \Lambda_{ij,lm}(\hat{\mathbf{k}}) T_{ij}^*(\hat{\mathbf{k}}, \omega) T_{lm}(\hat{\mathbf{k}}, \omega),$$

where $T_{ij}(\hat{\mathbf{k}}, \omega)$ denotes the stress-energy tensor in Fourier space

$$T_{ij}(\hat{\mathbf{k}}, \omega) = \frac{1}{2\pi} \int dt e^{i\omega t} \int d^3x e^{-i\omega \hat{\mathbf{k}} \cdot \mathbf{x}} T_{ij}(\mathbf{x}, t),$$

and Λ is the projection tensor for the transverse-traceless part

$$\Lambda_{ij,lm}(\hat{\mathbf{k}}) = \delta_{il}\delta_{jm} - 2\hat{k}_j\hat{k}_m\delta_{il} + \frac{1}{2}\hat{k}_i\hat{k}_j\hat{k}_l\hat{k}_m - \frac{1}{2}\delta_{ij}\delta_{lm} + \frac{1}{2}\delta_{ij}\hat{k}_l\hat{k}_m + \frac{1}{2}\delta_{lm}\hat{k}_i\hat{k}_j.$$

$$\Omega_{GW*} = \omega \frac{dE_{GW}}{d\omega} \frac{1}{E_{tot}} = \kappa^2 \left(\frac{H}{\beta}\right)^2 \left(\frac{\alpha}{\alpha+1}\right)^2 \Delta(\omega/\beta, v_b), \quad (13)$$

where we used the definition of the Hubble constant

$$H^2 = \frac{8\pi G \rho_{tot}}{3} = \frac{8\pi G (\rho_{vac} + \rho_{rad})}{3}, \quad (14)$$

and defined the dimensionless function Δ as

$$\Delta(\omega/\beta, v_b) = \frac{\omega^3}{\beta^3} \frac{3v_b^6 \beta^5}{2\pi V} \int d\hat{\mathbf{k}} \Lambda_{ij,lm} C_{ij}^* C_{lm}. \quad (15)$$

No dependence on H!

Integrations

$$\frac{dE_{GW}}{d\omega d\Omega} = 4G\rho_{vac}^2 \kappa^2 v_b^6 \omega^2 (|C_+|^2 + |C_-|^2), \quad (A2)$$

$$C_{\pm}(\omega) = \frac{1}{6\pi} \sum_n \int dt e^{i\omega(t-z_n)} (t - t_n)^3 A_{n,\pm}(\omega, t), \quad (A3)$$

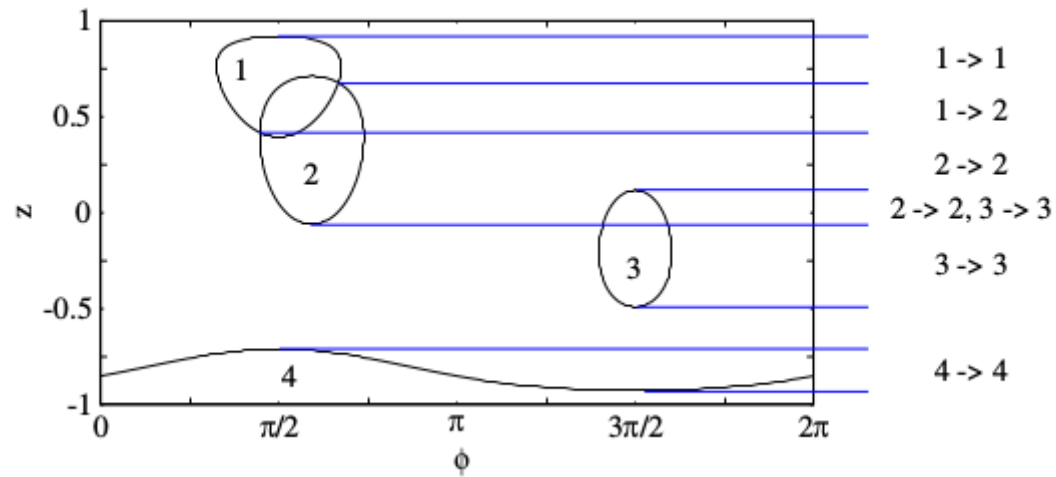
$$A_{n,\pm}(\omega, t) = \int_{-1}^1 dz e^{-iv_b \omega(t-t_n)z} B_{n,\pm}(z, t), \quad (A4)$$

$$B_{n,+}(z, t) = \frac{(1-z^2)}{2} \int_{S_n} d\phi \cos(2\phi), \quad B_{n,-}(z, t) = \frac{(1-z^2)}{2} \int_{S_n} d\phi \sin(2\phi). \quad (A5)$$

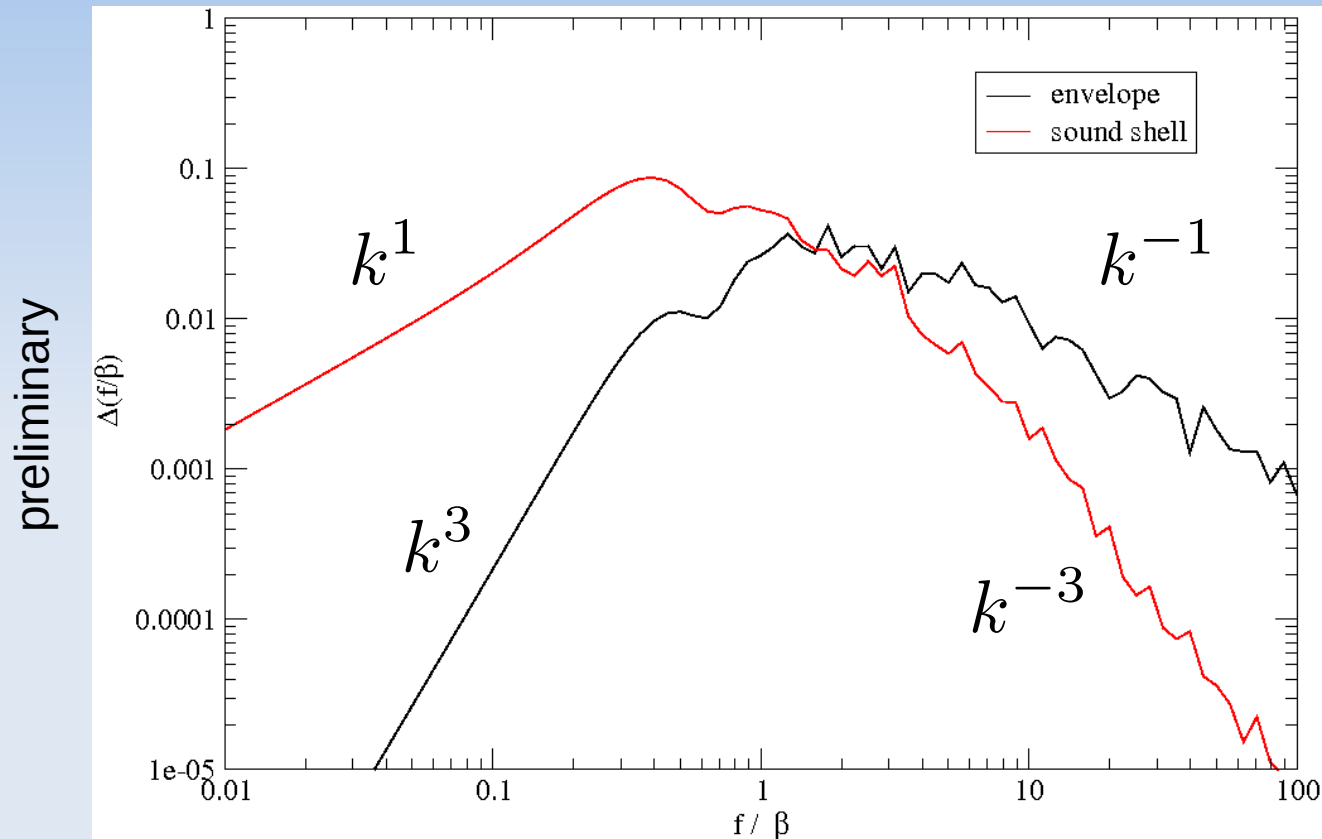
3 integrations, done 10^{11} times, 1 CPU hour

do the first integration analytically!

Integrations



Results



all integrals are finite

peak a little lower

amplitude a bit higher